



The *Stargate* collaboration

*past, present and future
of GRB follow-up at ESO/VLT*

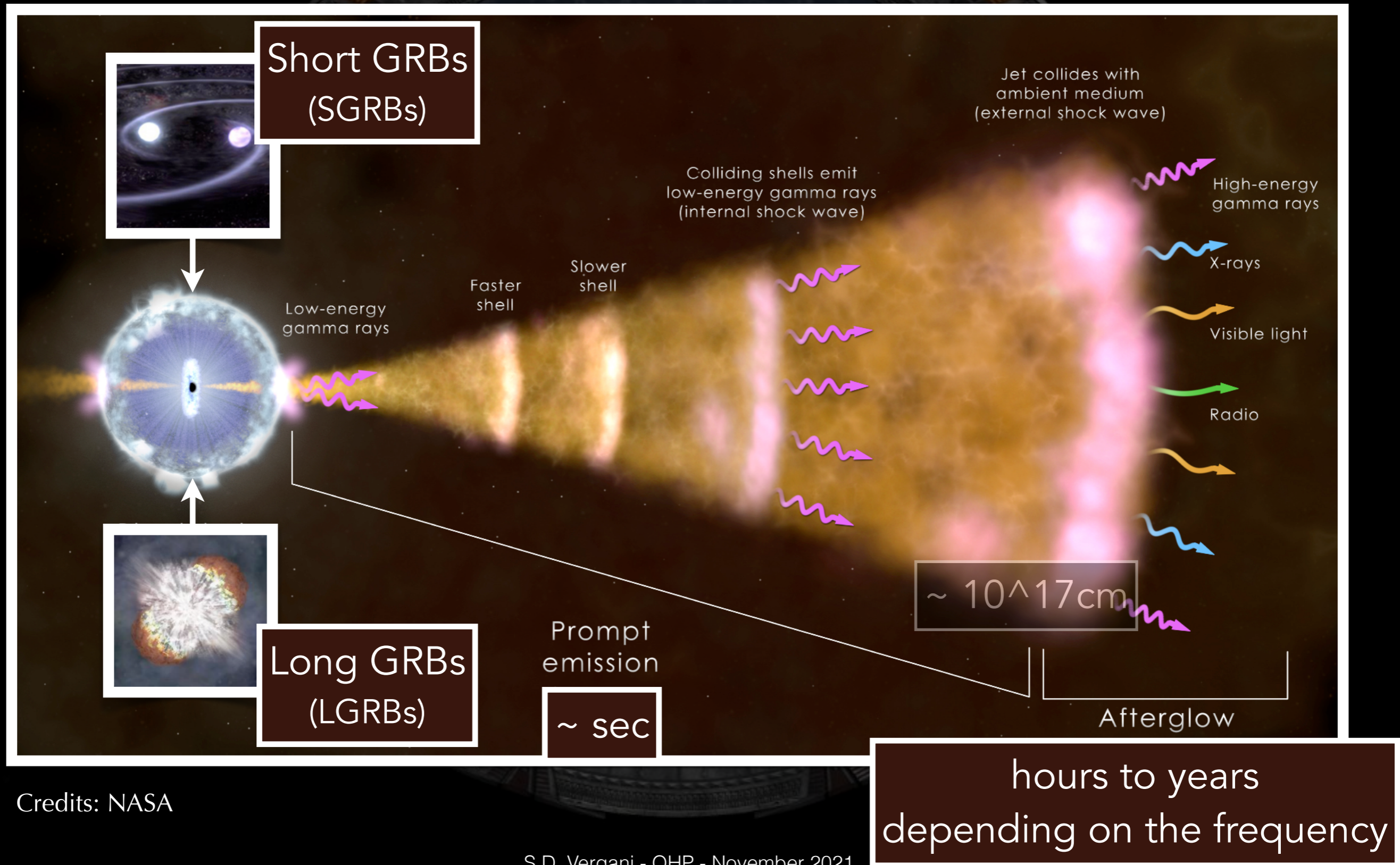


S.D. Vergani



Les Sursauts gamma (GRBs)

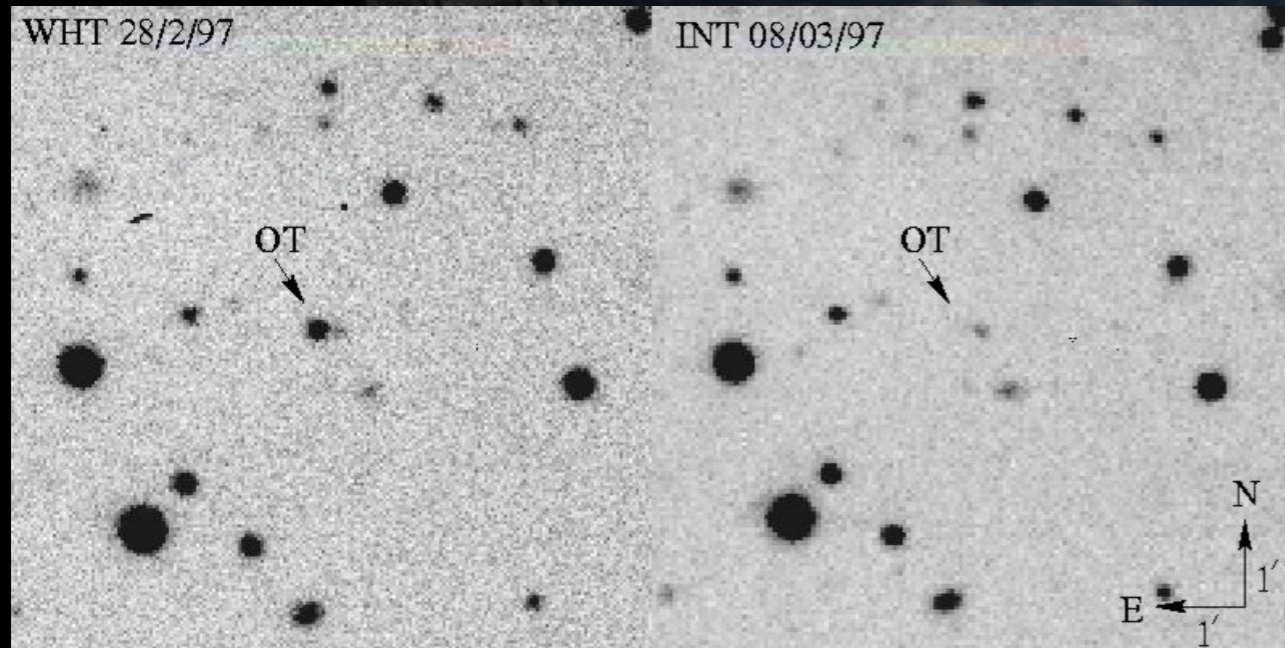
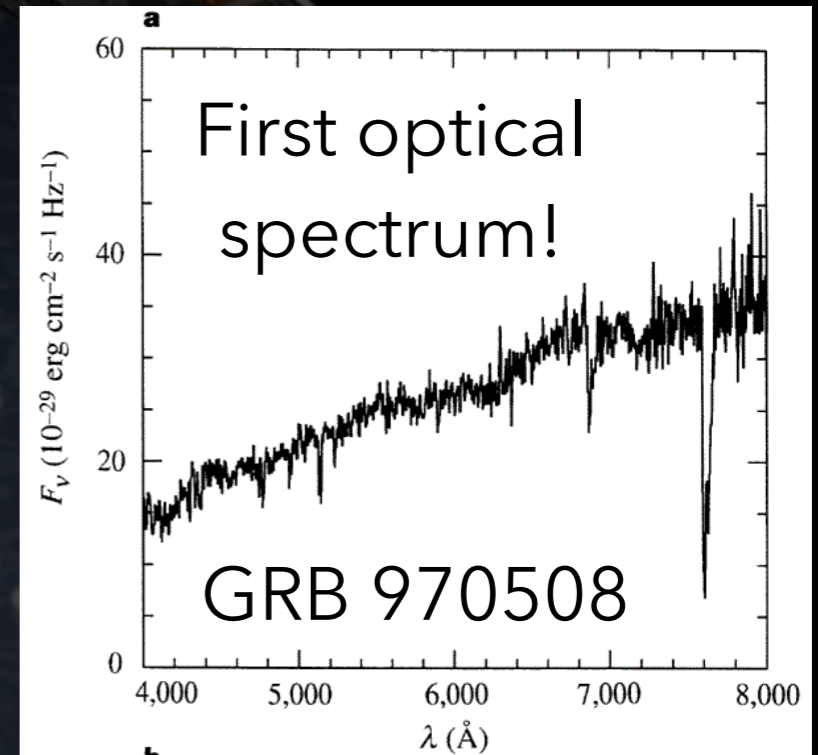
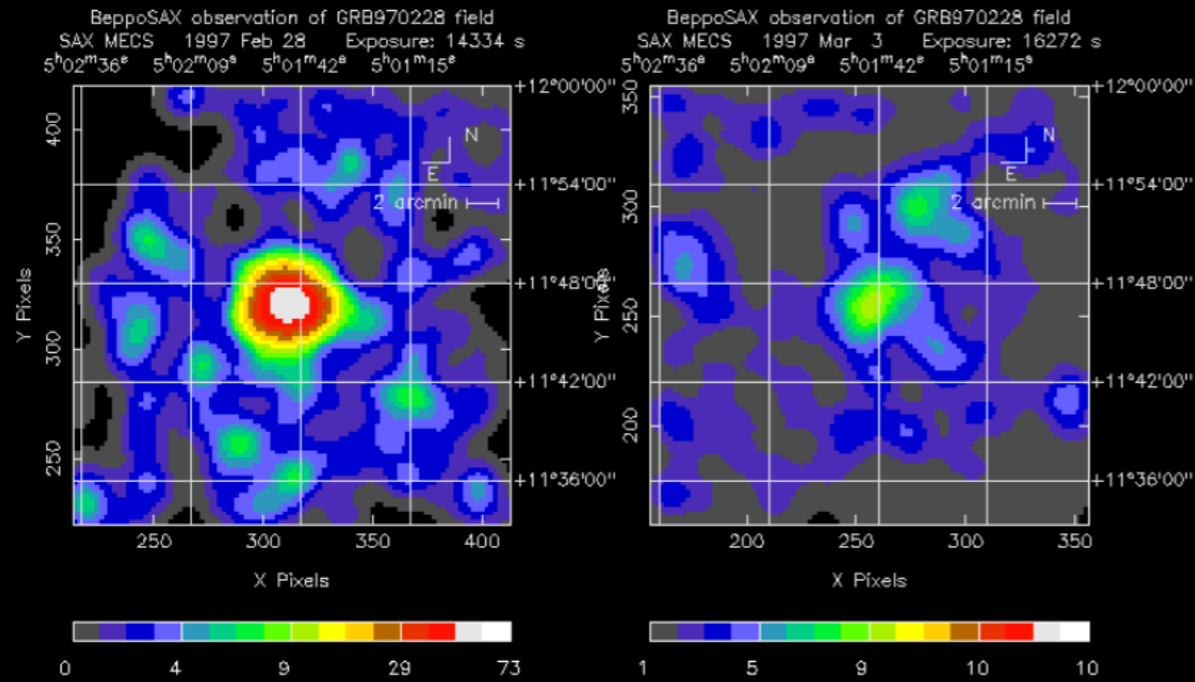
jets ultra-relativistes associés à la formation cataclysmique de trous noirs



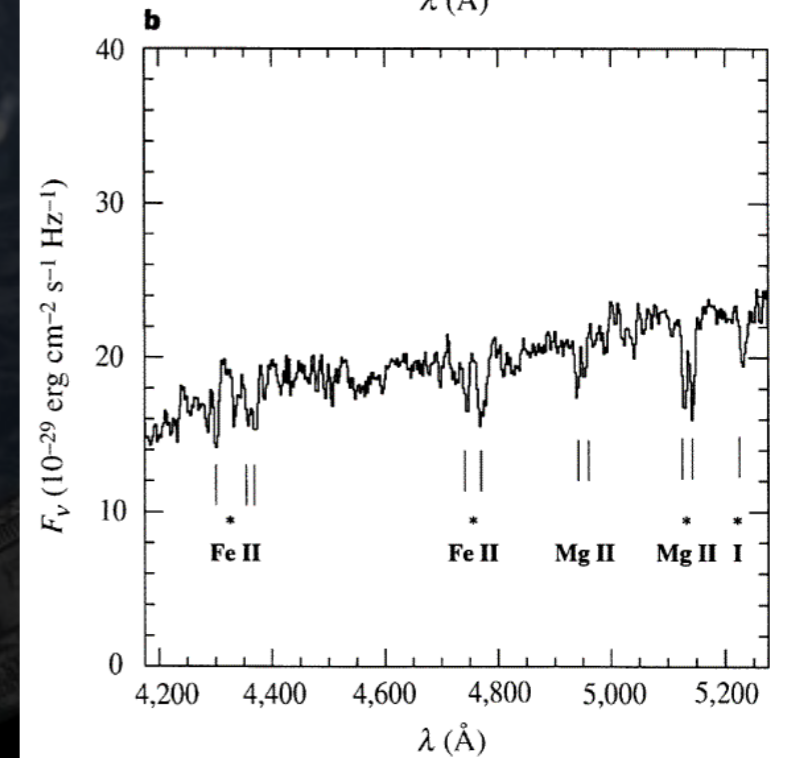
First afterglows

GRB970228 - Costa+97 - BeppoSAX

Metzger et al. 1997



GRB970228 - Groot+97





Very Large Telescope

1998: first light

1999: offered to the community

Target of Opportunity (ToO)



