GNATIOS D = 4 ?SUSY? SUPERSTRÍNGS ?

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We met in 1979 at Cargèse Summer School Since then 34 papers together & we are most frequent collaborators Topics range from topological string amplitudes 750 GeV diphoton excess to

But most important ... we are best friends





Going back to 1979: 5 At that time ... we worked on QCD & nobody way agking D = ? be cause you just open your eyes and see D = 4... and they are still there ... SUSY & SUPERSTRINGS were rather exoteric branches of mathematical physics because there was no physical (experimental) region to consider them seriously ... and that didn't change, either

But in the meantime, D became a parameter - as the central charge of world-sheet CFT arguably, D=10 is the preferred value Over the last 35 years, there has been extensive discussion on the shape and size of "extra" dimensions. In particular Ignation proneered the concept of large extra dimensions

But very few ideas why there are noncompact D = 4?

1. Brandenberger - Vafa string gas cosmology D = 1 / D = 42. T& Veneziano $D = 10 \rightarrow D = 4$ compactification driven by infrared divergences What's special about D=4?

4 is max D with IR divergences D=4: O probability of 1 & (brenstrahlung) D>4: finite probability of 1 & soft theorems What about SVSY? Here again, exact SUSY is preferred by strings D=4 SUSY ~ world-sheet SUSY Summary: D74 & SUSY are natural in string theory

Celestial holography:

hologram of D=4 world makes D = 2

D = 4 & SUSY unnatural

If invented in 1979 nobody would worry about D≠4 and SUSY ... but with holography, we had to wait until 1990's ('t Hooft Maldacena,...)

The idea of holography: Spacetimes boundary 4D physics (simpler) physics on lower-dimensional "boundary" What is the boundary of flat (Minkowski) or asymptotically flat (our Universe) spacetime?

You can see it on a clear night



CELESTIAL SPHERE CS_2

You can get there by following a light ray back to the past at "null infinity"

[]] = point on CS2 Direction of light ray = point 2 on complex plane $P(\infty)$ ZEC P(z)Stereographic projection $(S_{2} \rightarrow$ $p^2 = 0 \implies p^{\mu} = \omega q^{\mu}$ $q^{\mu} = \frac{1}{2} \left(1 + |z|^2, z + \bar{z}, -i(z - \bar{z}), 1 - |z|^2 \right)$ specifies direction of p



How does it look like from V~C Milleruin Falcon?

 $\mathbf{N} P(\infty)$ $\alpha = P(A)$ Stars "pile up" towards North Pole Z = 00 = dragged away from Jorith Pole Z=0 ° M Lorentz boost rapidity $7 \rightarrow \mathcal{C}$ = 2D dilation (scale)

In general: Lorentz transformation SO(1,3) = 2 D conformal symmetry $SL(2, \mathbb{C})$ $Z \rightarrow \frac{a\chi + b}{C\chi + d} det \begin{pmatrix} a & b \\ c & d \end{pmatrix} = 1$

Physics in D=4 asympt-flat gracetine encoded in 2D CFT on CS_2 \rightarrow use power of \mathbb{C}

 $(D=4):(D\neq4)$ 1:0

14 Suppose that we want to describe 4D scattering processes in terms of amplitudes on (S2: out out , in the

need to encode info not only about dividions Z but also about energies w,

in SL(2,C) invariant way?

15 Solution : replace plane waves characterized by p(w, z)by conformal wave packets Mellin dual characterized by A. T. 1 $\begin{array}{l} (\text{on formal wave - packets} \\ (x;p) = e^{\pm ipx} \\ \longrightarrow \\ (x;z,\bar{z}) = \int d\omega \ \omega^{-i} e^{\pm i\omega qx} \end{array}$ live $(xq)^{\Delta}$ $h=h=\frac{\Delta}{2}$ $h=h=\frac{\Delta}{2}$ $h=h=\frac{\Delta}{2}$ $h=h=\frac{\Delta}{2}$ $h=h=\frac{\Delta}{2}$ $h=h=\frac{\Delta}{2}$ normalizable only if Pastenski & Shao

j,6 We want to understand scattering amplitudes as two-dimensional conformal correlators $\langle in | out \rangle \rightarrow \langle O_1(x_1), \dots, O_N(x_N) \rangle$ Amplitudes: { in | out > A gauge byons gravitory m = 0 YM Einstein's GR $\langle p_{i},h_{i},a_{i},\ldots,p_{n},h_{n},a_{n}\rangle \equiv \mathcal{A}(p_{i},h_{i},a)$ $= (2\pi)^{4} i \delta^{(4)} (\Xi_{\text{pout}} - \Sigma_{\text{pui}}) \mathcal{M}$ Feynman, recursions, scattering egs, ...

17 $\mathcal{A}(z,\omega,...) \rightarrow \mathcal{A}(\Delta,z,...)$ $\widetilde{\mathcal{A}} = \int d\omega_{n} \omega^{n} \delta^{(4)}(\omega_{n}q_{1} + \omega_{n}q_{n}) \mathcal{M}(\omega_{i}, z, \overline{z}_{i})$ $\langle \mathcal{O}_{\Delta_1}(z_1), \ldots, \mathcal{O}_{\Delta_N}(z_N) \rangle \geq \mathcal{A}_{\Delta_1, z_1, \ldots, \Delta_N, z_N}$ correlation functions of conformal wave packets (primary fields) CELESTIAL AMPLITUDES E AMPLITUDES IN CONFORMAL BASIS In 2013, Andy Strominger showed that such correlators are "hiding" a symmetry much larger than the original Lorentz symmetry...

 $P(\infty)$ $\alpha = P(A)$ all possible (smooth) deformations \equiv diffeomorphisms of CS_2 $Z \longrightarrow f(Z)$ 2 D "local" unformal transformations generated by infinite-dimensional Virajoro algebra = BMS superrotations

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The goal is to construct

Clestial Conformal Field Theory in 2D (CCFT)

that describes 4D

elementary particle physics,

hopefully beyond the standard model

(gauge-gravity unification?)

By now a famous example of 4D - 2D dictionary: Soft theorems (D=4)simplest case: scalar QED

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 $\begin{array}{c} & \ddots & \\ & \ddots & \\ & \ddots & \\ & \ddots & \\ &$ ZK Pitk $\left(\sum_{i=1}^{N} Q_{i} \frac{EP_{i}}{kP_{i}}\right) \mathcal{M}(P_{i}, P_{i})$ $P_i \xrightarrow{k \to 0}$ leading singularity ~ 1/w K. universal ~ 1 w soft factor W



 $\widetilde{\mathcal{A}}(Z, \Delta \rightarrow 1; Z_1, \Delta_1, \dots, Z_N, \Delta_N)$ photon charged scalars



 \rightarrow soft photon = "Goldstone boyon" of asymptotic gauge symmetry $(D=4): (D \neq 4)$

2 (

What about SVSY?

{Q, Q}~P

 $D = 4 : P_{\mu} \sim \omega \left(1 + z\overline{z} , z + \overline{z} , -i(z - \overline{z}), 1 - z\overline{z} \right)$ (super)translations are non-holomorphic

NO 2D SUSY in Super BMS_D=4 (NO SUSY):(SUSY) 1:0



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Still, long way towards CCFT

Many things to be done... next workshop starts June 21 at CMSA/BHI/Harvard

Thank you, Ignatios for bling a wonderful friend & collaborator

and for inspiring all of us with your ground-breaking work.

Happy Birthday!

