## Introduction

Einstein theory in
4 spacetime
dimensions
Einstein theory in
5 spacetime
dimensions
More on the
nonlinearities of
the charges
Higher

## dimensions

Conclusions and comments

# The BMS algebra at spatial infinity ( $D=4$ and $D>4$ ) 

Marc Henneaux

Eurostrings 2022, Lyon
26 April 2022

## Introduction

The BMS algebra

# Why should one study the asymptotic structure 

 of gravity at spatial infinity in the asymptotically flat context?
## Introduction

The BMS algebra

# Why should one study the asymptotic structure 

 of gravity at spatial infinity in the asymptotically flat context?Will give here three reasons.

## Introduction

The BMS algebra at spatial infinity ( $D=4$ and
$D>4$ )
Marc Henneaux

## Introduction

Einstein theory in
4 spacetime
First reason
dimensions

Einstein theory in
5 spacetime
dimensions

More on the nonlinearities of the charges

Higher
dimensions

Conclusions and comments

## Introduction

```
The BMS algebra
at spatial infinity
    ( }D=4\mathrm{ and
        D>4)
Marc Henneaux
```


## Introduction

Einstein theory in
4 spacetime
dimensions

Einstein theory in
5 spacetime
dimensions

More on the nonlinearities of the charges

Higher
dimensions

## Conclusions and

 comments
## First reason

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## Introduction

The BMS algebra

## First reason

In order to better understand the role of the BMS group (first identified at null infinity) in the quantum theory, where physical states are usually defined on spacelike (Cauchy) hypersurfaces, it is important to unveil its action on spacelike (Cauchy) hypersurfaces, and thus, at spatial infinity.

## Introduction

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ ) Marc Henneaux

## Introduction

Einstein theory in
4 spacetime
dimensions

Einstein theory in
5 spacetime
dimensions

More on the nonlinearities of the charges

Higher
dimensions
Conclusions and comments
(First reason, continued)

## Introduction

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux

## Introduction

Einstein theory in
4 spacetime
dimensions
Einstein theory in
5 spacetime
dimensions
More on the nonlinearities of the charges

Higher
dimensions
Conclusions and comments
(First reason, continued)
This has been done recently in four spacetime dimensions, through a reconsideration of the boundary conditions at spatial infinity.

## Introduction

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux

## Introduction

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New, consistent boundary conditions have been proposed, which are invariant under the full infinite-dimensional BMS group,

## Introduction

(First reason, continued)
This has been done recently in four spacetime dimensions, through a reconsideration of the boundary conditions at spatial infinity.
New, consistent boundary conditions have been proposed, which are invariant under the full infinite-dimensional BMS group, providing a standard, non-trivial, canonical realization of the BMS symmetry.

## Introduction

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux

## Introduction

(First reason, continued)
This has been done recently in four spacetime dimensions, through a reconsideration of the boundary conditions at spatial infinity.

New, consistent boundary conditions have been proposed, which are invariant under the full infinite-dimensional BMS group, providing a standard, non-trivial, canonical realization of the BMS symmetry.
This establishes also the important fact that the BMS symmetry is a symmetry of the theory and not just a symmetry at null infinity.

## Introduction

The BMS algebra at spatial infinity

Second reason ( $D=4$ and $D>4)$ Marc Henneaux

## Introduction

Einstein theory in
4 spacetime
dimensions

Einstein theory in
5 spacetime
dimensions

More on the nonlinearities of the charges

Higher
dimensions

Conclusions and comments

## Introduction

The BMS algebra

Second reason
Another reason for investigating the asymptotic structure at spatial infinity is the need to understand the "matching" conditions of the fields and charges between $\mathscr{I}_{-}^{+}$and $\mathscr{I}_{+}^{-}$

## Introduction

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux

## Introduction

Einstein theory in
4 spacetime
dimensions

Einstein theory in
5 spacetime
dimensions

More on the nonlinearities of the charges

## Second reason

Another reason for investigating the asymptotic structure at spatial infinity is the need to understand the "matching" conditions of the fields and charges between $\mathscr{I}_{-}^{+}$and $\mathscr{I}_{+}^{-}$ which clearly involves "going through" spatial infinity $i^{0}$.

