

CEVO Task 4.2 progress report Since last Tech Forum

F.Bonnarel (CDS) as coordinator











Task 4.2 goals

- Implementation of FAIR principles for ESFRI data through the Virtual Observatory
- Push ESCAPE priorities at IVOA level
 - Help ESCAPE ESFRIs to use VO standards and expose the data
 - ESFRI feedback on VO standards and evolution of those.
- Community training events for scientists and data producers/providers
 - Vo schools
 - Data provider formats

• ASTRON : LOTSS DR2

(see tomorrow Yan Grange)

• JIVE : JIVE visibility data service release

(see next screenshots)

JIVE ObsTAP service with DataLink

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- ALMA : SIAV2 and ObsTAP services (see next screenshots)
- SKAO : preparing for the SRC network / strong integration of VO standards in the design.
- \rightarrow How to integrate ivoa Provenance + ...

ALMA SIA2 service

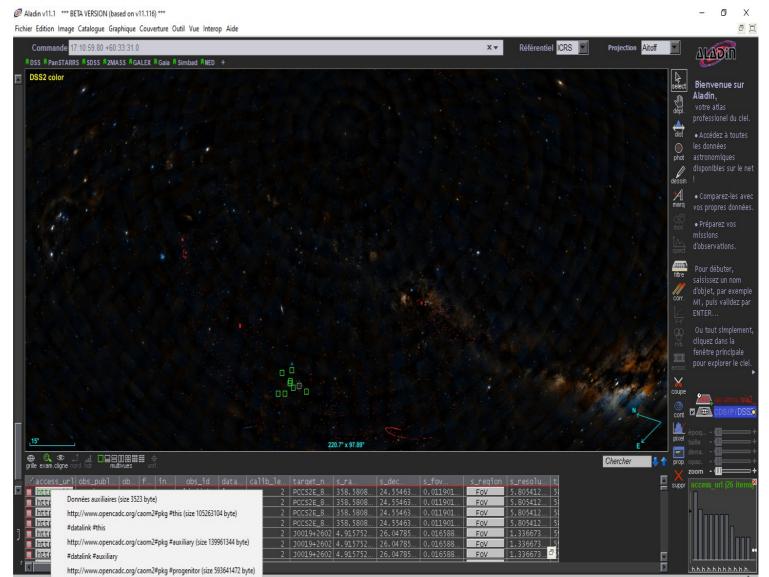
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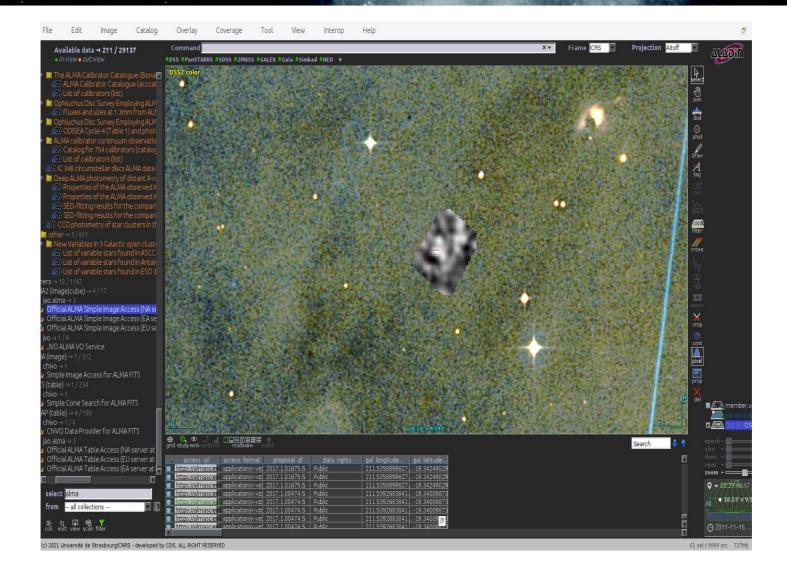
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ALMA DataLink



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ALMA loaded moment image



- All radio ESFRI projects + INAF radio archive + Nançay team participate a lot to IVOA radio interest group
 - Implementation note
 - Pulsar and FRB data + Single dish running meetings (Hack a thon SD metadata/ObsCore tomorrow)
 - ObsCore extension for visibility data



DRAFT - please do not distribute



IVOA Obscore Extension for Visibility data

Version 1.0

IVOA Note 2021-10-27

Working group

Data Model Working Group

This version

http://www.ivoa.net/documents/ObsCoreExtensionForVisibilityData/20211027

Latest version

http://www.ivoa.net/documents/ObsCoreExtensionForVisibilityData

Previous versions

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Editor(s)

François Bonnarel

Abstract

This is a proposed extension to the Obscore specification for description of visibility data

Status of this document

This is an IVOA Note expressing suggestions from and opinions of the authors. It is intended to share best practices, possible approaches, or other



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4.2 uv parameters

uv_distance_min and uv_distance_max are evaluated by fitting an ellipse on the visibilities present in the uv plane.

