

**OBSERVATORY** 

## Legacy Survey of Space and Time (LSST)





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SWINBURNE UNIVERSITY OF TECHNOLOGY

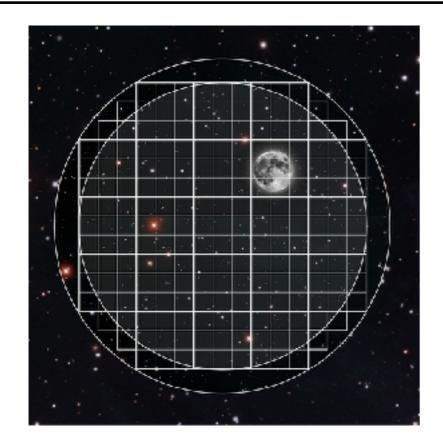
**2021 Programme National Hautes Energies** 

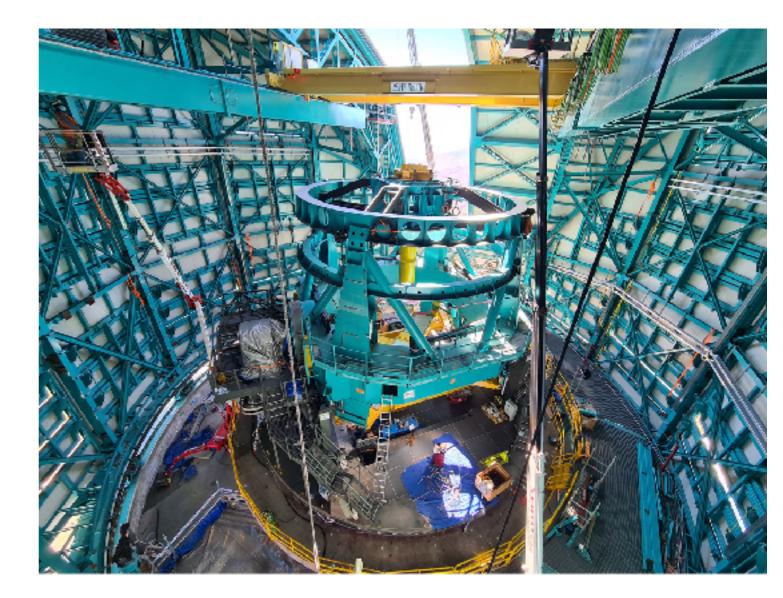


### **VERA C. RUBIN** observing facility to observe + store data OBSERVATORY

#### in a nutshell:

- telescope: 6.7-m equivalent
- world's largest CCD camera: 3.2 \* 10<sup>9</sup> pixels
- Resolves a golf ball 24 kilometers away!





LSST Project/NSF/AURA



# LSST a deep "video" of the sky

#### in numbers:

- 10-year photometric survey
- 6 optical bands: *ugrizy* (320-1050nm)
- 1,000 images/night = 15 TB/night
- 10 million transient candidates per night
- re-scheduled to start October 2023

#### TABLE 3: Various science metrics as functions of survey duration.

Quantity	Year 1	Y3	Y5	Y8	Year 10
$r_5 \operatorname{coadd}^a$	26.0	26.5	26.8	27.1	27.2
$\sigma$ (i=25) <sup>b</sup>	0.12	0.07	0.06	0.05	0.04
color vol. <sup>c</sup>	316	20	6	1.7	1
# of visits <sup>d</sup>	83	248	412	660	825
$\sigma_{\pi}$ (r=24) <sup>e</sup>	9.5	5.5	4.2	3.3	3.0
$\sigma_{\mu}$ (r=24) $^{f}$	32	6.1	2.8	1.4	1.0

 $^{a}$  The co-added depth in the r band (AB, 5 $\sigma$ ; point sources).

<sup>*b*</sup> The photometric error for a point source with i = 25.

<sup>c</sup> The volume of the 5-dimensional color error space, normalized by the final v <sup>d</sup> The number of visits per sky position (summed over all bands).

<sup>*e*</sup> The trigonometric parallax accuracy for a point source with r=24 (milliarcsec <sup>*f*</sup> The proper motion accuracy for a point source with r=24 (milliarcsec/yr).



# LSST a deep "video" of the sky

#### Wide-Fast-Deep survey (90% time)

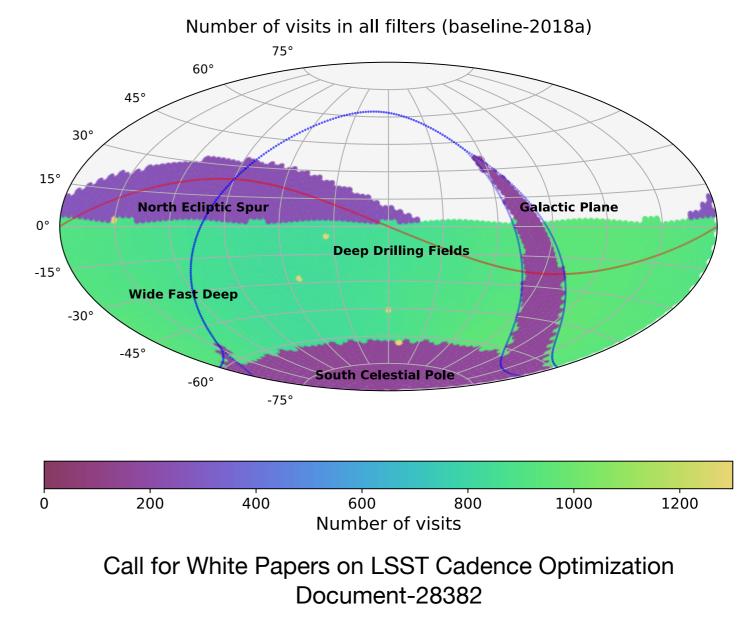
Baseline: 18,000 deg<sup>2</sup>, all visible sky in 3 nights At least 2 visits per night, +1 color

#### **Deep drilling fields (5%)**

4 fields selected (ELAIS-S1, XMM-LSS, CDF-S, and COSMOS), 1 in discussion (EDFS)

