

Identification and annotation of numerical simulations of galaxy collisions.



By Pierre
Galois

Proto-cluster from TNG50 simulation.



Observatoire **astronomique**

de Strasbourg | ObAS

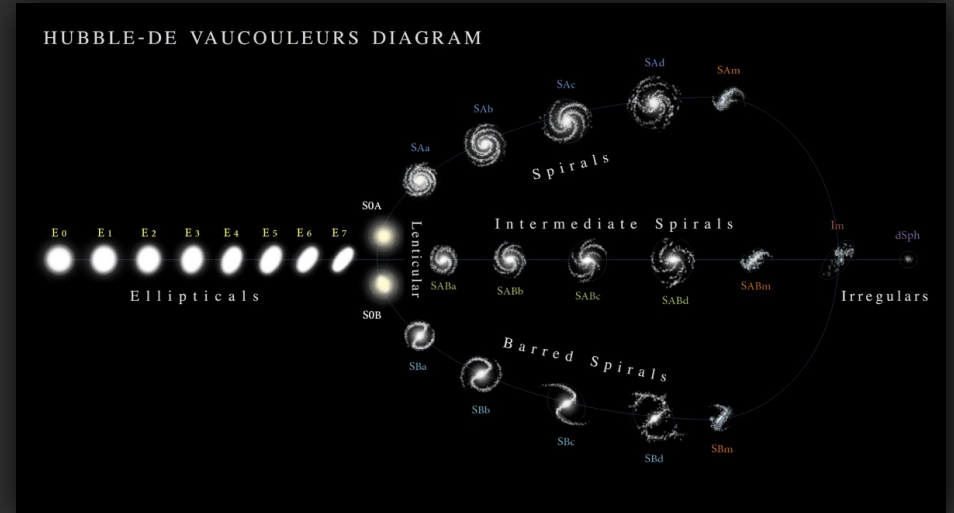
Supervised by Pierre-Alain Duc

Université

de Strasbourg

Introduction

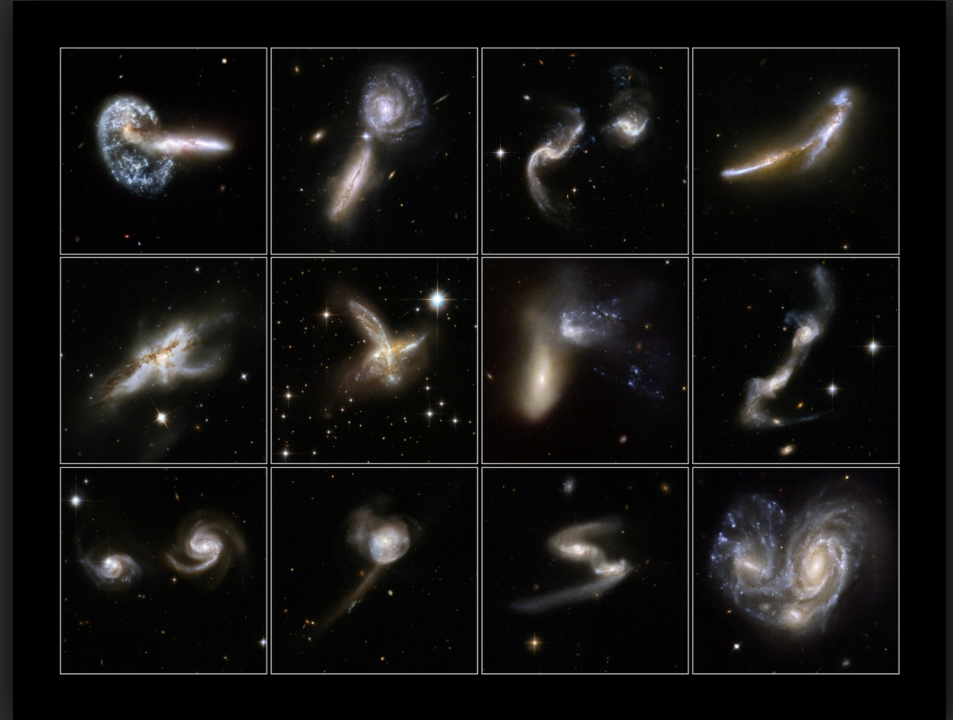
- Galaxy: stars, gas, dust, dark matter
- Two main families: spiral and elliptical
- Diversity of morphologies and internal properties



