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# The BMS algebra at spatial infinity (D = 4 and D > 4)

Marc Henneaux

Eurostrings 2022, Lyon 26 April 2022

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## Why should one study the asymptotic structure of gravity at spatial infinity in the asymptotically flat context?

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## Why should one study the asymptotic structure of gravity at spatial infinity in the asymptotically flat context?

Will give here three reasons.



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### First reason

In order to better understand the role of the BMS group (first identified at null infinity) in the quantum theory, where physical states are usually defined on spacelike (Cauchy) hypersurfaces,

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### First reason

In order to better understand the role of the BMS group (first identified at null infinity) in the quantum theory, where physical states are usually defined on spacelike (Cauchy) hypersurfaces, it is important to unveil its action on spacelike (Cauchy) hypersurfaces, and thus, at spatial infinity.

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### (First reason, continued)

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### (First reason, continued)

This has been done recently in four spacetime dimensions, through a reconsideration of the boundary conditions at spatial infinity.

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### (First reason, continued)

This has been done recently in four spacetime dimensions, through a reconsideration of the boundary conditions at spatial infinity.

New, consistent boundary conditions have been proposed, which are invariant under the full infinite-dimensional BMS group,

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This has been done recently in four spacetime dimensions, through a reconsideration of the boundary conditions at spatial infinity.

New, consistent boundary conditions have been proposed, which are invariant under the full infinite-dimensional BMS group,

providing a standard, non-trivial, canonical realization of the BMS symmetry.

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### (First reason, continued)

This has been done recently in four spacetime dimensions, through a reconsideration of the boundary conditions at spatial infinity.

New, consistent boundary conditions have been proposed, which are invariant under the full infinite-dimensional BMS group,

providing a standard, non-trivial, canonical realization of the BMS symmetry.

This establishes also the important fact that the BMS symmetry is a symmetry of the theory and not just a symmetry at null infinity.

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### Second reason

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### Second reason

Another reason for investigating the asymptotic structure at spatial infinity is the need to understand the "matching" conditions of the fields and charges between  $\mathscr{I}_{-}^{+}$  and  $\mathscr{I}_{+}^{-}$ 

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### Second reason

Another reason for investigating the asymptotic structure at spatial infinity is the need to understand the "matching" conditions of the fields and charges between  $\mathscr{I}_{-}^{+}$  and  $\mathscr{I}_{+}^{-}$  which clearly involves "going through" spatial infinity  $t^{0}$ .

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Another reason for investigating the asymptotic structure at spatial infinity is the need to understand the "matching" conditions of the fields and charges between  $\mathscr{I}_{-}^{+}$  and  $\mathscr{I}_{+}^{-}$  which clearly involves "going through" spatial infinity  $i^{0}$ . For instance, one knows that under very general initial conditions, the past limit of the Bondi  $m_{B}$  mass along future null infinity is equal to the ADM mass m.

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Similarly, the future limit of the Bondi mass  $m_B$  along past null infinity is also equal to the ADM mass m,

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Similarly, the future limit of the Bondi mass  $m_B$  along past null infinity is also equal to the ADM mass m,

implying the matching  $m_B(\mathscr{I}_{-}^+) = m = m_B(\mathscr{I}_{+}^-)$ .

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But the generator of time translations is only one of the Bondi-Metzner-Sachs (BMS) supertranslation generators.

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implying the matching  $m_B(\mathscr{I}^+_-) = m = m_B(\mathscr{I}^-_+)$ .

But the generator of time translations is only one of the Bondi-Metzner-Sachs (BMS) supertranslation generators.

Can one say more?

This requires understanding the action of all the BMS supertranslations at spatial infinity.

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### Yet another reason :

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### Yet another reason :

In five dimensions, the definition of null infinity is problematical (as it is in all odd spacetime dimensions).

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### Yet another reason :

In five dimensions, the definition of null infinity is problematical (as it is in all odd spacetime dimensions). But there exist soft theorems!

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### Yet another reason :

In five dimensions, the definition of null infinity is problematical (as it is in all odd spacetime dimensions).

### But there exist soft theorems!

Of which symmetries are these the Ward identities?

