

Improved near detectors and nuclear models for present and future neutrino oscillation experiments

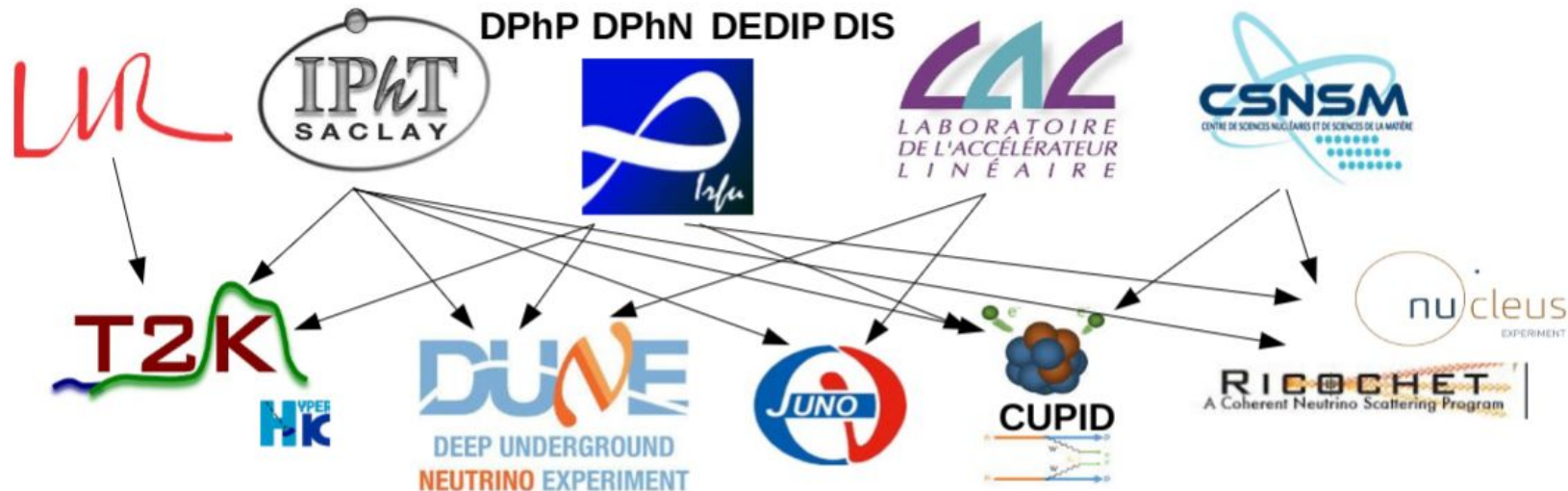
Jaafar Chakrani, for the BSM-Nu group
Laboratoire Leprince-Ringuet (LLR)
<jaafar.chakrani@polytechnique.edu>

Journée du LabEx P2IO
Laboratoire de Physique des Solides, Orsay
November 26th, 2021



- Introduction:
 - BSM-Nu project
 - Neutrino oscillations and long-baseline experiments
- T2K Near Detector Upgrade
- Improved nuclear models
- Summary and prospects

Introduction



BSM-Nu: federation of all the actors of neutrino physics inside the P2IO perimeter. Very new kind of group in the field: neutrino+nuclear physicist + theoreticians + engineers from different neutrino experiments working together and sharing expertise towards the precision era in nu physics.

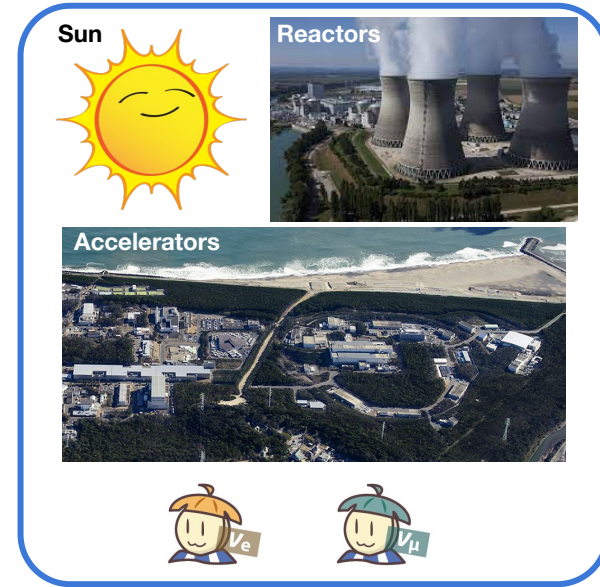
Goal: Neutrino nature (Majorana?) and neutrino mixing (MH? δ_{CP} ?) as a way to access BSM physics

- Neutrino flavor states: $\nu_e, \nu_\mu, \nu_\tau \rightarrow$ production and detection
- Neutrino mass states: $\nu_1, \nu_2, \nu_3 \rightarrow$ propagation

Neutrino flavors and masses mixing

6

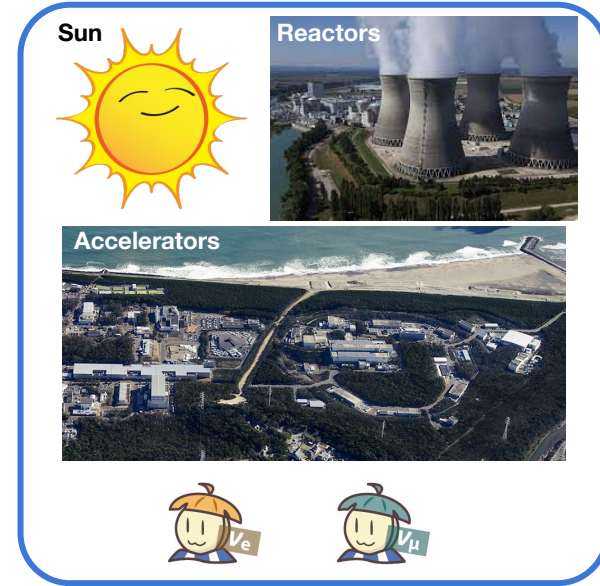
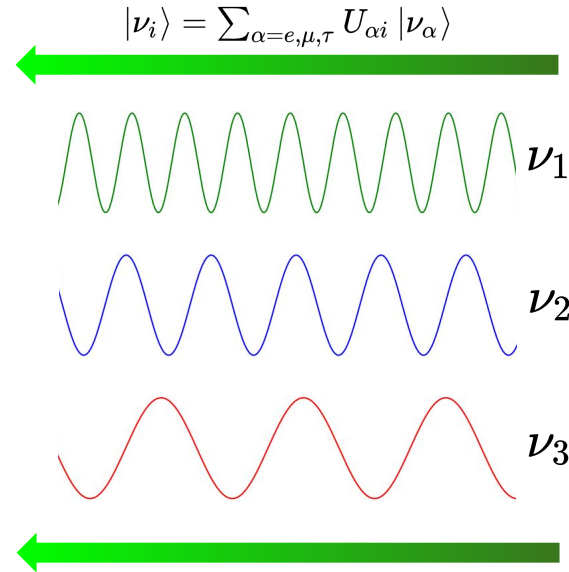
- Neutrino flavor states: $\nu_e, \nu_\mu, \nu_\tau \rightarrow$ **production** and detection
- Neutrino mass states: $\nu_1, \nu_2, \nu_3 \rightarrow$ propagation



Neutrino flavors and masses mixing

7

- Neutrino flavor states: $\nu_e, \nu_\mu, \nu_\tau \rightarrow$ **production** and detection
- Neutrino mass states: $\nu_1, \nu_2, \nu_3 \rightarrow$ **propagation**

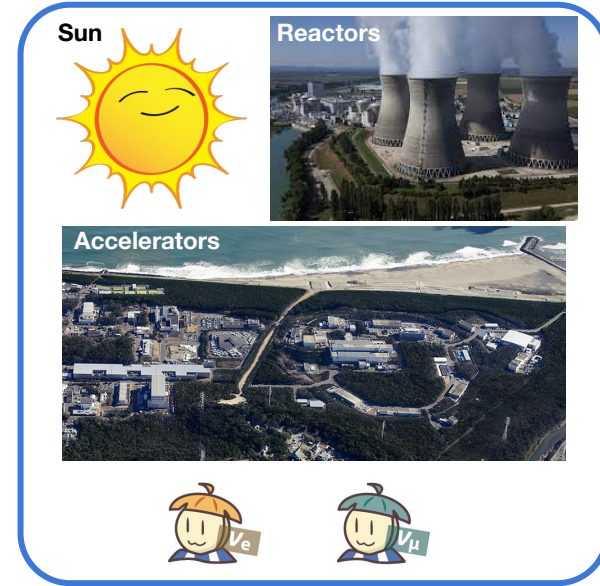
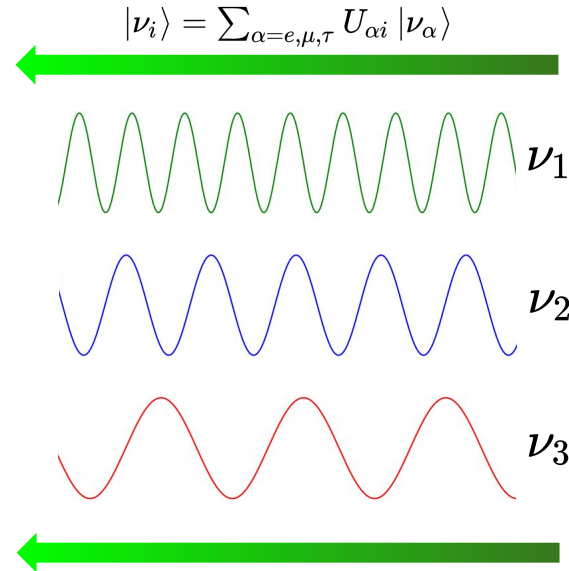


Neutrino flavors and masses mixing

8

- Neutrino flavor states: $\nu_e, \nu_\mu, \nu_\tau \rightarrow$ **production** and detection
- Neutrino mass states: $\nu_1, \nu_2, \nu_3 \rightarrow$ **propagation**

} **flavor oscillations**

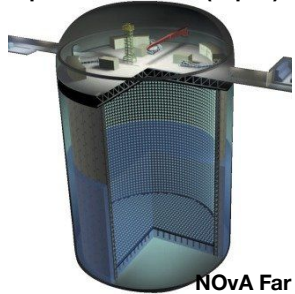


Neutrino flavors and masses mixing

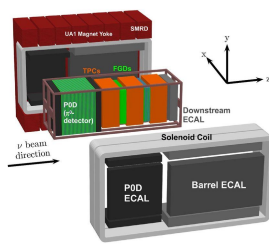
9

- Neutrino flavor states: $\nu_e, \nu_\mu, \nu_\tau \rightarrow$ **production** and **detection**
 - Neutrino mass states: $\nu_1, \nu_2, \nu_3 \rightarrow$ **propagation**
- } flavor oscillations

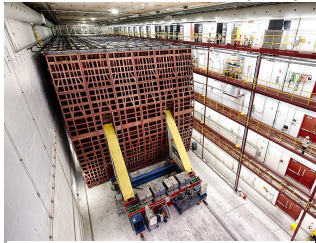
Super-Kamiokande (Japan)



