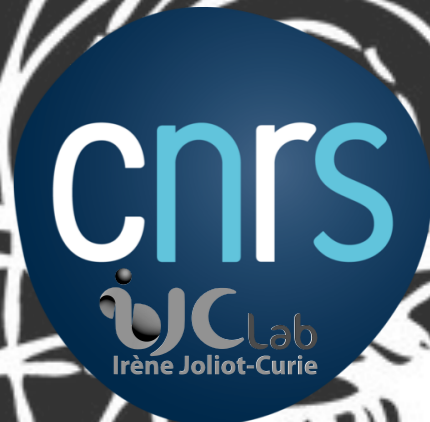


Neutrino Mass Ordering 5σ ?

IRN Meeting
June 2021



Anatael Cabrera
CNRS/IN2P3
IJCLab (Orsay)
LNCA (Chooz)



université
PARIS-SACLAY

FACULTÉ
DES SCIENCES
D'ORSAY

Université
de Paris

status on neutrino oscillation knowledge...

Standard Model (3 families)

[leptons & quarks]

&

unitary **PMNS**_{3x3}($\theta_{12}, \theta_{23}, \theta_{13}, \delta_{CP}$)

&

$\pm\Delta m^2$ & $+\delta m^2$

no conclusive sign of
any extension so far!!

(inconsistencies vs uncertainties)

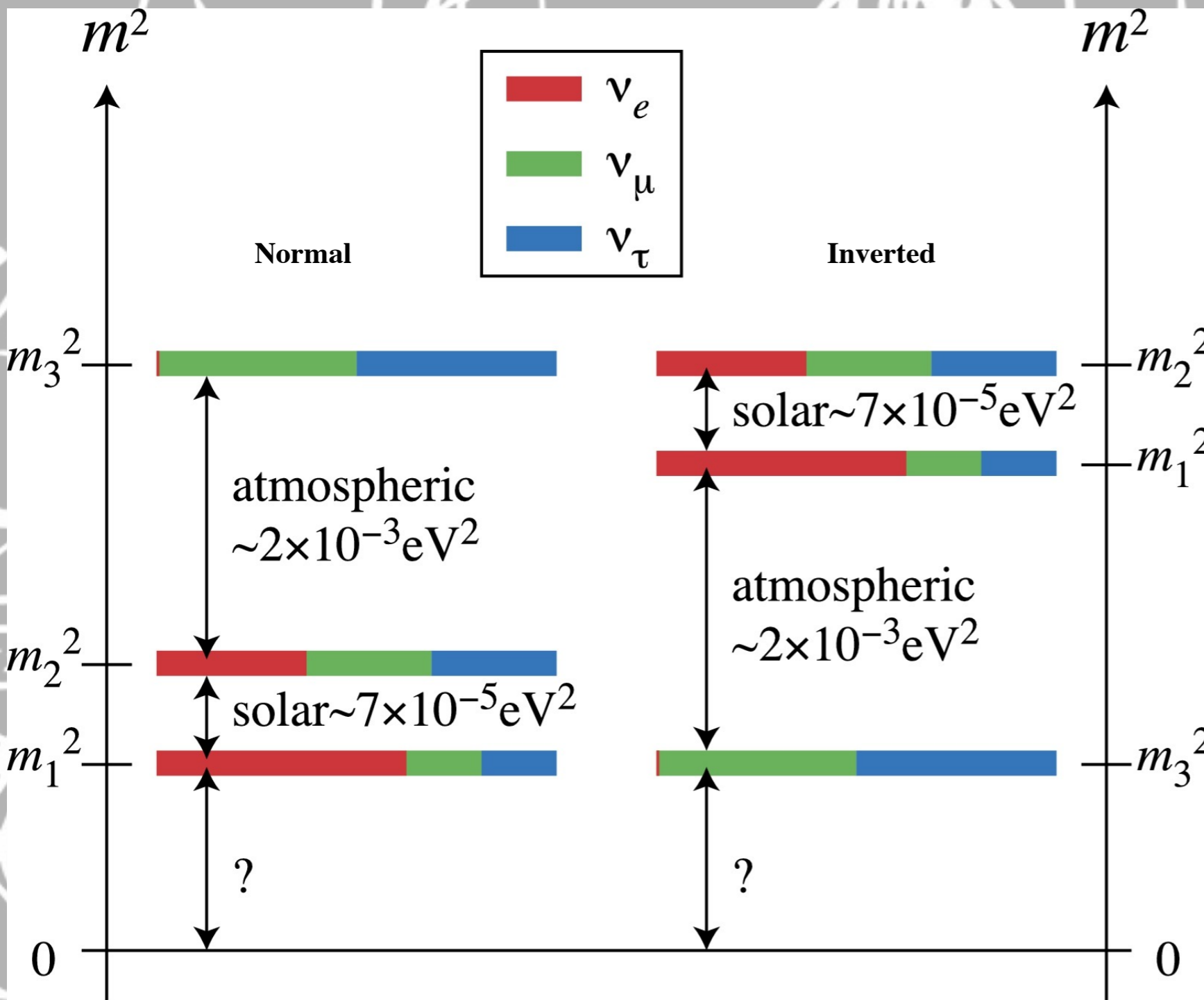
must measure all parameters → characterise & test (i.e. over-constrain) **Standard Model**

	today		
	best knowledge		NuFIT4.0
θ_{12}	3.0 %	SK⊕SNO	2.3 %
θ_{23}	5.0 %	NOvA+T2K	2.0 %
θ_{13}	1.8 %	DYB+DC+RENO	1.5 %
$+\delta m^2$	2.5 %	KamLAND	2.3 %
$ \Delta m^2 $	3.0 %	T2K+NOvA & DYB	1.3 %
Mass Ordering	unknown	SK et al	NMO ~3σ
CP Violation	unknown	T2K+NOvA	≈2σ

(now)

(reactor-beam)

soon JUNO⊕DUNE⊕HK will lead precision in the field → **sub-percent precision & CPV!**



Mass Ordering means...

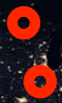
- **the lightest ν : m_1 or m_3 ?**
- **important consequences to...**
 - **Cosmology**
[ν role in Universe formation]
 - **Particle Physics**
[ex. $\beta\beta$ decay range]
- **discovery? test new physics!**
- **Standard Model: incomplete!**
[not known where it'd break first]

note: neutrino oscillations not sensitive to the ν absolute mass (other channels).

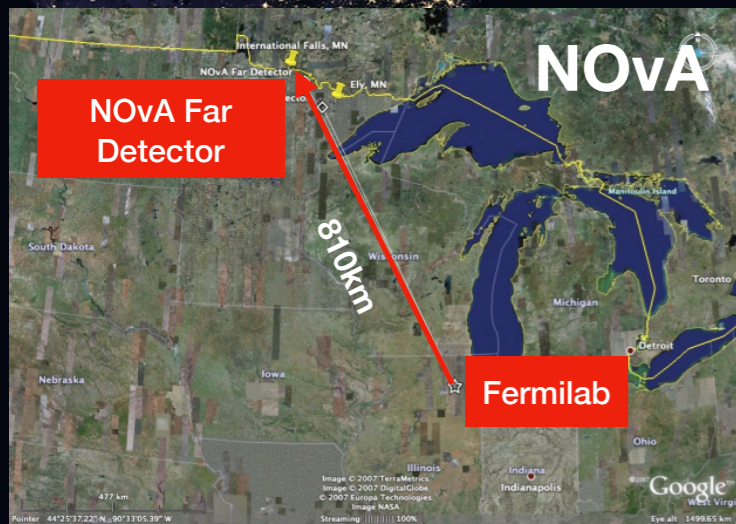
the Mass Ordering mystery...

running experiments...

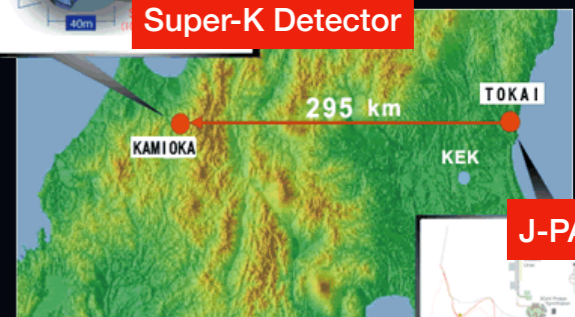
NOvA
(USA)



T2K
(Japan)



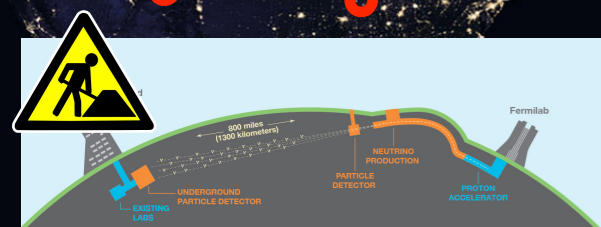
Super-K Detector



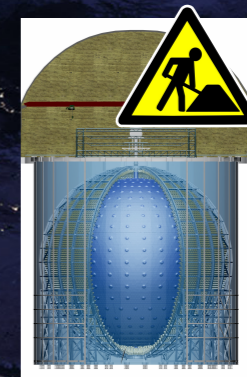
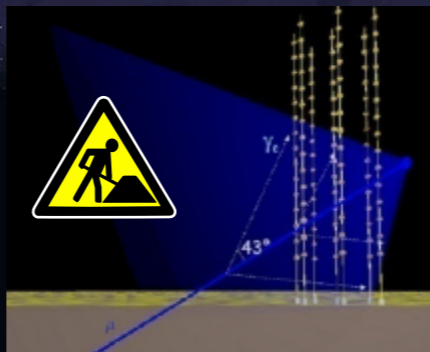
J-PARC

imminent experiments...

