

FREEZE-IN PRODUCED DARK MATTER IN THE ULTRA-RELATIVISTIC REGIME

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in collaboration with Jacopo Ghiglieri (arXiv 2012.09083, JCAP03(2021)075)



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University
of Basel

Department of Physics

FRAMING THE DARK MATTER MODEL

- Simplified DM models:
 \Rightarrow capture the d.o.f. and parameters needed to study DM phenomenology
- χ Majorana fermion singlet, $\chi \equiv$ DM particle
- η is charged under QCD and $U(1)_Y$, $\eta \equiv$ mediator with $M_\eta = M + \Delta M$

$$\begin{aligned} \mathcal{L} = & \mathcal{L}_{\text{SM}} + \frac{1}{2} \bar{\chi} (i\not{\partial} - M) \chi + (D_\mu \eta)^\dagger D^\mu \eta - M_\eta^2 \eta^\dagger \eta - \lambda_2 (\eta^\dagger \eta)^2 \\ & - \lambda_3 \eta^\dagger \eta \phi^\dagger \phi - y \eta^\dagger \bar{\chi} a_R q - y^* \bar{q} a_L \chi \eta \end{aligned}$$

same model and freeze-in see M. Garny and J. Heisig (2018) and G. Bélanger et al (2018)

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$$y \lesssim \mathcal{O}(10^{-8}), T_{\text{INI.}} \gg M_\eta \text{ AND } T > 150 \text{ GEV}$$

- DM χ **never reaches** thermal equilibrium $f_\chi(t, \mathbf{k}) \ll n_F(k^0)$
- η and q maintained **in equilibrium** by SM interactions
- χ accumulates over the thermal history through processes like $\eta \rightarrow \chi q$

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- we shall address the high-temperature dynamics
- multiple soft scatterings and $2 \rightarrow 2$ process

$T \gg M_\eta$ can be very important even for renormalizable interactions

PRODUCTION RATE AND RATE EQUATION

GENERAL APPROACH

- Given a field χ **weakly coupled** to a an **equilibrated bath**, with internal couplings g
[T. Asaka, M. Laine and M. Shaposhnikov (2006), M. Laine and A. Vuorinen (2017), D. Bödeker, M. Sangel and M. Wörmann (2016)]
- at leading order in y and **all orders in g** one can prove D. Bödeker, M. Sangel and M. Wörmann (2016)

$$\dot{f}_\chi(t, \mathbf{k}) = \Gamma(k)[n_F(k^0) - f_\chi(t, \mathbf{k})], \quad \Gamma(k) = \frac{|y|^2}{2k^0} \int d^4X e^{iK \cdot X} \langle [J(X), J(0)] \rangle$$

- $f_\chi(t, \mathbf{k})$ is the single-particle phase-space distribution; J **made of bath fields**



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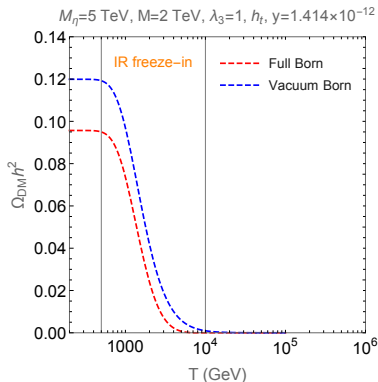
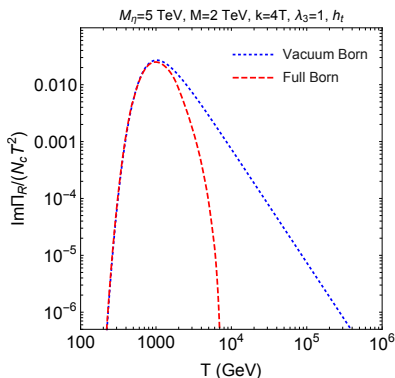
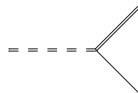
- when doing perturbative expansions \Rightarrow **Boltzmann equation is recovered**
- general framework to include:
resummation and NLO computations, non-perturbative and thermal effects

