

Liouville gravity as dilaton gravity

Thomas Mertens

Ghent University

Based on [arXiv:2006.07072](https://arxiv.org/abs/2006.07072) with G.J. Turiaci
[arXiv:2109.07770](https://arxiv.org/abs/2109.07770) with Y. Fan



Introduction

Goal: study lower-dimensional models of quantum gravity:

Introduction

Goal: study lower-dimensional models of quantum gravity:

Interesting class = 2d dilaton gravity models: [Grumiller-Kummer-Vasilevich '02](#)

$$S = \frac{1}{16\pi G} \int d^2x \sqrt{-g} (\Phi R + W(\Phi)) + \frac{1}{8\pi G} \oint d\tau \sqrt{-\gamma} \Phi_{bdy} K$$

Introduction

Goal: study lower-dimensional models of quantum gravity:

Interesting class = 2d dilaton gravity models: [Grumiller-Kummer-Vasilevich '02](#)

$$S = \frac{1}{16\pi G} \int d^2x \sqrt{-g} (\Phi R + W(\Phi)) + \frac{1}{8\pi G} \oint d\tau \sqrt{-\gamma} \Phi_{bdy} K$$

JT gravity = specify to $W(\Phi) = -\Lambda\Phi$

$$S_{JT} = \frac{1}{16\pi G} \int d^2x \sqrt{-g} \Phi (R - \Lambda) + \frac{1}{8\pi G} \oint d\tau \sqrt{-\gamma} \Phi_{bdy} K$$

[Teitelboim '83](#), [Jackiw '85](#)

Negative cosmological constant $\Lambda = -2 \rightarrow \text{AdS}_2$ space

Introduction

Goal: study lower-dimensional models of quantum gravity:

Interesting class = 2d dilaton gravity models: [Grumiller-Kummer-Vasilevich '02](#)

$$S = \frac{1}{16\pi G} \int d^2x \sqrt{-g} (\Phi R + W(\Phi)) + \frac{1}{8\pi G} \oint d\tau \sqrt{-\gamma} \Phi_{bdy} K$$

JT gravity = specify to $W(\Phi) = -\Lambda\Phi$

$$S_{JT} = \frac{1}{16\pi G} \int d^2x \sqrt{-g} \Phi (R - \Lambda) + \frac{1}{8\pi G} \oint d\tau \sqrt{-\gamma} \Phi_{bdy} K$$

[Teitelboim '83](#), [Jackiw '85](#)

Negative cosmological constant $\Lambda = -2 \rightarrow \text{AdS}_2$ space

Features:

- ▶ Appears as near-horizon theory of near-extremal higher-dimensional black holes
- ▶ Describes low-energy sector of all (known) SYK-like models
- ▶ With suitable boundary conditions, reduces to Schwarzian action [Almheiri-Polchinski '15](#), [Jensen '16](#), [Maldacena-Stanford-Yang '16](#), [Engelsöy-TM-Verlinde '16](#)
- ▶ **Exactly solvable to large extent**

Review - JT disk partition function

Exact quantum solution of amplitudes: $\int \frac{[\mathcal{D}g_{\mu\nu}][\mathcal{D}\Phi]}{\text{Vol}(\text{Diff})} \dots e^{-S_{\text{JT}}}$

Review - JT disk partition function

Exact quantum solution of amplitudes: $\int \frac{[\mathcal{D}g_{\mu\nu}][\mathcal{D}\Phi]}{\text{Vol}(\text{Diff})} \dots e^{-S_{\text{JT}}}$

Disk Partition function Cotler et al. '16, Maldacena-Stanford '16, Stanford-Witten '17

$$Z(\beta) = \text{Disk} = \int_0^{+\infty} dk (k \sinh 2\pi k) e^{-\beta k^2}$$

Energy variable $E = k^2$, $\rho(E) = \sinh 2\pi\sqrt{E}$

AdS₂ Euclidean holographic boundary with length = β , filled in by path integrating over all metrics and dilatons

Review - JT disk partition function

Exact quantum solution of amplitudes: $\int \frac{[\mathcal{D}g_{\mu\nu}][\mathcal{D}\Phi]}{\text{Vol}(\text{Diff})} \dots e^{-S_{\text{JT}}}$

Disk Partition function Cotler et al. '16, Maldacena-Stanford '16, Stanford-Witten '17

$$Z(\beta) = \text{Disk} = \int_0^{+\infty} dk (k \sinh 2\pi k) e^{-\beta k^2}$$

Energy variable $E = k^2$, $\rho(E) = \sinh 2\pi\sqrt{E}$

AdS₂ Euclidean holographic boundary with length = β , filled in by path integrating over all metrics and dilatons

Thermodynamic limit (saddle):

$$\rho(E) \sim e^{2\pi\sqrt{E}} \Rightarrow \sqrt{E} = \frac{\pi}{\beta}$$


Matches with black hole first law (mass vs Hawking temperature) for classical AdS₂ JT black hole solution: Almheiri-Polchinski '15

$$ds^2 = -(r^2 - r_h^2)dt^2 + \frac{dr^2}{r^2 - r_h^2}, \quad \Phi(r) = r$$

Review - JT boundary two-point function

Boundary two-point function (of some boundary CFT matter operators $\mathcal{O}^h(\tau)$ coupled to gravity):


Bagrets-Altland-Kamenev '16, '17, TM-Turiaci-Verlinde '17, '18, Blommaert-TM-Verschelde '18, Yang '18, Kitaev-Suh '18, '19, Iliesiu-Pufu-Verlinde-Wang '19 ...

$$\langle \mathcal{O}^h(\tau_1) \mathcal{O}^h(\tau_2) \rangle_\beta = \tau_2 \text{---} \overset{h}{\text{---}} \tau_1 =$$


Review - JT boundary two-point function

Boundary two-point function (of some boundary CFT matter operators $\mathcal{O}^h(\tau)$ coupled to gravity):

Bagrets-Altland-Kamenev '16, '17, TM-Turiaci-Verlinde '17, '18, Blommaert-TM-Verschelde '18, Yang '18, Kitaev-Suh '18, '19, Iliesiu-Pufu-Verlinde-Wang '19 ...


$$\langle \mathcal{O}^h(\tau_1) \mathcal{O}^h(\tau_2) \rangle_\beta = \tau_2 \bullet \text{---} \overset{h}{\text{---}} \bullet \tau_1 =$$

$$\int_0^{+\infty} dk_1 (k_1 \sinh 2\pi k_1) \int_0^{+\infty} dk_2 (k_2 \sinh 2\pi k_2) e^{-\tau k_1^2 - (\beta - \tau) k_2^2} \underbrace{\frac{\Gamma(h \pm ik_1 \pm ik_2)}{\Gamma(2h)}}_{\text{vertex functions}}$$

$(\tau = \tau_2 - \tau_1)$
 \pm sign convention = take product of all cases

Review - JT boundary two-point function

Boundary two-point function (of some boundary CFT matter operators $\mathcal{O}^h(\tau)$ coupled to gravity):

Bagrets-Altland-Kamenev '16, '17, TM-Turiaci-Verlinde '17, '18, Blommaert-TM-Verschelde '18, Yang '18, Kitaev-Suh '18, '19, Iliesiu-Pufu-Verlinde-Wang '19 ...

$$\langle \mathcal{O}^h(\tau_1) \mathcal{O}^h(\tau_2) \rangle_\beta = \tau_2 \bullet \text{---} \overset{h}{\text{---}} \bullet \tau_1 =$$

$$\int_0^{+\infty} dk_1 (k_1 \sinh 2\pi k_1) \int_0^{+\infty} dk_2 (k_2 \sinh 2\pi k_2) e^{-\tau k_1^2 - (\beta - \tau) k_2^2} \underbrace{\frac{\Gamma(h \pm ik_1 \pm ik_2)}{\Gamma(2h)}}_{\text{vertex functions}}$$

$(\tau = \tau_2 - \tau_1)$

\pm sign convention = take product of all cases

Generalization to multi-point functions and OTOCs possible

TM-Turiaci-Verlinde '17, '18

JT gravity: summary

Importance of these calculations for gravity:

- ▶ **Boundary correlators:** interpretation in terms of gravitational physics (shockwaves, complexity = volume, bulk entanglement entropy, spectral occupation and Unruh heat bath physics for bulk operators . . .) [Maldacena-Stanford-Yang '16](#), [Lam-TM-Turiaci-Verlinde '18](#), [Yang '18](#), [TM '19](#), [Blommaert-TM-Verschelde '19](#), '20 . . .