DUNE
(USA)



ORCA
(France)

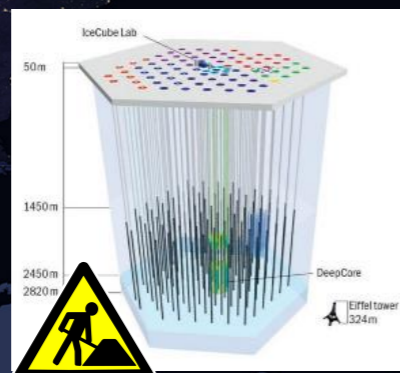
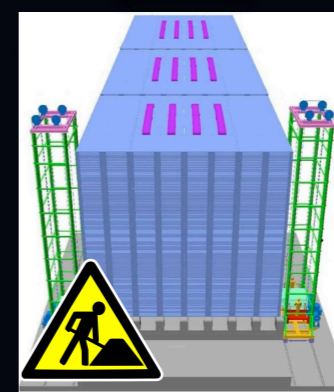


T2K → HyperK
(Japan)



JUNO
(China)

INO
(India)



IceCube → PINGU
(Antartica)

apologies: not all experiments mentioned



The fate of hints: updated global analysis of three-flavor neutrino oscillations

Ivan Esteban,^a M. C. Gonzalez-Garcia,^{a,b,c} Michele Maltoni,^d Thomas Schwetz,^e Albert Zhou^e

^aDepartament de Física Quàntica i Astrofísica and Institut de Ciències del Cosmos, Universitat de Barcelona, Diagonal 647, E-08028 Barcelona, Spain

^bInstitució Catalana de Recerca i Estudis Avançats (ICREA), Pg. Lluís Companys 23, 08010 Barcelona, Spain.

^cC.N. Yang Institute for Theoretical Physics, State University of New York at Stony Brook, Stony Brook, NY 11794-3840, USA

^dInstituto de Física Teórica UAM/CSIC, Calle de Nicolás Cabrera 13-15, Universidad Autónoma de Madrid, Cantoblanco, E-28049 Madrid, Spain

^eInstitut für Kernphysik, Karlsruher Institut für Technologie (KIT), D-76021 Karlsruhe, Germany

E-mail: ivan.esteban@fqa.ub.edu,

maria.gonzalez-garcia@stonybrook.edu, michele.maltoni@csic.es,

schwetz@kit.edu, albert.zhou@kit.edu

ABSTRACT: Our herein described combined analysis of the latest neutrino oscillation data presented at the Neutrino2020 conference shows that previous hints for the neutrino mass ordering have significantly decreased, and normal ordering (NO) is favored only at the 1.6σ level. Combined with the χ^2 map provided by Super-Kamiokande for their atmospheric neutrino data analysis the hint for NO is at 2.7σ . The CP conserving value $\delta_{CP} = 180^\circ$ is within 0.6σ of the global best fit point. Only if we restrict to inverted mass ordering, CP violation is favored at the $\sim 3\sigma$ level. We discuss the origin of these results – which are driven by the new data from the T2K and NOvA long-baseline experiments–, and the relevance of the LBL-reactor oscillation frequency complementarity. The previous 2.2σ tension in Δm_{21}^2 preferred by KamLAND and solar experiments is also reduced to the 1.1σ level after the inclusion of the latest Super-Kamiokande solar neutrino results. Finally we present updated allowed ranges for the oscillation parameters and for the leptonic Jarlskog determinant from the global analysis.

KEYWORDS: neutrino oscillations, solar and atmospheric neutrinos

today's world data leads to...

NMO favoured to $\sim 2.7\sigma$ (2020)

- **Super-Kamiokande** (most info so far)
- **1.6σ** (NOvA \oplus T2K & DC \oplus DYB \oplus RENO)

what are the leading experiments?

what's going to happen next?

arXiv:2007.14792v1 [hep-ph] 27 Jul 2020

today's NMO status...

Earliest Resolution to the Neutrino Mass Ordering?

Anatael Cabrera^{*1,2,4}, Yang Han^{†1,2}, Michel Obolensky¹, Fabien Cavalier², João Coelho², Diana Navas-Nicolás², Hiroshi Nunokawa^{‡2,7}, Laurent Simard², Jianming Bian³, Nitish Nayak³, Juan Pedro Ochoa-Ricoux³, Bedřich Roskovec³, Pietro Chimenti⁵, Stefano Dusini^{6a}, Marco Grassi^{6b}, Mathieu Bongrand^{8,2}, Rebin Karaparambil⁸, Victor Lebrin⁸, Benoit Viaud⁸, Frederic Yermia⁸, Lily Asquith⁹, Thiago J. C. Bezerra⁹, Jeff Hartnell⁹, Pierre Lasorak⁹, Jiajie Ling¹⁰, Jiajun Liao¹⁰, and Hongzhao Yu¹⁰

¹APC, CNRS/IN2P3, CEA/IRFU, Observatoire de Paris, Sorbonne Paris Cité University, 75205 Paris Cedex 13, France

²IJCLab., Université Paris-Saclay, CNRS/IN2P3, 91405 Orsay, France

³Department of Physics and Astronomy, University of California at Irvine, Irvine, California 92697, USA

⁴LNCA Underground Laboratory, CNRS/IN2P3 - CEA, Chooz, France

⁵Departamento de Física, Universidade Estadual de Londrina, 86051-990, Londrina – PR, Brazil

^{6a}INFN, Sezione di Padova, via Marzolo 8, I-35131 Padova, Italy

^{6b}Dipartimento di Fisica e Astronomia, Università di Padova, via Marzolo 8, I-35131 Padova, Italy

⁷Department of Physics, Pontifícia Universidade Católica do Rio de Janeiro, Rio de Janeiro, RJ, 22451-900, Brazil

⁸SUBATECH, CNRS/IN2P3, Université de Nantes, IMT-Atlantique, 44307 Nantes, France

⁹Department of Physics and Astronomy, University of Sussex, Falmer, Brighton BN1 9QH, United Kingdom

¹⁰Sun Yat-sen University, NO. 135 Xingang Xi Road, Guangzhou, China, 510275

August 27, 2020 – v3.5

when can we **resolve** ($\geq 5\sigma$) the **neutrino Mass Order**?

[earliest time scale]

which experiments (i.e. the minimal set) to yield the **full resolution**?

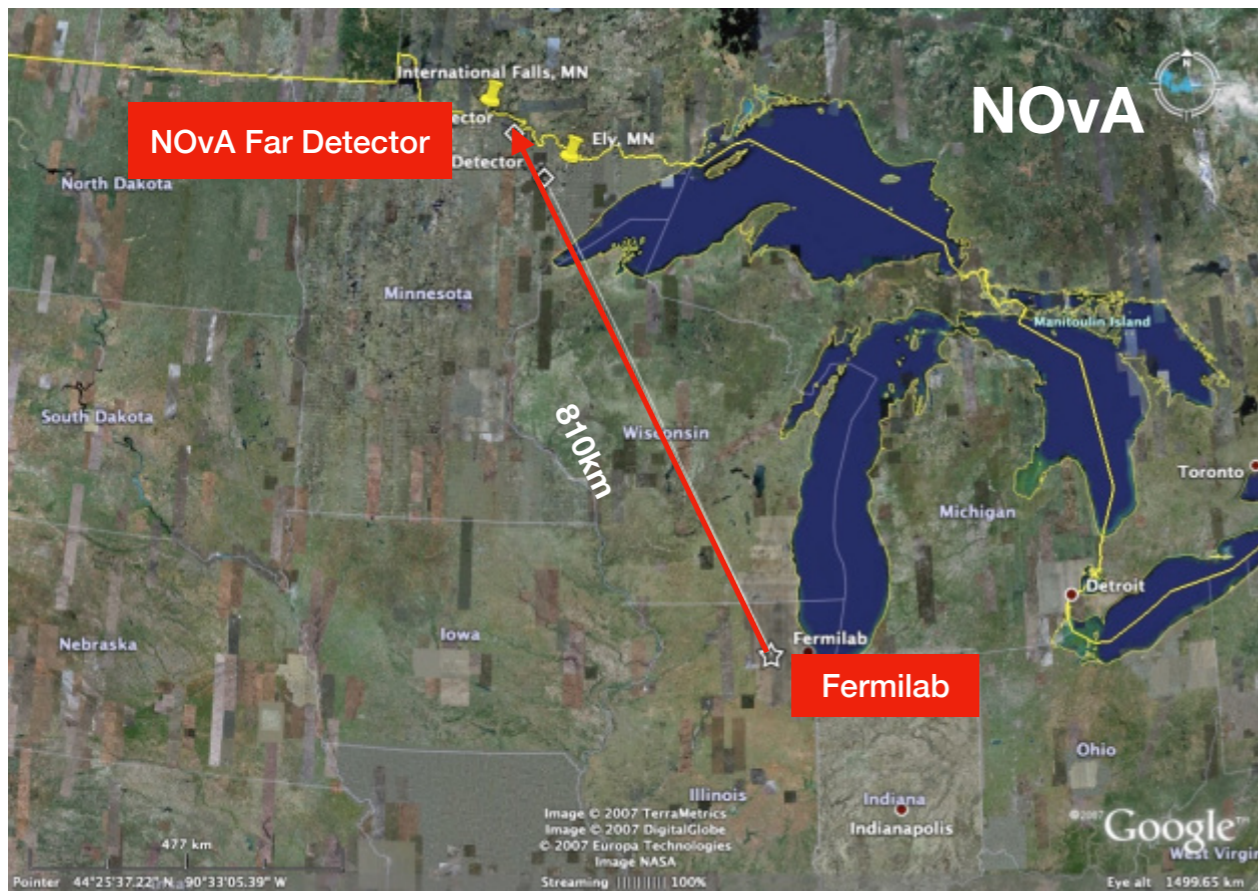
what physics exploited to yield the **full resolution**?

MO to probe **new physics?** (discovery potential)

our studies goal...