BORN RATE AND THERMAL MASSES

- Born rate $\eta \rightarrow \chi + q$ **with** and **without** thermal masses (recall m_q is purely thermal)

for thermal masses see also L. Darmé, A. Hryczuk, D. Karamitros and L. Roszkowski (2019)

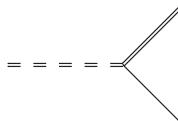
$$\dot{n}_{\text{DM}} + 3Hn_{\text{DM}} = 2|y|^2 \int_{\mathbf{k}} \frac{n_{\text{F}}(k^0)}{k^0} \text{Im}\Pi_R^{\text{Born}}$$



LPM RESUMMATION FOR LIGHT-CONE KINEMATICS

HIGH-TEMPERATURES $\pi T \gg M_\eta$

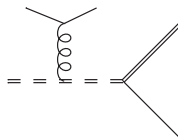
- all the particles are seen as *massless*,
- momenta of external particles
 $p \sim \pi T$,
- particles are ultra-relativistic, and gT
is a soft scale
 \Rightarrow **collinear kinematics \approx high T**



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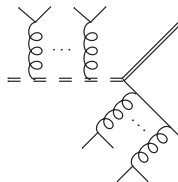
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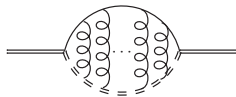
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n soft scatterings: LPM resummation

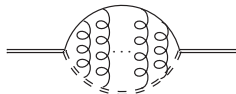
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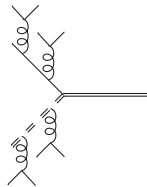
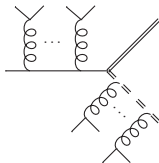
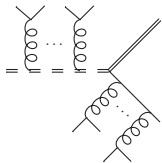
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- at $T \gg M_\eta$ three effective processes contribute to the production of χ [$1 + n \leftrightarrow 2 + n$]

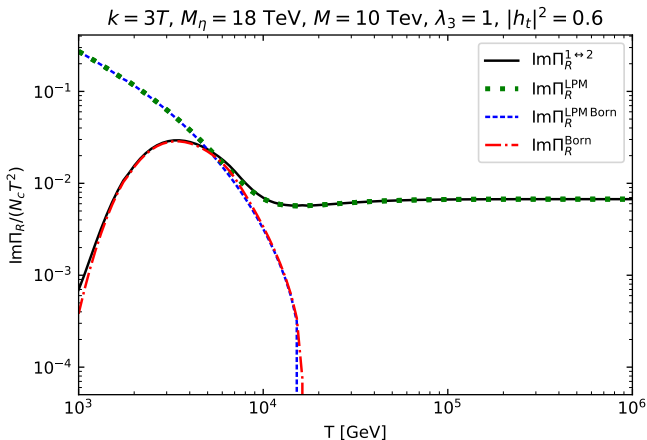
$$\eta \rightarrow \chi + q, \quad q \rightarrow \chi + \eta, \quad q + \eta \rightarrow \chi$$



LPM RESULTS

- prescription for any temperature (see I. Ghisoiu and M. Laine (2014))

$$\text{Im}\Pi_R^{1\leftrightarrow 2} = \text{Im}\Pi_R^{\text{LPM}} - \text{Im}\Pi_R^{\text{LPM Born}} + \text{Im}\Pi_R^{\text{Born}}$$



$2 \rightarrow 2$ SCATTERINGS

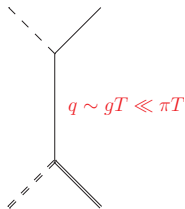
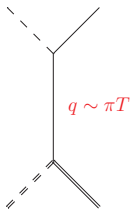
$$\mathcal{M}_a = \text{[Diagram 1]} + \text{[Diagram 2]} \quad \mathcal{M}_b = \text{[Diagram 3]} + \text{[Diagram 4]}$$