JT gravity: summary

Importance of these calculations for gravity:

- ▶ **Boundary correlators:** interpretation in terms of gravitational physics (shockwaves, complexity = volume, bulk entanglement entropy, spectral occupation and Unruh heat bath physics for bulk operators . . .) [Maldacena-Stanford-Yang '16](#), [Lam-TM-Turiaci-Verlinde '18](#), [Yang '18](#), [TM '19](#), [Blommaert-TM-Verschelde '19](#), '20 . . .
- ▶ **Higher genus and multi-boundary amplitudes:** important to understand very-late time correlators, explain Page curve with replica wormholes [Saad-Shenker-Stanford '19](#), [Saad '19](#), [Almheiri-Hartman-Maldacena-Shaghoulian-Tajdini '19](#), [Penington-Shenker-Stanford-Yang '19](#)
(not in this talk)

JT gravity: summary

Importance of these calculations for gravity:

- ▶ **Boundary correlators:** interpretation in terms of gravitational physics (shockwaves, complexity = volume, bulk entanglement entropy, spectral occupation and Unruh heat bath physics for bulk operators . . .) [Maldacena-Stanford-Yang '16](#),
[Lam-TM-Turiaci-Verlinde '18](#), [Yang '18](#), [TM '19](#), [Blommaert-TM-Verschelde '19](#), '20 . . .
- ▶ **Higher genus and multi-boundary amplitudes:** important to understand very-late time correlators, explain Page curve with replica wormholes [Saad-Shenker-Stanford '19](#), [Saad '19](#),
[Almheiri-Hartman-Maldacena-Shaghoulian-Tajdini '19](#), [Penington-Shenker-Stanford-Yang '19](#)
(not in this talk)

→ Would be interesting to extend our class of solvable models to find out how generic these lessons are

JT gravity: summary

Importance of these calculations for gravity:

- ▶ **Boundary correlators:** interpretation in terms of gravitational physics (shockwaves, complexity = volume, bulk entanglement entropy, spectral occupation and Unruh heat bath physics for bulk operators ...) [Maldacena-Stanford-Yang '16](#), [Lam-TM-Turiaci-Verlinde '18](#), [Yang '18](#), [TM '19](#), [Blommaert-TM-Verschelde '19](#), '20 ...
- ▶ **Higher genus and multi-boundary amplitudes:** important to understand very-late time correlators, explain Page curve with replica wormholes [Saad-Shenker-Stanford '19](#), [Saad '19](#), [Almheiri-Hartman-Maldacena-Shaghoulian-Tajdini '19](#), [Penington-Shenker-Stanford-Yang '19](#)
(not in this talk)

→ Would be interesting to extend our class of solvable models to find out how generic these lessons are

GOAL: discuss 2d Liouville gravity amplitudes in same language

→ interpret as specific model of **dilaton gravity**

Liouville gravity: definition

Liouville gravity = non-critical string = 2d matter CFT coupled to gravity, or critical string with a 2d Liouville + matter + ghost CFT

Polyakov '81, David '88, Distler-Kawai '89 . . .

Total action: $S_L + S_M + S_{gh}$

with conformal anomaly constraint $c_M + c_L + c_{gh} = 0$

Liouville gravity: definition

Liouville gravity = non-critical string = 2d matter CFT coupled to gravity, or critical string with a 2d Liouville + matter + ghost CFT

Polyakov '81, David '88, Distler-Kawai '89 . . .

Total action: $S_L + S_M + S_{gh}$

with conformal anomaly constraint $c_M + c_L + c_{gh} = 0$

► Liouville action: $S_L = \frac{1}{4\pi} \int_{\Sigma} \left[(\hat{\nabla}\phi)^2 + Q\hat{R}\phi + 4\pi\mu e^{2b\phi} \right]$

$$Q = b + b^{-1}, \quad c_L = 1 + 6Q^2 > 25$$

Arises from conformal factor $g_{\mu\nu} = e^{2b\phi} \hat{g}_{\mu\nu}$ of 2d gravity

Liouville gravity: definition

Liouville gravity = non-critical string = 2d matter CFT coupled to gravity, or critical string with a 2d Liouville + matter + ghost CFT

Polyakov '81, David '88, Distler-Kawai '89 . . .

Total action: $S_L + S_M + S_{gh}$

with conformal anomaly constraint $c_M + c_L + c_{gh} = 0$

► Liouville action: $S_L = \frac{1}{4\pi} \int_{\Sigma} \left[(\hat{\nabla}\phi)^2 + Q\hat{R}\phi + 4\pi\mu e^{2b\phi} \right]$

$$Q = b + b^{-1}, \quad c_L = 1 + 6Q^2 > 25$$

Arises from conformal factor $g_{\mu\nu} = e^{2b\phi} \hat{g}_{\mu\nu}$ of 2d gravity

► For most of talk: $S_M =$ arbitrary CFT with $c_M < 1$

particular choice to keep in mind:

→ timelike Liouville CFT (Liouville with $b \rightarrow ib, \phi \rightarrow -i\chi$)

Liouville gravity: definition

Liouville gravity = non-critical string = 2d matter CFT coupled to gravity, or critical string with a 2d Liouville + matter + ghost CFT

Polyakov '81, David '88, Distler-Kawai '89 . . .

Total action: $S_L + S_M + S_{gh}$

with conformal anomaly constraint $c_M + c_L + c_{gh} = 0$

- ▶ Liouville action: $S_L = \frac{1}{4\pi} \int_{\Sigma} \left[(\hat{\nabla}\phi)^2 + Q\hat{R}\phi + 4\pi\mu e^{2b\phi} \right]$
 $Q = b + b^{-1}$, $c_L = 1 + 6Q^2 > 25$
Arises from conformal factor $g_{\mu\nu} = e^{2b\phi} \hat{g}_{\mu\nu}$ of 2d gravity
- ▶ For most of talk: $S_M =$ arbitrary CFT with $c_M < 1$
particular choice to keep in mind:
→ timelike Liouville CFT (Liouville with $b \rightarrow ib$, $\phi \rightarrow -i\chi$)
- ▶ S_{gh} is usual bc -ghost theory with $c_{gh} = -26$

Liouville gravity: Fixed-length boundaries

Phrase in same language as JT gravity:

→ Reinterpret worldsheet topology as the 2d Euclidean spacetime of a gravitational model

Liouville gravity: Fixed-length boundaries

Phrase in same language as JT gravity:

→ Reinterpret worldsheet topology as the 2d Euclidean spacetime of a gravitational model

Interested in holography → 2d manifold with boundary of fixed length β ; in this talk only disk topology

Liouville gravity: Fixed-length boundaries

Phrase in same language as JT gravity:

→ Reinterpret worldsheet topology as the 2d Euclidean spacetime of a gravitational model

Interested in holography → 2d manifold with boundary of fixed length β ; in this talk only disk topology

Boundary conditions:

- ▶ When viewing the theory as 2d quantum gravity, Liouville field related to metric $g_{\mu\nu}$: $ds^2 = e^{2b\phi} dzd\bar{z}$
⇒ Boundary length $\equiv \ell = \oint e^{b\phi}$ is fixed
= Fourier transform of FZZT-brane boundary (Neumann-like)

Liouville gravity: Fixed-length boundaries

Phrase in same language as JT gravity:

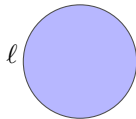
→ Reinterpret worldsheet topology as the 2d Euclidean spacetime of a gravitational model

Interested in holography → 2d manifold with boundary of fixed length β ; in this talk only disk topology

Boundary conditions:

- ▶ When viewing the theory as 2d quantum gravity, Liouville field related to metric $g_{\mu\nu}$: $ds^2 = e^{2b\phi} dzd\bar{z}$
⇒ Boundary length $\equiv \ell = \oint e^{b\phi}$ is fixed
= Fourier transform of FZZT-brane boundary (Neumann-like)
- ▶ Matter + ghost: vacuum brane boundary (Dirichlet)

Disk partition function



Disk partition function



Fixed length amplitude:

$$Z(\ell) \sim \int_0^\infty ds \sinh(2\pi bs) \sinh\left(\frac{2\pi s}{b}\right) e^{-\ell \kappa \cosh(2\pi bs)} \quad \kappa \equiv \frac{\sqrt{\mu}}{\sqrt{\sin \pi b^2}}$$

Disk partition function



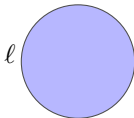
Fixed length amplitude:

$$Z(\ell) \sim \int_0^\infty ds \sinh(2\pi bs) \sinh\left(\frac{2\pi s}{b}\right) e^{-\ell \kappa \cosh(2\pi bs)} \quad \kappa \equiv \frac{\sqrt{\mu}}{\sqrt{\sin \pi b^2}}$$

► **JT limit:** $b \rightarrow 0$, $\ell \sim \frac{\ell_{\text{JT}}}{\kappa b^4} \rightarrow +\infty$ Saad-Shenker-Stanford '19, TM-Turiaci '20

$$Z(\ell) \rightarrow \int_0^\infty dk (k \sinh 2\pi k) e^{-\ell_{\text{JT}} k^2}, \quad s = bk$$

Disk partition function



Fixed length amplitude:

$$Z(\ell) \sim \int_0^\infty ds \sinh(2\pi bs) \sinh\left(\frac{2\pi s}{b}\right) e^{-\ell \kappa \cosh(2\pi bs)} \quad \kappa \equiv \frac{\sqrt{\mu}}{\sqrt{\sin \pi b^2}}$$

