

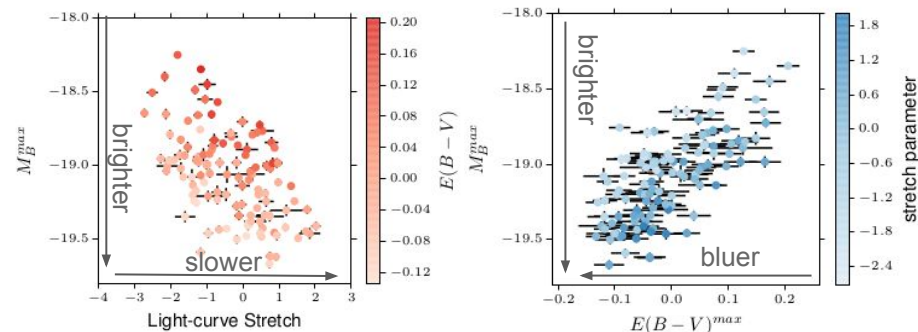
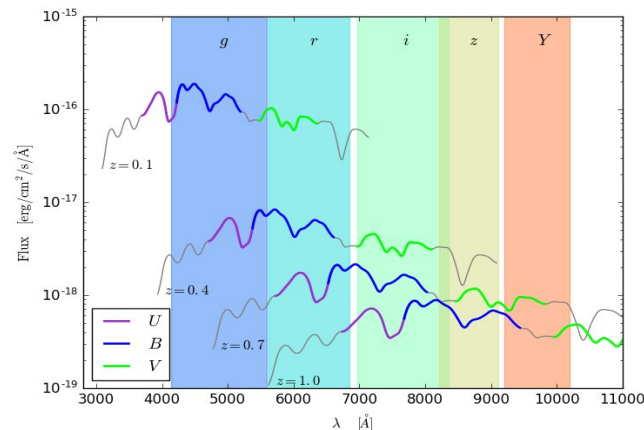
NaCl : Nouveaux algorithmes de Courbes de lumière