## Introduction

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux

## Introduction

Einstein theory in
4 spacetime
dimensions

Einstein theory in
5 spacetime
dimensions

More on the nonlinearities of the charges

Higher
dimensions

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Another reason for investigating the asymptotic structure at spatial infinity is the need to understand the "matching" conditions of the fields and charges between $\mathscr{I}_{-}^{+}$and $\mathscr{I}_{+}^{-}$ which clearly involves "going through" spatial infinity $i^{0}$.

For instance, one knows that under very general initial conditions, the past limit of the Bondi $m_{B}$ mass along future null infinity is equal to the ADM mass $m$.

## Introduction

The BMS algebra at spatial infinity ( $D=4$ and D $>4$ )

Marc Henneaux

## Introduction

Einstein theory in
4 spacetime
dimensions

Einstein theory in
5 spacetime
dimensions

More on the nonlinearities of the charges

Higher
dimensions

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Another reason for investigating the asymptotic structure at spatial infinity is the need to understand the "matching" conditions of the fields and charges between $\mathscr{I}_{-}^{+}$and $\mathscr{I}_{+}^{-}$ which clearly involves "going through" spatial infinity $i^{0}$.
For instance, one knows that under very general initial conditions, the past limit of the Bondi $m_{B}$ mass along future null infinity is equal to the ADM mass $m$.
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## Introduction

The BMS algebra at spatial infinity ( $D=4$ and D $>4$ )

Marc Henneaux

## Introduction

Einstein theory in
4 spacetime
dimensions

Einstein theory in
5 spacetime
dimensions

More on the nonlinearities of the charges

Higher
dimensions
Conclusions and comments

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## Introduction

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux

## Introduction

Einstein theory in 4 spacetime dimensions

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## Introduction

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux

## Introduction

Einstein theory in 4 spacetime dimensions

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Can one say more?

## Introduction

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux

## Introduction

Einstein theory in 4 spacetime dimensions

Einstein theory in 5 spacetime dimensions

More on the nonlinearities of the charges

Higher dimensions comments

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But the generator of time translations is only one of the Bondi-Metzner-Sachs (BMS) supertranslation generators.
Can one say more?
This requires understanding the action of all the BMS supertranslations at spatial infinity.

## Introduction

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux

## Yet another reason :

## Introduction

Einstein theory in
4 spacetime
dimensions

Einstein theory in
5 spacetime
dimensions

More on the nonlinearities of the charges

Higher
dimensions

Conclusions and

## comments

## Introduction

The BMS algebra
$D>4$ )
Marc Henneaux

Introduction
Einstein theory in
4 spacetime
dimensions
Einstein theory in
5 spacetime
dimensions
More on the nonlinearities of the charges

Higher
dimensions
Conclusions and
comments

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In five dimensions, the definition of null infinity is problematical (as it is in all odd spacetime dimensions).

## Introduction

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But there exist soft theorems!

## Introduction

## Introduction

Einstein theory in
4 spacetime
dimensions

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In five dimensions, the definition of null infinity is problematical (as it is in all odd spacetime dimensions).
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Of which symmetries are these the Ward identities?

## Introduction

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux

## Introduction

Einstein theory in
4 spacetime
dimensions
Einstein theory in
5 spacetime
dimensions
More on the

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## Introduction

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It turns out that while the answer to this question is not immediate at null infinity, the analysis at spatial infinity raises no conceptual problem
and directly leads to the infinite dimensional symmetry " $\mathrm{BMS}_{5}$ group",

## Introduction

The BMS algebra at spatial infinity ( $D=4$ and D $>4$ )

Marc Henneaux

## Introduction

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and directly leads to the infinite dimensional symmetry " $\mathrm{BMS}_{5}$ group",
the realization of which exhibits (somewhat unexpectedly) a very interesting nonlinear structure.

## Introduction

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ ) Marc Henneaux

## Introduction

Einstein theory in 4 spacetime dimensions

Einstein theory in 5 spacetime dimensions

More on the nonlinearities of the charges

Higher
dimensions

## Conclusions and

 comments
## Introduction

The purpose of this talk is to provide the general ideas of the asymptotic analysis at spatial infinity in $D=4$ and $D=5$ spacetime dimensions.

## Introduction

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux

## Introduction

## Einstein theory in

4 spacetime
dimensions

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The study will be carried on spacelike hypersurfaces that are asymptotically flat hyperplanes, using Hamiltonian methods.

