

# Modeling of High Column Density Systems (HCDs) in the Ly- $\alpha$ Forests

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# Outline

1. Introduction to Lyman- $\alpha$  forests
2. Modeling of HCDs
3. Summary

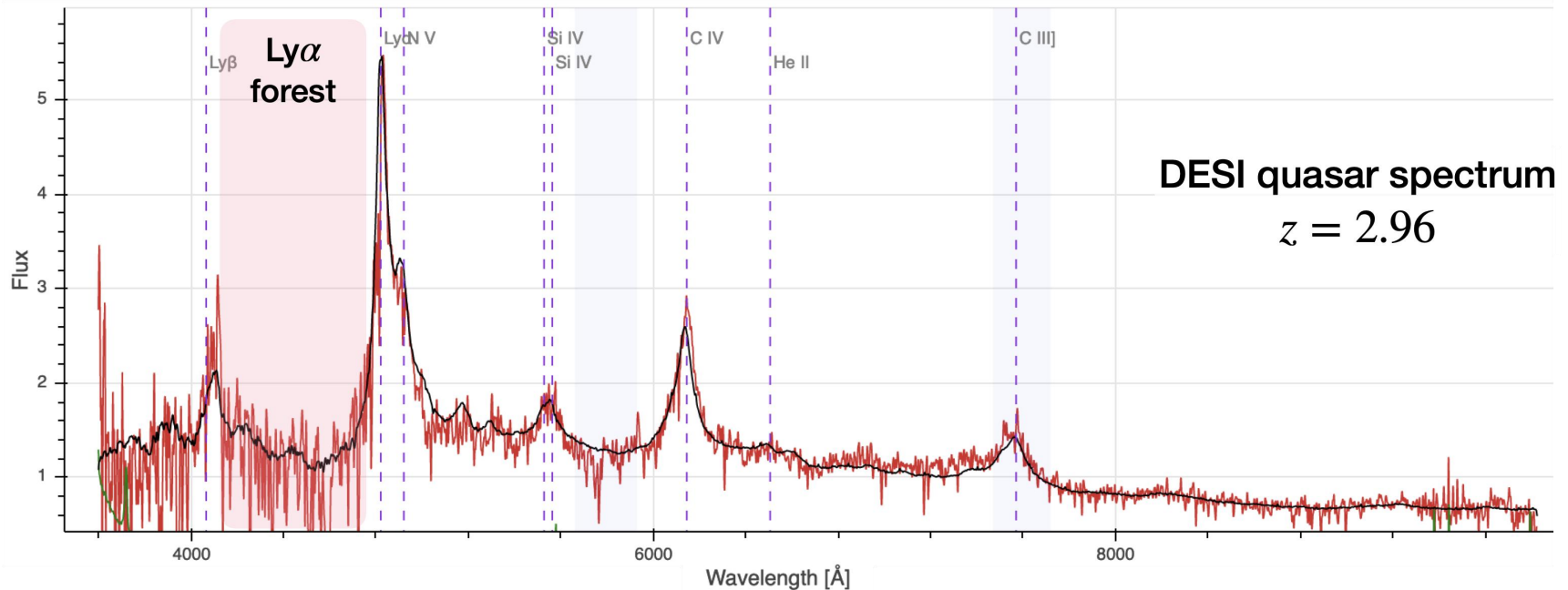
# 1. Introduction to Ly- $\alpha$ forests

# Quasars

Quasars:

- The most luminous objects in the universe
- Supermassive black holes
- Accretion disks of matter

$$\lambda_{RF, Ly\alpha} = 1216 \text{ \AA}$$

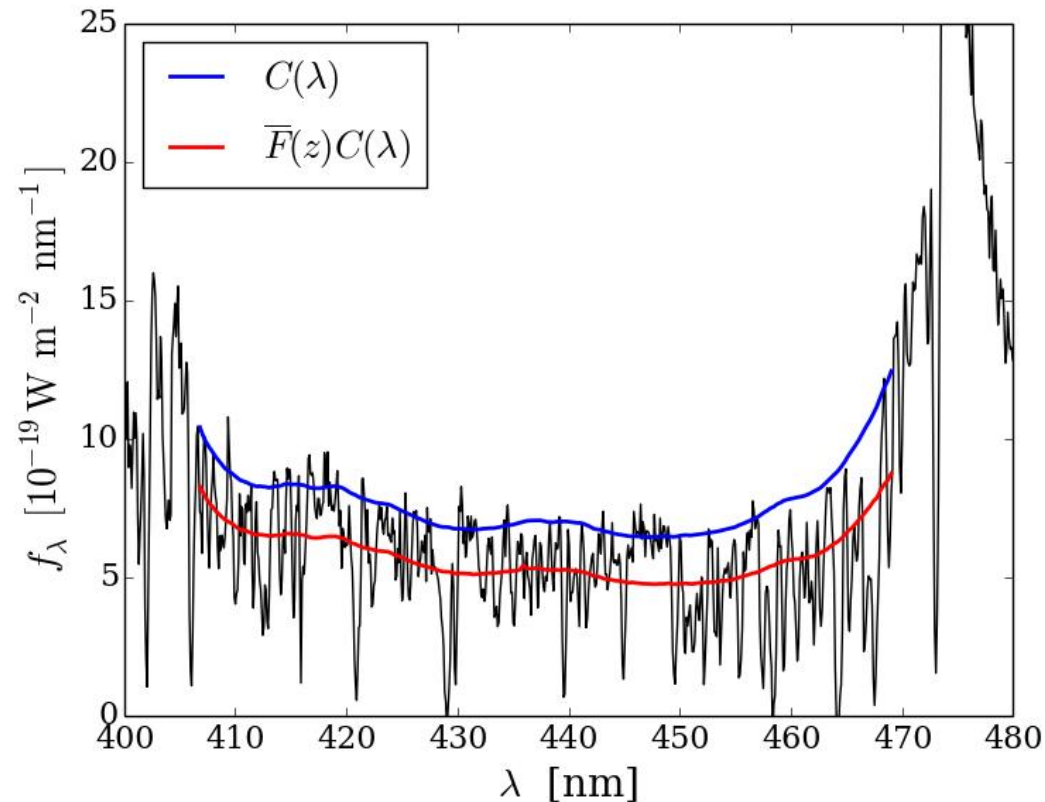


Quasar spectrum from DESI.

# The Lyman-alpha forest

The Lyman-alpha forest :

- QSO continuum: unabsorbed spectrum.
- Transmitted flux field: flux/continuum.
- Trace the density fluctuations and velocity-gradient fluctuations.