Guy Augarde, Nicolas Regnault

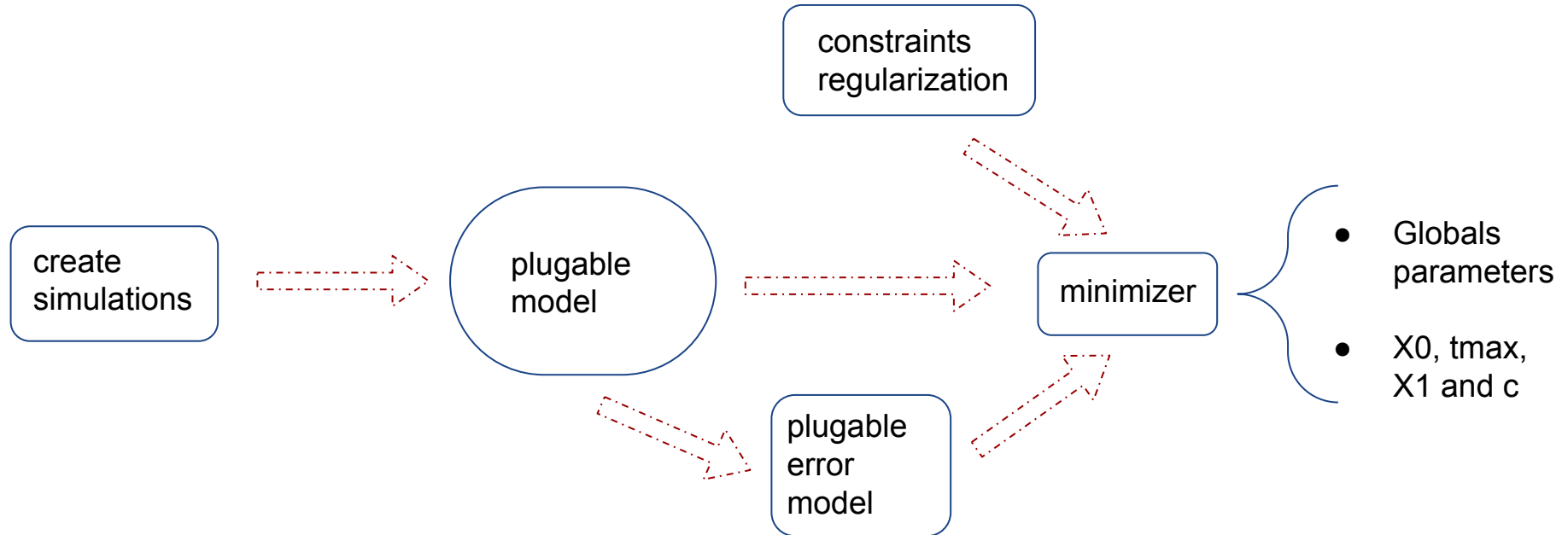
Light curve fitter

- To construct a Hubble Diagram, every magnitude must be expressed in the same restframe band : band B (by convention) at peak : m_B^*
- To minimize the dispersion the Hubble diagram residuals : extraction of parameter link to the stretch, x_1 , and color, c .

$$\mu = m_B^* - M_B + \alpha X_1 - \beta c \pm 15\%$$

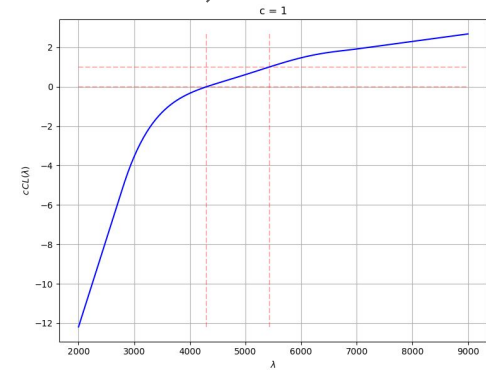
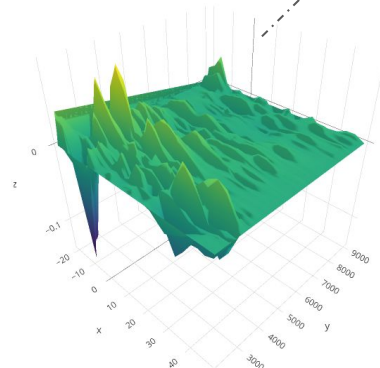
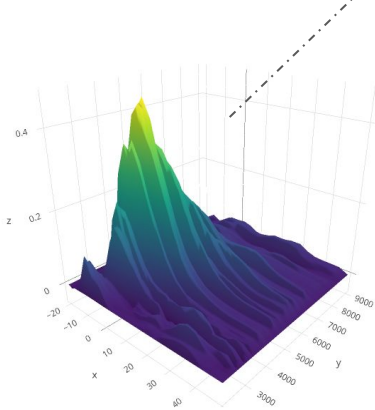


- Why a new generation ?
 - Number of SN is much higher :
 - ✧ scaling code
 - ✧ fast training procedure
 - Clear propagation of systematics :
 - ✧ Propagate the systematics on the time of maximum
 - ✧ Propagate the calibration uncertainties
 - Construction of a flexible pipeline in order to test new modelization
 - Study of model bias



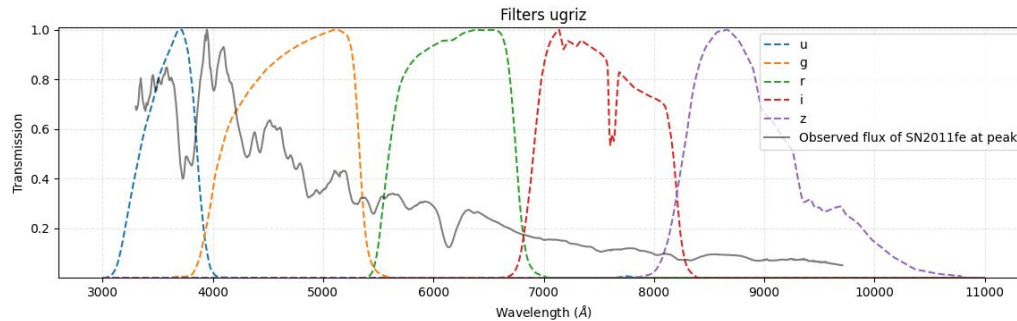
NaCl : Spectrophotometric model

$$S^{SN}(\lambda, p) = [M_0(p, \lambda) + X_1 M_1(p, \lambda)] e^{cCL(\lambda)}$$