<https://en.wikipedia.org/wiki/File:Hubble-Vaucouleurs.png>

Introduction

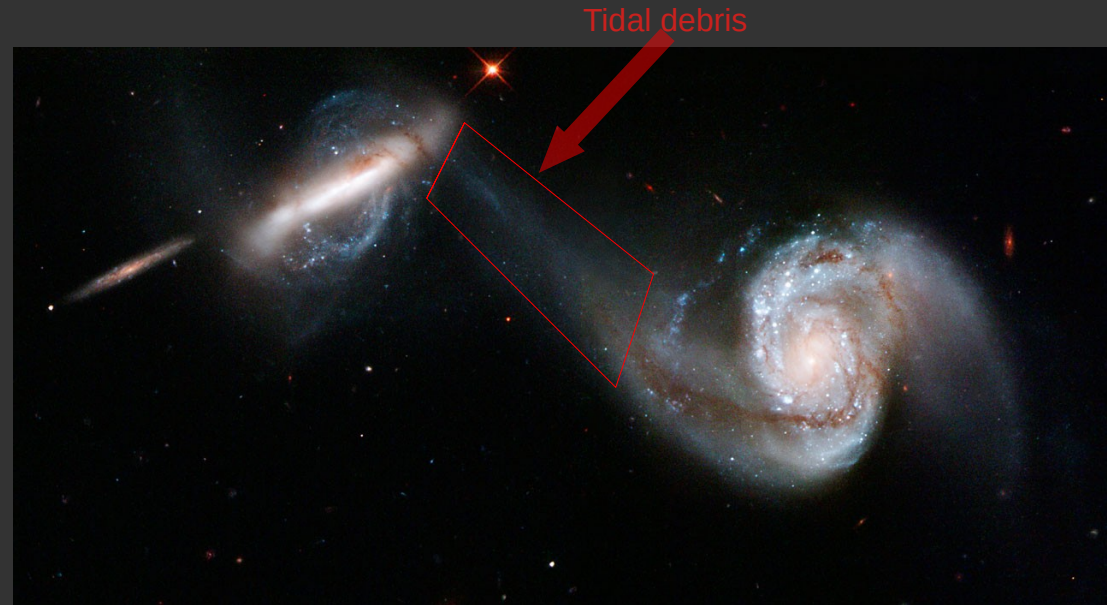
- Tend to interact, collide and merge
- Source of the observed diversity
- Gravitational interaction:
 - Deformation of galaxies
 - Apparition of tidal debris



NASA/ESA

Introduction

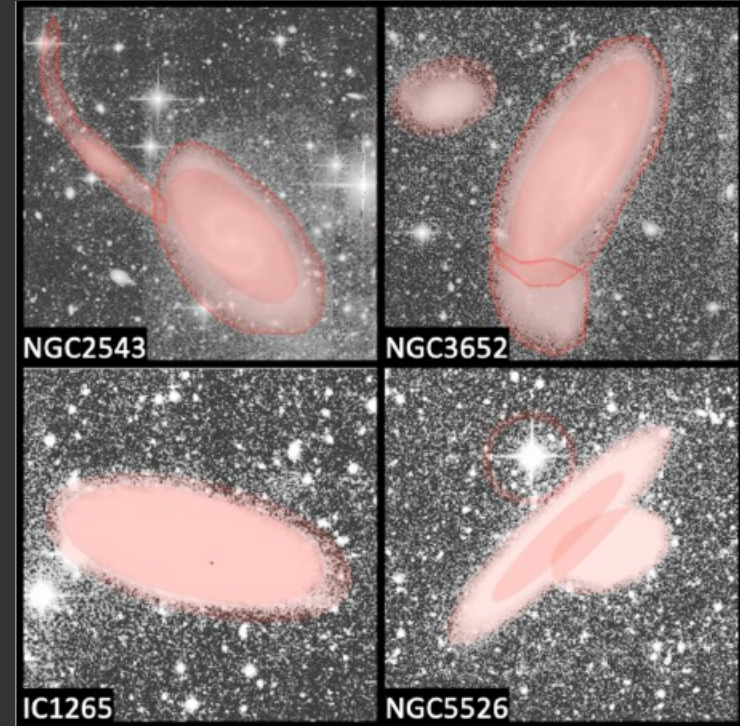
- Tidal debris: directly linked to galaxies history
 - Problem:
 - Faint and extended
 - Low contrast
- ⇒ Hard to observe
- Each image of galaxy must be characterized by hand: long process



Arp 87, NASA/ESA

Annotation and previous works

- Development of an annotation tool
- Intuitive and visual annotation
- Goal:
 - Identify, highlight and label tidal structures on galaxy images
 - Extract properties of this debris
- Large number of annotations already performed