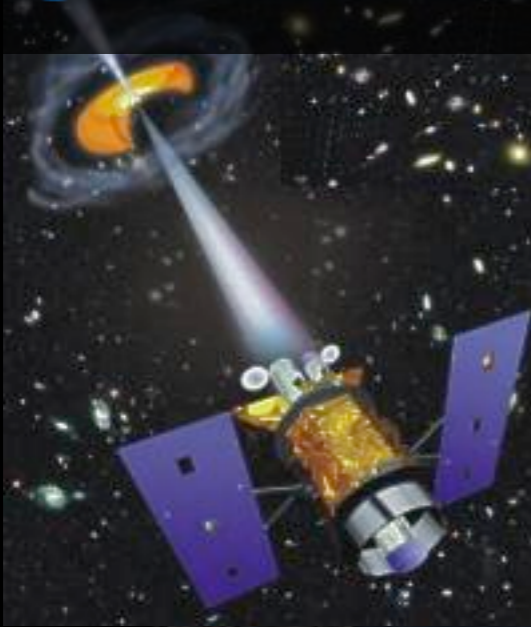
Very Large Telescope

NO *Swift* → No rapid (X-ray) afterglow detection → position

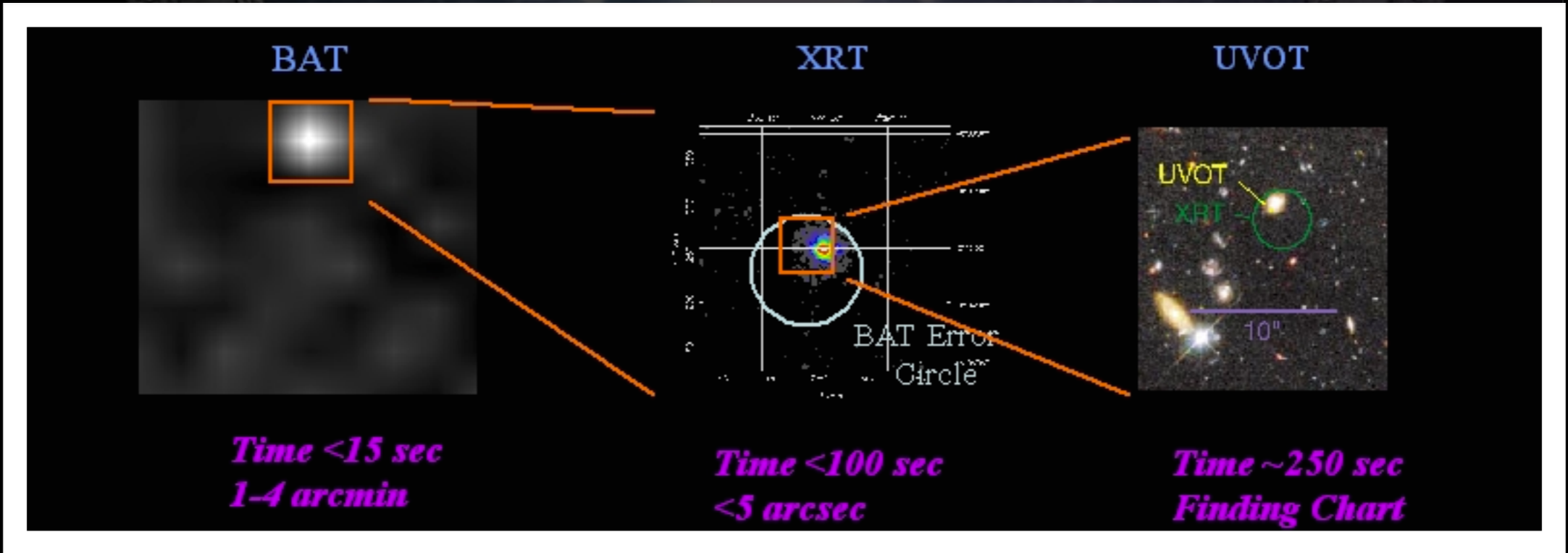
NO robotic telescopes

2004

Swift

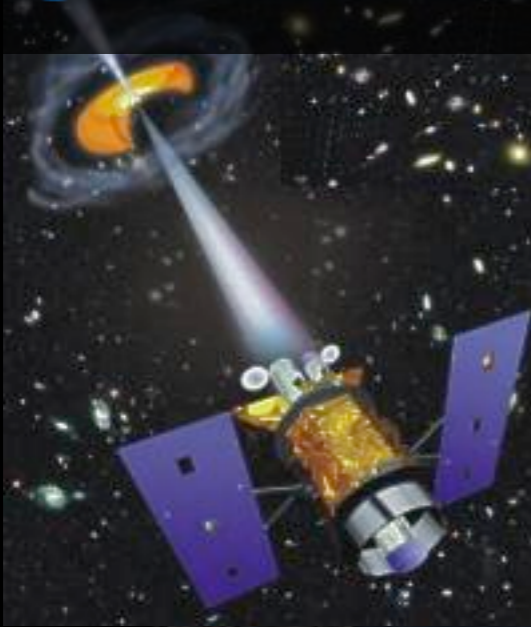


+



2004

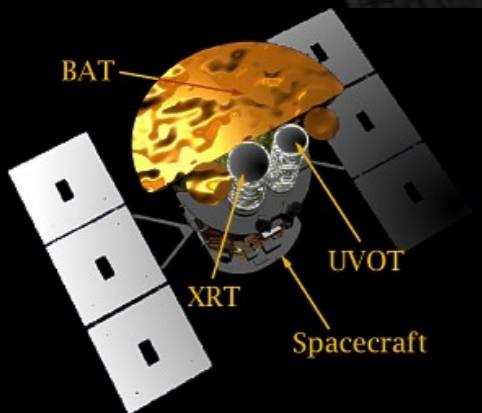
Swift



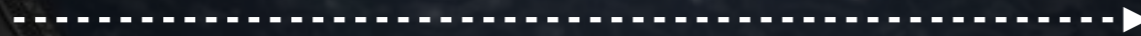
+



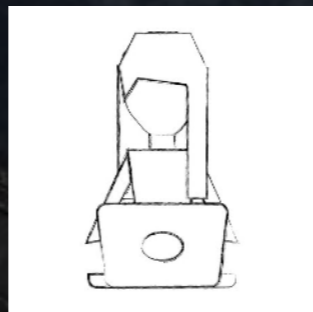
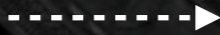
Alert



RRM (few minutes)



Sec



ToO (hours)



Observations



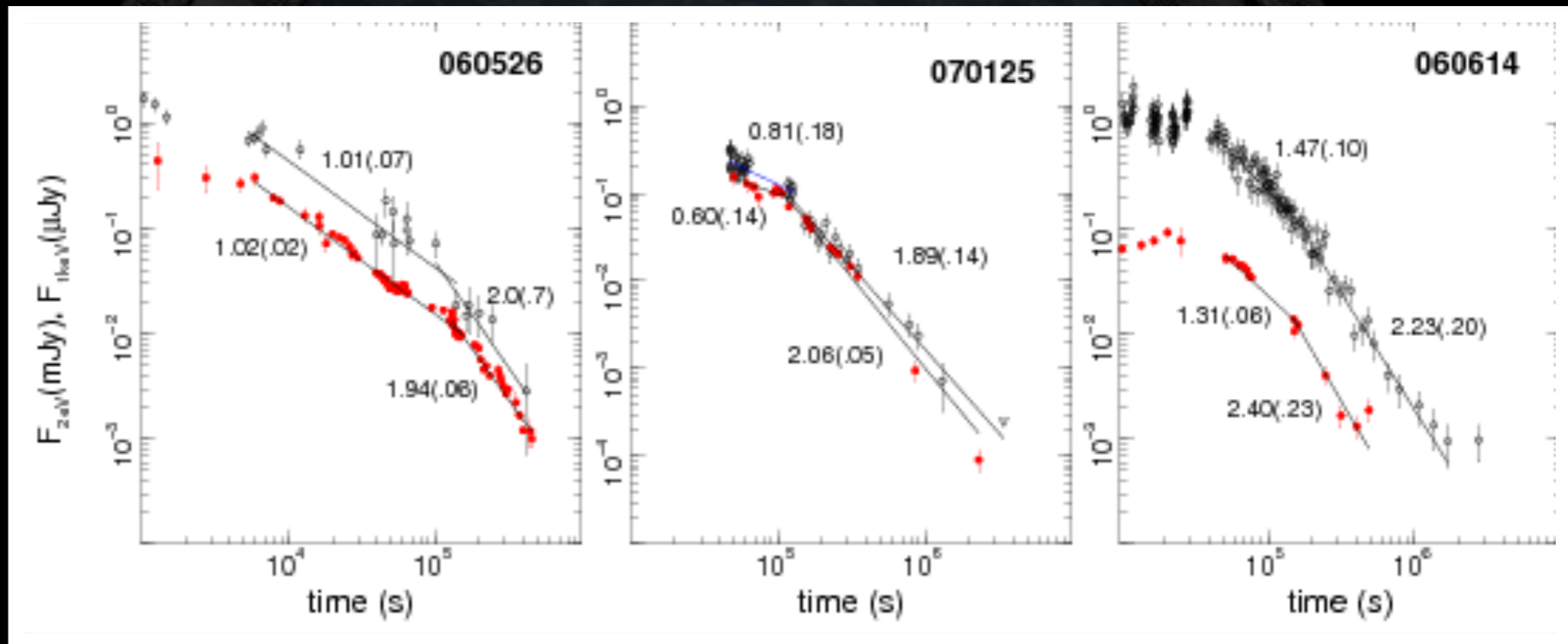


MISTICI (linked to *Swift*)
GRACE (the others)

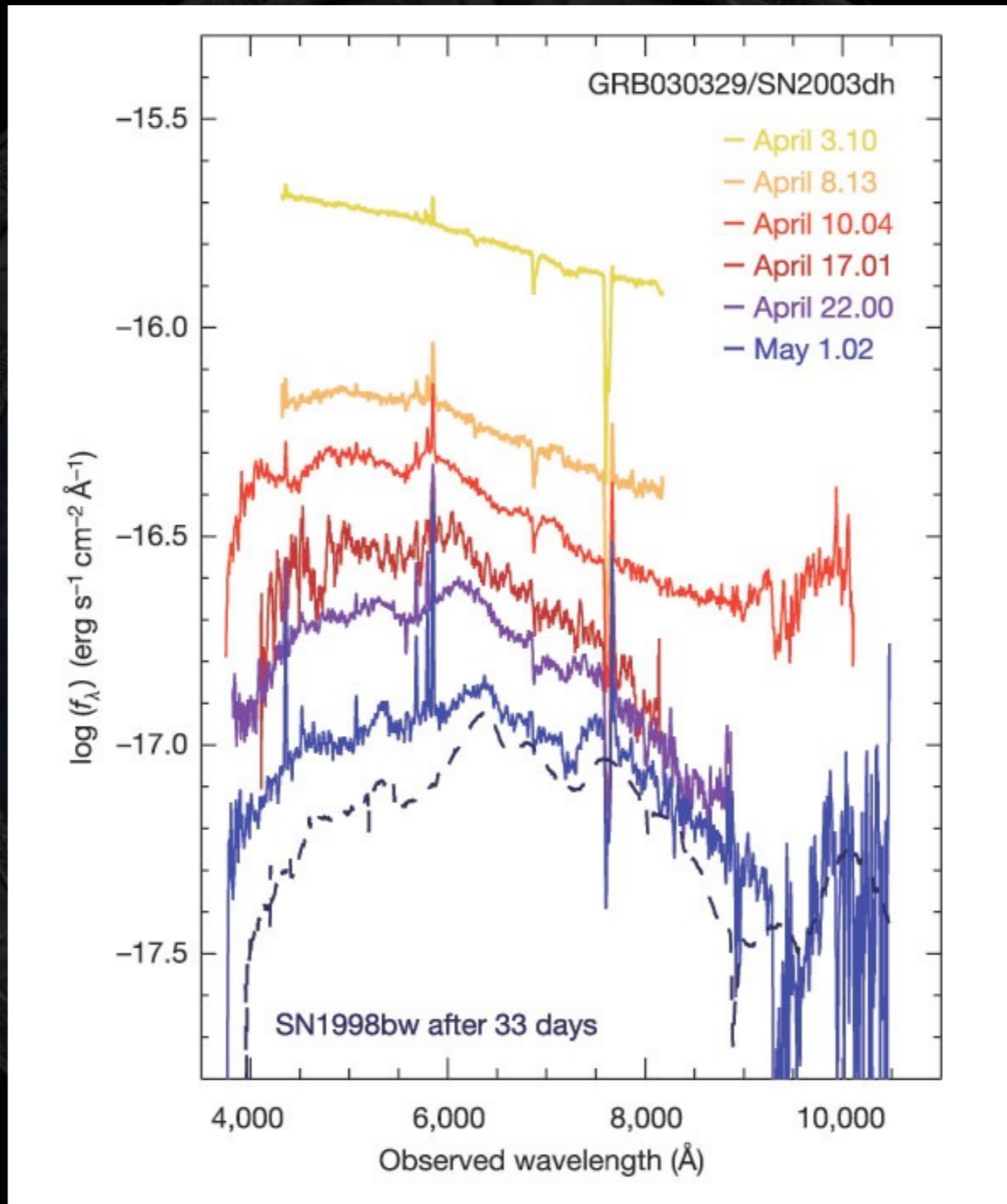
ESO experimented many policies :)