To compute the ellipse's eccentricity of the UV distribution a principal component analysis (PCA) with 2 components is performed over the data points sampling the UV plane to select the main axis of data scattering. The first component is used to rotate the distribution of UV in a way that the major variation of the distribution is leaning towards the x axis of a bi dimensional xy Cartesian plane. The major axis length and the minor axis length of the ellipse are therefore defined as the semi distance between the most positive point along the x/y axis and the most negative point among the y axis. For instance, if the range of the rotated UV will cover on the $x \in [-10, 10]$ the major axis distance would be 10, a similar procedure is done on the y axis. This procedure allows the definition of the UV distribution eccentricity:

uv distribution exc) computed as follows:

$$uv_distribution_exc = \sqrt{1 - \frac{b^2}{a^2}}$$

(1)

where a is the major axis length and b is the minor axis length. The filling factor of the UV plane (hereafter uv_distribution_fill) is computed as the average number of samples found in a $N^{uv}_{samples} \times N^{we}_{samples}$ equispaced grid enclosing the rotated ellipse. In formulas, the boundaries of a cell (i,j) are defined by the boundaries

$$u \in [u_{min} + \frac{u_{max} - u_{min}}{N_{samples}^{uv} + i, u_{min} + \frac{u_{max} - u_{min}}{N_{samples}^{uv}} \cdot (i+1)]$$
(2)

and

$$v \in [v_{min} + \frac{v_{max} - v_{min}}{N_{samples}^{uv}} \cdot j, v_{min} + \frac{v_{max} - v_{min}}{N_{samples}^{uv}} \cdot (j + 1)]$$
 (3)

where u_{max}/v_{max} are the respective maximum u/v of the uv sample and u_{min}/v_{min} is the minimum u/v of the uv sample.

Given the above boundaries the number of samples within a cell (i,j) will be $n_{i,j}^{w}$ and uv_distribution_fill will be then computed as

$$uv_distribution_fill = \frac{\sum_{i=1}^{N_{uv}^{uv} maples} \sum_{j=1}^{N_{uv}^{uv} maples} n_{i,j}^{uv}}{(N_{samples}^{uv})^2},$$
 (4)

in the preliminary analysis $N_{samples}^{uv} = 1000$.

4.3 time parameters

t_exp_min and t_exp_max are added because of strong variation in the individual time stamps duration.



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Wiki Home WebChanges WebTopicList WebStatistics	On pulsar data and other time oriented radio data in the VO													
wiki Meta & Help /OA inow fain aindbox	Organized by Brent Miszalski, Mark Cresitello-Dittmar, (TDIG) Mark Lacy, François Bonnarel (<u>RadiolG</u>) Ada Nebot (as CSP chair)													
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Topic revision: r6 - 2022-01-26 - FrancoisBonnarel

Committees

Stds&Procs

Procs

www.ivoa.net

- CTA + KM3Net
 - Find out use cases for gamma-ray neutrino interoperability and VO exposition
 - Several meetings organized to understand the issues and pave the way to solutions (see next screenshot)
 - Hack a Thon this afternoon at 2 30 PM

25 June 2021	IVOA metadata and High Energy Astrophysics	preliminary Working meeting	face to face meeting in Meudon	Agendaa Minutes by M.Louys and M.Servillat@
9 July 2021	SKAO CEVO provenance meeting	Working meeting	(Virtual)	Agenda@ General ProvSDP introduction by S.Sanchez@ IVOA ProvDM presentation by M.Servillat@ SKAO feedback on IVOA data model by J.Garrido@ Minutes@
1 September 2021	CEVO (CTAO + High energy) datamodels and standards for interoperability	Kickoff meeting	virtual	Agenda Introduction June25thView DL3 datamodel proposal minutes
13. September 2021	HiPS in the Data Lake	ESCAPE event	On-line event	
29 September 2021	ESCAPE General Assembly	Meeting	online	Link
24-28 October 2021	ADASS 2021	Conference	Cape Town and online	Link Ø
02-04 November 2021	IVOA Interop	Conference	Videocon Meeting	On-line
23-26 November 2021	Hands-on workshop for Data Providers	Workshop	On-line (gather.town)	event page 🖗
08 December 2021	CTA/KM3Net CEVO "datamodels for interoperability" workshop	Workshop	on-line	Agenda+Intro@ CTA DL3 draft datamodel@ KM3Net@ GADF@ Legacy work 1@ Legacy Work 2@ IVOA DM/specs for HEA@ minutes@
22-24 February 2022	2nd Science with interoperable data school	Hands-on School	Strasbourg (or on-line if necessary)	Event page Ø

- EGO/VIRGO :
 - visualization using MOC and AladinLite (see Giuseppe Greco talk tomorrow)
- ESO :
 - use of ESO services in Aladin/CASSIS prototype for spectral cubes visualisation by CDS/OMP (see ADASS slides next)

ESO VO interfaces to data

MUSE

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Figure 1. Discovery of MUSE cubes in Orion area with ESO ObsTAP service and SODA interface to obtain a cutout on one of the cubes.

- EST :
 - Mapping solarnet metadata / UCD + a few VEPs
 - 3 EPN-TAP services at ROB : USET sunspots and SPOCA coronal holes.
 - Study for VO integratiojn of SPASE event database

Uset Drawing service (ROB-EST)

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