## Introduction

The purpose of this talk is to provide the general ideas of the asymptotic analysis at spatial infinity in $D=4$ and $D=5$ spacetime dimensions.
The study will be carried on spacelike hypersurfaces that are asymptotically flat hyperplanes, using Hamiltonian methods. (Work done in collaboration with Cédric Troessaert, with also Oscar Fuentealba, Sucheta Majumdar, Javier Matulich and Turmoli Neogi who joined more recently)

## Canonical action

The BMS algebra at spatial infinity ( $D=4$ and
$D>4$ )
Marc Henneaux

Introduction
Einstein theory in 4 spacetime dimensions

## Einstein theory in

 5 spacetime dimensionsMore on the nonlinearities of the charges

## Higher

dimensions

## Conclusions and

 comments
## Canonical action

The BMS algebra
$D>4$ )
Marc Henneaux
Introduction
Einstein theory in 4 spacetime dimensions

## Einstein theory in

5 spacetime
dimensions
More on the
dimensions

## Conclusions and

comments

## A central role in the analysis will be played by the gravitational action

## Canonical action

## Introduction

Einstein theory in 4 spacetime dimensions

## A central role in the analysis will be played by the gravitational action <br> which reads, in Hamiltonian form (Dirac, ADM),

## Canonical action

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux

## Introduction

Einstein theory in 4 spacetime dimensions

Einstein theory in
5 spacetime
dimensions
More on the nonlinearities of the charges

## Higher

dimensions

A central role in the analysis will be played by the gravitational action
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$$
S\left[g_{i j}, \pi^{i j}, N, N^{i}\right]=\int d t\left\{\int d^{d} x\left(\pi^{i j} \partial_{t} g_{i j}-N^{i} \mathscr{H}_{i}^{\text {grav }}-N \mathscr{H}^{\rho g r a v}\right)-B_{\infty}\right\}
$$

## Canonical action

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux

## Introduction

Einstein theory in 4 spacetime dimensions

Einstein theory in
5 spacetime
dimensions
More on the nonlinearities of the charges

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where $B_{\infty}$ is a boundary term at infinity and where

## Canonical action

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux

## Introduction

Einstein theory in 4 spacetime dimensions

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S\left[g_{i j}, \pi^{i j}, N, N^{i}\right]=\int d t\left\{\int d^{d} x\left(\pi^{i j} \partial_{t} g_{i j}-N^{i} \mathscr{H}_{i}^{g r a v}-N \mathscr{H}^{g r a v}\right)-B_{\infty}\right\}
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where $B_{\infty}$ is a boundary term at infinity and where

$$
\mathscr{H}^{\text {grav }}=-\sqrt{g} R+\frac{1}{\sqrt{g}}\left(\pi^{i j} \pi_{i j}-\frac{1}{d-1} \pi^{2}\right) \approx 0, \quad \mathscr{H}_{i}^{\text {grav }}=-2 \nabla_{j} \pi_{i}^{j} \approx 0
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## Canonical action

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux

## Introduction

## Einstein theory in

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$$

The definition of the theory is completed by providing boundary conditions on the dynamical variables (definition of phase space), which are assumed to make the off-shell action finite.

## Standard boundary conditions - Parity conditions

## Introduction

## Einstein theory in

 4 spacetime dimensions
## Einstein theory in

5 spacetime
dimensions
More on the nonlinearities of the charges

Higher
dimensions

## Conclusions and

comments

## Standard boundary conditions - Parity conditions

$D>4$ )
Marc Henneaux

## Introduction

Einstein theory in 4 spacetime dimensions

Einstein theory in
5 spacetime
dimensions
More on the nonlinearities of the charges

Higher
dimensions
Conctusions and comments

The standard boundary conditions for Einstein gravity in four spacetime dimensions are (Regge-Teitelboim 1974)

## Standard boundary conditions - Parity conditions

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux

## Introduction

Einstein theory in 4 spacetime dimensions

Einstein theory in
5 spacetime
dimensions
More on the

The standard boundary conditions for Einstein gravity in four spacetime dimensions are (Regge-Teitelboim 1974)

$$
h_{i j} \equiv g_{i j}-\delta_{i j}=\frac{\bar{h}_{i j}\left(\mathbf{n}^{k}\right)}{r}+O\left(\frac{1}{r^{2}}\right), \quad \bar{h}_{i j}\left(-\mathbf{n}^{k}\right)=\bar{h}_{i j}\left(\mathbf{n}^{k}\right)
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\pi^{i j}=\frac{\bar{\pi}^{i j}\left(\mathbf{n}^{k}\right)}{r^{2}}+O\left(\frac{1}{r^{3}}\right), \quad \bar{\pi}^{i j}\left(-\mathbf{n}^{k}\right)=-\bar{\pi}^{i j}\left(\mathbf{n}^{k}\right)
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## Standard boundary conditions - Parity conditions