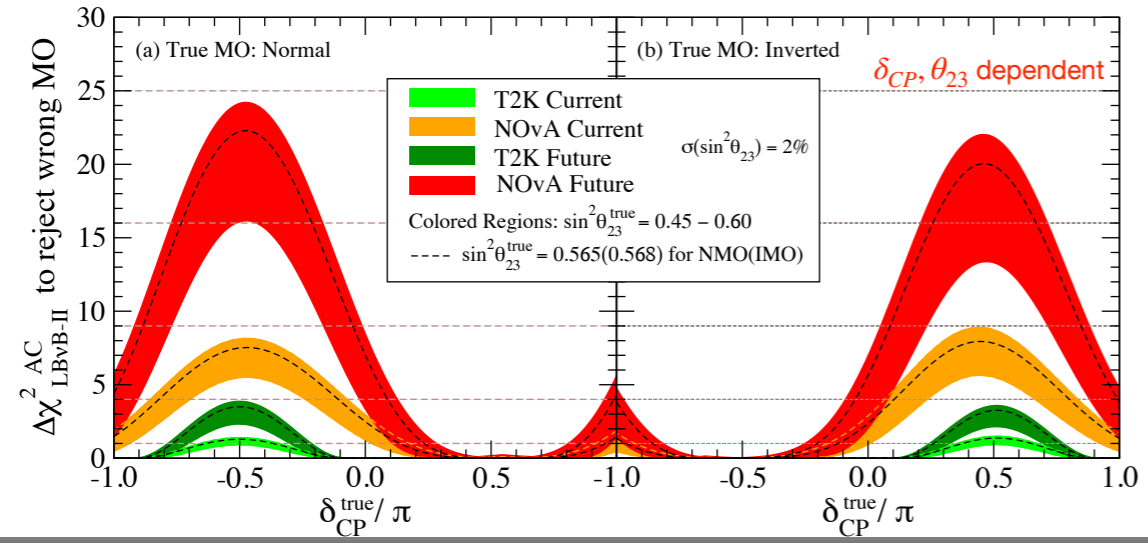
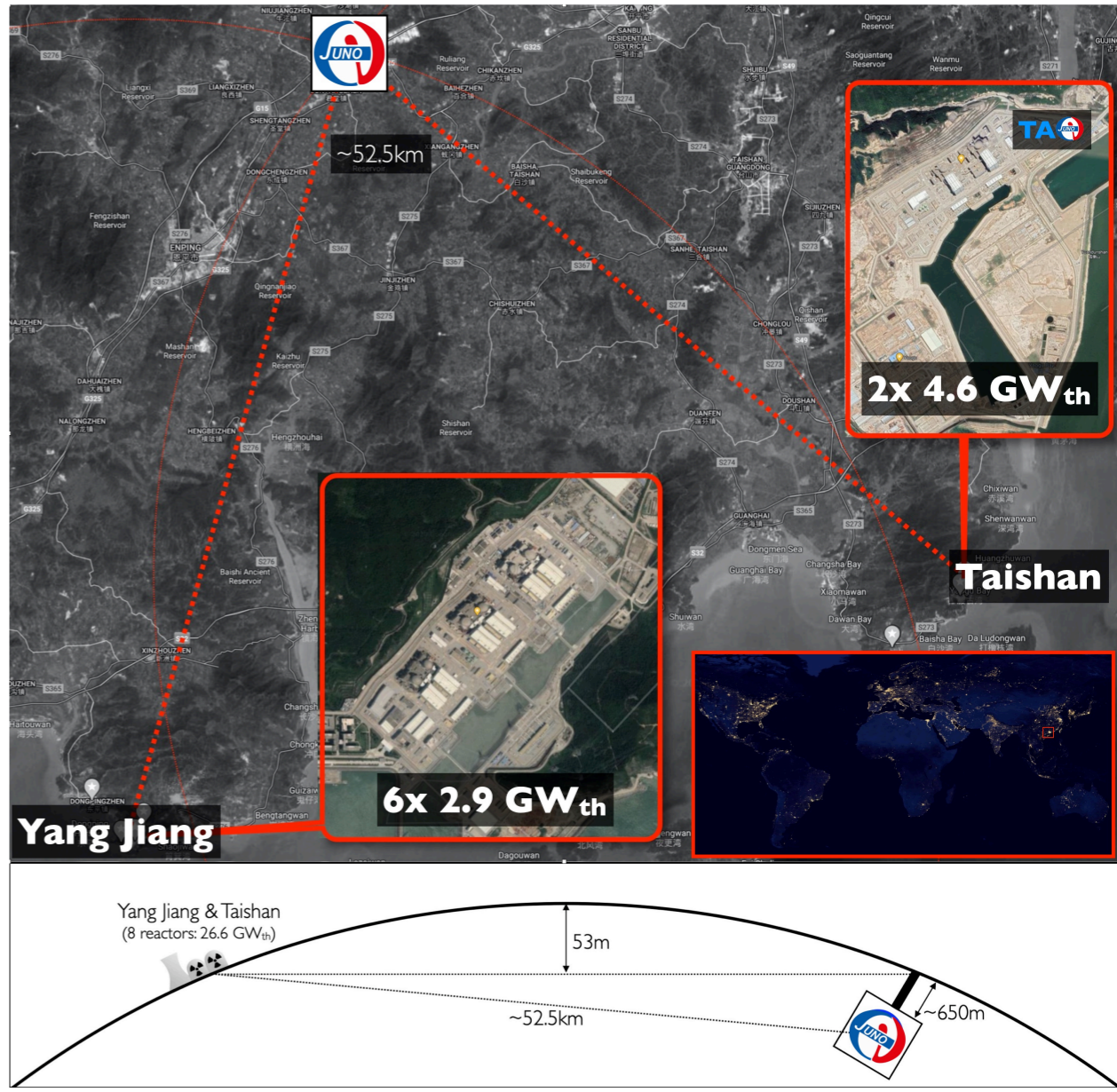
Matter Effects Oscillations

(CP experiments → **fake** CP-violation)



Vacuum Oscillations

(no CP-violation)

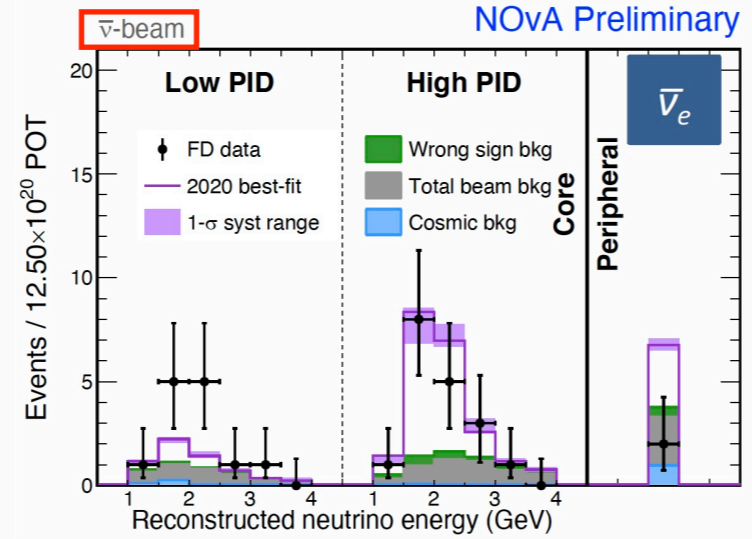
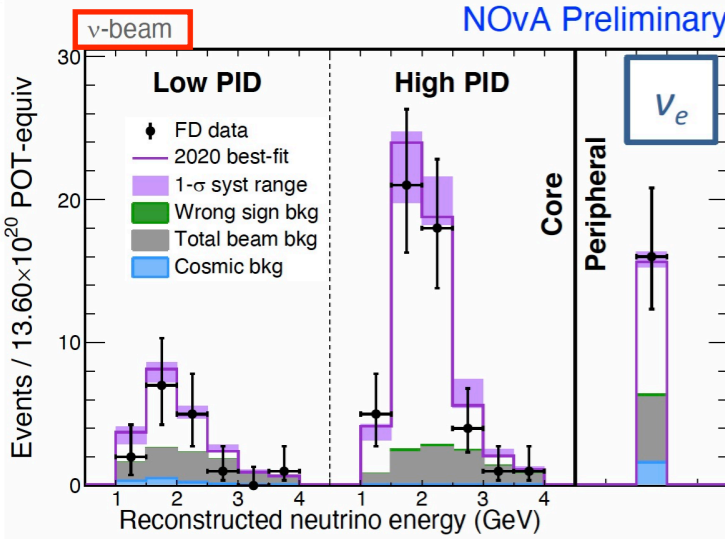


arXiv:2008.11280

only 2 ways to measure...

NOvA & T2K: direct comparison of oscillation with neutrino & anti-neutrino

ν_e and $\bar{\nu}_e$ Data at the Far Detector

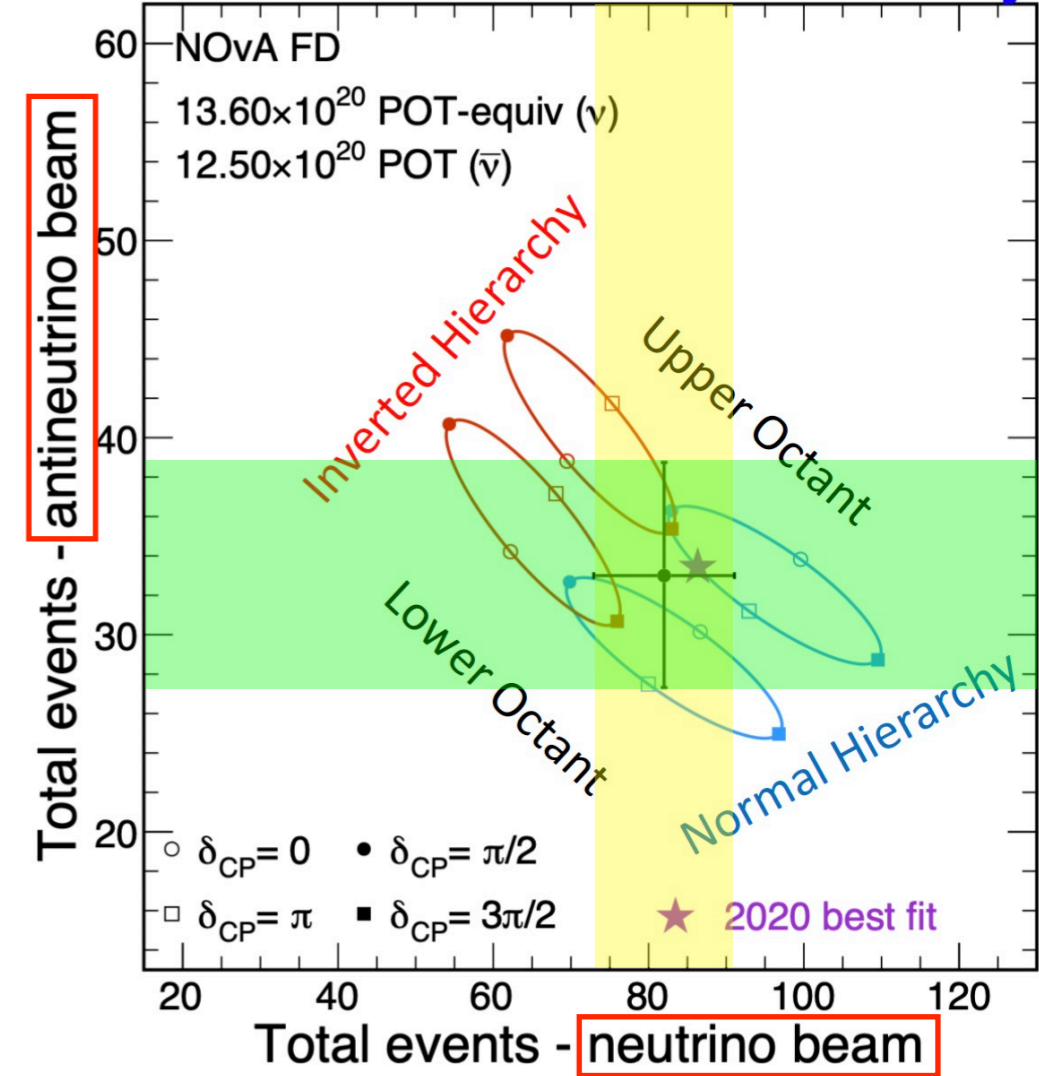


Total Observed	82	Range
Total Prediction	85.8	52-110
Wrong-sign	1.0	0.6-1.7
Beam Bkgd.	22.7	
Cosmic Bkgd.	3.1	
Total Bkgd.	26.8	26-28

