



PLANCK 2022

Towards excluding
a Light Z' explanation
of $b \rightarrow s\ell^+\ell^-$ anomalies

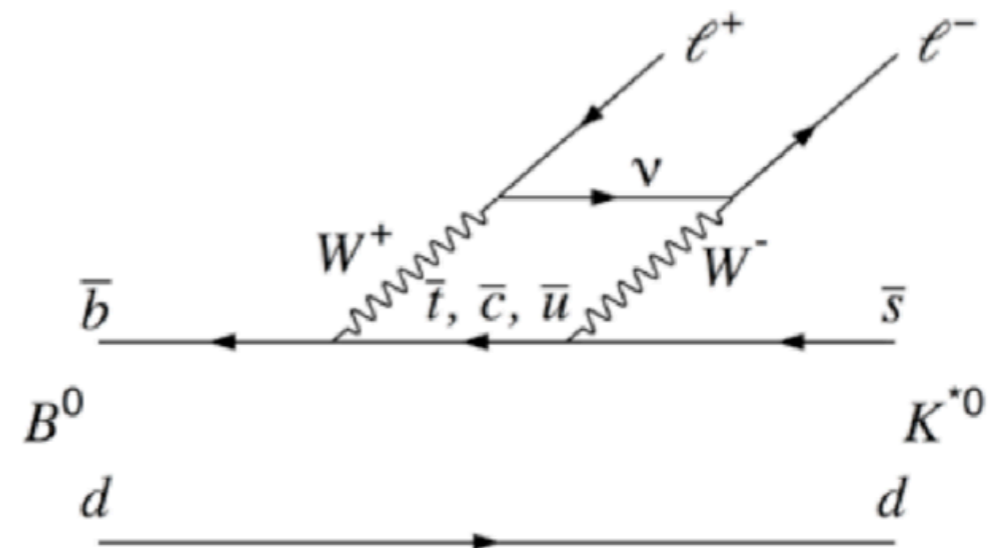
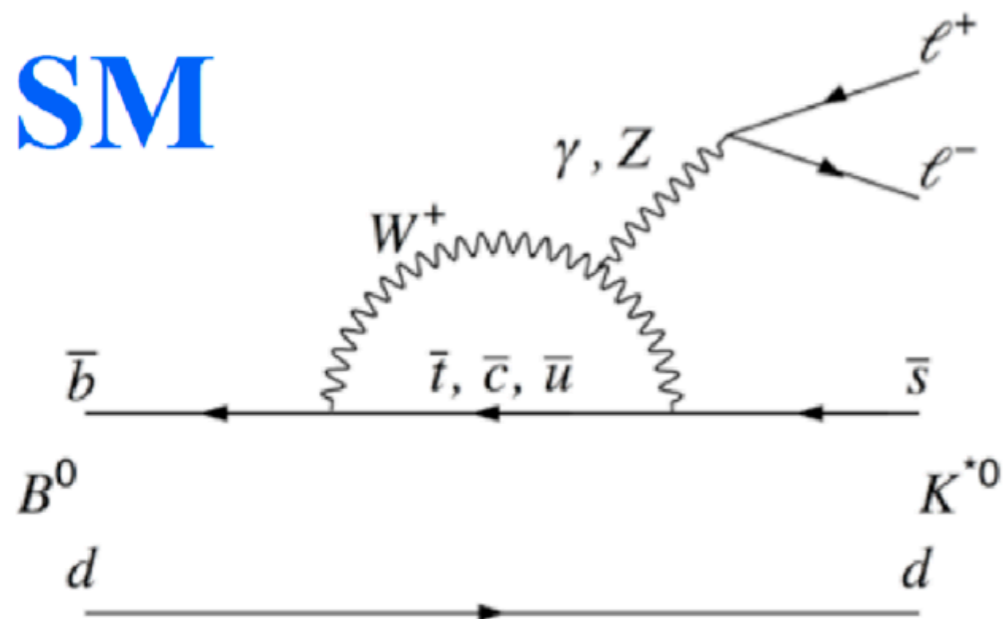


Neutral Current B Decays

Neutral Current B Decays

Many observables with the underlying process $b \rightarrow s \ell^+ \ell^-$ exhibit deviations from SM expectations.

SM



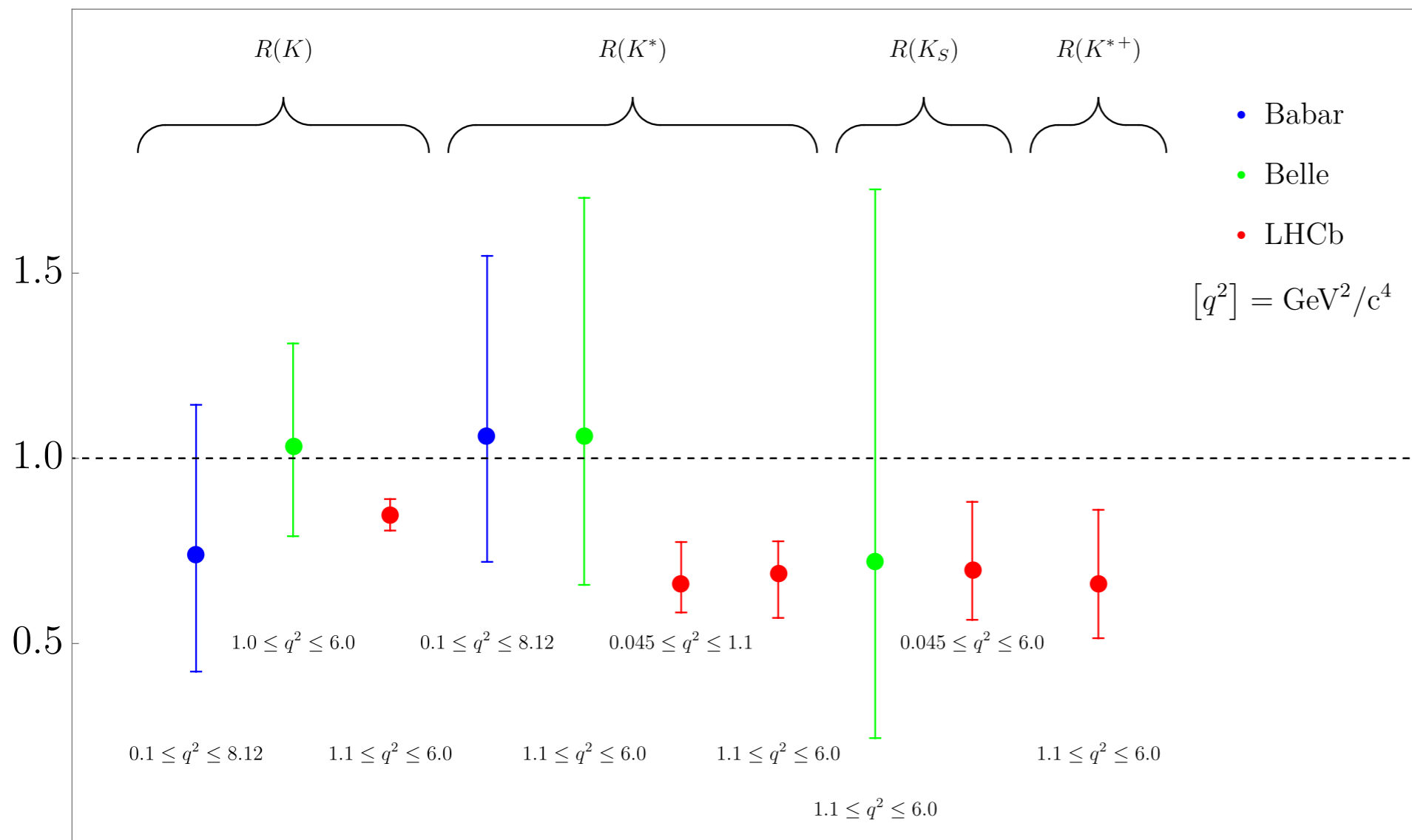
Due to their suppression in the SM, they have a high sensitivity to potential NP contributions.