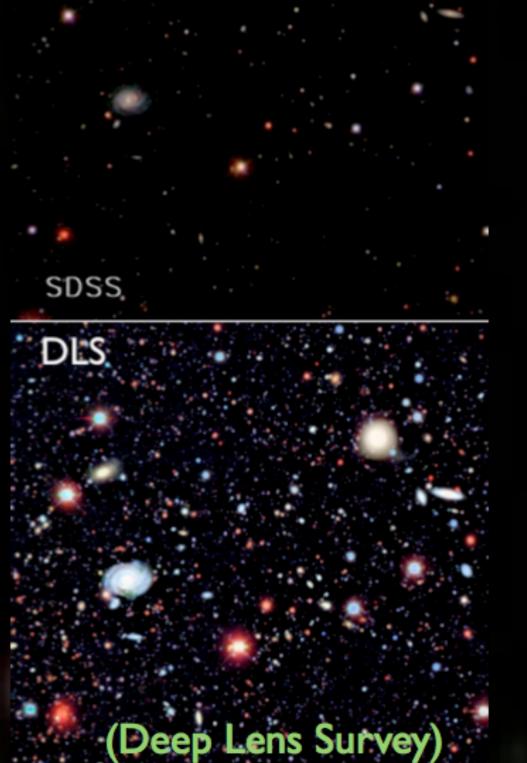
#### Candidate mini surveys (~5%)

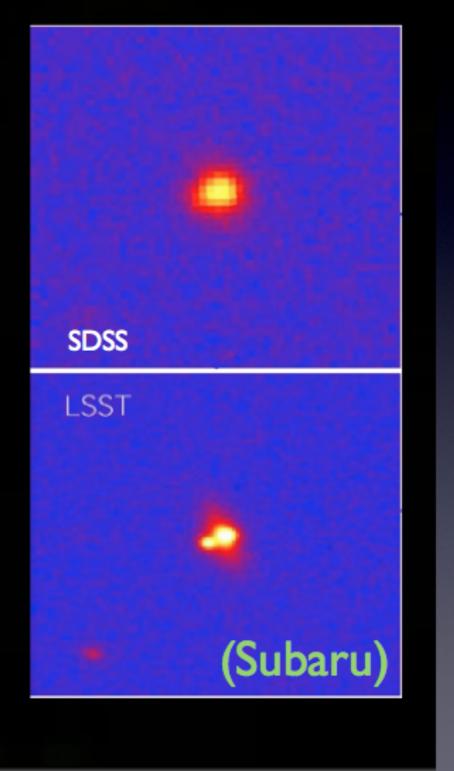
North Ecliptic Spur, the Galactic Plane, the South Celestial Pole, and five Deep Drilling mini surveys (Elais S1, XMM-LSS, Extended Chandra Deep Field-South, and COSMOS)



#### Ongoing optimization is based on a full simulation of 10 years of operations

# SDSS-LSST comparison: LSST=d(SDSS)/dt, LSST=SuperSDSS 7x7 arcmin, gri



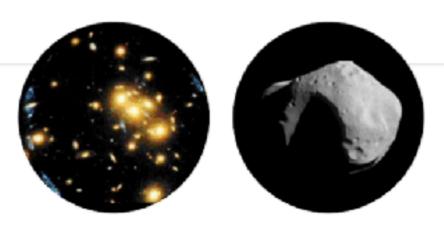


Slide from Ivezic

## Four Science Goals

#### Dark Matter, Dark Energy

- Weak Lensing
- Baryon acoustic oscillations
- Supernovae, Quasars



#### **Cataloging the Solar System**

- Potentially Hazardous Asteroids
- Near Earth Objects
- Object inventory of the Solar System

#### Milky Way Structure & Formation

- Structure and evolutionary history
- Spatial maps of stellar characteristics
- Reach well into the halo



#### **Exploring the Transient sky**

- Variable stars, Supernovae
- Fill in the variability phase-space
- Discovery of new classes of transients

"From Science Drivers to Reference Design", Ivezić et al. (2008), arXiv:0805.2366

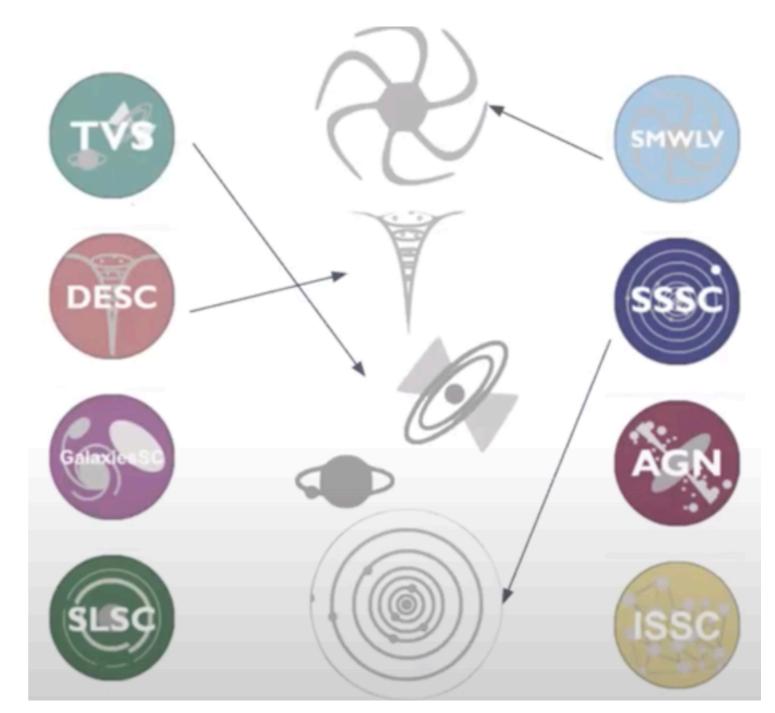
Vera C. Rubin Observatory | Joint Status Review | Aug 31 - Sept 3, 2020

Community: ~1500 researchers Non-US: ~400 international contributors France: ~80 members (~11 INSU ; ~70 IN2P3) info from E. Gangler 6



# LSST science collaborations

- Galaxies
- Stars, MW, local volume
- Solar system
- Dark Energy
- Active Galactic Nuclei
- Transients/variable stars
- Strong lensing
- Informatics and statistics





# LSST data products

#### VERA C. RUBIN OBSERVATORY

Raw Data Sequential 30s image, 20TB/night

37s

Now

**Prompt Data Product** Difference Image Analysis Alerts: up to 10 million per night

public in "real time"



**Prompt Products DataBase** Images, Object and Source catalogs from DIA Orbit catalog for ~6 million Solar System bodies



Annual Data Release Accessible via the LSST Science Platform & LSST Data Access Centers.