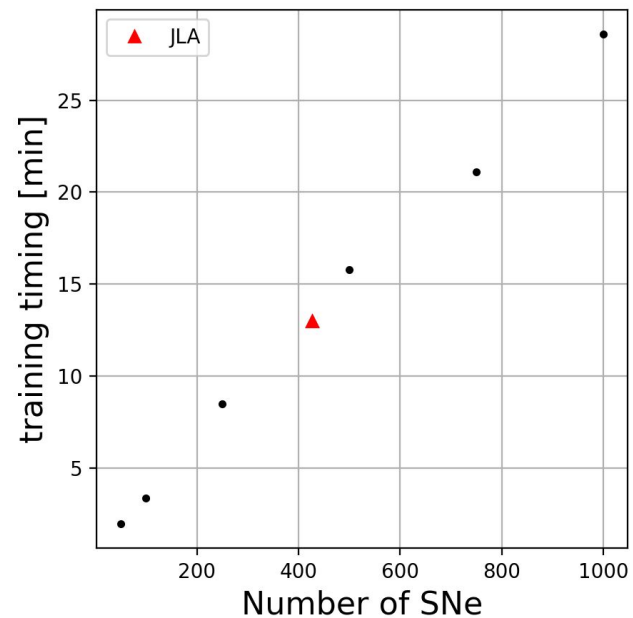
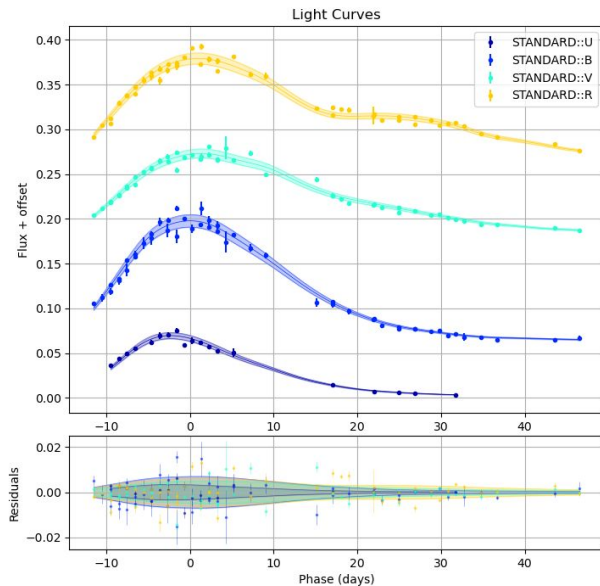
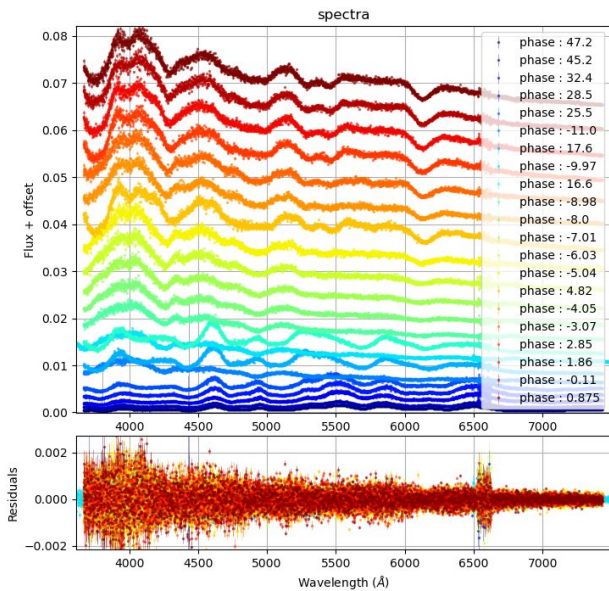


