

# No channel left behind

## Precision with $Vh$ at LHC and FCC-hh

Planck Conference 2022

*2 June 2022*

*Paris, France*

**Alejo N. Rossia**

*Department of Physics and Astronomy*

*University of Manchester*

With F. Bishara, S. De Curtis, L. Delle Rose, P. Englert, C. Grojean, M. Montull, G. Panico.

arXiv 2004.06122 (JHEP 07 (2020) 075)

arXiv 2011.13941 (JHEP 04 (2021) 154)

arXiv 22XX.YYYYYY



The University of Manchester

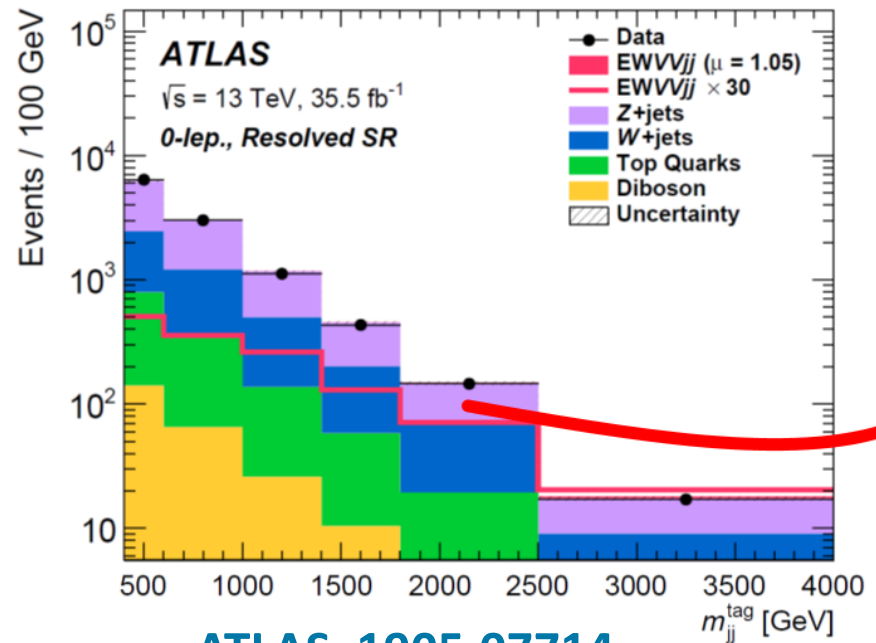


# A trick of the tail

Precision with hadron colliders? Yes!

# A trick of the tail

Precision with hadron colliders? Yes!



ATLAS, 1905.07714

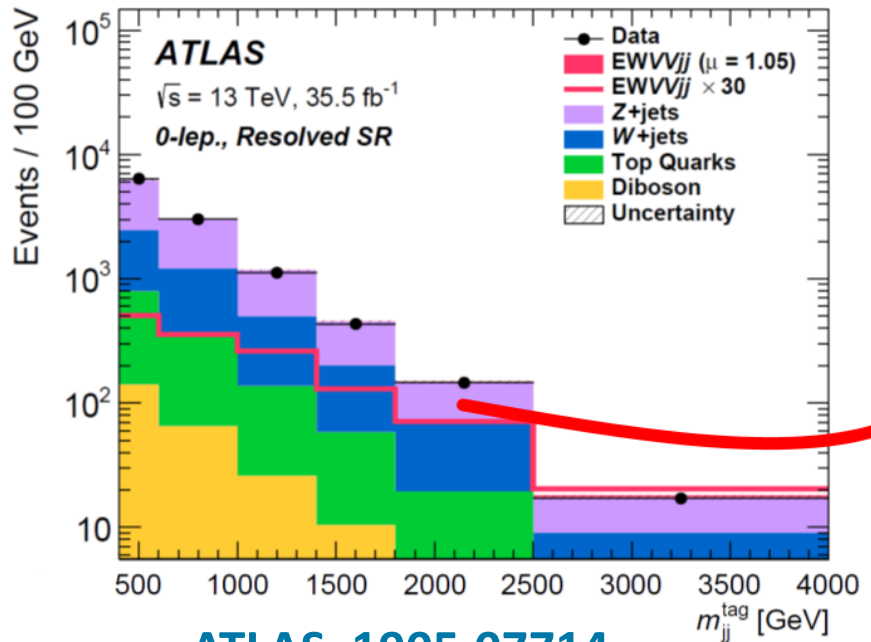
Clean channels + NP effects that grow with E



Tail hunting!

# A trick of the tail

Precision with hadron colliders? Yes!



Clean channels + NP effects that grow with E



Tail hunting!

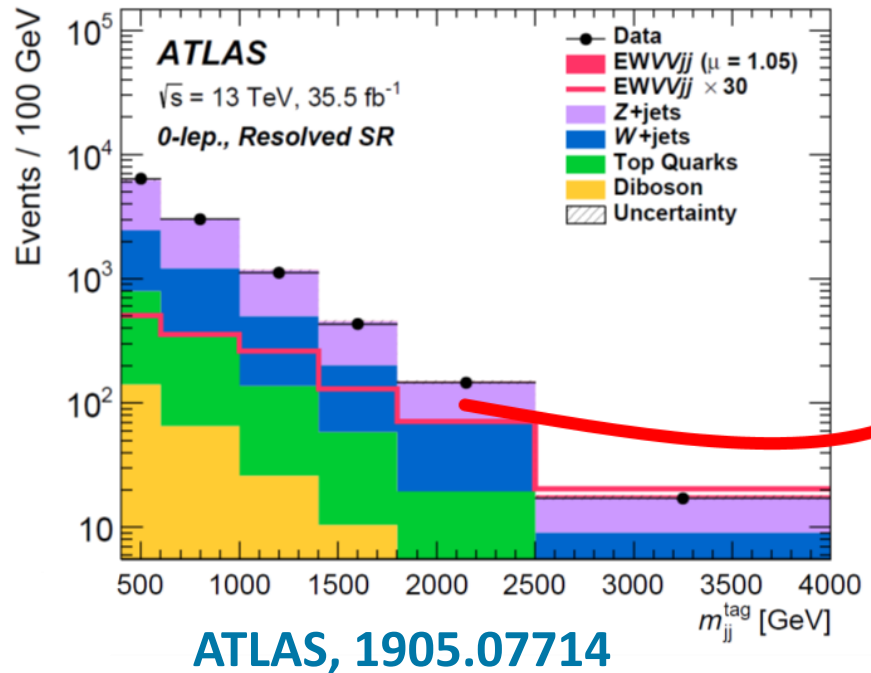
Heavy New Physics



Effective Field Theories

# A trick of the tail

Precision with hadron colliders? Yes!



Clean channels + NP effects that grow with E



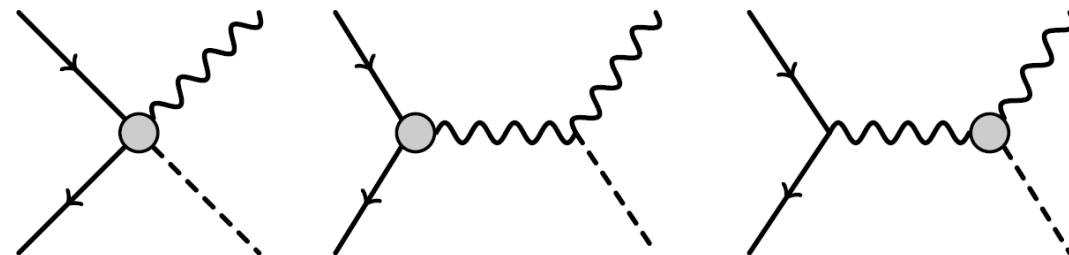
Tail hunting!

Heavy New Physics



Effective Field Theories

Diboson processes offer a window into EW and Higgs dynamics.



# New collider, new possibilities

For  $p_T^h > 550$  GeV:

$$pp \rightarrow W^\pm h$$

Higgs decay	Higgs BR	$n_{HL-LHC}$	$n_{HE-LHC}$	$n_{FCC-hh}$
$\bar{b}b$	$6 \cdot 10^{-1}$	$10^3$	$10^4$	$10^5$
$\tau\tau$	$6 \cdot 10^{-2}$	$10^2$	$10^3$	$10^4$
$\gamma\gamma$	$2 \cdot 10^{-3}$	$10^0$	$10^2$	$10^3$
$4l$	$2 \cdot 10^{-3}$	$10^0$	$10^2$	$10^3$
$\mu\mu$	$4 \cdot 10^{-4}$	$10^0$	$10^1$	$10^2$

# New collider, new possibilities

For  $p_T^h > 550$  GeV:

$$pp \rightarrow W^\pm h$$

	Higgs decay	Higgs BR	$n_{HL-LHC}$	$n_{HE-LHC}$	$n_{FCC-hh}$	$\frac{s}{\sqrt{s+b}} \ll 1$
<b>Today</b>	$\bar{b}b$	$6 \cdot 10^{-1}$	$10^3$	$10^4$	$10^5$	
	$\tau\tau$	$6 \cdot 10^{-2}$	$10^2$	$10^3$	$10^4$	
	$\gamma\gamma$	$2 \cdot 10^{-3}$	$10^0$	$10^2$	$10^3$	
	$4l$	$2 \cdot 10^{-3}$	$10^0$	$10^2$	$10^3$	
	$\mu\mu$	$4 \cdot 10^{-4}$	$10^0$	$10^1$	$10^2$	

# New collider, new possibilities

For  $p_T^h > 550$  GeV:

$$pp \rightarrow W^\pm h$$

	Higgs decay	Higgs BR	$n_{HL-LHC}$	$n_{HE-LHC}$	$n_{FCC-hh}$	$\frac{s}{\sqrt{s+b}} \ll 1$
<b>Today</b>	$\bar{b}b$	$6 \cdot 10^{-1}$	$10^3$	$10^4$	$10^5$	$\frac{s}{\sqrt{s+b}} \ll 1$
	$\tau\tau$	$6 \cdot 10^{-2}$	$10^2$	$10^3$	$10^4$	
<b>Future</b>	$\gamma\gamma$	$2 \cdot 10^{-3}$			$10^3$	$\frac{s}{\sqrt{s+b}} \approx 1$
	$4l$	$2 \cdot 10^{-3}$	$10^0$	$10^2$	$10^3$	
	$\mu\mu$	$4 \cdot 10^{-4}$	$10^0$	$10^1$	$10^2$	



# New collider, new possibilities

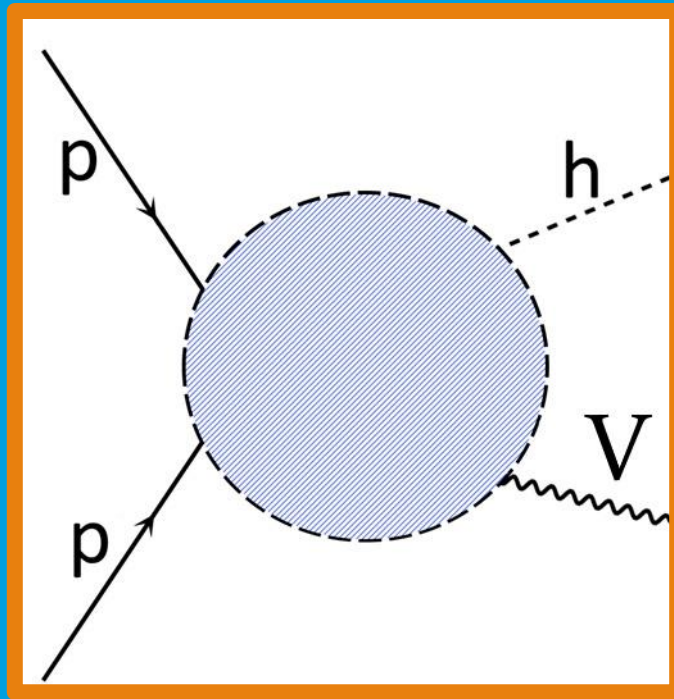
For  $p_T^h > 550$  GeV:

$$pp \rightarrow W^\pm h$$

	Higgs decay	Higgs BR	$n_{HL-LHC}$	$n_{HE-LHC}$	$n_{FCC-hh}$	$\frac{s}{\sqrt{s+b}} \ll 1$
<b>Today</b>	$\bar{b}b$	$6 \cdot 10^{-1}$	$10^3$	$10^4$	$10^5$	
	$\tau\tau$	$6 \cdot 10^{-2}$	$10^2$	$10^3$	$10^4$	
<b>Future</b>	$\gamma\gamma$	$2 \cdot 10^{-3}$	[shaded]		$10^3$	
	$4l$	$2 \cdot 10^{-3}$	$10^0$	$10^2$	$10^3$	
	$\mu\mu$	$4 \cdot 10^{-4}$	$10^0$	$10^1$	$10^2$	

**FCC-hh will open new channels. How will they compare with the known ones?**

# The process of interest.



$$pp \rightarrow Vh$$

# What New Physics can we probe?

Assumptions: SMEFT + Dim. 6 op. in Warsaw basis

$$\frac{c_{\varphi q}^{(1)}}{\Lambda^2} (\bar{Q}_L \gamma^\mu Q_L) \left( i H^\dagger \overleftrightarrow{D}_\mu H \right)$$

$$\frac{c_{\varphi u}}{\Lambda^2} (\bar{u}_R \gamma^\mu u_R) \left( i H^\dagger \overleftrightarrow{D}_\mu H \right)$$

$$\frac{c_{\varphi d}}{\Lambda^2} (\bar{d}_R \gamma^\mu d_R) \left( i H^\dagger \overleftrightarrow{D}_\mu H \right)$$

$$\frac{c_{\varphi q}^{(3)}}{\Lambda^2} (\bar{Q}_L \sigma^a \gamma^\mu Q_L) \left( i H^\dagger \sigma^a \overleftrightarrow{D}_\mu H \right)$$

$$\frac{c_{\varphi W}}{\Lambda^2} H^\dagger H W^{a, \mu\nu} W_{\mu\nu}^a$$

$$\frac{c_{\varphi \widetilde{W}}}{\Lambda^2} H^\dagger H W^{a, \mu\nu} \widetilde{W}_{\mu\nu}^a$$

$$\widetilde{W}^{a, \mu\nu} \equiv \frac{1}{2} \epsilon^{\mu\nu\rho\sigma} W_{\rho\sigma}^a$$

# What New Physics can we probe?

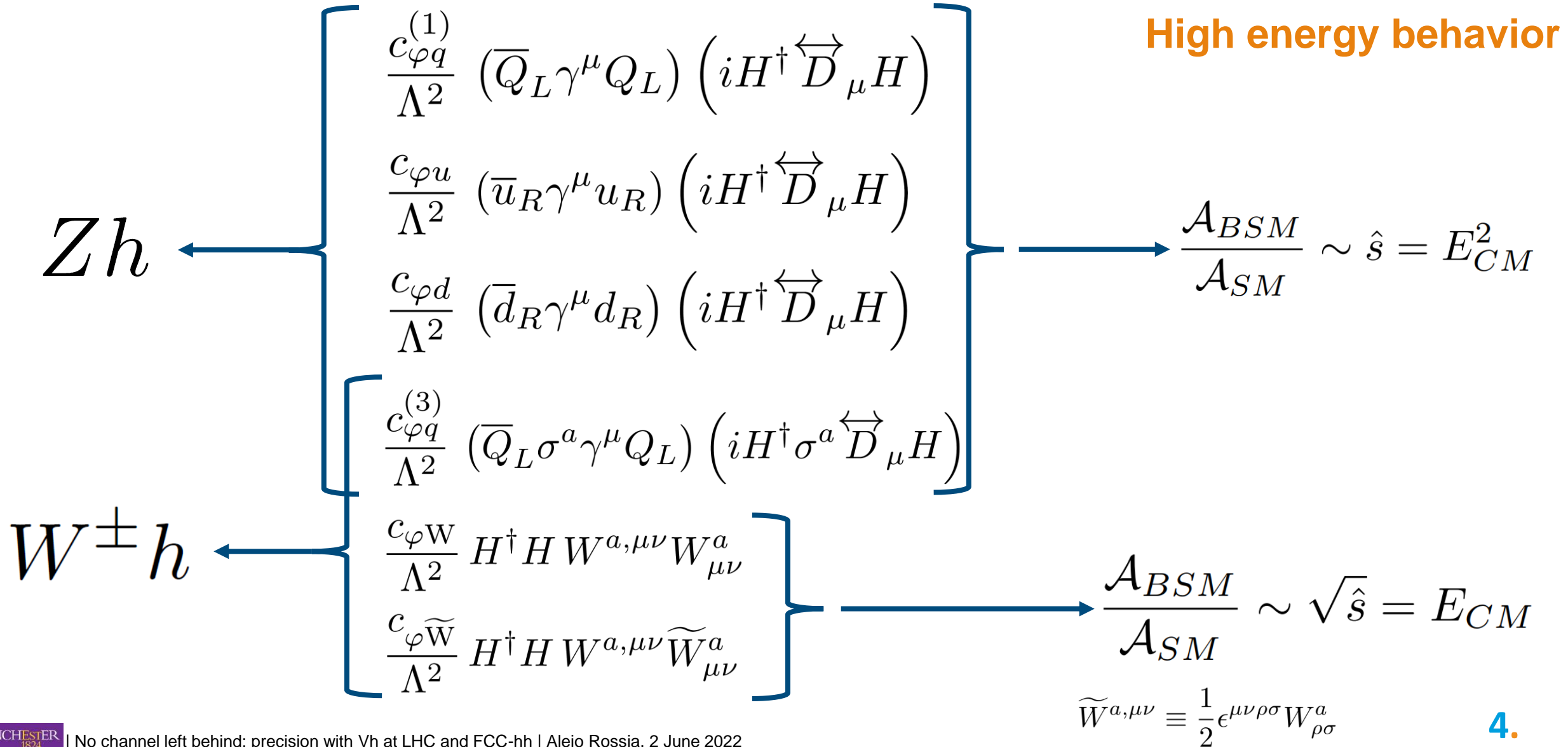
Assumptions: SMEFT + Dim. 6 op. in Warsaw basis

$$\begin{array}{l}
 Zh \\
 \\
 \\
 \\
 W^\pm h
 \end{array}
 \left\{
 \begin{array}{l}
 \frac{c_{\varphi q}^{(1)}}{\Lambda^2} (\bar{Q}_L \gamma^\mu Q_L) \left( i H^\dagger \overleftrightarrow{D}_\mu H \right) \\
 \frac{c_{\varphi u}}{\Lambda^2} (\bar{u}_R \gamma^\mu u_R) \left( i H^\dagger \overleftrightarrow{D}_\mu H \right) \\
 \frac{c_{\varphi d}}{\Lambda^2} (\bar{d}_R \gamma^\mu d_R) \left( i H^\dagger \overleftrightarrow{D}_\mu H \right) \\
 \frac{c_{\varphi q}^{(3)}}{\Lambda^2} (\bar{Q}_L \sigma^a \gamma^\mu Q_L) \left( i H^\dagger \sigma^a \overleftrightarrow{D}_\mu H \right) \\
 \frac{c_{\varphi W}}{\Lambda^2} H^\dagger H W^{a,\mu\nu} W_{\mu\nu}^a \\
 \frac{c_{\varphi \widetilde{W}}}{\Lambda^2} H^\dagger H W^{a,\mu\nu} \widetilde{W}_{\mu\nu}^a
 \end{array}
 \right.$$

$$\widetilde{W}^{a,\mu\nu} \equiv \frac{1}{2} \epsilon^{\mu\nu\rho\sigma} W_{\rho\sigma}^a$$