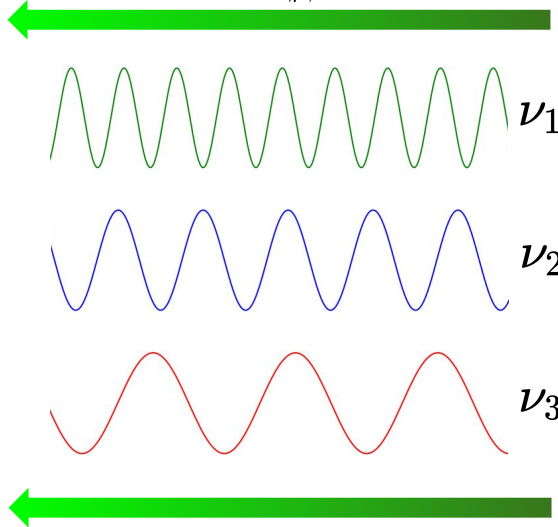
T2K Near Detector (Japan)



NOvA Far Detector (USA)



$$|\nu_i\rangle = \sum_{\alpha=e,\mu,\tau} U_{\alpha i} |\nu_\alpha\rangle$$



Sun



Reactors



Accelerators



- Mass and flavor states mixing: $|\nu_i\rangle = \sum_{\alpha=e,\mu,\tau} U_{\alpha i} |\nu_\alpha\rangle$

$$U = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta_{CP}} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta_{CP}} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \quad \begin{aligned} c_{ij} &= \cos(\theta_{ij}) \\ s_{ij} &= \sin(\theta_{ij}) \end{aligned}$$

- Long-baseline experiments are sensitive to:
 - Atmospheric parameters $(\theta_{23}, \Delta m_{32}^2)$ through $\nu_\mu/\bar{\nu}_\mu$ disappearance

$$P(\bar{\nu}_\mu \rightarrow \bar{\nu}_\mu) \approx 1 - \sin^2 2\theta_{23} \sin^2 \left(\frac{\Delta m_{32}^2 L}{4E} \right)$$

- $(\delta_{CP}, \theta_{23})$ through $\nu_e/\bar{\nu}_e$ appearance

$$P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e) \approx \sin^2 2\theta_{13} \sin^2 \theta_{23} \sin^2 \left(\frac{\Delta m_{32}^2 L}{4E} \right) (\mp) O(\delta_{CP})$$

- Mass and flavor states mixing: $|\nu_i\rangle = \sum_{\alpha=e,\mu,\tau} U_{\alpha i} |\nu_\alpha\rangle$

$$U = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta_{CP}} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta_{CP}} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \quad \begin{matrix} c_{ij} = \cos(\theta_{ij}) \\ s_{ij} = \sin(\theta_{ij}) \end{matrix}$$

- Long-baseline experiments are sensitive to:

- Atmospheric parameters $(\theta_{23}, \Delta m_{32}^2)$ through $\nu_\mu/\bar{\nu}_\mu$ disappearance

$$P(\bar{\nu}_\mu \rightarrow \bar{\nu}_\mu) \approx 1 - \sin^2 2\theta_{23} \sin^2 \left(\frac{\Delta m_{32}^2 L}{4E} \right)$$

- $(\delta_{CP}, \theta_{23})$ through $\nu_e/\bar{\nu}_e$ appearance

$$P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e) \approx \sin^2 2\theta_{13} \sin^2 \theta_{23} \sin^2 \left(\frac{\Delta m_{32}^2 L}{4E} \right) (\mp) O(\delta_{CP})$$

If $P(\nu_\mu \rightarrow \nu_e) \neq P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)$
then matter and anti-matter
could behave differently in
the lepton sector
→ CP violation!

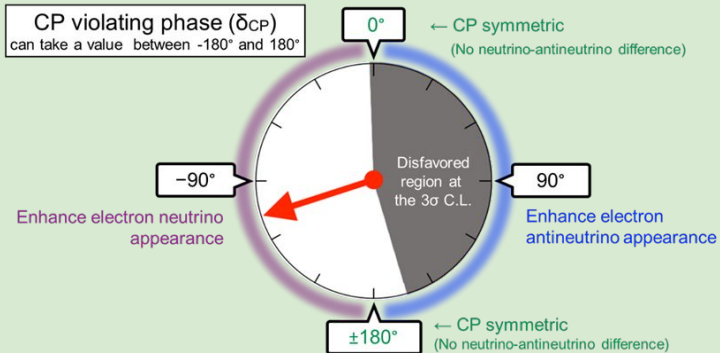
This could shed light on
the matter/anti-matter
asymmetry in the Universe

First hints by T2K

April 2020



- CP conservation is excluded at a 2σ level
- Preference for a \sim maximal CP violation



$$U_{\alpha i} |\nu_{\alpha}\rangle$$

$$\begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \quad \begin{aligned} c_{ij} &= \cos(\theta_{ij}) \\ s_{ij} &= \sin(\theta_{ij}) \end{aligned}$$

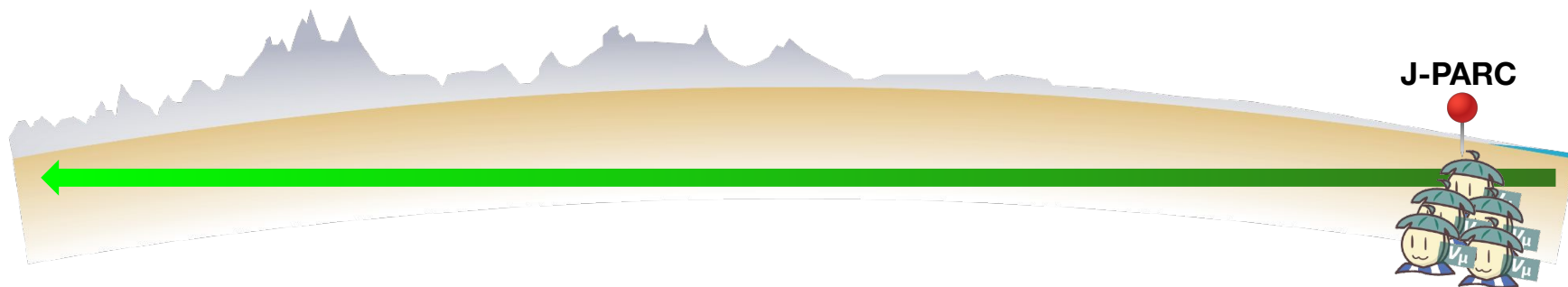
If $P(\nu_{\mu} \rightarrow \nu_e) \neq P(\bar{\nu}_{\mu} \rightarrow \bar{\nu}_e)$
then matter and anti-matter
could behave differently in
the lepton sector
→ CP violation!

This could shed light on
the matter/anti-matter
asymmetry in the Universe

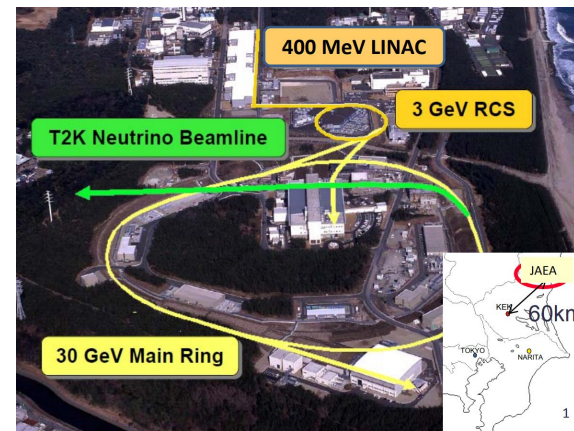
ance

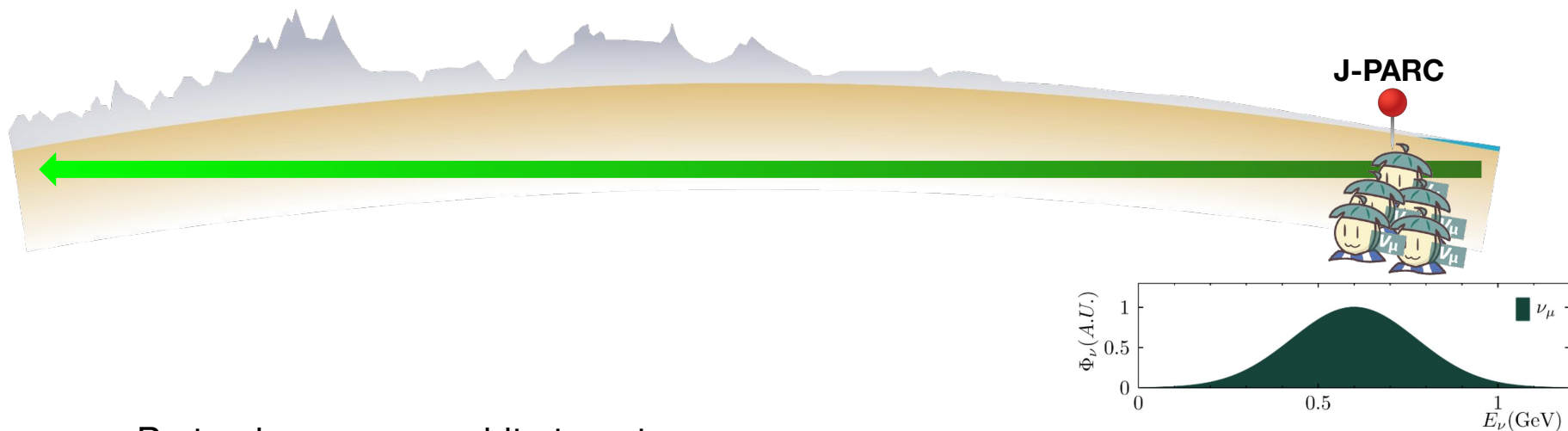
$$\frac{2}{32} \frac{L}{E}$$

$$\left(\mp \right) O(\delta_{CP})$$

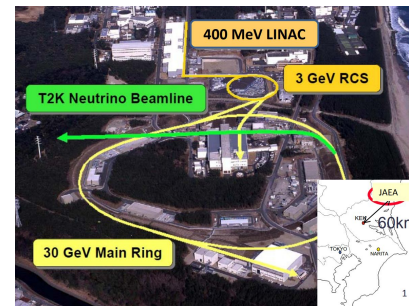


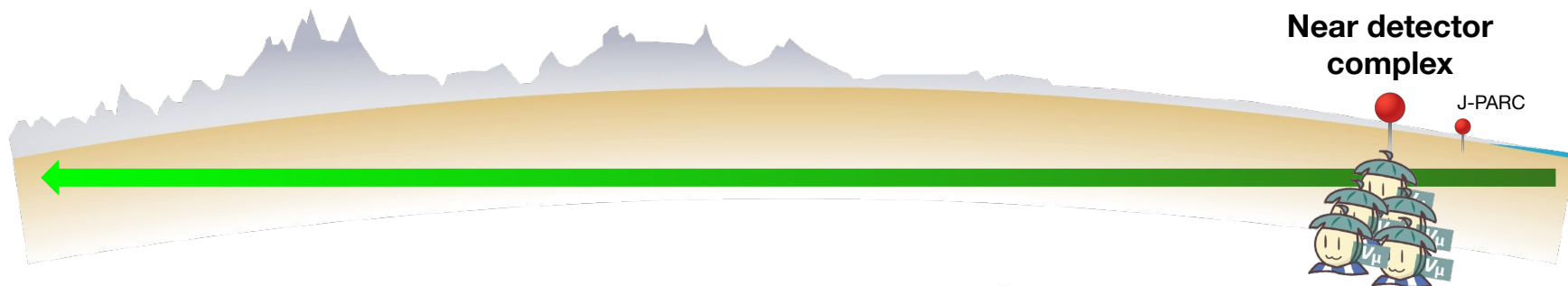
- Proton beam on graphite target
- Produced hadrons decay into muon (anti-)neutrinos