Neutral Current B Decays



$$R(A) = \frac{\int \frac{d\mathcal{B}}{dq^2}(B \rightarrow A\mu^+\mu^-)dq^2}{\int \frac{d\mathcal{B}}{dq^2}(B \rightarrow Ae^+e^-)dq^2}$$

Huge experimental effort from LHCb and Belle



There is a coherent pattern of deviations from the SM

Other Observables show coherent discrepancies with the SM

$B_s \rightarrow \mu^+\mu^-$ theoretically very clean but chirality suppress

Angular Observables: P'_5, \dots theoretically less clean

Neutral Current B Decays

To perform a global fit to all data we work within the model-independent approach of the effective Hamiltonian:

$$H_{\text{eff}} = -\frac{4G_F}{\sqrt{2}} V_{tb} V_{ts}^* \sum_i C_i O_i$$

$$O_9 = \frac{e}{16\pi^2} (\bar{s} \gamma_\mu P_L b) (\bar{\ell} \gamma^\mu \ell),$$

$$O_{10} = \frac{e}{16\pi^2} (\bar{s} \gamma_\mu P_L b) (\bar{\ell} \gamma^\mu \gamma_5 \ell),$$

$$O_S = \frac{e}{16\pi^2} (\bar{s} P_R b) (\bar{\ell} \ell),$$

$$O_P = \frac{e}{16\pi^2} (\bar{s} P_R b) (\bar{\ell} \gamma_5 \ell),$$

$$O'_9 = \frac{e}{16\pi^2} (\bar{s} \gamma_\mu P_R b) (\bar{\ell} \gamma^\mu \ell),$$

$$O'_{10} = \frac{e}{16\pi^2} (\bar{s} \gamma_\mu P_R b) (\bar{\ell} \gamma^\mu \gamma_5 \ell),$$

$$O'_S = \frac{e}{16\pi^2} (\bar{s} P_L b) (\bar{\ell} \ell),$$

$$O'_P = \frac{e}{16\pi^2} (\bar{s} P_L b) (\bar{\ell} \gamma_5 \ell).$$

This approach works also for a light Z' : $C_i(q^2)$!

Good Fit requires effects in the muon channel: $C_9^\mu, C_9^\mu = -C_{10}^\mu, \dots$

[2103.13370](#)

[2103.12738](#)

[2104.08921](#)

Light Z'

<https://inspirehep.net/literature/2039365>

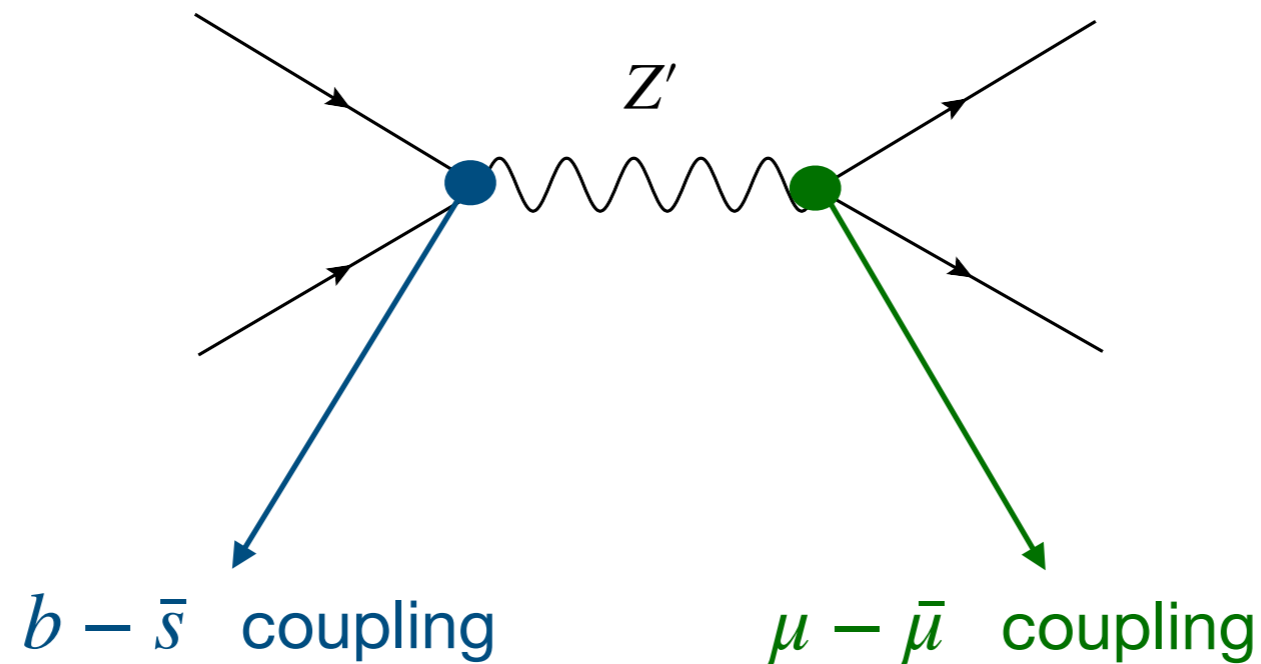
W. Altmannshofer, A. Crivellin, C.A.M., G. Inguglia, P. Feichtinger, J.M. Camalich

Light Z'