Photometry: Deep observations to look for jet breaks

Panaitescu et al. 2009



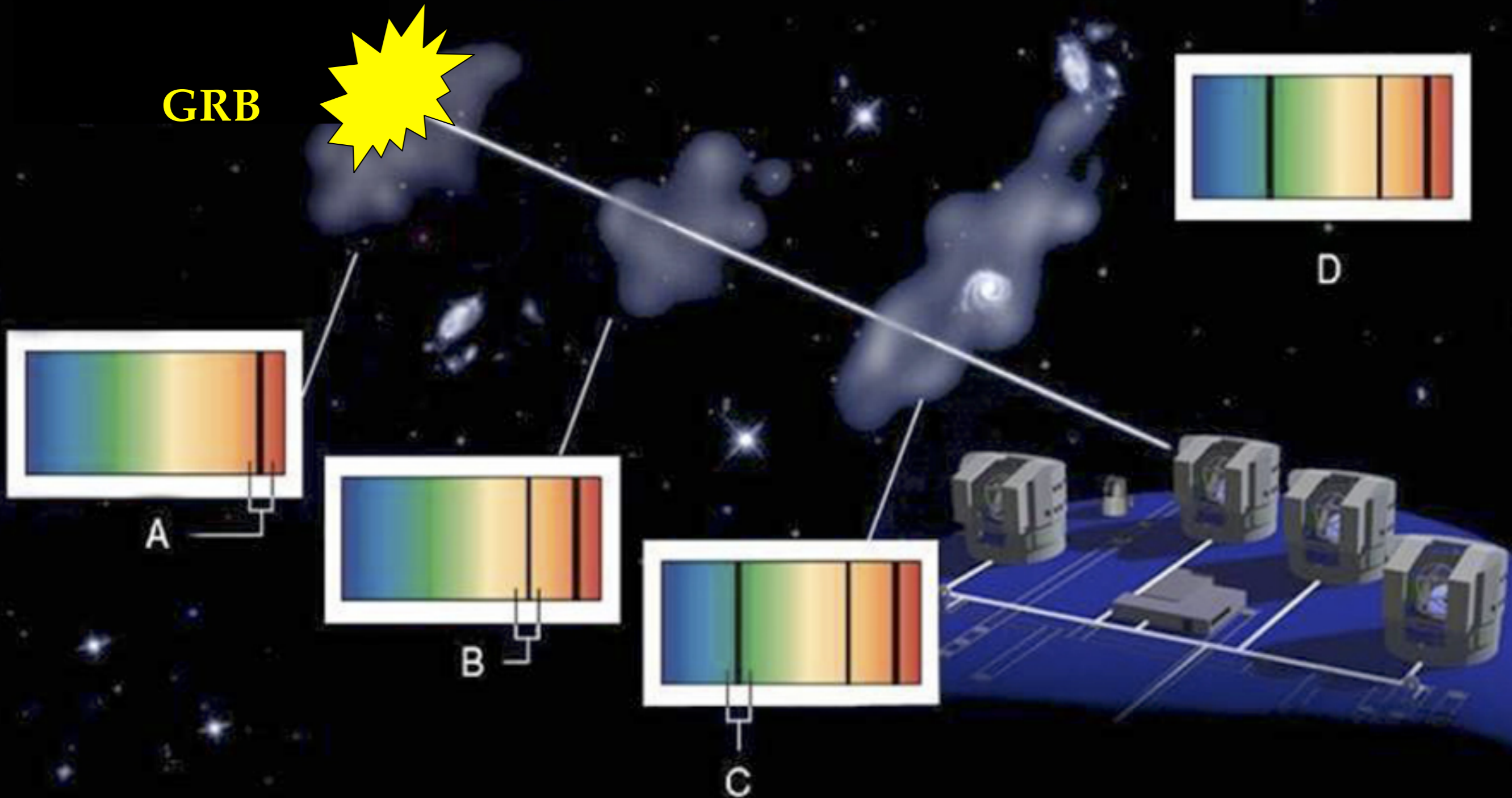
SN spectroscopy for cosmological GRBs



Hjorth+2003

The importance of GRB afterglow spectroscopy

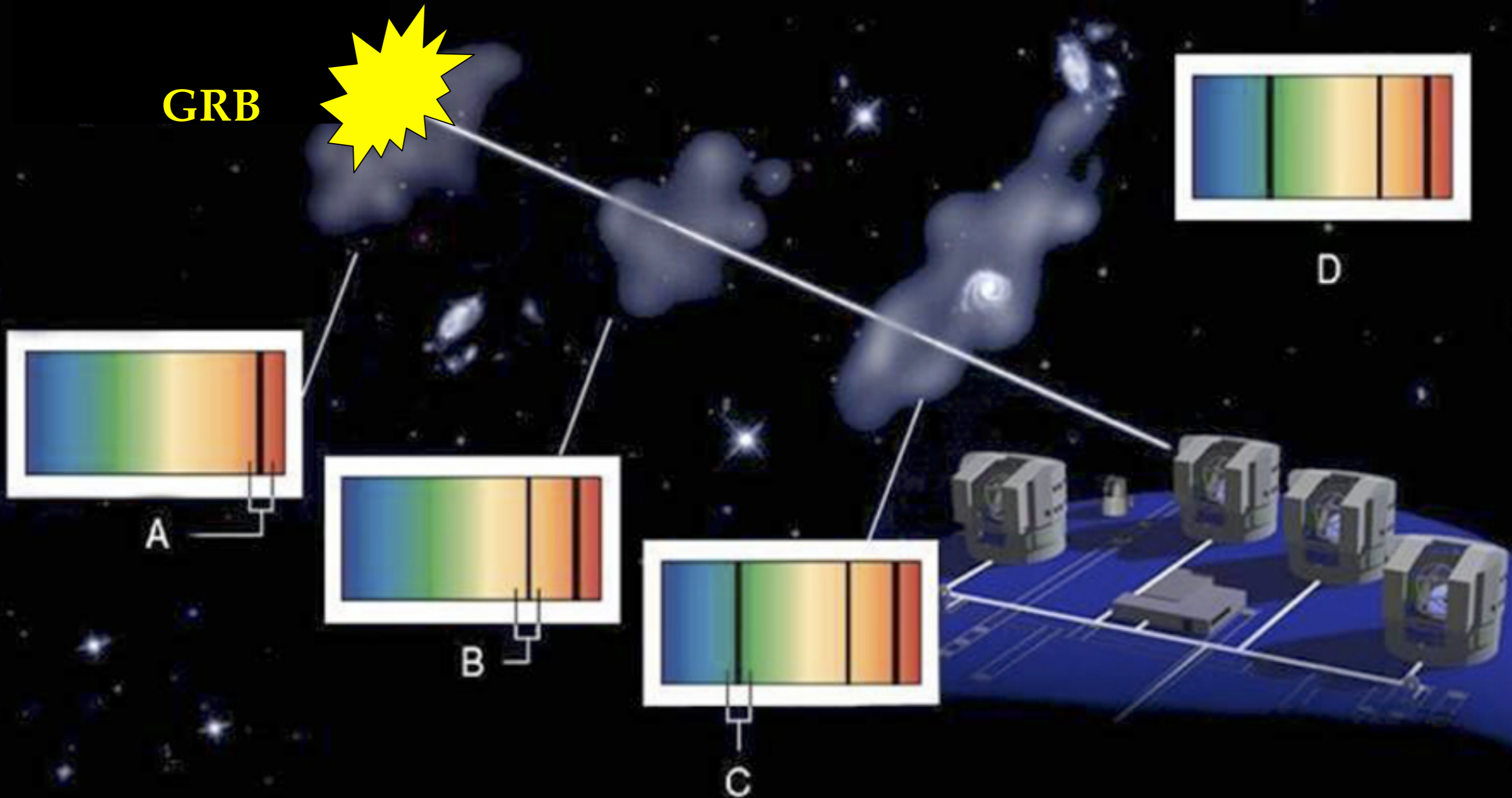
GRB



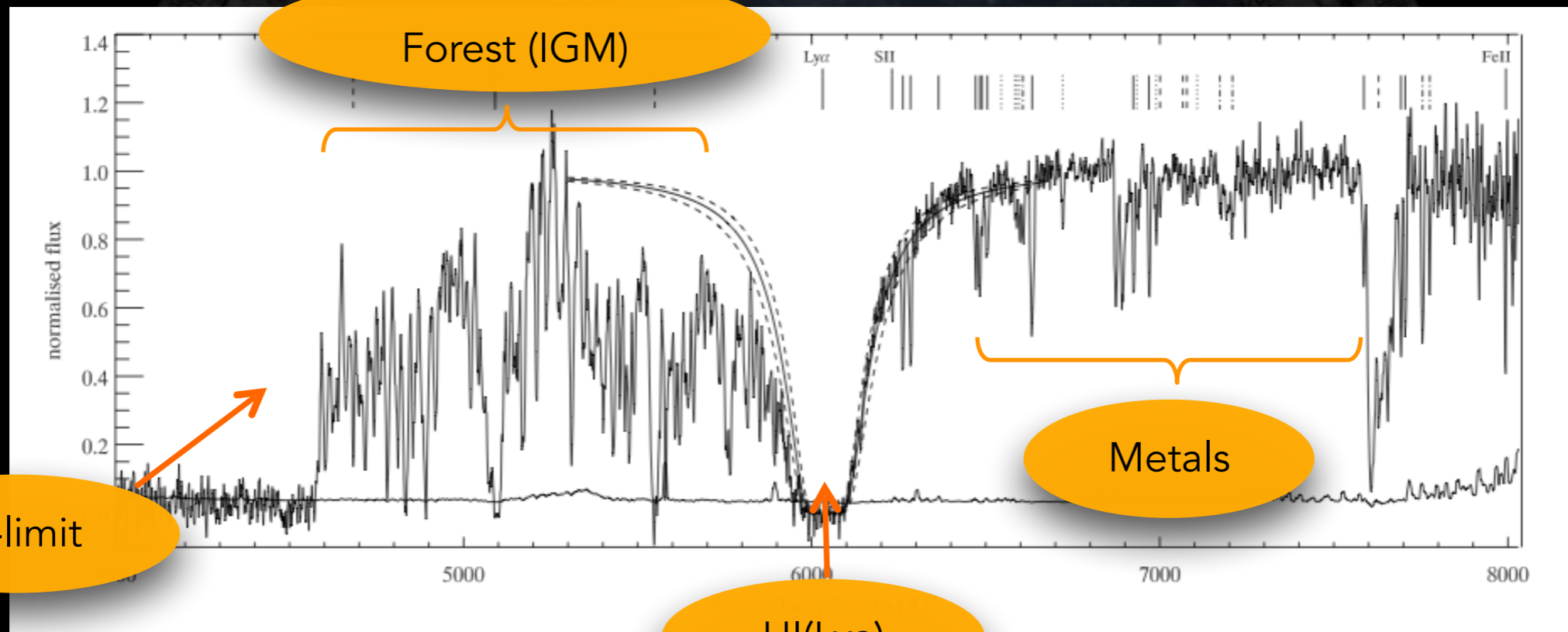
The importance of GRB afterglow spectroscopy