NaCl : Spectrophotometric model

$$\phi_X^{SN}(p) = \int X_0 S^{SN}(\lambda, p) T_X \left(\frac{\lambda}{1+z} \right) \frac{\lambda}{hc} d\lambda$$



sn1999dq z : 0.0143

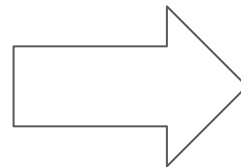


- The primary flux reference.
- The transfer of the primary calibrators to the secondary standards.
- The error measurements of the CALSPEC standards.
- The transfer from the CALSPEC standards to the tertiary standard

Numerical estimation
of the Jacobian of light curves
parameters wrt error on ZP

Photometric calibration
uncertainties matrix

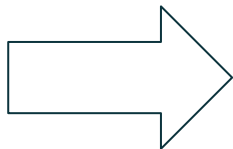
$$\mathbf{C}_{\text{cal}} = \mathbf{J} \mathbf{C}_{\eta} \mathbf{J}^{\dagger}$$



- add one latent parameter per survey band;
- dispersion given the intercalibration matrix.

$$\phi_X^{SN}(p) = \int \mathbf{X}_0 S^{SN}(\lambda, p) T_X \left(\frac{\lambda}{1+z} \right) \frac{\lambda}{hc} (1 + \eta_X) d\lambda$$

- broad-band colour uncertainties;
- Coherent offset between different points of a same light curve;
- No correlation between measurements in different bandpasses.

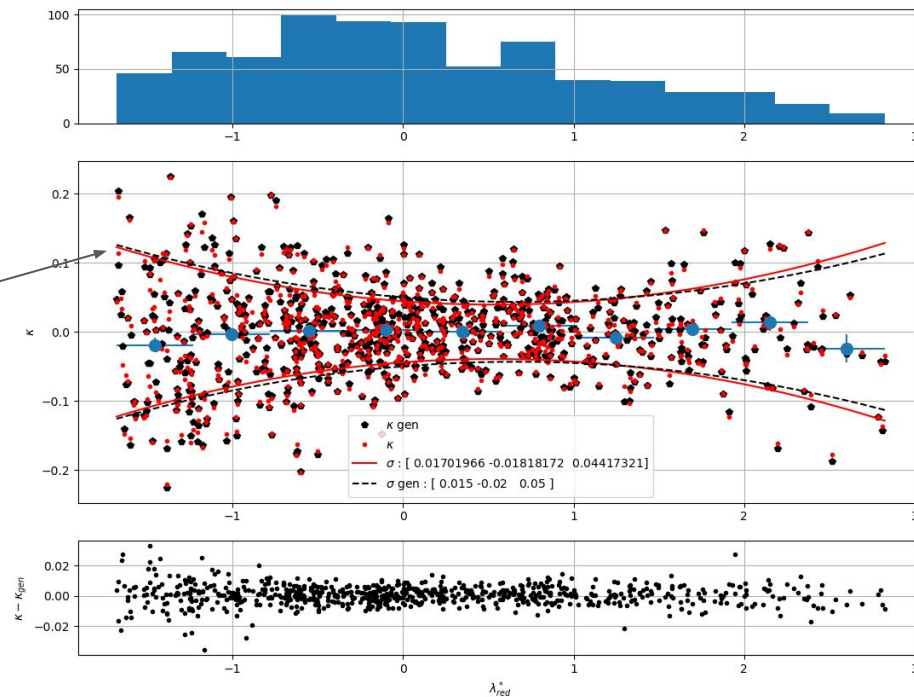


- adding one latent parameter to each light curve κ
- Hold by a gaussian prior during the minimization:
$$\kappa \sim N(0, \sigma_{\kappa}(\lambda))$$
- Estimation of the variance of the prior during the minimization

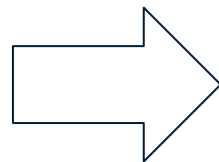
$$\phi_X^{SN}(p) = \int X_0 S^{SN}(\lambda, p) T_X \left(\frac{\lambda}{1+z} \right) \frac{\lambda}{hc} (1 + \eta_X)(1 + \kappa) d\lambda$$