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It turns out that while the answer to this question is not immediate at null infinity, the analysis at spatial infinity raises no conceptual problem

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It turns out that while the answer to this question is not immediate at null infinity, the analysis at spatial infinity raises no conceptual problem

and directly leads to the infinite dimensional symmetry "BMS $_{5}$  group",

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It turns out that while the answer to this question is not immediate at null infinity, the analysis at spatial infinity raises no conceptual problem

and directly leads to the infinite dimensional symmetry " $\mathrm{BMS}_5$  group",

the realization of which exhibits (somewhat unexpectedly) a very interesting nonlinear structure.

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Conclusions and comments The purpose of this talk is to provide the general ideas of the asymptotic analysis at spatial infinity in D = 4 and D = 5 spacetime dimensions.

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The purpose of this talk is to provide the general ideas of the asymptotic analysis at spatial infinity in D = 4 and D = 5 spacetime dimensions.

The study will be carried on spacelike hypersurfaces that are asymptotically flat hyperplanes, using Hamiltonian methods.

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Conclusions and comments The purpose of this talk is to provide the general ideas of the asymptotic analysis at spatial infinity in D = 4 and D = 5 spacetime dimensions.

The study will be carried on spacelike hypersurfaces that are asymptotically flat hyperplanes, using Hamiltonian methods. (Work done in collaboration with Cédric Troessaert, with also Oscar Fuentealba, Sucheta Majumdar, Javier Matulich and Turmoli Neogi who joined more recently)

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# A central role in the analysis will be played by the gravitational action

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# A central role in the analysis will be played by the gravitational action

which reads, in Hamiltonian form (Dirac, ADM),

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# A central role in the analysis will be played by the gravitational action

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$$S[g_{ij},\pi^{ij},N,N^i] = \int dt \left\{ \int d^d x \left( \pi^{ij} \partial_t g_{ij} - N^i \mathcal{H}_i^{grav} - N \mathcal{H}_i^{grav} \right) - B_{\infty} \right\}$$

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where  $B_{\infty}$  is a boundary term at infinity and where
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where  $B_{\infty}$  is a boundary term at infinity and where

$$\mathcal{H}^{grav} = -\sqrt{g}R + \frac{1}{\sqrt{g}}(\pi^{ij}\pi_{ij} - \frac{1}{d-1}\pi^2) \approx 0, \quad \mathcal{H}^{grav}_i = -2\nabla_j \pi^j_i \approx 0.$$

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Conclusions and comments A central role in the analysis will be played by the gravitational action which reads, in Hamiltonian form (Dirac, ADM),

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where  $B_{\infty}$  is a boundary term at infinity and where

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The definition of the theory is completed by providing boundary conditions on the dynamical variables (definition of phase space), which are assumed to make the off-shell action finite.

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Conclusions and comments The standard boundary conditions for Einstein gravity in four spacetime dimensions are (Regge-Teitelboim 1974)

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$$h_{ij} \equiv g_{ij} - \delta_{ij} = \frac{\overline{h}_{ij}(\mathbf{n}^k)}{r} + O(\frac{1}{r^2}), \quad \overline{h}_{ij}(-\mathbf{n}^k) = \overline{h}_{ij}(\mathbf{n}^k)$$
$$\pi^{ij} = \frac{\overline{\pi}^{ij}(\mathbf{n}^k)}{r^2} + O(\frac{1}{r^3}), \quad \overline{\pi}^{ij}(-\mathbf{n}^k) = -\overline{\pi}^{ij}(\mathbf{n}^k).$$

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and

$$\pi^{ij} = \frac{\overline{\pi}^{ij}(\mathbf{n}^k)}{r^2} + O(\frac{1}{r^3}), \quad \overline{\pi}^{ij}(-\mathbf{n}^k) = -\overline{\pi}^{ij}(\mathbf{n}^k).$$

They involve strict parity conditions under the antipodal map  $\mathbf{n}^k \rightarrow -\mathbf{n}^k$ , where  $\mathbf{n}^k$  is the unit normal to the sphere at infinity

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They involve strict parity conditions under the antipodal map  $\mathbf{n}^k \rightarrow -\mathbf{n}^k$ , where  $\mathbf{n}^k$  is the unit normal to the sphere at infinity  $(f(\mathbf{n}^k) \equiv f(\theta, \varphi))$ .

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Conclusions and comments To see the full BMS group, one must allow a "parity-twisted component" in the leading orders of the asymptotic metric and momenta.