$$m_{Z'} \lesssim 6 \text{ GeV}$$

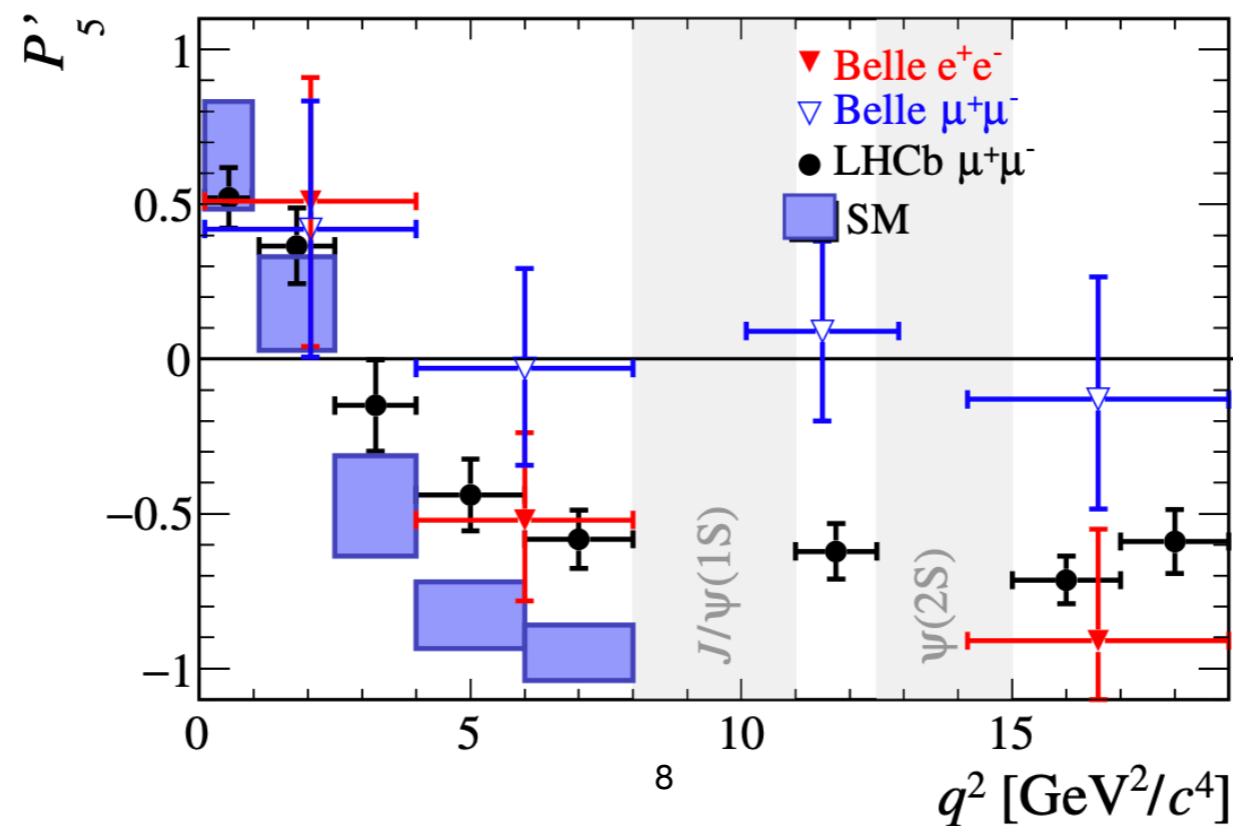
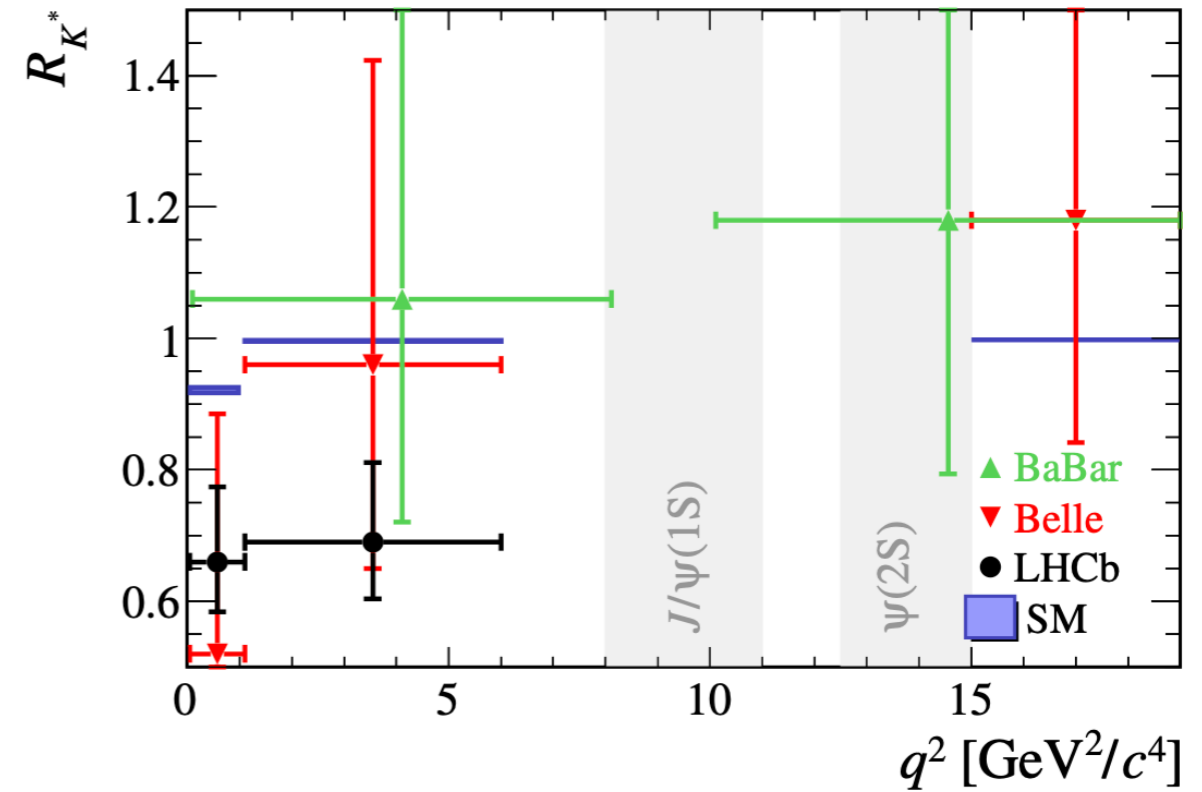
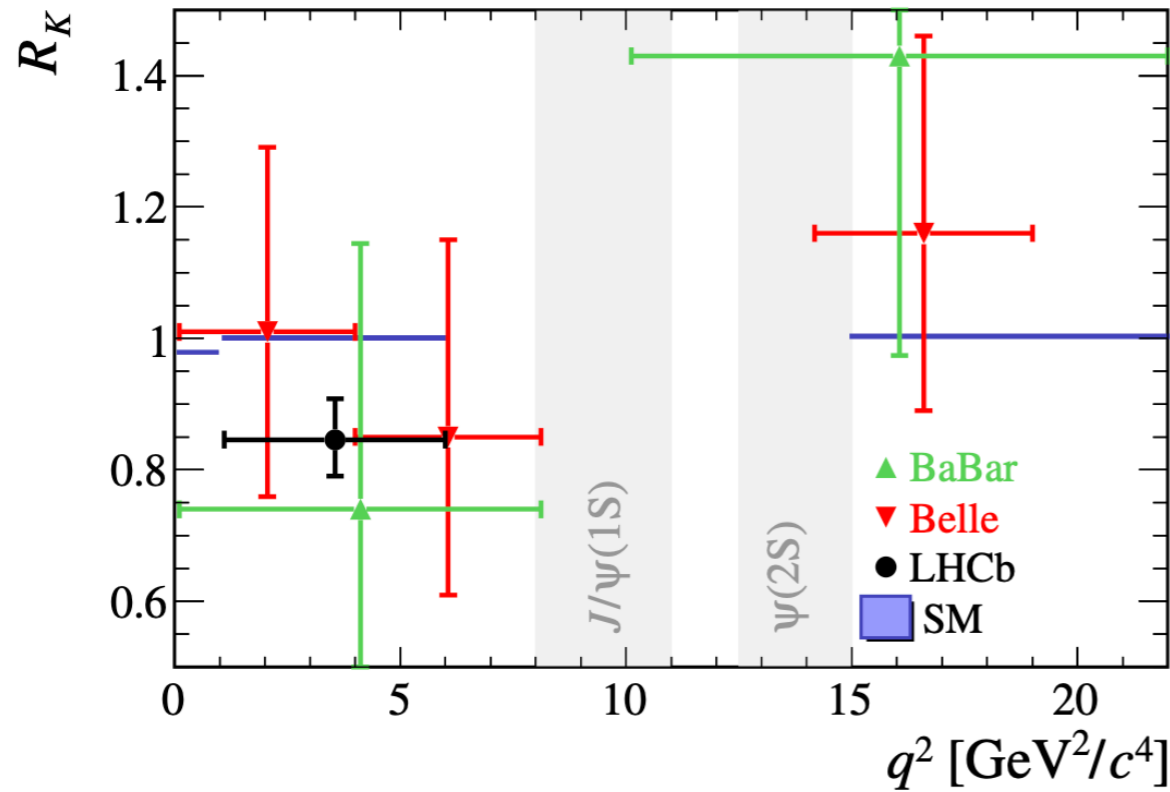
We want to investigate whether a solution to $b \rightarrow s\ell^+\ell^-$ can be excluded

What are the minimal ingredients we need?



Light Z'

$$m_{Z'} \lesssim 6 \text{ GeV}$$



Light Z'

Setup

$$m_{Z'} \lesssim 6 \text{ GeV}$$

$$\mathcal{L}_{Z'} \supset \left[\bar{\mu} \left(g_{\mu\mu}^V \gamma^\mu + g_{\mu\mu}^A \gamma^\mu \gamma_5 \right) \mu + g_{sb}^L \bar{s} \gamma^\mu P_L b + g_\chi \bar{\chi} \gamma^\mu \chi \right] Z'_\mu$$

Observables

- $b \rightarrow s \ell^+ \ell^-$
- $B \rightarrow K^{(*)} + \text{invisible}$
- $B_s - \bar{B}_s$ mixing
- $(g - 2)_\mu$
- $pp \rightarrow \mu^+ \mu^- + \text{invisible}$
- $e^+ e^- \rightarrow \mu^+ \mu^- + \text{invisible}$

