The Divine Master model

Julien Lesgourgues, 25.11.2021, LAPTh

Quintessential / Fluid / Fuzzy Dark Matter (2000-2003)

Galactic halos of fluid dark matter#1Alexandre Arbey (Annecy, LAPTH and Savoie U.), Julien Lesgourgues (CERN and Annecy, LAPTH), Pierre Salati (Annecy, LAPTH and Savoie U.) (Jan, 2003)Published in: Phys.Rev.D 68 (2003) 023511 • e-Print: astro-ph/0301533 [astro-ph]					
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Hui, Ostriker, Tremaine, Witten 2016

	Ultraligh	nt scalars a	as cosmological dark matter	#1
6	 Lam Hui (Columbia U.), Jeremiah P. Ostriker (Columbia U. and Princeton U. Observ.), Scott Tremaine (Princeton, Inst. Advanced Study), Edward Witten (Princeton, Inst. Advanced Study) 2016) Published in: <i>Phys.Rev.D</i> 95 (2017) 4, 043541 • e-Print: 1610.08297 [astro-ph.CO] 			Scott d Study) (Oct 26,
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 #4

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 #4

 Published in: Phys.Rev.D 64 (2001) 123528 • e-Print: astro-ph/0105564 [astro-ph]
 #4

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$$\mathcal{S} = \int \sqrt{-g} \, d^4 x \, \mathcal{L}\left\{\Phi, \partial_\mu \Phi\right\} = \int \sqrt{-g} \, d^4 x \, \left\{g^{\mu\nu} \, \partial_\mu \Phi^\dagger \, \partial_\nu \Phi \, - \, U\left(\Phi\right)\right\} \quad \text{with} \quad U = m^2 \, \Phi^\dagger \, \Phi$$



Cosmological constraints on quintessential halos

Alexandre Arbey (Annecy, LAPTH and Savoie U.), Julien Lesgourgues (CERN and Annecy, LAPTH), Pierre Salati (Annecy, LAPTH and Savoie U.) (Dec, 2001) Published in: *Phys.Rev.D* 65 (2002) 083514 • e-Print: astro-ph/0112324 [astro-ph]

$$\mathcal{L} = g^{\mu\nu} \partial_{\mu} \phi^{\dagger} \partial_{\nu} \phi - V(\phi) \qquad V = m^2 \phi^{\dagger} \phi + \lambda \left\{ \phi^{\dagger} \phi \right\}^2$$

Cosmological evolution of quintessential DM background, for $\lambda = 0$ and $m \sim 10^{-23}$ eV



Galactic halos of fluid dark matter

Alexandre Arbey (Annecy, LAPTH and Savoie U.), Julien Lesgourgues (CERN and Annecy, LAPTH), Pierre Salati (Annecy, LAPTH and Savoie U.) (Jan, 2003)

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for the formula for the for

$$\mathcal{L} = g^{\mu\nu} \partial_{\mu} \phi^{\dagger} \partial_{\nu} \phi - V(\phi) , \qquad V(\phi) = m^2 \varphi^{\dagger} \phi + \lambda \left\{ \phi^{\dagger} \phi \right\}^2$$