Thesis: Julianna Stermer

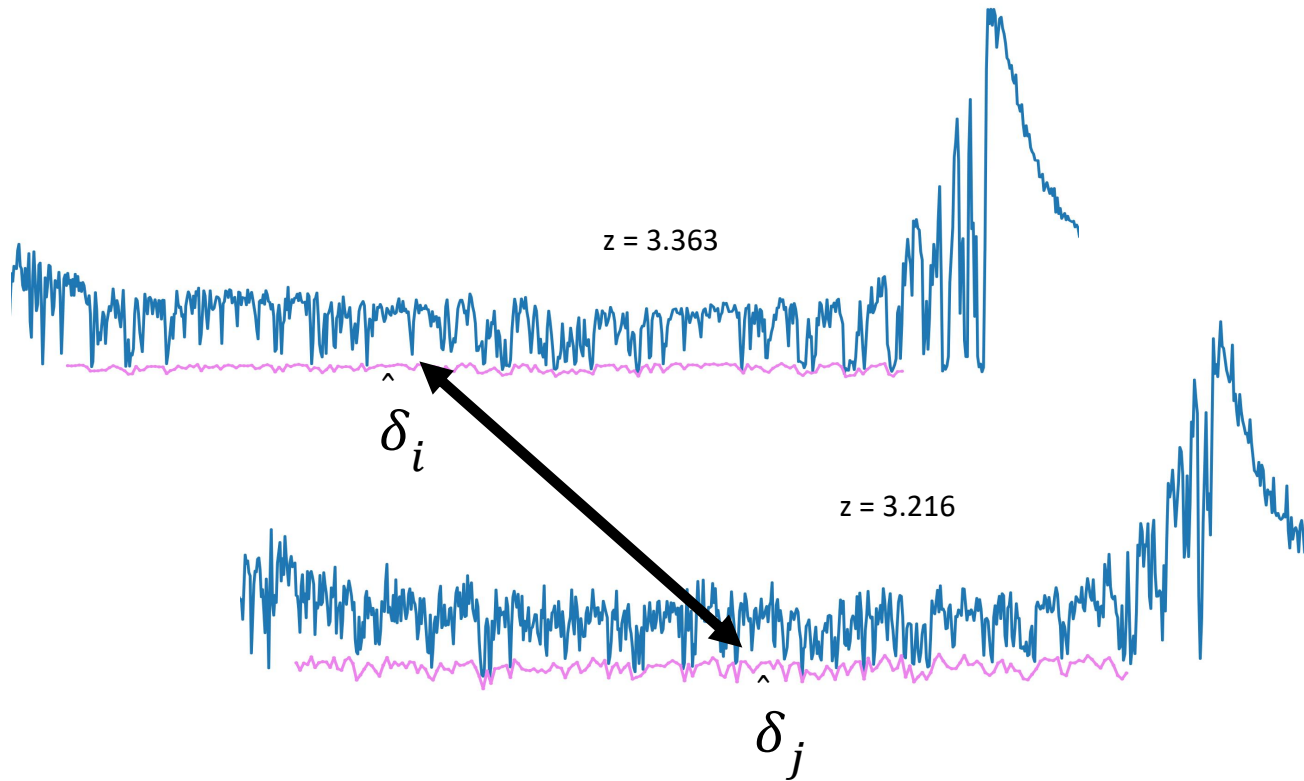
Transmitted flux fraction:

$$F_q(\lambda) = \frac{f_q(\lambda)}{C_q(\lambda)}$$

Observed flux  
Quasar continuum

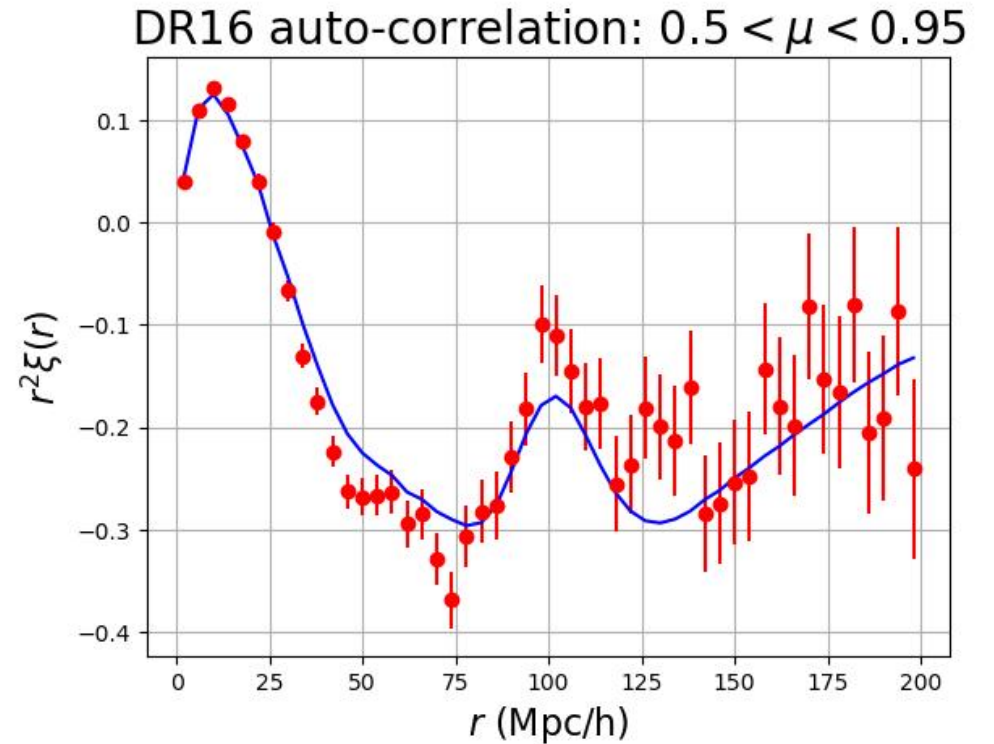
$$\text{Flux delta field: } \delta_F(\vec{x}) = \frac{F(\vec{x}) - \bar{F}}{\bar{F}} = \frac{f_q(\lambda)}{C_q(\lambda)\bar{F}} - 1$$

