# Fast and accurate gravitational-wave modelling with principal component regression

GdR OG 11/10/2021

Cyril Cano (Gipsa-lab) Nicolas Le Bihan (Gipsa-lab) Éric Chassande-Mottin (APC)

Preprint: https://tds.virgo-gw.eu/ql/?c=16877 Git: https://git.ligo.org/cyril.cano/gw-generation





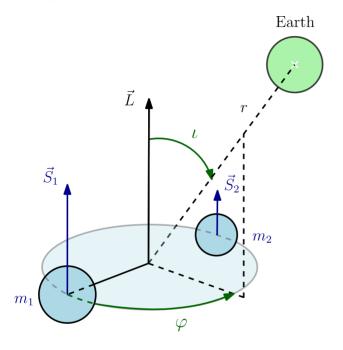




### **Motivations**

- Bayesian inference needs large amount of waveforms (~10<sup>5</sup>)
- Time domain GW generation is computationally expensive
- Need for fast and accurate generative model
- Reduced Order Models [Pürrer 2016]
- ML with Mixture of Experts (med. mismatch ~10<sup>-4</sup>) [Schmidt et al. 2020]
- Proposed model: **principal component regression**

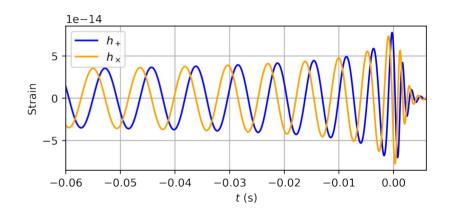
### **BBH** parameters



 $m_1,\,m_2,\,S_{1z},\,S_{2z}\,$  : binary parameters

 $\iota, \varphi$ : line of sight

 $\ ^{r}: \\ \text{luminosity distance}$ 



### **Generative model**

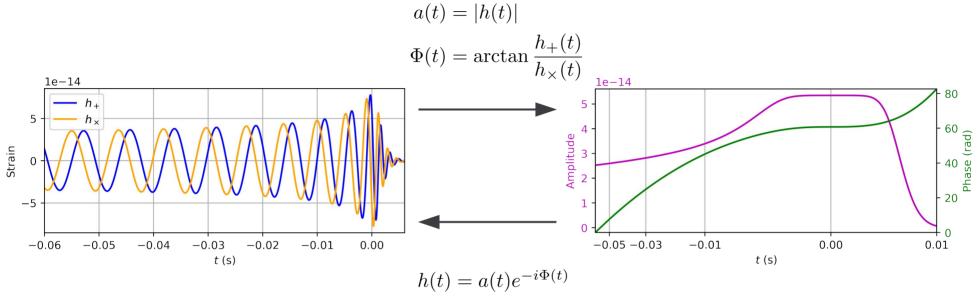
SEOBNRv4

Only (2,2) mode

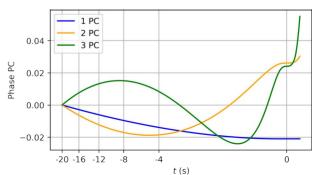
$$h(t;m_1, m_2, S_{1z}, S_{2z}, \iota, \varphi, r)$$

$$h(t) = h_{+} - ih_{\times}(t)$$

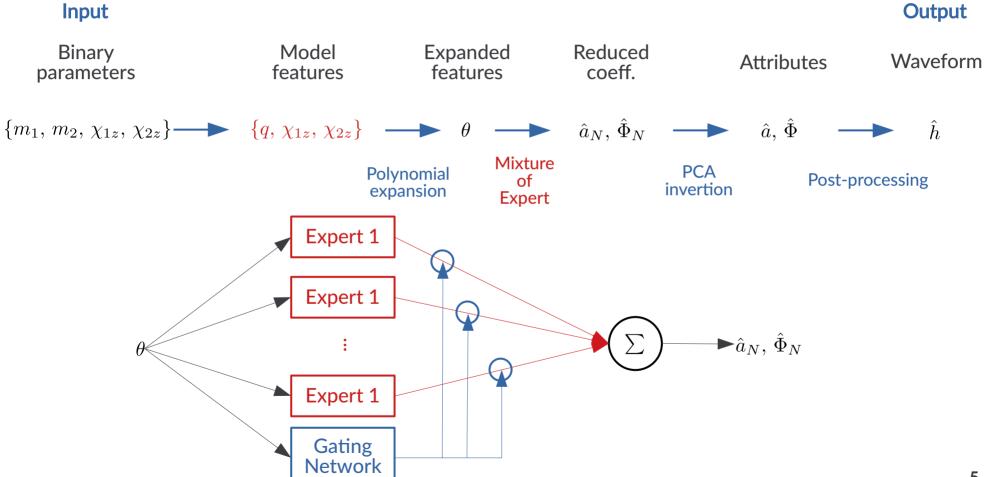
### Waveform attributes



- ML model generates amplitude and phase
- Non uniform time grid  $sign(t)|t|^{\frac{1}{\alpha}}$
- Needs dimension reduction: truncated PCA