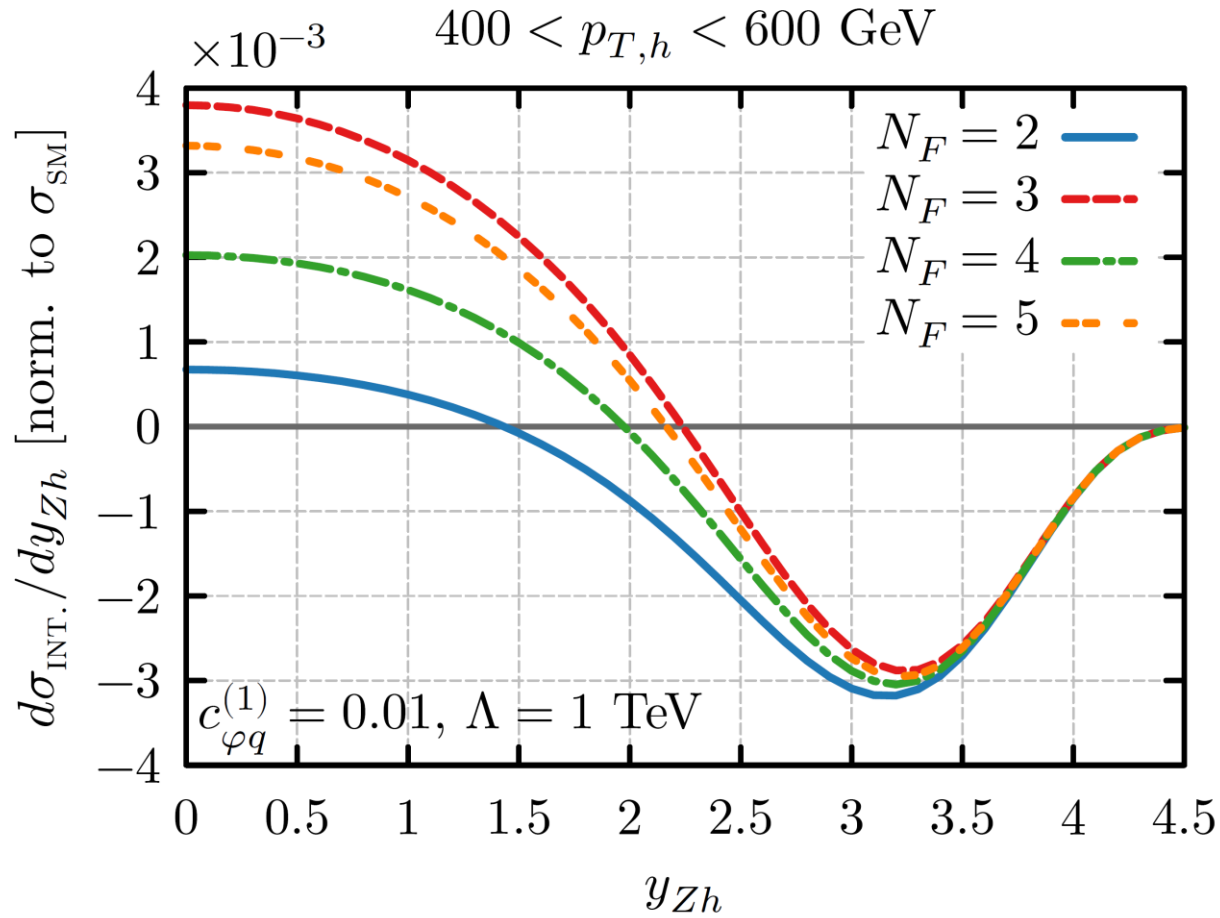
# What New Physics can we probe?

Assumptions: SMEFT + Dim. 6 op. in Warsaw basis



Zh.

# Double binning for the win

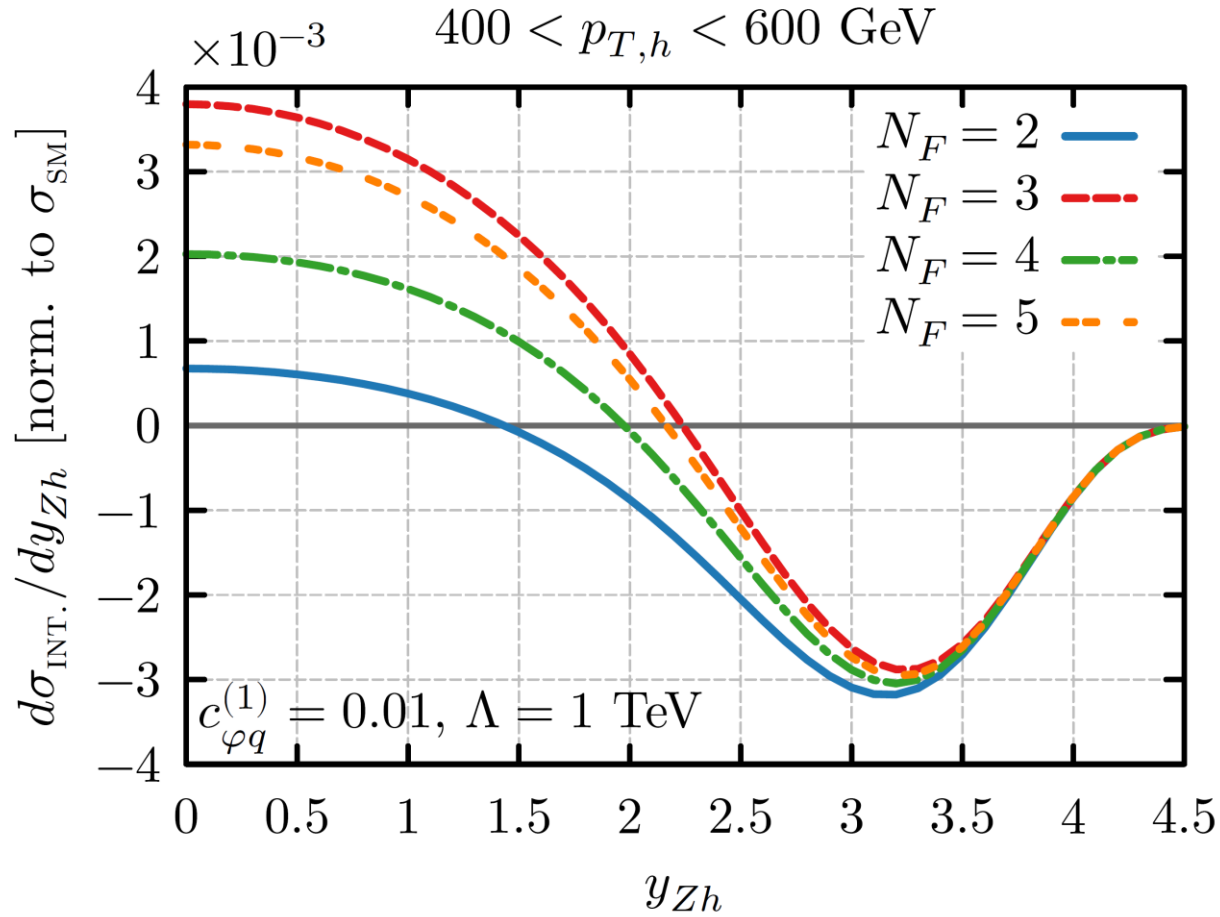


$$\sigma_{\mathcal{O}_{\varphi q}^{(1)}}^{\text{int}} \propto s_W^2 Q - T_3$$

**Cancellation of up and down contributions**

# Zh.

## Double binning for the win



$$\sigma_{\mathcal{O}_{\varphi q}^{(1)}}^{\text{int}} \propto s_W^2 Q - T_3$$

**Cancellation of up and down contributions**

**Differential in  $p_T$  and rapidity (only FCC-hh)**

$$\text{Min}\{p_T^h, p_T^Z\} \in \{200, 400, 600, 800, 1000, \infty\} \text{ GeV}$$

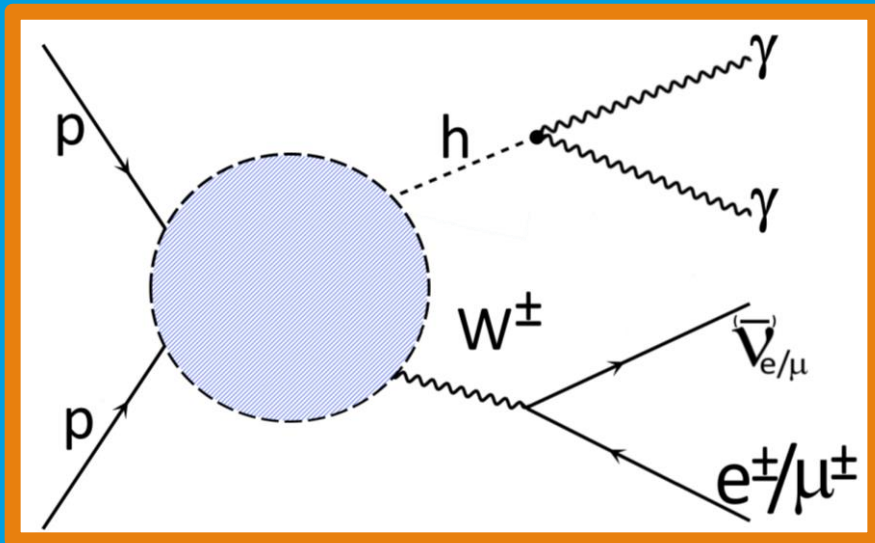
$$|y_{Zh}| \in [0, 2), [2, 6]$$

(Slightly different rapidity binning for  $Z \rightarrow \nu\bar{\nu}$ )



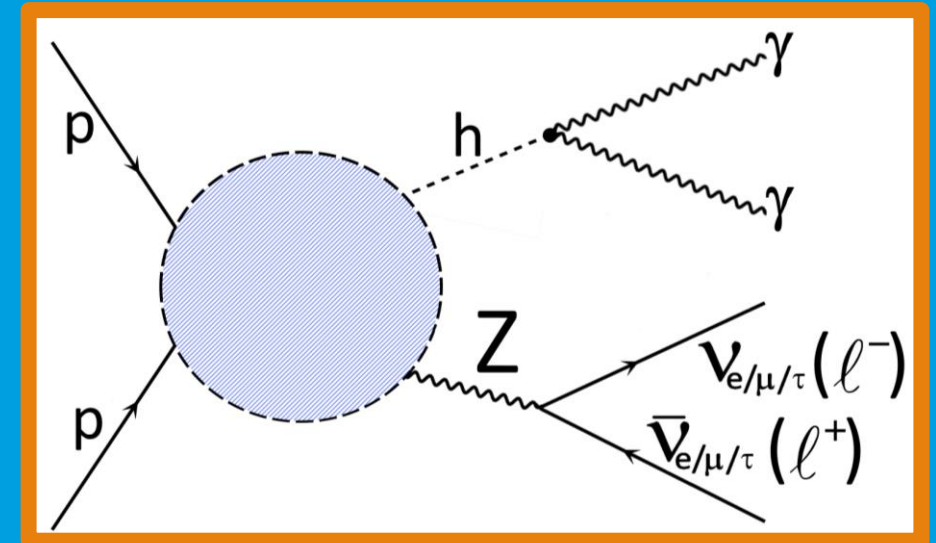
# Diphotonic Vh, a summary.

arXiv 2004.06122 (JHEP 07 (2020) 075)



$$pp \rightarrow W^\pm h \rightarrow l^\pm \nu \gamma \gamma$$

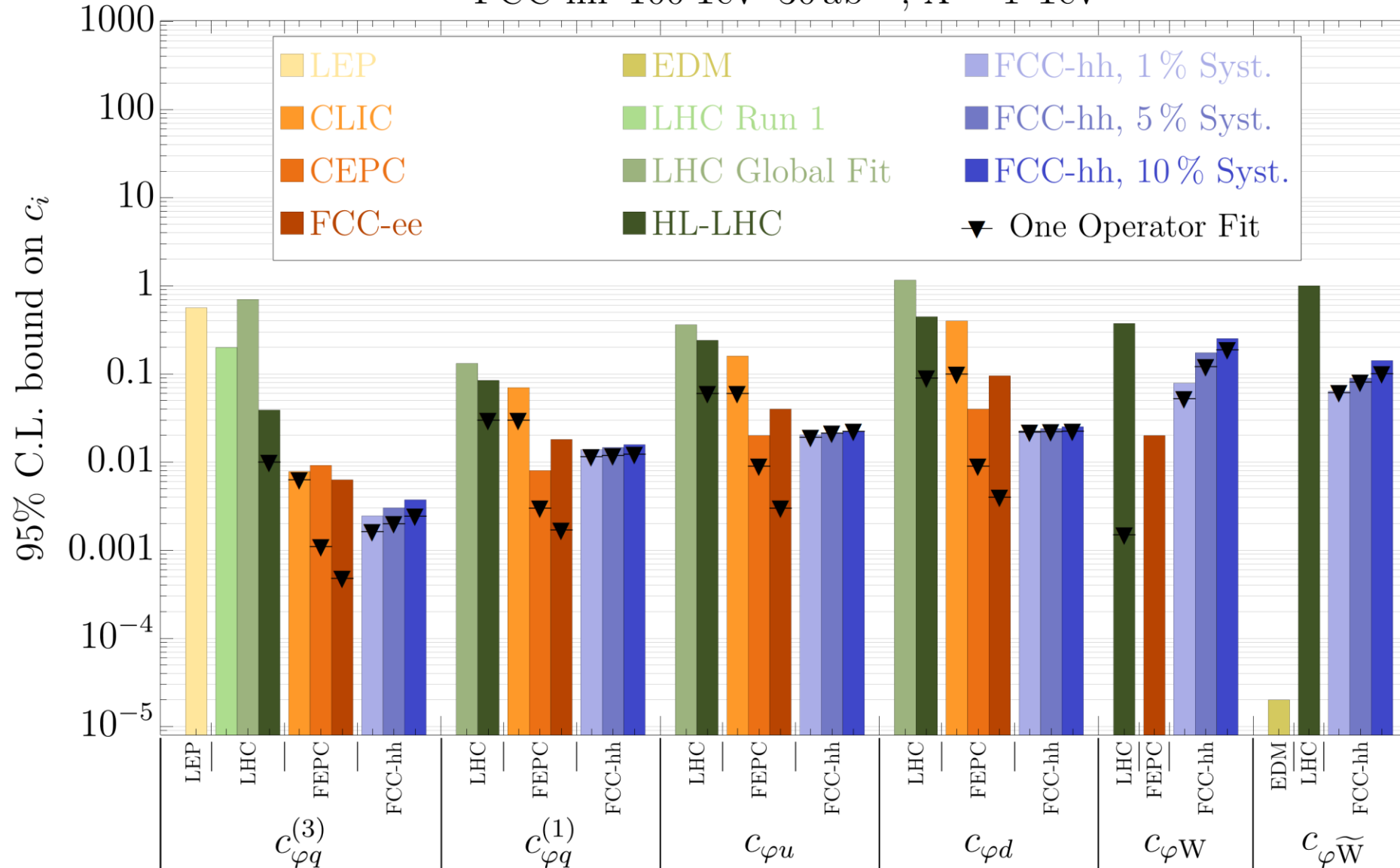
arXiv 2011.13941 (JHEP 04 (2021) 154)