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Conclusions and comments To see the full BMS group, one must allow a "parity-twisted component" in the leading orders of the asymptotic metric and momenta.

One thus imposes

$$h_{ij} \equiv g_{ij} - \delta_{ij} = h_{ij}^{RT} + U_{ij}, \qquad \pi^{ij} = \pi_{RT}^{ij} + V^{ij}$$

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 $U_{ij}$  and  $V^{ij}$  are the parity-twisted contributions that take the form of a gauge transformation (rewritten in Hamiltonian form).

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 $U_{ij}$  and  $V^{ij}$  are the parity-twisted contributions that take the form of a gauge transformation (rewritten in Hamiltonian form). They are of the same order as the leading terms in  $h_{ij}$  and  $\pi^{ij}$ (O(1/r) and  $O(1/r^2)$  respectively but have the opposite parity. For a canonical action of the Lorentz boosts, one imposes moreover that  $U_{ij}$  be parametrized by an O(1) odd function of the angles  $\overline{U}(\mathbf{n}^k) = O(1) = -U(-\mathbf{n}^k)$ . Furthermore,  $V^{ij}$  is parametrized by an O(1) even function of the angles  $V(\mathbf{n}^k) = V(-\mathbf{n}^k)$ .

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Conclusions and comments Do these relaxed parity conditions involving a twist lead to a consistent description (finite symplectic form, well-defined generators)?

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Conclusions and comments Do these relaxed parity conditions involving a twist lead to a consistent description (finite symplectic form, well-defined generators)? The answer is affirmative and requires some work (even though the idea is elementary).

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Conclusions and comments Do these relaxed parity conditions involving a twist lead to a consistent description (finite symplectic form, well-defined generators)? The answer is affirmative and requires some work (even though the idea is elementary).

One finds furthermore that the asymptotic symmetries are given by hypersurface deformations that behave asymptotically as

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Conclusions and comments Do these relaxed parity conditions involving a twist lead to a consistent description (finite symplectic form, well-defined generators)? The answer is affirmative and requires some work (even though the idea is elementary).

One finds furthermore that the asymptotic symmetries are given by hypersurface deformations that behave asymptotically as

$$\begin{aligned} \xi &= b_i x^i + T(\mathbf{n}) + O(r^{-1}) \\ \xi^i &= b_j^i x^j + W_i(\mathbf{n}) + O(r^{-1}), \quad b_{ij} = -b_{ji}, \quad W_i(\mathbf{n}) = \partial_i (rW(\mathbf{n})). \end{aligned}$$

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where *T* is even and *W* is odd.

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where *T* is even and *W* is odd. The terms  $b_i x^i$  and  $b^i_j x^j$  describe respectively boosts and spatial rotations.

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where T is even and W is odd.

The terms  $b_i x^i$  and  $b^i_j x^j$  describe respectively boosts and spatial rotations.

The zero mode of *T* and the first spherical harmonic component of *W* describe translations.

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In fact, the even function T and the odd function W combine to form a single arbitrary function of the angles, as in the null infinity description of the supertranslations.

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The symmetries are canonical transformations with generators  $P_{\xi}[g_{ij}, \pi^{ij}] = \int d^3x \left(\xi \mathcal{H} + \xi^i \mathcal{H}_i\right) + \mathcal{B}_{\xi}[g_{ij}, \pi^{ij}]$ 

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 $P_{\xi}[g_{ij},\pi^{ij}] = \int d^3x \left(\xi \mathcal{H} + \xi^i \mathcal{H}_i\right) + \mathcal{B}_{\xi}[g_{ij},\pi^{ij}]$ 

where  $\mathscr{B}_{\xi}[g_{ij}, \pi^{ij}]$  is a surface term, the explicit form of which can be found in MH and C. Troessaert.

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The algebra of the generators can be easily verified to be the BMS algebra.

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#### There is complete agreement with the null infinity results.

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In particular, the "matching conditions" of Strominger, which involve the antipodal map, are in fact a consequence of the boundary conditions at spatial infinity

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Only "parity up to a gauge transformation" conditions are allowed

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(and these are necessary to avoid log singularities at null infinity).
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As in 4 dimensions, one must include explicitly the improper gauge symmetries in the asymptotic form of the fields.

A new feature is that improper gauge terms are not at the same order in  $\frac{1}{r}$  as the Coulomb part,

while in 4 dimensions, they are at the same order but distinguished by parity conditions.