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux

## Introduction

Einstein theory in 4 spacetime dimensions

Einstein theory in
5 spacetime
dimensions
More on the

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They involve strict parity conditions under the antipodal map $\mathbf{n}^{k} \rightarrow-\mathbf{n}^{k}$, where $\mathbf{n}^{k}$ is the unit normal to the sphere at infinity

## Standard boundary conditions - Parity conditions

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux

## Introduction

Einstein theory in 4 spacetime dimensions

Einstein theory in 5 spacetime dimensions

More on the nonlinearities of the charges

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They involve strict parity conditions under the antipodal map $\mathbf{n}^{k} \rightarrow-\mathbf{n}^{k}$, where $\mathbf{n}^{k}$ is the unit normal to the sphere at infinity $\left(f\left(\mathbf{n}^{k}\right) \equiv f(\theta, \varphi)\right)$.

# Parity conditions twisted by an improper diffeomorphism 

The BMS algebra at spatial infinity ( $D=4$ and<br>$D>4$ )<br>Marc Henneaux

## Introduction

## Einstein theory in

 4 spacetime dimensions
## Einstein theory in

5 spacetime
dimensions
More on the nonlinearities of the charges

Higher
dimensions

## Conctusions and

comments

# Parity conditions twisted by an improper diffeomorphism 

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )<br>Marc Henneaux<br>\section*{Introduction}<br>Einstein theory in 4 spacetime dimensions

To see the full BMS group, one must allow a "parity-twisted component" in the leading orders of the asymptotic metric and momenta.

## Parity conditions twisted by an improper diffeomorphism

The BMS algebra

Marc Henneaux

## Introduction

Einstein theory in

To see the full BMS group, one must allow a "parity-twisted component" in the leading orders of the asymptotic metric and momenta.

One thus imposes

$$
h_{i j} \equiv g_{i j}-\delta_{i j}=h_{i j}^{R T}+U_{i j}, \quad \pi^{i j}=\pi_{R T}^{i j}+V^{i j}
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$U_{i j}$ and $V^{i j}$ are the parity-twisted contributions that take the form of a gauge transformation (rewritten in Hamiltonian form).

## Parity conditions twisted by an improper diffeomorphism

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux

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$U_{i j}$ and $V^{i j}$ are the parity-twisted contributions that take the form of a gauge transformation (rewritten in Hamiltonian form). They are of the same order as the leading terms in $h_{i j}$ and $\pi^{i j}$ $\left(O(1 / r)\right.$ and $O\left(1 / r^{2}\right)$ respectively but have the opposite parity.

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$U_{i j}$ and $V^{i j}$ are the parity-twisted contributions that take the form of a gauge transformation (rewritten in Hamiltonian form).
They are of the same order as the leading terms in $h_{i j}$ and $\pi^{i j}$ $\left(O(1 / r)\right.$ and $O\left(1 / r^{2}\right)$ respectively but have the opposite parity.
For a canonical action of the Lorentz boosts, one imposes moreover that $U_{i j}$ be parametrized by an $O(1)$ odd function of the angles $\bar{U}\left(\mathbf{n}^{k}\right)=O(1)=-U\left(-\mathbf{n}^{k}\right)$. Furthermore, $V^{i j}$ is parametrized by an $O(1)$ even function of the angles $V\left(\mathbf{n}^{k}\right)=V\left(-\mathbf{n}^{k}\right)$.

## BMS group at spatial infinity

The BMS algebra at spatial infinity ( $D=4$ and
$D>4$ )
Marc Henneaux
Introduction
Einstein theory in 4 spacetime dimensions

## Einstein theory in

 5 spacetimedimensions
More on the nonlinearities of the charges

## Higher

dimensions

## Conclusions and

## comments

## BMS group at spatial infinity

The BMS algebra
$D>4$ )
Marc Henneaux

Introduction
Einstein theory in 4 spacetime dimensions

## Einstein theory in

5 spacetime
dimensions
More on the nonlinearities of the charges

Higher
dimensions
Conclusions and comments

## Do these relaxed parity conditions involving a twist lead to a consistent description

## BMS group at spatial infinity

The BMS algebra

## Introduction

Einstein theory in 4 spacetime dimensions

> Do these relaxed parity conditions involving a twist lead to a consistent description (finite symplectic form, well-defined generators)?