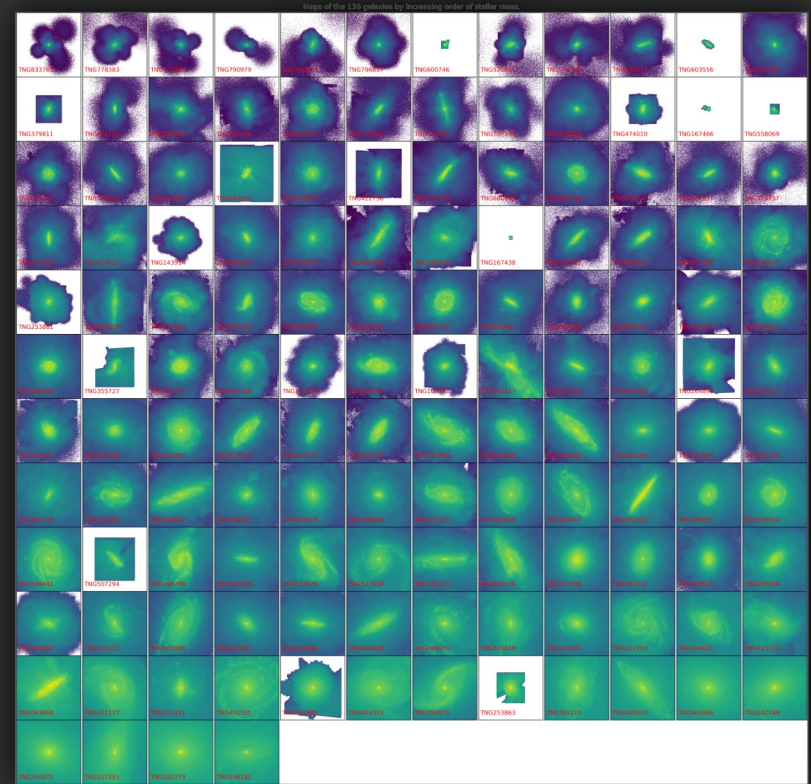
E.Sola et al.

Goals of the project

- Annotate images from simulation
- Interest of simulation:
 - Understand plurality of observations
 - Constrain models of galaxy evolution
- Adapt and characterize a sample of simulated galaxies, annotate them and extract properties of tidal debris

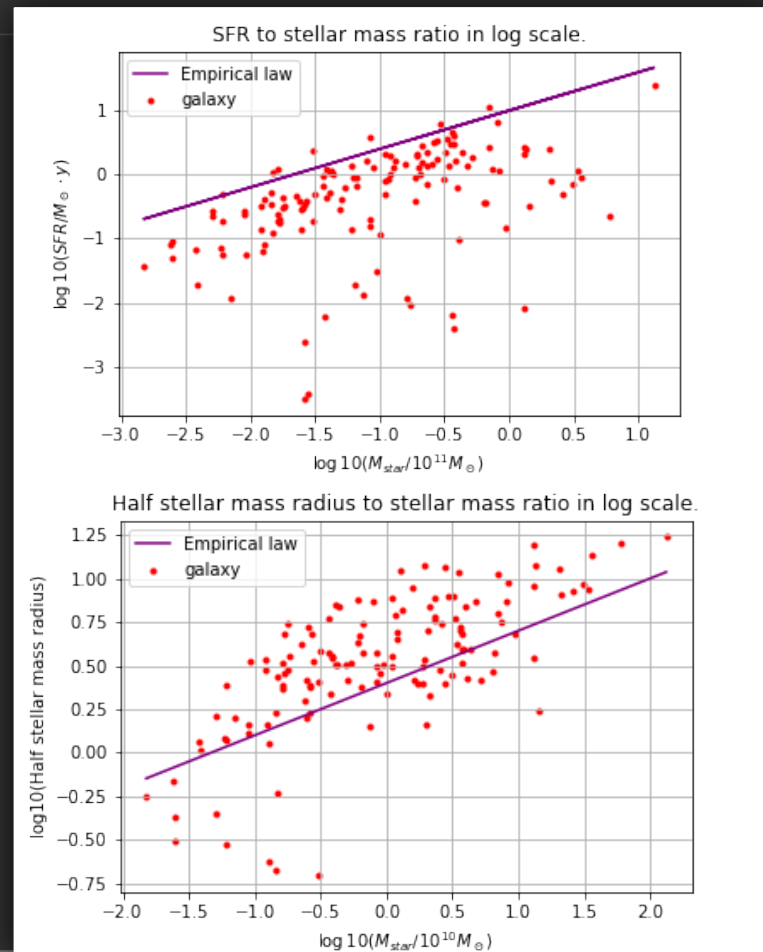
Characterization

- Our initial sample: 136 galaxies
- First step: check if our sample is consistent with observed galaxies
- Visual inspection:
 - Realistic aspect
 - Mainly spirals
 - Some very large galaxies



Characterization and scale relations

- Scale relation: link a “size” properties (radius, mass, etc...) to an other property
- Empirical relation.
- Our sample: consistent with this relation



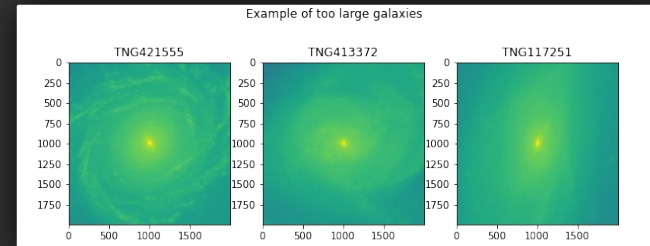
Characterization

- Advantages:

- Realistic aspects
- Realistic shapes
- Scale relations

- Counterparts:

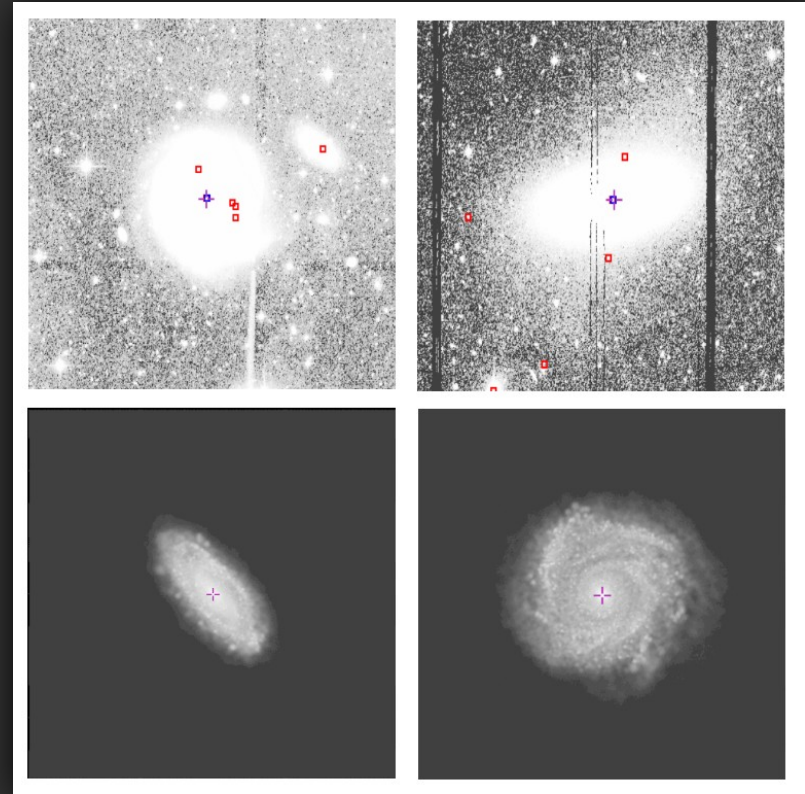
- Large number of spiral galaxies
- Small FOV



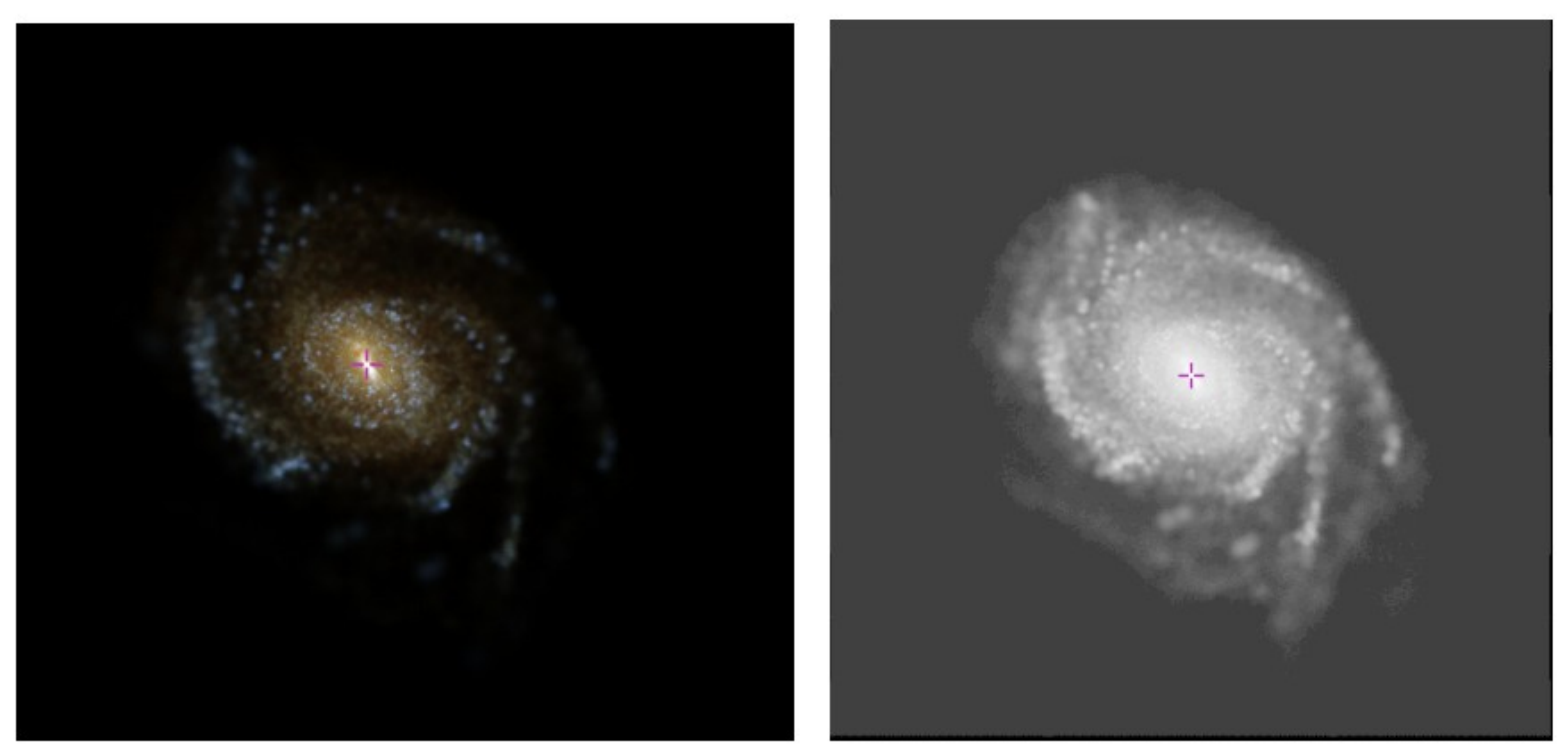
- Physical properties: consistent with observations