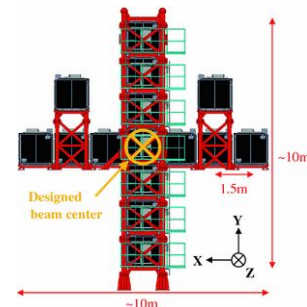
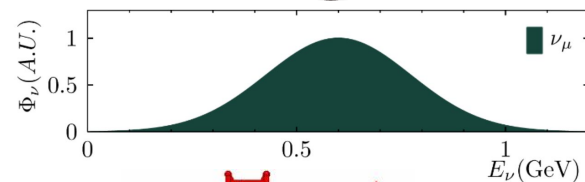
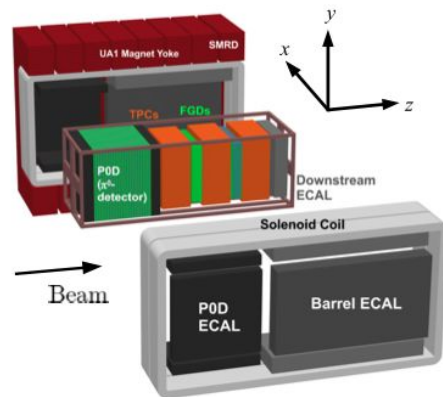
- Proton beam on graphite target
- Produced hadrons decay into muon (anti-)neutrinos

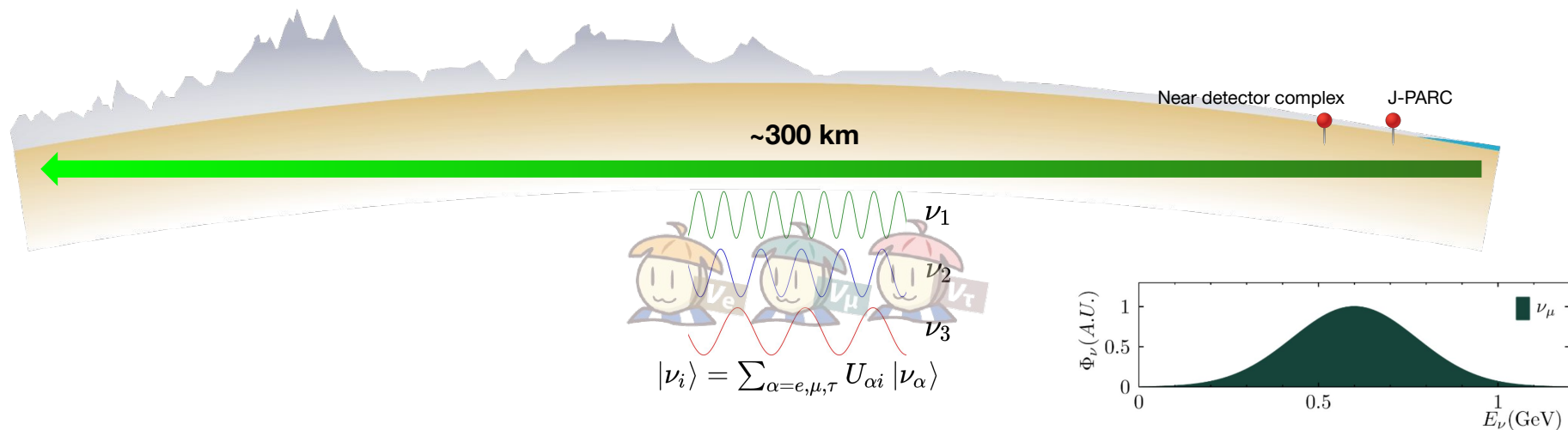


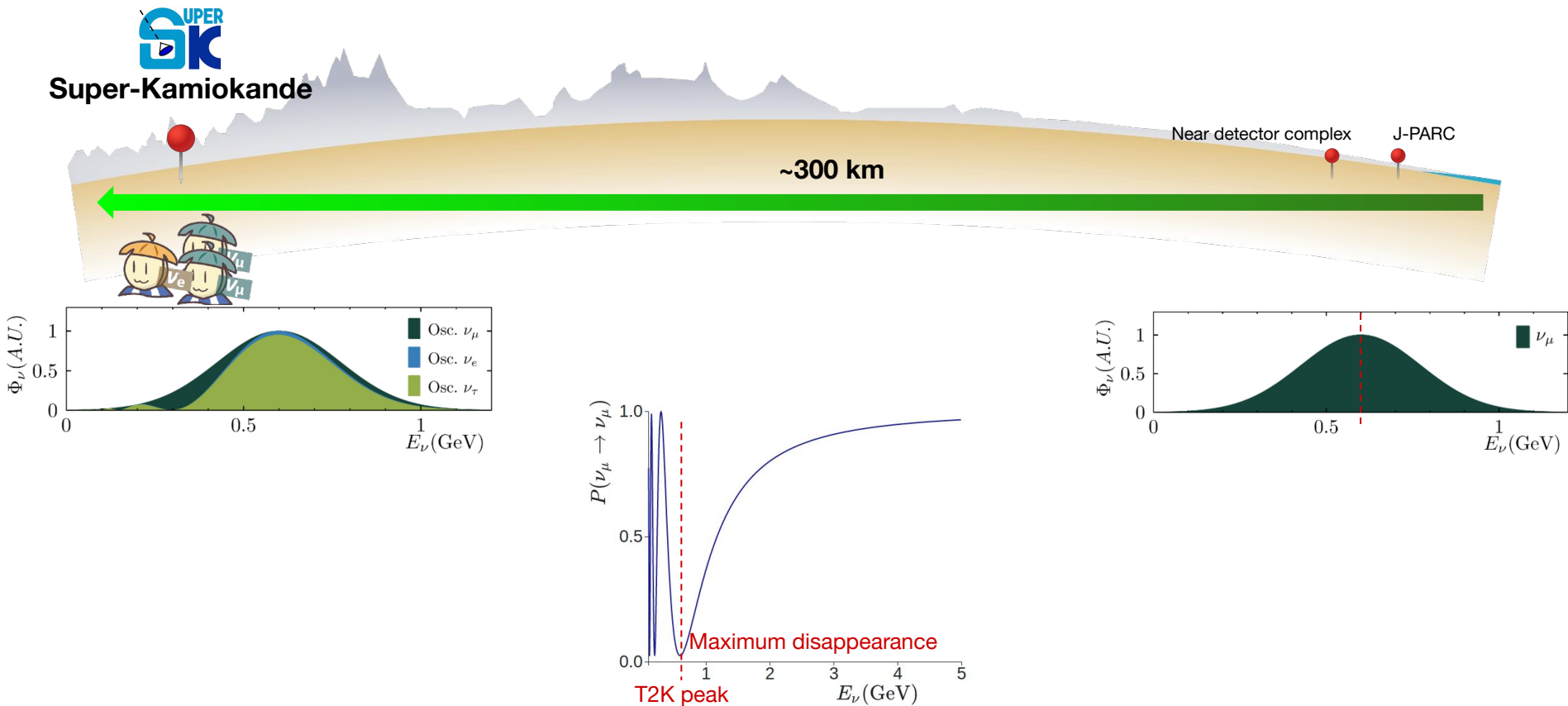


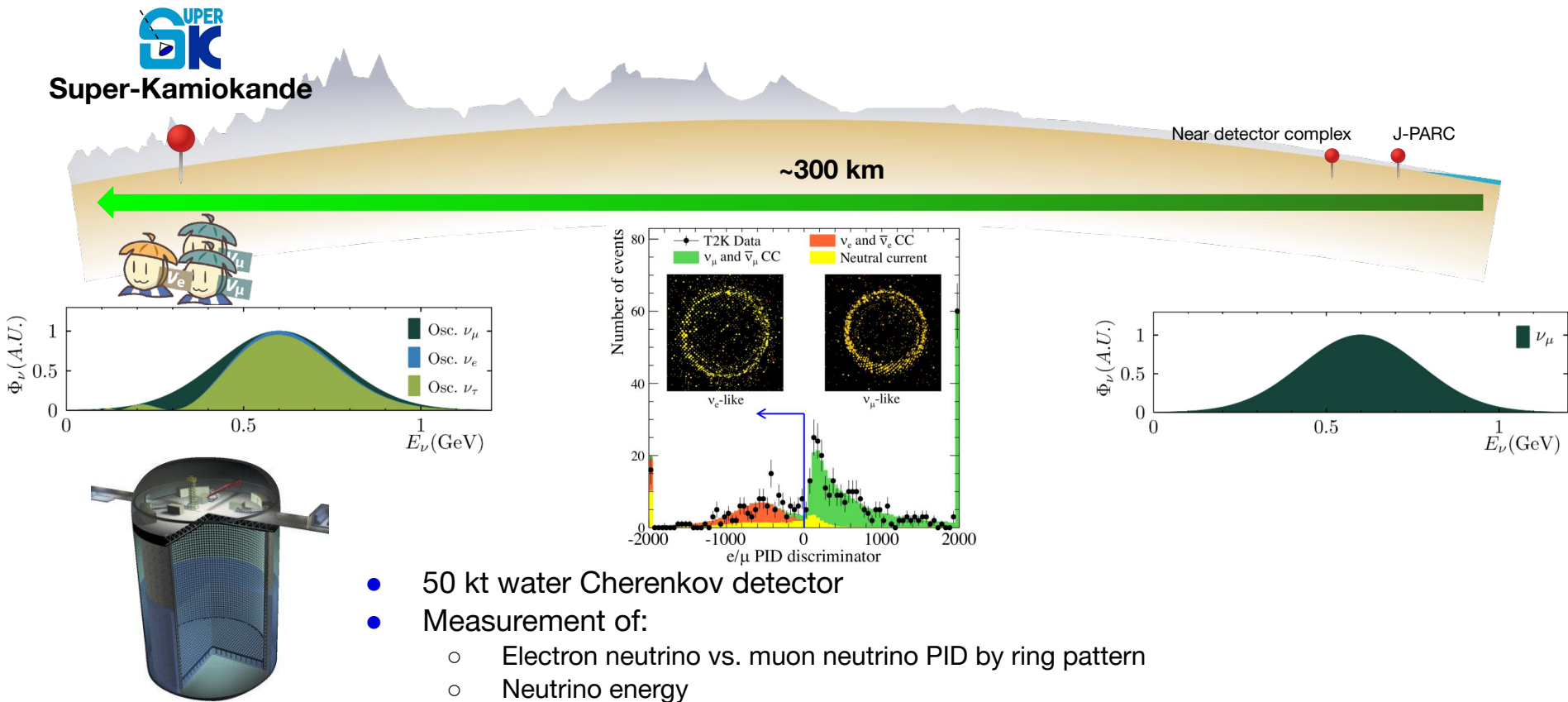
- Measure unoscillated neutrino flux:
 - Electron neutrino and wrong-sign contaminations
 - Neutrino-nucleus interactions