$$pp \rightarrow Zh \rightarrow l^+ l^- (\nu \bar{\nu}) \gamma \gamma$$

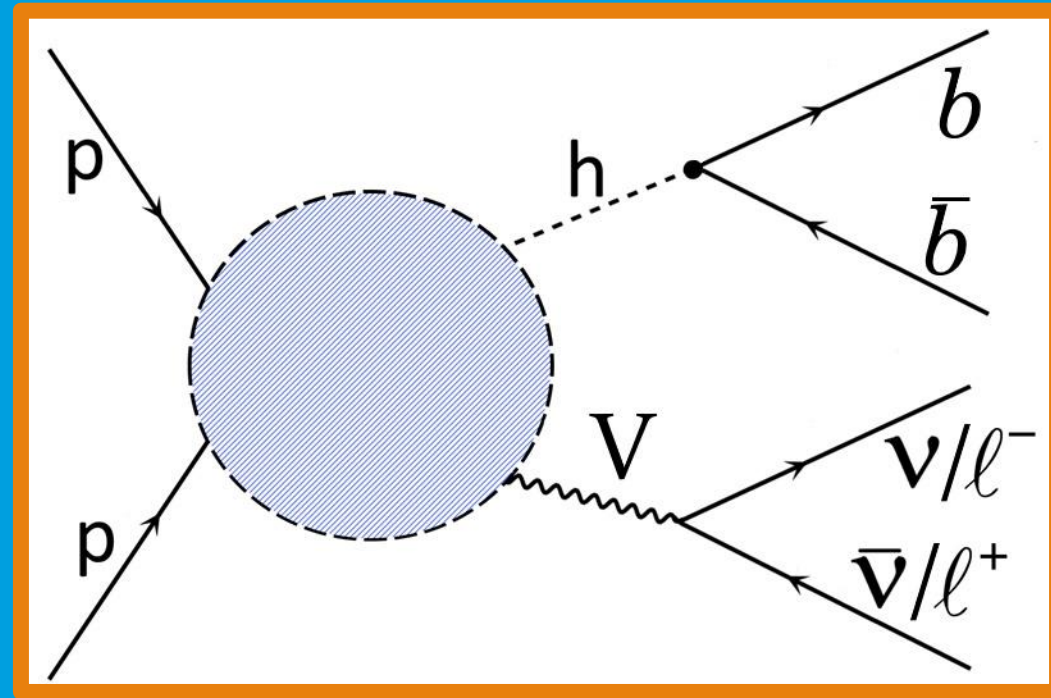
# Diphotonic channel in perspective

FCC-hh 100 TeV 30 ab<sup>-1</sup>,  $\Lambda = 1$  TeV



# Let them be quarks, $Vh$ .

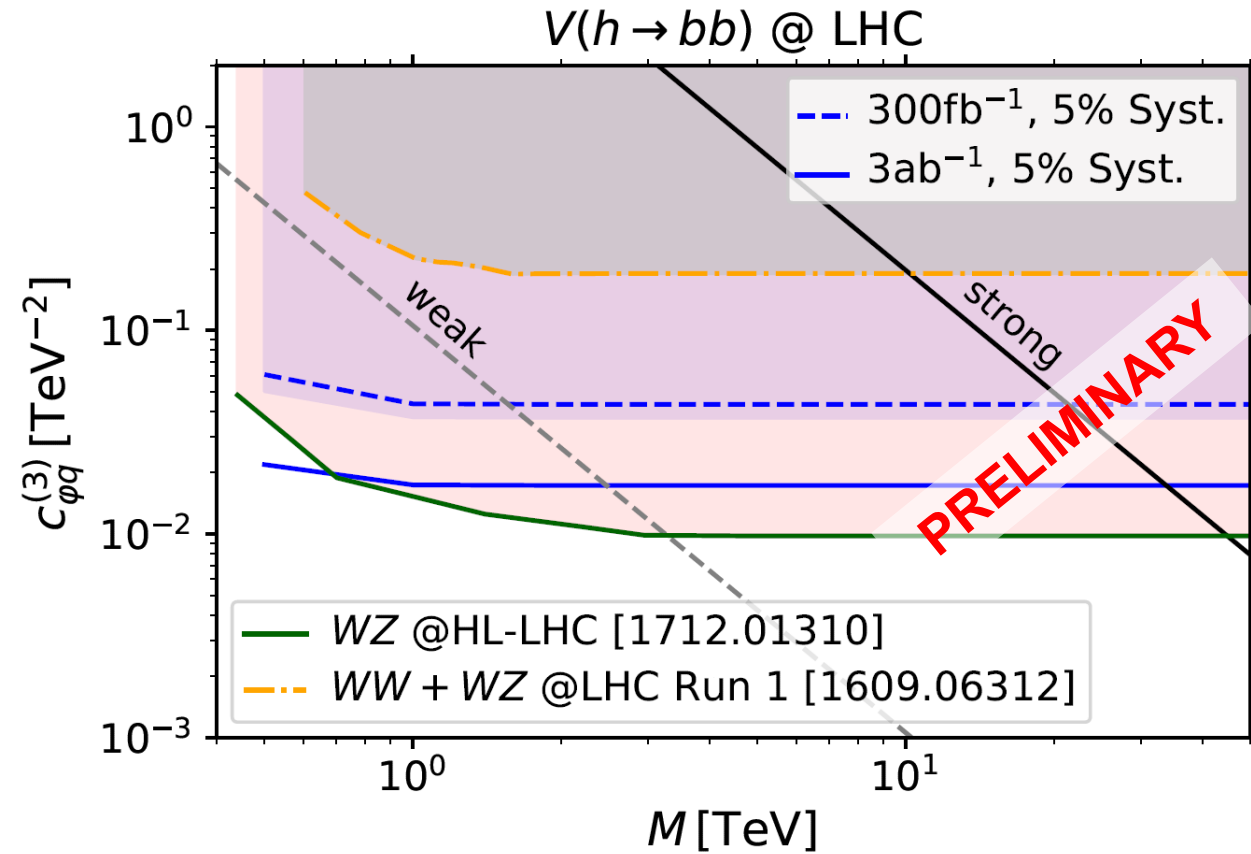
arXiv 22XX.YYYZZ



$$pp \rightarrow Vh \rightarrow \ell(\nu)\ell(\nu)b\bar{b}$$

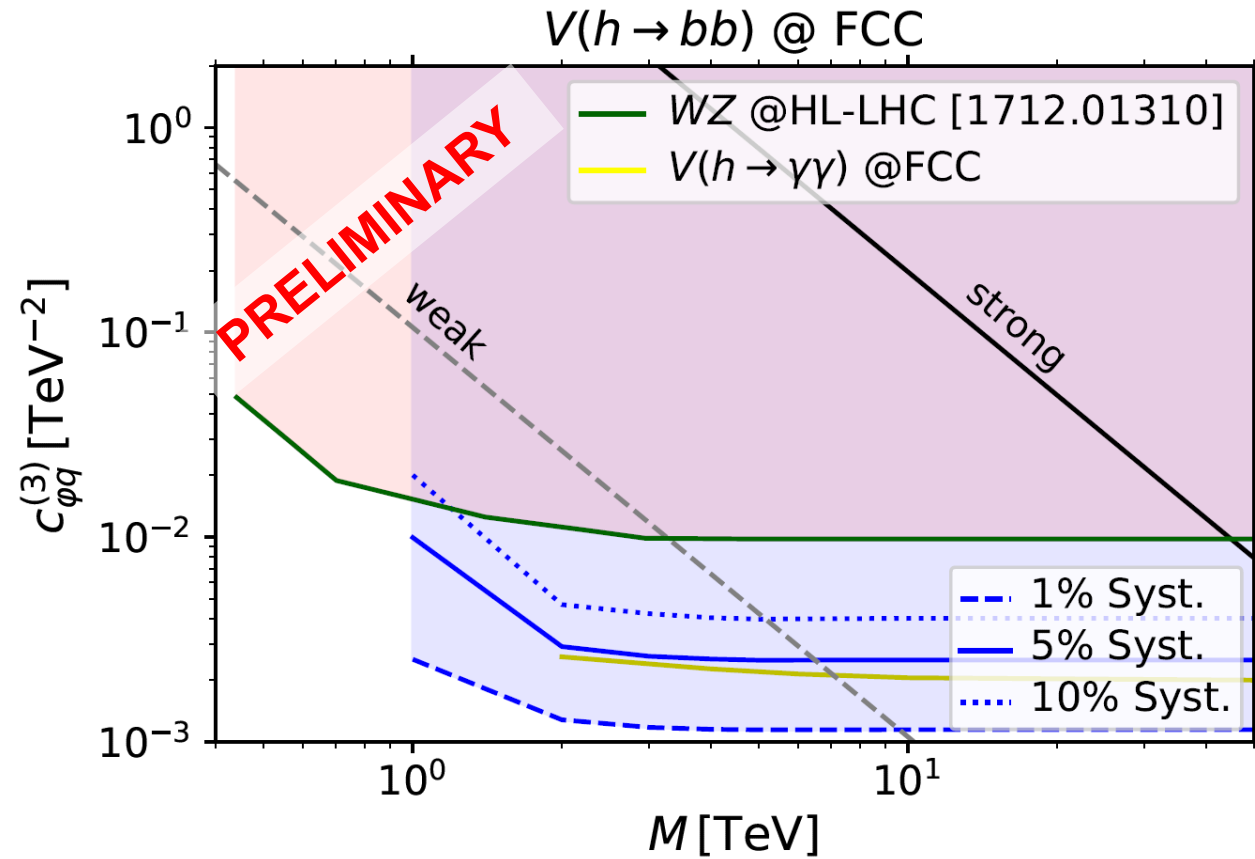
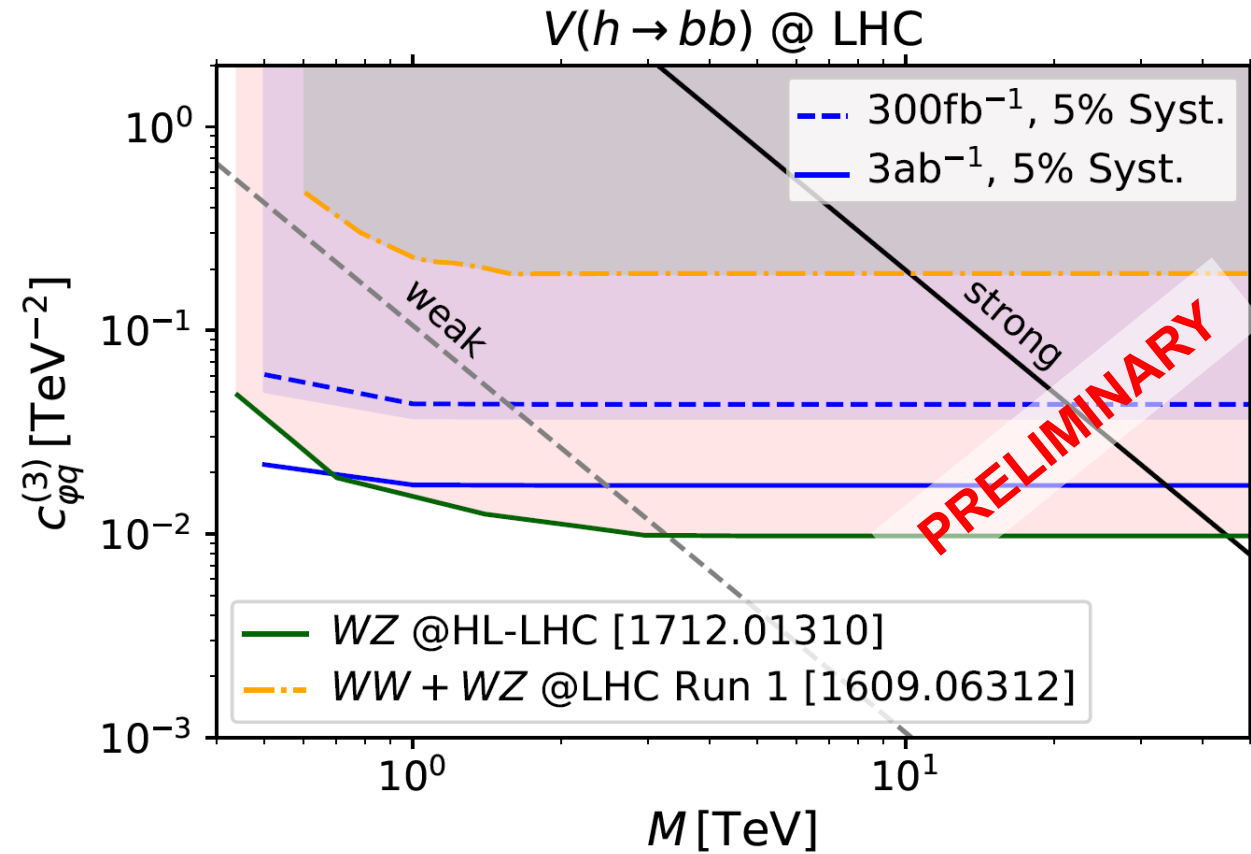
# Vh.

## 95% C.L. on $c_{\varphi q}^{(3)}$ with $h \rightarrow b\bar{b}$



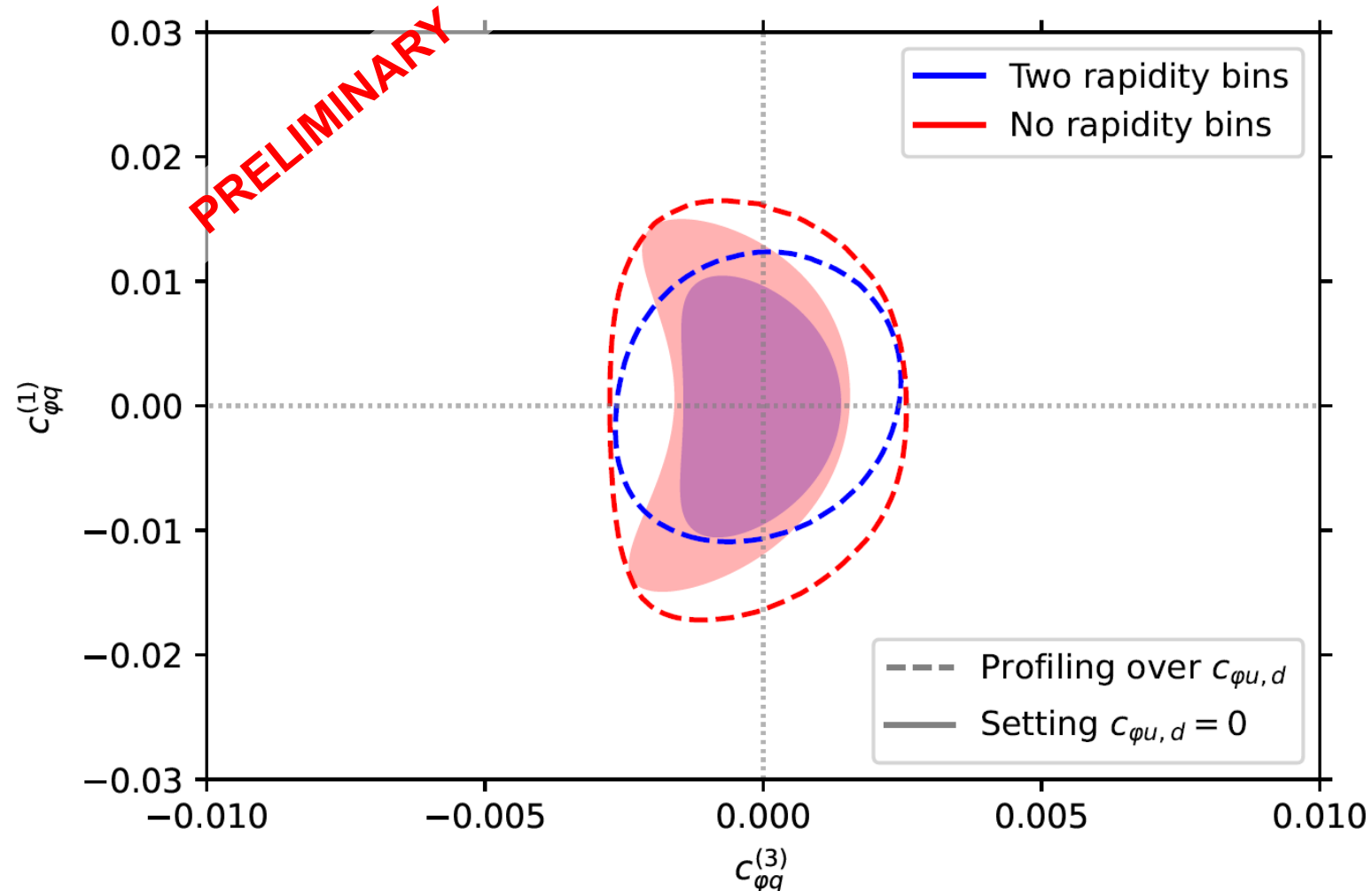
# Vh.