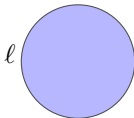
► **JT limit:** $b \rightarrow 0$, $\ell \sim \frac{\ell_{\text{JT}}}{\kappa b^4} \rightarrow +\infty$ Saad-Shenker-Stanford '19, TM-Turiaci '20

$$Z(\ell) \rightarrow \int_0^\infty dk (k \sinh 2\pi k) e^{-\ell_{\text{JT}} k^2}, \quad s = bk$$

► **Back to full Liouville gravity**, with interpretation of $\ell = \beta$:

$$Z(\beta) \sim \int_\kappa^\infty dE e^{-\beta E} \rho(E), \quad \rho(E) = \sinh\left(\frac{1}{b^2} \operatorname{arccosh} \frac{E}{\kappa}\right)$$

Disk partition function



Fixed length amplitude:

$$Z(\ell) \sim \int_0^\infty ds \sinh(2\pi bs) \sinh\left(\frac{2\pi s}{b}\right) e^{-\ell \kappa \cosh(2\pi bs)} \quad \kappa \equiv \frac{\sqrt{\mu}}{\sqrt{\sin \pi b^2}}$$

- ▶ **JT limit:** $b \rightarrow 0$, $\ell \sim \frac{\ell_{\text{JT}}}{\kappa b^4} \rightarrow +\infty$ Saad-Shenker-Stanford '19, TM-Turiaci '20

$$Z(\ell) \rightarrow \int_0^\infty dk (k \sinh 2\pi k) e^{-\ell_{\text{JT}} k^2}, \quad s = bk$$

- ▶ **Back to full Liouville gravity**, with interpretation of $\ell = \beta$:

$$Z(\beta) \sim \int_\kappa^\infty dE e^{-\beta E} \rho(E), \quad \rho(E) = \sinh\left(\frac{1}{b^2} \operatorname{arccosh} \frac{E}{\kappa}\right)$$

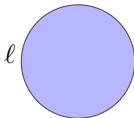
Thermodynamic limit (saddle):

$$\sqrt{E^2 - \kappa^2} = \frac{1}{b^2 \beta}$$

IR: $E = \kappa + E_{\text{JT}} \Rightarrow \sqrt{E_{\text{JT}}} \sim \beta^{-1}$, the JT black hole first law

UV: $E \sim \beta^{-1}$

Disk partition function



Fixed length amplitude:

$$Z(\ell) \sim \int_0^\infty ds \sinh(2\pi bs) \sinh\left(\frac{2\pi s}{b}\right) e^{-\ell \kappa \cosh(2\pi bs)} \quad \kappa \equiv \frac{\sqrt{\mu}}{\sqrt{\sin \pi b^2}}$$

- ▶ **JT limit:** $b \rightarrow 0$, $\ell \sim \frac{\ell_{\text{JT}}}{\kappa b^4} \rightarrow +\infty$ Saad-Shenker-Stanford '19, TM-Turiaci '20

$$Z(\ell) \rightarrow \int_0^\infty dk (k \sinh 2\pi k) e^{-\ell_{\text{JT}} k^2}, \quad s = bk$$

- ▶ **Back to full Liouville gravity**, with interpretation of $\ell = \beta$:

$$Z(\beta) \sim \int_\kappa^\infty dE e^{-\beta E} \rho(E), \quad \rho(E) = \sinh\left(\frac{1}{b^2} \operatorname{arccosh} \frac{E}{\kappa}\right)$$

Thermodynamic limit (saddle):

$$\sqrt{E^2 - \kappa^2} = \frac{1}{b^2 \beta}$$

IR: $E = \kappa + E_{\text{JT}} \Rightarrow \sqrt{E_{\text{JT}}} \sim \beta^{-1}$, the JT black hole first law

UV: $E \sim \beta^{-1}$

\rightarrow holographic UV/IR connection: not aAdS like JT gravity

Boundary two-point function

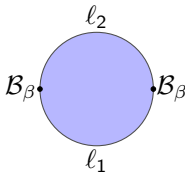
Boundary tachyon vertex operator: $\mathcal{B}_\beta = c \Phi_M e^{\beta\phi}$

→ Matter primary Φ_M gravitationally dressed by Liouville operator $e^{\beta\phi}$, and gauge-fixed

Boundary two-point function

Boundary tachyon vertex operator: $\mathcal{B}_\beta = c \Phi_M e^{\beta\phi}$

→ Matter primary Φ_M gravitationally dressed by Liouville operator $e^{\beta\phi}$, and gauge-fixed

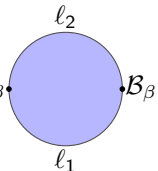
$$\text{Boundary 2-pt function} = \langle \mathcal{B}_\beta \mathcal{B}_\beta \rangle_{\ell_1, \ell_2} = \mathcal{B}_\beta \bullet \text{---} \text{---} \text{---} \bullet \mathcal{B}_\beta$$


Boundary two-point function

Boundary tachyon vertex operator: $\mathcal{B}_\beta = c \Phi_M e^{\beta\phi}$

→ Matter primary Φ_M gravitationally dressed by Liouville operator $e^{\beta\phi}$, and gauge-fixed

$$\text{Boundary 2-pt function} = \langle \mathcal{B}_\beta \mathcal{B}_\beta \rangle_{\ell_1, \ell_2} = \mathcal{B}_\beta \bullet$$



$$= \int_0^{+\infty} ds_1 ds_2 \rho(s_1) \rho(s_2) e^{-\ell_1 \kappa \cosh 2\pi b s_1} e^{-\ell_2 \kappa \cosh 2\pi b s_2} \frac{S_b(\beta_M \pm i s_1 \pm i s_2)}{S_b(2\beta_M)}$$

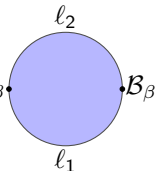
where $\rho(s) = \sinh(2\pi b s) \sinh\left(\frac{2\pi}{b}s\right)$ and $\beta_M = b - \beta$ TM-Turiaci '20

Boundary two-point function

Boundary tachyon vertex operator: $\mathcal{B}_\beta = c \Phi_M e^{\beta\phi}$

→ Matter primary Φ_M gravitationally dressed by Liouville operator $e^{\beta\phi}$, and gauge-fixed

$$\text{Boundary 2-pt function} = \langle \mathcal{B}_\beta \mathcal{B}_\beta \rangle_{\ell_1, \ell_2} = \mathcal{B}_\beta \bullet \text{---} \text{---} \bullet \mathcal{B}_\beta$$



$$= \int_0^{+\infty} ds_1 ds_2 \rho(s_1) \rho(s_2) e^{-\ell_1 \kappa \cosh 2\pi b s_1} e^{-\ell_2 \kappa \cosh 2\pi b s_2} \frac{S_b(\beta_M \pm i s_1 \pm i s_2)}{S_b(2\beta_M)}$$

where $\rho(s) = \sinh(2\pi b s) \sinh\left(\frac{2\pi}{b}s\right)$ and $\beta_M = b - \beta$ TM-Turiaci '20

Technicality: Liouville piece = $\langle e^{\beta_1 \phi(x)} e^{\beta_2 \phi(0)} \rangle \sim \frac{\delta(\beta_1 - \beta_2)}{|x|^{2\Delta_\beta}} \mathcal{A}_{\beta_1, \beta_2}$

→ $\delta(0)$ cancelled by modding out CKG

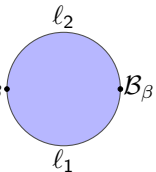
→ Matter + ghost cancels out worldsheet coordinate dependence

Boundary two-point function

Boundary tachyon vertex operator: $\mathcal{B}_\beta = c \Phi_M e^{\beta\phi}$

→ Matter primary Φ_M gravitationally dressed by Liouville operator $e^{\beta\phi}$, and gauge-fixed

$$\text{Boundary 2-pt function} = \langle \mathcal{B}_\beta \mathcal{B}_\beta \rangle_{\ell_1, \ell_2} = \mathcal{B}_\beta \bullet$$



$$= \int_0^{+\infty} ds_1 ds_2 \rho(s_1) \rho(s_2) e^{-\ell_1 \kappa \cosh 2\pi b s_1} e^{-\ell_2 \kappa \cosh 2\pi b s_2} \frac{S_b(\beta_M \pm i s_1 \pm i s_2)}{S_b(2\beta_M)}$$

where $\rho(s) = \sinh(2\pi b s) \sinh\left(\frac{2\pi}{b}s\right)$ and $\beta_M = b - \beta$ TM-Turiaci '20

Technicality: Liouville piece = $\langle e^{\beta_1 \phi(x)} e^{\beta_2 \phi(0)} \rangle \sim \frac{\delta(\beta_1 - \beta_2)}{|x|^{2\Delta_\beta}} \mathcal{A}_{\beta_1, \beta_2}$

→ $\delta(0)$ cancelled by modding out CKG

→ Matter + ghost cancels out worldsheet coordinate dependence

JT limit ($b \rightarrow 0$, $\beta_M = bh$): $S_b(bx) \sim \Gamma(x)$:

$$\int_0^{+\infty} dk_1 dk_2 (k_1 \sinh 2\pi k_1) (k_2 \sinh 2\pi k_2) e^{-k_1^2 \ell_{JT1}} e^{-k_2^2 \ell_{JT2}} \frac{\Gamma(h \pm i k_1 \pm i k_2)}{\Gamma(2h)}$$