#### public yearly

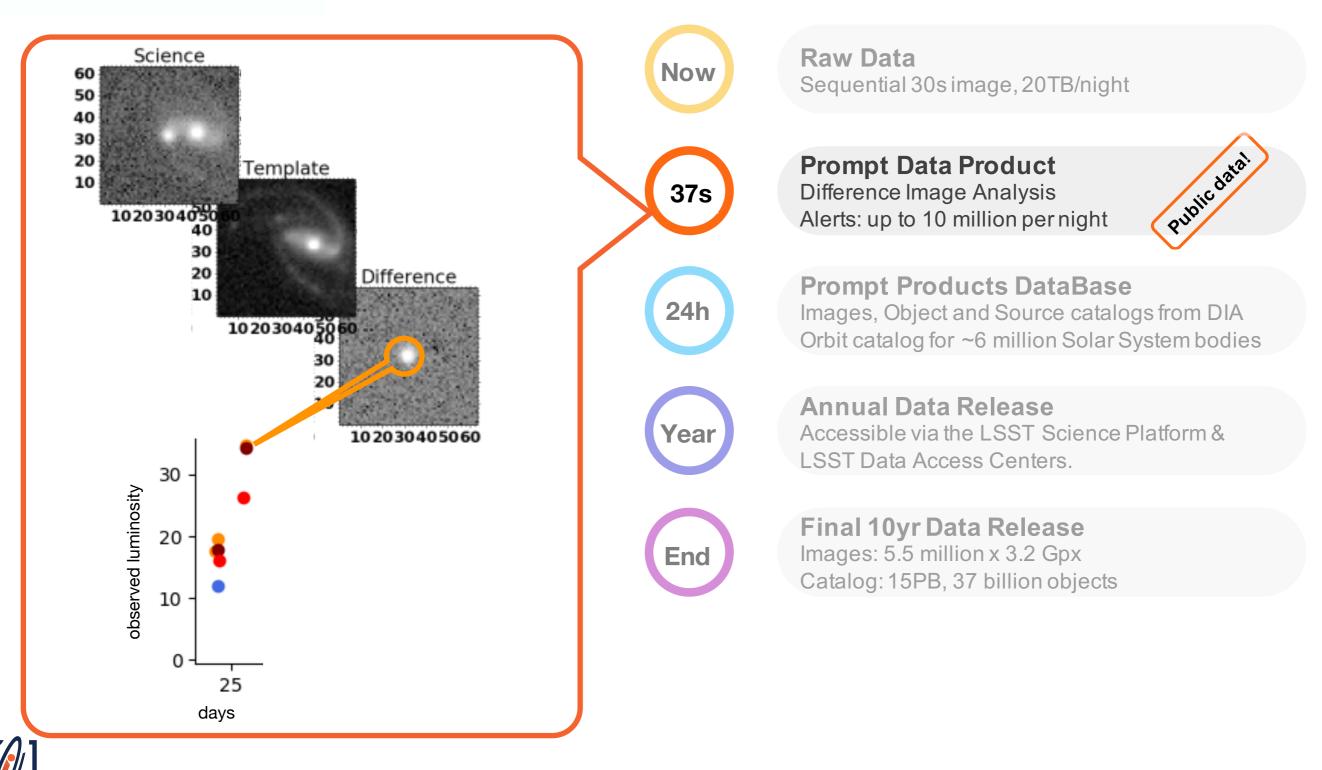


**Final 10yr Data Release** Images: 5.5 million x 3.2 Gpx Catalog: 15PB, 37 billion objects public at the end of the survey

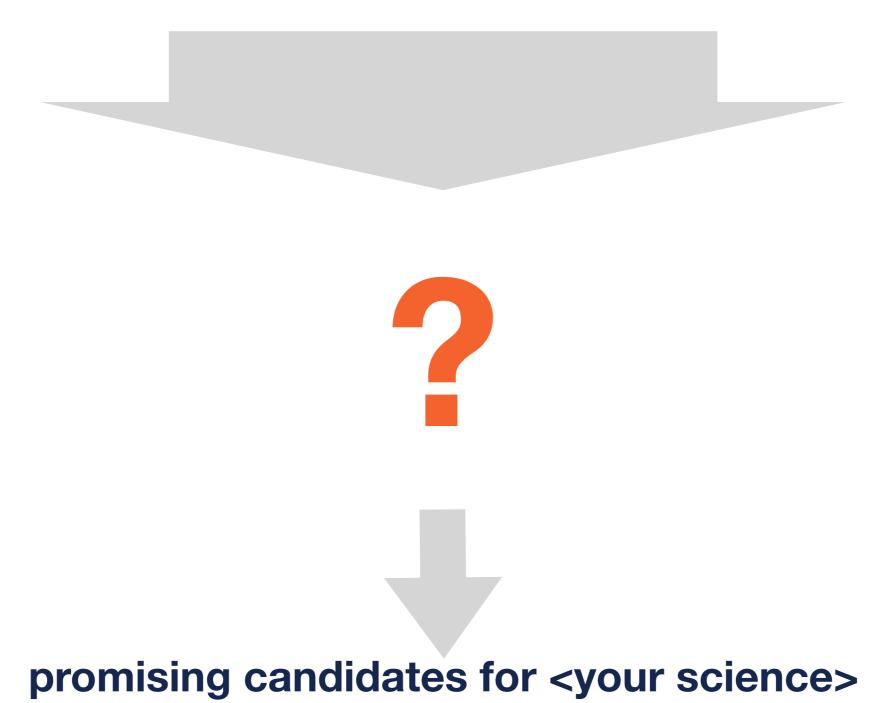


# LSST data products

#### VERA C. RUBIN OBSERVATORY



## LSST ~ 10 million transient alerts per night



+ connecting with multi-wavelength/messenger surveys?



## LSST ~ 10 million transient alerts per night



+ connecting with multi-wavelength/messenger surveys?



# The broker panorama





Fink is:

- a community driven effort, open to anyone
- a broker designed for the LSST alert stream

## **Goal: Multi-science broker**

#### MNRAS 2020, arXiv: 2009.10185 Anais Möller, Julien Peloton, Emille E. O. Ishida,

Chris Arnault, Etienne Bachelet , Tristan Blaineau, Dominique Boutigny, Abhishek Chauhan, Emmanuel Gangler, Fabio Hernandez, Julius Hrivnac, Marco Leoni, Nicolas Leroy, Marc Moniez, Sacha Pateyron, Adrien Ramparison, Damien Turpin, Réza Ansari, Tarek Allam Jr., Armelle Bajat , Biswajit Biswas, Alexandre Boucaud, Johan Bregeon, Jean-Eric Campagne, Johann Cohen-Tanugi, Alexis Coleiro, Damien Dornic, Dominique Fouchez, Olivier Godet, Philippe Gris, Sergey Karpov, Ada Nebot Gomez-Moran, Jérémy Neveu, Stephane Plaszczynski, Volodymyr Savchenko, Natalie Webb



