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# Deep learning jet modifications in heavy-ion collisions

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with Daniel Pablos and Konrad Tywoniuk

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#### **Outline**

#### 1 Motivation

2 Deep learning jet energy loss

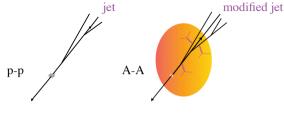
#### 3 Application

- Sensitivity of jet observables to in-medium modification
- Jet tomography

#### 4 Conclusion and outlook

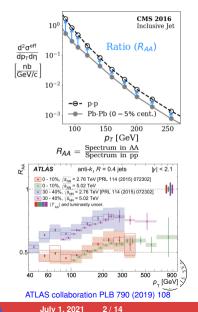


## Jets in the medium

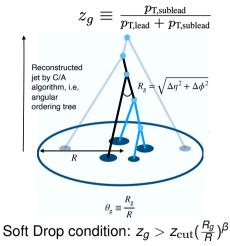


J. Brewer, HP'20

- Quark-gluon plasma (QGP) created in heavy ion collision: deconfined phase, hot dense medium
- Jets, collimated sprays of energetic particles, serving as hard probe to medium properties
- Jets are quenched and modified in the medium via parton energy loss



### Jet substructures



J. Mulligan, HP'20

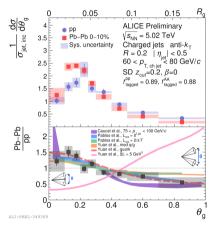
 Detector as camera: positions, energies of particles

All jet constituents are reclustered in angular ordering.

SoftDrop: find the first *hard* splitting between two subjets satisfying  $z_g > z_{cut}\theta^{\beta}$  with momentum sharing  $z_g$  and angle of branching  $R_g$ .



## Jet modifications: ambiguous interpretations



Ratio of jet observables distr. between medium and vacuum, BOTH with  $p_T^{jet} > p_T^{cut}$  ■ Interplay: jet substructures, e.g., *R<sub>g</sub>*, could

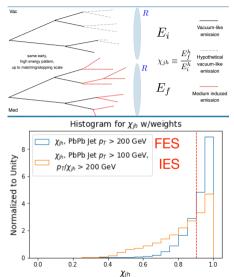
- be modified during the passage through the medium and/or
- affect the amount of jet energy loss and then this jet doesn't pass the  $p_T$  cut in the selection, i.e., selection bias.
- Jets produce emissions with smaller R<sub>g</sub> in medium than in vacuum: presumes medium scale dominates
- Jets with larger R<sub>g</sub> in vacuum are more suppressed in medium: presumes vacuum scale dominates

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Can we disentangle these two effects with knowledge of the degree of quenching for eac individual measured jets?

Yi-Lun Du

## Energy loss ratio & Jet selections



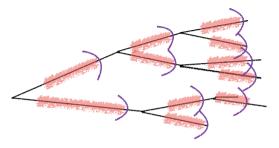
Study jet observables for jets that belong to 2 different quenching classes:

- Unquenched class:  $\chi_{ih} > 0.9$ .
- Quenched class:  $\chi_{ih} < 0.9$ .
- **p** piets:  $p_T > 200 \text{ GeV}$
- PbPb jets:
  - Final Energy Selection (FES): impose  $p_T$  cut on final energy  $p_T > 200 \text{ GeV} \rightarrow \text{Steeply falling energy}$ loss dist. Biased by little guenched samples!
  - Initial Energy Selection (IES): impose  $p_T$  cut on *initial energy* via  $\chi_{ib}$ ,  $p_T/\chi_{ib} > 200 \text{ GeV} \& p_T > 100$  $GeV \rightarrow More$  support of fairly guenched jets in the quenched class. More distinguishable!



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#### Hybrid model



- PYTHIA8 down to hadronization scale
- Strongly coupled energy loss at every stage
- Hadrons from the hydro. wake (medium response)

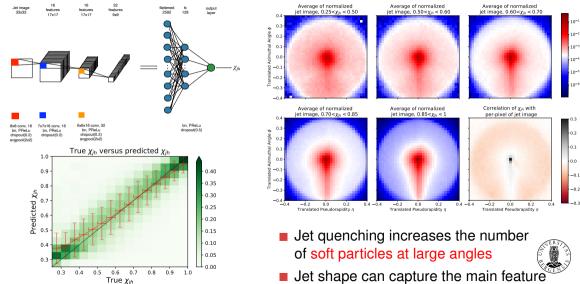
Casalderrey-Solana, Gulhan, Milhano, Daniel Pablos, Rajagopal JHEP '15,'16,'17

- Vacuum jets using  $\hat{p}_{T,\min} = 50$ GeV, with oversampling power  $p_T^4$ .
- PbPb collisions in 0-5% centrality at  $\sqrt{s} = 5.02$  ATeV.
- Reconstructed jets with anti- $k_T$ , R = 0.4, required to be  $|\eta| < 2$  and  $p_T^{\text{jet}} > 100 \,\text{GeV}$ .
- ~ 250,000 jets. 80% for training and 20% for validation.