Group theory interpretation of JT gravity

Example: JT boundary two-point function:

$$\int dk_1 (k_1 \sinh 2\pi k_1) \int dk_2 (k_2 \sinh 2\pi k_2) e^{-\tau k_1^2 - (\beta - \tau) k_2^2} \frac{\Gamma(h \pm i k_1 \pm i k_2)}{\Gamma(2h)}$$

Group theory interpretation of JT gravity

Example: JT boundary two-point function:

$$\int dk_1 (k_1 \sinh 2\pi k_1) \int dk_2 (k_2 \sinh 2\pi k_2) e^{-\tau k_1^2 - (\beta - \tau) k_2^2} \frac{\Gamma(h \pm i k_1 \pm i k_2)}{\Gamma(2h)}$$

→ Measure and energies match with Plancherel measure and Casimir of continuous irreps of (modification of) $SL(2, \mathbb{R})$

Group theory interpretation of JT gravity

Example: JT boundary two-point function:

$$\int dk_1 (k_1 \sinh 2\pi k_1) \int dk_2 (k_2 \sinh 2\pi k_2) e^{-\tau k_1^2 - (\beta - \tau) k_2^2} \frac{\Gamma(h \pm i k_1 \pm i k_2)}{\Gamma(2h)}$$

→ Measure and energies match with Plancherel measure and Casimir of continuous irreps of (modification of) $SL(2, \mathbb{R})$

→ Vertex function is 3j-symbol² with two such continuous irreps (states) and one discrete lowest weight irrep (operator):

$$\int dg R_{j_1, m_1 n_1}(g) R_{j_2, m_2 n_2}(g) R_{j_3, m_3 n_3}(g) = \begin{pmatrix} j_1 & j_2 & j_3 \\ m_1 & m_2 & m_3 \end{pmatrix} \begin{pmatrix} j_1 & j_2 & j_3 \\ n_1 & n_2 & n_3 \end{pmatrix}$$

where $R_{j, mn}(g)$ are group representation matrices = $\langle j, m | g | j, n \rangle$
E.g. Wigner D-functions for $SU(2)$

Group theory interpretation of JT gravity

Example: JT boundary two-point function:

$$\int dk_1 (k_1 \sinh 2\pi k_1) \int dk_2 (k_2 \sinh 2\pi k_2) e^{-\tau k_1^2 - (\beta - \tau) k_2^2} \frac{\Gamma(h \pm i k_1 \pm i k_2)}{\Gamma(2h)}$$

→ **Measure** and **energies** match with Plancherel measure and Casimir of continuous irreps of (modification of) $SL(2, \mathbb{R})$

→ **Vertex function** is 3j-symbol² with two such continuous irreps (states) and one discrete lowest weight irrep (operator):

$$\int dg R_{j_1, m_1 n_1}(g) R_{j_2, m_2 n_2}(g) R_{j_3, m_3 n_3}(g) = \begin{pmatrix} j_1 & j_2 & j_3 \\ m_1 & m_2 & m_3 \end{pmatrix} \begin{pmatrix} j_1 & j_2 & j_3 \\ n_1 & n_2 & n_3 \end{pmatrix}$$

where $R_{j, mn}(g)$ are **group representation matrices** = $\langle j, m | g | j, n \rangle$
E.g. Wigner D-functions for $SU(2)$

Applied to our case: $\int dx R_{k_1, 00}(x) R_{h, 00}(x) R_{k_2, 00}(x)$

$R_{k, 00}(x) = K_{2ik}(e^x)$ are representation matrices in mixed (parabolic) basis = Whittaker functions in math literature

Group theory interpretation of JT gravity

Example: JT boundary two-point function:

$$\int dk_1 (k_1 \sinh 2\pi k_1) \int dk_2 (k_2 \sinh 2\pi k_2) e^{-\tau k_1^2 - (\beta - \tau) k_2^2} \frac{\Gamma(h \pm i k_1 \pm i k_2)}{\Gamma(2h)}$$

→ **Measure** and **energies** match with Plancherel measure and Casimir of continuous irreps of (modification of) $SL(2, \mathbb{R})$

→ **Vertex function** is 3j-symbol² with two such continuous irreps (states) and one discrete lowest weight irrep (operator):

$$\int dg R_{j_1, m_1 n_1}(g) R_{j_2, m_2 n_2}(g) R_{j_3, m_3 n_3}(g) = \begin{pmatrix} j_1 & j_2 & j_3 \\ m_1 & m_2 & m_3 \end{pmatrix} \begin{pmatrix} j_1 & j_2 & j_3 \\ n_1 & n_2 & n_3 \end{pmatrix}$$

where $R_{j, mn}(g)$ are **group representation matrices** = $\langle j, m | g | j, n \rangle$
E.g. Wigner D-functions for $SU(2)$

Applied to our case: $\int dx R_{k_1, 00}(x) R_{h, 00}(x) R_{k_2, 00}(x)$

$R_{k, 00}(x) = K_{2ik}(e^x)$ are representation matrices in mixed (parabolic) basis = Whittaker functions in math literature

$$\int_{-\infty}^{+\infty} dx K_{2ik_1}(e^x) e^{2hx} K_{2ik_2}(e^x) \sim \frac{\Gamma(h \pm i k_1 \pm i k_2)}{\Gamma(2h)}$$

Group theory interpretation of JT gravity

Example: JT boundary two-point function:

$$\int dk_1 (k_1 \sinh 2\pi k_1) \int dk_2 (k_2 \sinh 2\pi k_2) e^{-\tau k_1^2 - (\beta - \tau) k_2^2} \frac{\Gamma(h \pm i k_1 \pm i k_2)}{\Gamma(2h)}$$

→ **Measure** and **energies** match with Plancherel measure and Casimir of continuous irreps of (modification of) $SL(2, \mathbb{R})$

→ **Vertex function** is 3j-symbol² with two such continuous irreps (states) and one discrete lowest weight irrep (operator):

$$\int dg R_{j_1, m_1 n_1}(g) R_{j_2, m_2 n_2}(g) R_{j_3, m_3 n_3}(g) = \begin{pmatrix} j_1 & j_2 & j_3 \\ m_1 & m_2 & m_3 \end{pmatrix} \begin{pmatrix} j_1 & j_2 & j_3 \\ n_1 & n_2 & n_3 \end{pmatrix}$$

where $R_{j, mn}(g)$ are **group representation matrices** = $\langle j, m | g | j, n \rangle$
E.g. Wigner D-functions for $SU(2)$

Applied to our case: $\int dx R_{k_1, 00}(x) R_{h, 00}(x) R_{k_2, 00}(x)$

$R_{k, 00}(x) = K_{2ik}(e^x)$ are representation matrices in mixed (parabolic) basis = Whittaker functions in math literature

$$\int_{-\infty}^{+\infty} dx K_{2ik_1}(e^x) e^{2hx} K_{2ik_2}(e^x) \sim \frac{\Gamma(h \pm i k_1 \pm i k_2)}{\Gamma(2h)}$$

Explanation: 1st order $SL(2, \mathbb{R})$ BF formulation of JT gravity

Quantum group interpretation of Liouville gravity (1)

Liouville gravity amplitudes arise from (modification of)
 $\mathcal{U}_q(\mathfrak{sl}(2, \mathbb{R}))$, $q = e^{\pi i b^2}$ ($b \rightarrow 0 \equiv q \rightarrow 1$ is undeformed algebra)

Quantum group interpretation of Liouville gravity (1)

Liouville gravity amplitudes arise from (modification of)
 $\mathcal{U}_q(\mathfrak{sl}(2, \mathbb{R}))$, $q = e^{\pi i b^2}$ ($b \rightarrow 0 \equiv q \rightarrow 1$ is undeformed algebra)

Continuous self-dual irreps Ponsot-Teschner '99 . . .:

- ▶ Casimir operator $C_s \equiv \cosh 2\pi b s$
- ▶ Plancherel measure: $d\mu(s) = ds \sinh 2\pi b s \sinh \frac{2\pi s}{b}$

Quantum group interpretation of Liouville gravity (1)

Liouville gravity amplitudes arise from (modification of)
 $\mathcal{U}_q(\mathfrak{sl}(2, \mathbb{R}))$, $q = e^{\pi i b^2}$ ($b \rightarrow 0 \equiv q \rightarrow 1$ is undeformed algebra)

Continuous self-dual irreps Ponsot-Teschner '99 . . .:

- ▶ Casimir operator $C_s \equiv \cosh 2\pi b s$
- ▶ Plancherel measure: $d\mu(s) = ds \sinh 2\pi b s \sinh \frac{2\pi s}{b}$

