#### LAWRENCE BERKELEY NATIONAL LABORATORY

# DIRECT DARK MATTER SEARCHES WITH THE LZ EXPERIMENT

#### **UPDATE AND STATUS**

WEDNESDAY, JUNE 2, 2020 - GDR DUPHY

### QUENTIN RIFFARD







## LIQUID XENON TIME PROJECTION CHAMBER



- High ionization yield ( $W \sim 13.7 \,\mathrm{eV}$ )
- High scintillation yield (> 50,000 photons/MeV)
- **Transparent** to its own scintillation light
- **Self-shielded**
- High intrinsic radio-purity
- **Axion Like Particles**
- Exotic Dark Matter candidates (mirror dark matter, solar axions, ...)
- Neutrino physics & astrophysics

#### Liquid Xenon is a good target for **WIMP Dark Matter searches...**

#### High density

### ... but not only

-  $0\nu\beta\beta$  search using <sup>134</sup>Xe & <sup>136</sup>Xe



## LIQUID XENON TIME PROJECTION CHAMBER



#### Liquid Xenon TPC

- Type of signals
  - S1: Primary scintillation (light)
  - S2: Secondary scintillation (charge)
- 3D position reconstruction
  - XY position:  $\mathcal{O}(mm)$  from S2 light pattern
  - Z position:  $\mathcal{O}(\mu m)$  from S1-S2 timing

#### LZ TPC

- 1.5 m in diameter and height
- 10 tonnes of total mass, 7 tonnes active and 5.6 tonnes fiducial volume
- Electric field: 300 V/cm
- 247 (top) + 247 (bottom) PMTs to monitor the TPC



## LZ TIME PROJECTION CHAMBER IN PICTURES













### LZ DETECTOR SYSTEMS



- 4850 feet (1.48 km) underground
- Muon flux reduced by  $\mathcal{O}(10^7)$



### LZ DETECTOR SYSTEMS









#### Water tank in the Davis cavern at SURF



![](_page_6_Picture_7.jpeg)

### LZ VETO SYSTEMS

![](_page_7_Figure_1.jpeg)

### LZ has two veto systems

- Two tonnes of LXe surrounding the TPC
  - PMTs at top and bottom of the skin region
  - Lined with PTFE to maximize light collection efficiency
- Anti-coincidence detector for γ-rays

#### **Outer Detector**

- 17 tonnes Gd-loaded liquid scintillator in acrylic vessels
  - 120 8" PMTs mounted in the water tank
  - Anti-coincidence detector for γ-rays and neutrons
  - Observe  $\sim 8 \,\text{MeV} \,\gamma$ -rays from thermal neutron capture

![](_page_7_Picture_12.jpeg)

- **Geant4-based simulation** for background studies (Astro.Phys. 125 102480 (2020))

![](_page_8_Figure_4.jpeg)

![](_page_8_Picture_6.jpeg)

### **BACKGROUND REDUCTION**

### **Signal/Background discrimination**

ER and NR events can be distinguished from their different S2/S1 ratio

![](_page_9_Figure_3.jpeg)

1000 days run

- Before discrimination 1131 ER events and 10.4 NR events
- After discrimination 5.97 ER events and 0.51 NR events (with 99.5% ER discrimination, 50% NR efficiency)

![](_page_9_Picture_7.jpeg)

![](_page_9_Figure_8.jpeg)

![](_page_9_Picture_11.jpeg)

### **BACKGROUND REDUCTION**

### Fiducialisation and Veto cut (NR background)

**Before veto cuts** 

![](_page_10_Figure_3.jpeg)

**12.31 cts / 1000 days** 

![](_page_10_Picture_5.jpeg)

After veto cuts

![](_page_10_Figure_7.jpeg)

1.03 cts / 1000 days

![](_page_10_Picture_9.jpeg)

## WIMP DARK MATTER PROJECTED SENSITIVITY

![](_page_11_Figure_1.jpeg)

LZ sensitivity Paper (Phys. Rev. D 101, 052002)

![](_page_11_Picture_3.jpeg)

# $0\nu\beta\beta$ Decay of $^{136}$ Xe

![](_page_12_Figure_1.jpeg)

- **No enrichments** (8.8%)
- <sup>136</sup>Xe  $Q_{\beta\beta} = 2458 \,\text{keV}$
- Fiducial volume: 1 tonne
- $T_{1/2}(90 C.L.) > 1.06 \times 10^{26}$  years in 1000 live days

Projected sensitivity to the  $0\nu\beta\beta$  decay of <sup>136</sup>Xe (Phys. Rev. C 102, 014602 (2020))

![](_page_12_Figure_7.jpeg)

![](_page_12_Picture_8.jpeg)

### SUMMARY

### Good progress in assembly and integration of detector and associated systems

- **Expecting first data later this year**
- **Dark Matter detection** 
  - WIMP projected sensitivity:  $1.4 \times 10^{-48} \text{ cm}^2$  at  $40 \text{ GeV/c}^2$
  - Low mass WIMPs sensitivity (arXiV:2101.08753)
  - Sensitivity to new physics via low-energy electron recoils (arXiv:2102.11740)
- **Neutrino physics** 
  - $0\nu\beta\beta$  decay of <sup>136</sup>Xe (Phys. Rev. C 102, 014602 (2020))
  - $2\nu\beta\beta$  and  $0\nu\beta\beta$  decays of <sup>134</sup>Xe (arXiv:2104.13374)

### Stay tuned!

![](_page_13_Picture_16.jpeg)

# THE LZ (LUX-ZEPLIN) COLLABORATION

34 Institutions: 250 scientists, engineers, and technical staff

- Black Hills State University
- Brandeis University
- Brookhaven National Laboratory
- Brown University
- Center for Underground Physics
- Edinburgh University
- Fermi National Accelerator Lab.
- Imperial College London
- Lawrence Berkeley National Lab.
- Lawrence Livermore National Lab.
- LIP Coimbra
- Northwestern University
- Pennsylvania State University
- Royal Holloway University of London
- SLAC National Accelerator Lab.
- South Dakota School of Mines & Tech
- South Dakota Science & Technology Authority
- STFC Rutherford Appleton Lab.
- Texas A&M University
- University of Albany, SUNY
- University of Alabama
- University of Bristol
- University College London
- University of California Berkeley
- University of California Davis
- University of California Santa Barbara
- University of Liverpool
- University of Maryland
- University of Massachusetts, Amherst
- University of Michigan
- University of Oxford
- University of Rochester
- University of Sheffield
- University of Wisconsin, Madison

![](_page_14_Picture_37.jpeg)

Thanks to our sponsors and participating institutions!

![](_page_14_Picture_39.jpeg)

![](_page_14_Picture_40.jpeg)

![](_page_14_Picture_41.jpeg)

U.S. Department of Energy Office of Science

![](_page_14_Picture_43.jpeg)

![](_page_14_Picture_44.jpeg)

Science and Technology Facilities Council

![](_page_14_Picture_46.jpeg)

![](_page_14_Picture_47.jpeg)

![](_page_14_Picture_48.jpeg)

@lzdarkmatter

### **THANK YOU**

![](_page_15_Picture_1.jpeg)

![](_page_15_Picture_3.jpeg)

U.S. Department of Energy Office of Science

![](_page_15_Picture_5.jpeg)

### Thanks to our sponsors and 34 participating institutions!

![](_page_15_Picture_7.jpeg)

![](_page_15_Picture_8.jpeg)

![](_page_15_Picture_9.jpeg)

![](_page_15_Picture_10.jpeg)

![](_page_15_Picture_11.jpeg)