Preparation of the images

- Annotation tool: support only a specific extension \Rightarrow Need to convert raw images.
- Adapt to look like real images \Rightarrow Find optimal parameters.
- We kept 53 images from the 136.

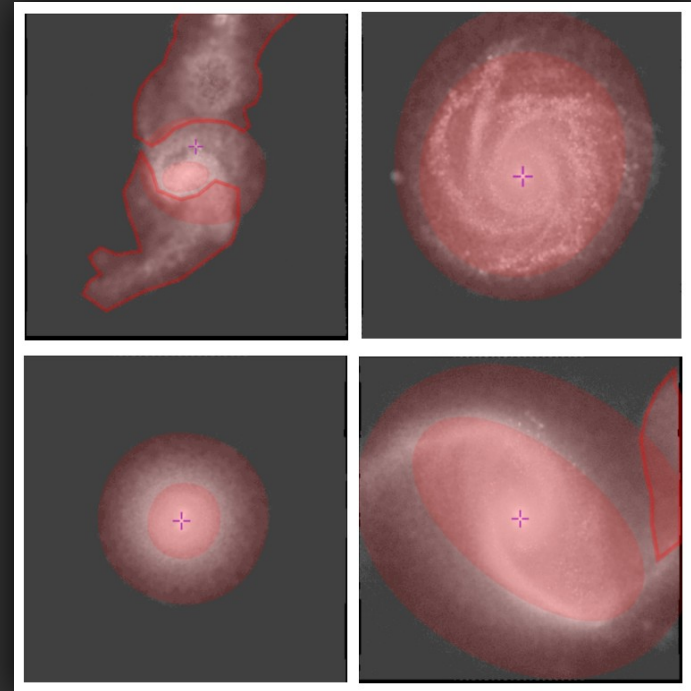


Preparation of the images

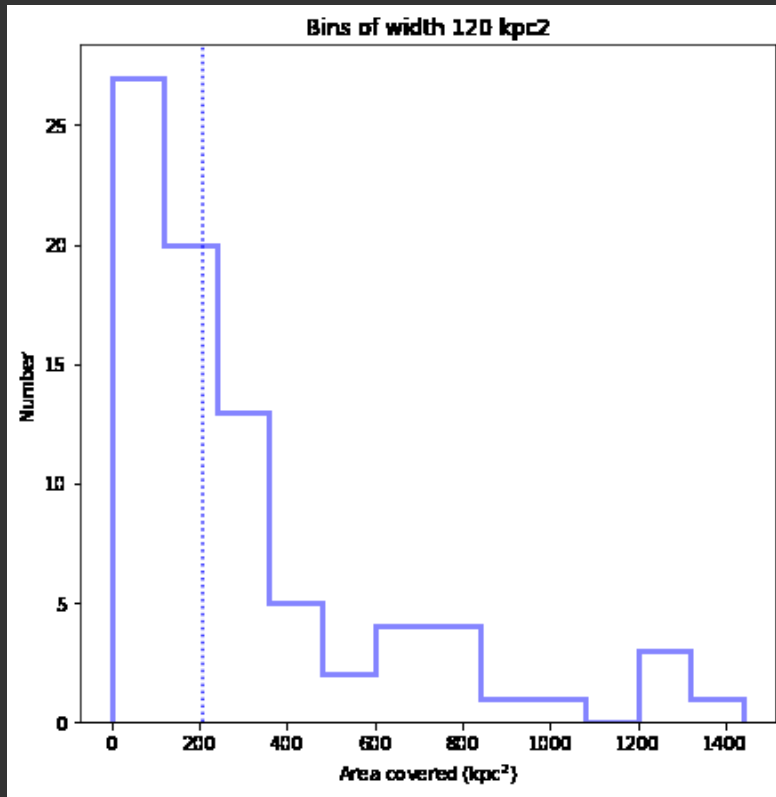


Annotation and preliminary results

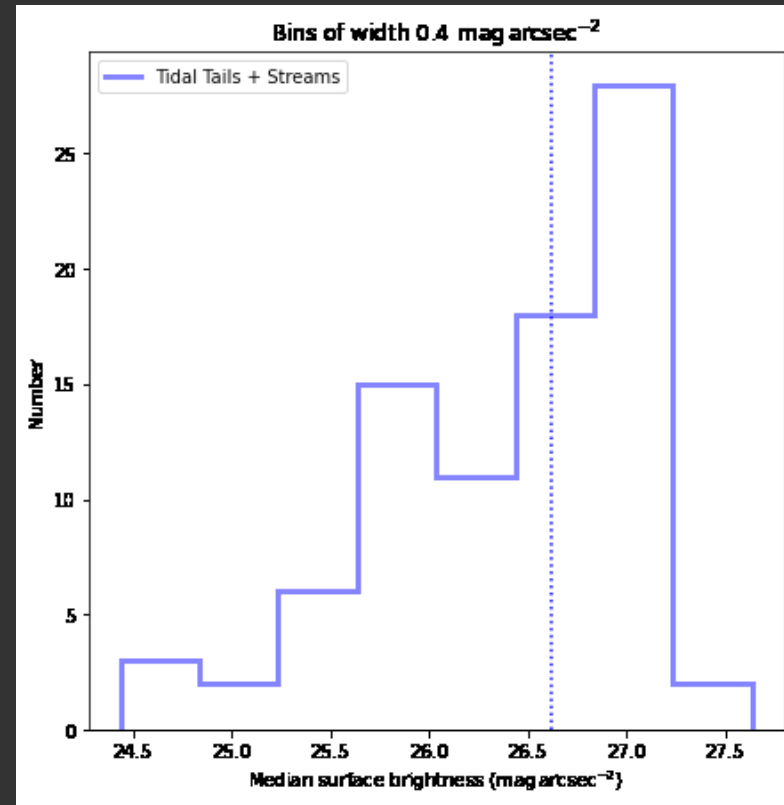
- Highlight main galactic components and tidal features:
 - Brightest part
 - Halo
 - Tails/plumes
- Allow access to their main properties:
 - Area
 - Size
 - Surface brightness
- Quantitative analysing of tidal debris



Annotation and preliminary results



E.Sola



E.Sola

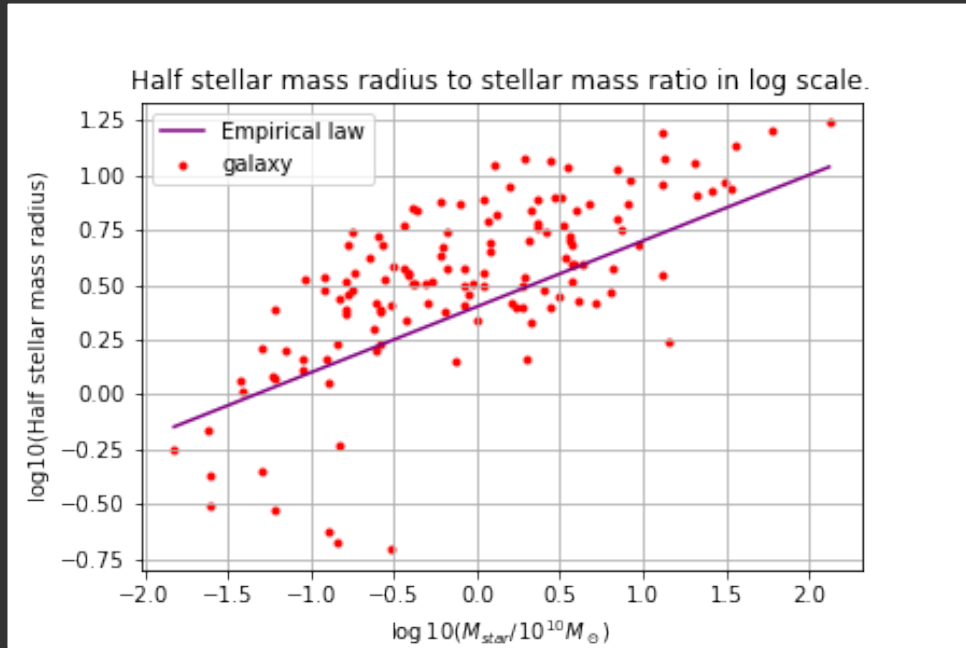
Conclusion and future prospects

- Heart of the project:
 - Characterize a sample of simulated galaxies
 - Annotation
 - Extraction of informations
- Properties of tidal debris \Rightarrow Redraw the history of the galaxy