Light Z'

$$m_{Z'} \lesssim 6 \text{ GeV}$$

$$b \rightarrow s\ell^+\ell^- \text{ \& } (g-2)_\mu$$

$$\mathcal{L}_{Z'} \supset \left[\bar{\mu} \left(g_{\mu\mu}^V \gamma^\mu + g_{\mu\mu}^A \gamma^\mu \gamma_5 \right) \mu + g_{sb}^L \bar{s} \gamma^\mu P_{L,R} b + g_\chi \bar{\chi} \gamma^\mu \chi \right] Z'_\mu$$

$$b \rightarrow s\ell^+\ell^-$$

$$(g-2)_\mu$$

$$C_9 = \frac{g_{sb}^L g_{\mu\mu}^V}{q^2 - m_{Z'}^2 + im_{Z'} \Gamma_{Z'}}$$

$$C_{10} = \frac{g_{sb}^L g_{\mu\mu}^A}{q^2 - m_{Z'}^2 + im_{Z'} \Gamma_{Z'}}$$

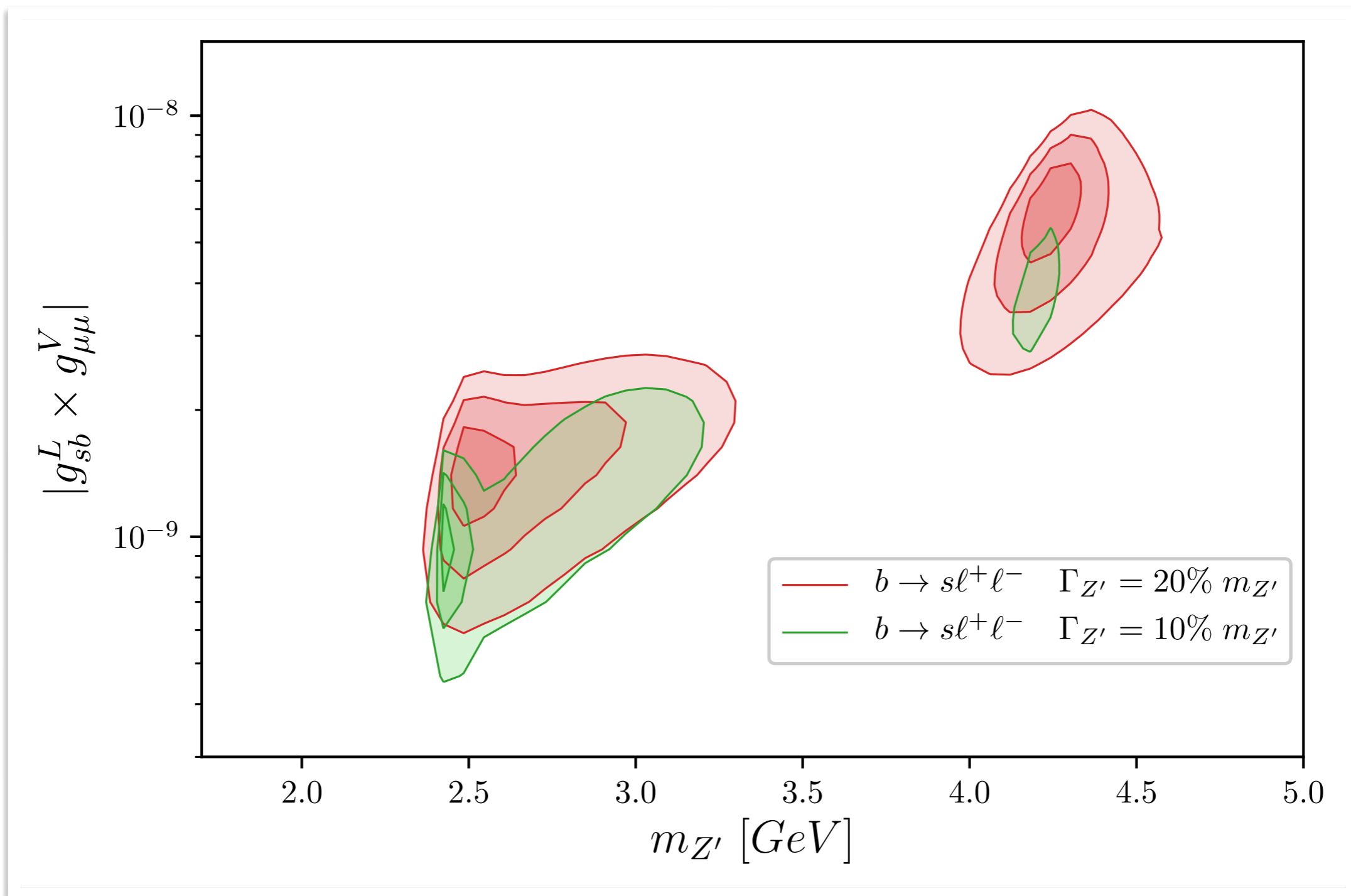
$$\Delta a_\mu = \frac{m_\mu^2}{12\pi^2 M_{Z'}^2} \text{Re} \left[(g_{\mu\mu}^V)^2 - 5(g_{\mu\mu}^A)^2 \right]$$

$$g_V = -\sqrt{5} g_A$$

Light Z'

$$b \rightarrow s\ell^+\ell^-$$

$$C_9 = -\sqrt{5}C_{10}$$



Light Z'

$$B \rightarrow K^{(*)} \nu \nu$$

$$\mathcal{B}(B \rightarrow K^{(*)} \chi \bar{\chi}) = \mathcal{B}(Z' \rightarrow \bar{\chi} \chi) \int_{s_{\min}}^{s_{\max}} ds \Gamma_{Z'}(s) \text{BW}(s) \mathcal{B}(B \rightarrow K^{(*)} Z')(s)$$

$$s_{\min} = 4m_{\chi}^2$$

$$s_{\max} = (m_B - m_{K^{(*)}})^2$$

$$\text{BW}(s) = \frac{\sqrt{s}}{\pi} \frac{1}{(s - m_{Z'}^2)^2 + \Gamma_{Z'}(s)^2 m_{Z'}^2}$$

$$\Gamma_{Z'}(s) = \frac{g_{\chi}^2}{12\pi\sqrt{s}} \sqrt{1 - 4\frac{m_{\chi}^2}{s}} (s + 2m_{\chi}^2)$$

Belle II:

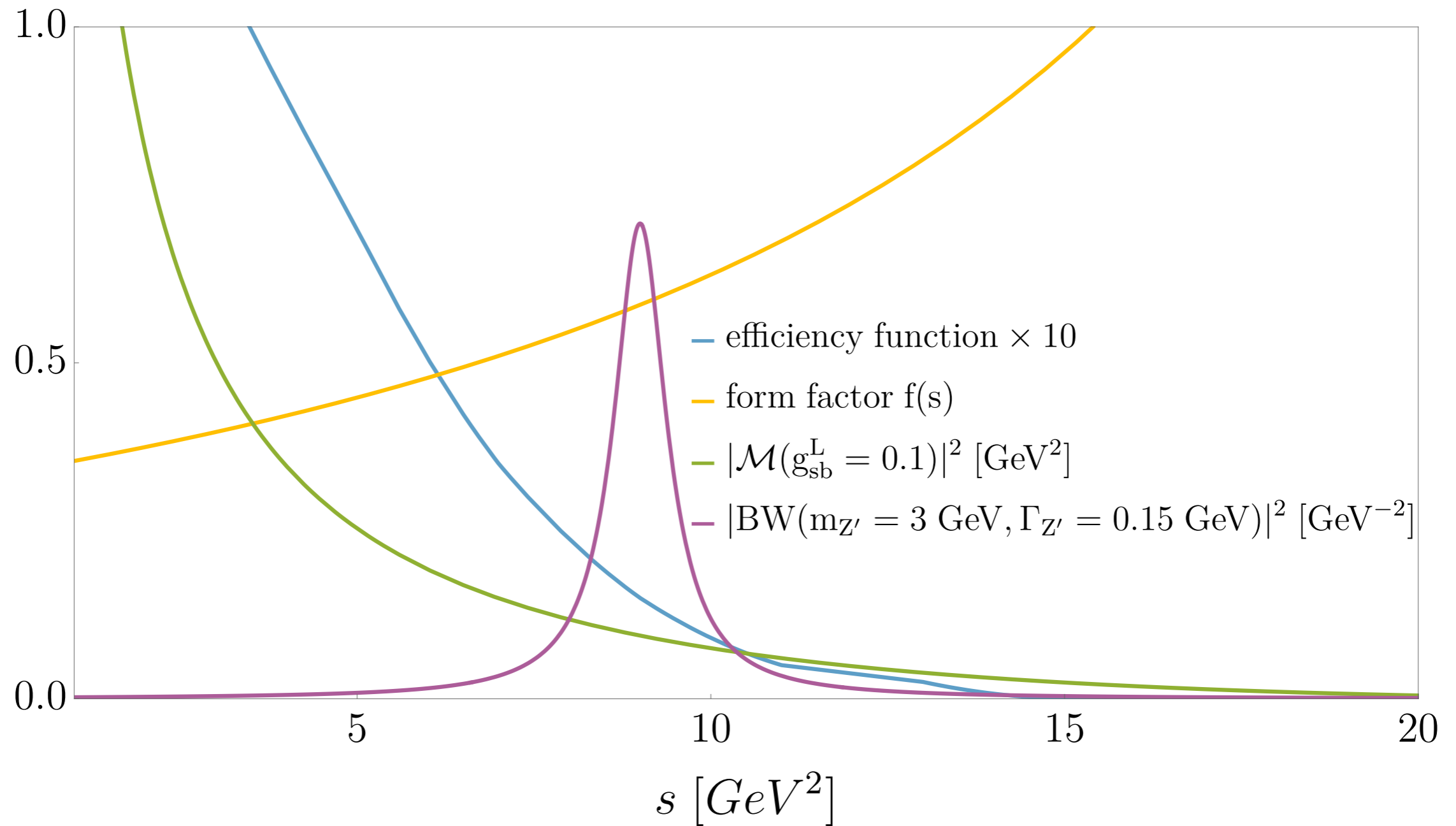
$$\mathcal{B}(B^+ \rightarrow K^+ \nu \bar{\nu}) < 4.1 \times 10^{-5}$$