### Schmidt's model [Schmidt et al. 2020]



### Overview of our model

#### Generation

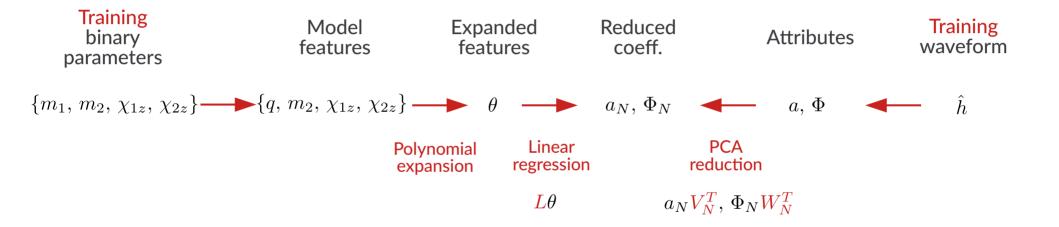
**Output** Input Expanded Reduced Binary Model **Attributes** Waveform features coeff. parameters features  $\{m_1, m_2, \chi_{1z}, \chi_{2z}\}$   $\theta$   $\hat{a}_N, \hat{\Phi}_N$   $\hat{a}, \hat{\Phi}$ Linear Polynomial PCA Post-processing invertion regression expansion  $\hat{a}_N V_N, \hat{\Phi}_N W_N$  $L\theta$ 

#### Goodness-of-fit metric:

$$\operatorname{mismatch}(h,g) = \min_{\tau \in \mathbb{R}} \left[ 1 - \frac{|\langle h_{\tau}, g \rangle|}{||h_{\tau}||||g||} \right] \quad \text{with} \quad \langle f, g \rangle = \int \frac{h(f)g^*(f)}{S(f)} \, df$$

### Overview of our model

### **Fitting**



### Goodness-of-fit metric:

$$\operatorname{mismatch}(h,g) = \min_{\tau \in \mathbb{R}} \left[ 1 - \frac{|\langle h_{\tau}, g \rangle|}{||h_{\tau}||||g||} \right] \qquad \text{with} \qquad \langle f, g \rangle = \int \frac{h(f)g^*(f)}{S(f)} \, df$$

# Hyperparameters tuning

#### Feature set

Tested features:

$$m_1, m_2, \chi_{1z}, \chi_{2z}, q, \mathcal{M}, \chi_{\text{eff}}, m_1^{-1}, m_2^{-1}$$

Tested feature sets:

$$\{m_1\}$$
  $\{m_1, m_2\}$   $\{m_1, m_2, \chi_{1z}\}$  ...  $\{m_2\}$   $\{m_1, \chi_{1z}\}$   $\{m_1, m_2, q\}$  ...  $\vdots$   $\vdots$ 

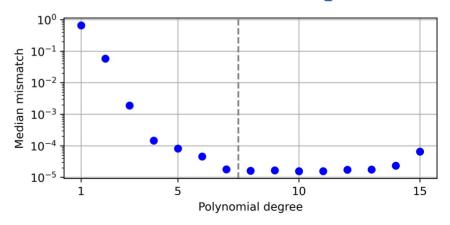
Several good choices:

$$\frac{\{\chi_{2z}, \, \chi_{\text{eff}}, \, \mathcal{M}\}}{\{q, m_2, \, \chi_{1z}, \, \chi_{2z}\}} \quad \begin{cases} q, m_2, \, \chi_{2z}, \, \chi_{\text{eff}} \} \\ \{\chi_{2z}, \, \chi_{\text{eff}}, \, m_1, \, m_2^{-1} \} \end{cases}$$
Selected

### Polynomial degree

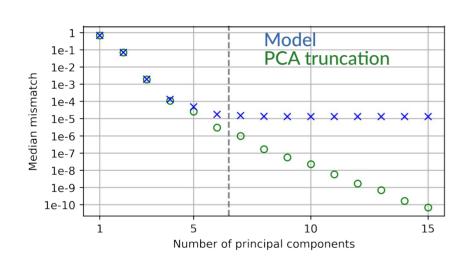
First order  $\{a,b\}$  Second order  $\{a,b,ab,a^2,\,b^2\}$  Third order  $\{a,b,ab,a^2,\,b^2,\,a^2b,\,ab^2,\,a^3,\,b^3\}$   $\vdots$ 

