Introduction 000

On the role of the initial stage concerning the dynamics of heavy-ion collisions using EPOS and PHSD

## Mahbobeh Jafarpour

Ph.D Supervisor: Klaus Werner with: Elena Bratkovskaya & Vadym Voronyuk Collaboration of Nantes-Frankfurt-Dubna groups GDR-QCD, l'île d'Oléron, France

23-25 May 2022





• At high T and  $\mu_B \to 0$ , Big Bang

Figure: Phase diagram of nuclear matter [1].



Figure: Space-time evolution of HIC.

Current Accelerators:

- SPS & LHC, CERN
- RHIC, BNL, New York

Future Accelerators:

- FAIR, Germany

# ladder and strings Off-shell remnantes Saturation [2, 3].

- INITIAL CONDITION: A Gribov-Regge multiple scattering approach is employed (PBGRT).
- CORE-CORONA SEPARATION: based on momentum and density of string segments.
- VISCOUS HYDRODYNAMIC EXPANSION: Using core part and cross-over equation of state (EOS) compatible with lattice QCD.
- STATISTICAL HADRONIZATION: employing Microcanonical decay/Cooper-Frye procedure and equilibrium hadron distribution.
- FINAL STATE HADRONIC CASCADE: applying the UrQMD model.

#### PHSD: Parton Hadron String Dynamics [4, 5].

- INITIAL A+A COLLISION: leads to formation of strings that decays to pre-hadrons, done by PYTHIA.
- QGP FORMATION: based on local energy-density.

Introduction

- QGP STAGE: evolution based on off-shell transport eqs. derived by Kadanoff-Baym eqs. with the DQPM defining the parton spectral function i.e. masses and widths.
- HADRONIZATION: massive off-shell partons with broad spectral functions hadronize to off-shell baryons and mesons.
- HADRONIC PHASE: evolution based on the off-shell transport eqs. with hadron-hadron interaction.

References

EPOSi+PHSDe : Employing a sophisticated EPOS approach to determine the initial distribution of matter (EPOSi)(partons/hadrons) and then using PHSD for the evolution of matter in a non-equilibrium transport approach (PHSDe).



Purpose: Separate "initial" and "evolution" effects

# Introduction EPOS EPOS2PHSD PHSD Results Conclusion and Outlook References occorrection in EPOS:

# Parton Based Gribov Regge Theory (PBGRT) [6]:

- Hard/Soft processes, Energy conservation by multiple Pomeron exchange
- Calculation of elastic/inelastic Cross-Sections (uncut ladder, soft contribution)
- Particle production [7] (cut ladder, semi-hard/hard contribution)



◆□ → ◆□ → ◆ = → ◆ = → ○ へ ○ 5/31



#### Core-Corona Separation in EPOS:



Energy loss of each string segment at given time  $\tau$ :

$$P_t^{new} = P_t - f_{Eloss} \int_{\gamma} \rho dL$$

If  $P_t^{new} > 0 \rightarrow$  Corona particle If  $P_t^{new} < 0 \rightarrow$  Core particle

< □ > < @ > < ≧ > < ≧ > ≧ 9 Q 0 6/31

#### Core-Corona pre-hadrons in EPOSi+PHSDe:



- rope segments: longitudinal color field, consider in 3D, larger string tension and transverse momentum
- Core pre-hadrons : decay of rope segments/clusters based on Microcanonical treatment [8]
- Corona pre-hadrons = Corona particles



#### Inserting all pre-hadrons from EPOS into PHSD:



< □ > < □ > < □ > < Ξ > < Ξ > Ξ の < ⊖ <sub>8/31</sub>





- All EPOS core/corona pre-hadrons are inserted into PHSD arrays
- Core pre-hadrons melted into QGP with respect to the melting condition ( $\varepsilon > 0.5 GeV/fm^3$ )
- Energy Density is computed in the Comoving frame in the three models:  $T^{\mu\nu}(\vec{q}) = \int \frac{d^3p}{E} p^{\mu} p^{\nu} f(\vec{q}, \vec{p}), \quad \varepsilon = T^{00}$

<□ ▶ < □ ▶ < ■ ▶ < ■ ▶ < ■ ▶ ■ ⑦ Q (~ 9/31)



# Energy Density Evolution



Figure: Time evolution of the energy density in the x-y plane (at z=0) for Au-Au collisions at 200A GeV with an impact parameter of 7fm, for three models.