## BMS group at spatial infinity

The BMS algebra

## BMS group at spatial infinity

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux

## Introduction

Einstein theory in 4 spacetime dimensions

Einstein theory in
5 spacetime
dimensions
More on the nonlinearities of the charges

Higher

## dimensions

Do these relaxed parity conditions involving a twist lead to a consistent description
(finite symplectic form, well-defined generators)?
The answer is affirmative and requires some work (even though the idea is elementary).
One finds furthermore that the asymptotic symmetries are given by hypersurface deformations that behave asymptotically as

## BMS group at spatial infinity

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux

## Introduction

Einstein theory in 4 spacetime dimensions

Einstein theory in 5 spacetime dimensions

More on the nonlinearities of the charges

Higher

## dimensions

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\begin{aligned}
& \xi=b_{i} x^{i}+T(\mathbf{n})+O\left(r^{-1}\right) \\
& \xi^{i}=b_{j}^{i} x^{j}+W_{i}(\mathbf{n})+O\left(r^{-1}\right), \quad b_{i j}=-b_{j i}, \quad W_{i}(\mathbf{n})=\partial_{i}(r W(\mathbf{n}))
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## BMS group at spatial infinity

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux

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where $T$ is even and $W$ is odd.

## BMS group at spatial infinity

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux

## Introduction

Einstein theory in 4 spacetime dimensions

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\begin{aligned}
& \xi=b_{i} x^{i}+T(\mathbf{n})+O\left(r^{-1}\right) \\
& \xi^{i}=b_{j}^{i} x^{j}+W_{i}(\mathbf{n})+O\left(r^{-1}\right), \quad b_{i j}=-b_{j i}, \quad W_{i}(\mathbf{n})=\partial_{i}(r W(\mathbf{n}))
\end{aligned}
$$

where $T$ is even and $W$ is odd.
The terms $b_{i} x^{i}$ and $b^{i}{ }_{j} x^{j}$ describe respectively boosts and spatial rotations.

## BMS group at spatial infinity

Do these relaxed parity conditions involving a twist lead to a consistent description
(finite symplectic form, well-defined generators)?
The answer is affirmative and requires some work (even though the idea is elementary).
One finds furthermore that the asymptotic symmetries are given by hypersurface deformations that behave asymptotically as

$$
\begin{aligned}
& \xi=b_{i} x^{i}+T(\mathbf{n})+O\left(r^{-1}\right) \\
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where $T$ is even and $W$ is odd.
The terms $b_{i} x^{i}$ and $b^{i}{ }_{j} x^{j}$ describe respectively boosts and spatial rotations.
The zero mode of $T$ and the first spherical harmonic component of $W$ describe translations.

## BMS group at spatial infinity

The BMS algebra at spatial infinity ( $D=4$ and
$D>4$ )
Marc Henneaux
Introduction
Einstein theory in 4 spacetime dimensions

## Einstein theory in

 5 spacetimedimensions
More on the nonlinearities of the charges

## Higher

dimensions

## Conclusions and

## comments

## BMS group at spatial infinity

$D>4$ )
Marc Henneaux

## Introduction

The higher spherical harmonics describe general supertranslations.

Einstein theory in 4 spacetime dimensions

## Einstein theory in

5 spacetime
dimensions

More on the nonlinearities of the charges

Higher
dimensions

## Conclusions and

 comments
## BMS group at spatial infinity

The higher spherical harmonics describe general supertranslations.
In fact, the even function $T$ and the odd function $W$ combine to form a single arbitrary function of the angles, as in the null infinity description of the supertranslations.

## BMS group at spatial infinity

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux

## Introduction

Einstein theory in 4 spacetime dimensions

The higher spherical harmonics describe general supertranslations.
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## BMS group at spatial infinity

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux

## Introduction

Einstein theory in 4 spacetime dimensions

Einstein theory in
5 spacetime
dimensions
More on the nonlinearities of the charges

Higher

## dimensions

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$$
P_{\xi}\left[g_{i j}, \pi^{i j}\right]=\int d^{3} x\left(\xi \mathscr{H}+\xi^{i} \mathscr{H}_{i}\right)+\mathscr{B}_{\xi}\left[g_{i j}, \pi^{i j}\right]
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## BMS group at spatial infinity

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux

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## BMS group at spatial infinity

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux

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The algebra of the generators can be easily verified to be the BMS algebra.