Total Observed	33	Range
Total Prediction	33.2	25-45
Wrong-sign	2.3	1.0-3.2
Beam Bkgd.	10.2	
Cosmic Bkgd.	1.6	
Total Bkgd.	14.0	13-15

>4 σ evidence of $\bar{\nu}_e$ appearance

NOvA Preliminary

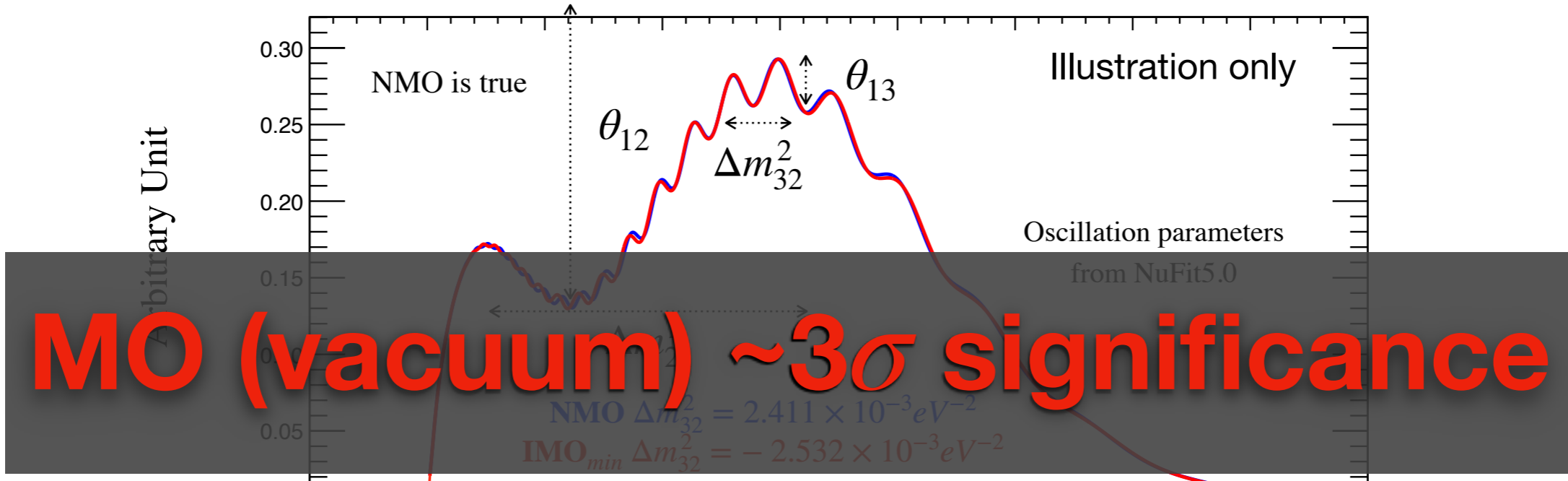


Appearance Channel [$\theta_{23} \oplus \theta_{13}, \delta_{CP}, \mathbf{MO}$]: $\nu_\mu \rightarrow \nu_e$ [ν and anti- ν]

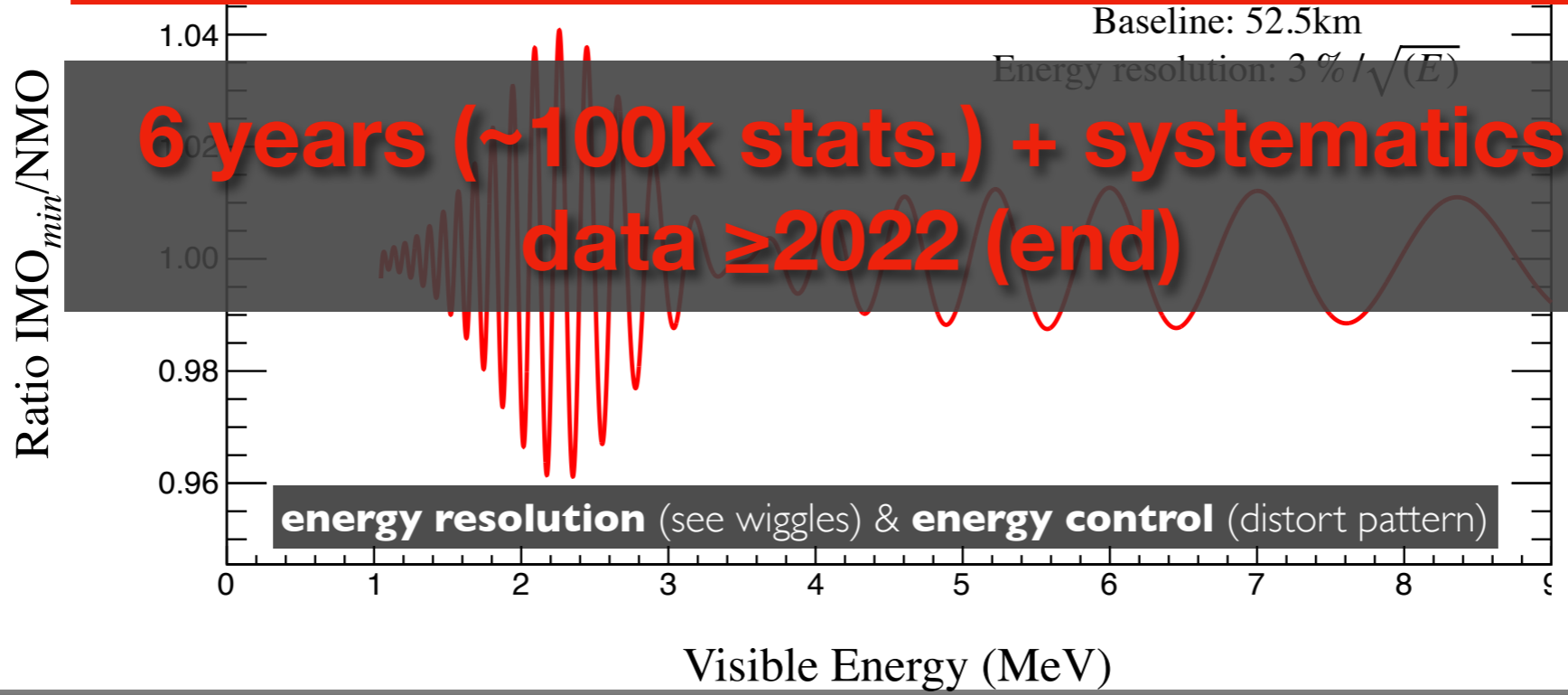
Disappearance Channel [$\theta_{23}, \Delta m^2_{32}$]: $\nu_\mu \rightarrow \nu_\mu$ "survival probability" (not shown)

NOvA/T2K observables...

JUNO ultra-precise oscillometry: 2 oscillations & interference terms (hard physics)



Disappearance Channel [$\theta_{12}, \delta m_{12}^2, \Delta m_{32}^2, MO - \theta_{13}$]: $\nu_e \rightarrow \nu_e$ [anti- ν]



the JUNO (hardest) way...

in 2020...

Super-Kamiokande — no!

T2K (≤ 2024) — no!

little $\leq 2\sigma \rightarrow$ T2K designed for cleanest $\delta(\text{CP})$

NOvA (≤ 2026) — no!

not bad !! $\leq 4\sigma$ (by 2026), only if lucky on $\delta(\text{CP})$!

JUNO (≥ 2022) — no!

not bad !! robust $\sim 3\sigma$ (by 2028) — careless of $\delta(\text{CP})$!!

by 2030...

DUNE ($\geq 2028?$) — yes!

stunning $> 5\sigma$ (by $\sim 2030?$) — careless of $\delta(\text{CP})$!!

Hyper-K ($\geq 2028?$) — no!