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## **CNN Prediction & Interpretability**



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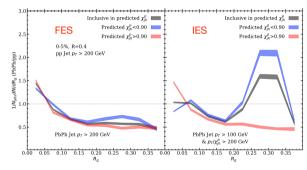
# Jet radius, R<sub>g</sub>

 $R_g$  ratio between PbPb and pp jets

■ FES: Selection bias towards jets with smaller *R<sub>g</sub>*, originated by *p<sub>T</sub>* cut.

#### IES:

- Unquenched class: still biased due to  $\chi_{jh}$  cut: to belong to this class, a jet had better to be with smaller  $R_g$ , compared with all pp jets.
- Quenched class presents features related to energy loss, compared with unquenched class: jet quenching leads to enhancement of large R<sub>g</sub>.





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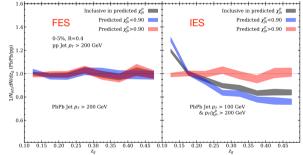
#### Groomed momentum sharing fraction, $z_g$

 $z_g$  ratio between PbPb and pp jets

 FES: No selection bias observed. Scale of emission isn't strongly dependent on splitting fraction z<sub>g</sub>.

IES:

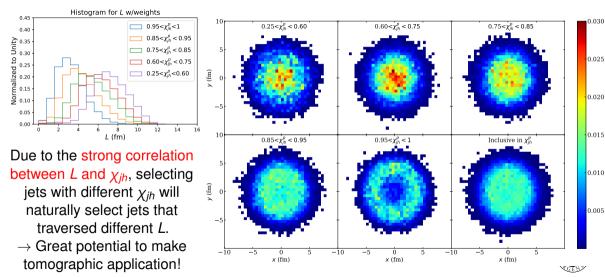
 Quenched class presents features related to energy loss, compared with unquenched class: jet quenching leads to enhancement of smaller z<sub>g</sub> subjets.



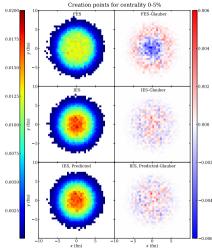
Y.-L. Du, D. Pablos, K. Tywoniuk, JHEP03(2021)206

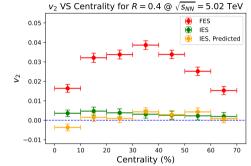


# Applications: Jet tomography, length VS $\chi_{jh}$



### **Applications: creation points & orientation**





- IES "removes" final state interactions (selection bias), since we record "all" jets.
- IES provides access to the genuine jet creation point distribution and jet orientation.

Y.-L. Du, D. Pablos, K. Tywoniuk, arXiv: 2106.11271

 $V_2 = \left\langle \frac{p_x^2 - p_y^2}{p_x^2 + p_y^2} \right\rangle$ 



#### **Conclusion and outlook**

- CNN can extract energy loss jet-by-jet from jet image with good performance
- Procedure generalisable to many jet quenching models
- Jet shape contains significant predictive power: angular distribution of soft particles
- Mitigate selection bias and reveal medium effects on various jet observables
- Open opportunity to make tomographic study

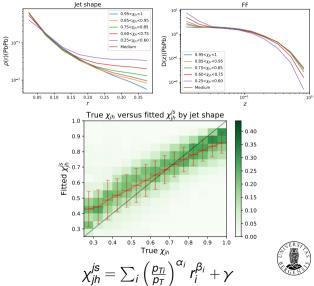
- Generalizability to other MC quenching models?
- Applicability to more realistic environment: fluctuating background?
- Better performance from other state-of-the-art neural networks?
- Extract traversed length with better precision?
- Unfold jet *initial* properties apart from jet energy?



## **Backup: Prediction performance with FCNN**

Input (size)	Output	Network	Loss
FF (10)	Xjh	FCNN	0.0058
Jet shape (8)	Xjh	FCNN	0.0033
FF, jet shape (18)	Xjh	FCNN	0.0032
FF, jet shape, features (25)	Xjh	FCNN	0.0028
Jet image & FF, jet shape, features (25)	Xjh	API: CNN&FCNN	0.0028

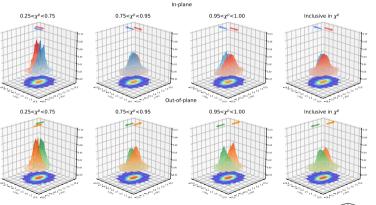
- Jet shape outperforms jet FF.
- Motivates construction from jet shape by 17-parameter fitting:
  - Still a bit worse than CNN
- Jet observables recover the performance by jet image with equivalent predictive power: interpretability!



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# Backup: Jet tomography with $\chi_{jh}$ & $v_2$

- $v_2 = \frac{\rho_x^2 \rho_y^2}{\rho_x^2 + \rho_y^2}$
- Top row: In-plane jets  $(v_2 > 0)$  going left  $(p_x < 0)$ and right  $(p_x > 0)$
- Bottom row: Out-of-plane jets ( $v_2 < 0$ ) going up ( $p_y > 0$ ) and down ( $p_y < 0$ )
- To get very quenched, jets have to travel longer in medium. So v<sub>2</sub> & p<sub>x,y</sub> are helpful for jet tomography.



Creation points density for centrality 30-40%,  $R = 0.4 \otimes \sqrt{s_{HN}} = 5.02$  TeV, FES,  $p_T > 100$  GeV