- Considered by M. Garny and J. Heisig (1809.10135) for $T \leq M$ (possibly some issues with IR of some processes)
- we look at $\pi T \gg M_\eta, m_\eta, m_q$: for s/t and u/t contributions from **both hard and soft momentum regions** D. Besak and D. Bodeker (2012), J. Ghiglieri and M. Laine (2016)

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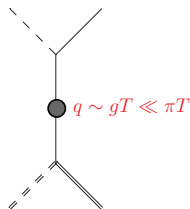
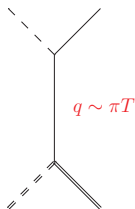
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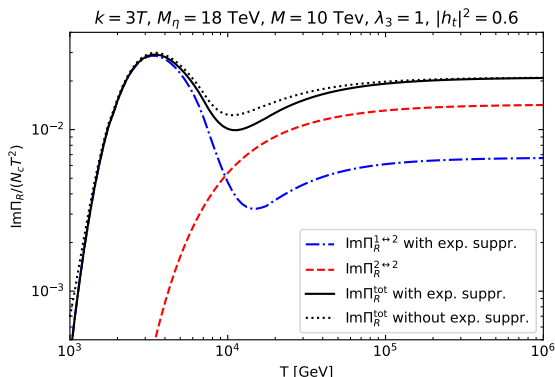
$$\begin{aligned} \text{Im}\Pi_R^{2\leftrightarrow 2} &= \frac{2}{(4\pi)^3 k} \int_k^\infty dq_+ \int_0^k dq_- \left\{ [n_F(q_0) + n_B(q_0 - k)] N_c (Y_q^2 g_1^2 + C_F g_3^2 + |h_q|^2) \Phi_{s2} \right\} \\ &+ \frac{2}{(4\pi)^3 k} \int_0^k dq_+ \int_{-\infty}^0 dq_- \left\{ [1 - n_F(q_0) + n_B(k - q_0)] N_c (Y_q^2 g_1^2 + C_F g_3^2 + |h_q|^2) \Phi_{t2} \right\} \\ &- \left[n_B(k) + \frac{1}{2} \right] N_c (Y_q^2 g_1^2 + C_F g_3^2 + |h_q|^2) \frac{k\pi^2 T^2}{q^2} \left\} + N_c \frac{m_q^2}{16\pi} \left[n_B(k) + \frac{1}{2} \right] \ln \left(1 + \frac{4k^2}{m_q^2} \right) \end{aligned}$$

SUMMARY OF THE RATES

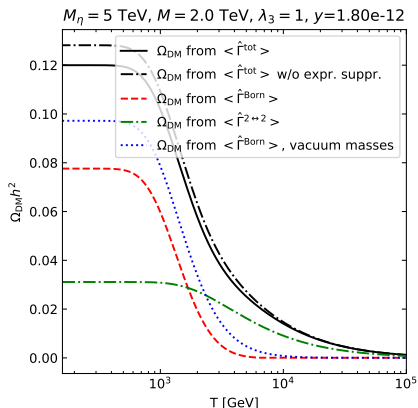
- Phenomenological switch off the high-temperature processes J. Ghiglieri and M. Laine 1605.07720

$$\kappa(M_\eta) = \frac{3}{\pi^2 T^3} \int_0^\infty dp p^2 n_B(E_\eta) [1 + n_B(E_\eta)], \quad \Pi_R^{\text{tot}} = \text{Im}\Pi_R^{1\leftrightarrow 2} + \text{Im}\Pi_R^{2\leftrightarrow 2}$$

- $\text{Im}\Pi_R^{1\leftrightarrow 2} = (\text{Im}\Pi_R^{\text{LPM}} - \text{Im}\Pi_R^{\text{LPM Born}}) \kappa(M_\eta) + \text{Im}\Pi_R^{\text{Born}}$