# The Lyman- $\alpha$ auto-Correlation function



Thesis: Julianna Stermer

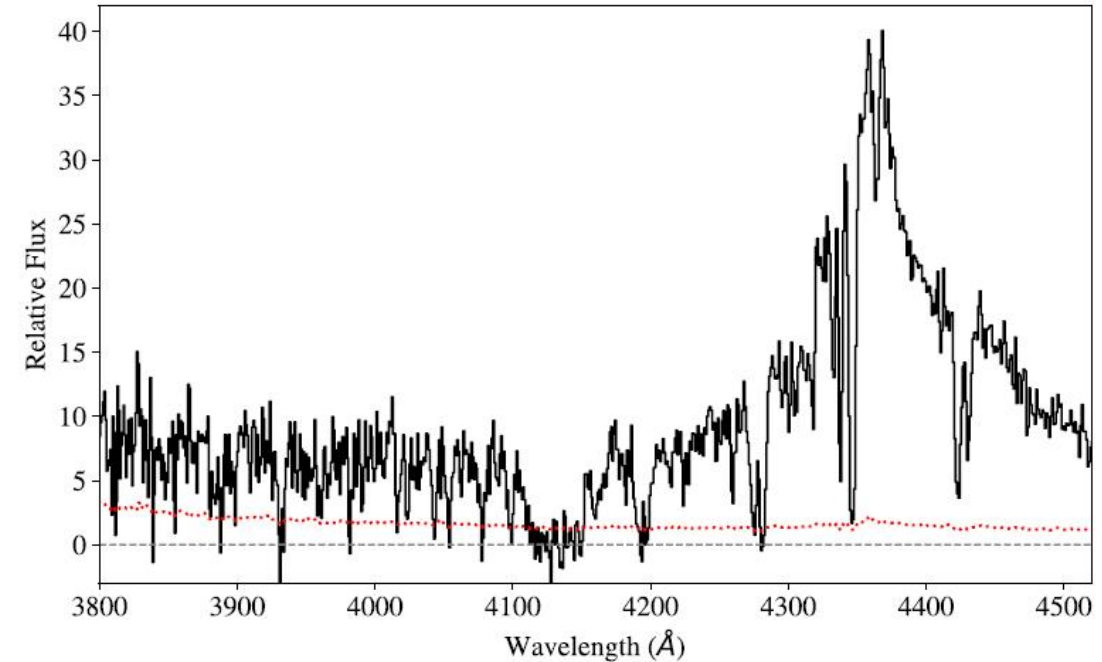
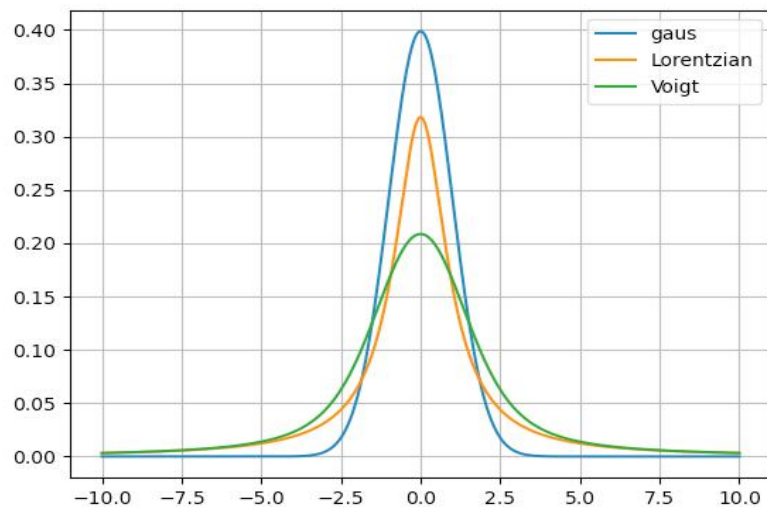
$$\xi_A = \frac{\sum_{(i,j) \in A} w_i w_j \delta_i \delta_j}{\sum_{(i,j) \in A} w_i w_j}$$



du Mas des Bourboux et al, arXiv:2007.08995

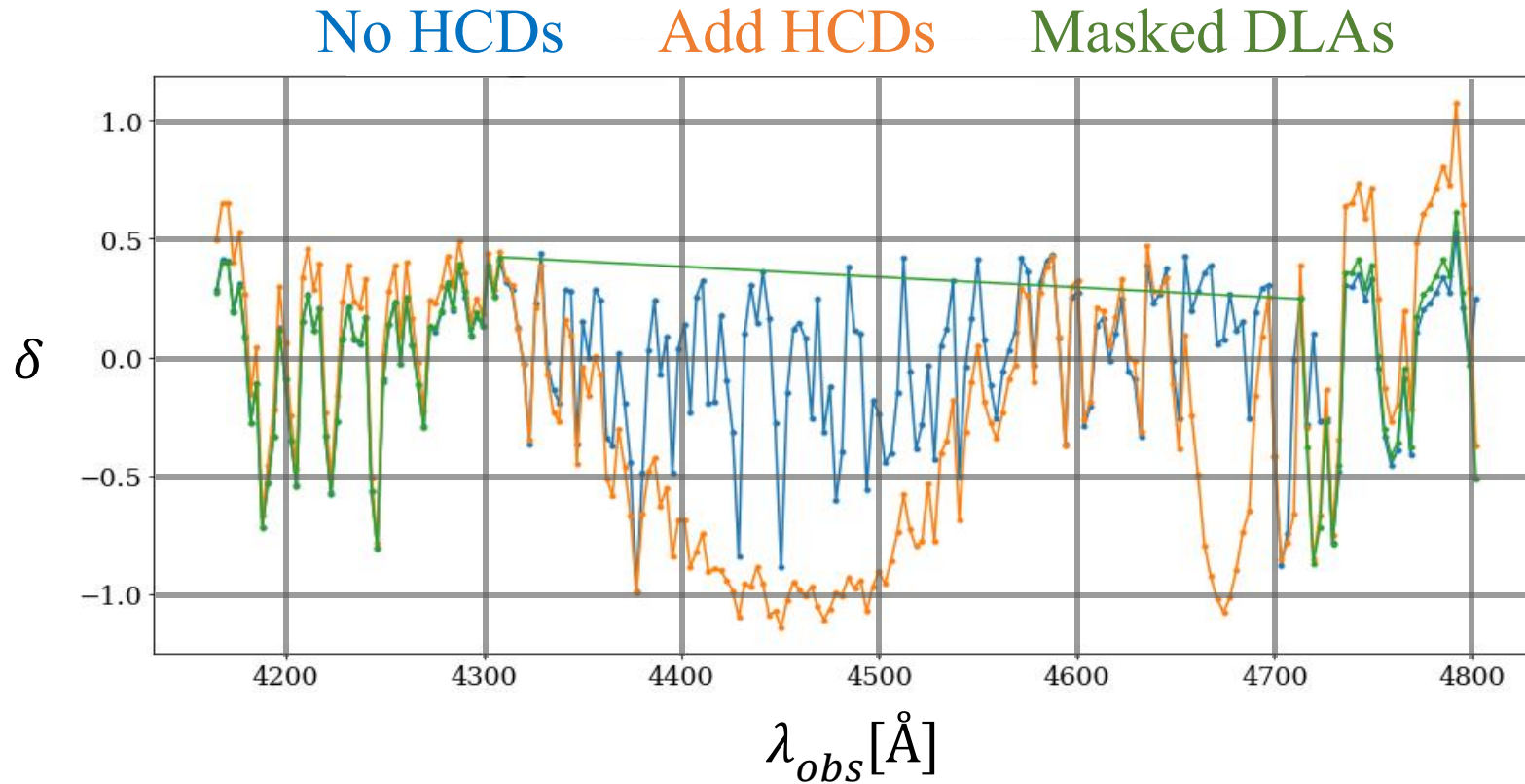
# Damped Lyman- $\alpha$ Systems (DLAs)