→ Reduce **systematic uncertainties**





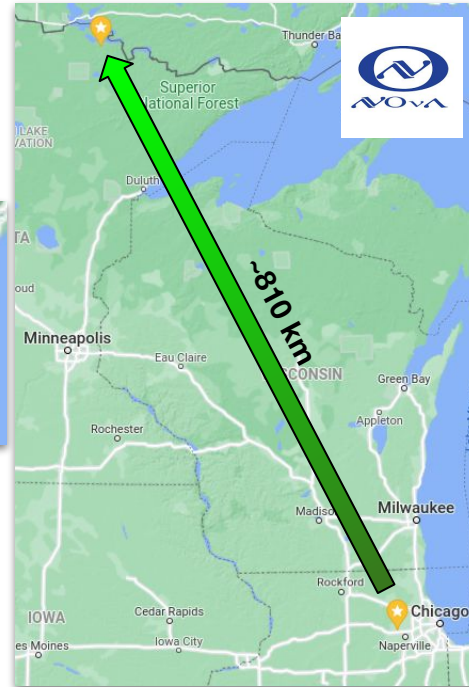




Long Baseline Experiments

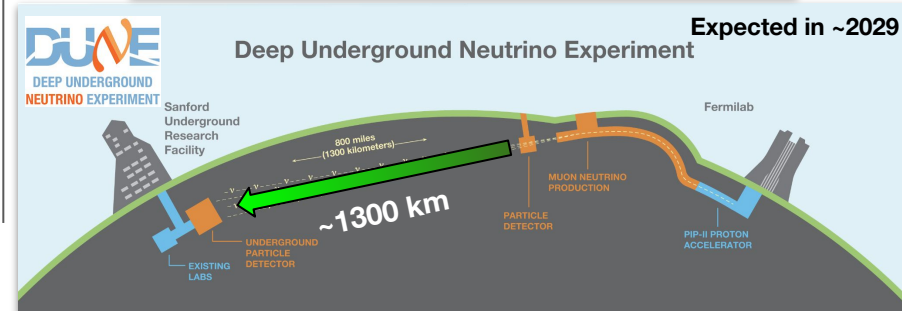
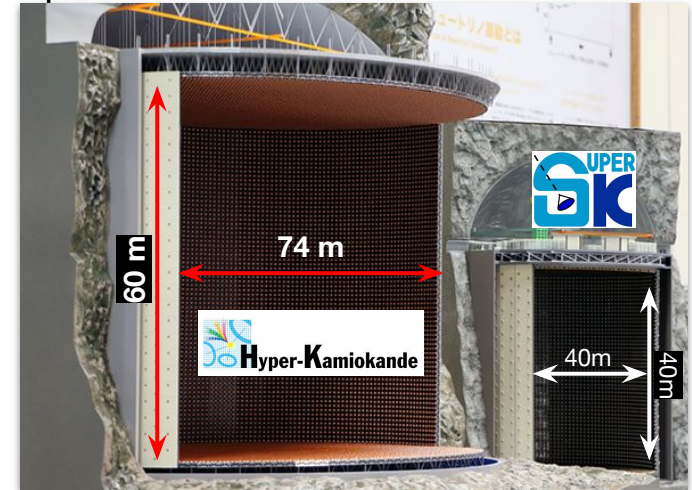
19

- Currently:



- Next generation:

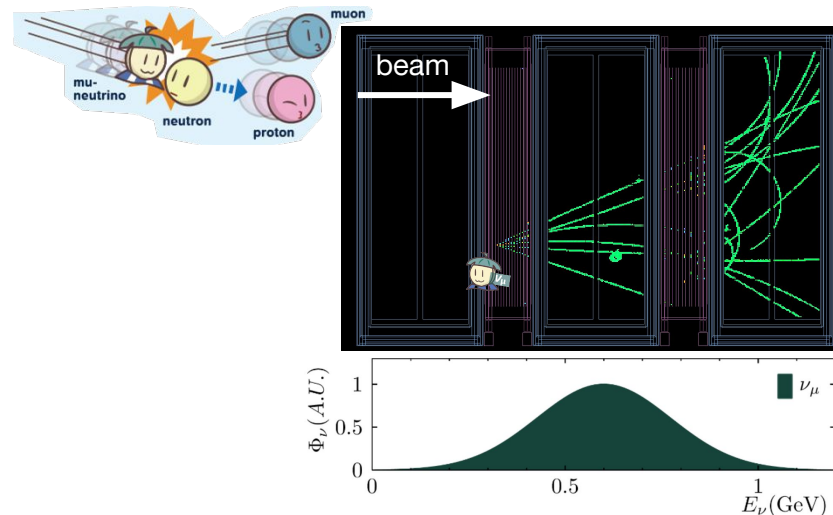
Expected in 2027



- To measure the oscillation parameters, the **neutrino energy** needs to be determined precisely

$$P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e) \approx \sin^2 2\theta_{13} \sin^2 \theta_{23} \sin^2 \left(\frac{\Delta m_{32}^2 L}{4E} \right) (\mp) O(\delta_{CP})$$

- An accurate reconstruction of the **neutrino energy** from the **outgoing particles** requires a precise **neutrino-nucleus interaction model**



- To measure the oscillation parameters, the **neutrino energy** needs to be determined precisely

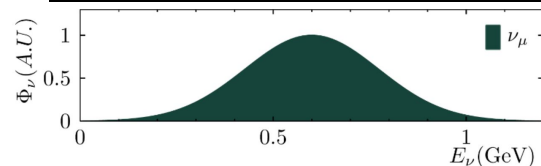
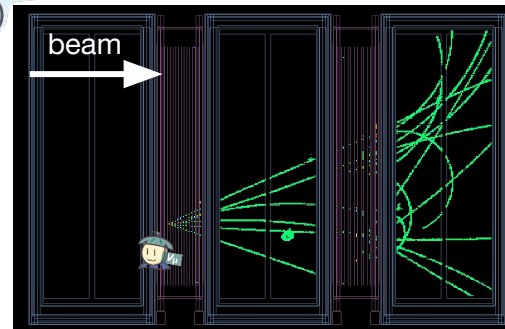
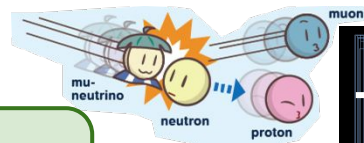
$$P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e) \approx \sin^2 2\theta_{13} \sin^2 \theta_{23} \sin^2 \left(\frac{\Delta m_{32}^2 L}{4E} \right) (\mp) O(\delta_{CP})$$

- An accurate reconstruction of the **neutrino energy** from the **outgoing particles** requires a precise **neutrino-nucleus interaction model**

Two ways to improve the neutrino energy reconstruction:

- Improve the detectors
- Improve the nuclear model

→ **Focus of BSM-Nu WP2**

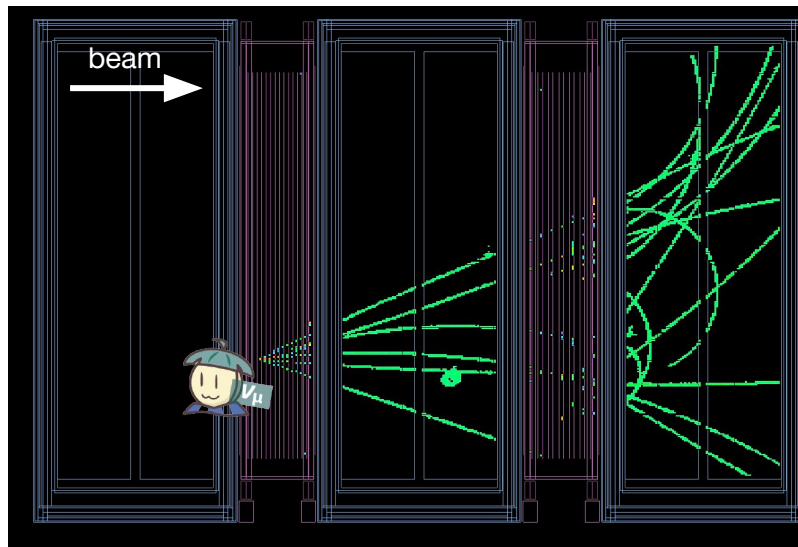
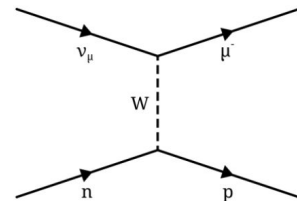
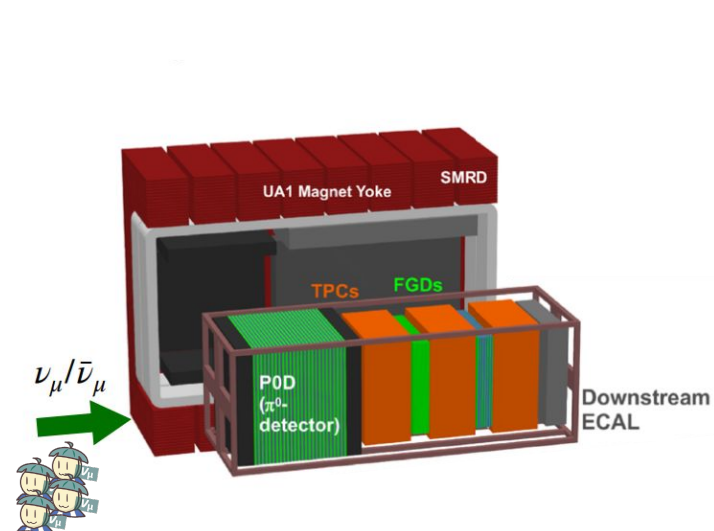




T2K Near Detector Upgrade (T2K-II)

Off-axis Near Detector at 280m (ND280)