### Seventh order used for regression

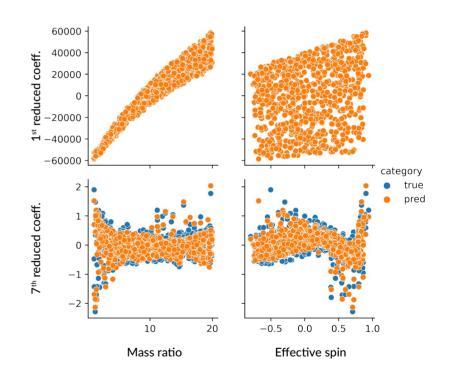


# Hyperparameters tuning

### Number of PC for the phase



Six PC used for dimension reduction



### Results on SEOBNRv4

### **Dataset properties**

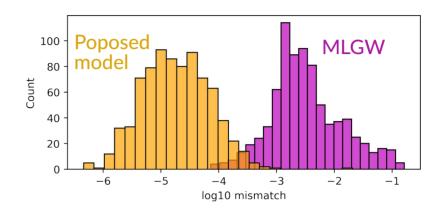
- Training size: 3200 Testing size: 800
- Mass ratio: U([1, 20])
- Dimensionless spins: U([-0.8, 0.95])

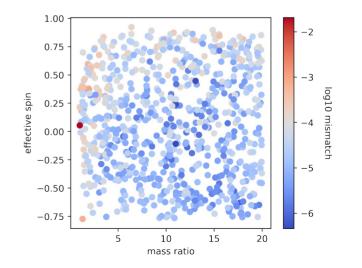
### **Accurate**

- $q_{50\%} = 2.10^{-5}$
- $95\% = 2.10^{-6}$
- $95\% = 1.5.10^{-4}$

### **Fast**

- ~100 times faster than SEOBNRv4
- Can be faster without interpolation (from non-uniform to regular time grid)





### Python library

#### **README**

Contact: cyril.cano@gipsa-lab.fr

#### Overview

This Git repo provides the library mlpgw.py that allows to train a machine-learning model able to regress gravitational-wave waveform from a set of examples as described in this article.

This Git repo includes several notebooks that allow to reproduce the results presented in the paper.

The learning part is mainly based on scikit-learn. This package is included in the required environment.

Take care that in the notebook a gravitational waveform h(t) is denoted There was an error rendering this math block but There was an error rendering this math block in the paper.

#### Installation

Clone this Git repo and create the environment gw-generation by running:

conda env create -f environment.yml

Activate the environment

conda activate gw-generation

... and run the following command line from this folder:

conda develop .

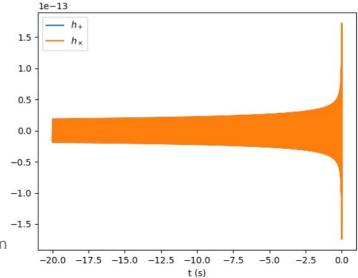
#### How to generate a waveform using a pre-computed ML model?

This notebook shows how to generate a waveform with a pre-computed ML model.

The pre-computed model is stored in a set of Pickle files (see data/)

### Python library

```
In [1]:
         1 import numpy as np
         2 import matplotlib.pyplot as plt
         3 import mlpgw
         5 %matplotlib notebook
In [2]:
         1 # Dowload the model
           model = mlpgw.load obj('../data/model')
In [3]: 1 # Make prediction
         2 h pred = model.predict(m1=15, m2=5, s1z=0.9, s2z=0.2)
In [4]:
         1 # Plot it
         2 plt.figure()
         3 plt.plot(h pred['time'], h pred['hp'], label=r'$h +$')
         4 plt.plot(h pred['time'], h pred['hc'], label=r'$h \times$')
         5 plt.xlabel('t (s)')
         6 plt.legend()
           plt.show()
```



Git: https://git.ligo.org/cyril.cano/gw-generation

# Conclusion/perspectives

- Take home messages: Fast and accurate GW generation with principal component regression
  - Applicability up to SNR ~ 288 (12 in the worst case) \*: mismatch  $< \frac{N}{2 \mathrm{SNR}^2}$
  - Non conventional features lead to better results
  - Simple method with off-the-shelf algorithms from scikit-learn

Perspectives:

- Subdominant modes
- Comparison with other ML state of the art algorithms
- Precessing BBH

## ${f R}^2$ scores

$$R^{2}(y, \hat{y}) = 1 - \frac{\sum_{i=1}^{n} (y_{i} - \hat{y}_{i})}{\sum_{i=1}^{n} (y_{i} - \bar{y})}$$

| PC | 1        | 2        | 3        | 4       | 5       | 6      |
|----|----------|----------|----------|---------|---------|--------|
| a  | 1.1e-05  | 0.00711  | 0.0591   | 0.0394  | 3.64    | 0.31   |
| Φ  | 1.95e-09 | 8.32e-07 | 0.000312 | 0.00674 | 0.00179 | 0.0643 |