- NHI column density of a gas concentration:  
High Column Density Systems (HCDs):  $N \geq 10^{17} / \text{cm}^2$   
Damped Lyman- $\alpha$  Systems (DLAs):  $N \geq 10^{20} / \text{cm}^2$
- HCD absorption parametrized with Voigt profile
- Voigt profile = Gaussian  $\otimes$  Lorentzian  
Gaussian: thermal Doppler broadening  
Lorentzian: cross-section



*D. Parks et al. (2018)*

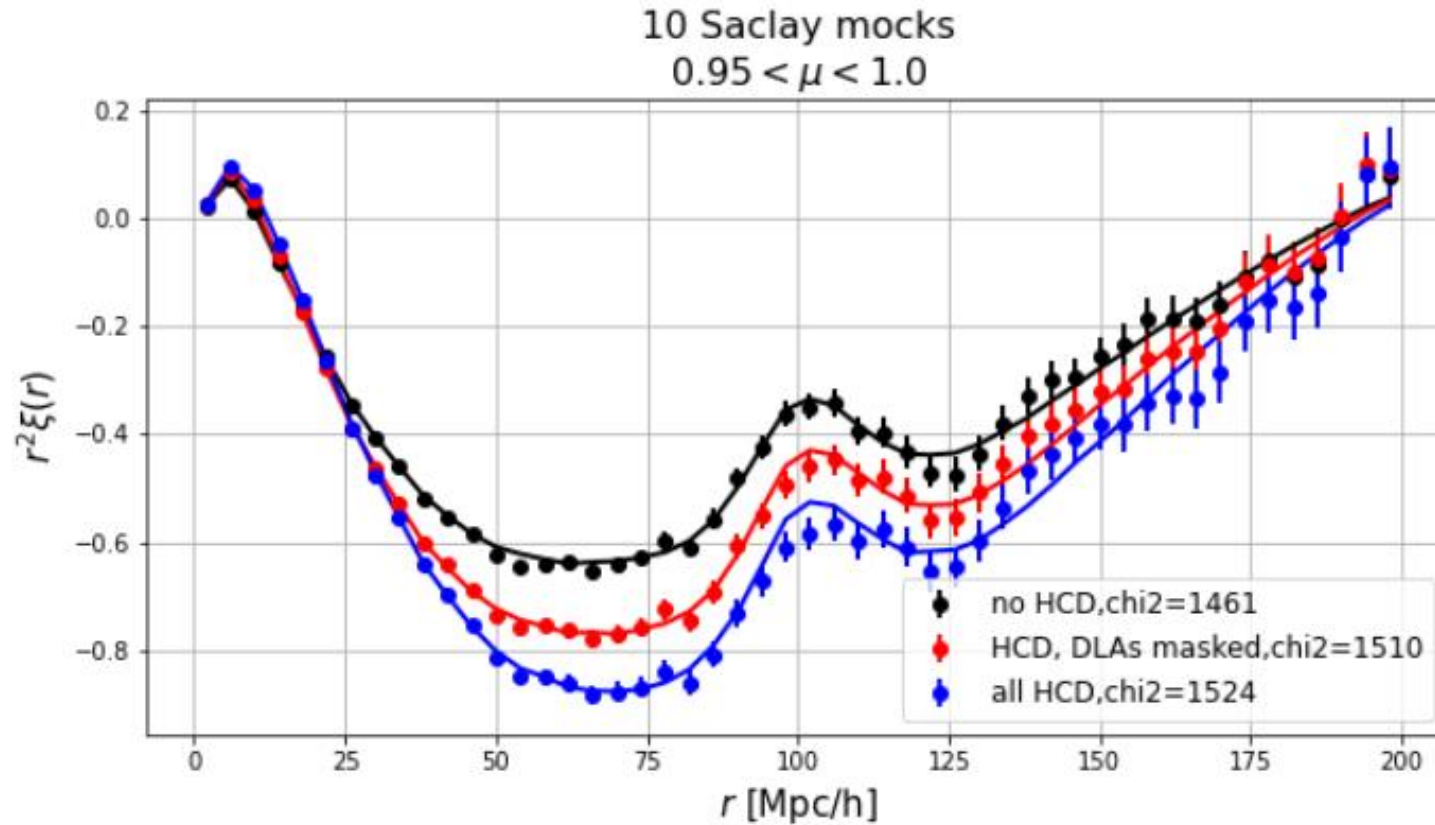
# HCD effect on the correlation function



- DLAs are detectable and can be masked.
- Smaller HCDs are not detectable but they smooth the delta field (cutoff in kpar) and thus affect the correlation function.



# HCD effect on the correlation function



d.o.f=1574

## 2. Modeling of HCDs

# Model Introduction

Modeling of Ly- $\alpha$  power spectrum:  $P_{Ly\alpha}(\mathbf{k}) = P_{QL}(\mathbf{k})D_{NL}(\mathbf{k})b_{Ly\alpha}^2(1 + \beta_{Ly\alpha}\mu^2)^2$

HCD model:

$$b'_{Ly\alpha} = b_{Ly\alpha} + b_{HCD}F_{HCD}(k_{\parallel})$$

$$b'_{Ly\alpha}\beta'_{Ly\alpha} = b_{Ly\alpha}\beta_{Ly\alpha} + b_{HCD}\beta_{HCD}F_{HCD}(k_{\parallel})$$

HCDs are absorbers



$b_{HCD} < 0$  like  $b_{LYA} < 0$

FT of HCDs profile



Cut-off at high  $k_{\parallel}$

# Model Introduction

Models for  $F_{\text{HCD}}$ :  $F_{\text{HCD}}^{\text{Rogers}}(k_{\parallel}) = \int (\widetilde{V} - 1)(k_{\parallel}, n) f(n) dn$  (Rogers et al. (2018))

$$F_{\text{HCD}}^{\text{exp}}(k_{\parallel}) = \exp(-L_{\text{HCD}} * k_{\parallel})$$

- In DR16:  $L_{\text{HCD}}=10$
- Best fit:  $L_{\text{HCD}}=3$

$$F_{\text{HCD}}^{L\beta\gamma}(k_{\parallel}) = \frac{1}{(1 + (L_{\text{HCD}} * k_{\parallel})^{\beta})^{\gamma}}$$