Figure: Time evolution of the energy density in the x-y plane (at z=0) for Au-Au collisions at 200A GeV with an impact parameter of 7fm, for three models.

◆□▶ ◆□▶ ◆臣▶ ◆臣▶ 三臣 - のへで





Figure: Time evolution of the energy density in the x-y plane (at z=0) for Au-Au collisions at 200A GeV with an impact parameter of 7fm, for three models.

- EPOS: system expands in the longitudinal direction
- EPOSi+PHSDe: nearly identical to the EPOS
- PHSD: begins later and has more ED than others





Figure: Time evolution of the energy density in the x-y plane (at z=0) for Au-Au collisions at 200A GeV with an impact parameter of 7fm, for three models.

◆□▶ ◆□▶ ◆三▶ ◆三▶ 三三 のへぐ





Figure: Time evolution of the energy density in the x-y plane (at z=0) for Au-Au collisions at 200A GeV with an impact parameter of 7fm, for three models.

◆□▶ ◆□▶ ◆ □▶ ◆ □▶ ○ □ ○ ○ ○ ○





Figure: Time evolution of the energy density in the x-y plane (at z=0) for Au-Au collisions at 200A GeV with an impact parameter of 7fm, for three models.

◆□▶ ◆□▶ ◆ □▶ ◆ □▶ ○ □ ○ ○ ○ ○





Figure: Time evolution of the energy density in the x-y plane (at z=0) for Au-Au collisions at 200A GeV with an impact parameter of 7fm, for three models.

▲ロト ▲周ト ▲ヨト ▲ヨト ヨー のくぐ





Figure: Time evolution of the energy density in the x-y plane (at z=0) for Au-Au collisions at 200A GeV with an impact parameter of 7fm, for three models.

▲□▶ ▲□▶ ▲臣▶ ▲臣▶ 三臣 - のへで





Figure: Time evolution of the energy density in the x-y plane (at z=0) for Au-Au collisions at 200A GeV with an impact parameter of 7fm, for three models.

◆□▶ ◆□▶ ◆ □▶ ◆ □▶ ○ □ ○ ○ ○ ○





Figure: Time evolution of the energy density in the x-y plane (at z=0) for Au-Au collisions at 200A GeV with an impact parameter of 7fm, for three models.

◆□▶ ◆□▶ ◆三▶ ◆三▶ 三三 のへぐ





Figure: Time evolution of the energy density in the x-y plane (at z=0) for Au-Au collisions at 200A GeV with an impact parameter of 7fm, for three models.

↓ □ ▶ ↓ ● ▶ ↓ ■ ▶ ↓ ■ ♪ ○ ○ 20/31





Figure: Time evolution of the energy density in the x-y plane (at z=0) for Au-Au collisions at 200A GeV with an impact parameter of 7fm, for three models.

- EPOS: strong transverse expansion leads to more transverse flows
- EPOSi+PHSDe: less transverse expansion than EPOS, same forms as pure PHSD
- PHSD: more ED than others and expands spherically 21/31

Dynamical description of strongly interacting system in PHSD

PHSD



EPOS	EPOS2PHSD	PHSD	Results	Conclusion and Outlook	References
			00000		

# RESULTS

Comparing the Particle Production, Transverse Momentum  $(p_T)$ , Anisotropic Flow  $(v_2 \text{ and } v_3)$  for Au-Au@200GeV With different simulations: EPOSi+PHSD, EPOS, and pure PHSD

<□ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □





24/31

## Transverse Momentum spectra: Au-Au@200GeV



25/31



Event Plane method, Fourier series:  $E \frac{d^3 N}{d^3 p} = \frac{1}{2\pi} \frac{d^2 N}{p d p_t d y} (1 + \sum_{n=1}^{\infty} 2v_n \cos(n(\phi - \Psi_{EP})))$   $v_n(p_t, y) = < \cos(n(\phi - \Psi_{EP})) >$  $v_2 = \text{elliptic flow, } v_3 = \text{triangular flow, } v_4 = \text{quadrangular flow}$ 