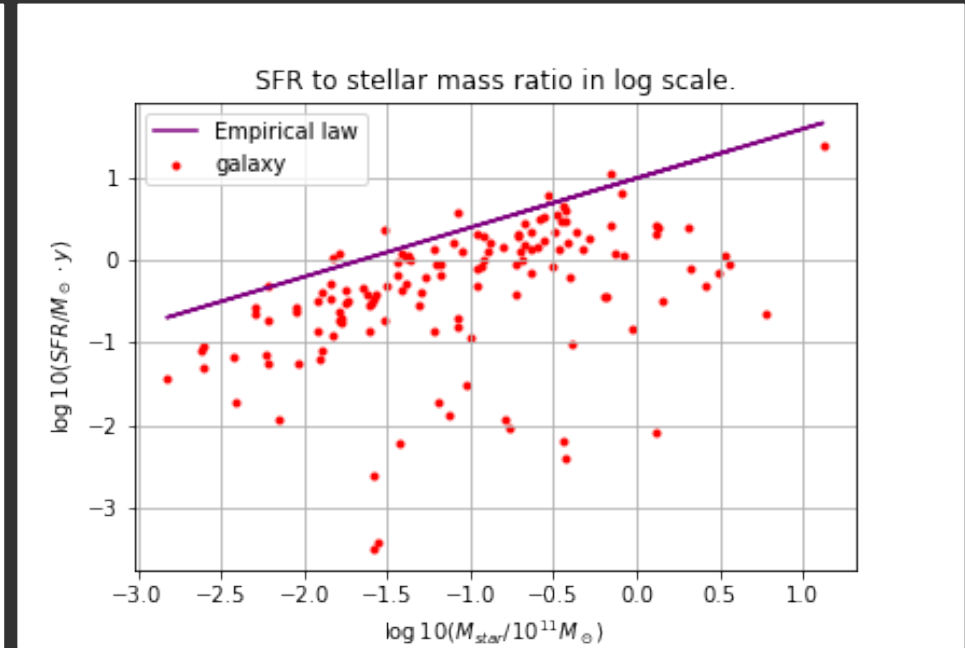
Conclusion and future prospects

- Next steps:
 - Better sample, much representative
 - Annotate more galaxies
 - Compare with annotations from real images
- Final goals: machine learning and automatization of the process

Annexe: Scale relations

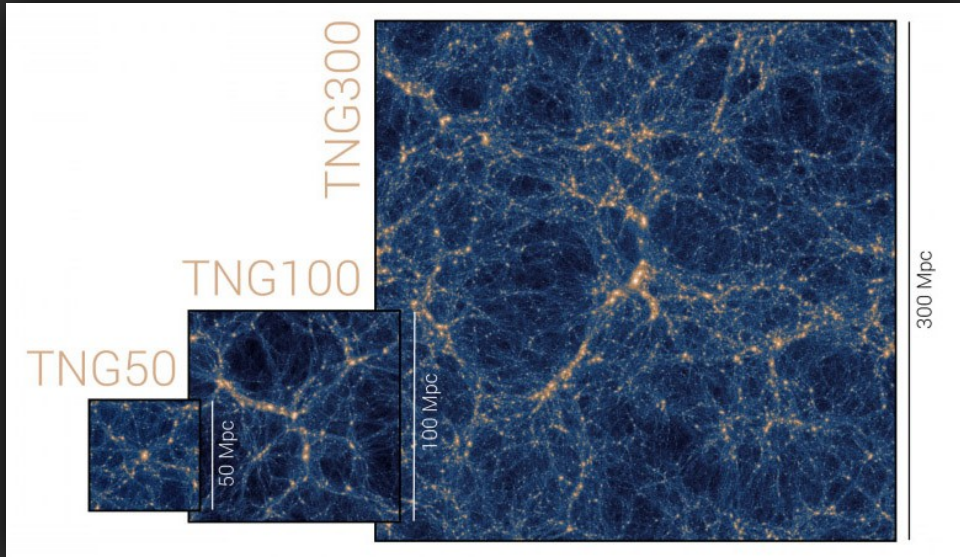


$$\log_{10}\left(\frac{R_d}{1 \text{ kpc}}\right) = H + \xi \log_{10}\left(\frac{M_{star}}{10^{10} M_{sun}}\right), H \approx 0,4, \kappa \approx 0,6$$



$$\log_{10}\left(\frac{SFR}{M_{sun} \text{ year}^{-1}}\right) = Y + \kappa \log_{10}\left(\frac{M_{star}}{10^{11} M_{sun}}\right), Y \approx 1, \xi \approx 0,2$$