23

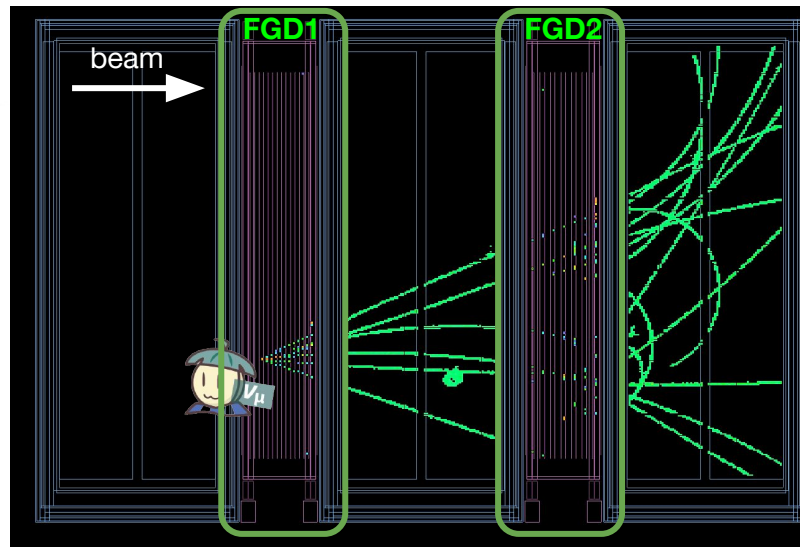
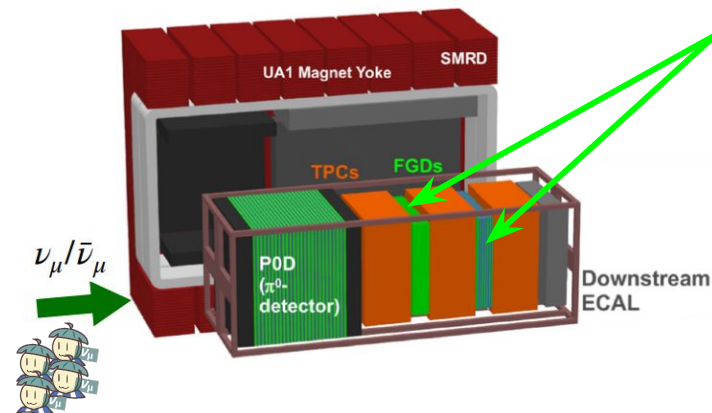
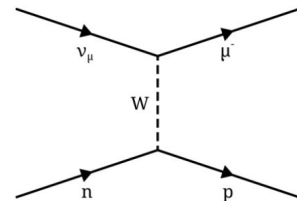


Off-axis Near Detector at 280m (ND280)

24

Fine Grained Detectors (FGDs):

- Plastic scintillator tracker
- Target for neutrinos 
- FGD2 has water target layers

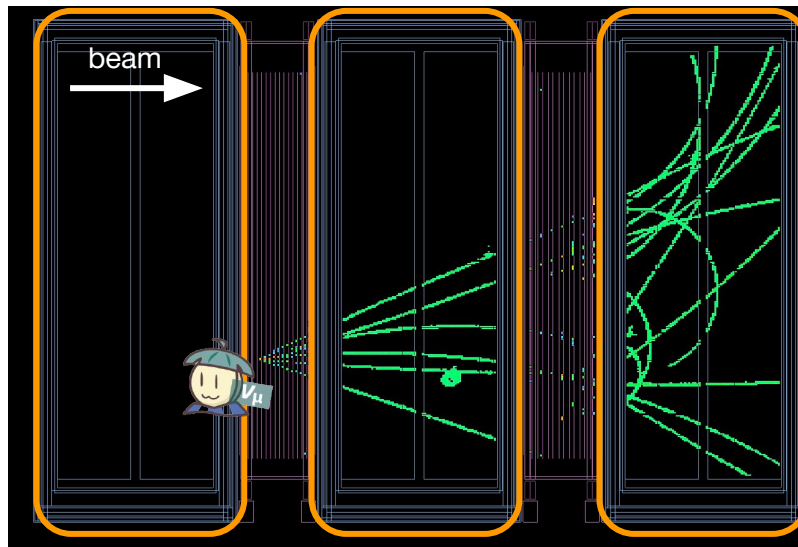
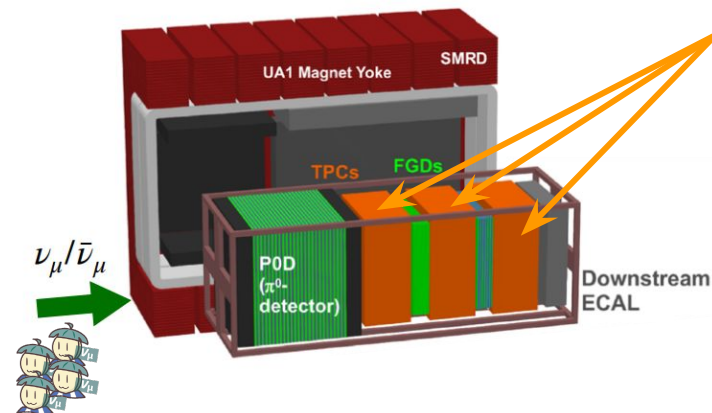
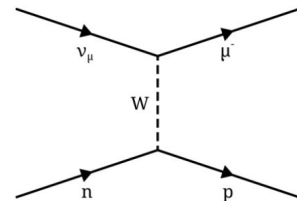


Off-axis Near Detector at 280m (ND280)

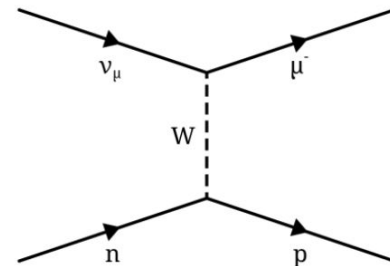
25

Time Projection Chambers (TPCs)

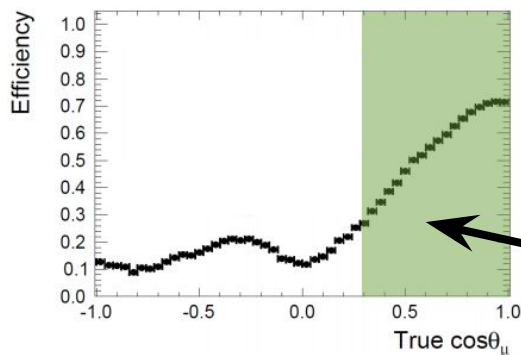
- Tracking detectors
- Charged particle momentum
- Particle ID



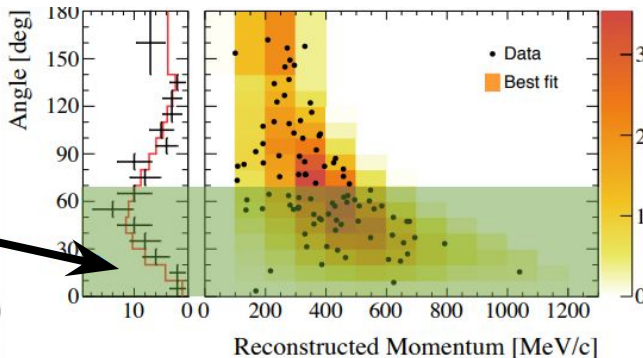
- Non-isotropic efficiency (unlike Super-Kamiokande)
- High momentum proton threshold (~ 450 MeV/c)
- For the oscillation analysis, neutrino interactions are characterized in muon kinematics only



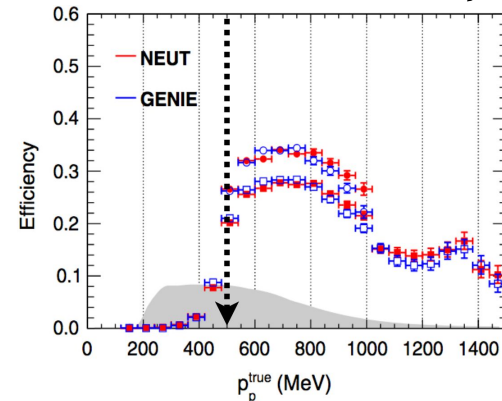
Muon detection efficiency at ND280



ν_e candidates at Super-K



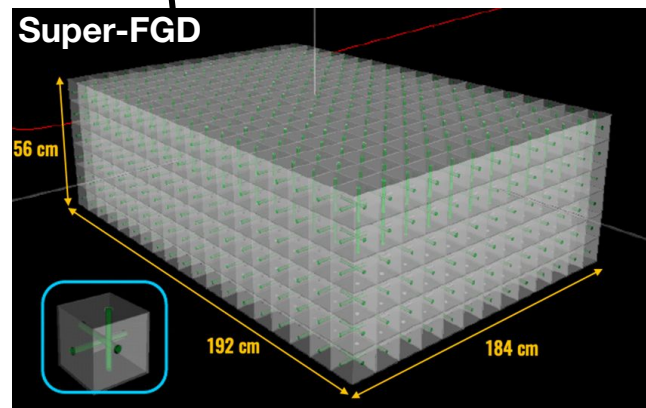
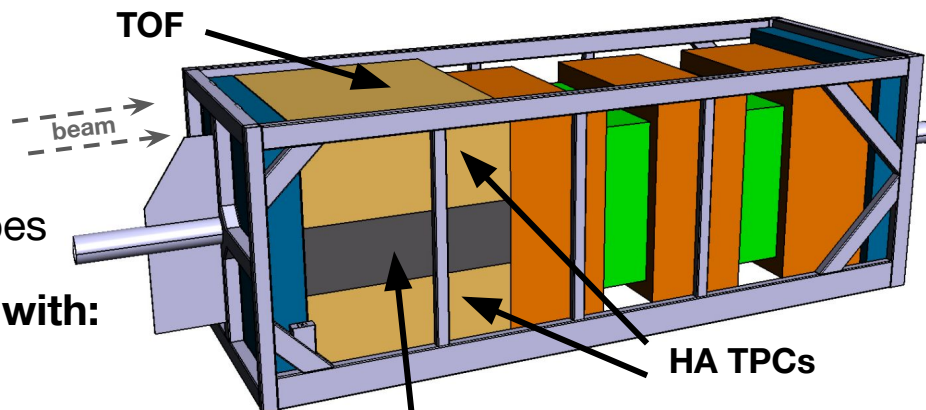
Proton detection efficiency



- New high-angle TPCs
- New Time Of Flight detector
- Super-FGD: 2.10^6 1 cm^3 scintillator cubes

The goal is to reduce the ND systematics with:

- Fully active target
- 4π acceptance for charged particles
- Lower proton momentum threshold ($\sim 300 \text{ MeV}/c$)
- Neutron detection



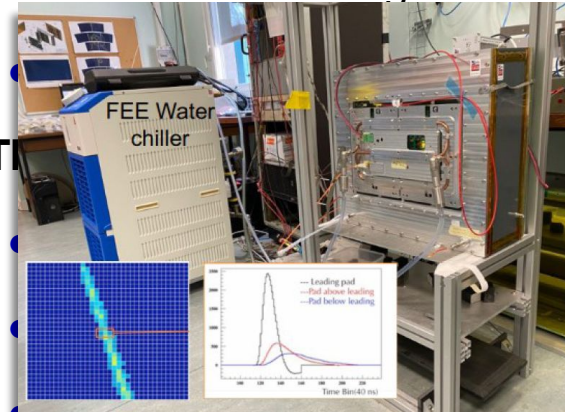


New high-angle TPCs

Beam test at CERN in 2018
 Prototype tested at DESY in 2019 & 2021
 Ongoing tests at CERN