## 95% C.L. on $c_{\varphi q}^{(3)}$ with $h \rightarrow b\bar{b}$



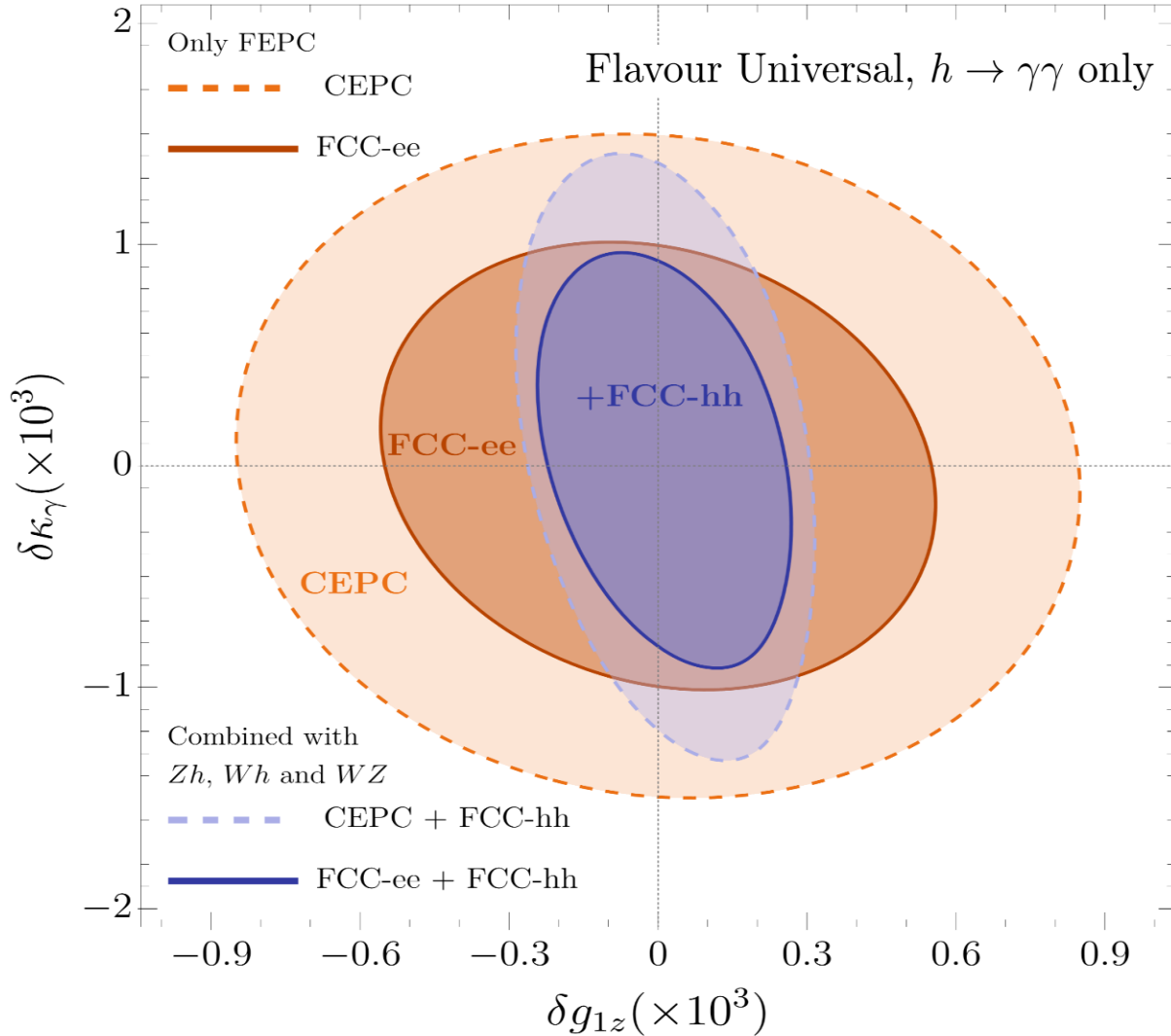
# Rapidity binning effects

$Zh$  at FCC-hh 100 TeV,  $30 \text{ ab}^{-1}$ , 1% syst.



# Sizeable impact on aTGC bounds

FCC-hh 100 TeV 30 ab<sup>-1</sup>, 95% C.L., 5% Syst.



$$c_{\varphi q}^{(3)} = + \frac{\Lambda^2}{4m_W^2} g^2 (\delta g_L^{Zu} - \delta g_L^{Zd} - c_W^2 \delta g_{1z})$$

$$c_{\varphi q}^{(1)} = - \frac{\Lambda^2}{4m_W^2} g^2 \left( \delta g_L^{Zu} + \delta g_L^{Zd} + \frac{1}{3} (t_W^2 \delta \kappa_\gamma - s_W^2 \delta g_{1z}) \right)$$

$$c_{\varphi u} = - \frac{\Lambda^2}{2m_W^2} g^2 \left( \delta g_R^{Zu} + \frac{2}{3} (t_W^2 \delta \kappa_\gamma - s_W^2 \delta g_{1z}) \right)$$

$$c_{\varphi d} = - \frac{\Lambda^2}{2m_W^2} g^2 \left( \delta g_R^{Zd} - \frac{1}{3} (t_W^2 \delta \kappa_\gamma - s_W^2 \delta g_{1z}) \right)$$

# Conclusions

- At FCC-hh, there will be new diboson channels to do precision measurements, such as  $(W, Z) h$  with  $h \rightarrow \gamma\gamma$ .
- With a simple  $p_T$  binning, they offer competitive sensitivity to  $\mathcal{O}_{\varphi q}^{(3)}$ .
- In  $Zh$ , a binning in rapidity improves the sensitivity to  $\mathcal{O}_{\varphi q}^{(1)}$ .
- $h \rightarrow b\bar{b}$  allows to perform these studies at (HL-)LHC, but with limitations.
- $h \rightarrow \gamma\gamma$  and  $h \rightarrow b\bar{b}$  at FCC-hh achieve similar results in different ways.
- $Wh$  and  $Zh$  with  $h \rightarrow \gamma\gamma$  are not exploration channels, but important to probe different directions.



# Thank you for your attention

## Contact



The University of Manchester

[www.manchester.ac.uk](http://www.manchester.ac.uk)

Alejo N. Rossia

HEP Theory Group

Dept. Of Physics and Astronomy

The University of Manchester

E-mail: [alejo dot rossia at manchester dot ac dot uk](mailto:alejo_dot_rossia_at_manchester_dot_ac_dot_uk)

<http://www.hep.man.ac.uk/>

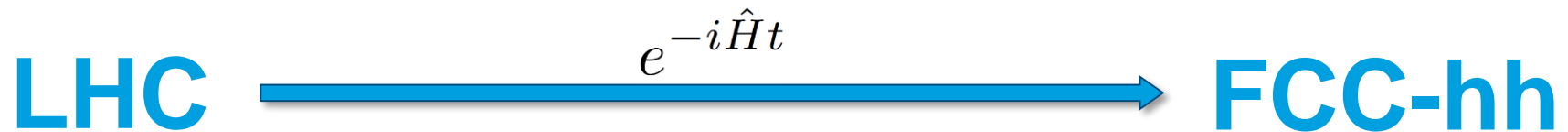
# Appendix.

For even more details, read our papers or contact us.

# Diboson in the present

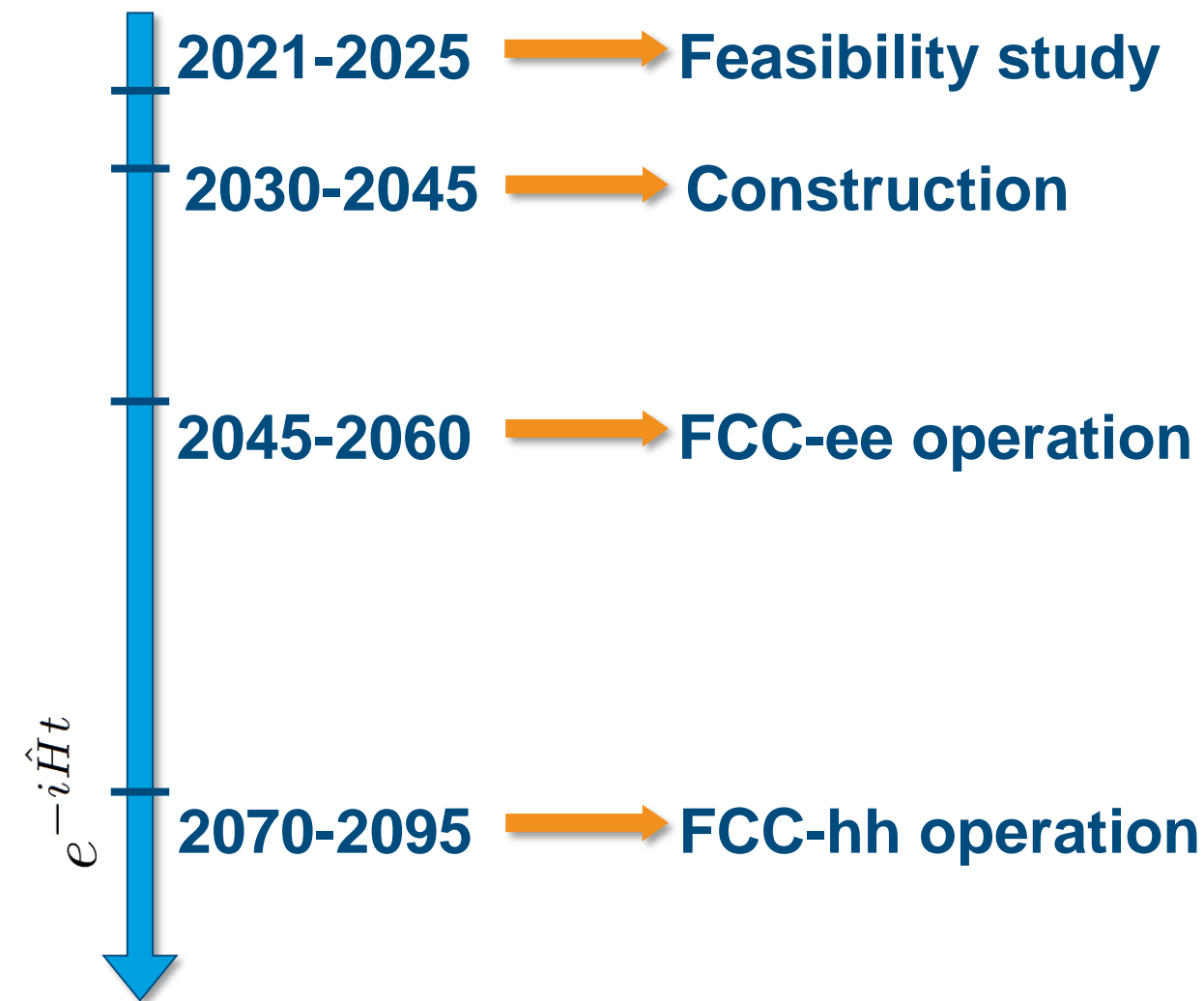
$(W/Z)h$  @ LHC {  
ATLAS, *Eur. Phys. J. C* 81 (2021) 2, 178, ArXiv: 2007.02873  
ATLAS, *Phys. Lett. B* 816 (2021) 136204, ArXiv: 2008.02508  
CMS, *JHEP*07 (2021) 027, ArXiv: 2103.06956

## What will change in the future?

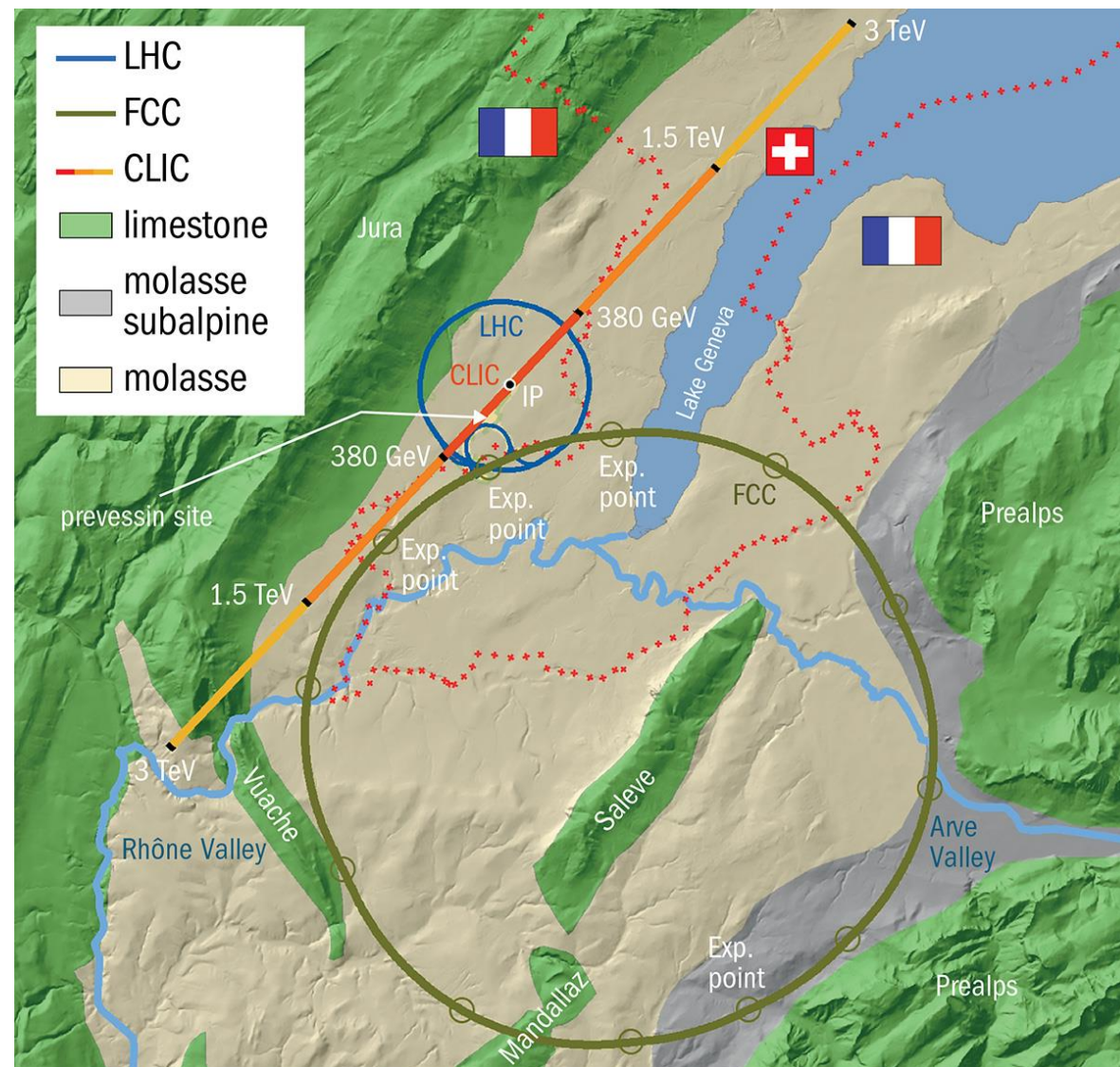


	HL-LHC	FCC-hh
C.o.M. energy	14 TeV	100 TeV
Int. Luminosity	3 ab <sup>-1</sup>	30 ab <sup>-1</sup>

# FCC-hh: The LHC of the future

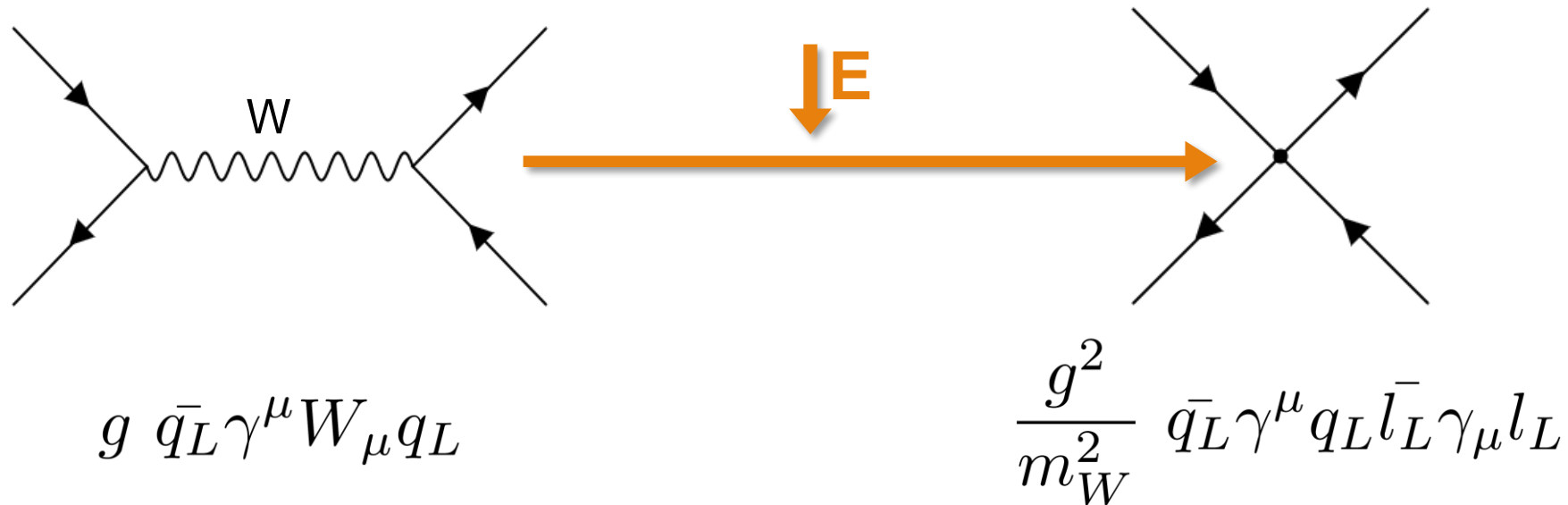


Timeline from talk by M. Benedikt (CERN) at FCC Workshop 2022



# Why Effective Field Theories?

- The main idea behind EFTs is in all fields of Physics.
- NP at a higher scale affect the interactions seen at a lower scale.



- Operators with dimension > 4 encode the NP effects in the EFT.
- Offer a more model-independent way of searching for NP.