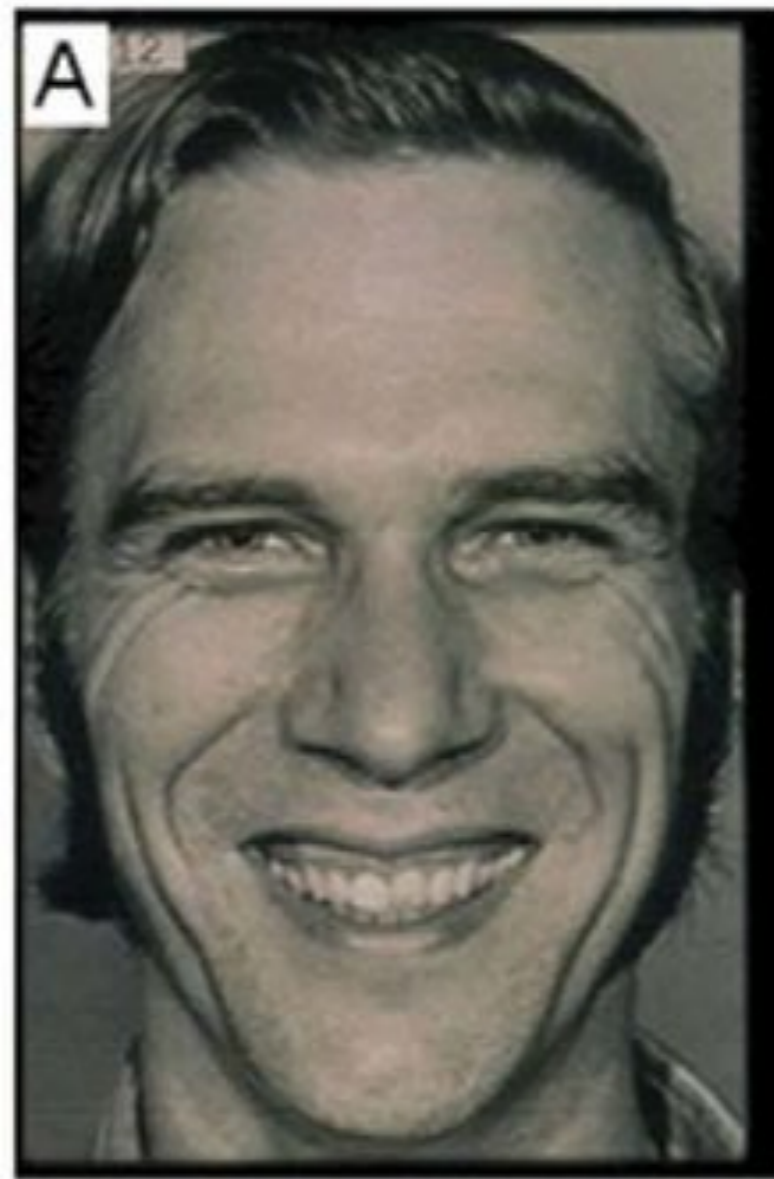
(like T2K) targets the cleanest $\delta(\text{CP})$ [minimal matter effects]

atmospheric neutrino — critical! (extra info maybe $\sim 5\sigma$)

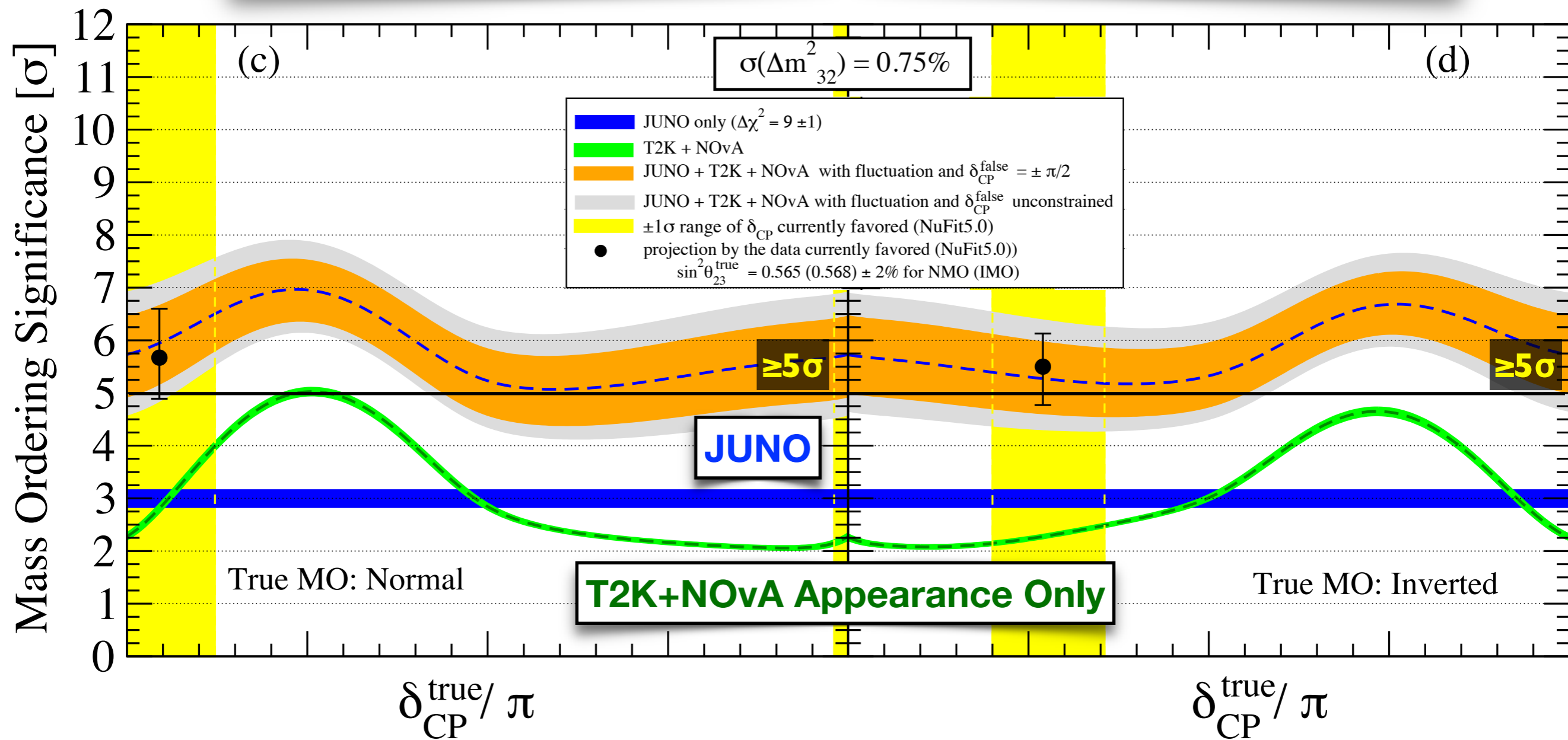
atmospheric not addressed in our analysis (complex) \rightarrow reinforce our conclusions

resolution ($\approx 5\sigma$) anybody...?

- **T2K Appearance (≤ 2024) — no!**
 - **NOvA Appearance (≤ 2026) — no!**
 - **JUNO (≥ 2022) — no!**
- \Rightarrow **T2K + NOvA + JUNO = yes? \rightarrow no!**
(just adding)
- \Rightarrow **T2K \oplus NOvA \oplus JUNO = yes!**
(synergies: appearance & disappearance)

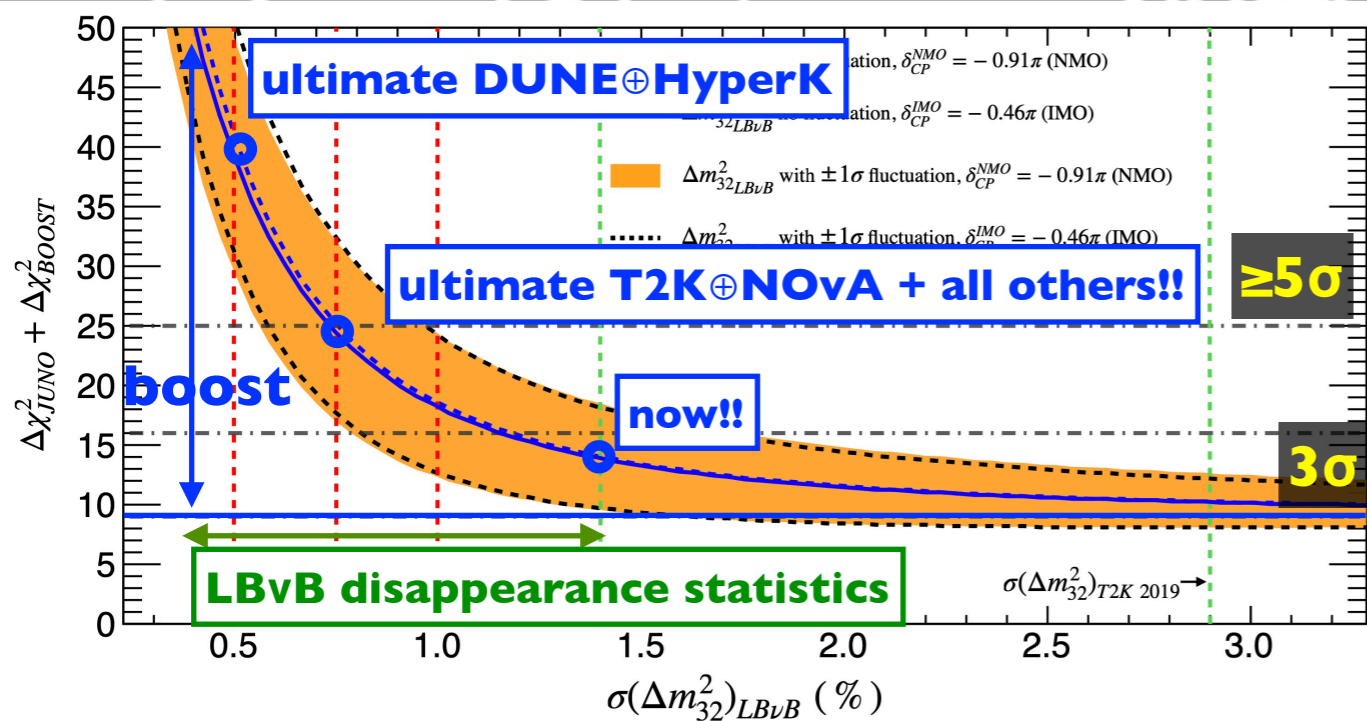


still, $\sim 5\sigma$ before 2030...