Example: boundary two-point function:

$$\int_0^{+\infty} ds_1 ds_2 \rho(s_1) \rho(s_2) e^{-\ell_1 \cosh 2\pi b s_1} e^{-\ell_2 \cosh 2\pi b s_2} \frac{S_b(\beta_M \pm i s_1 \pm i s_2)}{S_b(2\beta_M)}$$

Measure and energies again match with Plancherel measure and Casimir operator

Quantum group interpretation of Liouville gravity (1)

Liouville gravity amplitudes arise from (modification of)
 $\mathcal{U}_q(\mathfrak{sl}(2, \mathbb{R}))$, $q = e^{\pi i b^2}$ ($b \rightarrow 0 \equiv q \rightarrow 1$ is undeformed algebra)

Continuous self-dual irreps [Ponsot-Teschner '99](#) . . . :

- ▶ Casimir operator $C_s \equiv \cosh 2\pi b s$
- ▶ Plancherel measure: $d\mu(s) = ds \sinh 2\pi b s \sinh \frac{2\pi s}{b}$

Example: boundary two-point function:

$$\int_0^{+\infty} ds_1 ds_2 \rho(s_1) \rho(s_2) e^{-\ell_1 \cosh 2\pi b s_1} e^{-\ell_2 \cosh 2\pi b s_2} \frac{S_b(\beta_M \pm i s_1 \pm i s_2)}{S_b(2\beta_M)}$$

Measure and energies again match with Plancherel measure and Casimir operator

Parabolic matrix element of $\mathcal{U}_q(\mathfrak{sl}(2, \mathbb{R}))$ $R_{s,00}^\epsilon(x)$ [Kharchev et al. '01](#):

$$e^{\pi i 2 s x} \int_{-\infty}^{+\infty} \frac{d\zeta}{(2\pi b)^{-2i\zeta/b - 2is/b}} S_b(-i\zeta) S_b(-i2s - i\zeta) e^{-\pi i \epsilon (\zeta^2 + 2s\zeta)} e^{2\pi i \zeta x}$$

Leads to correct vertex function as 3j-symbol with two such insertions and one discrete rep insertion: [TM-Turiaci '20](#), [Fan-TM '21](#)

$$\int_{-\infty}^{+\infty} dx R_{s_1,00}^\epsilon(x) R_{s_2,00}^{\epsilon*}(x) e^{2\beta_M \pi x} \sim \frac{S_b(\beta_M \pm i s_1 \pm i s_2)}{S_b(2\beta_M)}$$

Quantum group interpretation of Liouville gravity (2)

? **Explanation** by some q -BF formulation of Liouville gravity?

→ Open problem, we will present a different gauge theory perspective further on

Quantum group interpretation of Liouville gravity (2)

? **Explanation** by some q -BF formulation of Liouville gravity?

→ Open problem, we will present a different gauge theory perspective further on

Extension: $\mathcal{N} = 1$ Liouville supergravity analogously has

$\mathcal{U}_q(\mathfrak{osp}(1|2, \mathbb{R}))$ quantum supergroup structure [Fan-TM '21](#)

Proposal for parabolic representation matrix element for

$\mathcal{U}_q(\mathfrak{osp}(1|2, \mathbb{R}))$:

Quantum group interpretation of Liouville gravity (2)

? **Explanation** by some q -BF formulation of Liouville gravity?

→ Open problem, we will present a different gauge theory perspective further on

Extension: $\mathcal{N} = 1$ Liouville supergravity analogously has

$\mathcal{U}_q(\mathfrak{osp}(1|2, \mathbb{R}))$ quantum supergroup structure [Fan-TM '21](#)

Proposal for parabolic representation matrix element for

$\mathcal{U}_q(\mathfrak{osp}(1|2, \mathbb{R}))$:

$$R_{s,00}^{\epsilon,\pm}(x) = e^{\pi i s x} \int_{-\infty}^{+\infty} \frac{d\zeta}{(4\pi b)^{-2i\zeta/b-2is}} e^{-\pi i \frac{\epsilon}{2}(\zeta^2+2s\zeta)} e^{\pi i \zeta x} \\ \times [S_{\text{NS}}(-i\zeta) S_{\text{R}}(-i2s - i\zeta) \pm S_{\text{R}}(-i\zeta) S_{\text{NS}}(-i2s - i\zeta)]$$

computed using representation theory of quantum supergroup

Liouville gravity as a dilaton gravity model

Rewrite Liouville gravity as a dilaton gravity model with specific potential [Seiberg-Stanford \(unpublished\)](#), [TM-Turiaci '20](#)

Liouville gravity as a dilaton gravity model

Rewrite Liouville gravity as a dilaton gravity model with specific potential [Seiberg-Stanford \(unpublished\)](#), [TM-Turiaci '20](#)

Assumption: Describe matter as timelike Liouville:

$$S = S_L[\phi] + S_M[\chi]$$

$$S_L[\phi] = \frac{1}{4\pi} \int_{\Sigma} \left[(\hat{\nabla} \phi)^2 + 4\pi\mu e^{2b\phi} \right]$$

$$S_M[\chi] = \frac{1}{4\pi} \int_{\Sigma} \left[-(\hat{\nabla} \chi)^2 - 4\pi\mu e^{2b\chi} \right]$$

Liouville gravity as a dilaton gravity model

Rewrite Liouville gravity as a dilaton gravity model with specific potential [Seiberg-Stanford \(unpublished\)](#), [TM-Turiaci '20](#)

Assumption: Describe matter as timelike Liouville:

$$S = S_L[\phi] + S_M[\chi]$$

$$S_L[\phi] = \frac{1}{4\pi} \int_{\Sigma} \left[(\hat{\nabla} \phi)^2 + 4\pi \mu e^{2b\phi} \right]$$

$$S_M[\chi] = \frac{1}{4\pi} \int_{\Sigma} \left[-(\hat{\nabla} \chi)^2 - 4\pi \mu e^{2b\chi} \right]$$

Field redefinition: $\phi = b^{-1}\rho - b\pi\Phi$ and $\chi = b^{-1}\rho + b\pi\Phi$:

$$\rightarrow S = - \int \partial\Phi \cdot \partial\rho + \int e^{2\rho} (\mu e^{-2\pi b^2\Phi} - \mu e^{2\pi b^2\Phi})$$

Liouville gravity as a dilaton gravity model

Rewrite Liouville gravity as a dilaton gravity model with specific potential [Seiberg-Stanford \(unpublished\)](#), [TM-Turiaci '20](#)

Assumption: Describe matter as timelike Liouville:

$$S = S_L[\phi] + S_M[\chi]$$

$$S_L[\phi] = \frac{1}{4\pi} \int_{\Sigma} \left[(\hat{\nabla} \phi)^2 + 4\pi \mu e^{2b\phi} \right]$$

$$S_M[\chi] = \frac{1}{4\pi} \int_{\Sigma} \left[-(\hat{\nabla} \chi)^2 - 4\pi \mu e^{2b\chi} \right]$$

Field redefinition: $\phi = b^{-1}\rho - b\pi\Phi$ and $\chi = b^{-1}\rho + b\pi\Phi$:

$$\rightarrow S = - \int \partial\Phi \cdot \partial\rho + \int e^{2\rho} (\mu e^{-2\pi b^2\Phi} - \mu e^{2\pi b^2\Phi})$$

Setting $ds^2 = e^{2\rho} dz d\bar{z}$ ($R \sim e^{-2\rho} \partial\bar{\partial}\rho$) \rightarrow first term = $\int \sqrt{g} R \Phi$

Second term = $W(\Phi) \sim \sinh(2\pi b^2\Phi)$

The limit for small b gives back JT gravity

\Rightarrow **Liouville gravity = sinh dilaton gravity**

Classical dilaton gravity and asymptotics (1)

Let's now investigate holography in sinh dilaton gravity

Classical dilaton gravity and asymptotics (1)

Let's now investigate holography in sinh dilaton gravity

General dilaton gravity: [Gegenberg-Kunstatter-Louis-Martinez '94](#), [Witten '20](#)

$$S = \frac{1}{2} \int d^2x \sqrt{-g} (\Phi R + W(\Phi))$$

Up to bulk diffeo's, the [general classical solution](#) can be written as:

$$ds^2 = -A(r)dt^2 + \frac{dr^2}{A(r)}, \quad \Phi(r) = r \text{ with } A(r) = \int_{r_h}^r dr' W(r')$$

Classical dilaton gravity and asymptotics (1)

Let's now investigate holography in sinh dilaton gravity

General dilaton gravity: [Gegenberg-Kunstatter-Louis-Martinez '94](#), [Witten '20](#)

$$S = \frac{1}{2} \int d^2x \sqrt{-g} (\Phi R + W(\Phi))$$

Up to bulk diffeos, the [general classical solution](#) can be written as:

$$ds^2 = -A(r)dt^2 + \frac{dr^2}{A(r)}, \quad \Phi(r) = r \text{ with } A(r) = \int_{r_h}^r dr' W(r')$$

Describes BH geometry with horizon at $r = r_h$

Classical dilaton gravity and asymptotics (1)