DM ENERGY DENSITY WITH HIGH-T EFFECTS

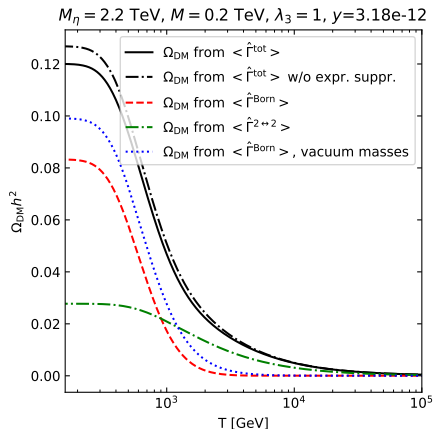
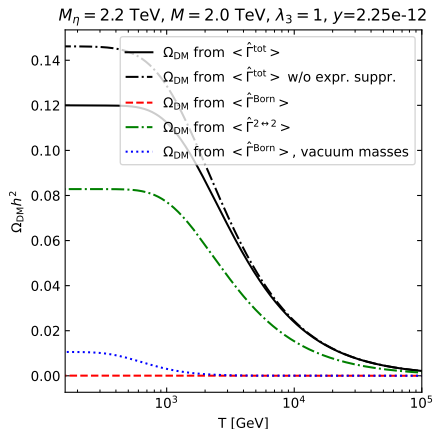


- at $T \gg M_\eta, M$
 \rightarrow **non-negligible production of χ !**
- important high-temperature window
 $2 \rightarrow 2$ and effective $1 \leftrightarrow 2$

- Born rate with vacuum masses \Rightarrow **20% reduction** of $\Omega_{\text{DM}} h^2$ with respect to Π_R^{tot}
- **30%** when including **thermal masses** but excluding $2 \rightarrow 2$ and effective $1 \leftrightarrow 2$
- estimation of theoretical error: LPM with and without $\kappa(M_\eta)$, here $\sim 10\%$ effect

LARGE AND SMALL MASS SPLITTINGS

- the smaller $\Delta M/M$ the larger the effect of thermal masses, LPM and $2 \rightarrow 2$
Left plot: $\Delta M/M = 0.1$; Right plot: $\Delta M/M = 10$
- also other parameters of the model are relevant (h_q, λ_3)



SUMMARY

- we studied the impact of the **ultra-relativistic regime** on the production of a feebly interacting DM particle
- Even in models with renormalizable interactions
high-temperature $1 \leftrightarrow 2$, $2 \rightarrow 2$ can give $\mathcal{O}(1)$ contribution

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high-temperature $1 \leftrightarrow 2, 2 \rightarrow 2$ can give $\mathcal{O}(1)$ contribution
- simplified dark matter model:
 χ Majorana fermion DM and η mediator charged under $SU(3) \otimes U(1)_Y$
- **large impact from $1 \leftrightarrow 2, 2 \rightarrow 2$**

$$M = 2 \text{ TeV}, \Delta M = 0.2 \text{ TeV} \Rightarrow \frac{(\Omega h^2)_{\text{full}}}{(\Omega h^2)_{\text{Born}}} \simeq 10$$

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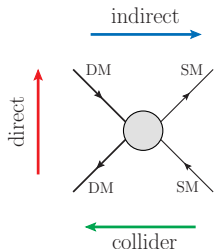
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- similar effects can affect **other models** if DM comes from particles in equilibrium
- **Our main uncertainty comes from the lack of NLO rates**
Extend existing results for one massive state

M.Laine (2013); I. Ghisoiu and M. Laine (2014); J. Ghiglieri and G. D. Moore (2014); G. Jackson (2019)

FREEZE-IN PRODUCTION MECHANISM



- DM as a particle: many candidates see review G. Bertone 2016
- non-interacting with photons, absolutely stable or long-lived $\sim \tau_{\text{Universe}}$
- Any model has to comply with

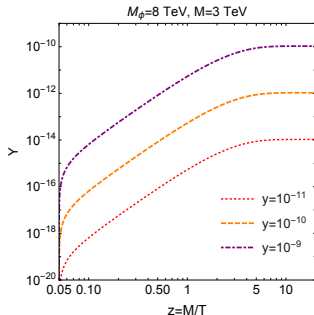
$$\Omega_{\text{DM}} h^2(M_{\text{DM}}, M_{\text{DM}'}, \alpha_{\text{DM}}, \alpha_{\text{SM}}) = 0.1200 \pm 0.0012$$