JUNO \oplus LB ν B-Disappearance [$\delta(\Delta m^2)=0.75\%$] \oplus LB ν B-Appearance


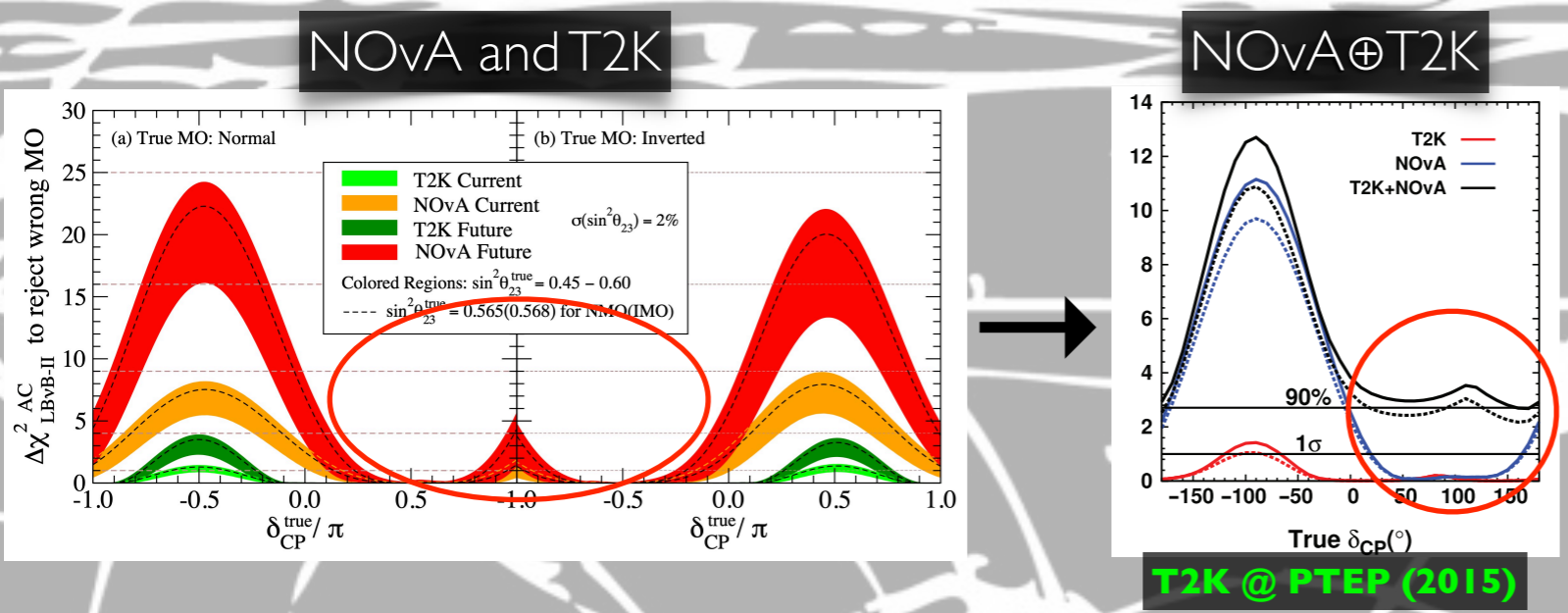
the power of synergies...

synergy I (JUNO vs NOvA ⊕ T2K): high precision disappearance Δm^2_{32} measurement



JUNO: unique vacuum oscillations (≥5σ!!!)

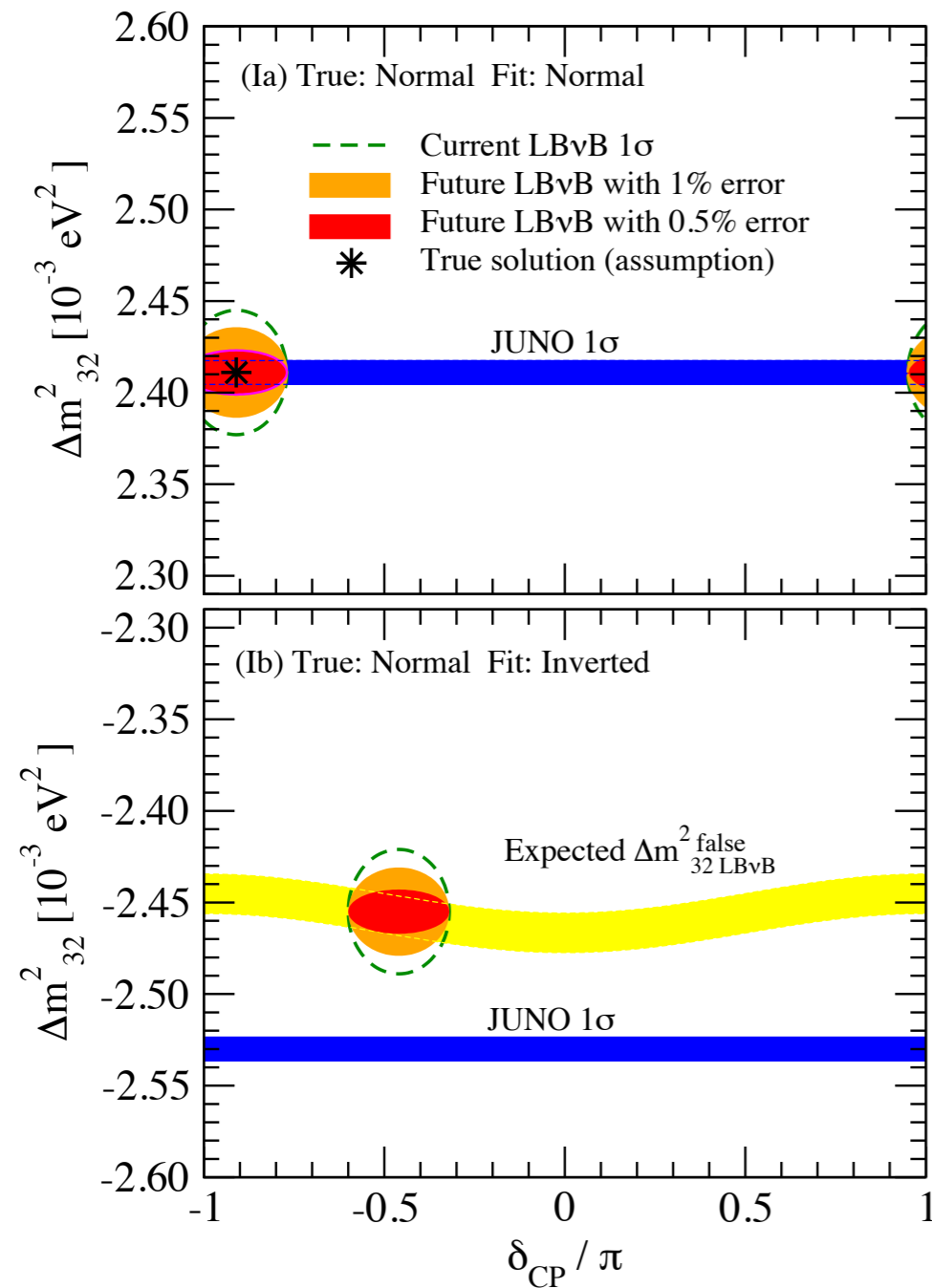
synergy II (NOvA vs T2K): MO ⊕ CPV complementary phase space discrimination



**NOvA: strong matter effects
T2K: clean CPV information**

arXiv:2008.11280

Mass Ordering: JUNO ⊕ NOvA ⊕ T2K...



$\Delta m_{32}^2[\text{reactor}]$ vs $\Delta m_{32}^2[\text{LBvB}]$

• **JUNO alone $\sim 3\sigma$**

standalone JUNO (intrinsic): “self Δm_{32}^2 ”

JUNO alone

• ≥ 2026 : **JUNO $\oplus \Delta m_{32}^2[\text{NOvA} \oplus \text{T2K}] \sim 5\sigma$**

JUNO exploits $\Delta m_{32}^2[\text{NOvA} \oplus \text{T2K}]$ via **PDG**

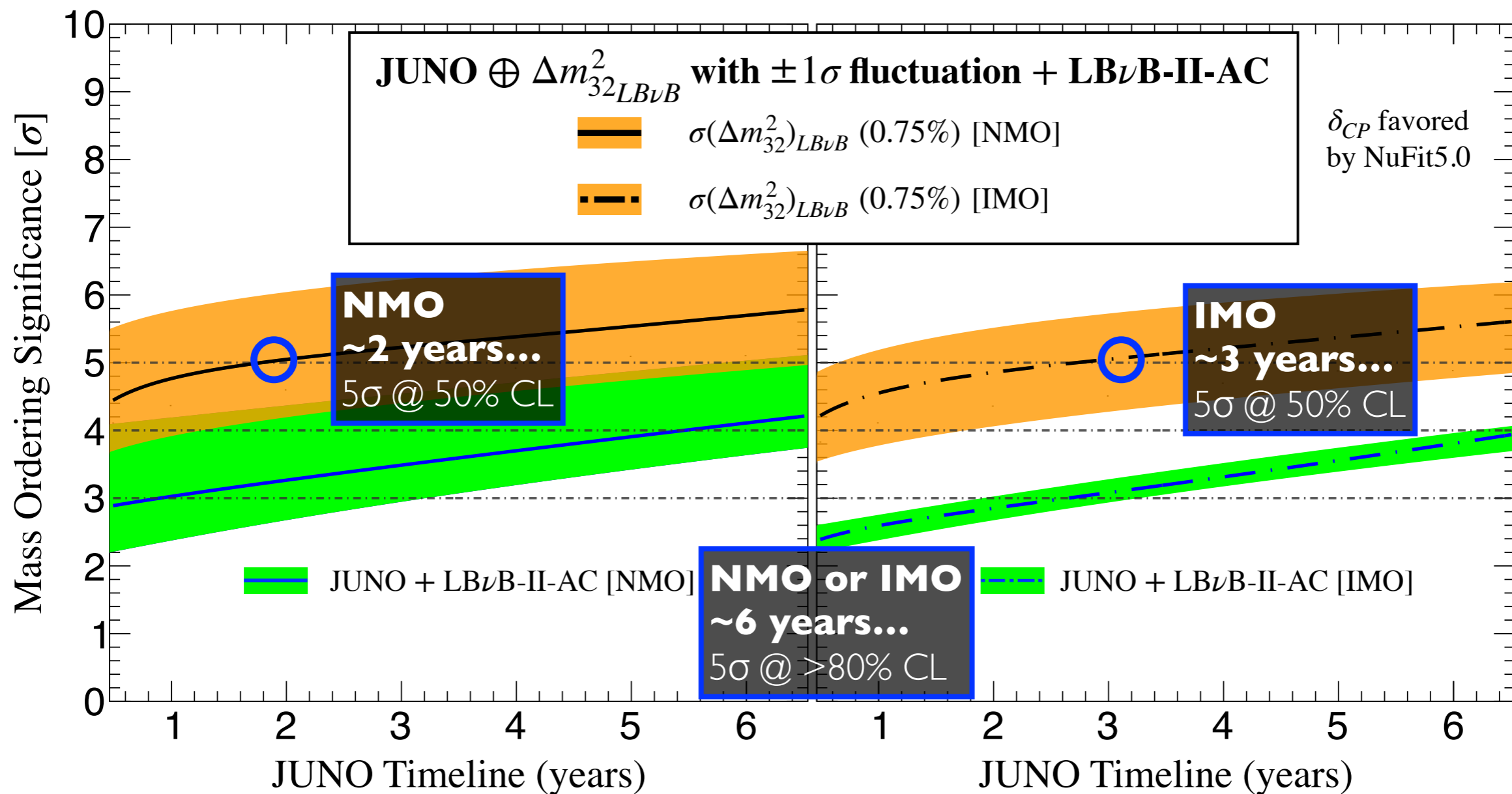
JUNO marginalised

• ≥ 2030 : **JUNO $\oplus \Delta m_{32}^2[\text{DUNE} \oplus \text{HyperK}] > 5\sigma$**