Best fit:  $L_{\text{HCD}}=13.4$

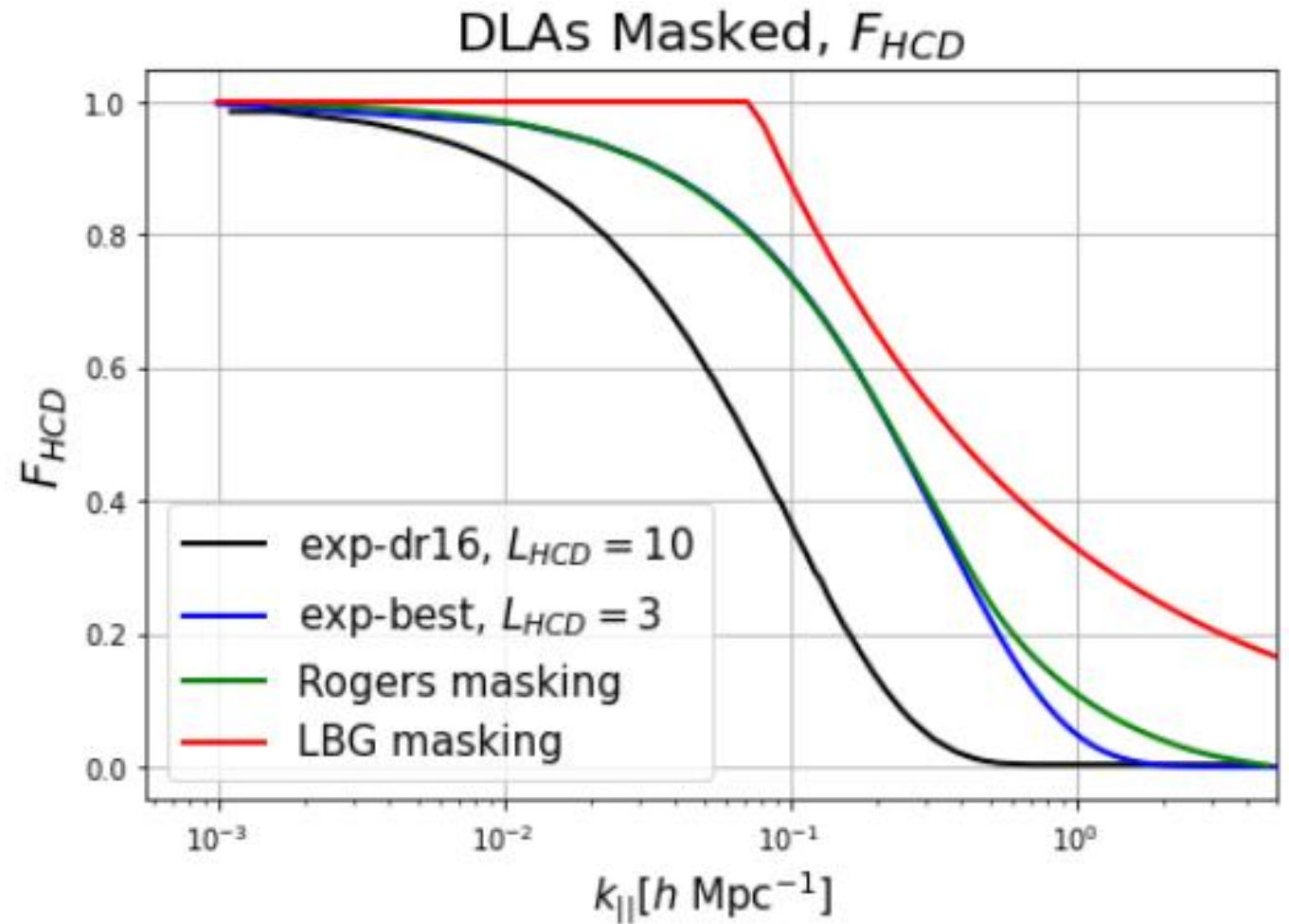
$\beta = 153.9$

$\gamma = 0.002776$

# $F_{\text{HCD}}$ function

$L\beta\gamma$  model:

- $F_{\text{HCD}} \sim 1$  for  $k_{\parallel} < 1/L_{\text{HCD}}$   
 $F_{\text{HCD}} \sim (Lk_{\parallel})^{-0.42}$  for  $k_{\parallel} > 1/L_{\text{HCD}}$
- Sharp kink at  $k_{\parallel} \sim 1/L_{\text{HCD}}$



# DR16, DLAs $>20.3$ masked

exp model:

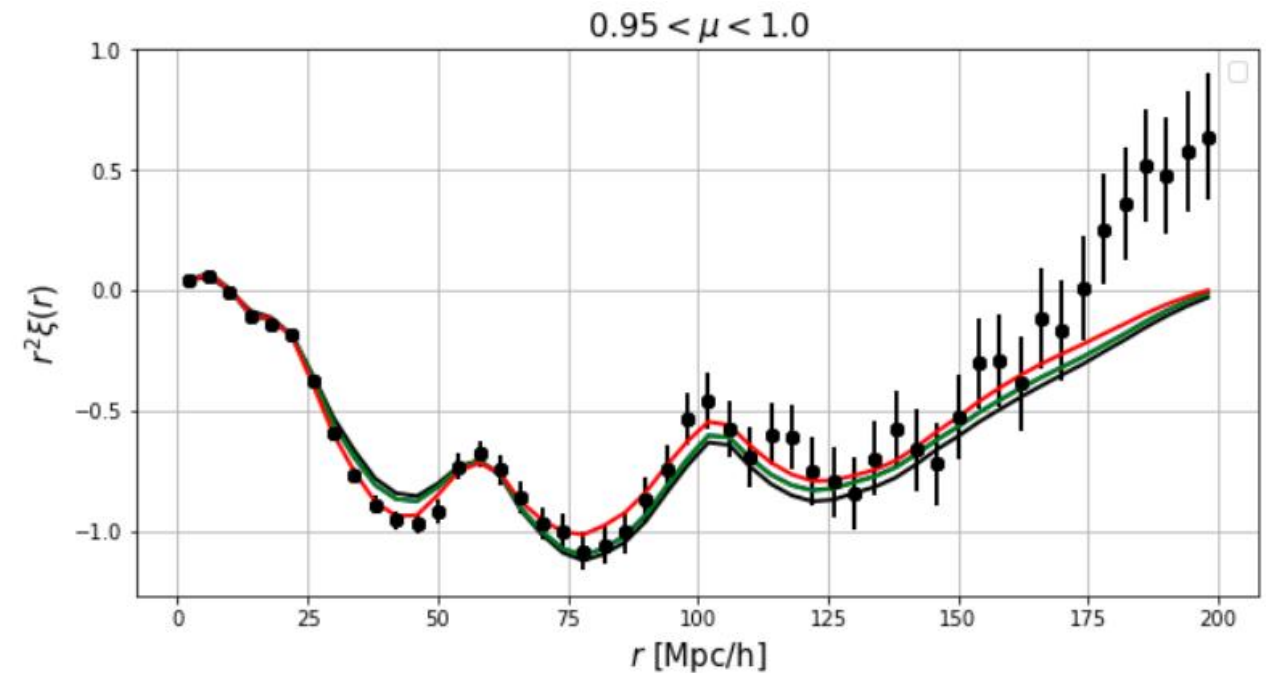
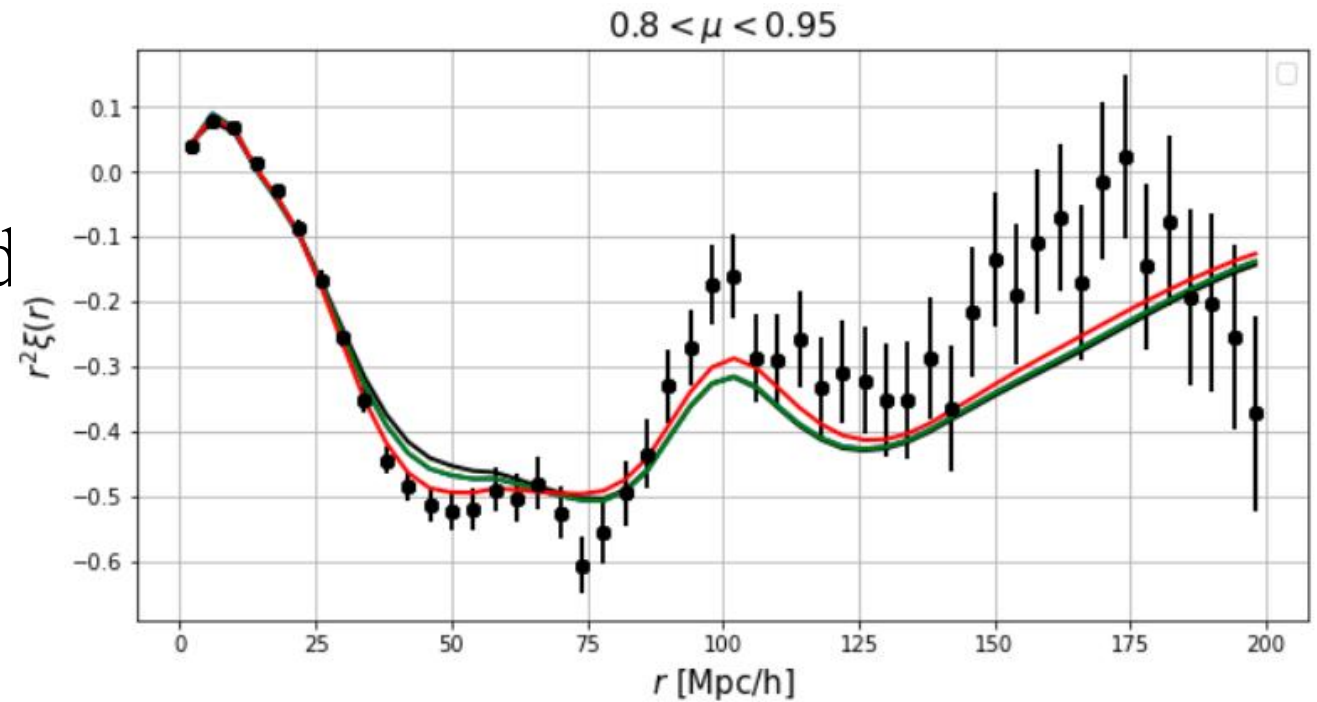
- $L_{\text{HCD}}=10$ :  $\chi^2=1604.73$
- $L_{\text{HCD}}=3$ :  $\chi^2=1576.96$