### https://fink-broker.org/members/

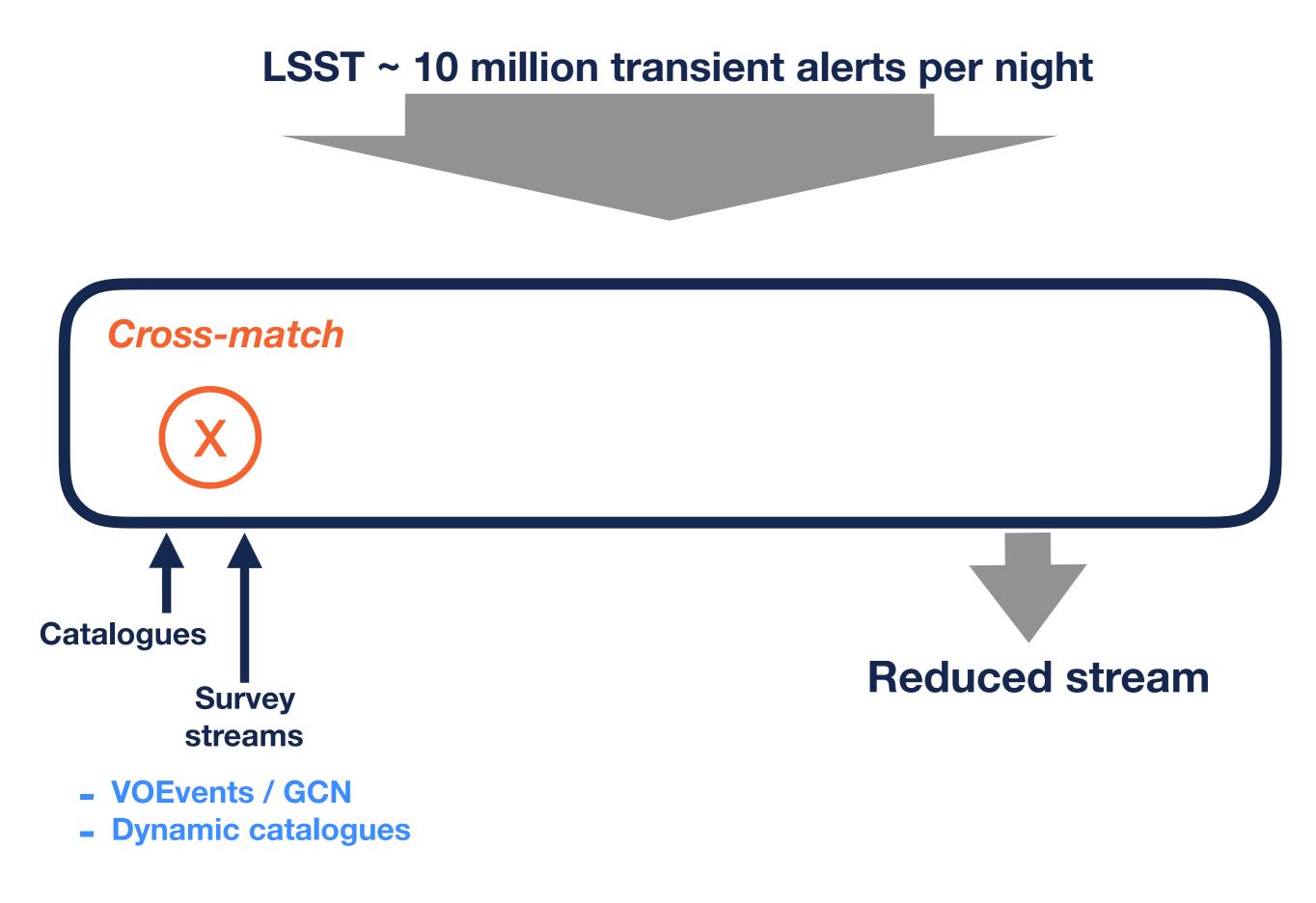
13

### LSST ~ 10 million transient alerts per night

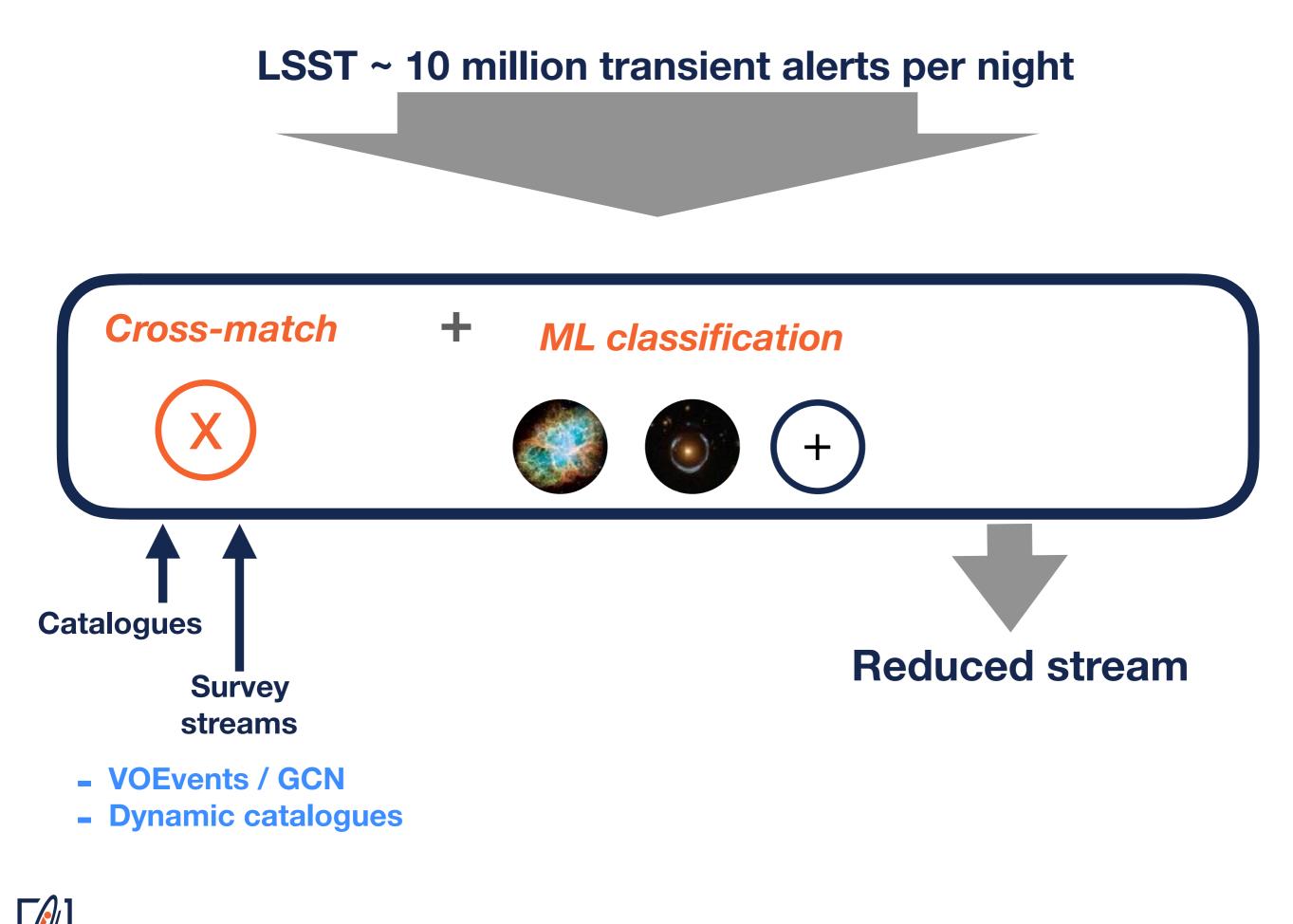