Figure: EPOS, EPOSi+PHSD, pure PHSD

うくで 26/31

Introduction EPOS EPOS2PHSD PHSD Results Conclusion and Outlook References oco

# Flow behavior $v_2$ , and $v_3$ : Au-Au@200GeV



Figure: EPOS, EPOSi+PHSD, pure PHSD

◆□ → ◆□ → ◆ ■ → ▲ ■ → ● ● ○ QC 27/31

Summary and Conclusion:

- Two different HIC models were successfully combined
- Comparison of space-time and energy density evolution by EPOSi+PHSDe with pure EPOS and pure PHSD
- Comparing observables like charged particles production,  $p_T$ ,  $v_2$ ,  $v_3$  in three different frameworks
- High  $p_T$  part has not been improved yet by EPOSi+PHSDe
- The distinctions between EPOS and PHSD are related to their "evolutions"
- In EPOSi+PHSDe and pure PHSD, the partons do not interact strongly enough to produce something equivalent to "strong pressure gradiants", which are reasonable to the transverse flow in EPOS
- The partonic scatterings do not provide sufficient "thermalization" in EPOSi+PHSDe

#### Current work:

• Investigation of dilepton production in EPOSi+PHSDe compared to pure PHSD.

Outlook:

- Employing the early hydrodynamical evolution from EPOS (EPOSh), then use the PHSD evolution (PHSDe) to study the production of particles in higher  $p_T$ .
- Checking the heavy-flavor particle behavior in EPOSi+PHSDe and comparing the results with two other models.
- Comparing EPOSi+PHSDe with different ranges energies from RHIC to LHC for various systems like p-p and Au-Au.
- Studying the inclusive photon yield in EPOSi+PHSDe compared to pure PHSD.





のへで 30/31

Introduction	EPOS	EPOS2PHSD	PHSD	Results	Conclusion and Outlook	References
000	000	00000000000000	0	00000		●
D.f	T					

## References I

- J. C. Collins and M. J. Perry, "Superdense matter: neutrons or asymptotically free quarks?," *Physical Review Letters*, vol. 34, no. 21, p. 1353, 1975.
- [2] K. Werner, F.-M. Liu, and T. Pierog, "Parton ladder splitting and the rapidity dependence of transverse momentum spectra in deuteron-gold collisions at the bnl relativistic heavy ion collider," *Physical Review C*, vol. 74, no. 4, p. 044902, 2006.
- [3] T. Pierog, I. Karpenko, J. M. Katzy, E. Yatsenko, and K. Werner, "Epos lhc: Test of collective hadronization with data measured at the cern large hadron collider," *Physical Review C*, vol. 92, no. 3, p. 034906, 2015.
- [4] W. Cassing and E. Bratkovskaya, "Parton transport and hadronization from the dynamical quasiparticle point of view," *Physical Review C*, vol. 78, no. 3, p. 034919, 2008.
- [5] W. Cassing and E. Bratkovskaya, "Parton-hadron-string dynamics: An off-shell transport approach for relativistic energies," *Nuclear Physics A*, vol. 831, no. 3-4, pp. 215–242, 2009.
- [6] H. J. Drescher, M. Hladik, S. Ostapchenko, T. Pierog, and K. Werner, "Parton-based gribov-regge theory," *Physics Reports*, vol. 350, no. 2-4, pp. 93–289, 2001.
- [7] K. Werner, B. Guiot, I. Karpenko, and T. Pierog, "Analyzing radial flow features in p-pb and p-p collisions at several tev by studying identified-particle production with the event generator epos3," *Physical Review C*, vol. 89, no. 6, p. 064903, 2014.
- [8] K. Werner and J. Aichelin, "Microcanonical treatment of hadronizing the quark gluon plasma," *Phys. Rev. C*, vol. 52, pp. 1584–1603, 1995.
- [9] R. Snellings, "Elliptic flow: a brief review," New Journal of Physics, vol. 13, no. 5, p. 055008, 2011.