Rogers model:

- $\chi^2=1577.64$
- Overlaps best exp model

$L\beta\gamma$  model:

- $\chi^2=1557.3$
- Fit well the region  $25 < r < 80$



# DLAs $>20.3$ masked

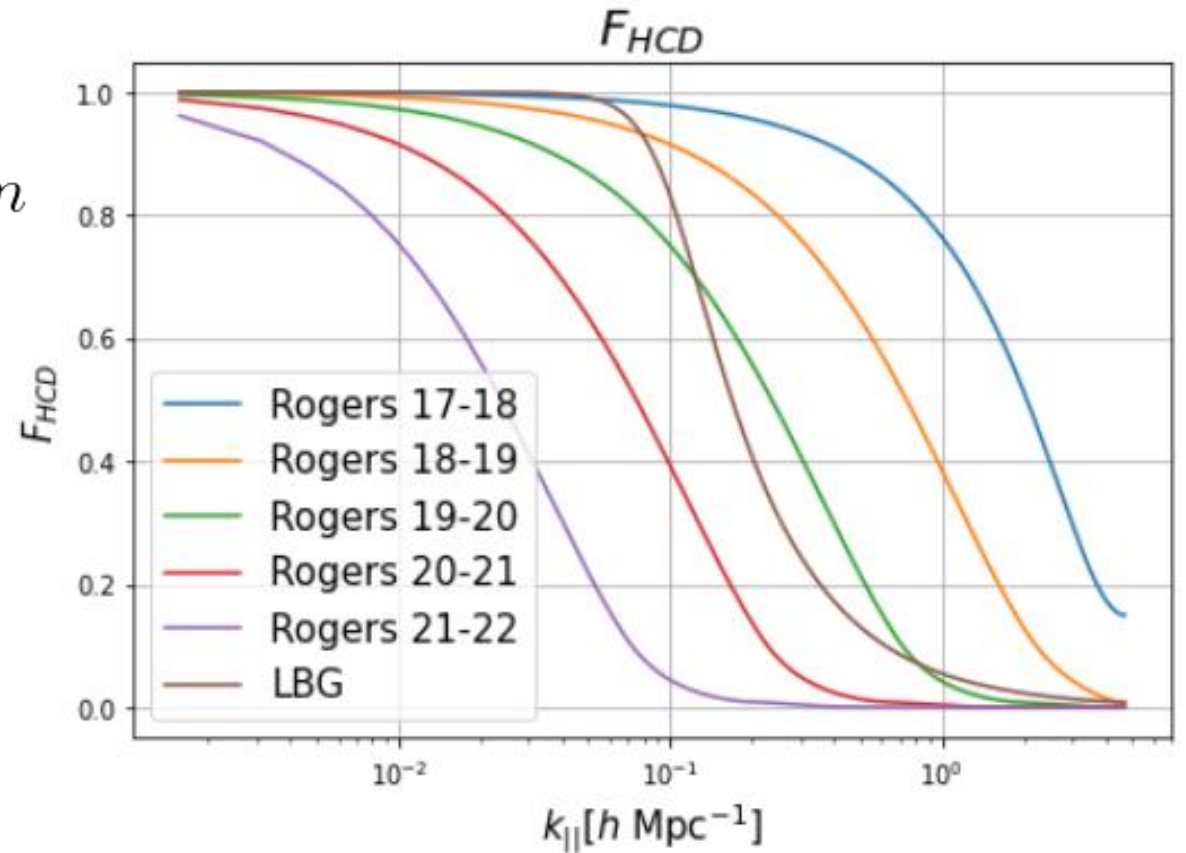
Data	Model	$\chi^2$	$\beta_{\text{LYA}}$	$\beta_{\text{HCD}}$	$b_{\text{LYA}}$	$b_{\text{HCD}}$
DR16 data	Rogers model	1577.64	2.49 $\pm$ 0.31	0.54 $\pm$ 0.08	-0.074 $\pm$ 0.018	-0.087 $\pm$ 0.007
	L $\beta\gamma$ model	1557.30	2.87 $\pm$ 0.40	0.54 $\pm$ 0.07	-0.061 $\pm$ 0.02	-0.088 $\pm$ 0.009
Saclay mocks	Rogers model	1511.86	1.62 $\pm$ 0.04	0.48 $\pm$ 0.09	-0.127 $\pm$ 0.04	-0.014 $\pm$ 0.004
	L $\beta\gamma$ model	1509.48	1.54 $\pm$ 0.03	0.49 $\pm$ 0.08	-0.136 $\pm$ 0.03	0.005 $\pm$ 0.003