[arXiv:2106.12634](https://arxiv.org/abs/2106.12634)

- New Time Of Flight detector

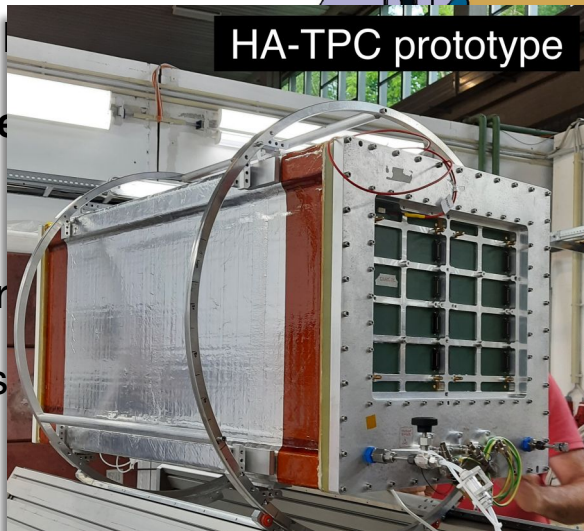


Scintill

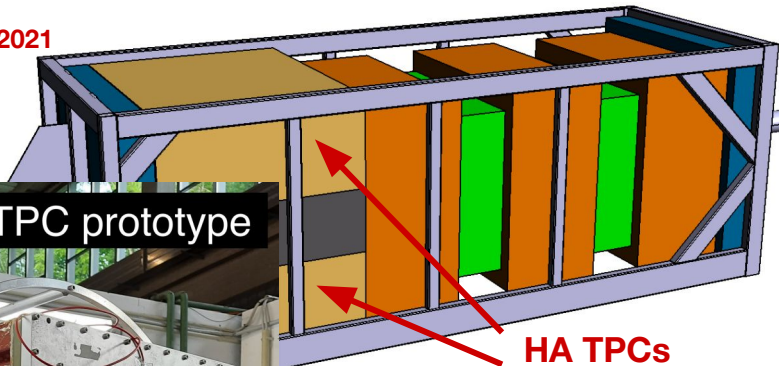
system

d par

thres

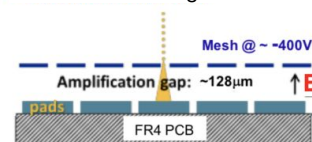


HA-TPC prototype

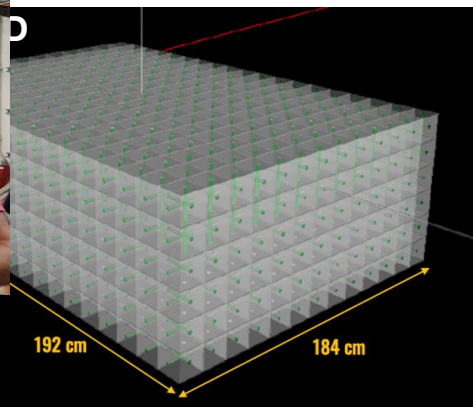
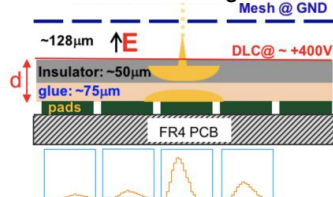


HA TPCs

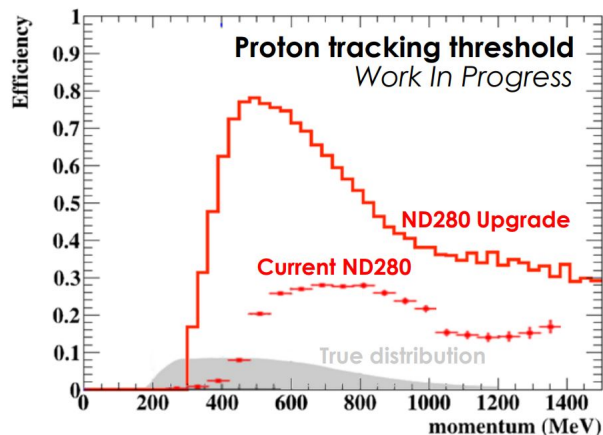
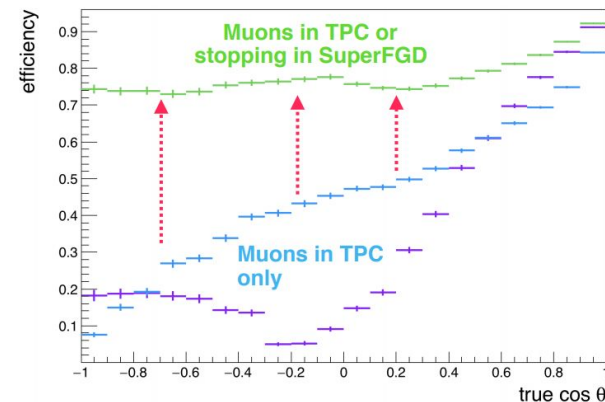
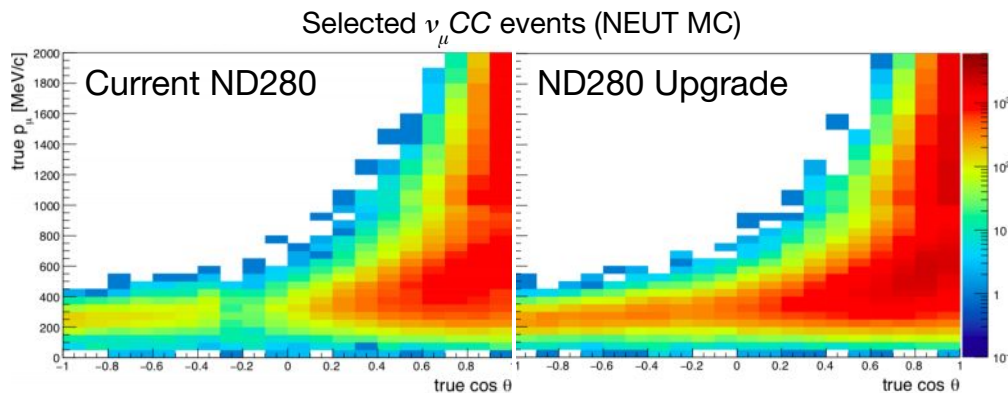
Standard Micromegas



Resistive Micromegas



- Improved reconstruction at high and backward angles \rightarrow better constraints on the neutrino interaction model
- Increased target mass (x2 current ND280) \rightarrow more statistics
- Better reconstruction of outgoing nucleons \rightarrow access to new observables
- Neutrino interaction measurements beyond p_μ , $\cos \theta_\mu$ (exclusive and multidimensional analyses)



- The neutrino energy can be estimated using the lepton kinematics only under the assumption of a quasi-elastic interaction on a static initial state nucleon

$$E_{\nu}^{QE} = \frac{m_p^2 - (m_n - E_b)^2 - m_{\mu}^2 + 2(m_n - E_b)E_{\mu}}{2(m_n - E_b - E_{\mu} + p_{\mu} \cos \theta_{\mu})}$$

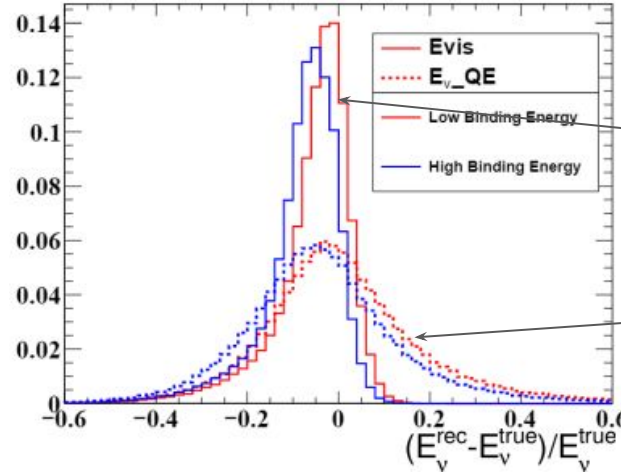
- When reconstructing the proton as well, the *visible energy* can be a better estimation of the neutrino energy

$$E_{\text{vis}} = E_{\mu} + T_p$$



[arXiv:2108.11779](https://arxiv.org/abs/2108.11779)

p.d.f.

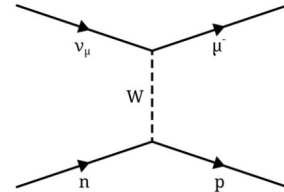


Using proton & lepton information

$$E_{\text{vis}} = E_{\mu} + T_p$$

Using lepton information only

$$E_{\nu}^{QE} = \frac{m_p^2 - (m_n - E_b)^2 - m_{\mu}^2 + 2(m_n - E_b)E_{\mu}}{2(m_n - E_b - E_{\mu} + p_{\mu} \cos \theta_{\mu})}$$

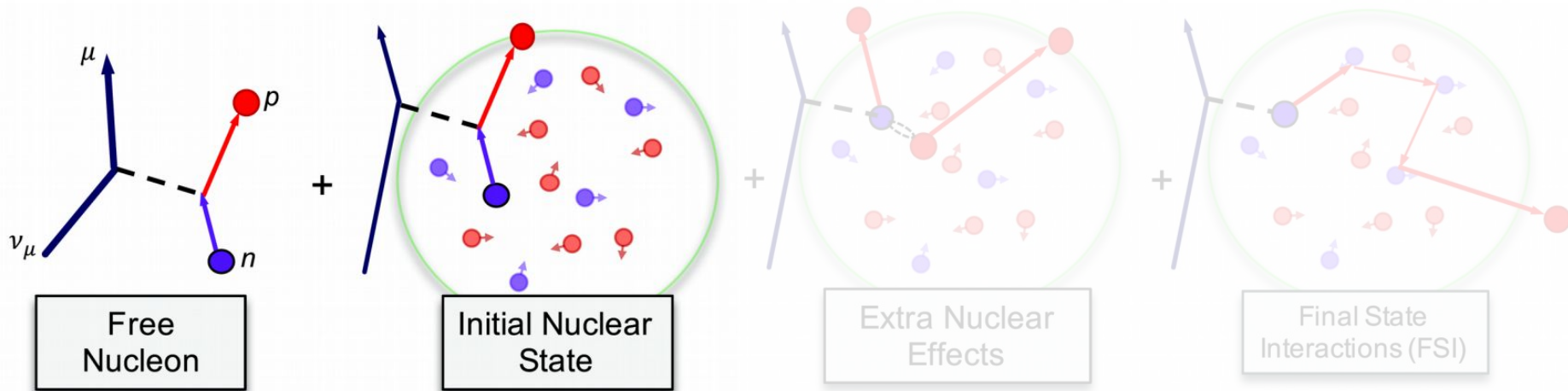


Improved nuclear models



$$E_\nu^{QE} = \frac{m_p^2 - (m_n - E_b)^2 - m_\mu^2 + 2(m_n - E_b)E_\mu}{2(m_n - E_b - E_\mu + p_\mu \cos \theta_\mu)}$$

is an accurate estimation of the neutrino energy using muon kinematics only



Nucleons are bound within nuclei (Carbon & Oxygen in T2K).