## BMS group at spatial infinity

The BMS algebra at spatial infinity ( $D=4$ and
$D>4$ )
Marc Henneaux
Introduction
Einstein theory in 4 spacetime dimensions

## Einstein theory in

 5 spacetimedimensions
More on the nonlinearities of the charges

## Higher

dimensions

## Conclusions and

## comments

## BMS group at spatial infinity

$D>4$ )
Marc Henneaux

## Introduction

Einstein theory in 4 spacetime dimensions

## Einstein theory in

5 spacetime
dimensions
More on the nonlinearities of the charges

Higher
dimensions

## Conclusions and

comments

There is complete agreement with the null infinity results.

## BMS group at spatial infinity

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux

## Introduction

Einstein theory in
4 spacetime dimensions

Einstein theory in
5 spacetime
dimensions
More on the
nonlinearities of
the charges
Higher
dimensions
Conctusions and comments

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## BMS group at spatial infinity

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux

## Introduction

Einstein theory in 4 spacetime dimensions

Einstein theory in
5 spacetime
dimensions
More on the nonlinearities of the charges

Higher

## dimensions

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## BMS group at spatial infinity

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux

## Introduction

Einstein theory in 4 spacetime dimensions

Einstein theory in
5 spacetime
dimensions
More on the

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## BMS group at spatial infinity

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux

## Introduction

Einstein theory in 4 spacetime dimensions

Einstein theory in
5 spacetime
dimensions
More on the

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## BMS group at spatial infinity

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux
Introduction
Einstein theory in 4 spacetime dimensions

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Only "parity up to a gauge transformation" conditions are allowed
(and these are necessary to avoid log singularities at null infinity).

## 5 spacetime dimensions

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux

Introduction

Einstein theory in
4 spacetime
dimensions

Einstein theory in 5 spacetime dimensions

More on the nonlinearities of the charges

Higher
dimensions

## Conclusions and

 comments
## 5 spacetime dimensions

The BMS algebra
$D>4$ )
Marc Henneaux

## Introduction

Binstein theory in
4 spacetime
dimensions
Einstein theory in
5 spacetime
dimensions
More on the
nonlinearities of
the charges
Higher
dimensions
Conclusions and
comments

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## 5 spacetime dimensions

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As in 4 dimensions, one must include explicitly the improper gauge symmetries in the asymptotic form of the fields.
A new feature is that improper gauge terms are not at the same order in $\frac{1}{r}$ as the Coulomb part,
while in 4 dimensions, they are at the same order but distinguished by parity conditions.

## 5 spacetime dimensions

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux

Introduction
Einstein theory in
4 spacetime
dimensions

Einstein theory in 5 spacetime dimensions

More on the nonlinearities of the charges

Higher
dimensions
Conclusions and comments

## 5 spacetime dimensions

The BMS algebra

## Introduction

Einstein theory in
4 spacetime
dimensions
Einstein theory in 5 spacetime dimensions

More on the nonlinearities of the charges

Higher
dimensions
Conclusions and comments

In 5 dimensions, the improper gauge terms are at order $\frac{1}{r}$ (for the metric), corresponding to a diffeomorphism parameter of order $\mathscr{O}(1)$, whereas "the rest" is at order $r^{-2}$ (cf Schwarschild in 5D)

## 5 spacetime dimensions

The BMS algebra

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## 5 spacetime dimensions

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux

## Introduction

Einstein theory in
4 spacetime
dimensions
Einstein theory in 5 spacetime dimensions

More on the nonlinearities of the charges

## Higher

## dimensions

In 5 dimensions, the improper gauge terms are at order $\frac{1}{r}$ (for the metric), corresponding to a diffeomorphism parameter of order $\mathscr{O}(1)$, whereas "the rest" is at order $r^{-2}$ (cf Schwarschild in 5D) One has schematically :

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## 5 spacetime dimensions

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux

## Introduction

Einstein theory in
4 spacetime
dimensions
Einstein theory in 5 spacetime dimensions

More on the

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## 5 spacetime dimensions

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux

## Introduction

Einstein theory in
4 spacetime
dimensions
Einstein theory in 5 spacetime dimensions