- Rogers model works for mocks but not for data
- $b_{\text{HCD}}/b_{\text{LYA}}$  is ten times smaller in mocks than for data

# $F_{\text{HCD}}$ from Voigt profile and HCDs distribution $f(n)$

$$F_{\text{HCD}}^{\text{Rogers}}(k_{\parallel}) = \int (\overline{V - 1})(k_{\parallel}, n) f(n) dn$$

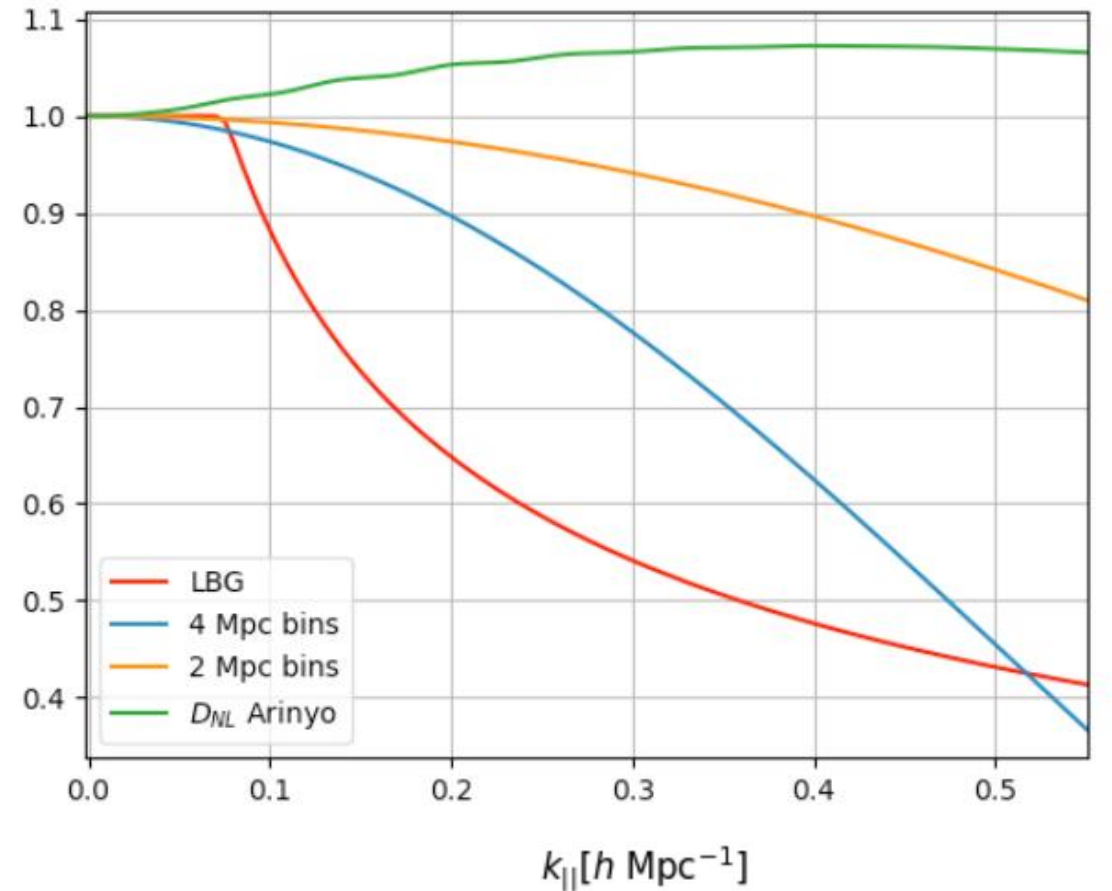
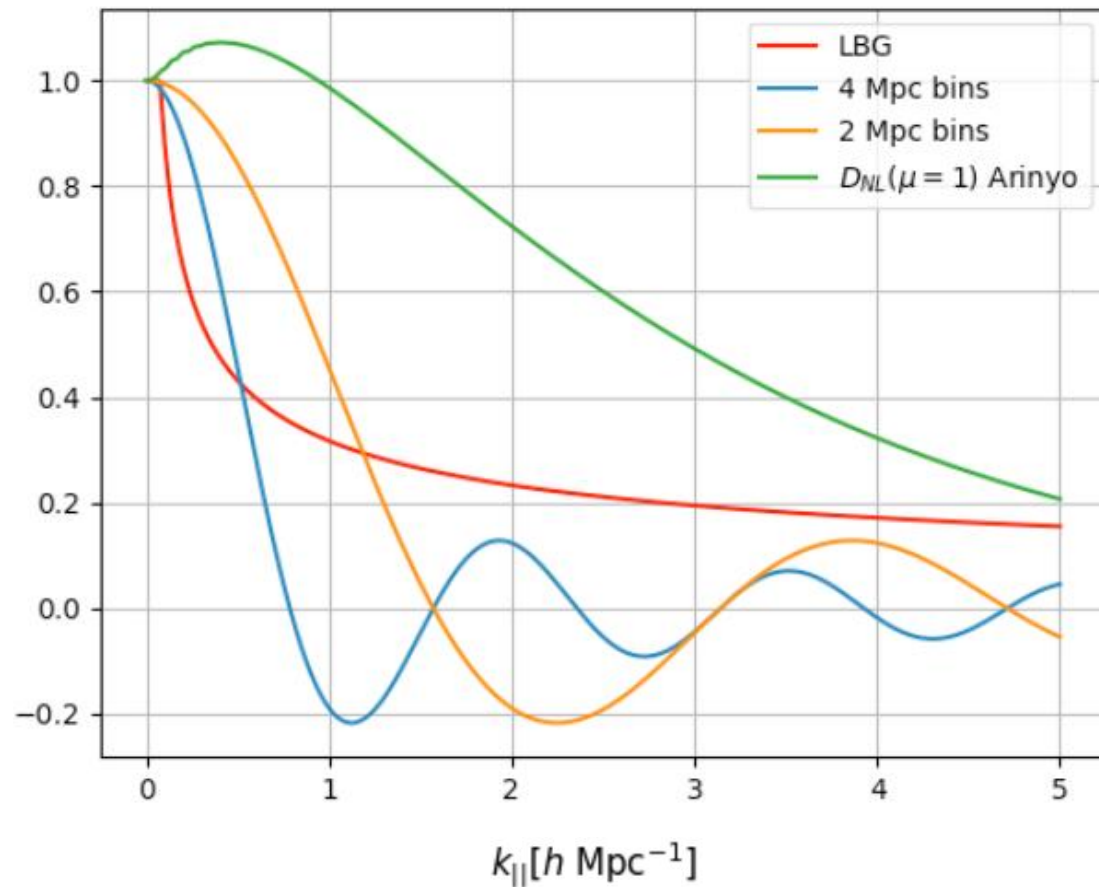
- no Voigt profile and HCDs distribution can reproduce  $L\beta\gamma$  model.
- $L\beta\gamma$  model captures something other than HCDs?