# Standard Model EFT (SMEFT) and Interference

- Field content and gauge symmetries of the SM and linearly realized EW sym.
- Add gauge invariant operators with dimension bigger than 4.


$$\mathcal{L} = \mathcal{L}_{SM} + \sum_i \frac{c_i^{(6)}}{\Lambda^2} \mathcal{O}_i^{(6)} + \sum_i \frac{c_i^{(8)}}{\Lambda^4} \mathcal{O}_i^{(8)} + \dots$$

- Leading deviations from the SM appear at dimension 6.

$$\sigma = |\mathcal{M}_{SM}|^2 + 2\text{Re}(\mathcal{M}_{SM} \mathcal{M}_{BSM}^*) + |\mathcal{M}_{BSM}|^2$$



$\propto c_i^{(6)} / \Lambda^2$



$\propto (c_i^{(6)} / \Lambda^2)^2$

**Interference**

# Interference patterns

## High energy behaviour

$V$ polarization	SM	$\mathcal{O}_{\varphi f}$	$\mathcal{O}_{\varphi W}$	$\mathcal{O}_{\varphi \tilde{W}}$
$\lambda = 0$	1	$\frac{\hat{s}}{\Lambda^2}$	$\frac{M_W^2}{\Lambda^2}$	0
$\lambda = \pm$	$\frac{M_W}{\sqrt{\hat{s}}}$	$\frac{\sqrt{\hat{s}} M_W}{\Lambda^2}$	$\frac{\sqrt{\hat{s}} M_W}{\Lambda^2}$	$\frac{\sqrt{\hat{s}} M_W}{\Lambda^2}$

$V = W, Z$

$\mathcal{O}_{\varphi f} = \mathcal{O}_{\varphi q}^{(3)}, \mathcal{O}_{\varphi q}^{(1)}, \mathcal{O}_{\varphi u}, \mathcal{O}_{\varphi d}$

Differential in  $p_T$

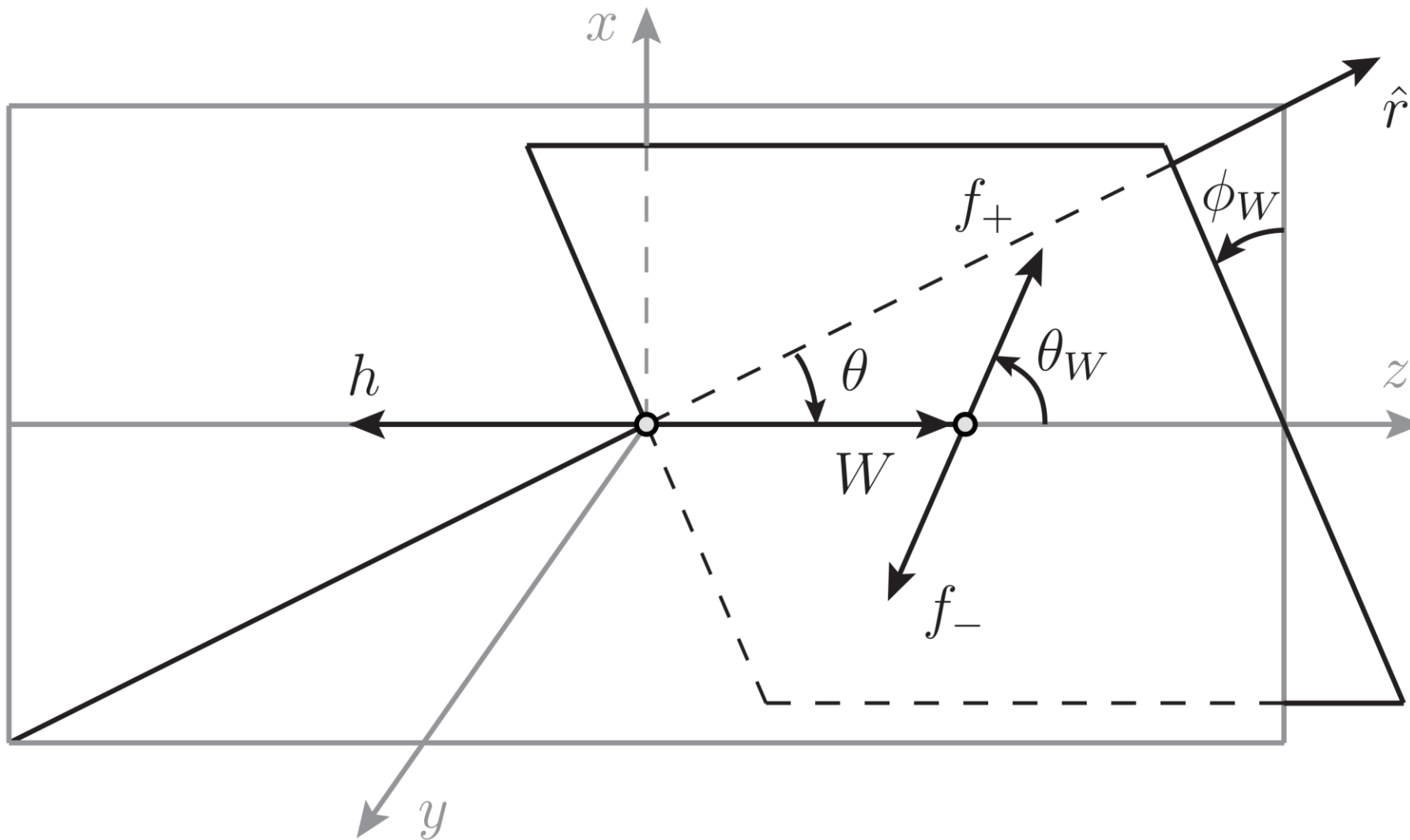


Interference between same polarisation

# Wh.

# Interference patterns

## Measuring angles resurrects interference

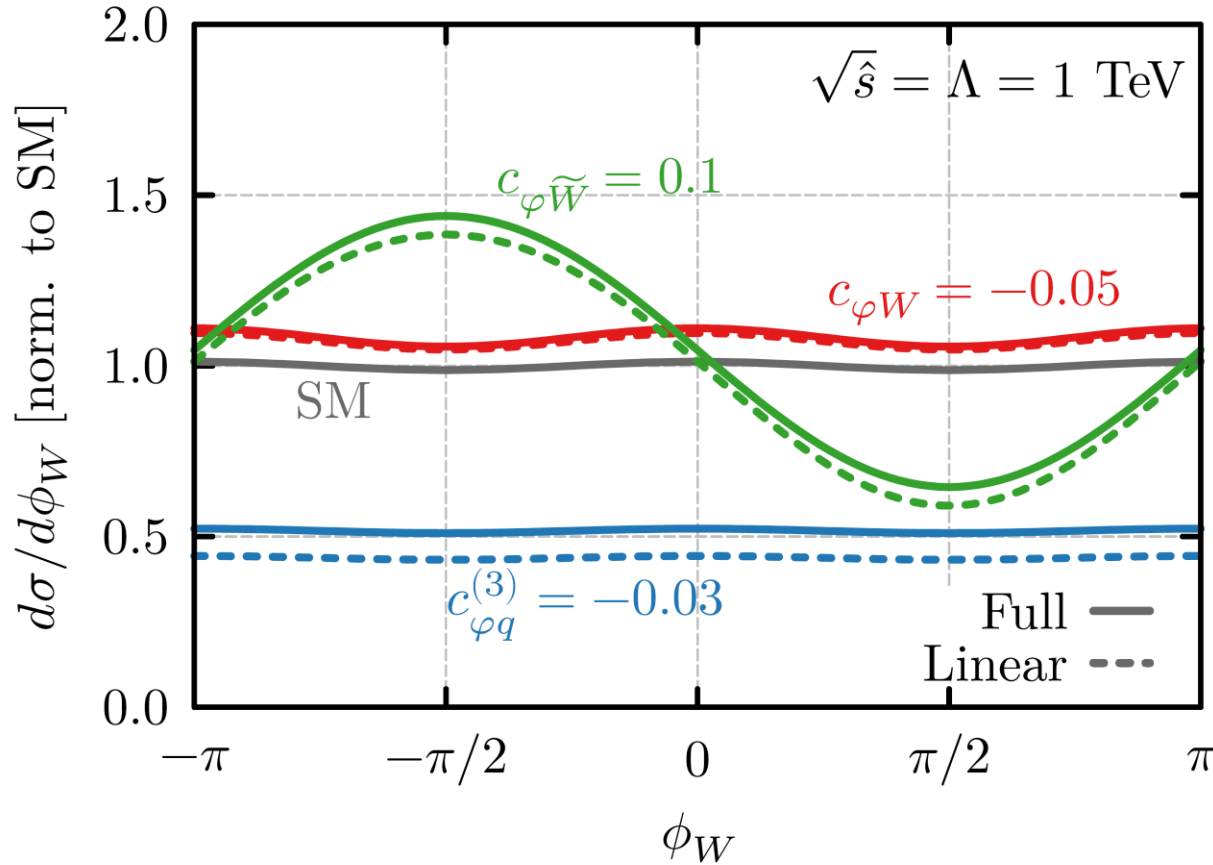




# Wh.

# Interference patterns

With  $\nu$  reconstruction ambiguity



## Differential in $p_T^h$ and $\phi_W$

$$\sigma_{\mathcal{O}_{\phi q}^{(3)}}^{int} \sim \frac{\hat{s}}{\Lambda^2} \quad \nu \text{ reconstruction}$$

$$(\phi_W \rightarrow \pi - \phi_W)$$

$$\sigma_{\mathcal{O}_{\phi W}}^{int} \sim \frac{\sqrt{\hat{s}} M_W}{\Lambda^2} \cos(\phi_W)$$

$$\sigma_{\mathcal{O}_{\phi\tilde{W}}}^{int} \sim \frac{\sqrt{\hat{s}} M_W}{\Lambda^2} \sin(\phi_W)$$

$$p_T^h \in \{200, 400, 600, 800, 1000, \infty\} \text{ GeV}$$

$$\phi_W \in [-\pi, 0], [0, \pi]$$

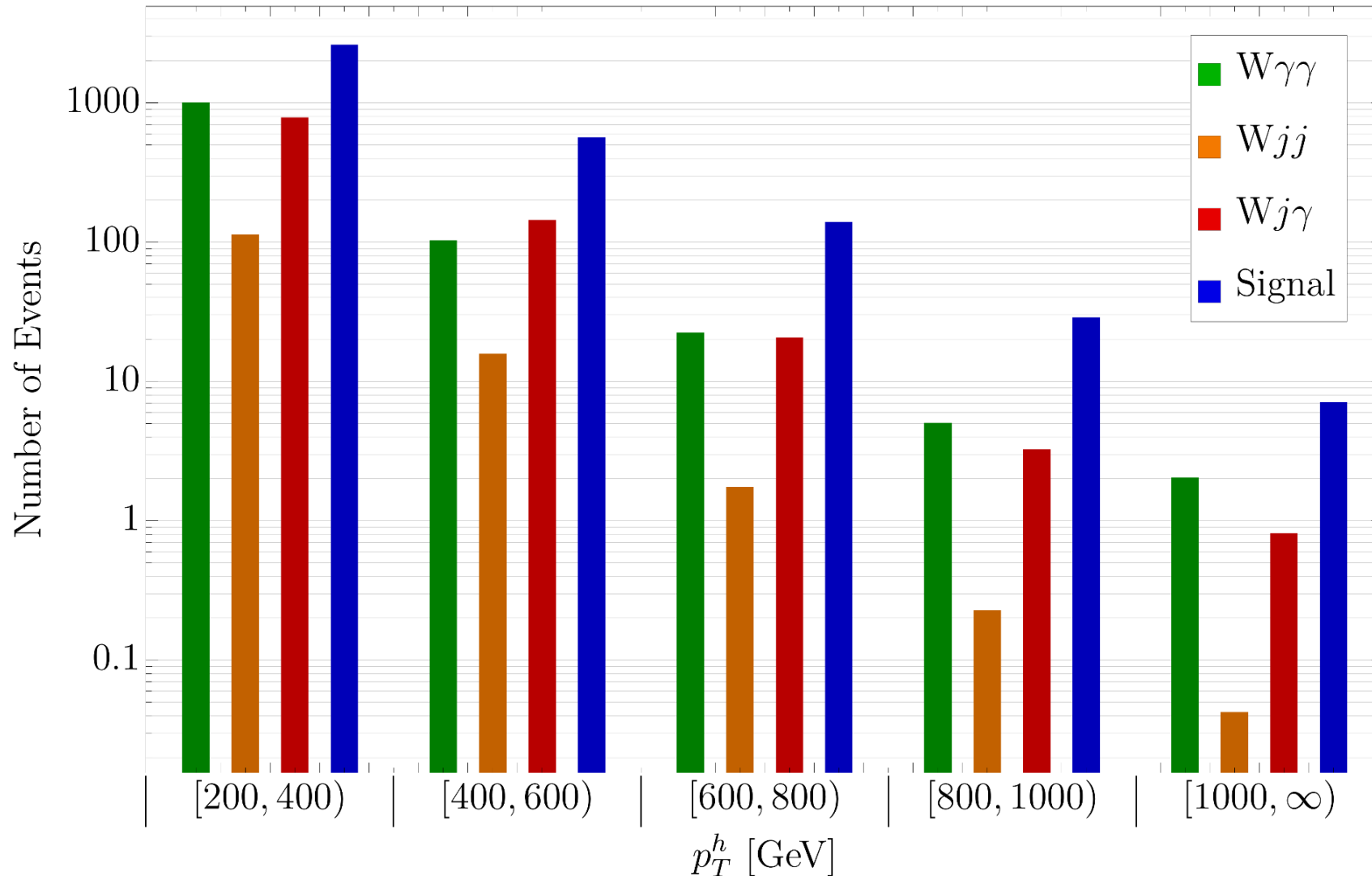
Categories		Variable	(HL-)LHC	FCC-hh
0-lepton	boosted	$p_{T,\min}$ [GeV]	$\{0, 300, 350, \infty\}$	$\{0, 200, 400, 600, 800, \infty\}$
	resolved		$\{0, 160, 200, 250, \infty\}$	$\{0, 200, 400, 600, 800, \infty\}$
1-lepton	boosted	$p_T^h$ [GeV]	$\{0, 175, 250, 300, \infty\}$	$\{0, 200, 400, 600, 800, \infty\}$
	resolved		$\{0, 175, 250, \infty\}$	$\{0, 200, 400, 600, \infty\}$
2-lepton	boosted	$p_{T,\min}$ [GeV]	$\{250, \infty\}$	$\{0, 200, 400, 600, \infty\}$
	resolved		$\{175, 200, \infty\}$	$\{0, 200, 400, 600, \infty\}$

# Wh.

## How big is the background?

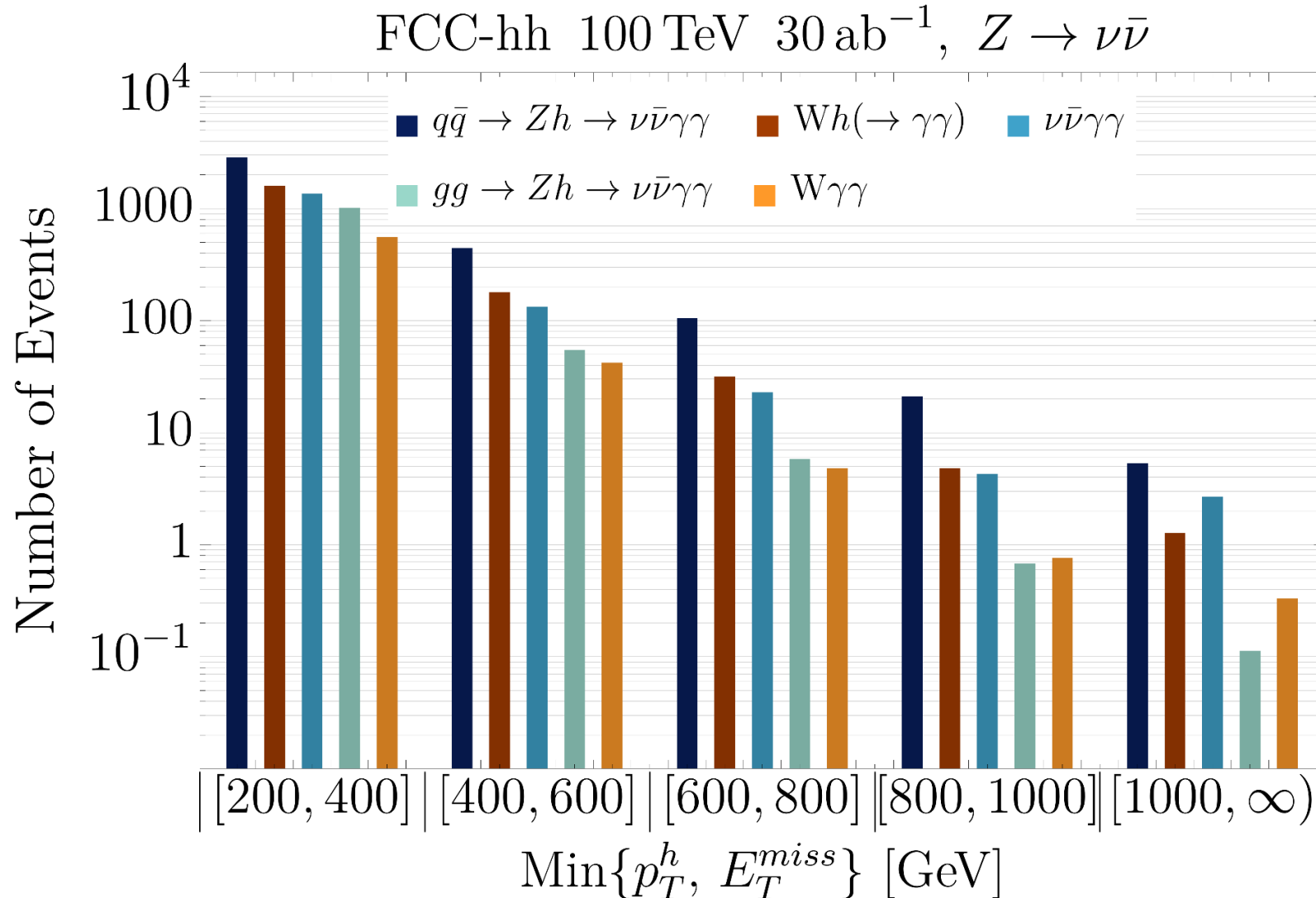
- Events per bin for the relevant processes