Let's now investigate holography in sinh dilaton gravity

General dilaton gravity: [Gegenberg-Kunstatter-Louis-Martinez '94](#), [Witten '20](#)

$$S = \frac{1}{2} \int d^2x \sqrt{-g} (\Phi R + W(\Phi))$$

Up to bulk diffeos, the [general classical solution](#) can be written as:

$$ds^2 = -A(r)dt^2 + \frac{dr^2}{A(r)}, \quad \Phi(r) = r \text{ with } A(r) = \int_{r_h}^r dr' W(r')$$

Describes BH geometry with horizon at $r = r_h$

$$\text{and } E = \frac{1}{2} \int^{W^{-1}(4\pi T_H)} W(\Phi) d\Phi$$

Classical dilaton gravity and asymptotics (1)

Let's now investigate holography in sinh dilaton gravity

General dilaton gravity: [Gegenberg-Kunstatter-Louis-Martinez '94](#), [Witten '20](#)

$$S = \frac{1}{2} \int d^2x \sqrt{-g} (\Phi R + W(\Phi))$$

Up to bulk diffeos, the [general classical solution](#) can be written as:

$$ds^2 = -A(r)dt^2 + \frac{dr^2}{A(r)}, \quad \Phi(r) = r \text{ with } A(r) = \int_{r_h}^r dr' W(r')$$

Describes BH geometry with horizon at $r = r_h$

$$\text{and } E = \frac{1}{2} \int^{W^{-1}(4\pi T_H)} W(\Phi) d\Phi$$

Choosing $W(\Phi) \sim \sinh 2\pi b^2 \Phi$ leads to $\sqrt{E^2 - \kappa^2} = T_H/b^2$

→ [matches with first law derived in semi-classical regime from disk partition function](#)

Classical dilaton gravity and asymptotics (1)

Let's now investigate holography in sinh dilaton gravity

General dilaton gravity: [Gegenberg-Kunstatter-Louis-Martinez '94, Witten '20](#)

$$S = \frac{1}{2} \int d^2x \sqrt{-g} (\Phi R + W(\Phi))$$

Up to bulk diffeo's, the [general classical solution](#) can be written as:

$$ds^2 = -A(r)dt^2 + \frac{dr^2}{A(r)}, \quad \Phi(r) = r \text{ with } A(r) = \int_{r_h}^r dr' W(r')$$

Describes BH geometry with horizon at $r = r_h$

$$\text{and } E = \frac{1}{2} \int^{W^{-1}(4\pi T_H)} W(\Phi) d\Phi$$

Choosing $W(\Phi) \sim \sinh 2\pi b^2 \Phi$ leads to $\sqrt{E^2 - \kappa^2} = T_H/b^2$

→ [matches with first law derived in semi-classical regime from disk partition function](#)

$$ds^2 = -\frac{\cosh 2\pi b^2 r - \cosh 2\pi b^2 r_h}{2\pi b^2 \sinh \pi b^2} dt^2 + \frac{2\pi b^2 \sinh \pi b^2}{\cosh 2\pi b^2 r - \cosh 2\pi b^2 r_h} dr^2$$

For $r, r_h \ll 1/b^2$ we get JT black hole

Classical dilaton gravity and asymptotics (1)

Let's now investigate holography in sinh dilaton gravity

General dilaton gravity: [Gegenberg-Kunstatter-Louis-Martinez '94, Witten '20](#)

$$S = \frac{1}{2} \int d^2x \sqrt{-g} (\Phi R + W(\Phi))$$

Up to bulk diffeo's, the [general classical solution](#) can be written as:

$$ds^2 = -A(r)dt^2 + \frac{dr^2}{A(r)}, \quad \Phi(r) = r \text{ with } A(r) = \int_{r_h}^r dr' W(r')$$

Describes BH geometry with horizon at $r = r_h$

$$\text{and } E = \frac{1}{2} \int^{W^{-1}(4\pi T_H)} W(\Phi) d\Phi$$

Choosing $W(\Phi) \sim \sinh 2\pi b^2 \Phi$ leads to $\sqrt{E^2 - \kappa^2} = T_H/b^2$

→ [matches with first law derived in semi-classical regime from disk partition function](#)

$$ds^2 = -\frac{\cosh 2\pi b^2 r - \cosh 2\pi b^2 r_h}{2\pi b^2 \sinh \pi b^2} dt^2 + \frac{2\pi b^2 \sinh \pi b^2}{\cosh 2\pi b^2 r - \cosh 2\pi b^2 r_h} dr^2$$

For $r, r_h \ll 1/b^2$ we get JT black hole

For $r \rightarrow \infty$, $R \sim e^{2\pi b^2 r} \rightarrow$ Curvature singularity at the boundary

[q-deformed holography?](#)

Classical dilaton gravity and asymptotics (2)

Boundary behavior of fields:

Recall our field redefinition: $\phi = b^{-1}\rho - b\pi\Phi$ and $\chi = b^{-1}\rho + b\pi\Phi$

Classical dilaton gravity and asymptotics (2)

Boundary behavior of fields:

Recall our **field redefinition**: $\phi = b^{-1}\rho - b\pi\Phi$ and $\chi = b^{-1}\rho + b\pi\Phi$

Liouville length can be written as: $\ell_L \equiv \int e^{b\phi} dt = \int e^\rho e^{-\pi b^2 \Phi} dt$

Classical dilaton gravity and asymptotics (2)

Boundary behavior of fields:

Recall our **field redefinition**: $\phi = b^{-1}\rho - b\pi\Phi$ and $\chi = b^{-1}\rho + b\pi\Phi$

Liouville length can be written as: $\ell_L \equiv \int e^{b\phi} dt = \int e^\rho e^{-\pi b^2 \Phi} dt$

Using asymptotics from classical metric

$e^\rho|_{r \rightarrow \infty} = \lim_{r \rightarrow +\infty} \frac{e^{\pi b^2 r}}{\sqrt{\pi b^2 \sin \pi b^2}}$ and dilaton

$\Phi(r)|_{r \rightarrow \infty} = \lim_{r \rightarrow +\infty} r$, we learn that:

Classical dilaton gravity and asymptotics (2)

Boundary behavior of fields:

Recall our **field redefinition**: $\phi = b^{-1}\rho - b\pi\Phi$ and $\chi = b^{-1}\rho + b\pi\Phi$

Liouville length can be written as: $\ell_L \equiv \int e^{b\phi} dt = \int e^\rho e^{-\pi b^2 \Phi} dt$

Using asymptotics from classical metric

$e^\rho|_{r \rightarrow \infty} = \lim_{r \rightarrow +\infty} \frac{e^{\pi b^2 r}}{\sqrt{\pi b^2 \sin \pi b^2}}$ and dilaton

$\Phi(r)|_{r \rightarrow \infty} = \lim_{r \rightarrow +\infty} r$, we learn that:

- ▶ $\ell_L \sim \ell$ where ℓ is the length in the dt metric
 - Liouville length ℓ_L and length ℓ measured in the dilaton gravity boundary metric are essentially equal
 - Explains (in part) why JT gravity is found in the limit

Classical dilaton gravity and asymptotics (2)

Boundary behavior of fields:

Recall our **field redefinition**: $\phi = b^{-1}\rho - b\pi\Phi$ and $\chi = b^{-1}\rho + b\pi\Phi$
Liouville length can be written as: $\ell_L \equiv \int e^{b\phi} dt = \int e^\rho e^{-\pi b^2 \Phi} dt$

Using asymptotics from classical metric

$$e^\rho|_{r \rightarrow \infty} = \lim_{r \rightarrow +\infty} \frac{e^{\pi b^2 r}}{\sqrt{\pi b^2 \sin \pi b^2}} \text{ and dilaton}$$

$\Phi(r)|_{r \rightarrow \infty} = \lim_{r \rightarrow +\infty} r$, we learn that:

- ▶ $\ell_L \sim \ell$ where ℓ is the length in the dt metric
→ Liouville length ℓ_L and length ℓ measured in the dilaton gravity boundary metric are essentially equal
→ Explains (in part) why JT gravity is found in the limit
- ▶ $e^{b\chi}|_{\partial} \sim \lim_{r \rightarrow +\infty} e^{2\pi b^2 r} \rightarrow +\infty \Rightarrow$ Matter sector needs to be described by identity brane (i.e. ZZ) boundary condition, which was used in the Liouville gravity calculations indeed

Poisson sigma model description (1)

How do we see the quantum group structure within this action?

First-order form of 2d dilaton gravity:

$$S = \frac{1}{2} \int d^2x \sqrt{-g} (\Phi R + W(\Phi)) = \int \left[\Phi d\omega + \frac{1}{4} W(\Phi) \epsilon^{ab} e_a \wedge e_b + X^a (de_a + \epsilon_a{}^b \omega \wedge e_b) \right]$$

Poisson sigma model description (1)

How do we see the quantum group structure within this action?