GRB redshift

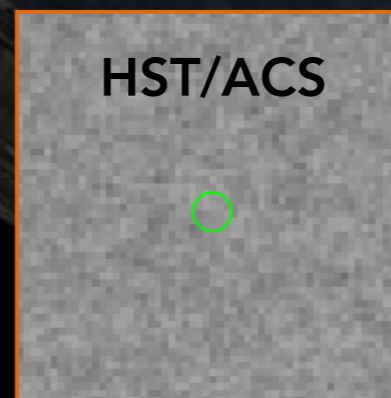
GRB



The importance of GRB afterglow spectroscopy



Courtesy of Nial Tanvir

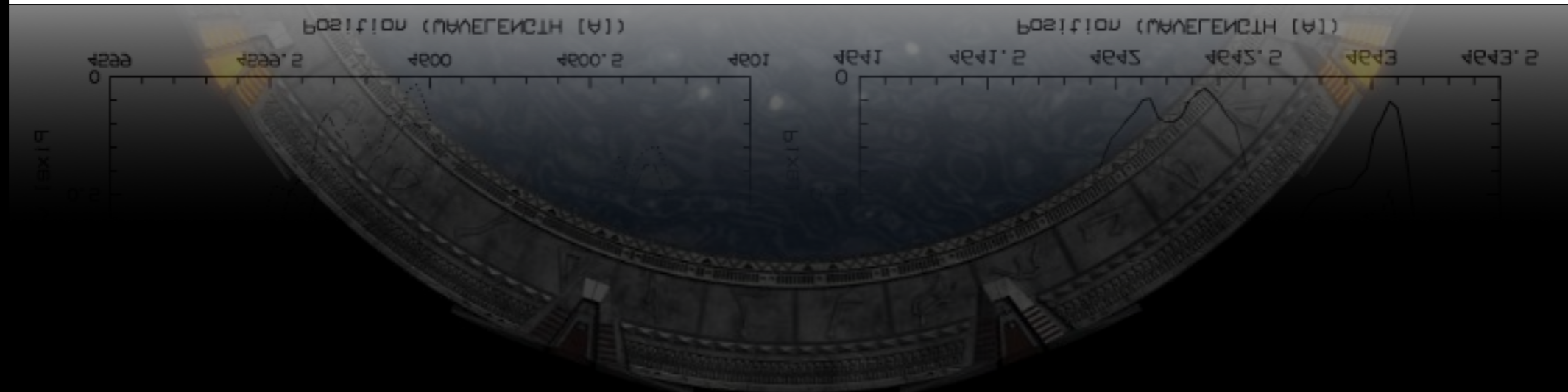
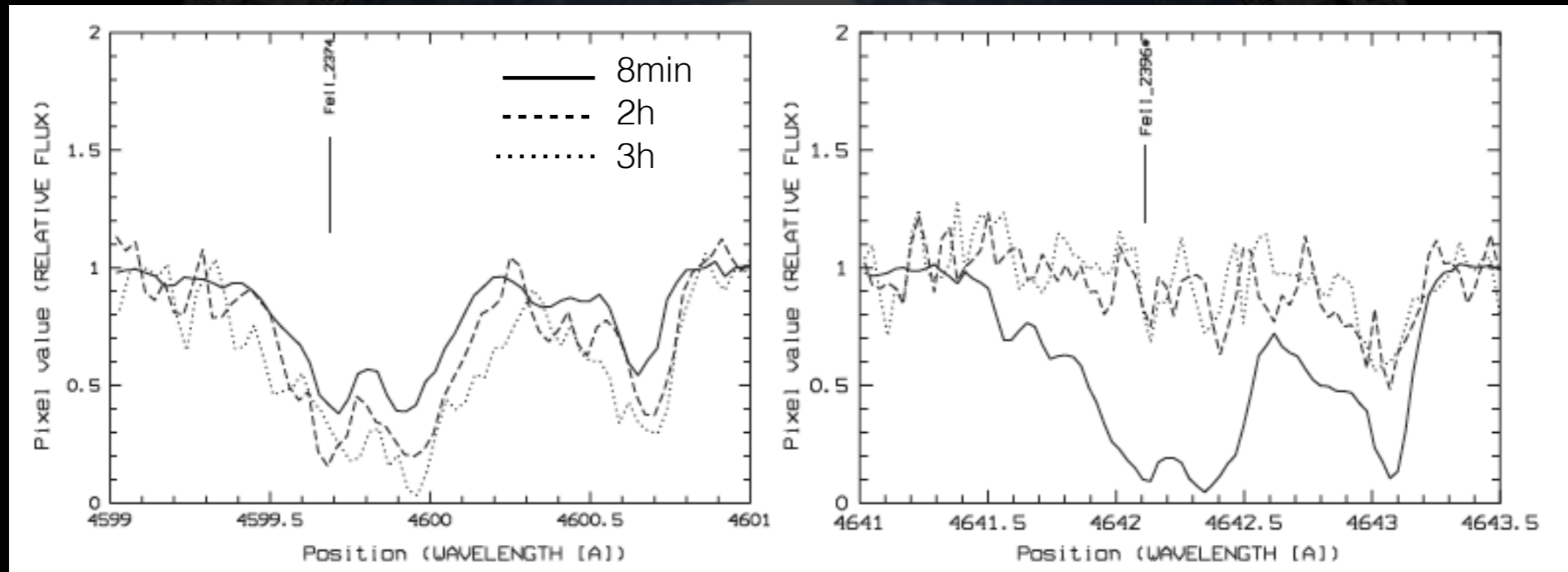


$z=3.97$
 $R>28.5$
Chen+ 2005; Starling+2005

Spectroscopy:

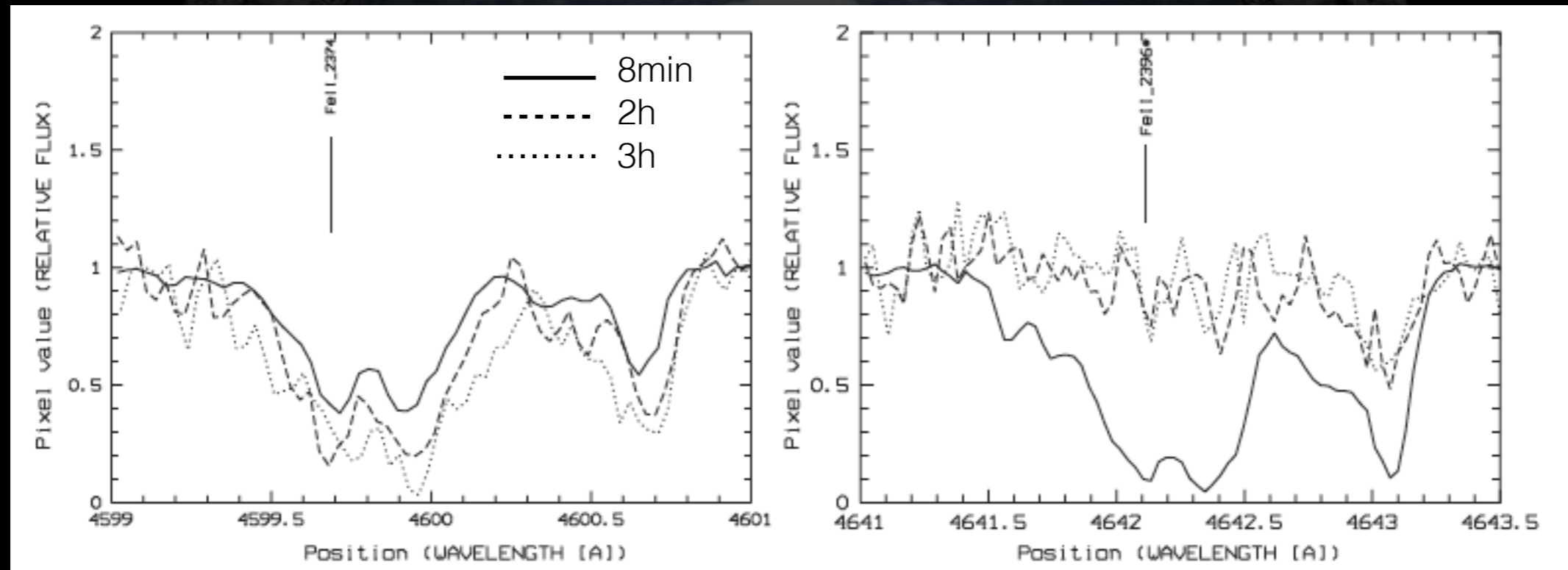
A at early times amazing set of absorption lines
Variation of fine structure lines

D'Elia+09



Variation of fine-structure and meta-stable lines due to UV-pumping caused by the GRB afterglow radiation

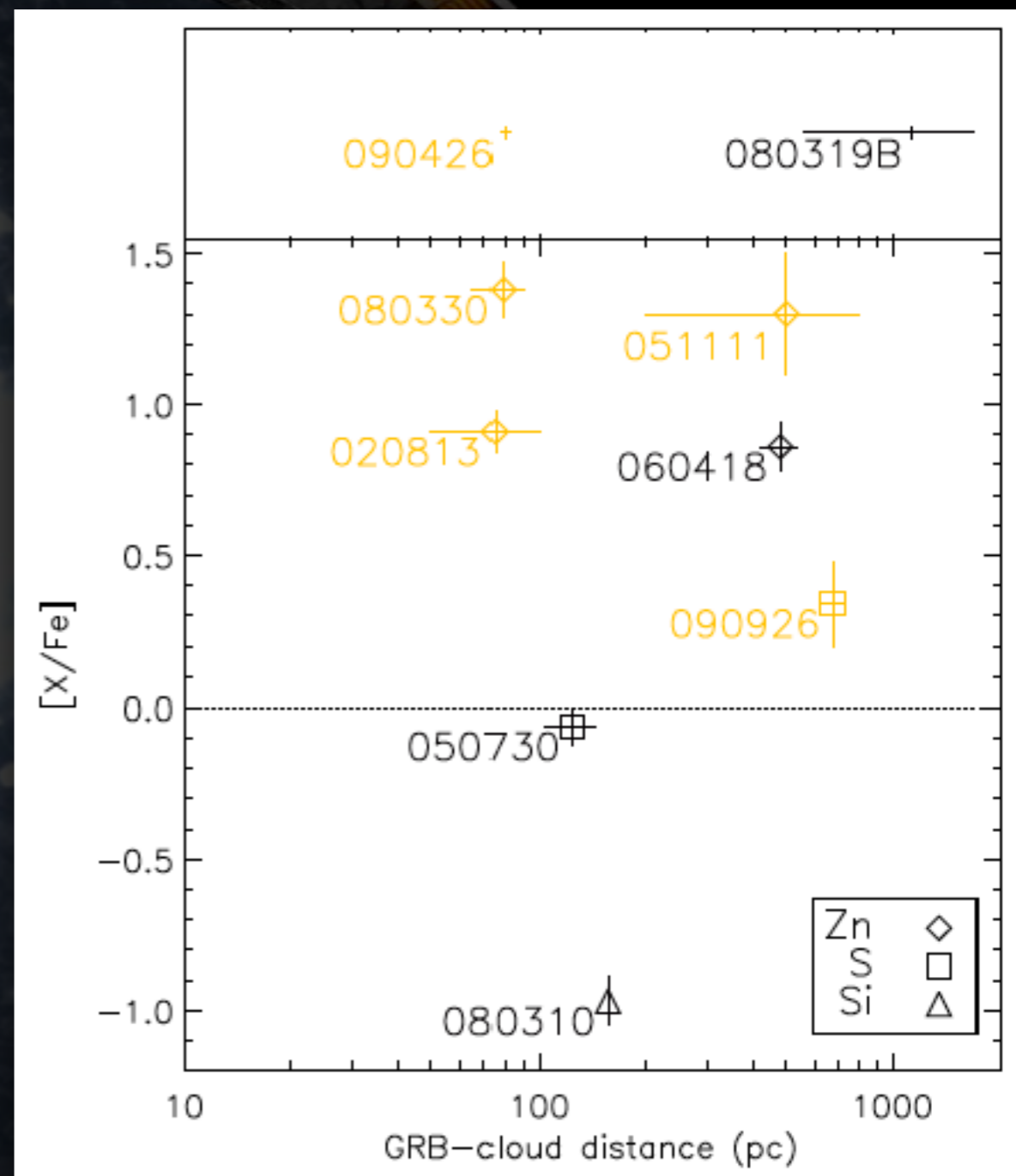
D'Elia+09



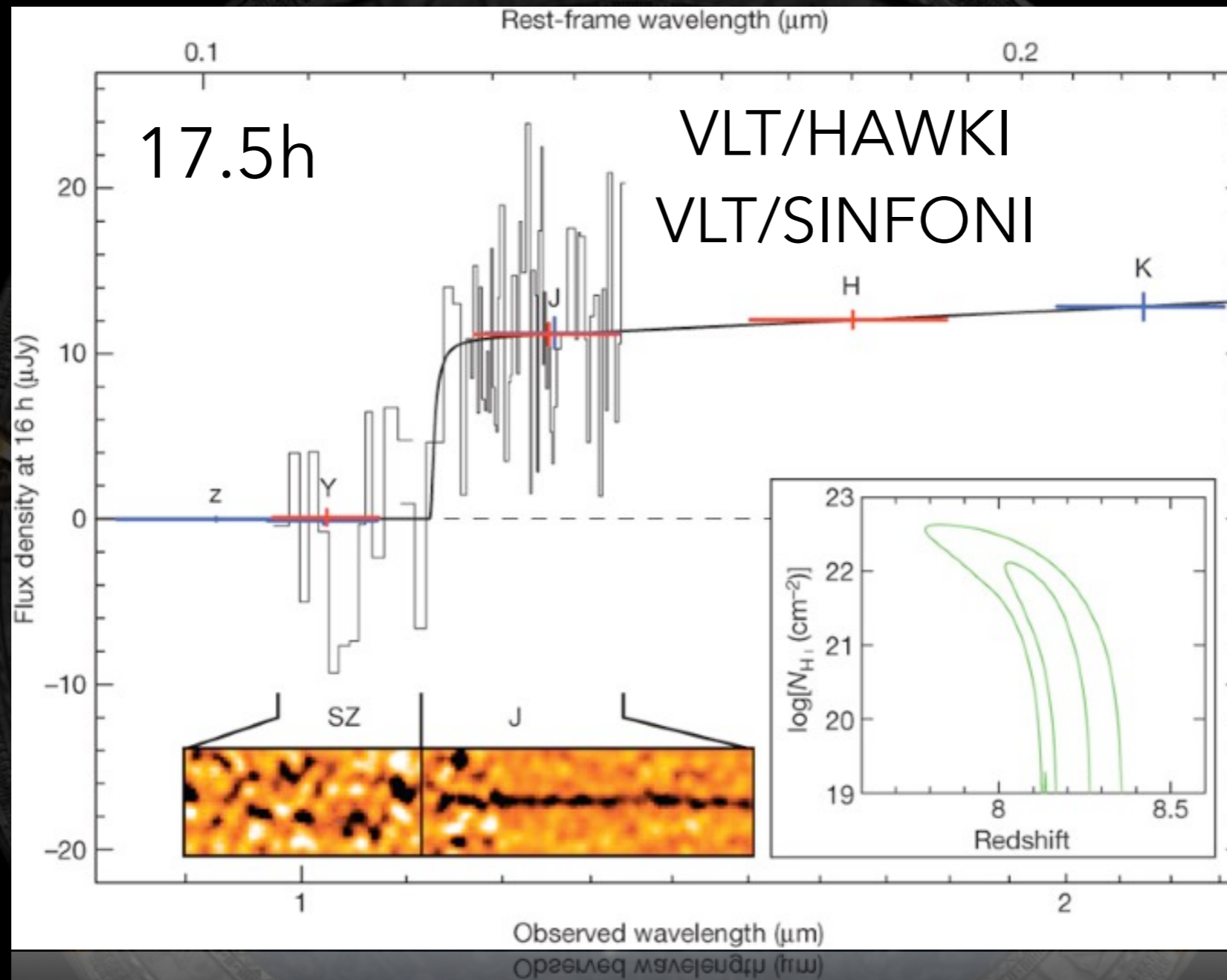
We can determine the distance of the closest clouds

We are probing the star-forming regions
(but not the circumburst environment)

Probing the ISM gas in detail
up to the highest redshift!



