



New Extension in Nançay Upgrading LOFAR







Radio @ OHP



BENEFIT:

no time loss: observations possible day&night if wind < 50 km/s optical obs hampered by unstable weather conditions: 68% clear nights/yr

GOALS:

- Undergraduate lectures @ (Bachelor and) Master in Physics: e.g.
 M1: Astrophysics (project based)
 - M2: Astronomical observations at OHP, next Instrument Scientist option
- …opening to R&D in µelectronics & telecom @ Ecole Centrale Marseille (L/M), Polytech Marseille (Cycle ingénieur), IUT (LP EPOCS)
- High-school lectures on (radio)astronomy research (Rectorat)
- 🛧 Outreach

Carlo Schimd, LAM - IPhU days, February 10-11, 2022



- Antenna diameter: 3 m
- Antenna type: Prime focus, parabolic
- Working frequency: 1420 MHz
- Bandwidth: 50 MHz
- Feed: H-FEED, optimized for 1420 MHz
- LNA: 2 high gain and stability 1420 MHz LNA units
- Polarization: circular, left and right hand
- Receiver: H142-One for 1420 MHz radio astronomy
- Spectrometer: 1024 channels (61 kHz each)
- Radiometer: 50 MHz instantaneous bandwidth
- A/D converter: 14 bits
- Mount: WP-100 computerized alt-azimuth
- Pier: C106-HEAVY High load capacity pier for concrete base
- Remote control: yes (internet + webcam)
- Maximum slewing speed: 90°/min (+anemometer)
- Control software: RadioUniversePRO (+ Stellarium)
- Weatherproof: Yes

Radio @ OHP

Carlo Schimd, LAM - IPhU days, January 11-12, 2021



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Example of high-res backend to be built "in house" by master students in Physics (Instrument Scientist) or µelectronics:

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UTCODE 100 Time 10 10 40

High speed – Radio Astronomy: to sample the entire band in 100,000 channels with a spectral resolution of 500 Hz, perfect for different studies in radio astronomy such as the **precise measurements of the differential Doppler effect** in the neutral hydrogen line in the Milky Way or **pulsar detection**.



Map Source: azel Size: 16.94 x 16.94* Step: 0.8* Aver: 20.0s Levels range: 643 Peak: 13249 FITS Filename: 20190602-075616_IMAGE-PROJ01-CASA_01#_01#.ftts UTC: 2019-06-02T07:56:16

RADIO2SPACE

RCPU-100

RADIOZSPACE

1



Most brilliant 21-cm sources in Northern hemisphere, observable with SPIDER-300A













Name	Nature of the contribution to the project	#PM 2022	#PM 2023	#PM 2024
Carlo Schimd	Management (acquisition, installation, tests). Teaching M1/M2 "FunPhys". Outreach.	3	3	3
Christophe Adami	Teaching M1/M2 "FunPhys". Member of the Time Allocation Committee (TAC).	0.25	0.25	0.25
Alexandre Beelen	Management assistant. Member of the Time Allocation Committee (TAC).	0.25	0.25	0.25
Stephane Basa	Management assistant (acquisition, installation).	0.5	-	-
Benoît Epinat	Member of the Time Allocation Committee (TAC).	0.25	0.25	0.25
Guilaine Lagache	Teaching M1/M2 "FunPhys".	0.25	0.25	0.25
Alessia Longobardi	Management assistant (tests). Member of the Time Allocation Committee (TAC).	0.5	1	1
Delphine Russeil	Teaching M1/M2 "FunPhys". Member of the Time Allocation Committee (TAC).	0.25	0.25	0.25
Patrice Theule	Management assistant (installation, tests). Teaching M1/M2 "FunPhys".	0.5	0.5	0.5
William Gillard	Development of projects for the 'Instrument Scientist' option of Master "FunPhys".	0.25	0.5	0.5
August Le Van Suu	OHP local contact. Relation with Centre d'Astronomie de Saint-Michel.	0.5	-	-
Jean Strajnic	Link with Direction régionale académique de l'enseignement supérieur – Région PACA	0.5	0.5	0.5



https://www.radio2space.com



BUDGET 2022 (update): 74580 € (+10% w.r.t. 12/2020)



BUDGET 2022 (updated): 74580 € ...priorities?