Others: Babar, Belle

Light Z'

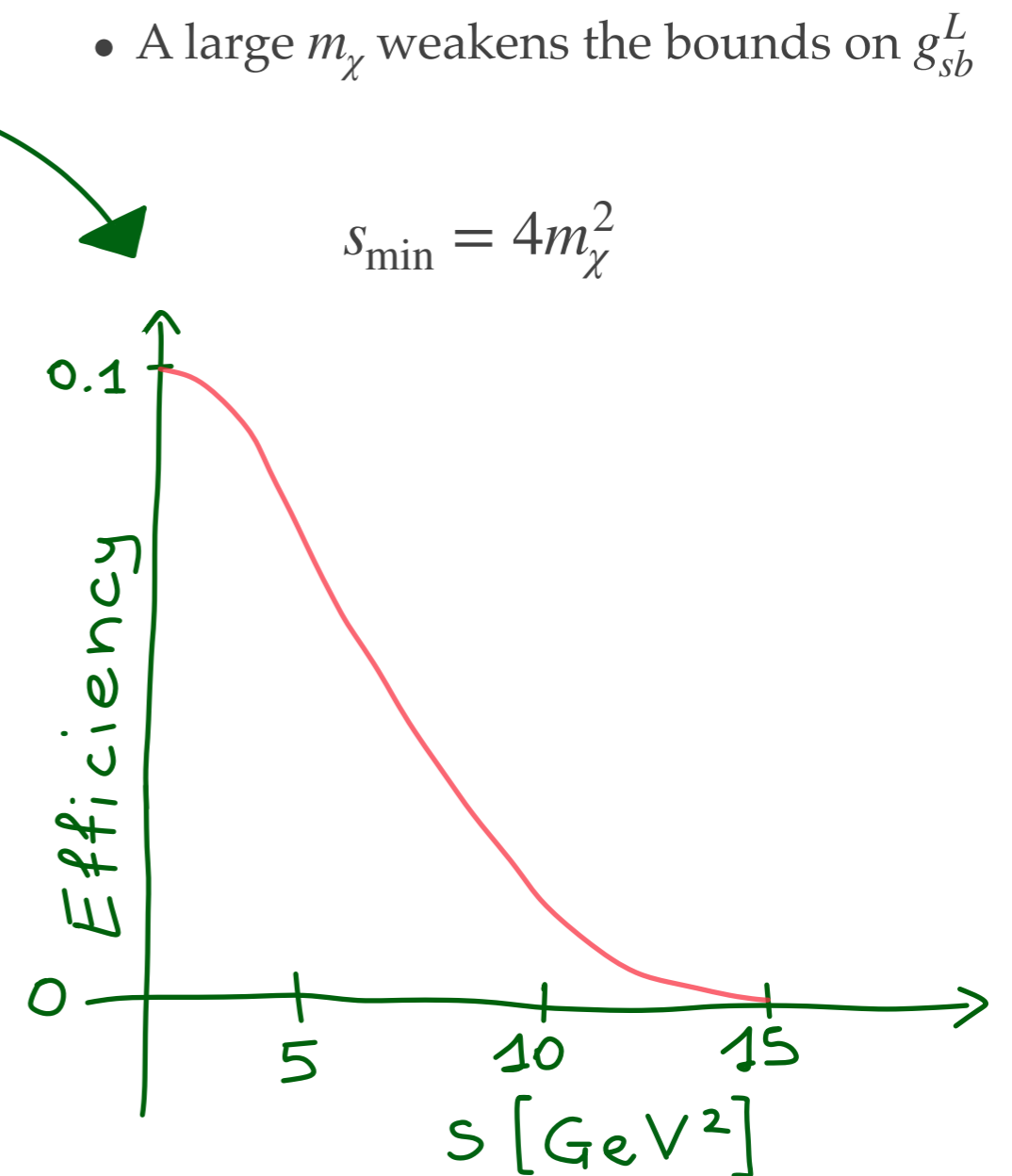
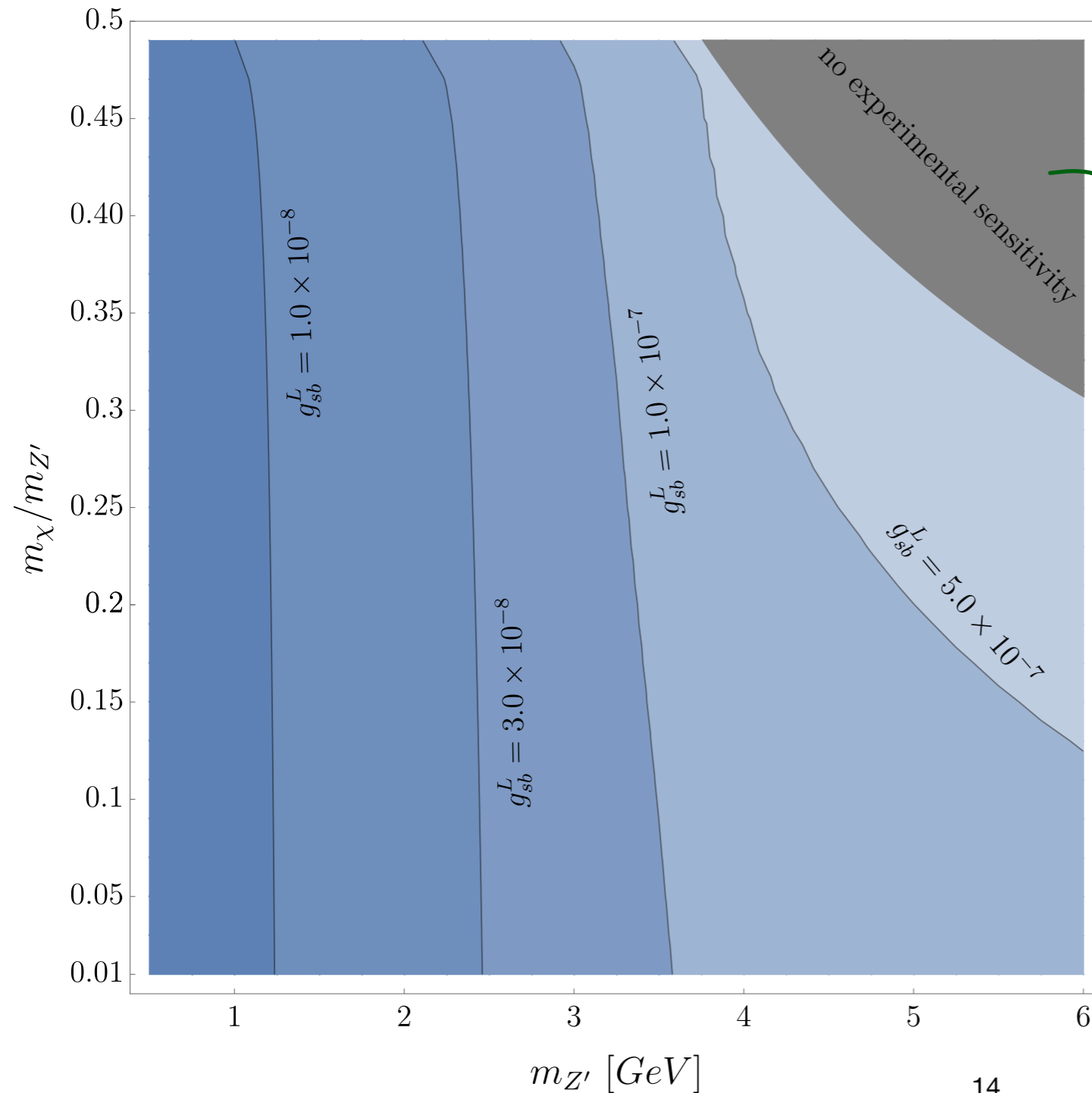
$$B \rightarrow K^{(*)} \nu \nu$$

$$B \rightarrow K \nu \nu$$



Light Z'