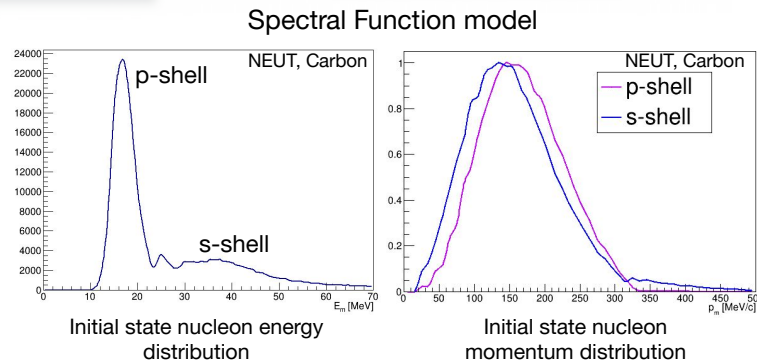
Need to account for:

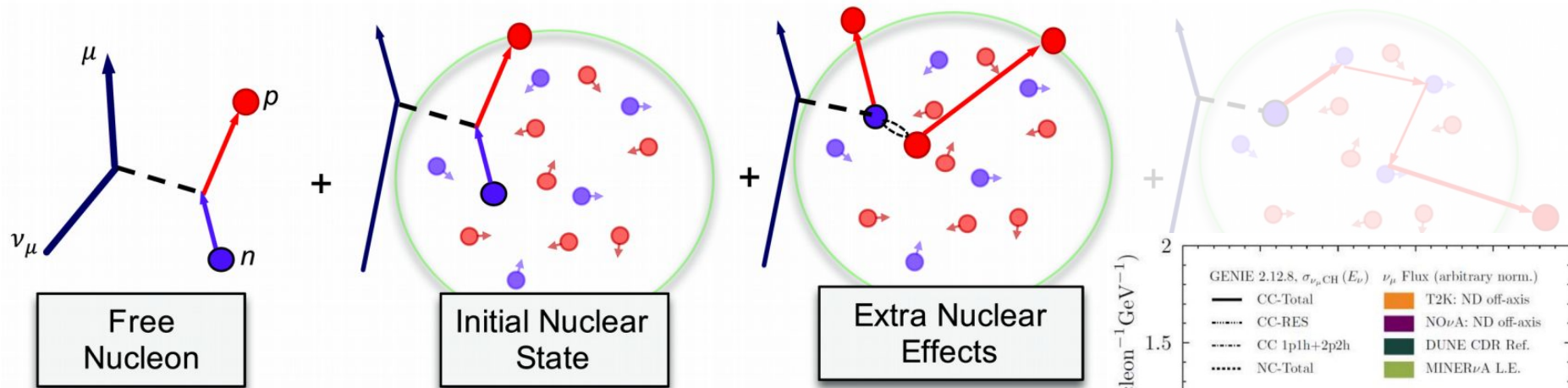
- Removal energy
- Fermi motion

T2K now uses the sophisticated **Spectral Function model**

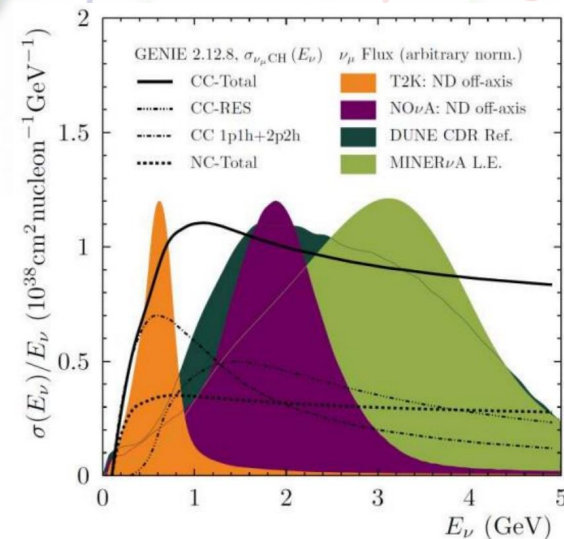
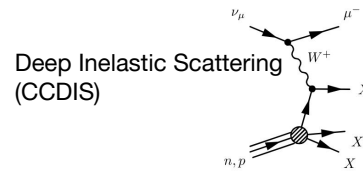
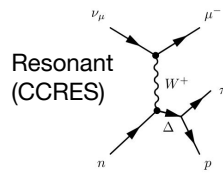
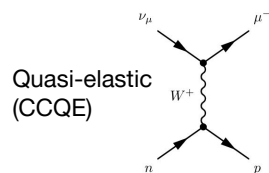


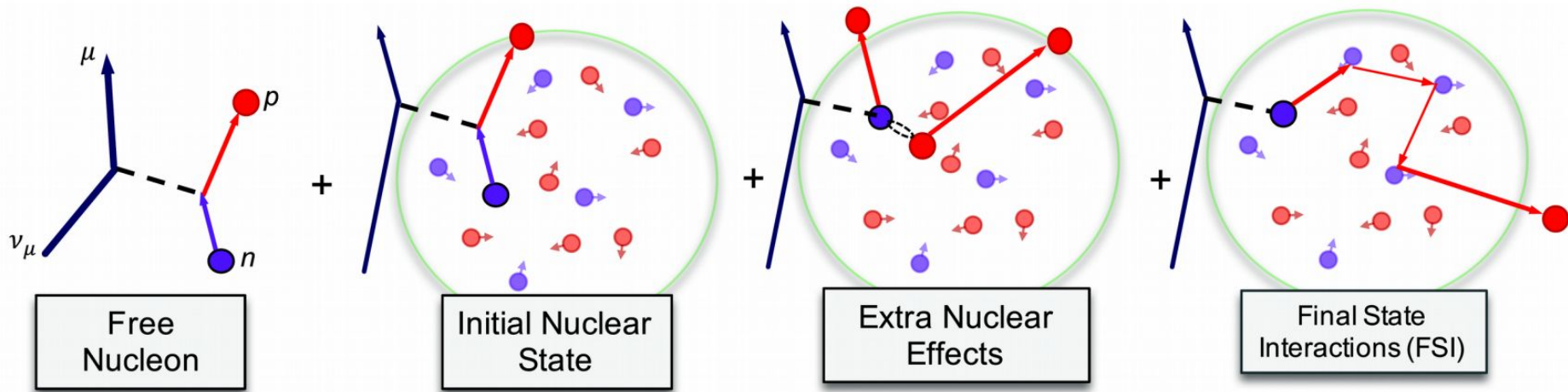
→ Development of systematic uncertainties in this new framework for the oscillation analysis





- Other nuclear effects can bias the reconstruction of neutrino energy like short-range correlations, 2p2h interactions...





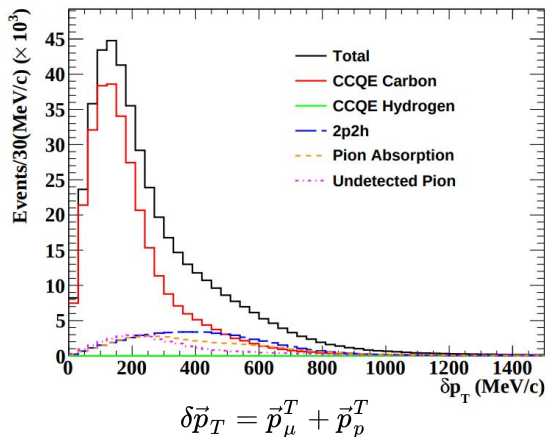
- The outgoing hadrons from a neutrino-nucleon interaction can re-interact with the remaining nucleus
- Ongoing work to improve the nuclear model for FSI of nucleons in neutrino simulations using INCL, one of the most predictive existing intranuclear-cascade models
→ crucial for T2K Near Detector Upgrade & next generation experiments



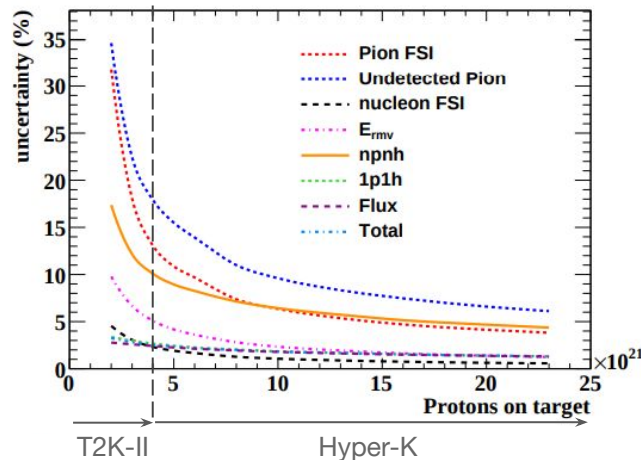


[arXiv:2108.11779](https://arxiv.org/abs/2108.11779)

Transverse momentum imbalance



Sensitivity to syst. uncertainties



- Using the outgoing proton information (eg. *visible energy & transverse momentum imbalance*) allows a better reconstruction of the neutrino energy and constraint on systematic uncertainties
- The dominant systematic uncertainties can be constrained down to the few-% level as required by future oscillation analyses of T2K-II and next generation experiments



Summary and prospects

- BSM-Nu project regroups several neutrino experiments around a common goal: access BSM physics through the characterisation of neutrino nature and mixing
- Long-baseline experiments aim at measuring neutrino oscillation parameters
- To probe CP violation, we need to improve the current detector technologies as well as the modeling of neutrino-nucleus interactions
- Contributions of BSM-Nu WP2:
 - Detectors: T2K Near Detector upgrade → High-Angle TPCs
 - Neutrino-nucleus interaction model: improvement of the initial state nuclear model (Spectral Function model) and ongoing work to improve the modeling of final state interactions
- Prospects of BSM-Nu WP2:
 - Preparation of a new fitting framework for oscillation analyses with new samples and observables in the era of T2K-II and beyond
 - Installation and commissioning of T2K Near Detector Upgrade at the end of 2022





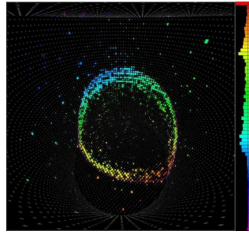
Back-up

Neutrino flavors and masses mixing

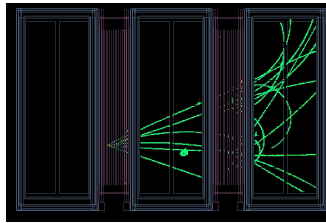
40

- Neutrino flavor states: $\nu_e, \nu_\mu, \nu_\tau \rightarrow$ **production** and **detection**
- Neutrino mass states: $\nu_1, \nu_2, \nu_3 \rightarrow$ **propagation**