Annexe: The *Illustris* TNG simulations

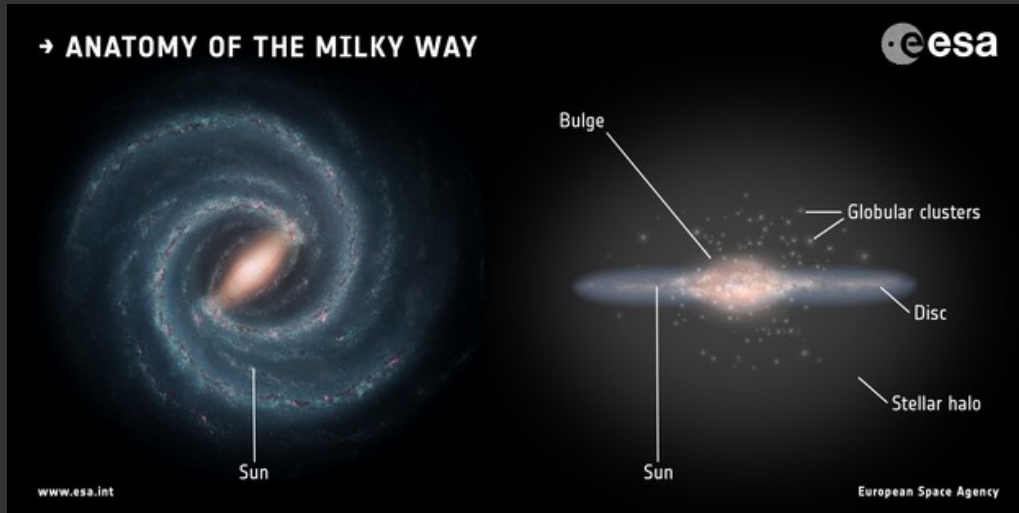


Illustris TNG

- Cosmological simulations of galaxies formation
- Hydrodynamical simulation
- 18 simulations
- Different size of simulated universe
- TNG50: $51,7^3$ Mpc

Reminder: 1pc=3.26 al

Annexe: Substructures of a galaxy.



ESA/NASA

- Different sub-structures:
 - Main galaxy
 - Stellar halo
 - Dark matter halo
- Components:
 - Dark matters
 - Stars
 - Gas
 - Dust

Annexe: optimization of conversion

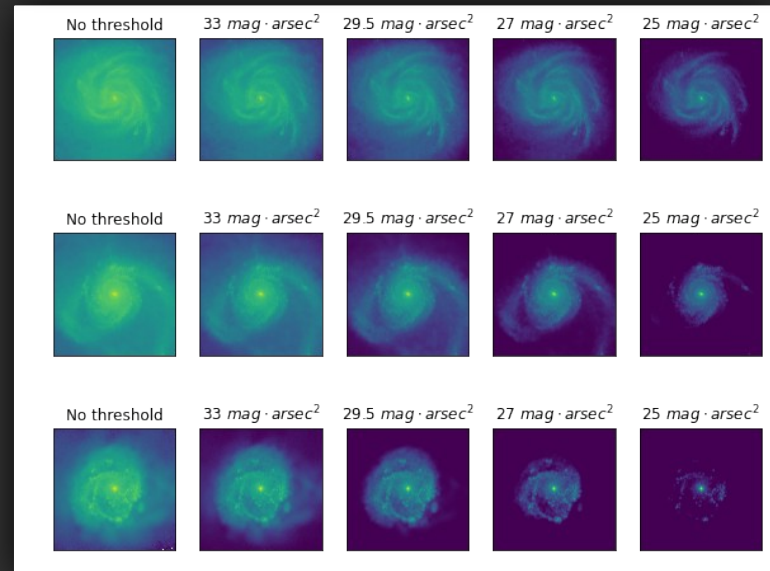
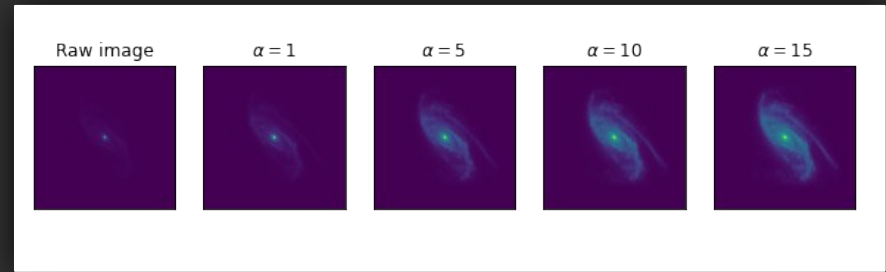
- Conversion:

$$y = \log_{10}(\alpha(x - \beta) + \sqrt{\alpha^2(x - \beta)^2 + 1}), \alpha, \beta \in \mathbb{R}$$

- Contrast and dynamics of images

- Here: $\alpha=10$, $\beta=0$

- Adapt the brightness threshold of the images



Annexe: Glitch and artefacts