Recasted Belle II analysis



Light Z'

Other bounds

$B_s - \bar{B}_s$ mixing

- Light Z' masses: OPE in $m_{Z'}/m_B \rightarrow$ weaker than $B \rightarrow K^{(*)}\nu\nu$
- For $m_{Z'} \sim m_B$: no reliable treatment yet

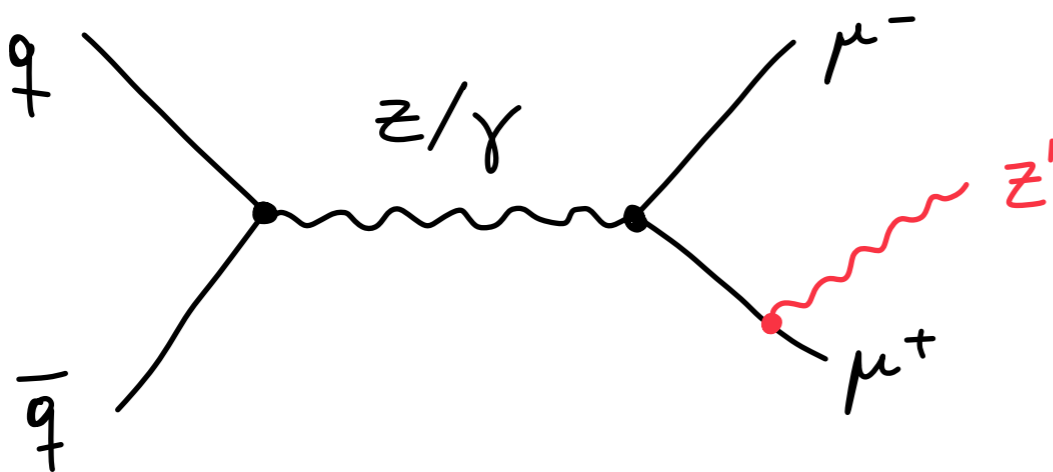
$e^+e^- \rightarrow \mu^+\mu^- + \text{invisible}$

- Recasted Belle II analysis for a sizeable Z' width:
 - 276 pb⁻¹ [hep/1912.11276](https://arxiv.org/abs/hep/1912.11276)
 - 50 fb⁻¹
 - 5 ab⁻¹

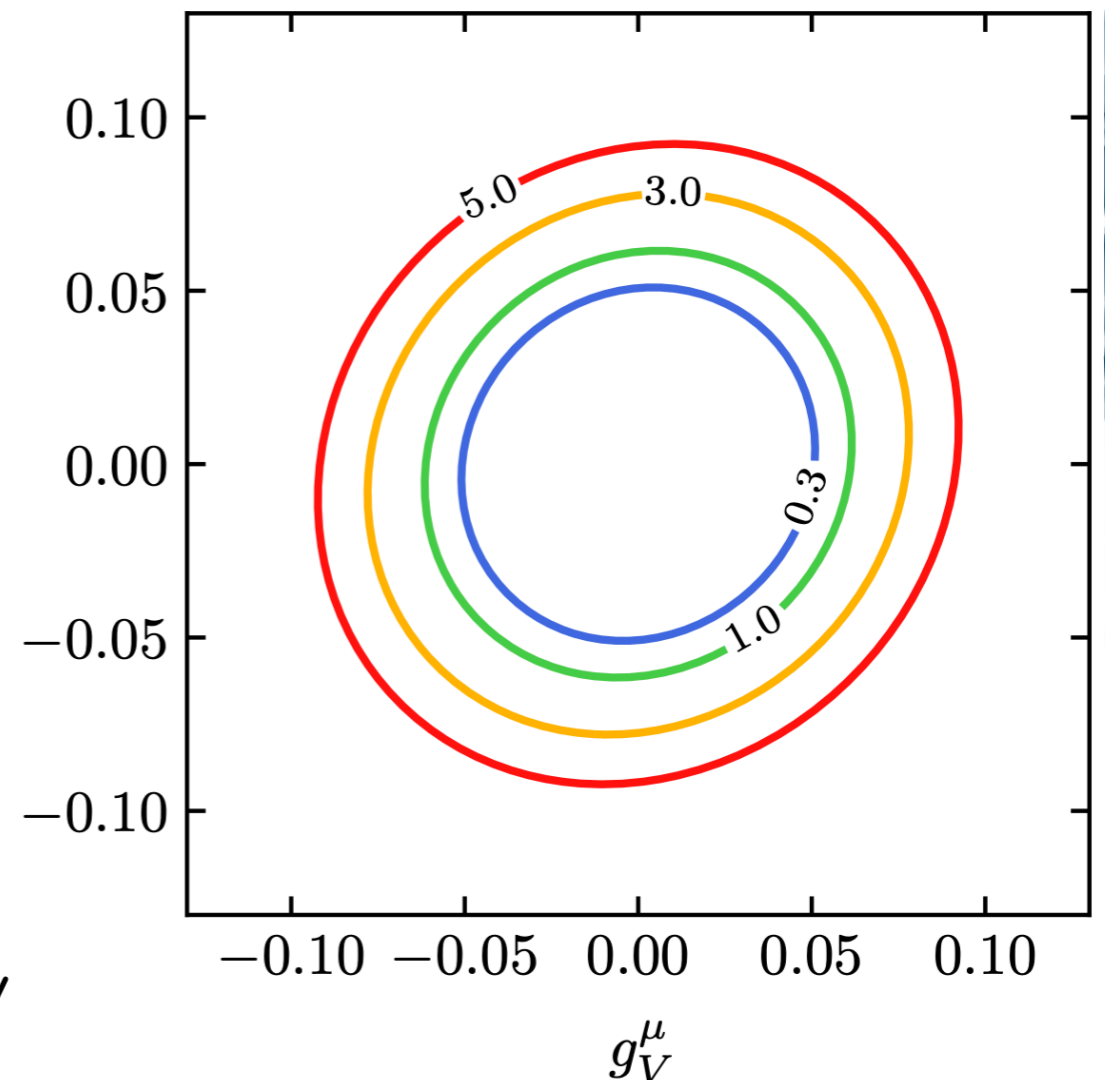
Limits on $g_{\mu\mu}^V$ of the order $10^{-1} - 10^{-2}$