Color Scatter reconstruction

$$\sigma_{\kappa}(\lambda_c) = \sum_{i=0}^{N_{\sigma}} \sigma_i \lambda^{N_{\sigma}-i}$$



- 100 realizations
- same sampling
- same S_{ne} parameters
- Varying Noise, calibration term, color scatter term and additional variability



- Model
 - Error model
 - without Color Scatter
 - Propagation of calibration uncertainties
- We keep light curves parameters

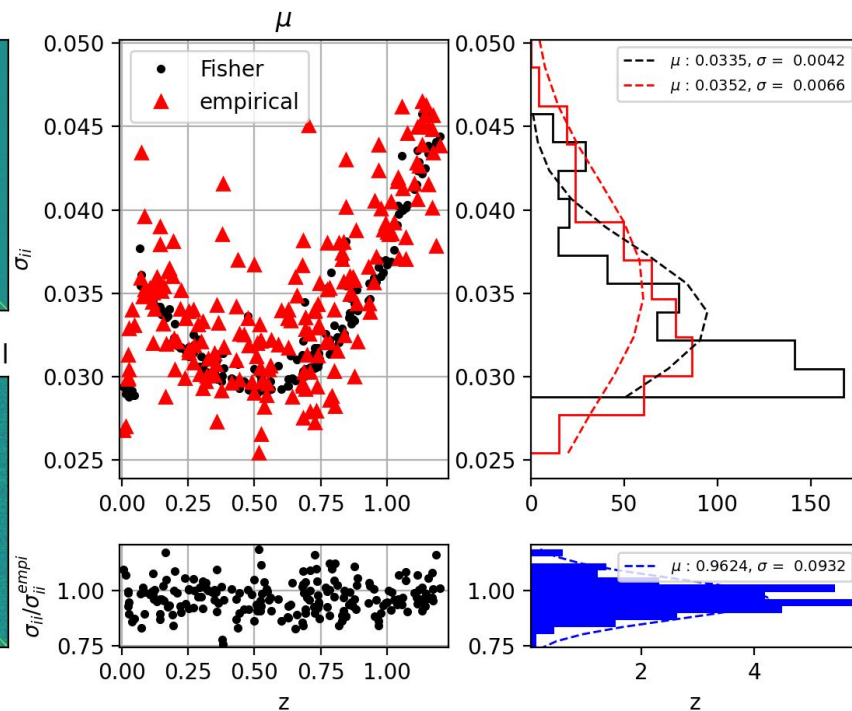
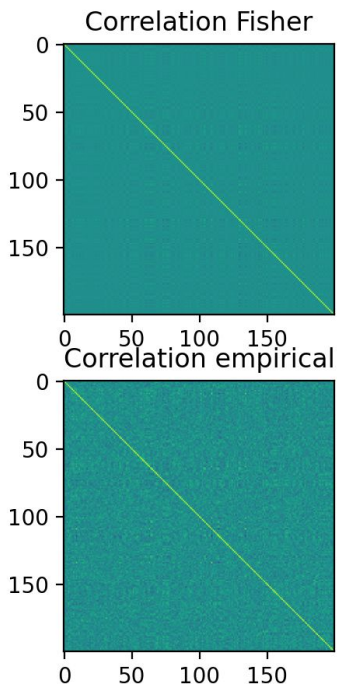
$$\mu = -2.5 \log(x_0) + \alpha x_1 - \beta c - M_B$$

Uncertainties on distances

Fisher Matrix

$$Cov = \frac{\partial \mu}{\partial p} H^{-1} \frac{\partial \mu}{\partial p}^T$$

$$Cov_{empi} = \frac{(\mu - \bar{\mu})^T (\mu - \bar{\mu})}{N - 1}$$



Conclusion & roadmap

Fast full-fledged minimization procedure :

- Fit time maximum luminosity;
 - SN intrinsic residuals variability model;
 - Single-step comprehensive fit;
 - Propagation of photometric calibration uncertainties;
 - Bias study on simulations.
- Release of the full training code, where the model is interchangeable to investigate new standardization parameters