More on the

In 5 dimensions, the improper gauge terms are at order $\frac{1}{r}$ (for the metric), corresponding to a diffeomorphism parameter of order $\mathscr{O}(1)$, whereas "the rest" is at order $r^{-2}$ (cf Schwarschild in 5D) One has schematically:

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## 5 spacetime dimensions

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux

## Introduction

Einstein theory in
4 spacetime
dimensions

Einstein theory in 5 spacetime dimensions

More on the nonlinearities of the charges

Higher
dimensions
Conclusions and comments

In 5 dimensions, the improper gauge terms are at order $\frac{1}{r}$ (for the metric), corresponding to a diffeomorphism parameter of order $\mathscr{O}(1)$, whereas "the rest" is at order $r^{-2}$ (cf Schwarschild in 5D) One has schematically :

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## 5 spacetime dimensions

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux

## Introduction

Einstein theory in 4 spacetime dimensions

Einstein theory in 5 spacetime dimensions

More on the nonlinearities of the charges

In 5 dimensions, the improper gauge terms are at order $\frac{1}{r}$ (for the metric), corresponding to a diffeomorphism parameter of order $\mathscr{O}(1)$, whereas "the rest" is at order $r^{-2}$ (cf Schwarschild in 5D) One has schematically:

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## 5 spacetime dimensions

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux

Introduction
Einstein theory in
4 spacetime
dimensions

Einstein theory in 5 spacetime dimensions

More on the nonlinearities of the charges

Higher
dimensions

## Conclusions and

 comments
## 5 spacetime dimensions

The BMS algebra
$D>4$ )
Marc Henneaux

## Introduction

Einstein theory in
4 spacetime
dimensions
Einstein theory in
5 spacetime
dimensions
More on the
nonlinearities of
the charges
Higher
dimensions
Conclusions and
comments

Even though the ideas are identical, there are striking new features in 5 dimensions :

## 5 spacetime dimensions

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ ) Marc Henneaux

## Introduction

Binstein theory in
4 spacetime
dimensions
Einstein theory in 5 spacetime dimensions

More on the nonlinearities of the charges

Higher
dimensions
Conctusions and comments

Even though the ideas are identical, there are striking new features in 5 dimensions :

- The size of $\mathrm{BMS}_{5}$ is bigger than expected. More precisely, the supertranslations depend on four functions of the angles.


## 5 spacetime dimensions

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux

## Introduction

Einstein theory in
4 spacetime
dimensions

Einstein theory in 5 spacetime dimensions

More on the nonlinearities of the charges

Higher dimensions

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## 5 spacetime dimensions

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux

## Introduction

Einstein theory in
4 spacetime
dimensions

Einstein theory in 5 spacetime dimensions

More on the nonlinearities of the charges

Higher
dimensions
Conclusions and comments

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## 5 spacetime dimensions

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux

## Introduction

Einstein theory in
4 spacetime
dimensions
Einstein theory in 5 spacetime dimensions

More on the nonlinearities of the charges

Higher
dimensions
Conclusions and comments

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## 5 spacetime dimensions

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux

## Introduction

Einstein theory in 4 spacetime dimensions

Einstein theory in 5 spacetime dimensions

More on the nonlinearities of the charges

Higher
dimensions
Conclusions and comments

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See O. Fuentealba, M. Henneaux, J. Matulich and C. Troessaert, e-Print : 2111.09664 [hep-th]


## Explicit expression of the energy

The BMS algebra at spatial infinity ( $D=4$ and
$D>4$ )
Marc Henneaux
Introduction
Einstein theory in
4 spacetime
dimensions
Einstein theory in 5 spacetime
dimensions
More on the nonlinearities of the charges

Higher
dimensions

## Conclusions and

comments

## Explicit expression of the energy

The BMS algebra
$D>4$ )
Marc Henneaux

## Introduction

Einstein theory in
4 spacetime
dimensions
Einstein theory in
5 spacetime
dimensions
More on the nonlinearities of the charges

## Higher

dimensions
Conctusions and
comments

Contrary to the familiar ADM expression for the energy in 4D,

## Explicit expression of the energy

## Introduction

Einstein theory in
4 spacetime
dimensions
Einstein theory in
5 spacetime
dimensions
More on the nonlinearities of the charges

## Higher

dimensions

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## Explicit expression of the energy