$pp \rightarrow \mu^+\mu^- + \text{invisible}$

F. Bishara, U. Haisch, P.F. Monni [hep/1705.03465](https://arxiv.org/abs/hep/1705.03465)



g_A^μ



ATLAS 3.2 fb⁻¹ $\sqrt{s} = 13$ TeV

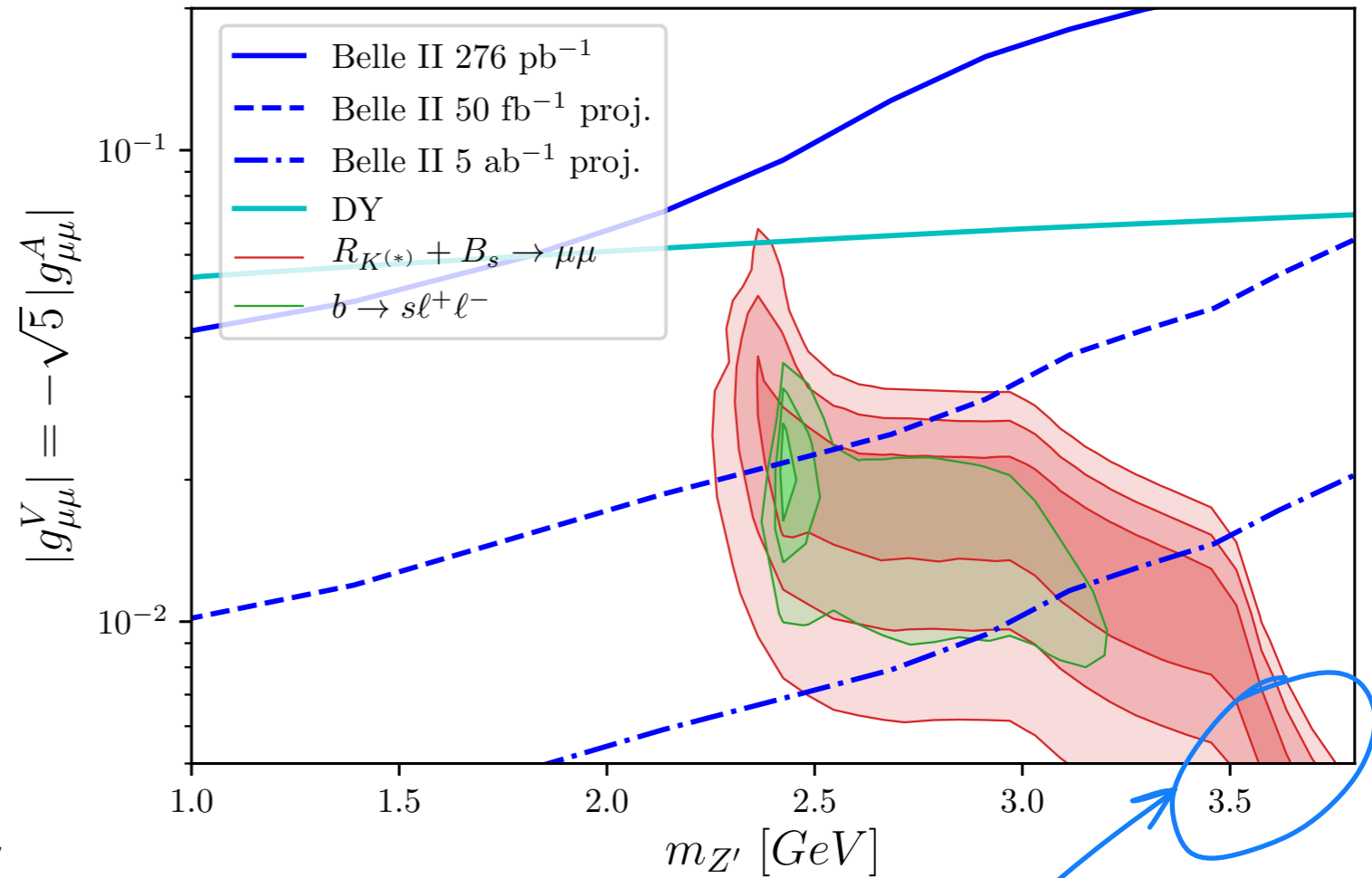
Light Z'

Bounds on $g_{\mu\mu}^V$

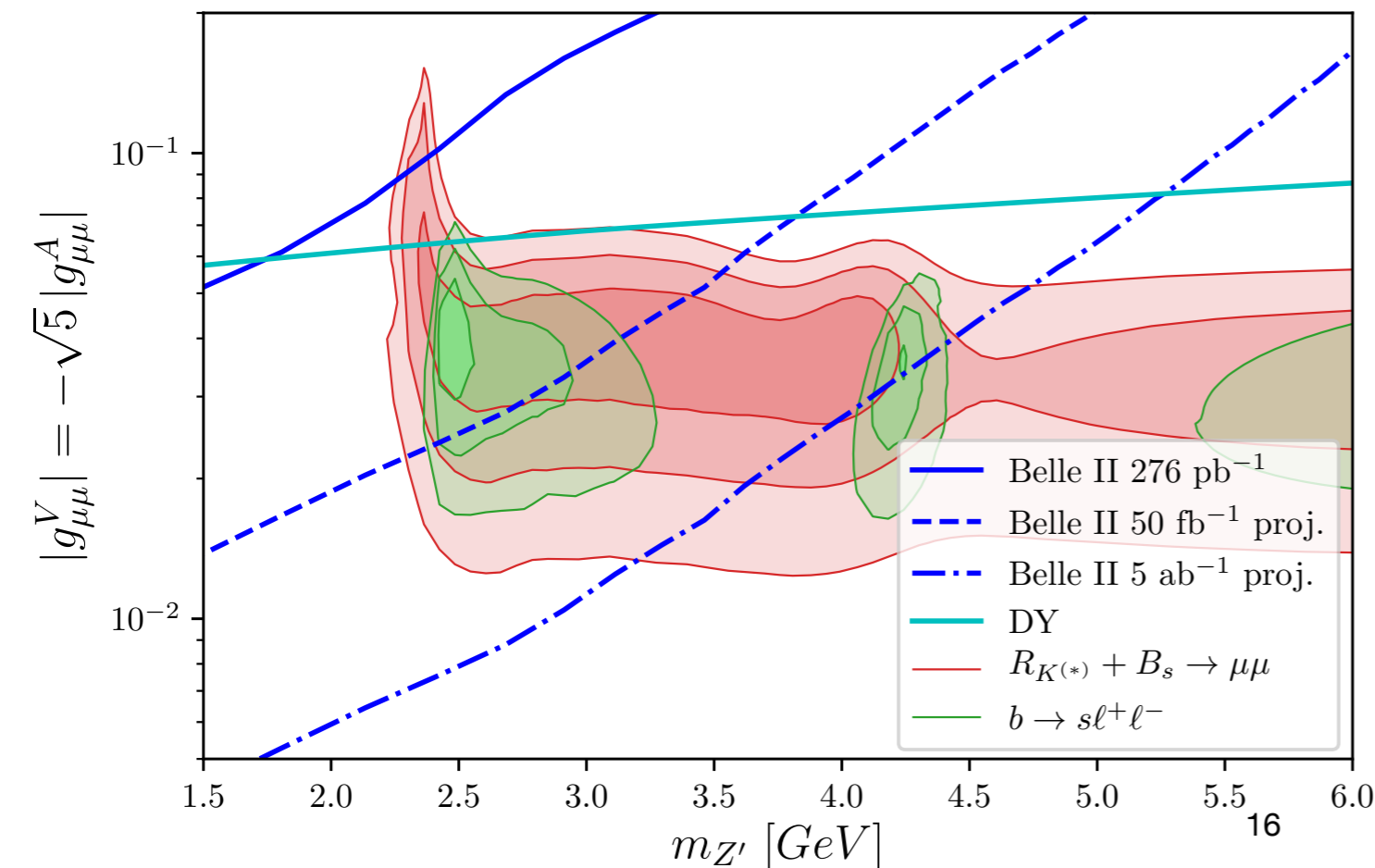
Maximum g_{bs}^L allowed by $B \rightarrow K^{(*)}\nu\nu$