# Comparison of binning, $D_{NL}$ and $L\beta\gamma$ scales

$D_{NL}$ : cutoff from physical nonlinearities at small scales



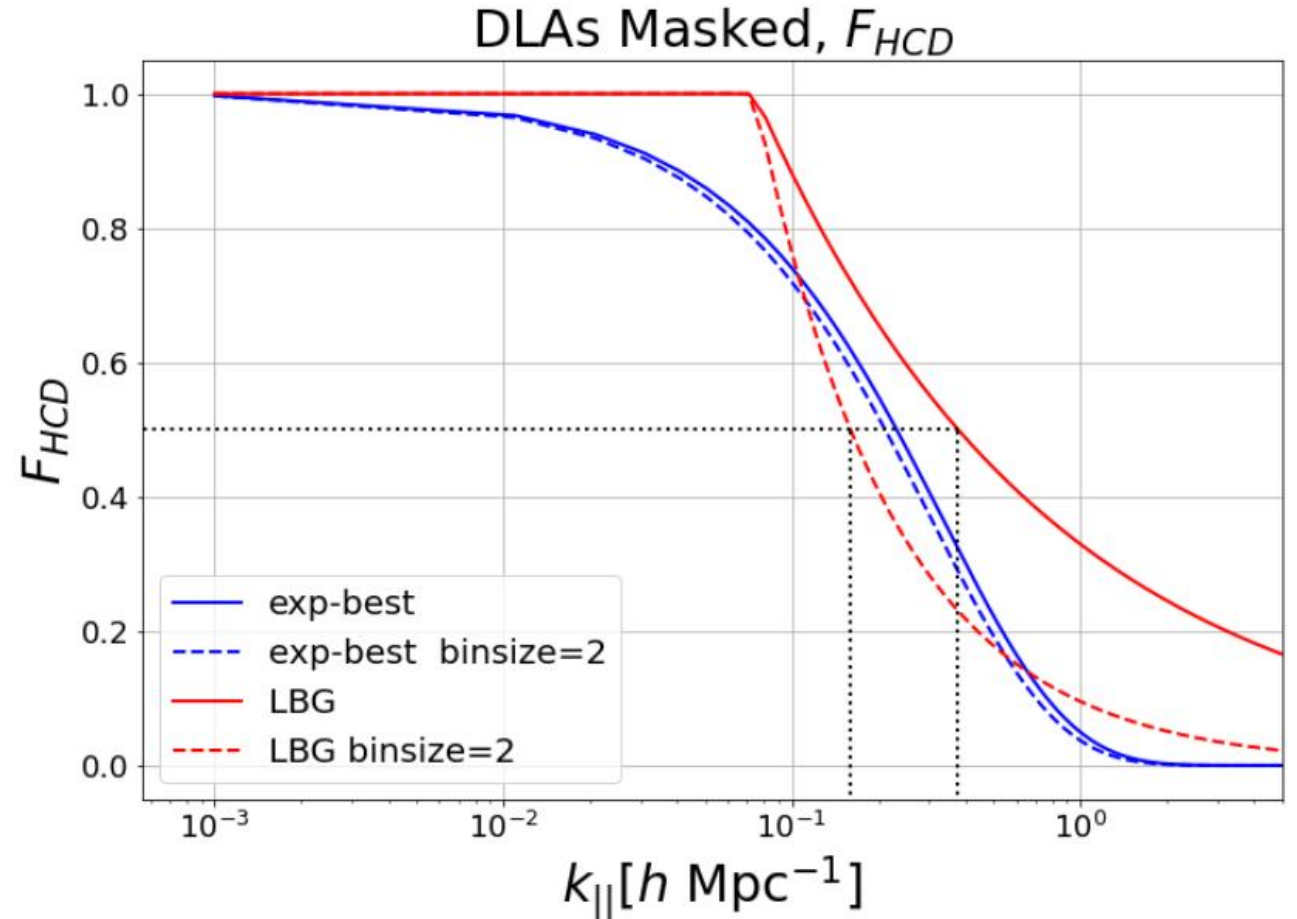
# Correlation function with binsize=2 Mpc/h

exp model:

- $L_{\text{HCD}}=3$
- No difference for binsize=4 or 2 Mpc/h

L $\beta$  $\gamma$  model:

- $L_{\text{HCD}}=13.5$
- the kink at  $k_{\parallel} \sim 1/L_{\text{HCD}}$ , independent of binning
- $\frac{1}{2} F_{\text{HCD}}$  at  $k_{\parallel} \sim L_{\text{bin}}/L_{\text{HCD}}$ , sensitive to binning



# Correlation function with binsize=2 Mpc/h

Data	Model	$\chi^2/\text{d.o.f}$	$\beta_{\text{LYA}}$	$\beta_{\text{HCD}}$	$b_{\text{LYA}}$	$b_{\text{HCD}}$
DR16 data binsize=4	exp Best	1577/1590	$2.48 \pm 0.30$	$0.54 \pm 0.08$	$-0.074 \pm 0.018$	$-0.087 \pm 0.007$
	L $\beta\gamma$ model	1557/1590	$2.87 \pm 0.40$	$0.54 \pm 0.07$	$-0.061 \pm 0.02$	$-0.088 \pm 0.009$
DR16 data binsize=2	exp Best	6244/6342	$2.74 \pm 0.85$	$0.48 \pm 0.09$	$-0.069 \pm 0.012$	$-0.09 \pm 0.022$
	L $\beta\gamma$ model	6219/6342	$2.03 \pm 0.43$	$0.51 \pm 0.09$	$-0.098 \pm 0.024$	$-0.051 \pm 0.023$

- $b_{\text{HCD}}/b_{\text{LYA}}$  more reasonable, but still ten times larger than for mocks

# Conclusion:

## Modeling of HCDs:

1.  $L\beta\gamma$  model gives smaller  $\chi^2$  (compared to voigt profile based models).
2.  $L\beta\gamma$  model goes through the points around  $r = 50$ .
3. No combination of Voigt profiles gives the  $L\beta\gamma$  model.
4.  $b_{\text{HCD}}$  is much larger for data than for mocks.
5.  $L\beta\gamma$  model captures something else than HCD?

## TODO:

- Physical understanding of  $L\beta\gamma$  model
- Understand effect of binning on  $L\beta\gamma$  model
- check whether the same problem is present in DESI data