GRB090423 @ $z=8.2$ (Tanvir+2009)



+ limited by FORS spectral coverage and UVES sensitivity and coverage



X-Shooter

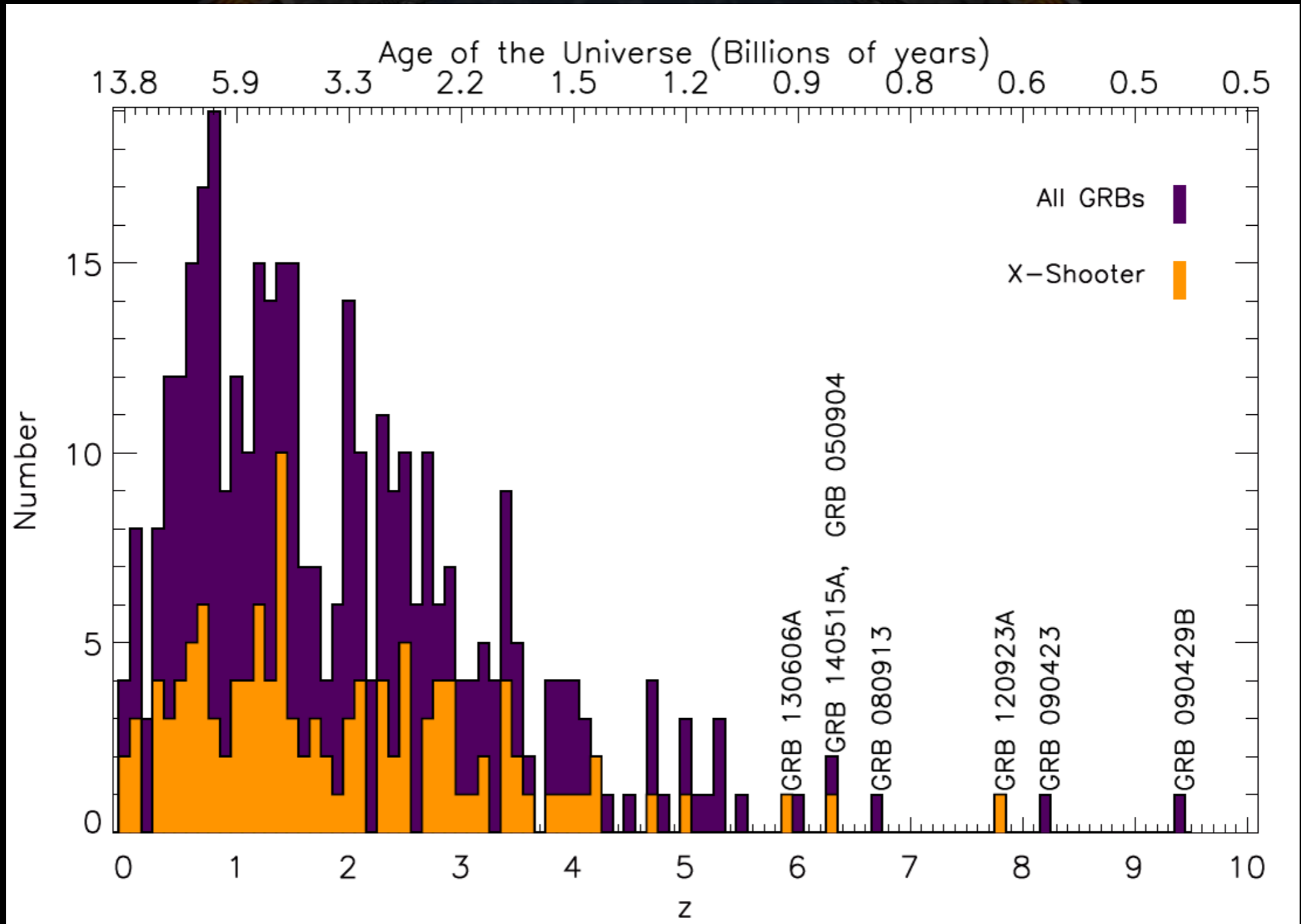
- 3 echelle spectrographs
- Full range 3000 - 24000 Å in a single shot
- Resolution 5000 - 10000
- Slit length 11"
- $m_{AB} \sim 21$ (1h, SNR=10; K \sim 19)



Different groups:

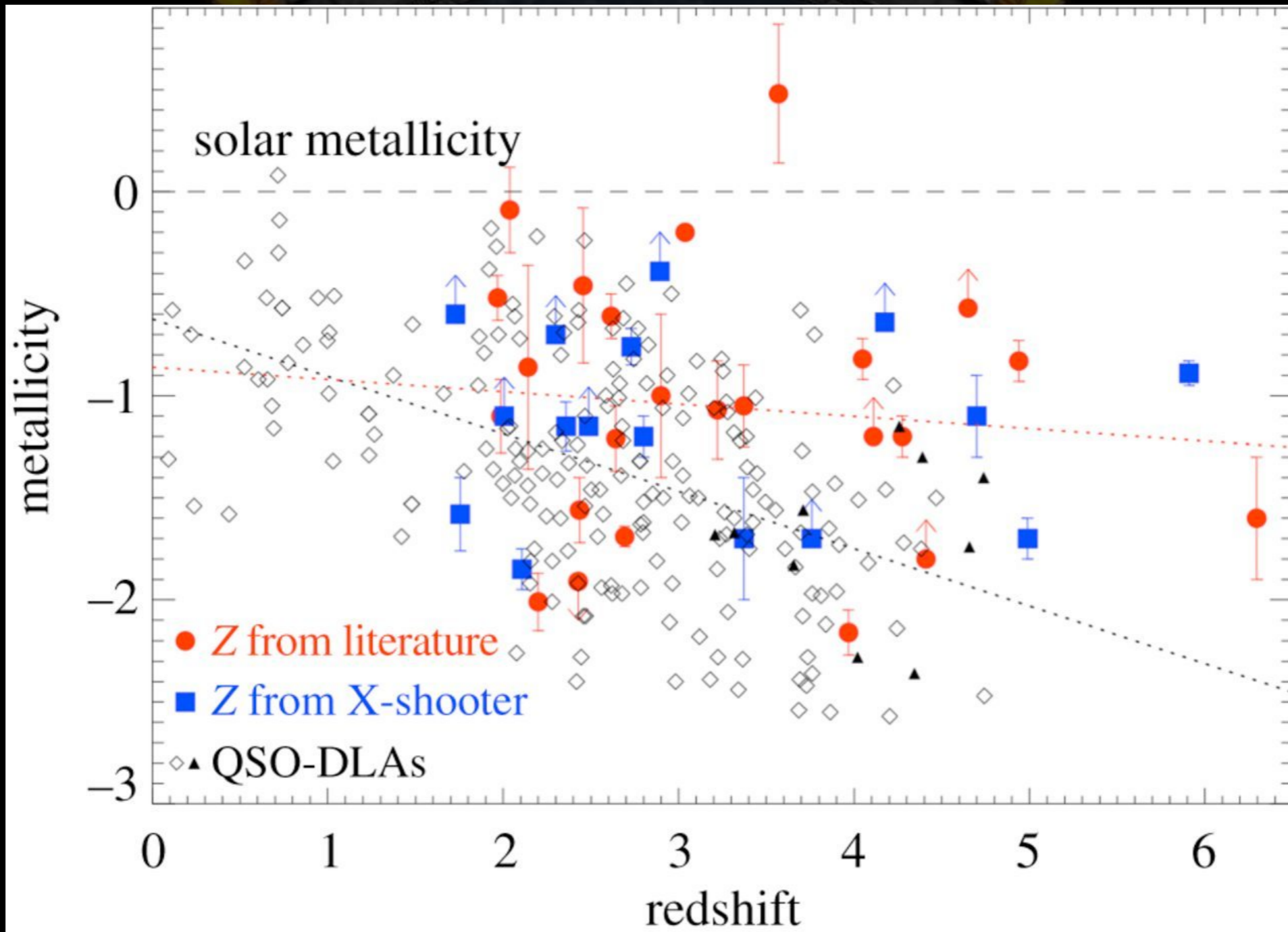
- X-shooter consortium (GTO + ToO/RRM afterglow spectroscopy PI: J. Fynbo)
- short GRBs
- low-z SNe
- other groups