Cannot exclude a light Z' yet

$$\Gamma_{Z'} = 10\% m_{Z'} \quad m_\chi = 0.49 \times m_{Z'}$$



$$\Gamma_{Z'} = 15\% m_{Z'} \quad m_\chi = 0.10 \times m_{Z'}$$



Belle efficiency drops
 $\Rightarrow g_{bs}^L$ can be very large

- Future Belle II analysis can exclude

$$m_{Z'} \sim 2.5 \text{ GeV}$$

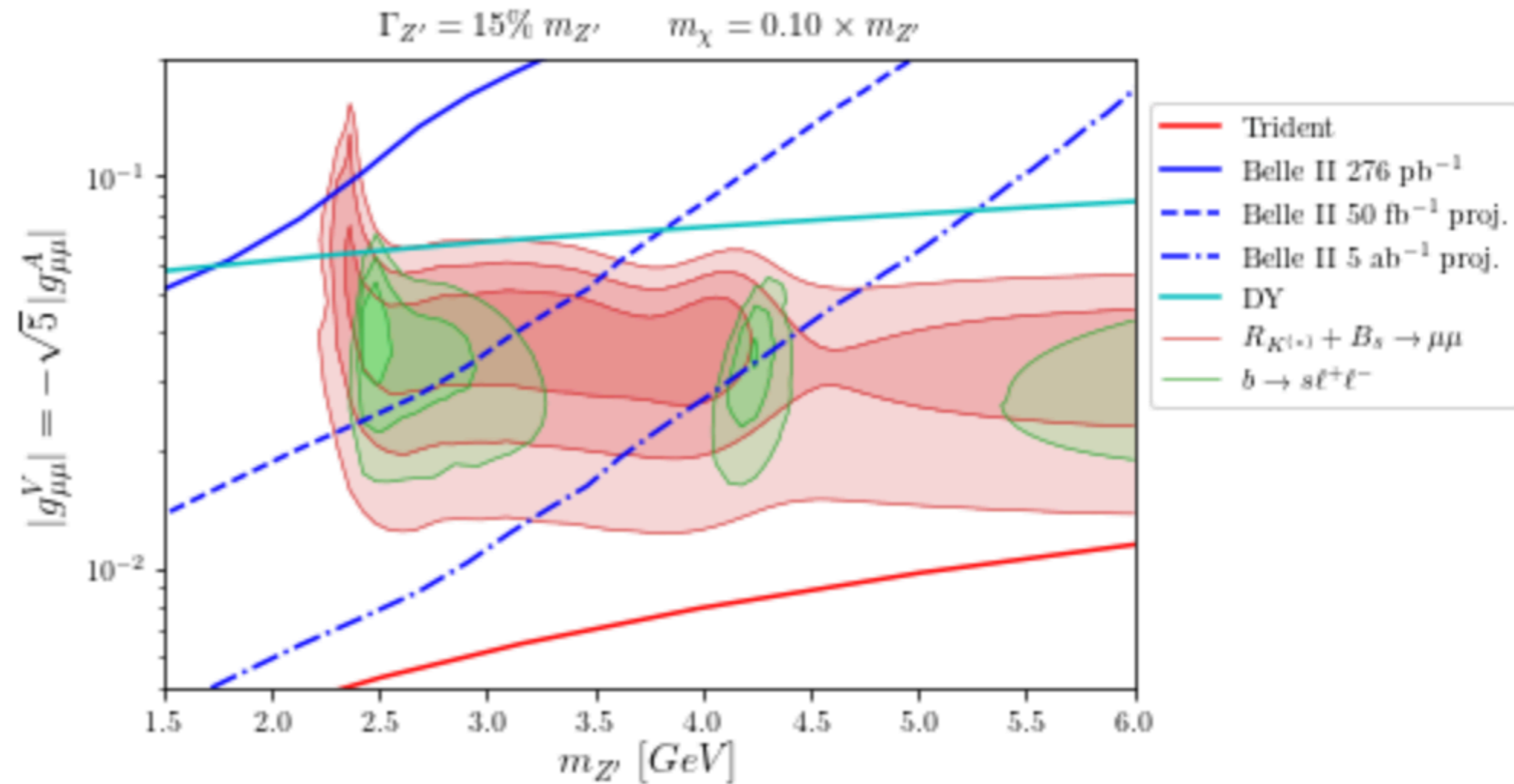
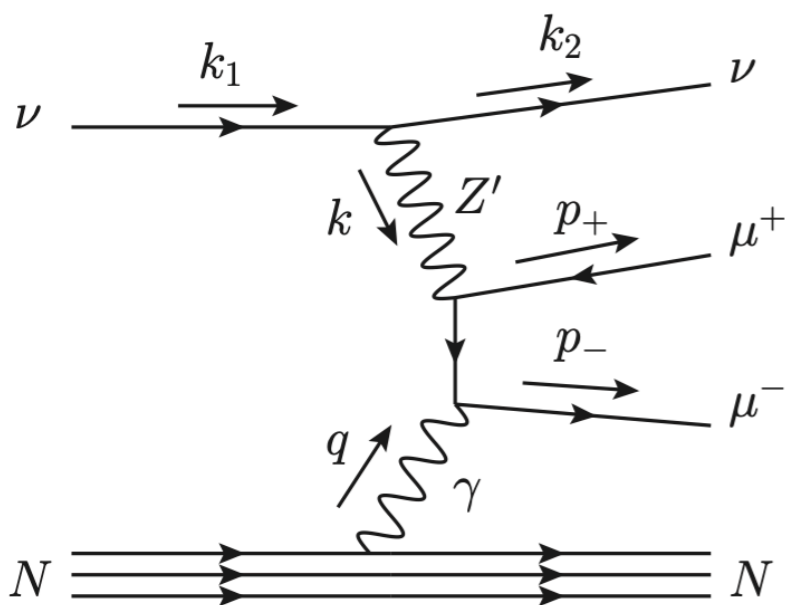
- $m_{Z'} \sim 4 \text{ GeV}$ would not lead to

$$R(K^{(*)}) > 1 \text{ in high-}q^2 \text{ bins}$$

Light Z'

Bounds on $g_{\mu\mu}^V$

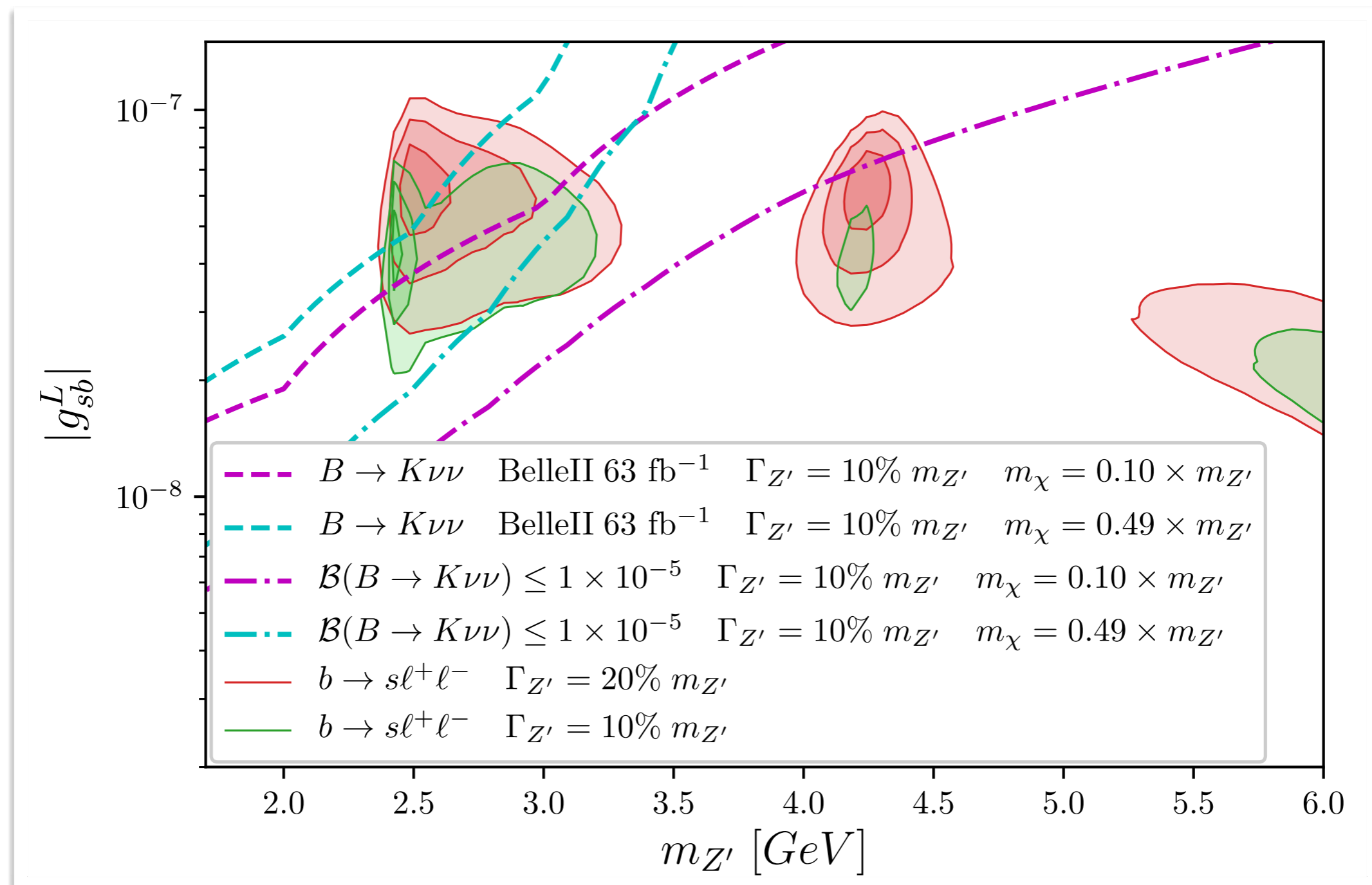
What about Neutrino Trident Production?



Light Z'

Bounds on g_{sb}^L

Alternatively we can maximize $g_{\mu\mu}^V$ from $e^+e^- \rightarrow \mu^+\mu^- + \text{invisible}$



Thank you for your attention!

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