FCC-hh 100 TeV 30 ab<sup>-1</sup>



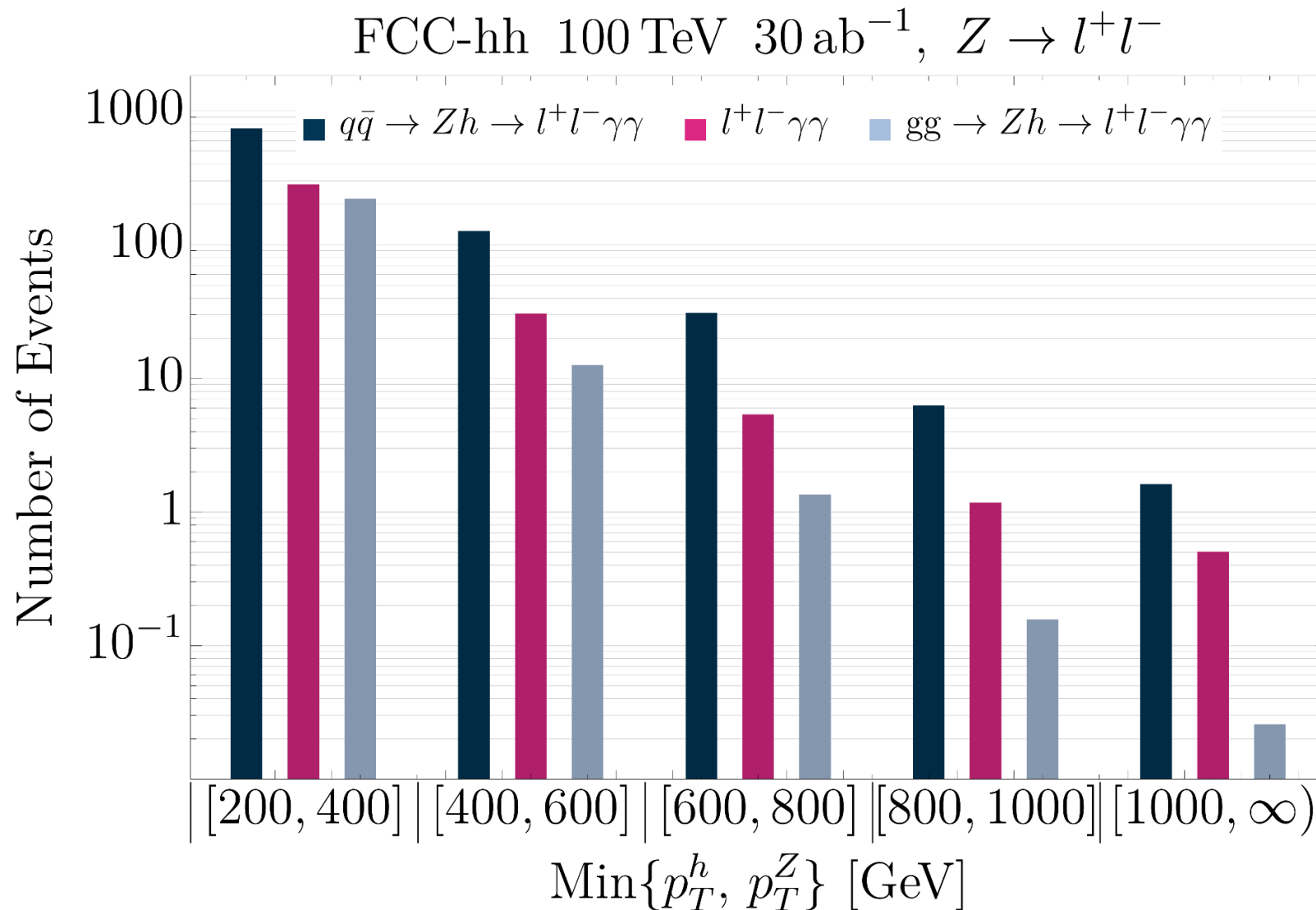
# Signal and background

- Wh is part of the signal because it is affected by  $\theta_{\varphi q}^{(3)}$ .



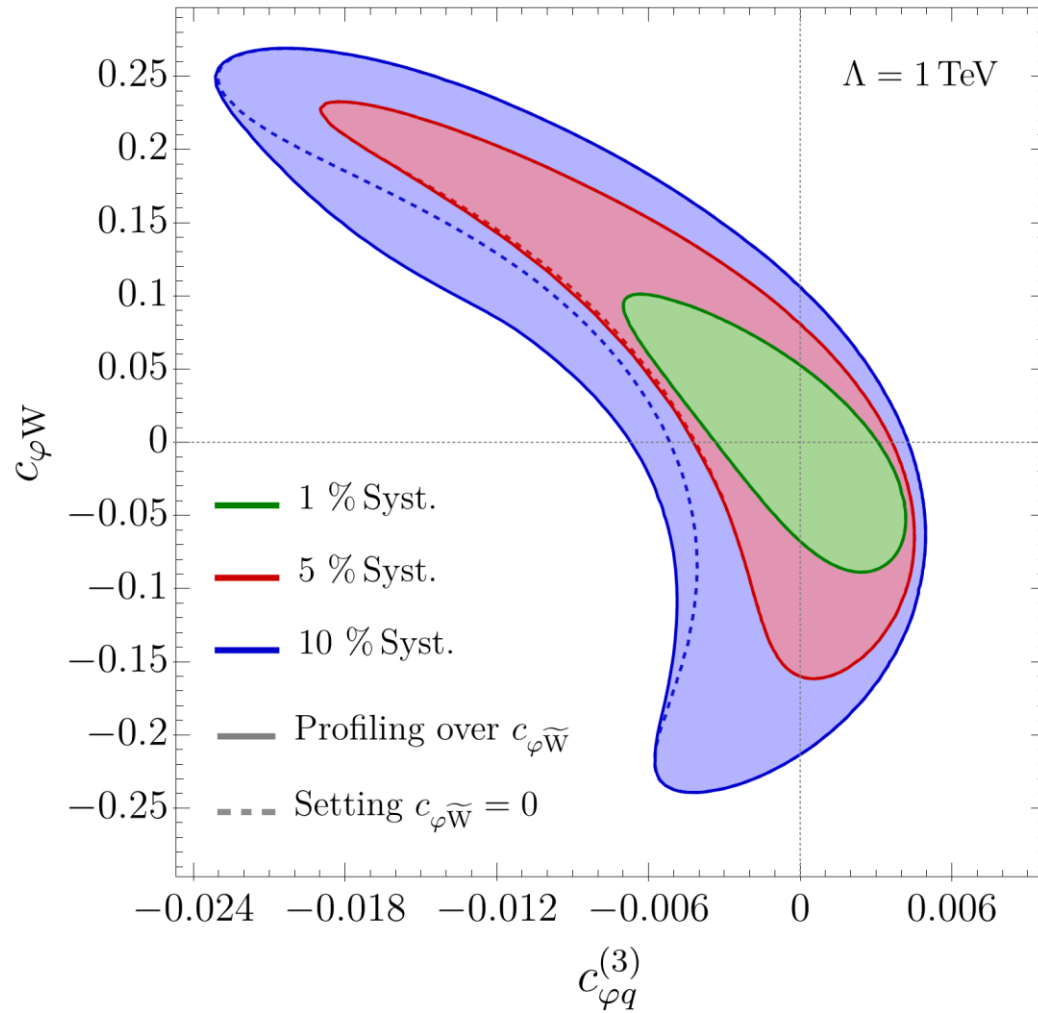
## More results

- Events per bin for the relevant processes in the leptonic channel.

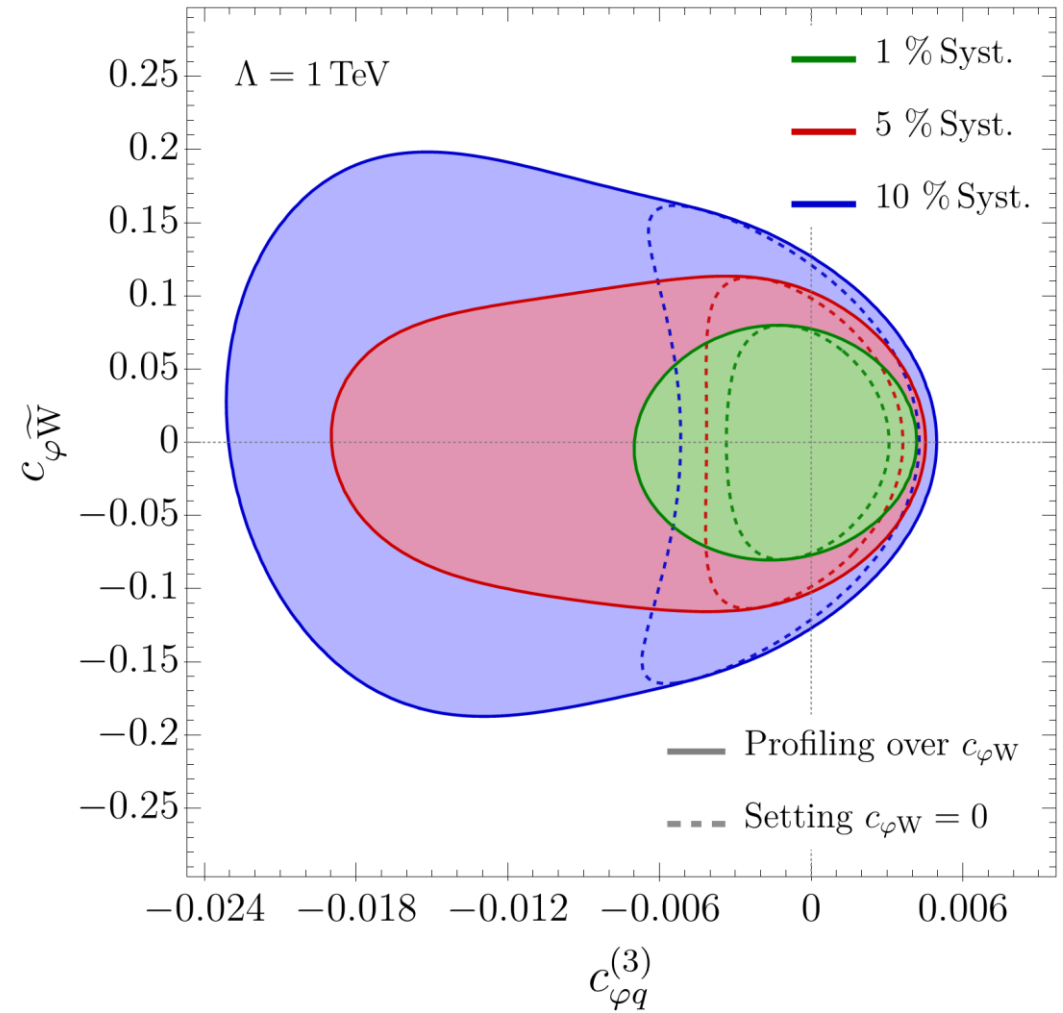


- 95% CL bounds

FCC-hh 100 TeV 30 ab<sup>-1</sup>

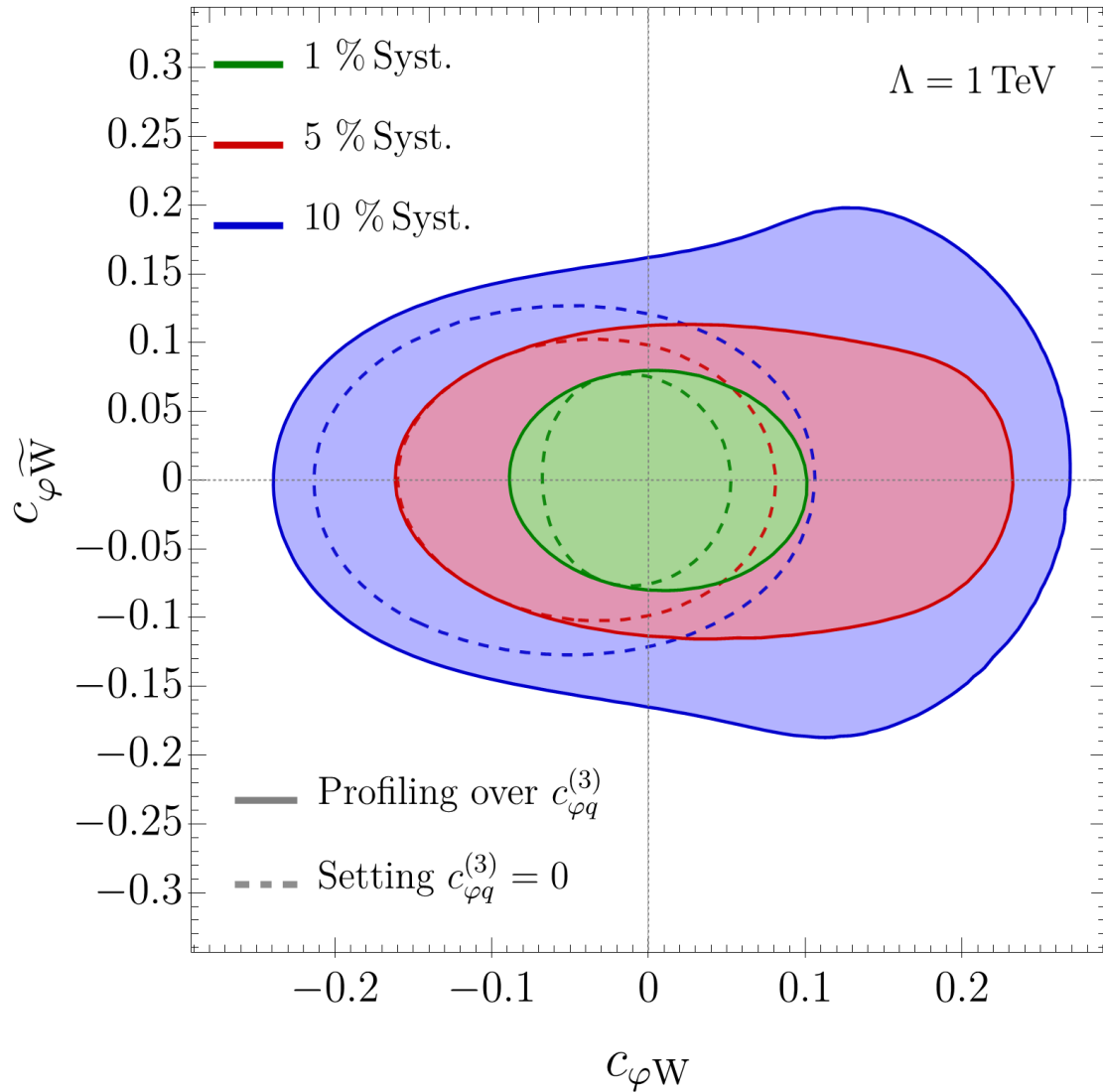


FCC-hh 100 TeV 30 ab<sup>-1</sup>

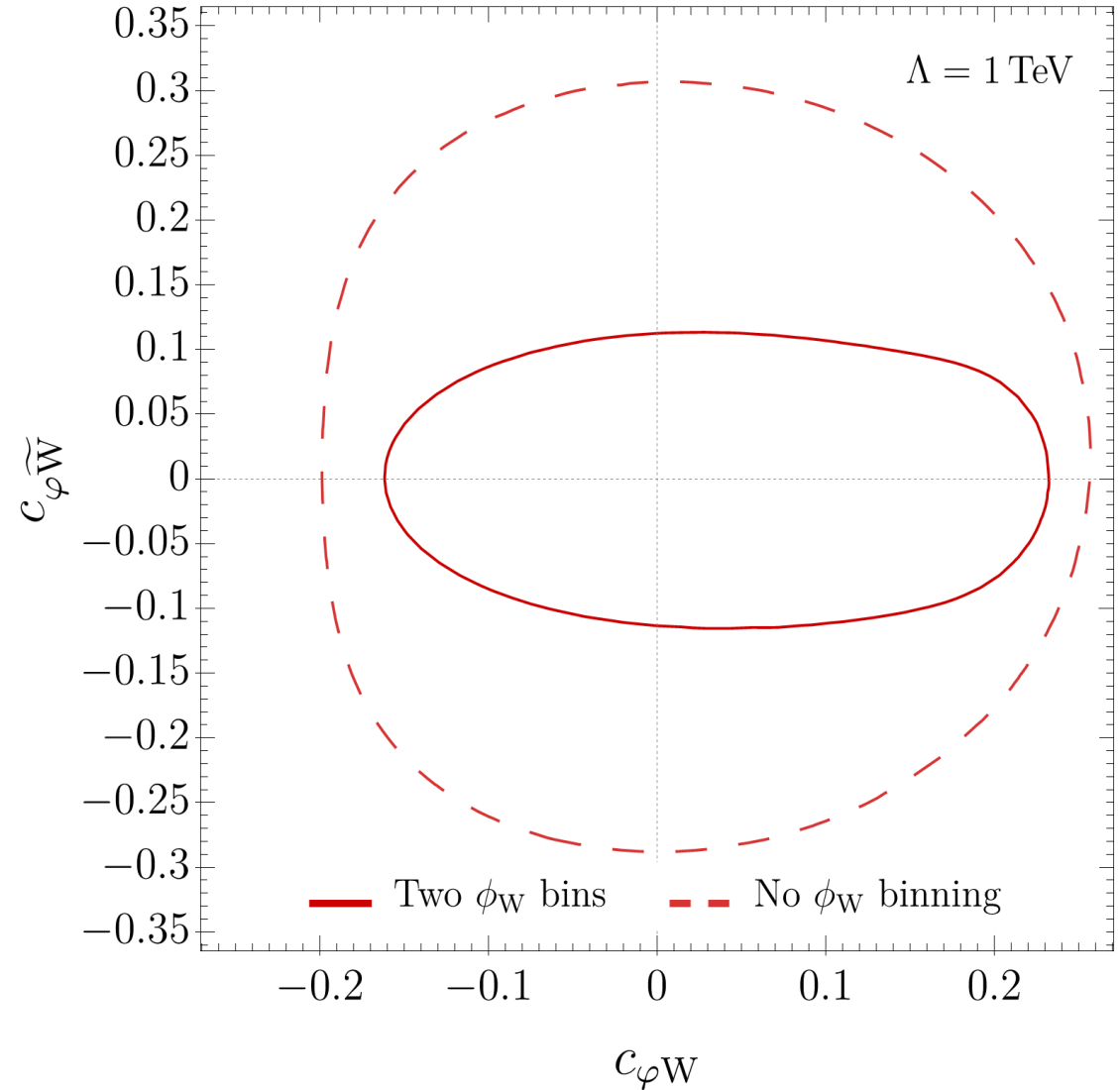


# Wh. 95% C.L. on the bosonic operators

FCC-hh 100 TeV 30 ab<sup>-1</sup>



FCC-hh 100 TeV 30 ab<sup>-1</sup>, 5% Syst.