# **Reduced stream**

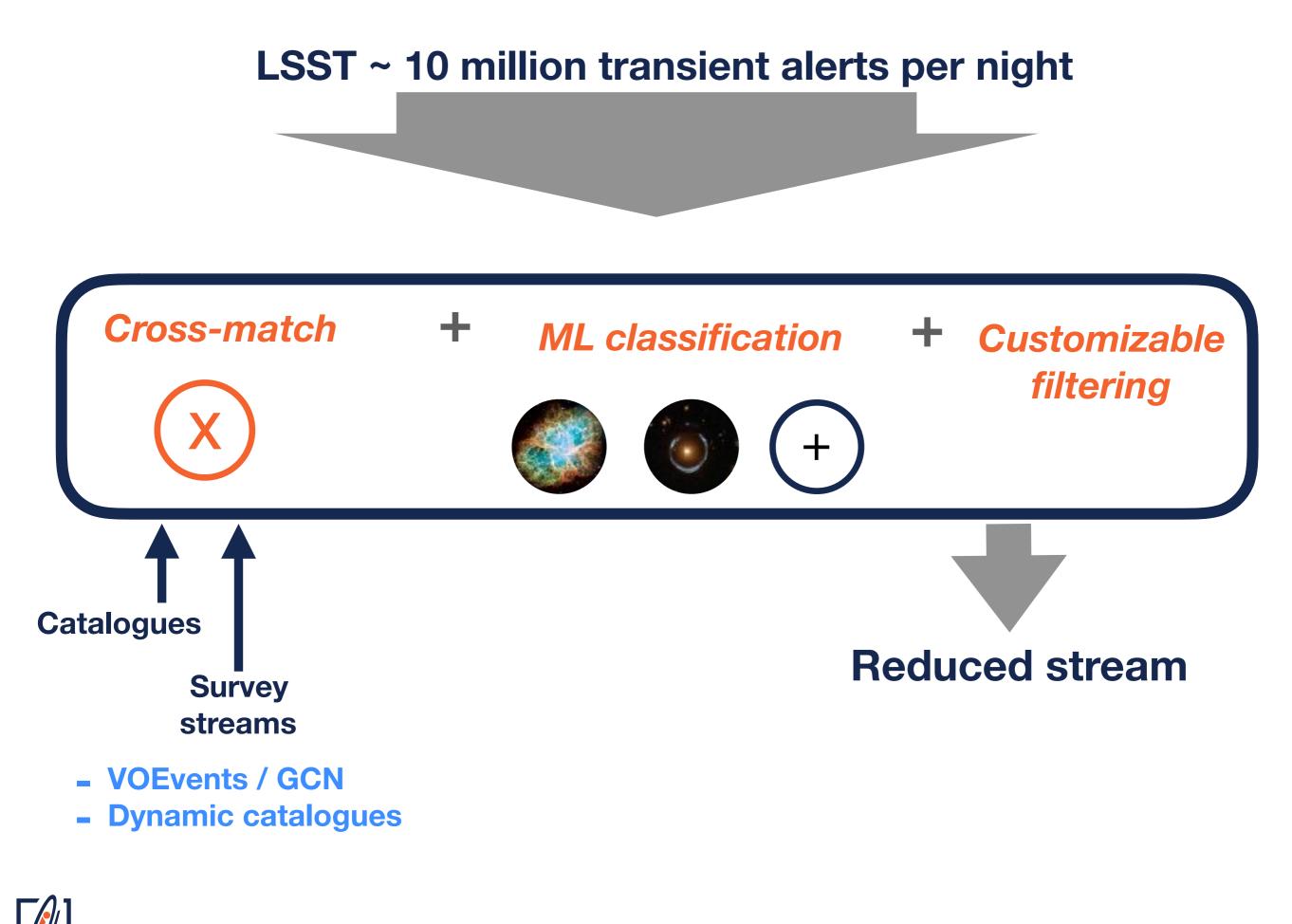




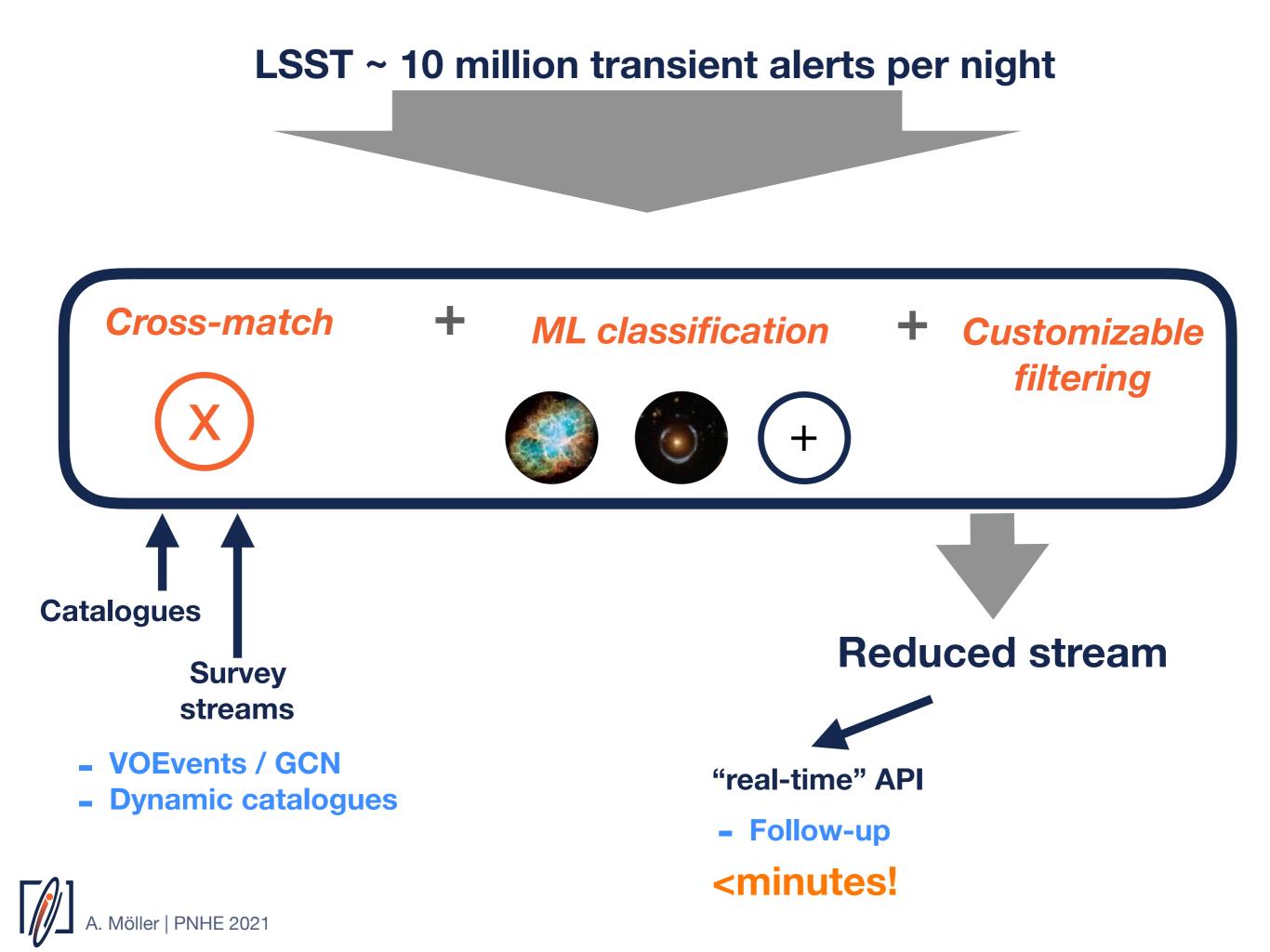


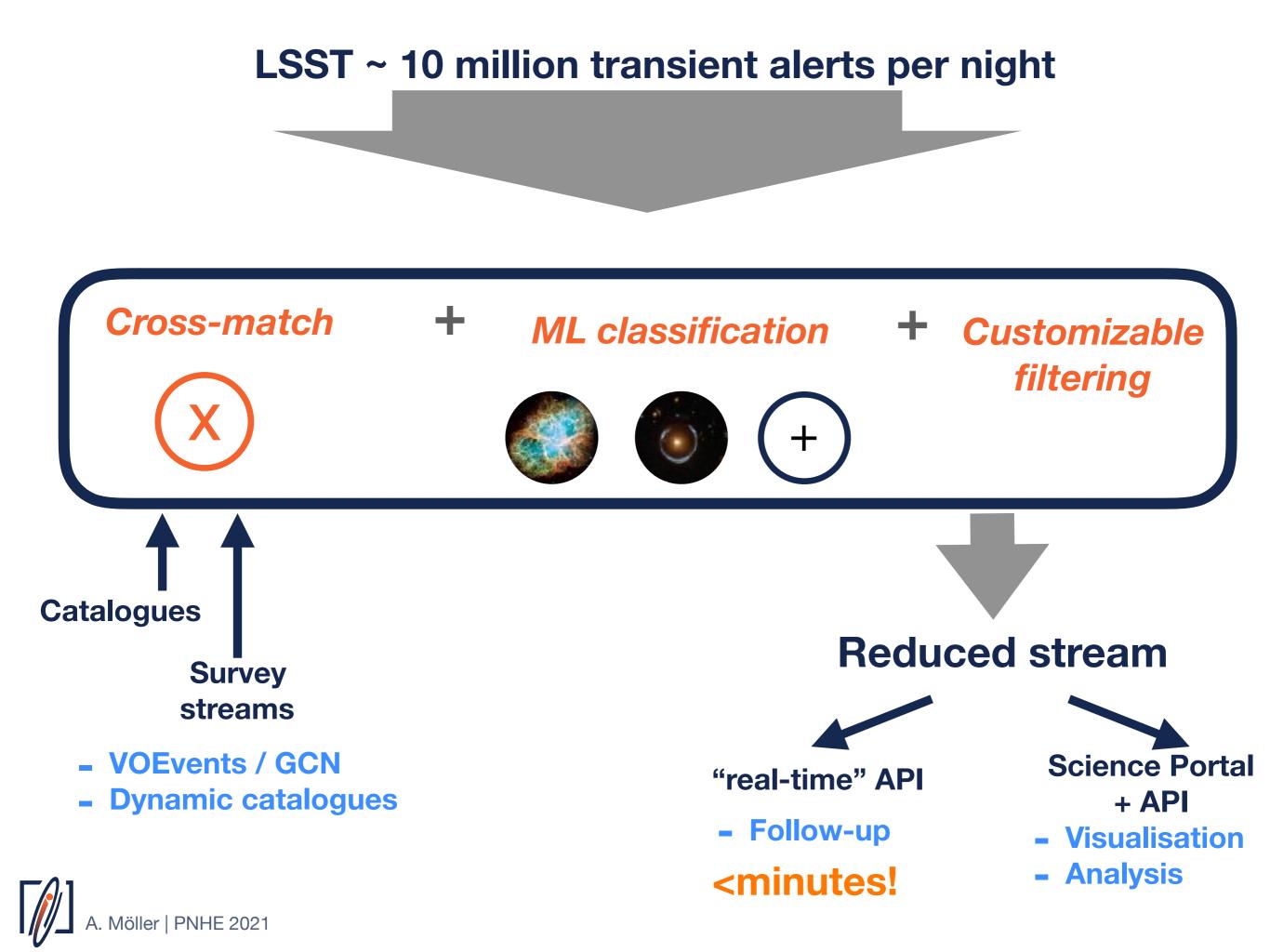


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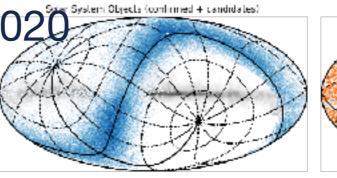
# **F//NK deployment with ZTF**

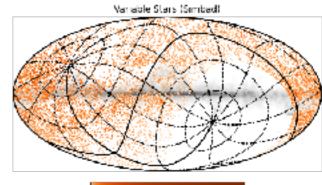
MoU with Zwicky Transient Facility (ZTF).