Redshift determination



Courtesy of A. De Cia

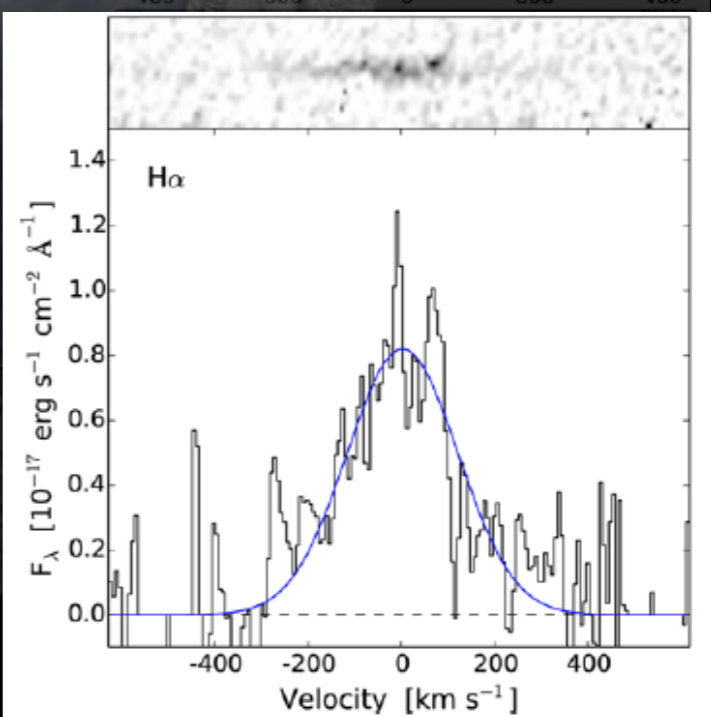
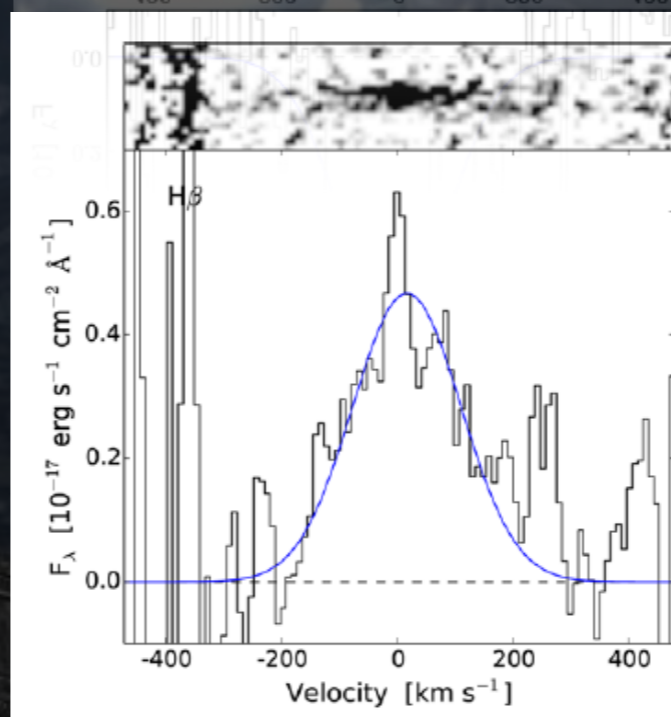
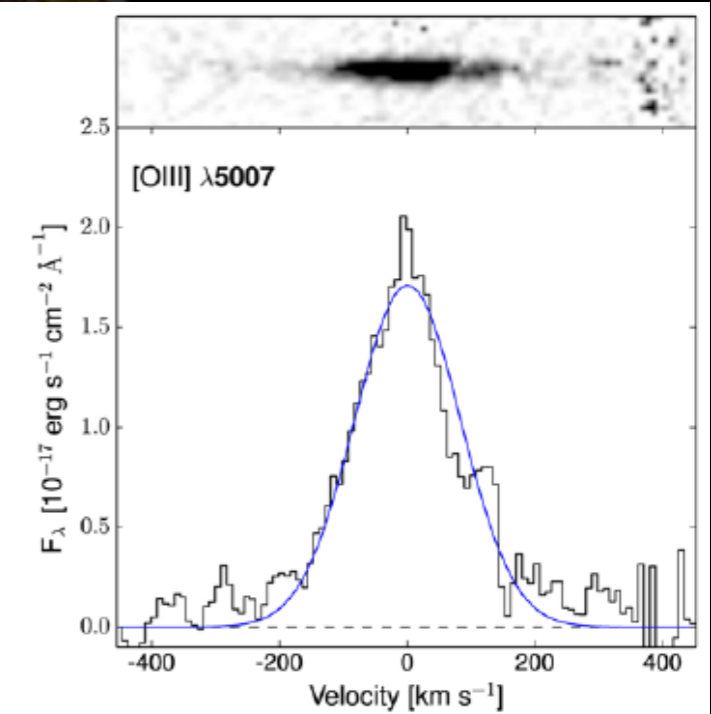
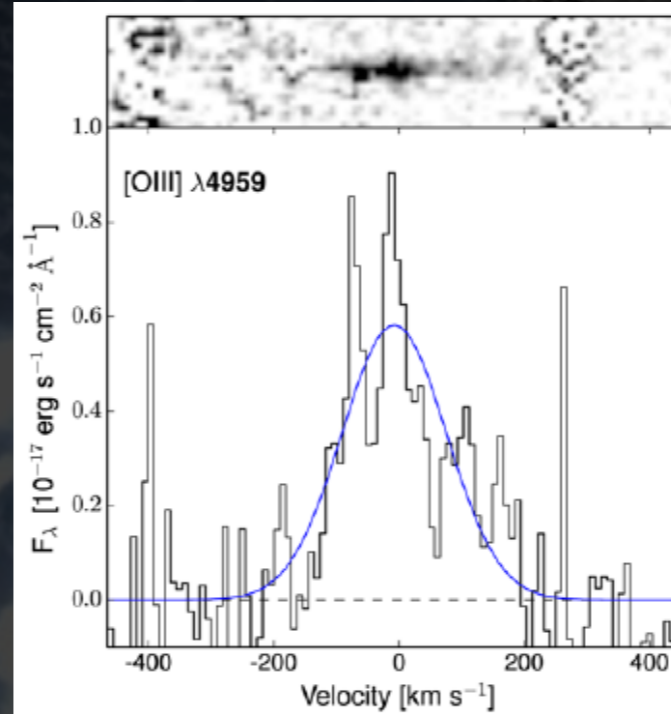
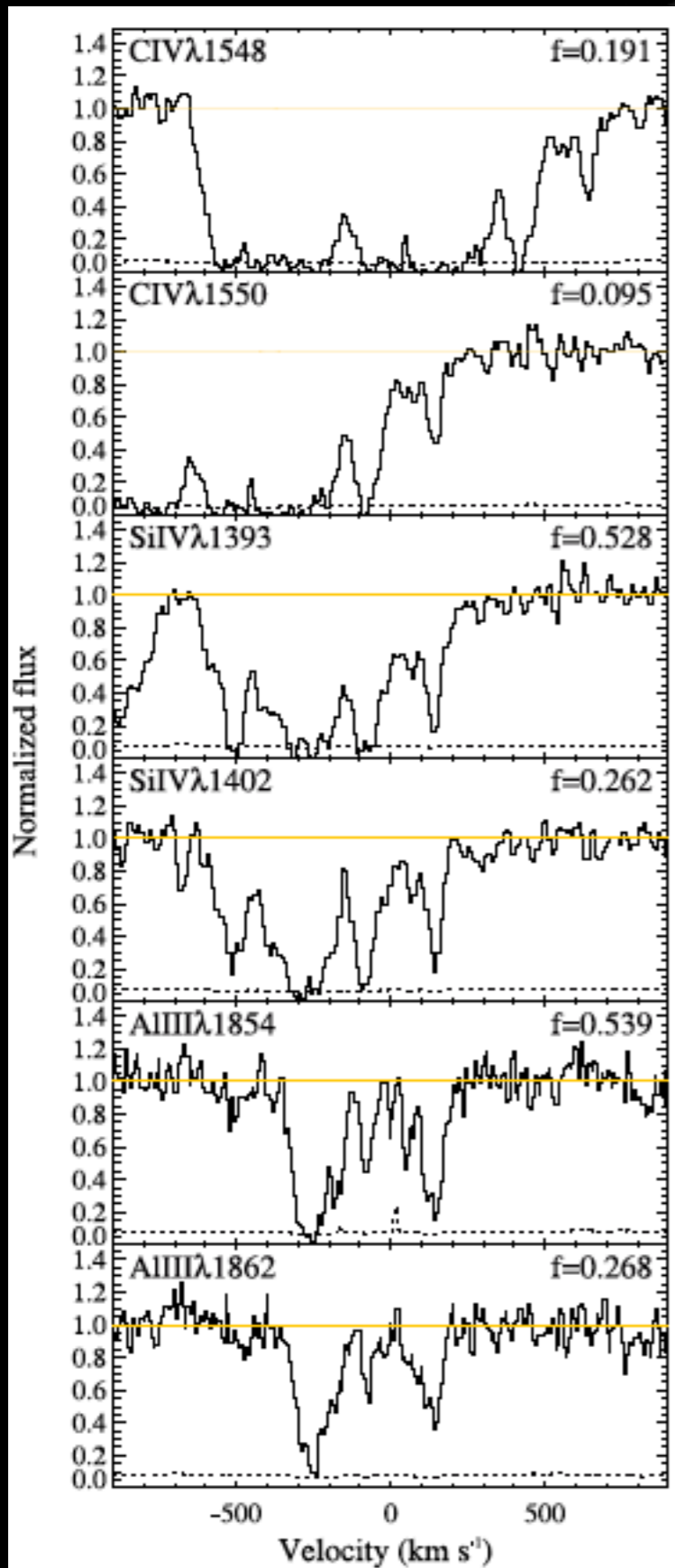
Metallicity



Adapted from Thöne+13

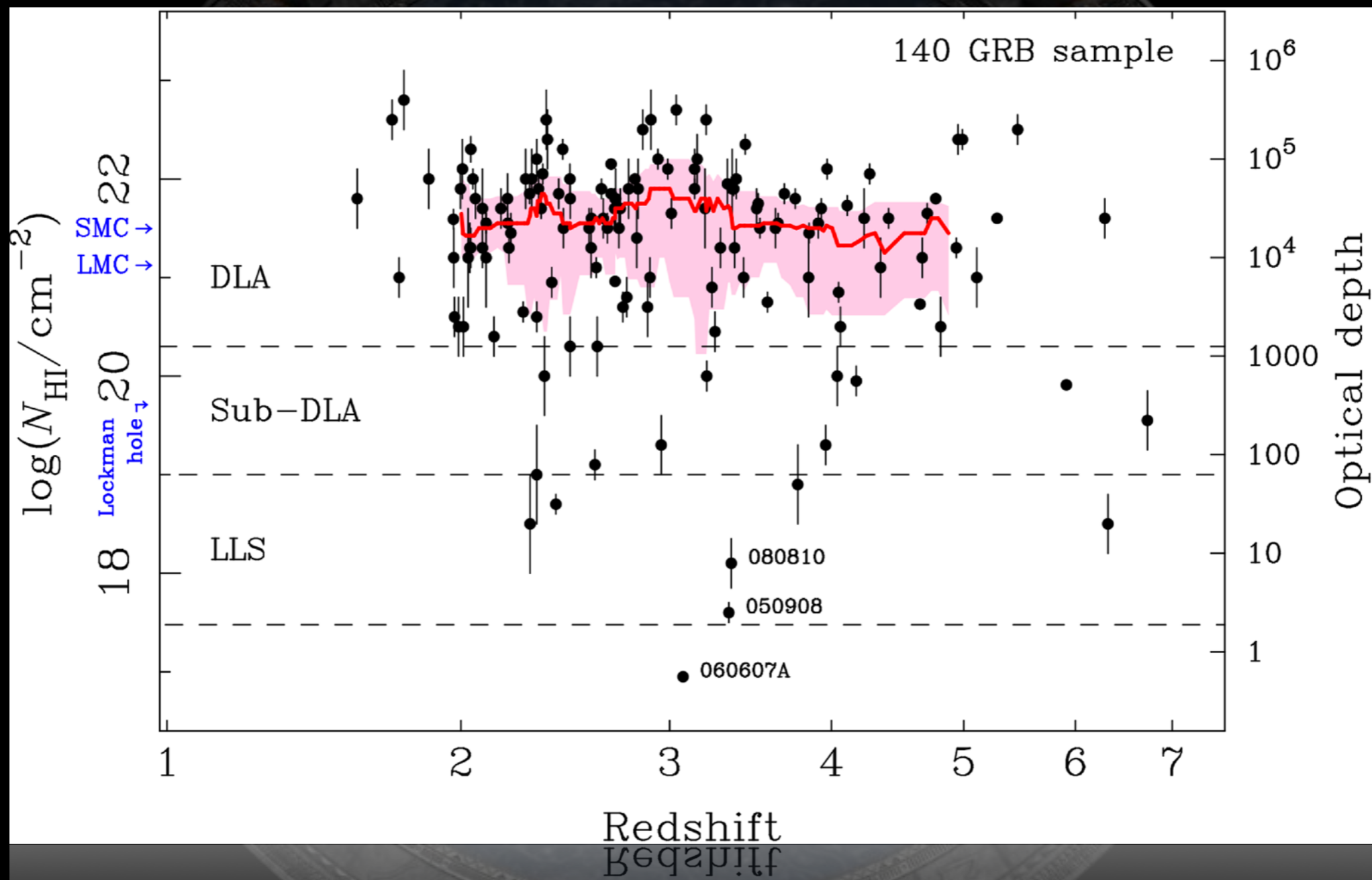
Emission AND absorption lines in the same spectrum