FREEZE-IN MECHANISM J. McDONALD (2002)

- DM never reach thermal equilibrium
- DM from decay and/or annihilations of **equilibrated** species
- for a simple model $\mathcal{L}_{\text{int}} = -y\phi\bar{\chi}\chi$, $\phi \rightarrow \chi\chi$

$$\frac{dn_{\chi}}{dt} + 3Hn_{\chi} = \langle \Gamma_{\phi \rightarrow \chi\chi} \rangle n_{\phi}^{\text{eq}}, \quad Y = n_{\chi}/s$$

- $\Omega_{\text{DM}} h^2 = \frac{M}{\text{GeV}} \frac{Y_{\text{fin}}}{3.645 \times 10^{-9}}$



THE BORN RATE WITH VANISHING THERMAL MASSES

- Let us look at the model at hand

$$\left(\frac{\partial}{\partial t} - Hk_i \frac{\partial}{\partial k_i} \right) f_\chi(t, \mathbf{k}) = \Gamma(k)[n_F(k^0) - f_\chi(t, \mathbf{k})],$$
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$$\begin{aligned} \Pi^E(K) &\equiv \text{Tr} \left\{ i \not{K} \left[\int_X e^{iK \cdot X} a_R \langle (\eta^\dagger q)(X) (\bar{q} \eta)(0) \rangle a_L \right] \right\} \\ &= N_c \int_{\mathbf{p}} T \sum_n \frac{-i \not{p} a_L}{p_n^2 + E_q^2} \frac{1}{(p_n + k_n)^2 + E_\eta^2} \end{aligned}$$

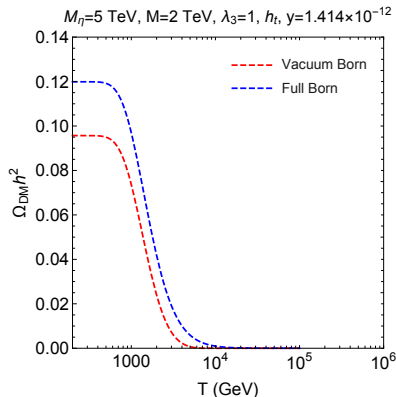
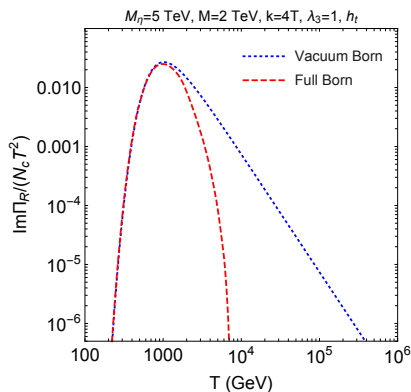
- with $E_q = |\mathbf{p}| = p$ and $E_\eta = \sqrt{(\mathbf{p} + \mathbf{k})^2 + M_\eta^2}$

see M. Laine and A. Vuorinen (2017)

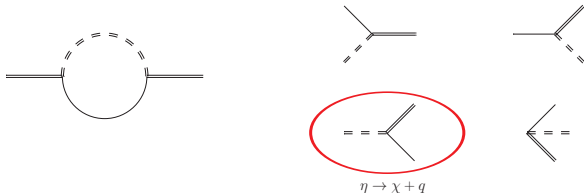
IN-VACUUM VERSUS FINITE THERMAL MASSES

- Scalar mass: $\mathcal{M}_\eta^2 = M_\eta^2 + m_\eta^2$, for $\eta \rightarrow \chi + q$ and $E_p = \sqrt{p^2 + m_q^2}$

$$\text{Im}\Pi_{R,\eta \rightarrow \chi q}^{\text{Born}} = \frac{N_c}{16\pi k} \int_{p_{\min}}^{p_{\max}} dp [\mathcal{M}_\eta^2 - M^2 - m_q^2 - 2k^0(E_p - p)] [n_B(k^0 + E_p) + n_F(E_p)]$$