>110 million alerts collected since 2020

Reduced streams:

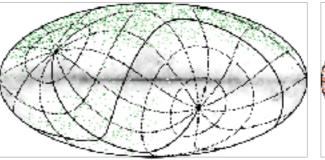
- Supernovae
- Microlensing
- Variable stars
- Solar System objects
- GRB afterglows
- Kilonovae (new)
- Early type la supernovae (new)

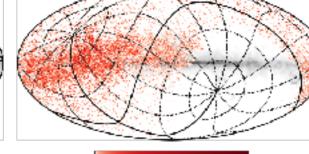


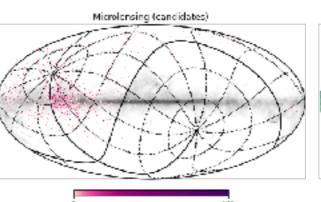


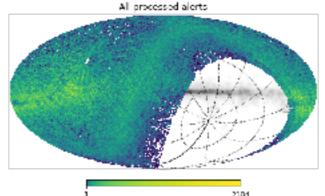












ZTF alerts November 2019 to June 2020



LSST stream: 10,000 alerts / 37 s

Cross-match Footprint w. survey stream: LIGO/Virgo



LSST stream: 10,000 alerts / 37 s

Cross-match

Footprint w. survey stream: LIGO/Virgo

Contextual: e.g. host-galaxy not a known object



Other multi-messenger/wavelength surveys



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### **Cross-match**

Footprint w. survey stream: LIGO/Virgo

Contextual: e.g. host-galaxy not a known object

Other multi-messenger/wavelength surveys

### **ML classification**



**Discard** other types of astrophysical events Algorithm for kilonovae **identification** 



LSST stream: 10,000 alerts / 37 s

### **Cross-match**

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Other multi-messenger/wavelength surveys

### **ML classification**



Discard other types of astrophysical events Algorithm for kilonovae identification

**Filtering** 

Select promising kilonovae candidates
Communicate All automatically done in <minutes