ITEM	CATEGOR	Y	PRIORITY	SUB-TOTAL/PRIORITY	remarks	
Mission CEO PrimaLuceLAB: 3-days mission with lodging at OHP	(TE)	300 €	1	2 700 €		
Honoraire CEO PrimaLuceLAB (3 jours) for installation project (TTC)	(OB)	2 400 €	1	2700 C		$I \land M \land \forall i$
Masonry at OHP: basement for SPIDER 300A telescope (concrete, pad reinforced with rebar)	(E)	3 400 €	2	3 400 €	*no VAT. New offer: 4 320 €	LABORATOIRE D'ASTROPHYSIQUE DE MARSEILLE
SPIDER 300A advanced radio telescope	(E)	39 950 €	3			
Kit for installation of electric cables and optical fibres	(E)	750 €	3			RÉGION IIII 🦄
NSGen noise generator for absolute calibration of the SPIDER radio telescope	(E)	3 950 €	3			SUD
IlltraSonic Wind Sensor for SPIDER radio telescones					R2S quote: no VAT	ALPES CÔTE D'AZUR
(to automatically park the radio telescope in safety position when the wind	(F)	3 700 €	3			
exceed the safety limit 50 km/s)	T ÖC	DEBATI	ONS	53 030 €		10 kC missing
Delivery of radiotelescope and accessories to OHP	(E)	2 240 €				I9 K€ missing
					*w/ VAT.	$\hat{\mu}$
Safety barrier at OHP (including gate; masonry work not included)	(E)	740 €	3		New offer (w/	
					masonry): 2 856 €	
UPS (onduleur)	(E)	1 400 €	3		online estimation	
Training for the use of the radiotelescope by PrimaLuceLab (3 hours)	(OB)	300 €	3			00
All-sky camera for SPIDER radio telescopes to remotely see the radio	(E)	1 950 €	4			
telescope	(-/			6 900 €		
Networking Module for RadioUniversePRO software for Radioastronomy	(E)	4 950 €				
Virtual Classroom	(-/		4			011/
Kit radio over fiber for radiotelescope SPIDER and optical fibres (at	(E)	8 550 €	5	8 550 €		
least 30 meters)	(-)					







RadioUniversePRO



noise generator







for RadioUniversePRO ethernet

all-sky webcam

Ongoing and future urgent actions & open questions





- find additional funding sources:
- waiting for answer by Actions Spécifique LOFAR-SKA: one call/year, "usually in autumn" but last call in Jan 2021
 Maison SKA France?
- 3) AMU/TIGER?
- 4) Rotary??!!



x*Marseille

budget for Priority 1-2 (visit OHP + projet + masonry): OK budget for Priority 3 (instrument): -19 k€

...wait until P3 fully acquired?

... or start operations (P1+P2) now?

de Haute-Provence

(DL Région PACA: 27/10/2023)

- Project Assistant to manage contracts and payments: IPhU/CNRS/...?
- **Operating costs: OHP ?**
- website: hosted by IPhU webpage/wiki?

Possible UNIQUE improvements:

interferometry !

5m dish antenna



+ 20 k€



Examples for Virtual Classroom



Experiments: solar radio emission with SPIDER radio telescope

Solar radio emission can be studied by using SPIDER radio telescopes. The Sun is in fact one of the most interesting radio source in the sky. In fact the Sun not only emits visible light but also other frequencies in the electromagnetic spectrum. For example, everyone can feel the Sun heat on our skin, expression of infrared radiation. In this article, with step-by-step guide, we see how the SPIDER radio telescopes detect radio waves coming from the Sun and we show...



Cassiopeia A recorded with SPIDER 300A radio telescope

Cassiopeia A it's an important object for radio astronomy, a supernova remnant located in Cassiopeia constellation with a flux of 2400 Jansky flux at 1420 MHz. Thanks to the large parabolic antennas, high sensitivity of the H142-One receiver and the advanced features of the RadioUniversePRO software, the SPIDER radio telescopes are able to detect Cassiopeia A. In this article we describe how we detected Cassiopeia A by using the 3 meter diameter SPIDER 300A advanced radio telescope: starting with antenna alignment, we detected interferences...



Taurus A recorded with SPIDER 300A radio telescope

Taurus A is the radio source in Taurus constellation that corresponds to the Crab Nebula (M1), the supernova remnant exploded on July 4, 1054 and noted by Chinese and Arabian astronomers of the time. Since then, the gas cloud has expanded and today is over 6 light years large. In this article we see how the SPIDER 300A radio telescope "discovered" it by capturing the radio waves emitted by Taurus A and converting them into a radio map, a real...

Example of radioastronomy at school



Radio astronomy at school: solar flux at 10 cm

Even objects around us emit radio waves. For example, pointing the radio telescope towards a nearby building, you will notice an increase in the value of radio waves. Comparing this value with the one found by pointing the Sun or the Moon, it will be possible to determine their surface temperature! Thanks to the possibility of calibrating the signal, SPIDER control software lets you to monitor the emission of a radio source for a long time. For example, by focusing to the Sun, you can make one measurement per day for the period of time necessary to its revolution (about 25 days) and then will be able to correlate it with other phenomena such as the number of sunspots visible with an optical telescope.