GRB 121024A @ z=2.3



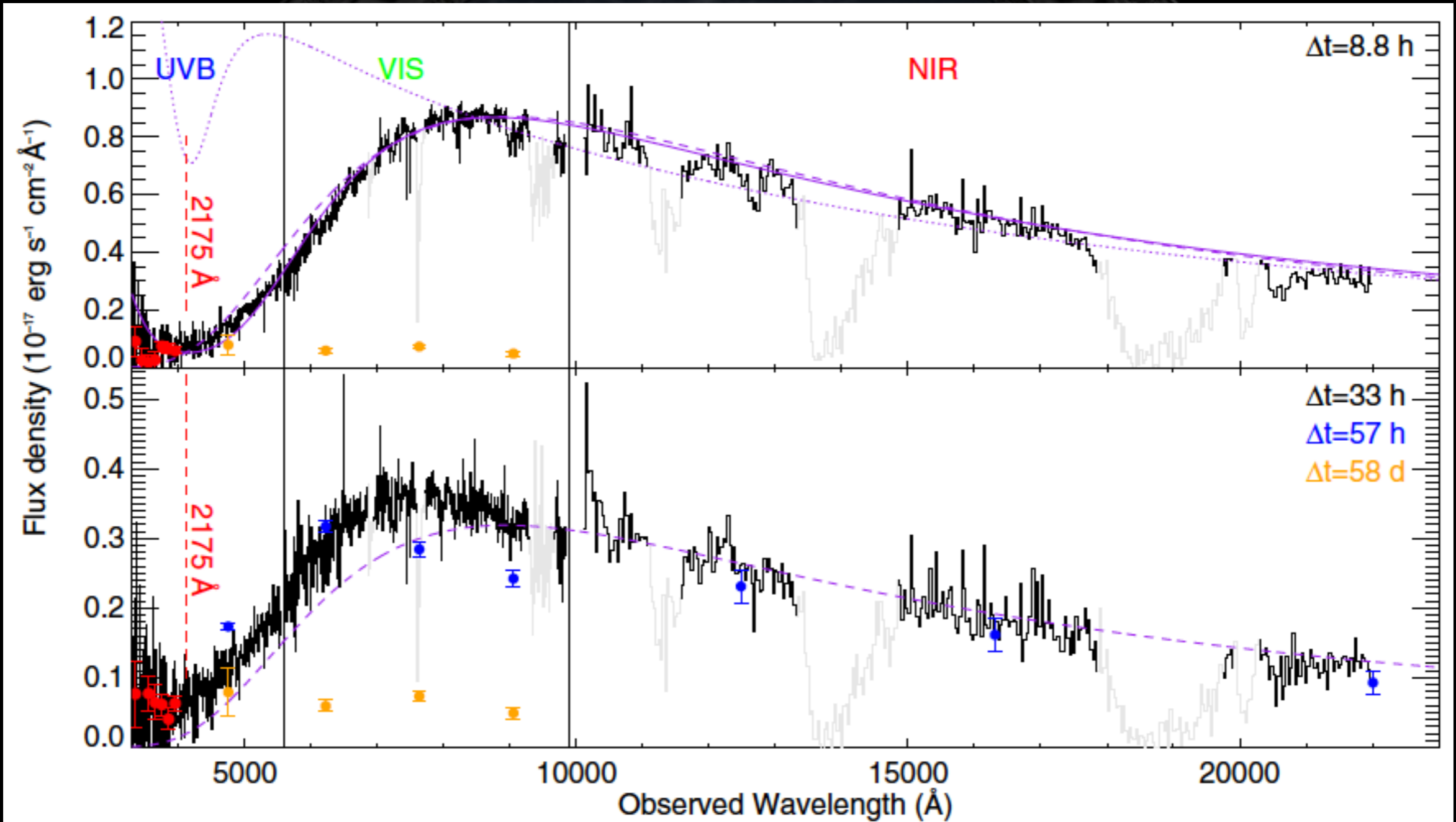
Statistical studies

Tanvir+19



Dust studies

GRB 140506A @ z=0.889



Fynbo+14

2000

10000

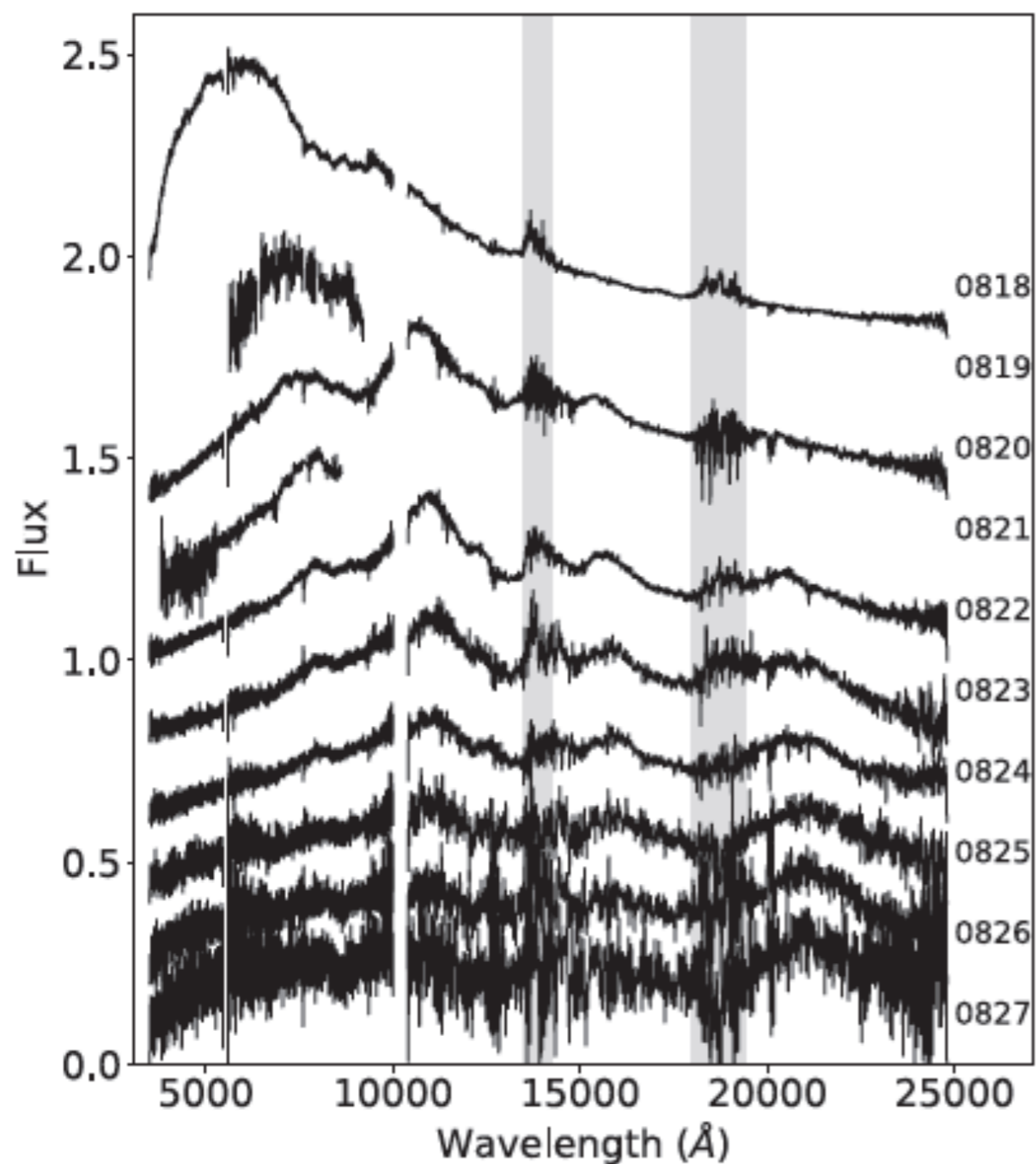
12000

50000

Multi-messenger astrophysics

ESO-VLT/X-Shooter

Pian et al. 2017, Nature



First spectral identification of a KN

- radioactive decay of **r-process nucleosynthesis**
- BNS merger **site for heavy element production in the Universe**

GW170817/GRB170817

Stargate

PIs: N. Tanvir, D. Malesani, S.D. Vergani



Altogether!

- Photometry
- Spectroscopy
- Polarimetry
- long & short GRBs

Low-z SNe separate

Stargate

PIs: N. Tanvir, D. Malesani, S.D. Vergani



Large Programme!
~30h/semester

Altogether!

- Photometry
- Spectroscopy
- Polarimetry
- long & short GRBs

Stargate

PIs: N. Tanvir, D. Malesani, S.D. Vergani



Altogether!

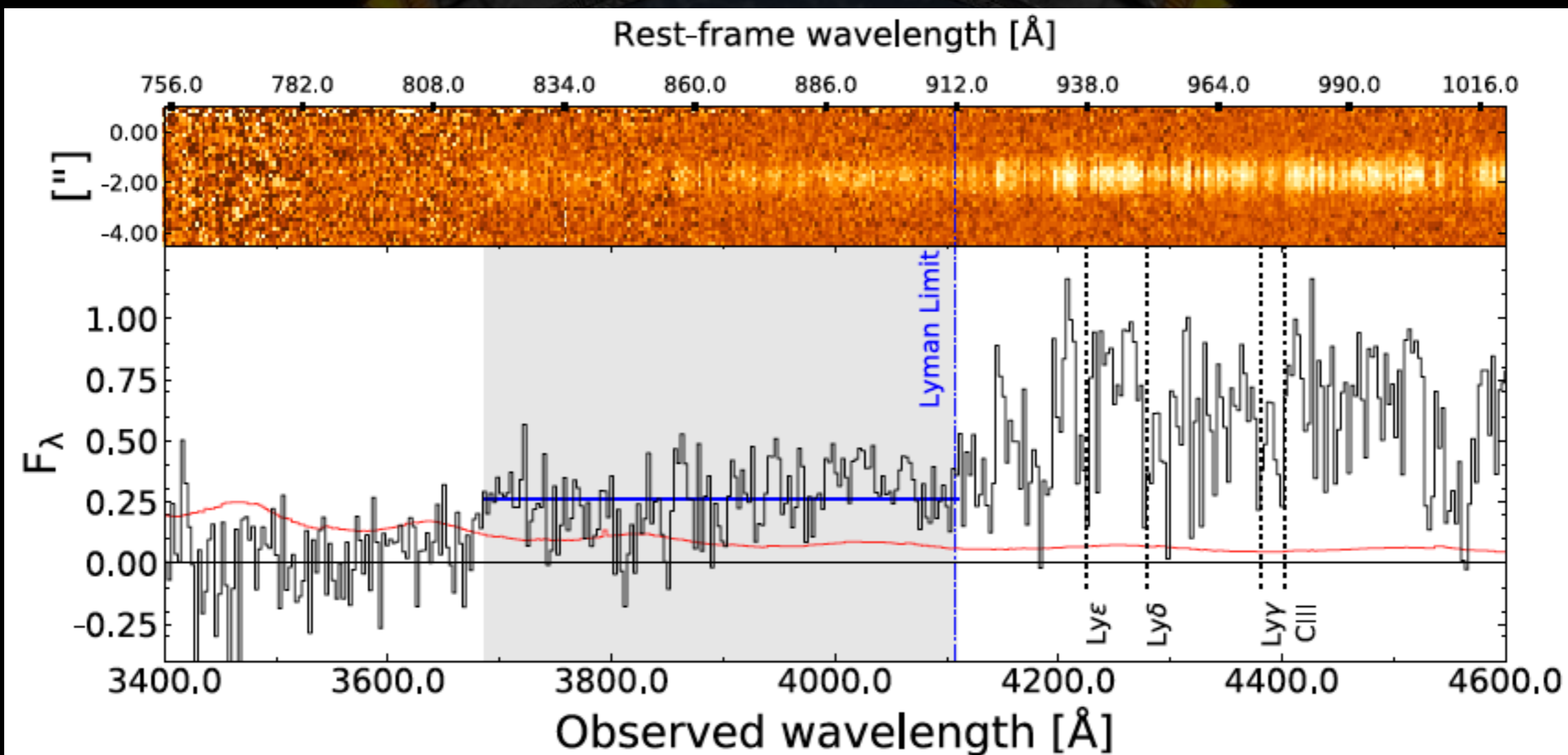
- Photometry
- Spectroscopy
- Polarimetry
- long & short GRBs