- 95% CL bounds summary

Coefficient	Profiled Fit		One Operator Fit	
$c_{\varphi q}^{(3)}$	$[-5.1, 3.4] \times 10^{-3}$	1% syst.	$[-2.7, 2.5] \times 10^{-3}$	1% syst.
	$[-11.6, 3.8] \times 10^{-3}$	5% syst.	$[-3.3, 2.9] \times 10^{-3}$	5% syst.
	$[-20.6, 4.1] \times 10^{-3}$	10% syst.	$[-4.0, 3.5] \times 10^{-3}$	10% syst.
$c_{\varphi W}$	$[-7.1, 7.9] \times 10^{-2}$	1% syst.	$[-5.3, 4.3] \times 10^{-2}$	1% syst.
	$[-13.0, 17.5] \times 10^{-2}$	5% syst.	$[-12.1, 6.8] \times 10^{-2}$	5% syst.
	$[-20.0, 25.2] \times 10^{-2}$	10% syst.	$[-18.8, 9.0] \times 10^{-2}$	10% syst.
$c_{\varphi \tilde{W}}$	$[-6.4, 6.4] \times 10^{-2}$	1% syst.	$[-6.1, 6.1] \times 10^{-2}$	1% syst.
	$[-9.0, 8.8] \times 10^{-2}$	5% syst.	$[-8.1, 8.1] \times 10^{-2}$	5% syst.
	$[-13.5, 14.2] \times 10^{-2}$	10% syst.	$[-10.1, 10.1] \times 10^{-2}$	10% syst.

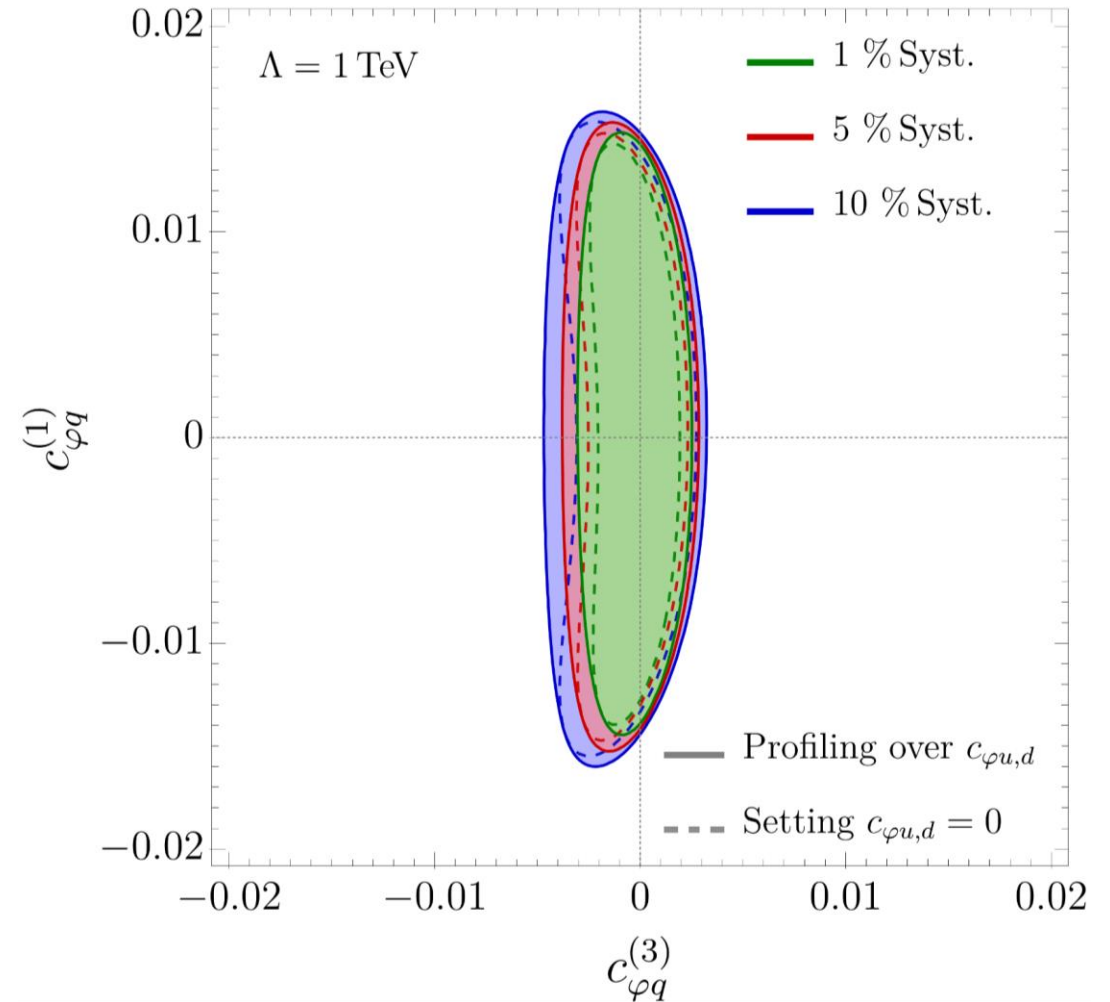
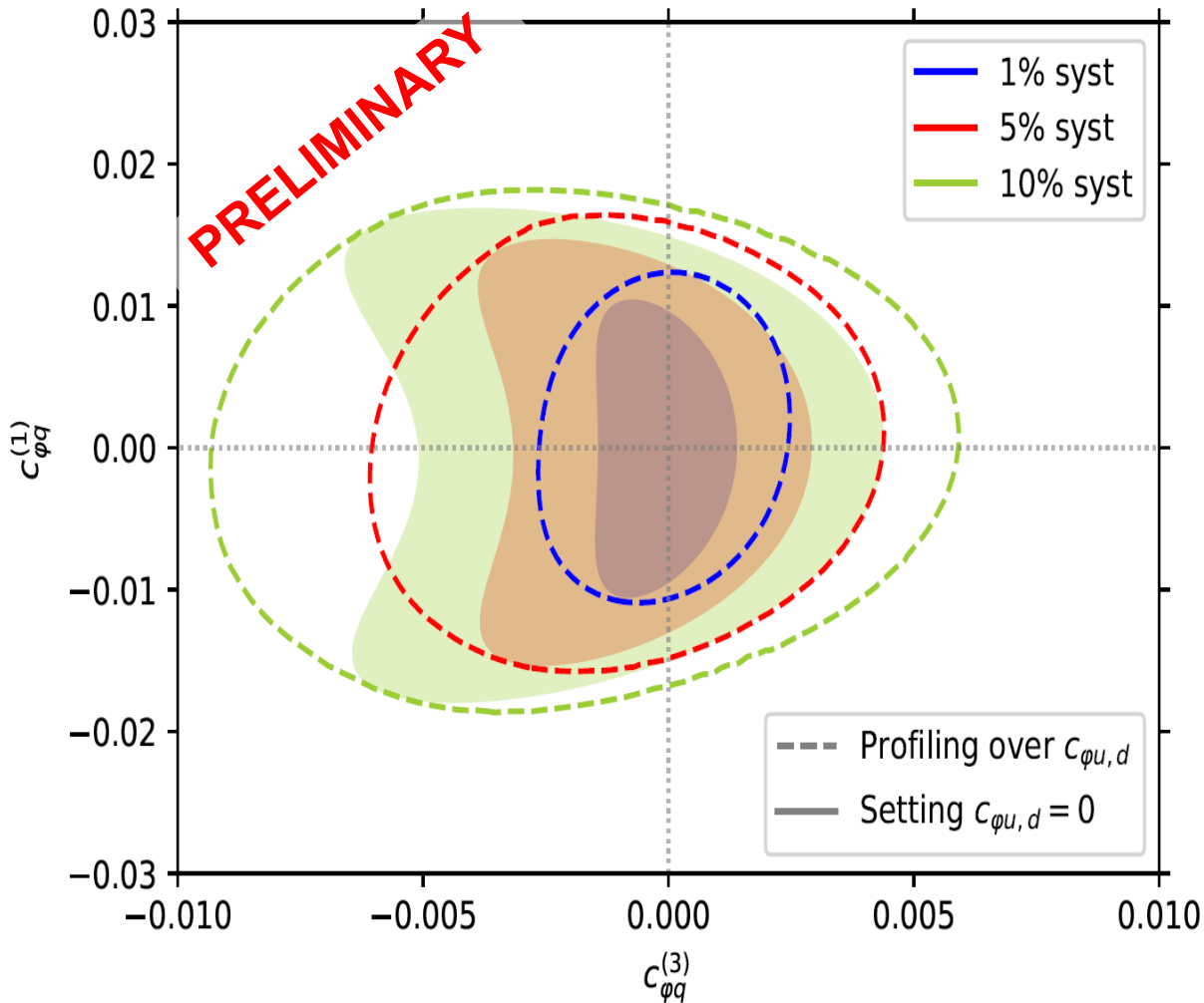


# At FCC-hh, photons or b-quarks?

FCC-hh 100 TeV  $30 \text{ ab}^{-1}$  ( $Zh + Wh$ )

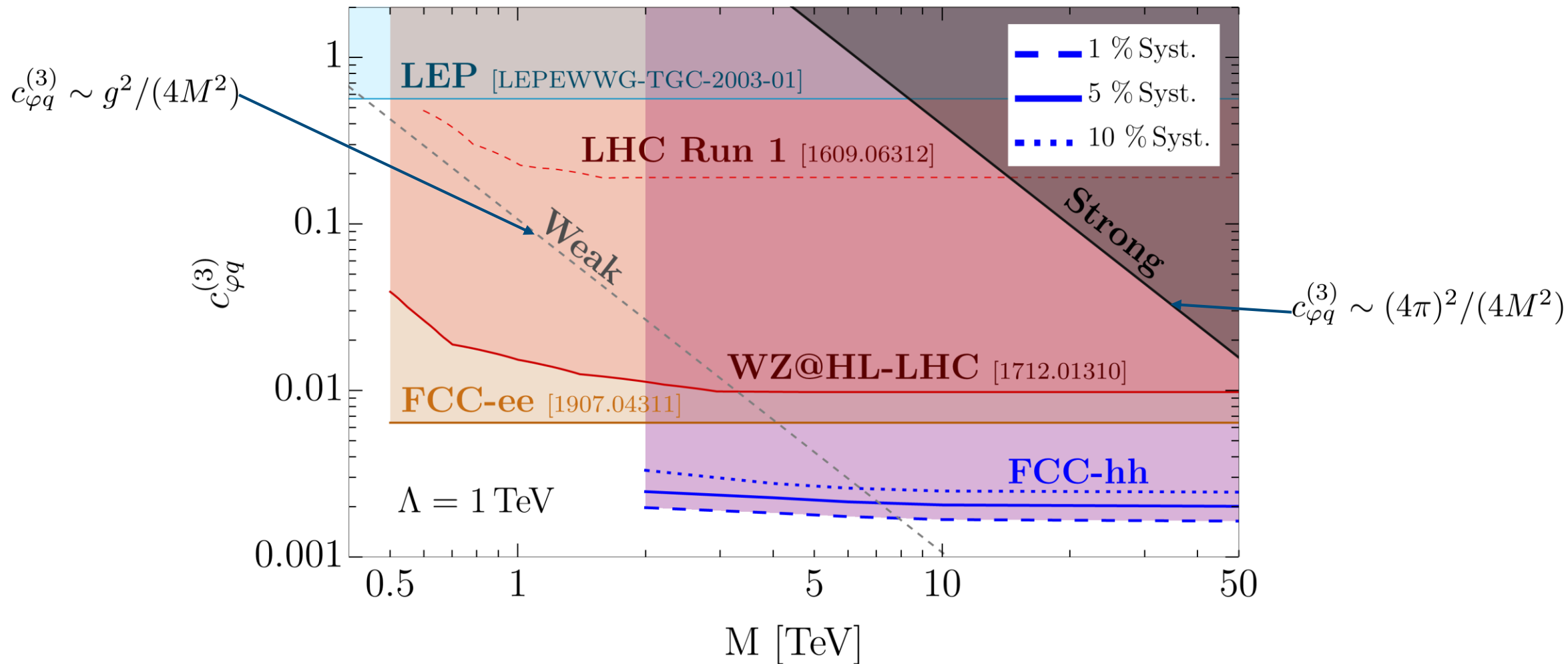
$h \rightarrow b\bar{b}$

$h \rightarrow \gamma\gamma$

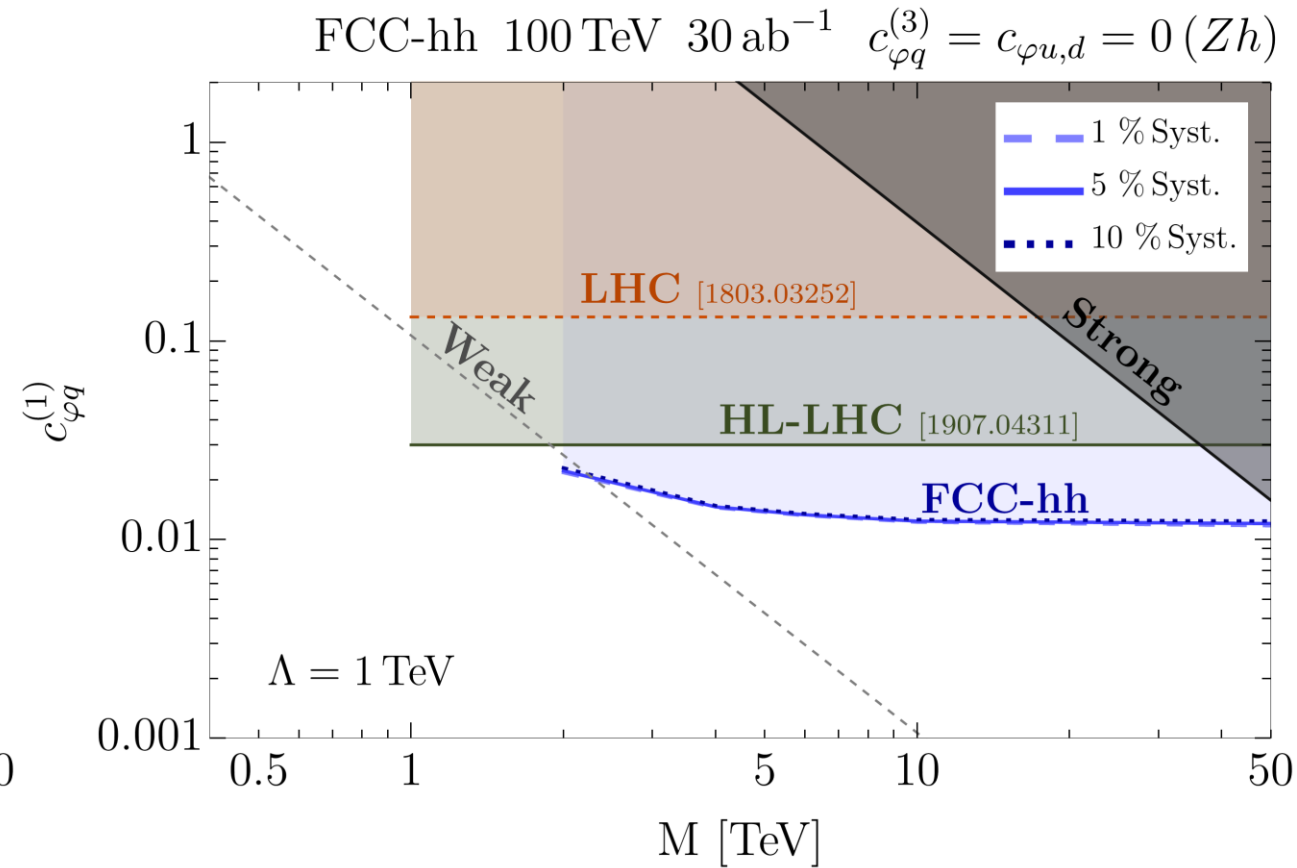
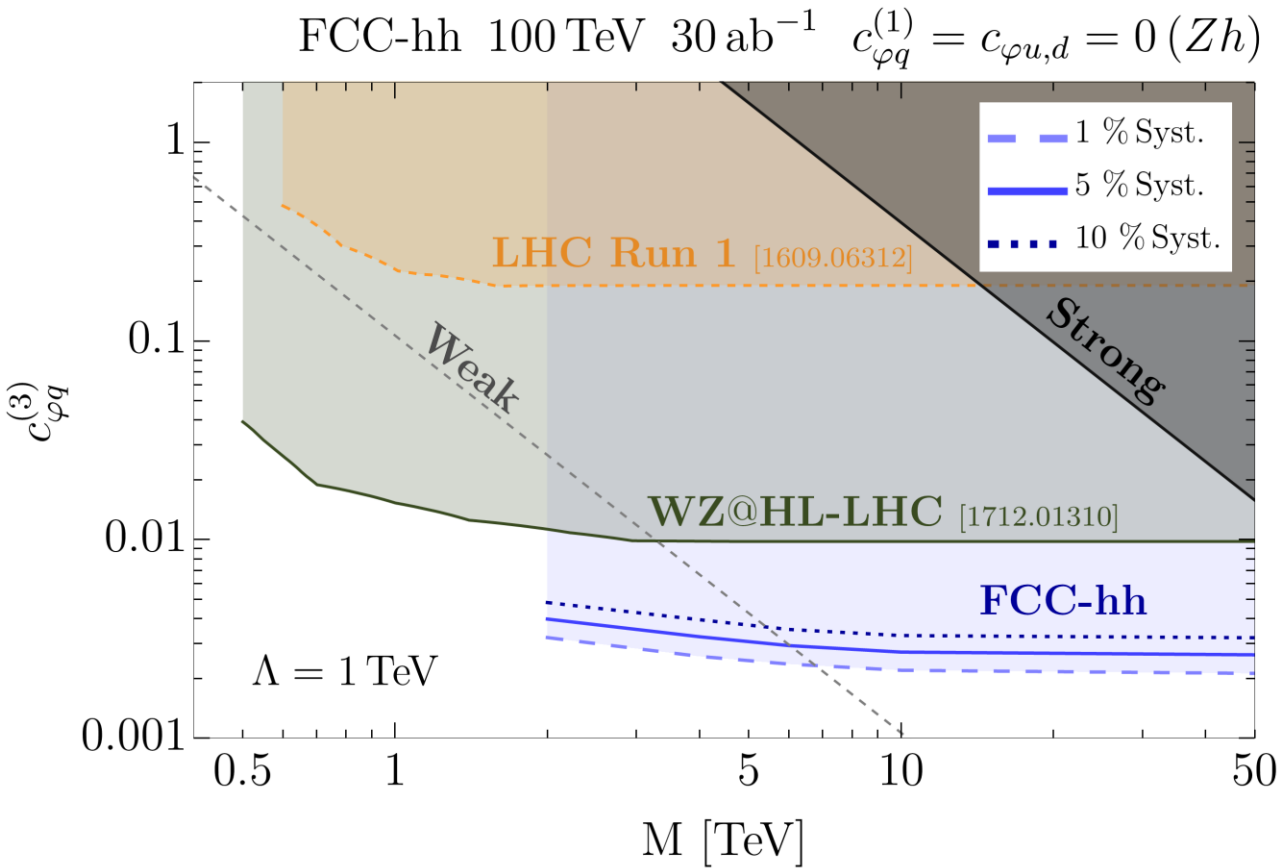