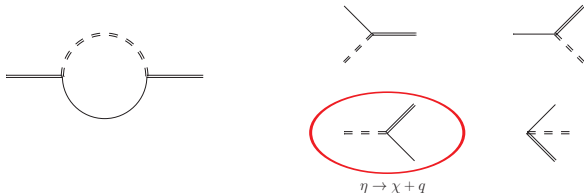
BORN TERM AND BOLTZMANN EQUATION



FOR US $M_\eta > M_\chi + M_q \dots$ (FIRST ROW: $M_\chi > M_\eta + M_q$, $M_q > M_\eta + M_\chi$)

$$\text{Im}\Pi_R^{\text{Born}} = \frac{N_c(M_\eta^2 - M^2)}{8n_F(k^0)} \int \frac{d^3\mathbf{p}_\eta}{(2\pi)^3} \int \frac{d^3\mathbf{p}_q}{(2\pi)^3} \frac{(2\pi)^4 \delta^4(\mathcal{P}_\eta - \mathcal{P}_q - \mathcal{K})}{E_\eta E_q} n_B(E_\eta)(1 - n_F(E_q))$$

BORN TERM AND BOLTZMANN EQUATION



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- $n_{\text{DM}} = 2 \int_{\mathbf{k}} f_\chi(t, \mathbf{k})$, for $\eta \rightarrow \chi + q$

$$\begin{aligned} \dot{n}_{\text{DM}} + 3Hn_{\text{DM}} &= 2|y|^2 \int_{\mathbf{k}} \frac{n_F(k^0)}{k^0} \text{Im}\Pi_R \\ &= 2|y|^2 N_c(M_\eta^2 - M^2) \int_{\mathbf{p}_\eta, \mathbf{p}_q, \mathbf{k}} \frac{(2\pi)^4 \delta^4(\mathcal{P}_\eta - \mathcal{P}_q - \mathcal{K})}{8E_\eta E_q k^0} n_B(E_\eta)[1 - n_F(E_q)] \end{aligned}$$

FIRST IMPROVEMENT: THERMAL MASSES

- at high temperatures, $\pi T \gg M_\eta$, **repeated interactions within the bath change the dispersion relations** \Rightarrow *asymptotic masses*

see also L. Darmé, A. Hryczuk, D. Karamitros and L. Roszkowski (2019)

- for $T > T_c \simeq 150$ GeV the quarks only have

$$m_q^2 = \frac{T^2}{4} (g_3^2 C_F + Y_q^2 g_1^2 + |h_q|^2)$$

- for the colored scalar

$$m_\eta^2 = \left(\frac{g_3^2 C_F + Y_q^2 g_1^2}{4} + \frac{\lambda_3}{6} \right) T^2$$

- no thermal mass correction for χ since $y \ll g$

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$$m_B^2 = \left(\frac{n_S}{6} + \frac{5n_G}{9} + \frac{Y_q^2 N_c}{3} \right) g_1^2 T^2, \quad m_g^2 = \left(\frac{N_c}{3} + \frac{n_G}{3} + \frac{1}{6} \right) g_3^2 T^2$$

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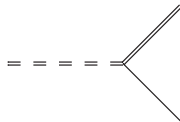
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$$\begin{aligned} m_\eta^2(M_\eta, p) &= \frac{Y_q^2 g_1^2 + C_F g_3^2}{2\pi^2} \int_0^\infty dq \left\{ \frac{n_B(E_\eta(q))}{E_\eta(q)} \left[q^2 + \frac{M_\eta^2 q}{2p} \ln \left(\frac{(p+q)^2}{(p-q)^2} \right) \right] + 2q n_B(q) \right\} \\ &+ \frac{\lambda_3}{\pi^2} \int_0^\infty dq q n_B(q) \end{aligned}$$