First-order form of 2d dilaton gravity:

$$S = \frac{1}{2} \int d^2x \sqrt{-g} (\Phi R + W(\Phi)) = \int \left[\Phi d\omega + \frac{1}{4} W(\Phi) \epsilon^{ab} e_a \wedge e_b + X^a (de_a + \epsilon_a{}^b \omega \wedge e_b) \right]$$

The general Poisson sigma model is of the form: Ikeda '93, Schaller-Strobl '94

$$S = \int (A_i \wedge dX^i - \frac{1}{2} A_i \wedge A_j P^{ji}(X))$$

where A_i is gauge connection, and X^i , $i = 1..m$ coordinatize a m -dimensional Poisson manifold target space:

$$\{X^i, X^j\}_{\text{PB}} = P^{ij}(X), \quad P^{ij} = -P^{ji}, \quad \partial_\ell P^{[ij} P^{\ell]k} = 0$$

Poisson sigma model description (1)

How do we see the quantum group structure within this action?

First-order form of 2d dilaton gravity:

$$S = \frac{1}{2} \int d^2x \sqrt{-g} (\Phi R + W(\Phi)) = \int \left[\Phi d\omega + \frac{1}{4} W(\Phi) \epsilon^{ab} e_a \wedge e_b + X^a (de_a + \epsilon_a{}^b \omega \wedge e_b) \right]$$

The general Poisson sigma model is of the form: Ikeda '93, Schaller-Strobl '94

$$S = \int (A_i \wedge dX^i - \frac{1}{2} A_i \wedge A_j P^{ji}(X))$$

where A_i is gauge connection, and X^i , $i = 1..m$ coordinatize a m -dimensional Poisson manifold target space:

$$\{X^i, X^j\}_{\text{PB}} = P^{ij}(X), \quad P^{ij} = -P^{ji}, \quad \partial_\ell P^{[ij]} P^{\ell[k]} = 0$$

Identifying $A_i = (e_0, e_1, \omega)$ and $X^i = (X^0, X^1, \Phi)$

→ 2d dilaton gravity is a special case of the PS model:

$$\{X^0, X^1\}_{\text{PB}} = \frac{W(X^2)}{2}, \quad \{X^a, X^2\}_{\text{PB}} = \epsilon^a{}_b X^b$$

Poisson sigma model description (1)

How do we see the quantum group structure within this action?

First-order form of 2d dilaton gravity:

$$S = \frac{1}{2} \int d^2x \sqrt{-g} (\Phi R + W(\Phi)) = \int [\Phi d\omega + \frac{1}{4} W(\Phi) \epsilon^{ab} e_a \wedge e_b + X^a (de_a + \epsilon_a{}^b \omega \wedge e_b)]$$

The general Poisson sigma model is of the form: Ikeda '93, Schaller-Strobl '94

$$S = \int (A_i \wedge dX^i - \frac{1}{2} A_i \wedge A_j P^{ji}(X))$$

where A_i is gauge connection, and X^i , $i = 1..m$ coordinatize a m -dimensional Poisson manifold target space:

$$\{X^i, X^j\}_{\text{PB}} = P^{ij}(X), \quad P^{ij} = -P^{ji}, \quad \partial_\ell P^{[ij]} P^{\ell[k]} = 0$$

Identifying $A_i = (e_0, e_1, \omega)$ and $X^i = (X^0, X^1, \Phi)$

→ 2d dilaton gravity is a special case of the PS model:

$$\{X^0, X^1\}_{\text{PB}} = \frac{W(X^2)}{2}, \quad \{X^a, X^2\}_{\text{PB}} = \epsilon^a{}_b X^b$$

or with lightcone coordinates $E^\pm \equiv -X^0 \pm X^1$ and $H \equiv X^2$:

$$\{H, E^\pm\}_{\text{PB}} = \pm E^\pm, \quad \{E^+, E^-\}_{\text{PB}} = W(H)$$

Poisson sigma model description (2)

Poisson algebra is “external” structure

Poisson sigma model description (2)

Poisson algebra is “external” structure \Rightarrow Identify the Poisson algebra with a symmetry algebra of the dynamical system

Poisson sigma model description (2)

Poisson algebra is “external” structure \Rightarrow Identify the Poisson algebra with a symmetry algebra of the dynamical system

$$S = \int (A_i \wedge dX^i - \frac{1}{2} A_i \wedge A_j P^{ji}(X))$$

Non-linear symmetry transformation:

$$\delta X^i = -\epsilon_j P^{ji}, \quad \delta A_i = -d\epsilon_i + A_j \epsilon_k \partial_i P^{kj}$$

Poisson sigma model description (2)

Poisson algebra is “external” structure \Rightarrow Identify the Poisson algebra with a symmetry algebra of the dynamical system

$$S = \int (A_i \wedge dX^i - \frac{1}{2} A_i \wedge A_j P^{ji}(X))$$

Non-linear symmetry transformation:

$$\delta X^i = -\epsilon_j P^{ji}, \quad \delta A_i = -d\epsilon_i + A_j \epsilon_k \partial_i P^{kj}$$

Conserved charges:

$$Q^i \equiv \int dx \delta^i X^j \pi_{X^j} = \dots = - \int_0^{+\infty} du A_{1j}(u) P^{ji}(X(u))$$

$$\text{with } \pi_{X^i}(x) \equiv \frac{\partial L}{\partial(\partial_0 X^i)} = -A_{1i}$$

Poisson sigma model description (2)

Poisson algebra is “external” structure \Rightarrow Identify the Poisson algebra with a symmetry algebra of the dynamical system

$$S = \int (A_i \wedge dX^i - \frac{1}{2} A_i \wedge A_j P^{ji}(X))$$

Non-linear symmetry transformation:

$$\delta X^i = -\epsilon_j P^{ji}, \quad \delta A_i = -d\epsilon_i + A_j \epsilon_k \partial_i P^{kj}$$

Conserved charges:

$$Q^i \equiv \int dx \delta^i X^j \pi_{X^j} = \dots = - \int_0^{+\infty} du A_{1j}(u) P^{ji}(X(u))$$

$$\text{with } \pi_{X^i}(x) \equiv \frac{\partial L}{\partial(\partial_0 X^i)} = -A_{1i}$$

charge algebra: [see e.g. Cattaneo-Felder '01](#)

$$\Rightarrow \boxed{\{Q^i, Q^j\} = P^{ij}(Q)}$$

\rightarrow same as original Poisson algebra, but now realized as a canonical phase-space algebra

Poisson sigma model description (3)

Upon quantization: $[\hat{Q}^i, \hat{Q}^j] \stackrel{?}{=} i\hbar \hat{P}^{ij}(\hat{Q})$

→ possible **ordering ambiguities** in Poisson tensor Fan-TM '21

Poisson sigma model description (3)

Upon quantization: $[\hat{Q}^i, \hat{Q}^j] \stackrel{?}{=} i\hbar \hat{P}^{ij}(\hat{Q})$

→ possible **ordering ambiguities** in Poisson tensor Fan-TM '21

For bosonic dilaton gravity \Rightarrow no issue

Poisson sigma model description (3)

Upon quantization: $[\hat{Q}^i, \hat{Q}^j] \stackrel{?}{=} i\hbar \hat{P}^{ij}(\hat{Q})$

→ possible **ordering ambiguities** in Poisson tensor Fan-TM '21

For bosonic dilaton gravity \Rightarrow no issue

For $\mathcal{N} = 1$ dilaton supergravity \Rightarrow important difference!

→ resulting algebra is **unique** (compatibility with Jacobi identity)

Poisson sigma model description (3)

Upon quantization: $[\hat{Q}^i, \hat{Q}^j] \stackrel{?}{=} i\hbar \hat{P}^{ij}(\hat{Q})$

→ possible **ordering ambiguities** in Poisson tensor Fan-TM '21

For bosonic dilaton gravity \Rightarrow no issue

For $\mathcal{N} = 1$ dilaton supergravity \Rightarrow important difference!

→ resulting algebra is **unique** (compatibility with Jacobi identity)

Main statements:

- **well-known**: For $W(H) = 2H$, the charge algebra is the $\mathfrak{sl}(2, \mathbb{R})$ Lie algebra
Matches with BF description of JT gravity

Poisson sigma model description (3)

Upon quantization: $[\hat{Q}^i, \hat{Q}^j] \stackrel{?}{=} i\hbar \hat{P}^{ij}(\hat{Q})$

→ possible **ordering ambiguities** in Poisson tensor Fan-TM '21

For bosonic dilaton gravity \Rightarrow no issue

For $\mathcal{N} = 1$ dilaton supergravity \Rightarrow important difference!

→ resulting algebra is **unique** (compatibility with Jacobi identity)

Main statements:

- ▶ **well-known**: For $W(H) = 2H$, the charge algebra is the $\mathfrak{sl}(2, \mathbb{R})$ Lie algebra

Matches with BF description of JT gravity

- ▶ **new**: For $W(H) \sim \sinh 2\pi b^2 H$, the charge algebra becomes the q -deformed algebra $U_q(\mathfrak{sl}(2, \mathbb{R}))$

Explains organization of Liouville gravity amplitudes in terms of $U_q(\mathfrak{sl}(2, \mathbb{R}))$

Poisson sigma model description (3)

Upon quantization: $[\hat{Q}^i, \hat{Q}^j] \stackrel{?}{=} i\hbar \hat{P}^{ij}(\hat{Q})$

→ possible **ordering ambiguities** in Poisson tensor Fan-TM '21

For bosonic dilaton gravity \Rightarrow no issue

For $\mathcal{N} = 1$ dilaton supergravity \Rightarrow important difference!