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$$g_{ij} = \delta_{ij} + \mathcal{G}_{ij} + h_{ij}^{core}$$

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# Even though the ideas are identical, there are striking new features in 5 dimensions :

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# Even though the ideas are identical, there are striking new features in 5 dimensions :

• The size of BMS<sub>5</sub> is bigger than expected. More precisely, the supertranslations depend on four functions of the angles.

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Conclusions and comments Even though the ideas are identical, there are striking new features in 5 dimensions :

- The size of BMS<sub>5</sub> is bigger than expected. More precisely, the supertranslations depend on four functions of the angles.
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- Supertranslation charges (including the energy) acquire non-linear contributions.
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- The size of BMS<sub>5</sub> is bigger than expected. More precisely, the supertranslations depend on four functions of the angles.
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- The asymptotic symmetry generators form a non-linear algebra. In particular, the brackets of the boosts acquire cubic contributions.
- There are central charges among the different types of supertranslation generators.

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See O. Fuentealba, M. Henneaux, J. Matulich and C. Troessaert, e-Print: 2111.09664 [hep-th]

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### Contrary to the familiar ADM expression for the energy in 4D,

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Explicitly,

$$E = 2 \oint_{\mathcal{S}_{\infty}^3} d^3x \sqrt{\overline{\gamma}} \Big[ (1/2)\overline{h}_A^A + D_A \overline{\lambda}^A + 3\overline{\lambda} - (1/8)\theta_{AB} \theta^{AB} \Big]$$

where

$$g_{rr} = 1 + \frac{2\overline{\lambda}}{r^2} + \frac{h_{rr}^{(2)}}{r^3} + \mathcal{O}\left(r^{-4}\right),$$
(4.1)

$$g_{rA} = \frac{\overline{\lambda}_A}{r} + \frac{h_{rA}^{(2)}}{r^2} + \mathcal{O}\left(r^{-3}\right), \qquad (4.2)$$

$$g_{AB} = r^2 \overline{g}_{AB} + r \theta_{AB} + \overline{h}_{AB} + \frac{h_{AB}^{(2)}}{r} + \mathcal{O}\left(r^{-2}\right).$$
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## as it follows by applying standard canonical methods.

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# These nonlinear terms are crucial for understanding the invariance properties of the energy

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(2013 lecture notes on energy by Piotr T. Chruściel,  $\alpha = 1$ )

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one finds  $\overline{h}_{AB} = c^2 \overline{\gamma}_{AB}$ ,  $\theta_{AB} = 2c \overline{\gamma}_{AB}$ ,  $\overline{\lambda} = 0$ ,  $\overline{\lambda}^A = 0$ 

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(first term : ADM contribution ; second term : nonlinear contribution)

## Higher dimensions

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# As one increases the dimension, the analysis becomes more and more technically intricate

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Conclusions and comments As one increases the dimension, the analysis becomes more and more technically intricate

because the gap between the pure (improper) diffeomorphism piece in the expansion of the fields (generated by O(1) vector fields) and the "Coulomb" piece widens by one power of 1/r as one increases the dimension by one.

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Non linearities (of increasing order) then proliferate.

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Non linearities (of increasing order) then proliferate.

(Note that these non-linearities are not seen in current null infinity treatments which linearize the theory at infinity)

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(Note that these non-linearities are not seen in current null infinity treatments which linearize the theory at infinity)

Preliminary analysis indicates that the size of the BMS group does not increase because the new terms in the diffeomorphism generators define proper gauge transformations.

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Conclusions and comments In order to reveal the action of the BMS group at spatial infinity, one needs to include an improper gauge transformation term in the asymptotic conditions.

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Such a term cannot be set to zero,

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Such a term cannot be set to zero,

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One can then construct a completely consistent canonical formulation of the BMS symmetry,

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One can then construct a completely consistent canonical formulation of the BMS symmetry,

not only in four spacetime dimensions where the description is in complete agreement with the null infinity results (providing furthermore new light on the matching conditions),

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One can then construct a completely consistent canonical formulation of the BMS symmetry,

not only in four spacetime dimensions where the description is in complete agreement with the null infinity results (providing furthermore new light on the matching conditions),

but also in five dimensions where there is currently no null infinity description.

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## THANK YOU!