LPM RESUMMATION FOR LIGHT-CONE KINEMATICS

$$\pi T \gg M_\eta$$

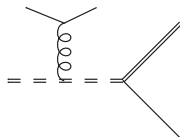
- all the particles are seen as *massless*,
- all the momenta of external particles are $p \sim \pi T$,
- particles are ultra-relativistic, and gT is a soft scale
 \Rightarrow **collinear kinematics \approx high T**



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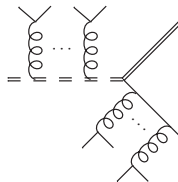
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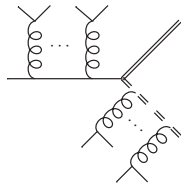
Long formation time $\sim \frac{1}{g^2 T}$

see e.g. J. Ghiglieri and G. D. Moore (2014)

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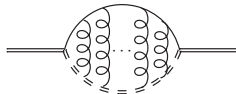
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- connection between LPM resummation and $\text{Im}\Pi_R^{\text{LPM}} \Rightarrow \chi$ self-energy
- notation and computational setting from

$$\hat{H} \equiv -\frac{M^2}{2k_0} + \frac{m_q^2 - \nabla_\perp^2}{2E_q} + \frac{\mathcal{M}_\eta^2 - \nabla_\perp^2}{2E_\eta} + i\Gamma(y), \quad y \equiv |\mathbf{y}_\perp|$$

$$\Gamma(y) = \frac{T}{2\pi} g_1^2 Y_q^2 \left[\ln\left(\frac{m_{By}}{2}\right) + \gamma_E + K_0(m_{By}) \right] + \frac{T}{2\pi} g_3^2 C_F \left[\ln\left(\frac{m_{gY}}{2}\right) + \gamma_E + K_0(m_{gY}) \right]$$

- \hat{H} enters the inhomogeneous equations for the functions $g(\mathbf{y})$ and $\mathbf{f}(\mathbf{y})$

$$(\hat{H} + i0^+)g(\mathbf{y}) = \delta^{(2)}(\mathbf{y}), \quad (\hat{H} + i0^+)\mathbf{f}(\mathbf{y}) = -\nabla_{\perp}\delta^{(2)}(\mathbf{y})$$

$$\begin{aligned} \text{Im}\Pi_{\text{R}}^{\text{LPM}} &= -\frac{N_c}{8\pi} \int_{-\infty}^{+\infty} dE_q \int_{-\infty}^{+\infty} dE_{\eta} \delta(k_0 - E_q - E_{\eta}) [1 - n_{\text{F}}(E_q) + n_{\text{B}}(E_{\eta})] \\ &\quad \frac{k_0}{E_{\eta}} \lim_{\mathbf{y} \rightarrow 0} \left\{ \frac{M^2}{k_0^2} \text{Im}[g(\mathbf{y})] + \frac{1}{E_q^2} \text{Im}[\nabla_{\perp} \cdot \mathbf{f}(\mathbf{y})] \right\} \end{aligned}$$

NUMERICAL STRATEGY AND LPM BORN

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ONCE $E_{\eta} = k^0 - E_q$ BY THE δ

- 1 $k^0 > E_q > 0$: this corresponds to the effective $2 \rightarrow 1$ process $\eta, q \rightarrow \chi$
- 2 $E_q < 0$: this corresponds to the effective $1 \rightarrow 2$ process $\eta \rightarrow q\chi$
 \Rightarrow LPM-Born with thermal masses is the $n = 0$ limit (no scatterings)
- 3 $E_q > k^0$: this corresponds to the effective $1 \rightarrow 2$ process $q \rightarrow \eta\chi$.

$$(\Omega_{\text{DM}} h^2)_{\text{obs.}} = (\Omega_{\text{DM}} h^2)_{\text{freeze-in}} + (\Omega_{\text{DM}} h^2)_{\text{super-WIMP}}$$

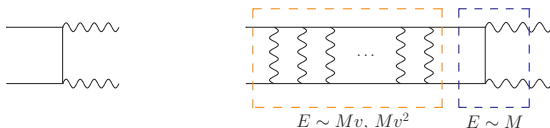
- η particles stays in chemical equilibrium till late times $T \sim M_\eta/25$
- there is a populations of η , as governed by freeze-out, which decays into χ