JUNO exploits $\Delta m_{32}^2[\text{DUNE} \oplus \text{HyperK}]$ via **PDG!**

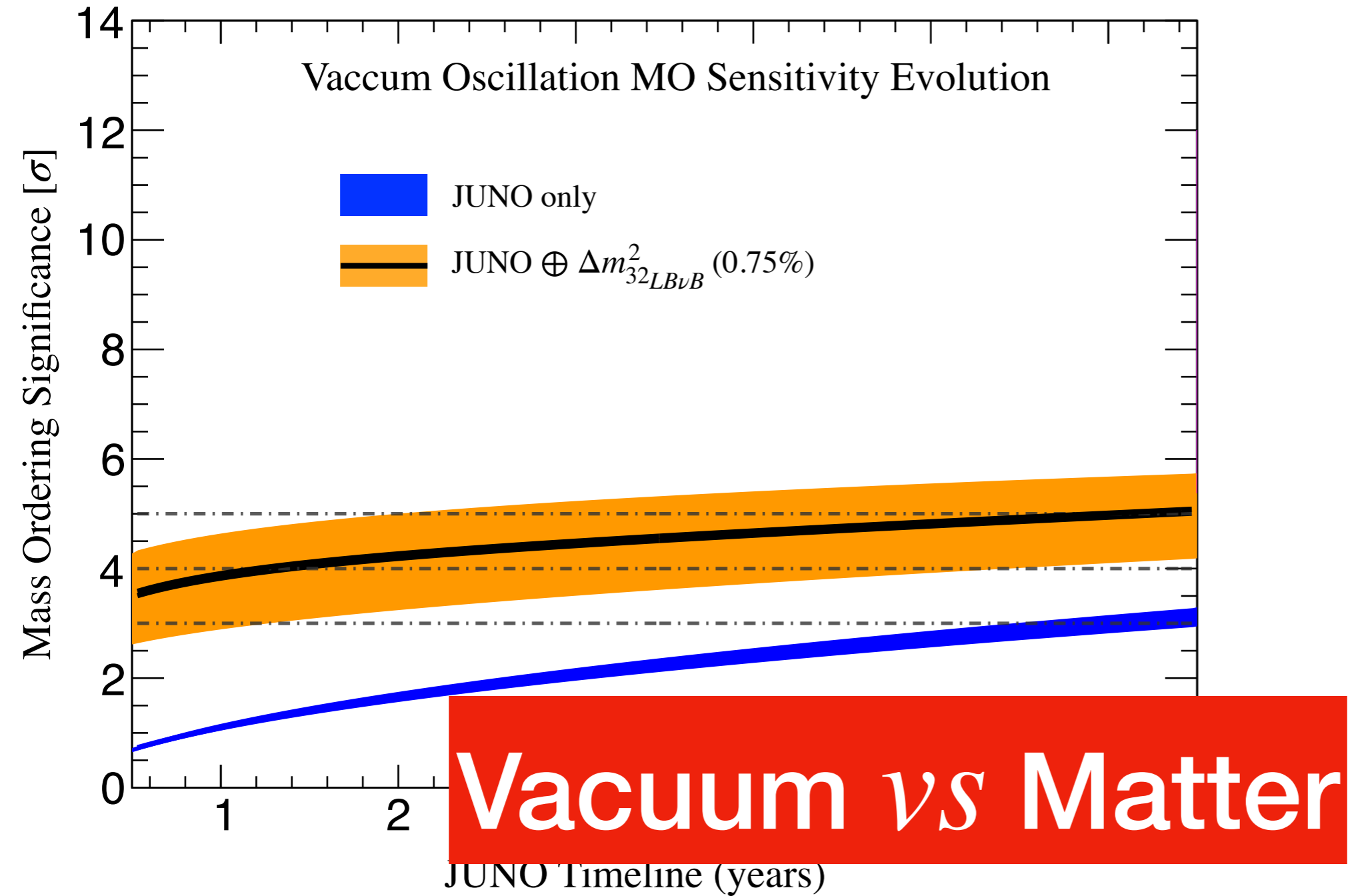
\Rightarrow **DUNE does not care about JUNO!! (too powerful)**

all about the Δm^2 synergy...



T2K data (2026) and NOvA data (2024) → release most precise Δm_{32}^2

~5 σ maybe even by $\geq 2026!!$ (if lucky)



first? MO @ $\geq 5\sigma$ possible ($\geq 90\%$ CL) — follow JUNO [2028]

discovery: physics BSM?

time evolution... **new physics?**

Earliest Resolution to the Neutrino Mass Ordering?

Anatael Cabrera^{*1,2,4}, Yang Han^{†1,2}, Michel Obolensky¹, Fabien Cavalier², João Coelho², Diana Navas-Nicolás², Hiroshi Nunokawa^{‡2,7}, Laurent Simard², Jianming Bian³, Nitish Nayak³, Juan Pedro Ochoa-Ricoux³, Bedřich Roskovec³, Pietro Chimenti⁵, Stefano Dusini^{6a}, Marco Grassi^{6b}, Mathieu Bongrand^{8,2}, Rebin Karaparambil⁸, Victor Lebrin⁸, Benoît Viaud⁸, Frederic Yermia⁸, Lily Asquith⁹, Thiago J. C. Bezerra⁹, Jeff Hartnell⁹, Pierre Lasorak⁹, Jiajie Ling¹⁰, Jiajun Liao¹⁰, and Hongzhao Yu¹⁰

¹APC, CNRS/IN2P3, CEA/IRFU, Observatoire de Paris, Sorbonne Paris Cité University, 75205 Paris Cedex 13, France

²IJCLab, Université Paris-Saclay, CNRS/IN2P3, 91405 Orsay, France

³Department of Physics and Astronomy, University of California at Irvine, Irvine, California 92697, USA

⁴LNCA Underground Laboratory, CNRS/IN2P3 - CEA, Chooz, France

⁵Departamento de Física, Universidade Estadual de Londrina, 86051-990, Londrina - PR, Brazil

^{6a}INFN, Sezione di Padova, via Marzolo 8, I-35131 Padova, Italy

^{6b}Dipartimento di Fisica e Astronomia, Università di Padova, via Marzolo 8, I-35131 Padova, Italy

⁷Department of Physics, Pontificia Universidade Católica do Rio de Janeiro, Rio de Janeiro, RJ, 22451-900, Brazil

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⁹Department of Physics and Astronomy, University of Sussex, Falmer, Brighton BN1 9QH, United Kingdom

¹⁰Sun Yat-sen University, NO. 135 Xingang Xi Road, Guangzhou, China, 510275

August 27, 2020 - v3.5

We hereby illustrate and numerically demonstrate via a simplified *proof of concept* calculation tuned to the latest average neutrino global data that the combined sensitivity of JUNO with NO ν A and T2K experiments has the potential to be the first fully resolved ($\geq 5\sigma$) measurement of neutrino Mass Ordering (MO) around 2028; tightly linked to the JUNO schedule. Our predictions account for the key ambiguities and the most relevant $\pm 1\sigma$ data fluctuations. In the absence of any concrete MO theoretical prediction and given its intrinsic binary outcome, we highlight the benefits of having such a resolved measurement in the light of the remarkable MO resolution ability of the next generation of long baseline neutrino beams experiments. We motivate the opportunity of exploiting the MO experimental framework to scrutinise the standard oscillation model, thus, opening for unique discovery potential, should unexpected discrepancies manifest. Phenomenologically, the deepest insight relies on the articulation of MO resolved measurements via at least the two possible methodologies *matter effects* and purely *vacuum* oscillations. Thus, we argue that the JUNO vacuum MO measurement may feasibly yield full resolution in combination to the next generation of long baseline neutrino beams experiments.

The discovery of *neutrino* (ν) *oscillations* phenomenon have completed a remarkable scientific endeavour lasting several decades that has changed forever our understanding of the phenomenology of the leptonic sector of the *standard model of elementary particles* (SM). A few modifications were accommodated to account for the new phenomenon [1]. This means the manifestation of massive neutrinos and leptonic mixing along with an embedded mechanism for the intrinsic difference between ν and $\bar{\nu}$ due to the violation of charge conjugation parity symmetry, or CP-violation (CPV); e.g. review [2].

Neutrino oscillations imply that the neutrino mass eigenstates (ν_1, ν_2, ν_3) spectrum is non-zero and non-degenerate, so at least two neutrinos are massive. Each mass eigenstate (ν_i ; with $i=1,2,3$) can be regarded as a

non-trivial mixture of the known neutrino flavour eigenstates (ν_e, ν_μ, ν_τ) linked to the three (e, μ, τ) respective charged leptons. Since no significant experimental evidence beyond three families exists so far, the mixing is characterised by the 3×3 so called *Pontecorvo-Maki-Nakagawa-Sakata* (PMNS) [3, 4] matrix, assumed unitary, thus parametrised by three independent mixing angles ($\theta_{12}, \theta_{23}, \theta_{13}$) and one CP phase (δ_{CP}). The neutrino mass spectra are indirectly known via the two measured *mass squared differences* indicated as $\delta m_{21}^2 (\equiv m_2^2 - m_1^2)$ and $\Delta m_{32}^2 (\equiv m_3^2 - m_2^2)$, respectively, related to the ν_2/ν_1 and ν_3/ν_2 pairs. The neutrino absolute mass is not directly accessible via neutrino oscillations and remains unknown, despite major active research [5].