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux

## Introduction

Einstein theory in
4 spacetime
dimensions
Einstein theory in 5 spacetime dimensions

More on the nonlinearities of the charges

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Explicitly,

$$
E=2 \oint_{S_{\infty}^{3}} d^{3} x \sqrt{\bar{\gamma}}\left[(1 / 2) \bar{h}_{A}^{A}+D_{A} \bar{\lambda}^{A}+3 \bar{\lambda}-(1 / 8) \theta_{A B} \theta^{A B}\right]
$$

where

$$
\begin{align*}
& g_{r r}=1+\frac{2 \bar{\lambda}}{r^{2}}+\frac{h_{r r}^{(2)}}{r^{3}}+\mathscr{O}\left(r^{-4}\right),  \tag{4.1}\\
& g_{r A}=\frac{\bar{\lambda}_{A}}{r}+\frac{h_{r A}^{(2)}}{r^{2}}+\mathscr{O}\left(r^{-3}\right),  \tag{4.2}\\
& g_{A B}=r^{2} \bar{g}_{A B}+r \theta_{A B}+\bar{h}_{A B}+\frac{h_{A B}^{(2)}}{r}+\mathscr{O}\left(r^{-2}\right) . \tag{4.3}
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## Explicit expression of the energy

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux

## Introduction

Einstein theory in
4 spacetime
dimensions

Einstein theory in 5 spacetime
dimensions

More on the nonlinearities of the charges

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as it follows by applying standard canonical methods.

## Energy of flat space

The BMS algebra at spatial infinity ( $D=4$ and
$D>4$ )
Marc Henneaux
Introduction
Einstein theory in
4 spacetime
dimensions
Binstein theory in 5 spacetime
dimensions
More on the nonlinearities of the charges

Higher
dimensions
Conclusions and comments

## Energy of flat space

## Introduction

Einstein theory in
4 spacetime
dimensions
Einstein theory in
5 spacetime
dimensions
More on the nonlinearities of the charges

## Higher

dimensions
Conclusions and
comments

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## Energy of flat space

The BMS algebra
$D>4$ )
Marc Henneaux

## Introduction

Einstein theory in
4 spacetime
dimensions
Einstein theory in
5 spacetime
dimensions
More on the nonlinearities of the charges

## Higher

dimensions
Conclusions and
comments

These nonlinear terms are crucial for understanding the
invariance properties of the energy
under coordinate transformations that decay "slowly" at infinity.

## Energy of flat space

## Introduction

Einstein theory in
4 spacetime
dimensions
Einstein theory in
5 spacetime
dimensions
More on the nonlinearities of the charges

## Higher

dimensions

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(first term : ADM contribution; second term : nonlinear contribution)

## Higher dimensions

The BMS algebra at spatial infinity ( $D=4$ and
$D>4$ )
Marc Henneaux

## Introduction

Einstein theory in
4 spacetime
dimensions
Einstein theory in
5 spacetime
dimensions
More on the nonlinearities of the charges

Higher
dimensions
Conclusions and comments

## Higher dimensions

## Introduction

As one increases the dimension, the analysis becomes more and more technically intricate
Binstein theory in
4 spacetime
dimensions
Einstein theory in
5 spacetime
dimensions
More on the nonlinearities of the charges

Higher
dimensions

## Higher dimensions

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because the gap between the pure (improper) diffeomorphism piece in the expansion of the fields (generated by $O(1)$ vector fields) and the "Coulomb" piece widens by one power of $1 / r$ as one increases the dimension by one.
Non linearities (of increasing order) then proliferate.

## Higher dimensions

The BMS algebra at spatial infinity ( $D=4$ and $D>4$ )

Marc Henneaux

## Introduction

Einstein theory in 4 spacetime dimensions

As one increases the dimension, the analysis becomes more and more technically intricate
because the gap between the pure (improper) diffeomorphism piece in the expansion of the fields (generated by $O(1)$ vector fields) and the "Coulomb" piece widens by one power of $1 / r$ as one increases the dimension by one.
Non linearities (of increasing order) then proliferate.
(Note that these non-linearities are not seen in current null infinity treatments which linearize the theory at infinity)

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(Note that these non-linearities are not seen in current null infinity treatments which linearize the theory at infinity)
Preliminary analysis indicates that the size of the BMS group does not increase because the new terms in the diffeomorphism generators define proper gauge transformations.

## Conclusions and comments

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More on the nonlinearities of the charges

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## THANK YOU !