2 people on duty (24/7)
1 week every ~ 2months

Lyman continuum leakage

GRB191004B @ $z=3.5$

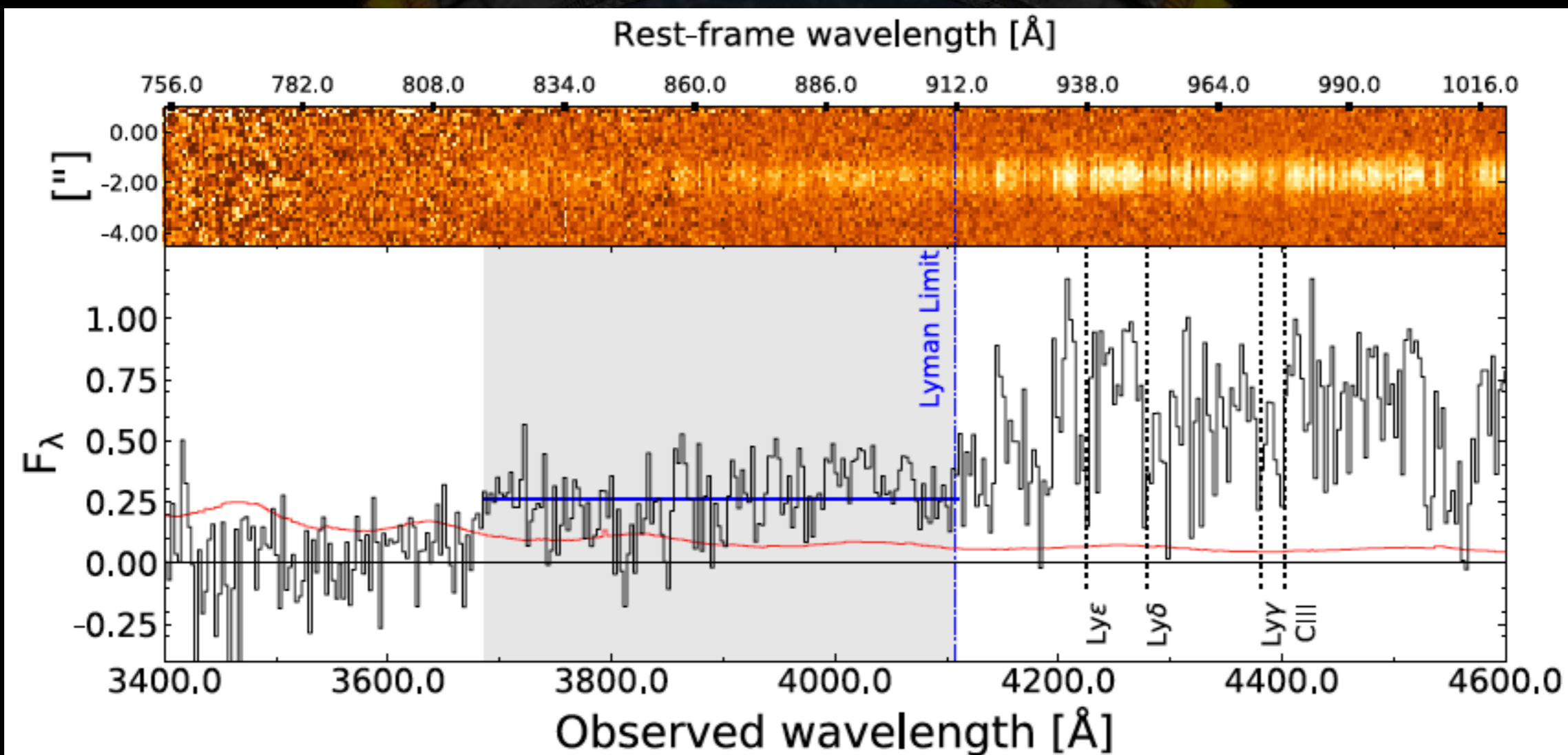
Vielfaure+2020



Lyman continuum leakage

GRB191004B @ $z=3.5$

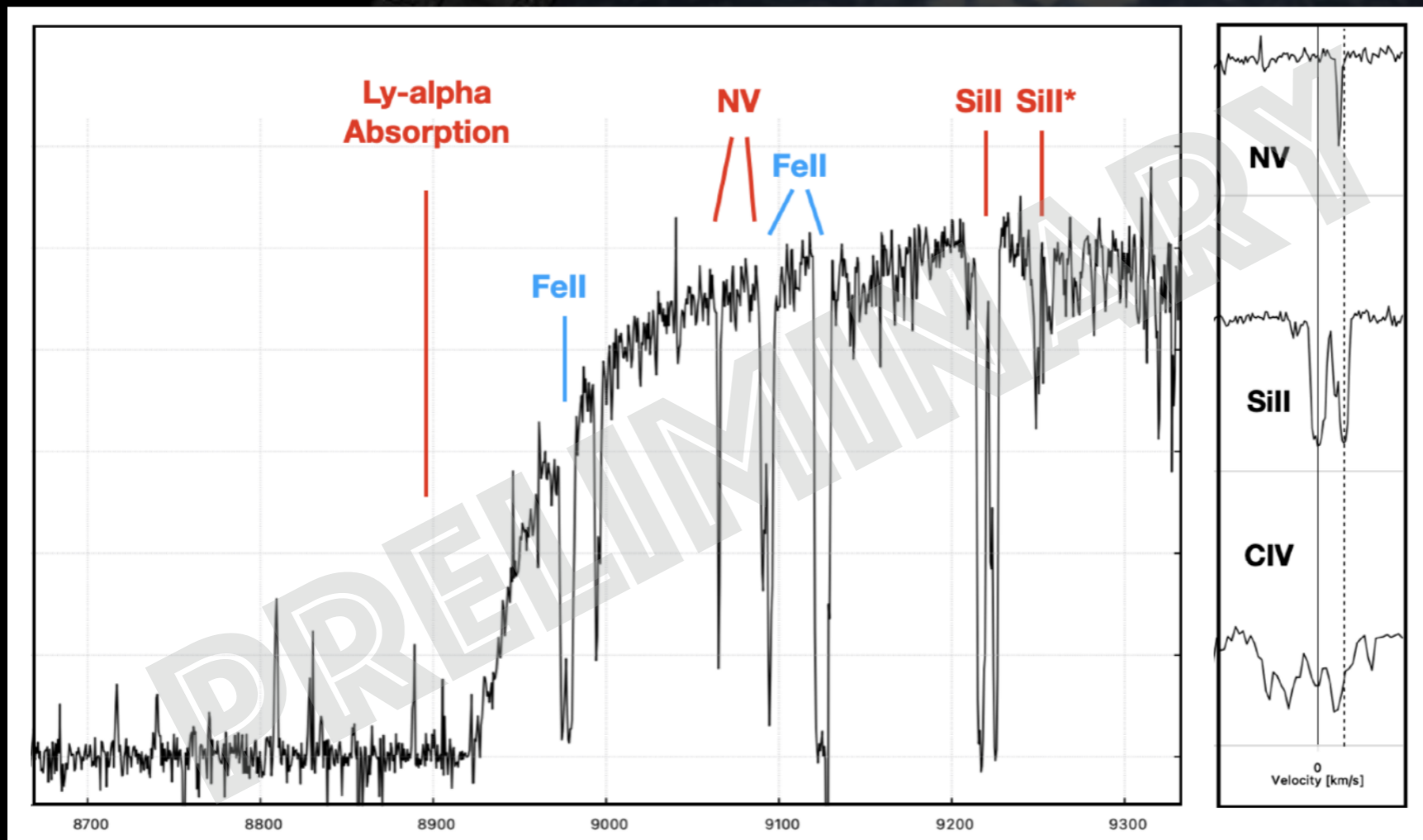
Vielfaure+2020



VHE GRB afterglow spectra
extremely rich of information

very high- z GRBs

X-shooter spectrum of $z=6.3$ GRB afterglow

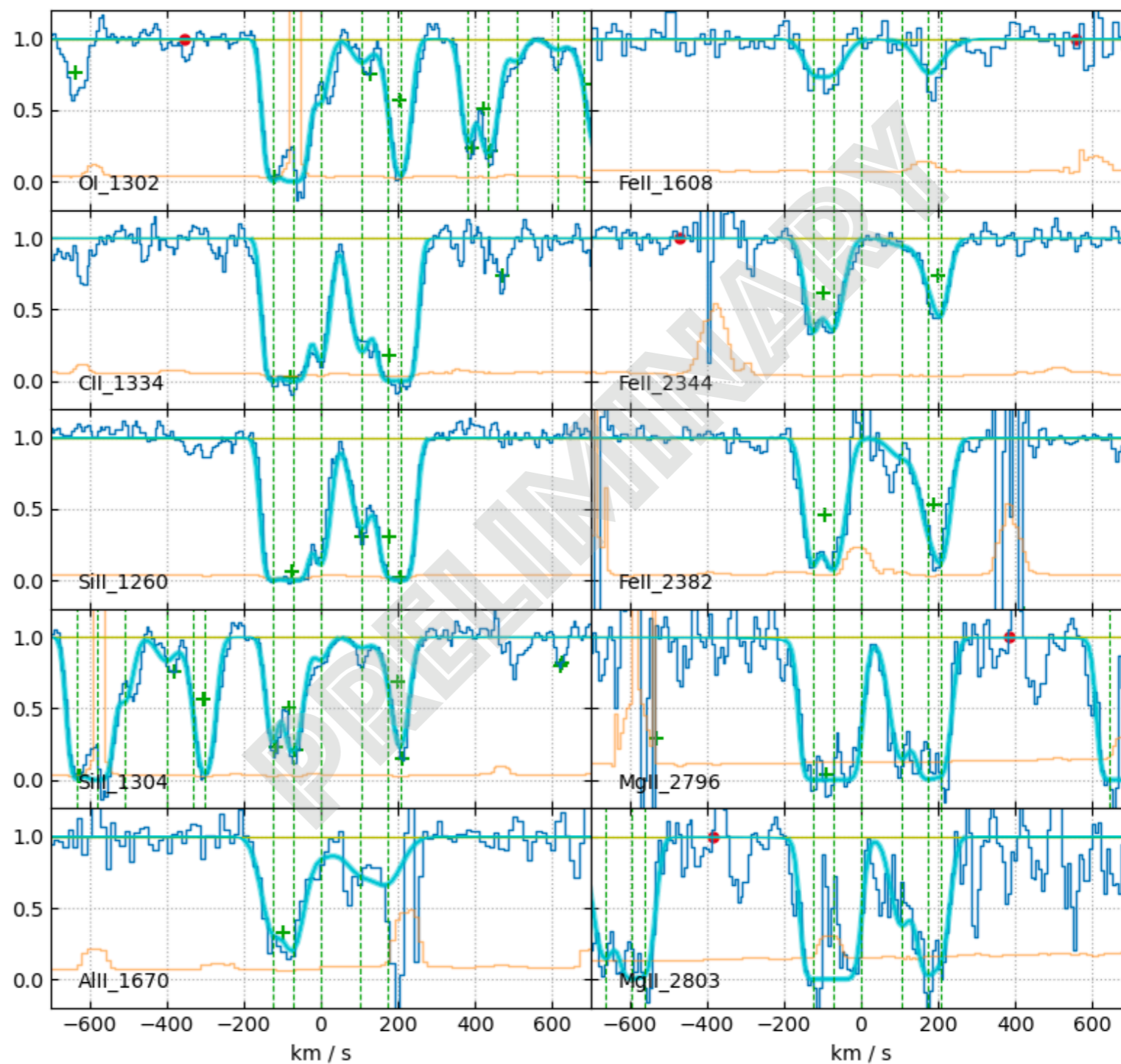


very high- z GRBs

X-shooter spectra



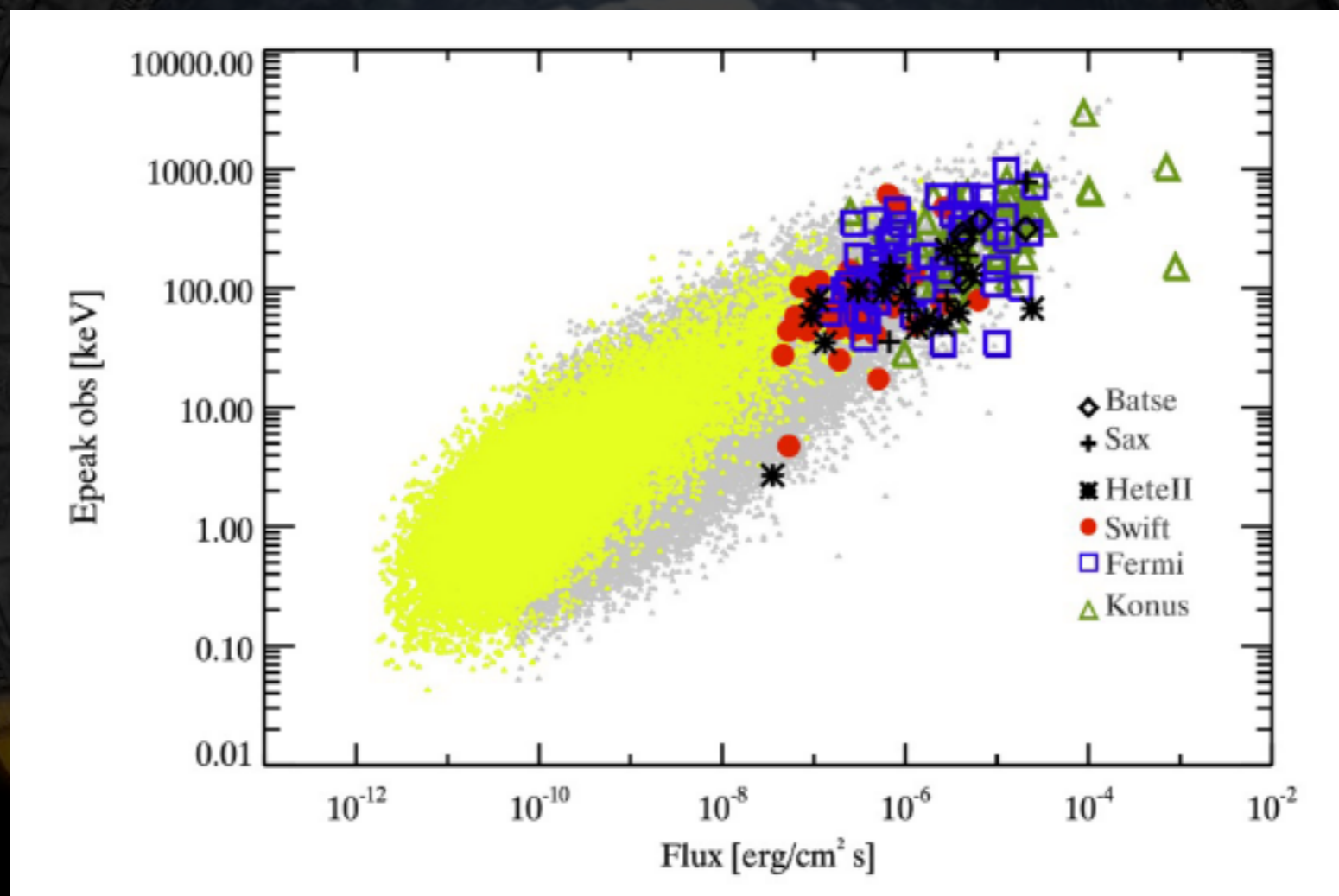
Saccardi+ in preparation



very high- z GRBs

We need to increase the number of high- z GRBs detected by the satellites

Ghirlanda+ 2015



$z > 5$ population

Some statistics

159 localized GRBs

35 redshift

16 VLT (the others: most of them too North, by GTC)

~4/5 no redshift :

VLT technical nights / visitors

No XRT (or delayed)

Too faint

Not observable

Bad weather

Too north

Some statistics

159 localized GRBs

35 redshift

16 VLT (the others: most of them too North, by GTC)

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No XRT (or delayed)

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Not observable

Bad weather

Too north

Open collaboration
Some SVOM members

If you are willing to help (and respect some basic rules)
you are welcome to join!

Data public under a certain period
(From months to 1yr)

GCN with redshift and basic info immediately released



Thank you!