# Bounds for $c_{\varphi q}^{(3)}$

FCC-hh 100 TeV 30  $\text{ab}^{-1}$ , 1-op. fit, ( $Zh + Wh$ )



- 95% CL bounds



- 95% CL bounds summary

Coefficient	Profiled Fit	One Operator Fit
$c_{\varphi q}^{(3)}$	$[-5.2, 3.1] \times 10^{-3}$ 1% syst.	$[-2.1, 2.0] \times 10^{-3}$ 1% syst.
	$[-6.7, 3.3] \times 10^{-3}$ 5% syst.	$[-2.6, 2.4] \times 10^{-3}$ 5% syst.
	$[-8.2, 3.7] \times 10^{-3}$ 10% syst.	$[-3.2, 2.8] \times 10^{-3}$ 10% syst.
$c_{\varphi q}^{(3)}$ (+Wh)	$[-2.5, 2.1] \times 10^{-3}$ 1% syst.	$[-1.6, 1.6] \times 10^{-3}$ 1% syst.
	$[-3.0, 2.4] \times 10^{-3}$ 5% syst.	$[-2.0, 1.9] \times 10^{-3}$ 5% syst.
	$[-3.7, 2.7] \times 10^{-3}$ 10% syst.	$[-2.4, 2.2] \times 10^{-3}$ 10% syst.
$c_{\varphi q}^{(1)}$	$[-1.3, 1.4] \times 10^{-2}$ 1% syst.	$[-1.1, 1.15] \times 10^{-2}$ 1% syst.
	$[-1.5, 1.5] \times 10^{-2}$ 5% syst.	$[-1.1, 1.2] \times 10^{-2}$ 5% syst.
	$[-1.6, 1.5] \times 10^{-2}$ 10% syst.	$[-1.2, 1.2] \times 10^{-2}$ 10% syst.
$c_{\varphi u}$	$[-2.0, 1.6] \times 10^{-2}$ 1% syst.	$[-1.9, 0.89] \times 10^{-2}$ 1% syst.
	$[-2.1, 1.7] \times 10^{-2}$ 5% syst.	$[-2.1, 0.96] \times 10^{-2}$ 5% syst.
	$[-2.2, 1.8] \times 10^{-2}$ 10% syst.	$[-2.2, 1.0] \times 10^{-2}$ 10% syst.
$c_{\varphi d}$	$[-2.1, 2.3] \times 10^{-2}$ 1% syst.	$[-1.4, 2.2] \times 10^{-2}$ 1% syst.
	$[-2.2, 2.4] \times 10^{-2}$ 5% syst.	$[-1.5, 2.2] \times 10^{-2}$ 5% syst.
	$[-2.3, 2.5] \times 10^{-2}$ 10% syst.	$[-1.5, 2.2] \times 10^{-2}$ 10% syst.

# Vh.

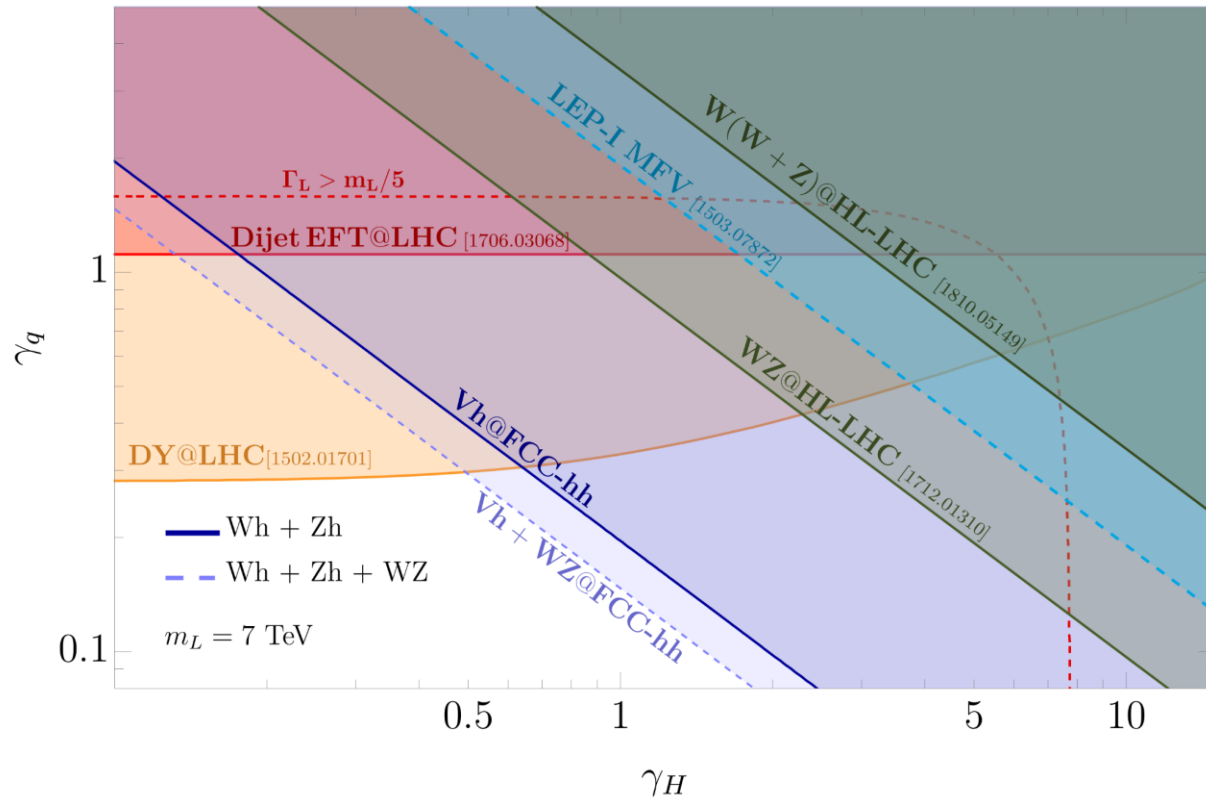
# UV models: Spin-1 triplets

$$L_\mu \sim (1, 3)_0$$

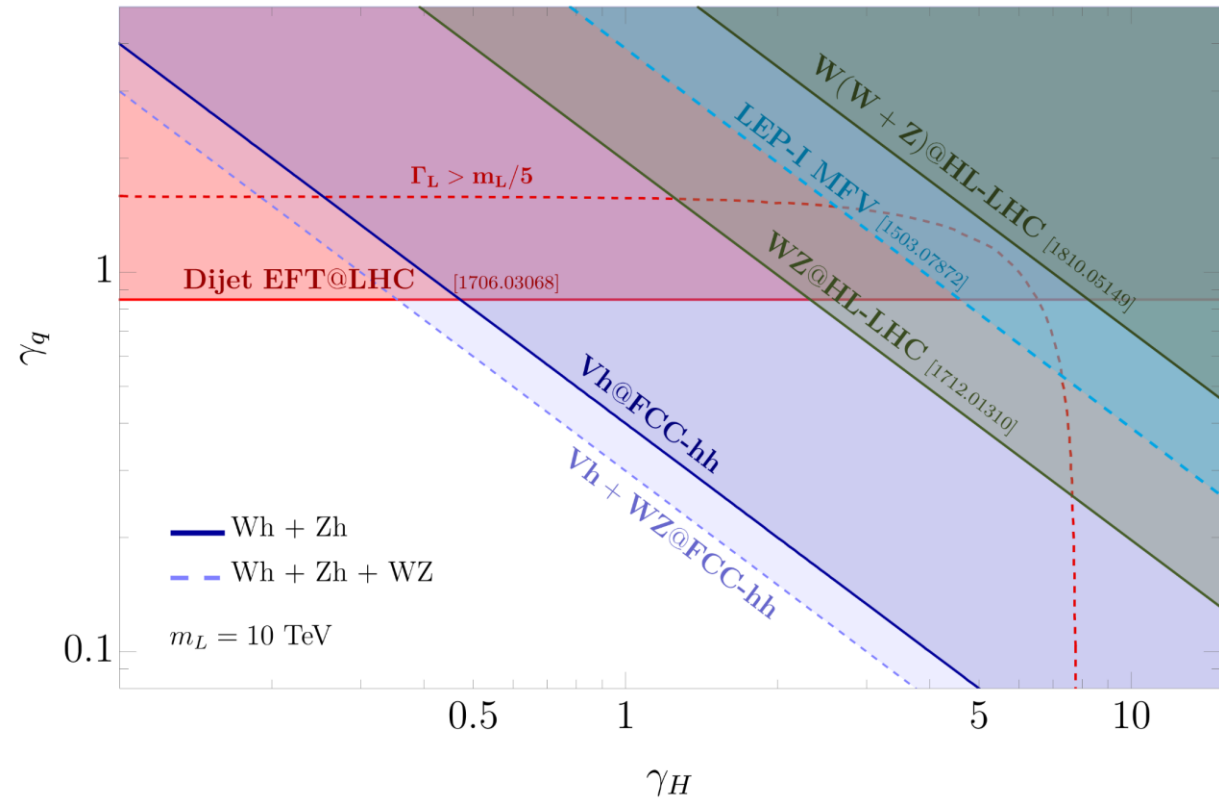
$$\mathcal{L}_{BSM} = \frac{1}{4} F_{L,\mu\nu}^a F_L^{a,\mu\nu} + \frac{m_L^2}{2} L_\mu L^\mu + \gamma_H L_\mu^a \frac{i}{2} H^\dagger \sigma^a \overleftrightarrow{D}^\mu H + \sum_f \gamma_f L_\mu^a \bar{f} \gamma^\mu \sigma^a f$$

$$c_{\varphi q}^{(3)} = -\frac{\gamma_H \gamma_q}{2m_L^2}$$

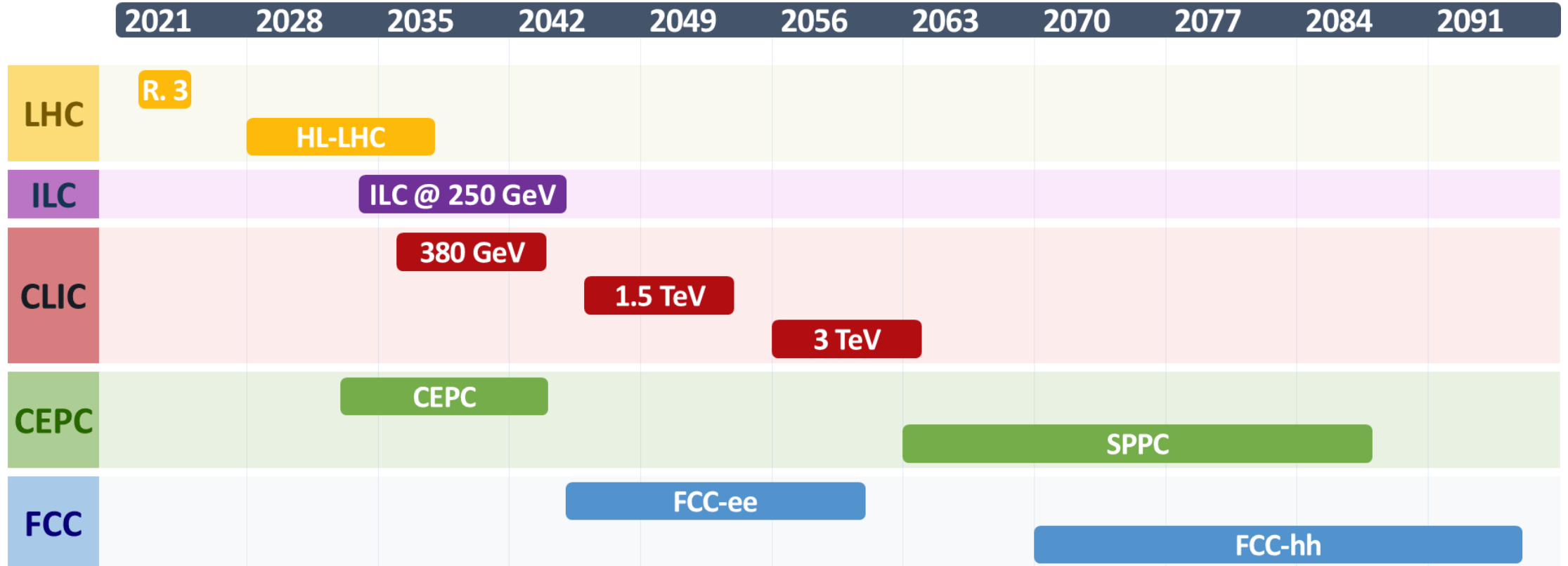
FCC-hh 100 TeV 30 ab<sup>-1</sup> 5% Syst.



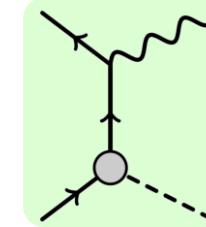
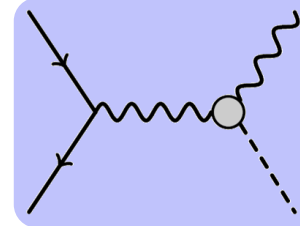
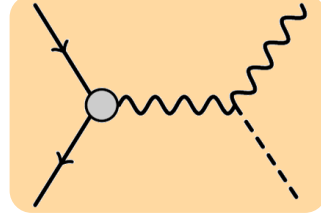
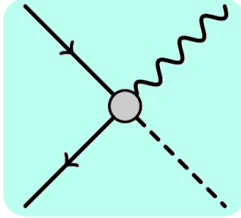
FCC-hh 100 TeV 30 ab<sup>-1</sup> 5% Syst.



# Future experiments timeline



## Dimension-6 operators in Vh



$$\mathcal{O}_{\varphi D} = (H^\dagger D^\mu H)^* (H^\dagger D_\mu H)$$

$$\mathcal{O}_{\varphi W} = H^\dagger H W^{a, \mu\nu} W_{\mu\nu}^a$$

$$\mathcal{O}_{\varphi \tilde{W}} = H^\dagger H W^{a, \mu\nu} \tilde{W}_{\mu\nu}^a$$

$$\mathcal{O}_{\varphi B} = H^\dagger H B^{\mu\nu} B_{\mu\nu}$$

$$\mathcal{O}_{\varphi \tilde{B}} = H^\dagger H B^{\mu\nu} \tilde{B}_{\mu\nu}$$

$$\mathcal{O}_{\varphi WB} = H^\dagger \sigma^a H B^{\mu\nu} W_{\mu\nu}^a$$

$$\mathcal{O}_{\varphi \tilde{W} B} = H^\dagger \sigma^a H B^{\mu\nu} \tilde{W}_{\mu\nu}^a$$

$$\mathcal{O}_{uW} = (\bar{q}_L \sigma^{\mu\nu} u_R) \tau^I \tilde{H} W_{\mu\nu}^I$$

$$\mathcal{O}_{dW} = (\bar{q}_L \sigma^{\mu\nu} d_R) \tau^I H W_{\mu\nu}^I$$

$$\mathcal{O}_{uB} = (\bar{q}_L \sigma^{\mu\nu} u_R) \tilde{H} B_{\mu\nu}$$

$$\mathcal{O}_{dB} = (\bar{q}_L \sigma^{\mu\nu} d_R) H B_{\mu\nu}$$

$$\mathcal{O}_{\varphi ud} = (\bar{u}_R \gamma^\mu d_R) \left( i H^\dagger \overleftrightarrow{D}_\mu H \right)$$

$$\mathcal{O}_{\varphi q}^{(3)} = (\bar{Q}_L \sigma^a \gamma^\mu Q_L) \left( i H^\dagger \sigma^a \overleftrightarrow{D}_\mu H \right)$$

$$\mathcal{O}_{\varphi q}^{(1)} = (\bar{Q}_L \gamma^\mu Q_L) \left( i H^\dagger \overleftrightarrow{D}_\mu H \right)$$

$$\mathcal{O}_{\varphi u} = (\bar{u}_R \gamma^\mu u_R) \left( i H^\dagger \overleftrightarrow{D}_\mu H \right)$$

$$\mathcal{O}_{\varphi d} = (\bar{d}_R \gamma^\mu d_R) \left( i H^\dagger \overleftrightarrow{D}_\mu H \right)$$

$$\mathcal{O}_{u\varphi} = H^\dagger H \left( \bar{q}_L \tilde{H} u_R \right)$$

$$\mathcal{O}_{d\varphi} = H^\dagger H \left( \bar{q}_L H d_R \right)$$

— MFV suppressed  
 — Sub-leading energy growth  
 — No interference with SM for massless quarks