→ resulting algebra is **unique** (compatibility with Jacobi identity)

Main statements:

- ▶ **well-known**: For $W(H) = 2H$, the charge algebra is the $\mathfrak{sl}(2, \mathbb{R})$ Lie algebra
Matches with BF description of JT gravity
- ▶ **new**: For $W(H) \sim \sinh 2\pi b^2 H$, the charge algebra becomes the q -deformed algebra $U_q(\mathfrak{sl}(2, \mathbb{R}))$
Explains organization of Liouville gravity amplitudes in terms of $U_q(\mathfrak{sl}(2, \mathbb{R}))$
- ▶ **new**: For prepotential $u(H) \sim \sinh 2\pi b^2 H$ in $\mathcal{N} = 1$ dilaton supergravity, the charge algebra becomes the q -deformed algebra $U_q(\mathfrak{osp}(1|2, \mathbb{R}))$

Deformation of JT and bigger picture (1)

Up to now: 2 dilaton potentials ($\sim \Phi$, $\sim \sinh 2\pi b^2 \Phi$)

Next: understand bigger picture for generic dilaton potentials

Deformation of JT and bigger picture (1)

Up to now: 2 dilaton potentials ($\sim \Phi$, $\sim \sinh 2\pi b^2 \Phi$)

Next: understand bigger picture for generic dilaton potentials

Class of deformations of the JT potential: [Maxfield-Turiaci '20](#), [Witten '20](#)

$$W(\Phi) = 2\Phi + \sum_i \epsilon_i e^{-\alpha_i \Phi}, \quad \pi < \alpha_i < 2\pi$$

Deformation of JT and bigger picture (1)

Up to now: 2 dilaton potentials ($\sim \Phi$, $\sim \sinh 2\pi b^2 \Phi$)

Next: understand bigger picture for generic dilaton potentials

Class of deformations of the JT potential: [Maxfield-Turiaci '20](#), [Witten '20](#)

$$W(\Phi) = 2\Phi + \sum_i \epsilon_i e^{-\alpha_i \Phi}, \quad \pi < \alpha_i < 2\pi$$

Interpretable as gas of (elliptic) defects in JT

→ Riemann surfaces with conical punctures

Deformation of JT and bigger picture (1)

Up to now: 2 dilaton potentials ($\sim \Phi$, $\sim \sinh 2\pi b^2 \Phi$)

Next: understand bigger picture for generic dilaton potentials

Class of deformations of the JT potential: [Maxfield-Turiaci '20](#), [Witten '20](#)

$$W(\Phi) = 2\Phi + \sum_i \epsilon_i e^{-\alpha_i \Phi}, \quad \pi < \alpha_i < 2\pi$$

Interpretable as gas of (elliptic) defects in JT

→ Riemann surfaces with conical punctures

Leads to a deformed density of states $\rho(k) \rightarrow \rho_{\text{def}}(k)$

Deformation of JT and bigger picture (1)

Up to now: 2 dilaton potentials ($\sim \Phi$, $\sim \sinh 2\pi b^2 \Phi$)

Next: understand bigger picture for generic dilaton potentials

Class of deformations of the JT potential: [Maxfield-Turiaci '20](#), [Witten '20](#)

$$W(\Phi) = 2\Phi + \sum_i \epsilon_i e^{-\alpha_i \Phi}, \quad \pi < \alpha_i < 2\pi$$

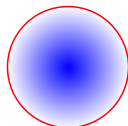
Interpretable as gas of (elliptic) defects in JT

→ Riemann surfaces with conical punctures

Leads to a deformed density of states $\rho(k) \rightarrow \rho_{\text{def}}(k)$

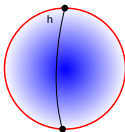
Geometric picture: With boundary asymptotics of classical $\Phi = r$

⇒ Does not modify $r \rightarrow +\infty$ asymptotics of JT gravity:



Deformation of JT and bigger picture (2)

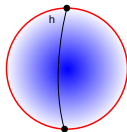
Similar argument works when boundary operator insertions are present:



$\rho(k) \rightarrow \rho_{\text{def}}(k)$ only, the vertex functions (Γ 's) are the same

Deformation of JT and bigger picture (2)

Similar argument works when boundary operator insertions are present:

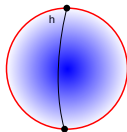


$\rho(k) \rightarrow \rho_{\text{def}}(k)$ only, the vertex functions (Γ 's) are the same

Intuition: Gas of defects does not reach the boundary where the 3-point vertices are located [Fan-TM '21](#)

Deformation of JT and bigger picture (2)

Similar argument works when boundary operator insertions are present:



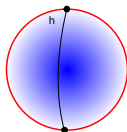
$\rho(k) \rightarrow \rho_{\text{def}}(k)$ only, the vertex functions (Γ 's) are the same

Intuition: Gas of defects does not reach the boundary where the 3-point vertices are located [Fan-TM '21](#)

Back to Liouville gravity: $\sinh 2\pi b^2 \Phi$ dilaton potential has different asymptotics

Deformation of JT and bigger picture (2)

Similar argument works when boundary operator insertions are present:

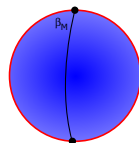


$\rho(k) \rightarrow \rho_{\text{def}}(k)$ only, the vertex functions (Γ 's) are the same

Intuition: Gas of defects does not reach the boundary where the 3-point vertices are located [Fan-TM '21](#)

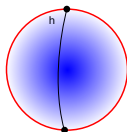
Back to Liouville gravity: $\sinh 2\pi b^2 \Phi$ dilaton potential has different asymptotics

Geometric picture: Gas of defects reaches bdy
 \Rightarrow Different vertex functions (S_b 's)



Deformation of JT and bigger picture (2)

Similar argument works when boundary operator insertions are present:



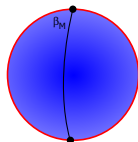
$\rho(k) \rightarrow \rho_{\text{def}}(k)$ only, the vertex functions (Γ 's) are the same

Intuition: Gas of defects does not reach the boundary where the 3-point vertices are located [Fan-TM '21](#)

Back to Liouville gravity: $\sinh 2\pi b^2 \Phi$ dilaton potential has different asymptotics

Geometric picture: Gas of defects reaches bdy
 \Rightarrow Different vertex functions (S_b 's)

Suggests classification of dilaton gravity models in different classes depending on the asymptotics of the dilaton potential



- ▶ **Observation 1:** Fixed length amplitudes in Liouville gravity have a JT limit where $b \rightarrow 0$

Summary

- ▶ **Observation 1:** Fixed length amplitudes in Liouville gravity have a JT limit where $b \rightarrow 0$
Conceptual advantage: JT gravity embedded in string theory, worldsheet genus expansion is reinterpreted as multi-universe expansion

Summary

- ▶ **Observation 1:** Fixed length amplitudes in Liouville gravity have a JT limit where $b \rightarrow 0$
Conceptual advantage: JT gravity embedded in string theory, worldsheet genus expansion is reinterpreted as multi-universe expansion
- ▶ **Observation 2:** Amplitudes display **quantum group** structure

Summary

- ▶ **Observation 1:** Fixed length amplitudes in Liouville gravity have a JT limit where $b \rightarrow 0$
Conceptual advantage: JT gravity embedded in string theory, worldsheet genus expansion is reinterpreted as multi-universe expansion
- ▶ **Observation 2:** Amplitudes display **quantum group** structure
Partial explanation from the Lagrangian perspective by rewriting Liouville gravity as a **dilaton gravity model with sinh potential** and using its **Poisson-sigma model description**
→ no full understanding (yet)

Summary

- ▶ **Observation 1:** Fixed length amplitudes in Liouville gravity have a JT limit where $b \rightarrow 0$
Conceptual advantage: JT gravity embedded in string theory, worldsheet genus expansion is reinterpreted as multi-universe expansion
- ▶ **Observation 2:** Amplitudes display **quantum group** structure
Partial explanation from the Lagrangian perspective by rewriting Liouville gravity as a **dilaton gravity model with sinh potential** and using its **Poisson-sigma model description**
→ no full understanding (yet)

Can we obtain a more general understanding along these lines of larger classes of 2d dilaton gravity models?

Summary

- ▶ **Observation 1:** Fixed length amplitudes in Liouville gravity have a JT limit where $b \rightarrow 0$
Conceptual advantage: JT gravity embedded in string theory, worldsheet genus expansion is reinterpreted as multi-universe expansion
- ▶ **Observation 2:** Amplitudes display quantum group structure
Partial explanation from the Lagrangian perspective by rewriting Liouville gravity as a dilaton gravity model with sinh potential and using its Poisson-sigma model description
→ no full understanding (yet)

Can we obtain a more general understanding along these lines of larger classes of 2d dilaton gravity models?

Thank you!