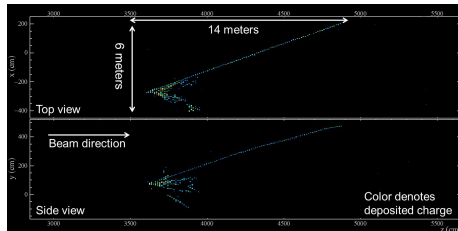
Super-Kamiokande



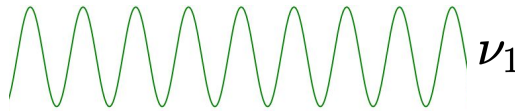
T2K Near Detector



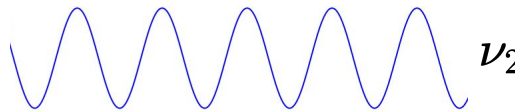
NOvA Far Detector



$$|\nu_i\rangle = \sum_{\alpha=e,\mu,\tau} U_{\alpha i} |\nu_\alpha\rangle$$



ν_1



ν_2



ν_3

Sun

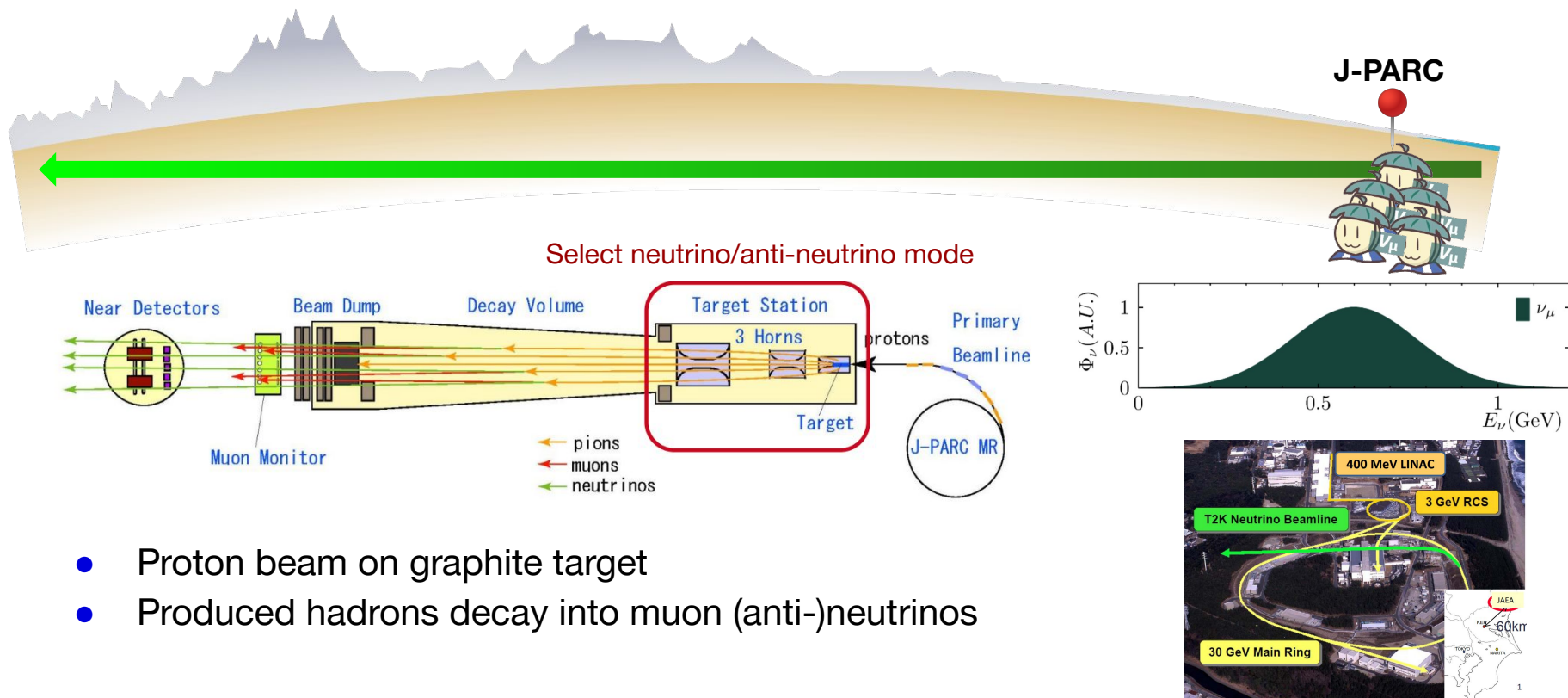


Reactors



Accelerators

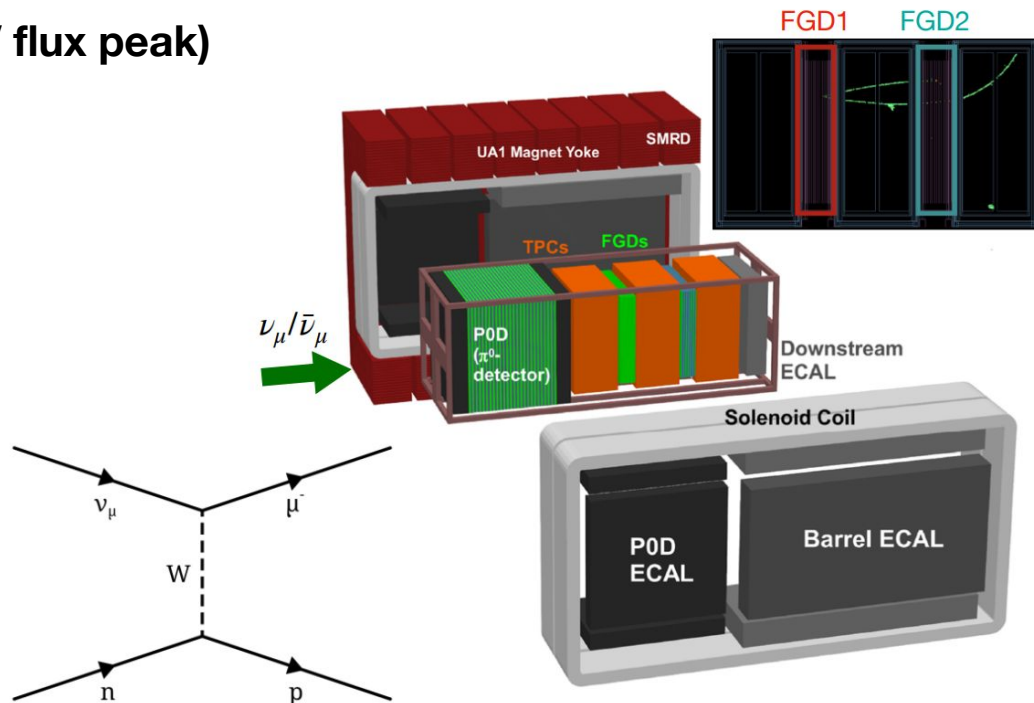




- Proton beam on graphite target
- Produced hadrons decay into muon (anti-)neutrinos

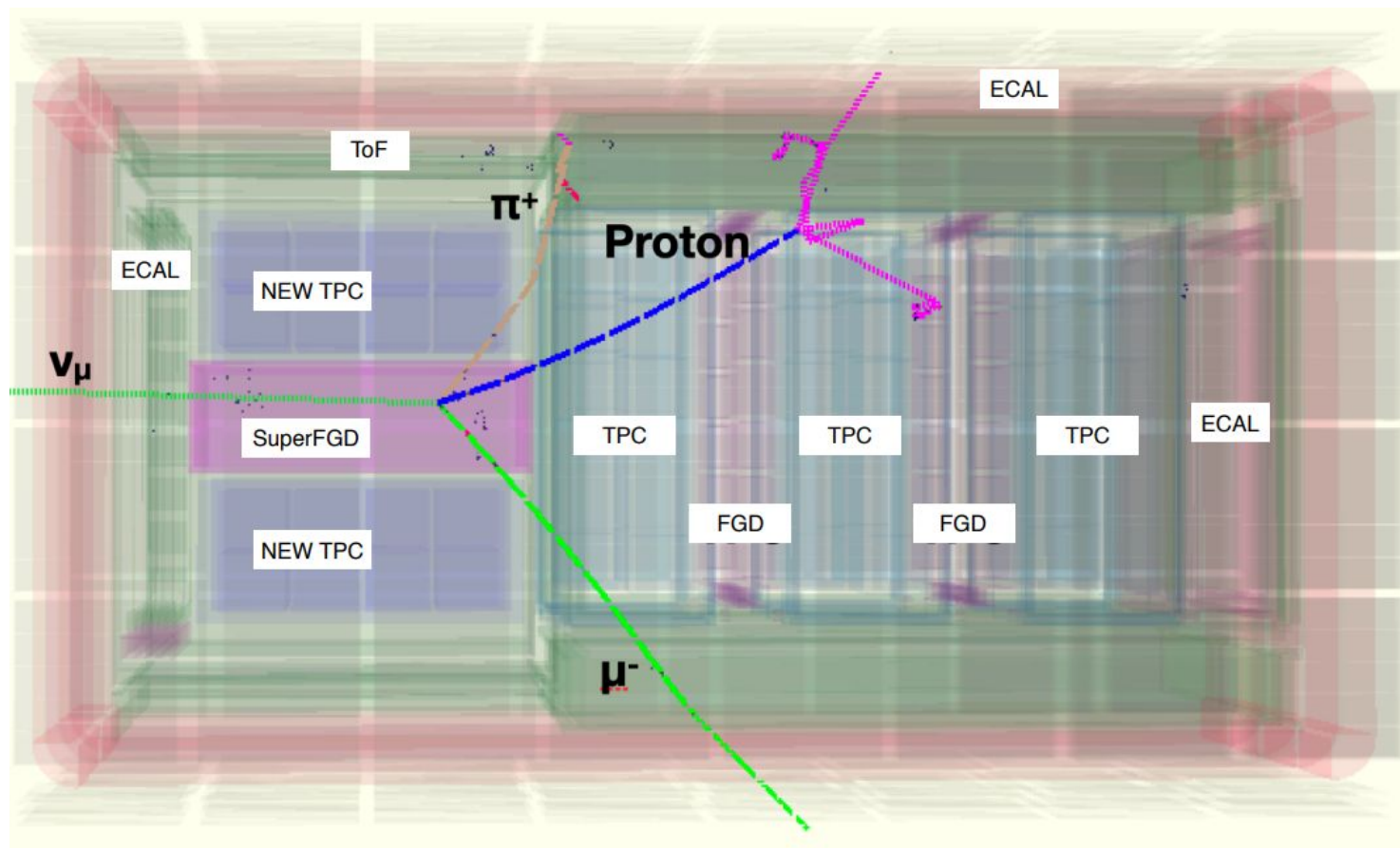
Off-axis detector (2.5 degrees, 0.6 GeV flux peak)

- Fine Grained Detectors (FGDs)
 - Plastic scintillator tracker
 - FGD1 & 2 carbon target (CH)
 - FGD2 has water target layers
- Time Projection Chambers (TPCs)
 - Tracking detectors
 - Charged particle momentum
 - Particle ID
- Very good measurements of muon kinematics



Event in the T2K Upgraded Near Detector

43



- Neutrons can be detected via their re-interaction within the detector
- If the path is long enough, neutron energy can be measured using time of flight
- Resolution: $\sim 15\text{-}30\%$