As of today, the field is well established both exper-

*Contact: anatael@in2p3.fr

†Contact: yang.han@apc.in2p3.fr

‡Contact: nunokawa@puc-rio.br

Physics potentials with a combined sensitivity of T2K-II, NO ν A extension and JUNO

S. Cao,^{1,*} A. Nath,^{2,†} T. V. Ngoc,^{3,4,‡} Ng. K. Francis,^{2,§} N. T. Hong Van,^{5,3,¶} and P. T. Quyen^{3,4,**}

¹High Energy Accelerator Research Organization (KEK), Tsukuba, Ibaraki, Japan

²Department of Physics, Tezpur University, Assam, India

³Institute For Interdisciplinary Research in Science and Education (IFIRSE), Quy Nhon, Vietnam

⁴Graduate University of Science and Technology,

Vietnam Academy of Science and Technology (VAST), Hanoi, Vietnam

⁵Institute of Physics, Vietnam Academy of Science and Technology (VAST), Hanoi, Vietnam

Abstract

Leptonic CP violation search, neutrino mass hierarchy determination, and precision measurement of oscillation parameters for an unitary test of the neutrino mixing matrix are among the major targets of the ongoing and future neutrino oscillation experiments. The work explores the physics reach for these targets by around 2027, when the 3rd generation of the neutrino experiments starts operation, with a combined sensitivity of three experiments T2K-II, NO ν A extension, and JUNO. It is shown that a joint analysis of these three experiments can conclusively determine the neutrino mass hierarchy. Also, it provides 5σ C.L. more or less to exclude CP conserving values if *true* $\delta_{CP} \sim \pm \frac{\pi}{2}$ and more than 50% fractional region of *true* δ_{CP} values can be explored with a significance of at least 3σ C.L. Besides, the joint analysis can provide unprecedented precision measurements of the atmospheric neutrino oscillation parameters and a great offer to solve the θ_{23} octant degeneracy in case of non-maximal mixing.

* cvson@post.kek.jp

† ankur04@tezu.ernet.in

‡ tranngocap06@ifirse.icise.vn

§ francis@tezu.ernet.in

¶ nhvan@iop.vast.ac.vn

** phantoquyen97@gmail.com

our results (end of August 2020)

confirmation (end of September 2020)
[poster @ Nu2020]

validation \leftrightarrow agreement...

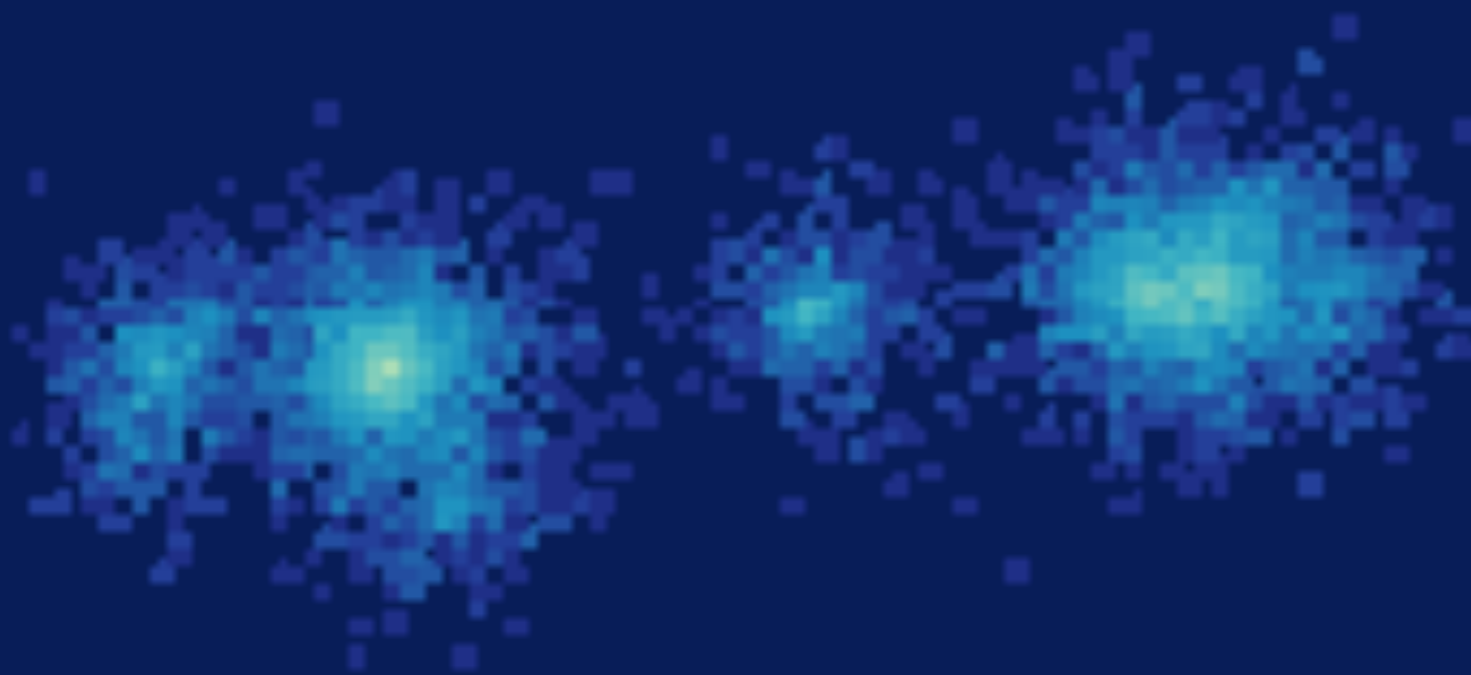
Neutrino Mass Ordering resolution...

- **fully resolved ($\geq 5\sigma$) by ≥ 2026 : **JUNO** \oplus **NOvA** \oplus **T2K****

(current and extra **atmospheric neutrino** \rightarrow even better)

- **first measurement** a mixture of **vacuum(JUNO)** \oplus **matter(NOvA, T2K, etc)**, **including atmospheric**

- ultimate **vacuum(JUNO) vs matter(DUNE): discovery?**



anatael@in2p3.fr

merci...

спасибі...

ありがとう...

danke...

고맙습니다...

obrigado...

Спасибо...

grazie...

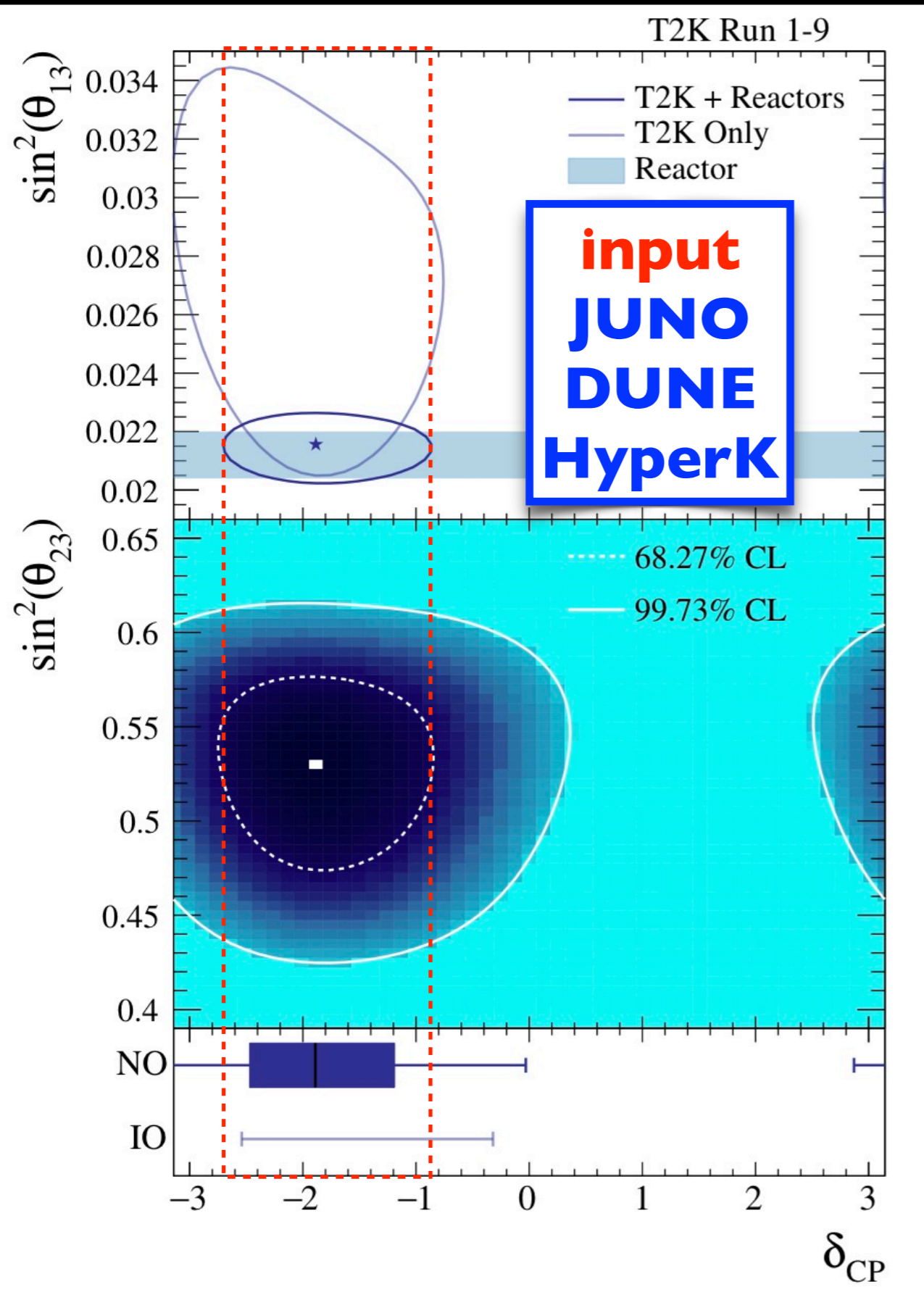
谢谢...

hvala...

gracias...

شكرا...

thanks...



θ_{13} implications

powerful constraint

CPV phase vs θ_{13}

[constrained by reactor]

CPV phase vs θ_{23}

[octant ambiguity]

CPV phase vs (Atmospheric) Mass Ordering