### Coordinating follow-up:

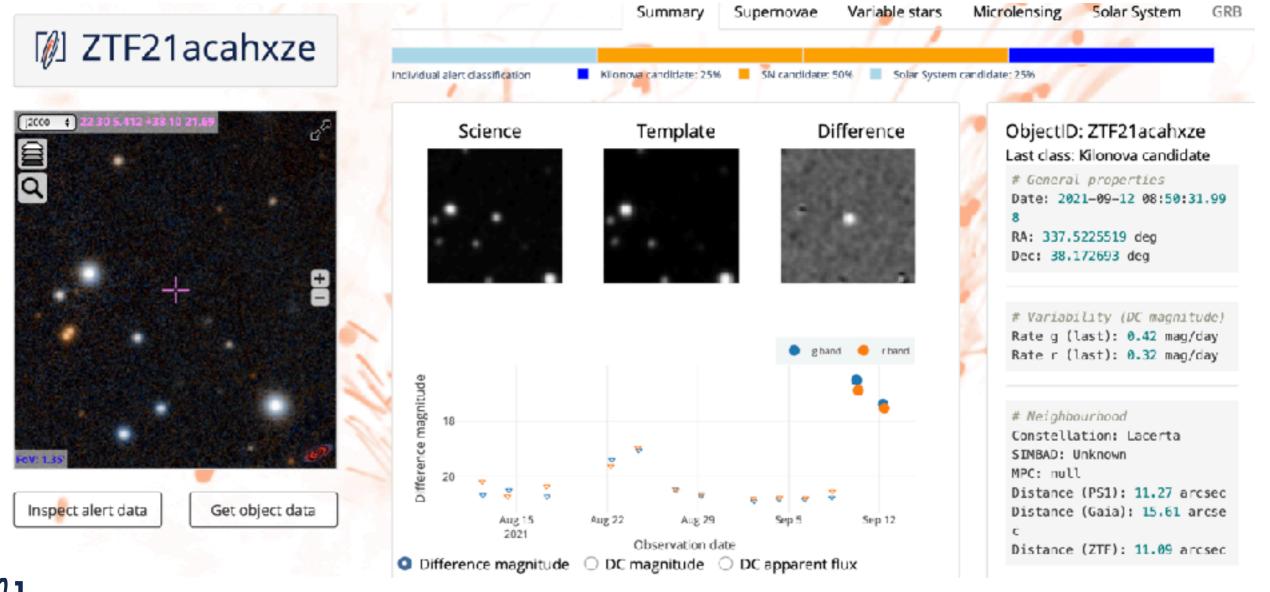
Möller | PNHE 2021

- ranked lists with promising candidates see Pierre

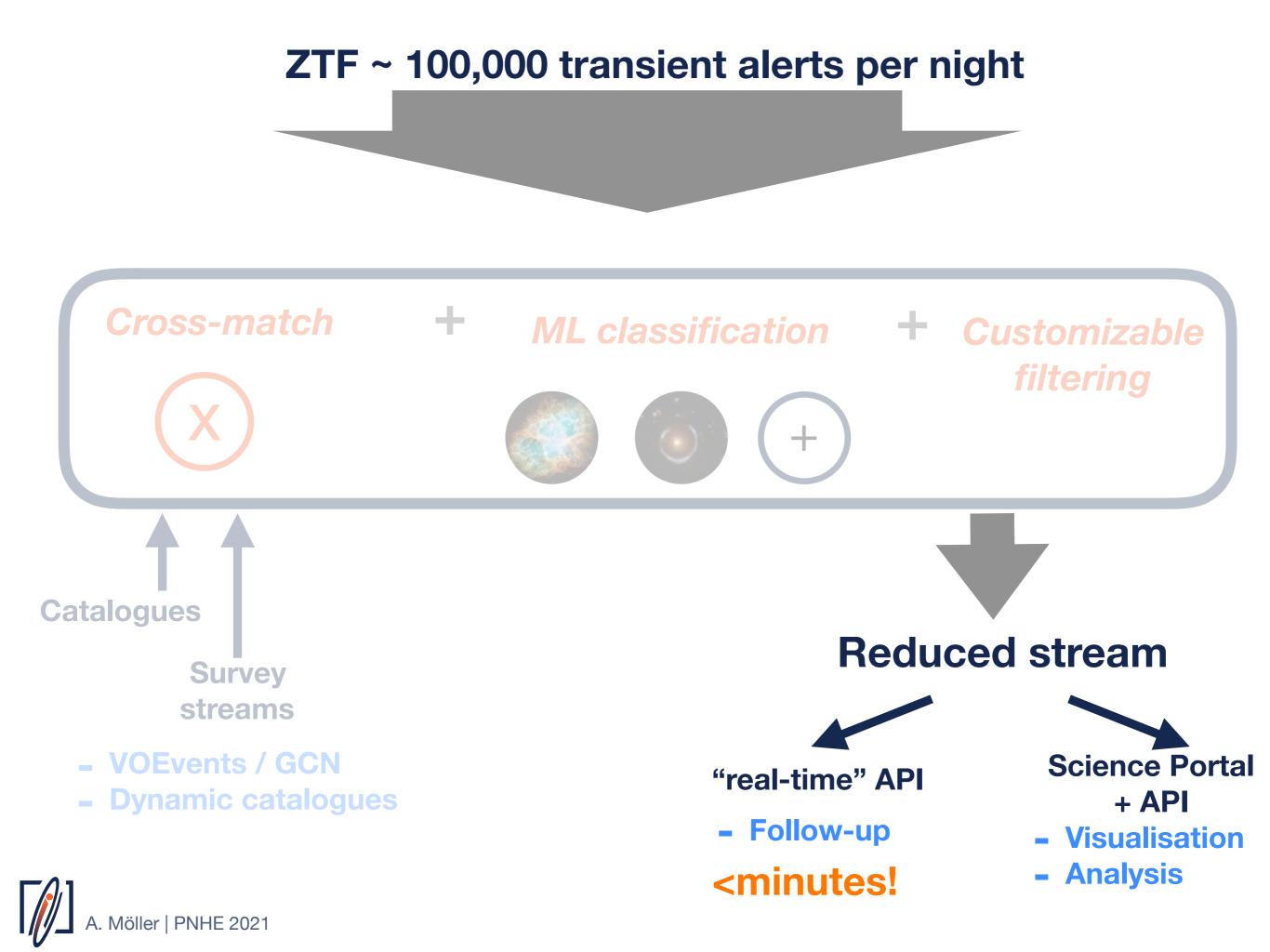
**Duverne's talk later** 

today!

- Downward stream < minutes</li>
- Ongoing collaboration with GRANDMA



With Juliette Vlieghe, Biswajit Biswas, Julien Peloton, GRANDMA team and many others!



Live demo science portal

https://fink-broker.org

# API

Fink Science portal 0.5 Explorer Xmatch API Info +

#### Fink API Summary of services

	oannie	ing of set fices		
HTTP Method URI		IURI	Action	Availability
	POST/GET	http://134.158.75.151:24000/api/v1/objects	Retrieve single object data from the Fink database	R 🥏
	POST/GET	http://134.158.75.151:24000/api/v1/explorer	Query the Fink alert database	Ø /
	POST/GET	http://134.158.75.151:24000/api/v1/latests	Get latest alerts by class	
	POST/GET	http://134.158.75.151:24000/api/v1/sso	Get Solar System Object data	R
	POST/GET	http://134.158.75.151:24000/api/v1/cutouts	Retrieve cutout data from the Fink database	EI.
	POST/GET	http://134.158.75.151:24000/api/v1/xmatch	Cross-match user-defined catalog with Fink alert data	×
	GET	http://134.158.75.151:24000/api/v1/classes	Display all Fink derived classification	
	GET	http://134.158.75.151:24000/api/v1/columns	Display all available alert fields and their type	

Retrieve object data

Query the database 0

Get latest alerts Get Solar System Objects Get Image data

ata Xmatch

#### Retrieve single object data

The list of arguments for retrieving object data can be found at http://134.158.75.151:24000/api/v1/objects.

In a unix shell, you would simply use

#### # Get data for ZTF19acnjwgm and save it in a CSV file

curl -H "Content-Type: application/json" -X POST -d '{"objectId":"ZTF19acnjwgm", "output-format":"csv"}' http://134.158.75.151:24000/api/v1/objects o ZTF19acnjwgm.csv

In python, you would use

import requests
import pandas as pd

# get data for ZTF19acnjwgm
r = requests.post(
 'http://134\_158\_75\_151.24200/api/v1/objects'



#### Fink case study: multi-messenger/wavelength astronomy

#### Goal

The goal of this notebook is to show how Fink can contribute to multi-messenger and multi-wavelength science. This notebook includes queries for science related to **Kionovae** and **gamma ray bursts (GRBs) counterparts**.

Real-time analysis for follow-up coordination will be better done through the Fink livestreams which processes and makes available the alert data within minutes of reception. Access through this notebook uses the API access which is not real-time based. We aggregate the data into our API and Science Portal database at the end of each observing night. API access can be useful to analyze last night and older alerts. For example in case of reprocessing analysis in the search for low signal-to-noise events in gravitational wave or other wavelength detectors.

#### What is behind?

This notebook will use:

- Fink cross-match with close by sources in the <u>SIMBAD (Wenger et al. 2020)</u> catalogue (within 1") using <u>xmatch service</u> provided by the Centre de Données astronomiques de Strasbourg (CDS).
- Fink's machine learning (ML) classification scores trained with (Muthukrishna et al. 2019) and (Stachie et al. 2019):
  - Random Forest to classify Kilonovae (Biswas et al. in prep)
  - SuperNNova (Möller & de Boissière 2019) to classify SNe at all light-curve epochs
  - Random Forest (Leoni et al. in prep) and (Ishida et al. 2019b) to classify early (pre-max) SN candidates
- Fink's extra features such as color g-r and rate when possible.

#### Environment set up

To run this notebook, you need to import the following libraries (already installed in colab):

#### In [1]: import requests

import pandas as pd
import numpy as np

import matplotlib.pyplot as plt import seaborn as sns sns.set\_context('talk')



- Discussions and work with teams from:
  - SVOM, GRANDMA, CTA, Integral, KM3NET, ...
  - Work on GRB (Fermi/SWIFT-ZTF) to prepare SVOM-LSST
  - GRANDMA: accepted CFHT observing proposal for Kilonovae search (2021A Coleiro+) using Fink's filtered ZTF alert stream.





VRO LSST starting end 2023! unveiling the Southern Sky like never before!



Rubin optical time-domain data will be a great asset for multimessenger/wavelength studies!

How to unravel it?

- Fink is already processing ZTF data stream (MoU 2020).
- First science modules deployed: SNe, GRB, microlensing, ...
- Automatic filtered streams <minutes for follow-up coordination</li>
- Analysis capabilities with the web portal and API (python callable)
- for VRO deployment at CC-in2p3
- We want to connect to new teams, collaborations and facilities!

Möller, Peloton, Ishida et al. 2021 MNRAS, arXiv: 2009.10185 <u>https://fink-broker.org</u> <u>contact@fink-broker.org</u>