M. Garny, J. Heisig (2018)

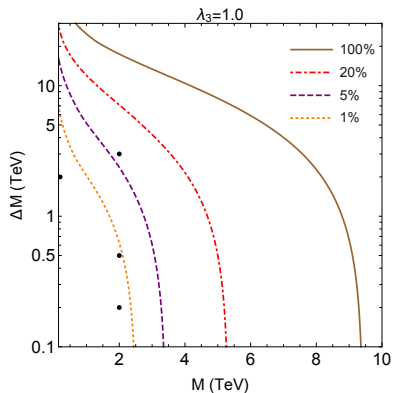
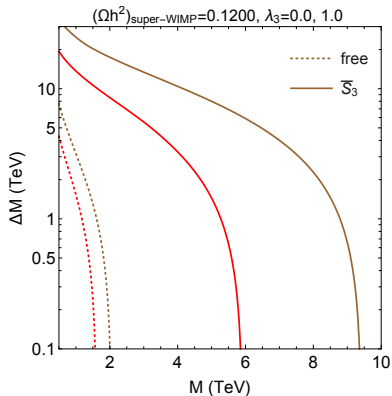
$$(\Omega_{\text{DM}} h^2)_{\text{super-WIMP}} = \frac{M}{M_\eta} (\Omega h^2)_\eta .$$

$$\frac{dn_\eta}{dt} + 3Hn_\eta = -\langle \sigma_{\text{eff}} v \rangle (n_\eta^2 - n_{\eta,\text{eq}}^2), \quad \langle \sigma_{\text{eff}} v \rangle = \frac{c_3 \bar{S}_3 + c_4 C_F \bar{S}_4}{N_c}$$

A. Mitridate et al (2017), S. B. and M. Laine (2018), S. B. and S. Vogl (2019), J. Hartz and K. Petraki (2018)

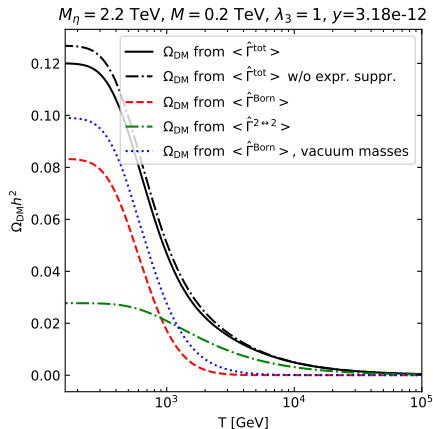
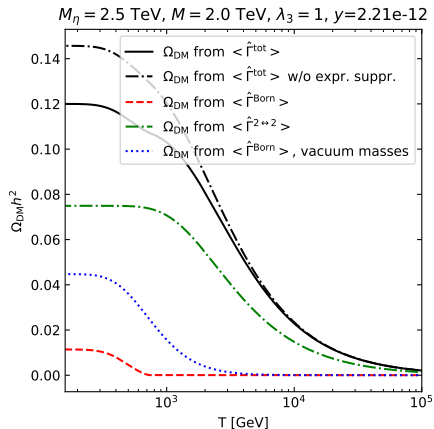


BENCHMARK VALUES WITH BOUND-STATE EFFECTS



- P0 ($M = 2.0$ TeV, $M_\eta = 5.0$ TeV); P1 ($M = 0.2$ TeV, $M_\eta = 2.2$ TeV);
P2 ($M = 2.0$ TeV, $M_\eta = 2.5$ TeV); P3 ($M = 2.0$ TeV, $M_\eta = 2.2$ TeV);

LARGE AND SMALL MASS SPLITTINGS II



- for $M_\eta = 2.2 \text{ TeV}$ the freeze-in production stops fairly close to $T_c \simeq 150 \text{ GeV}$
 \Rightarrow follow DM production in the SM broken phase
- CMS analysis provides us with $M_\eta > 1250 \text{ GeV}$ CMS-PAS-EXO-16-036