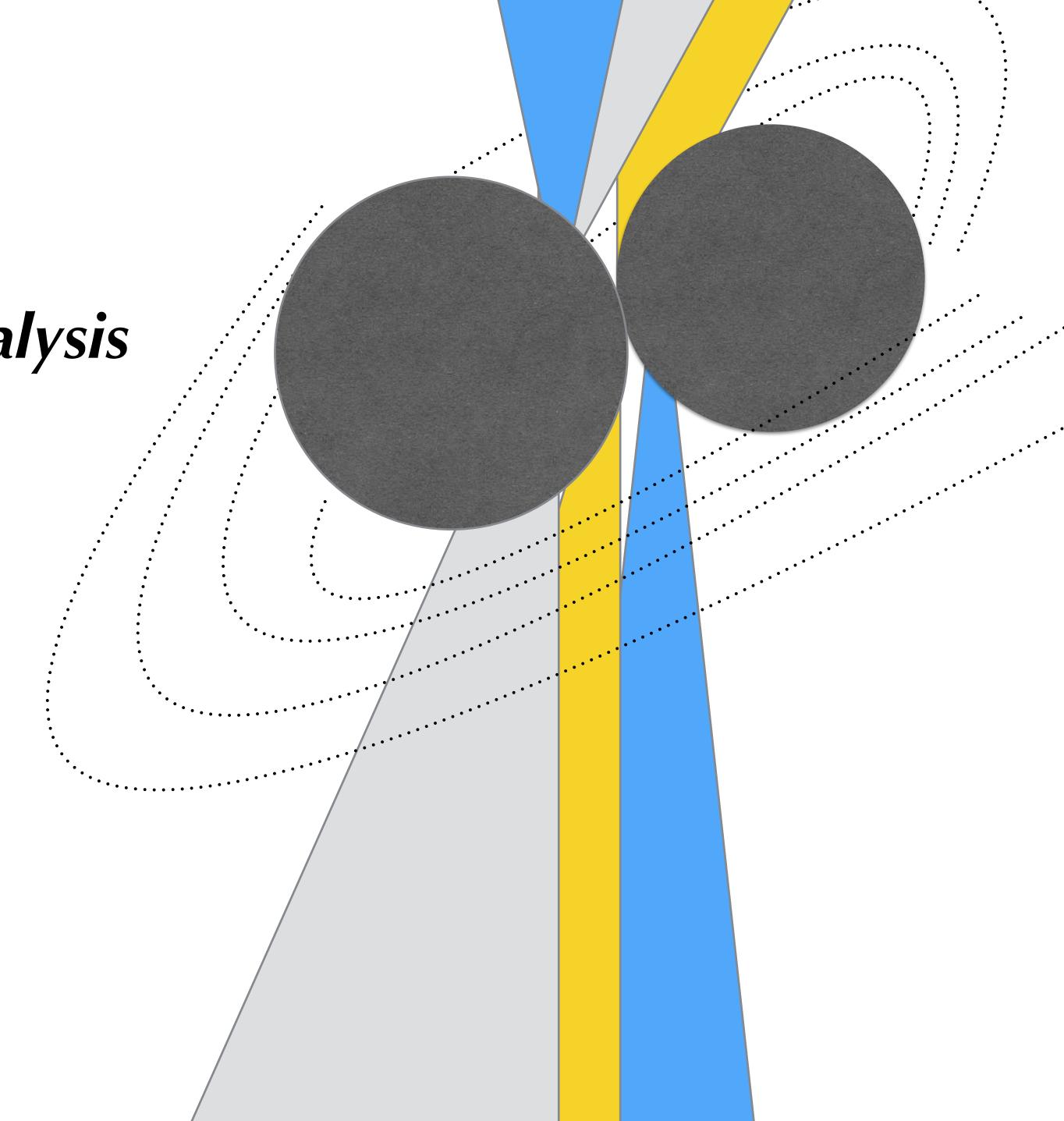
# Deep learning for LISA data analysis

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# **INTERFEROMETER**

## SPACE



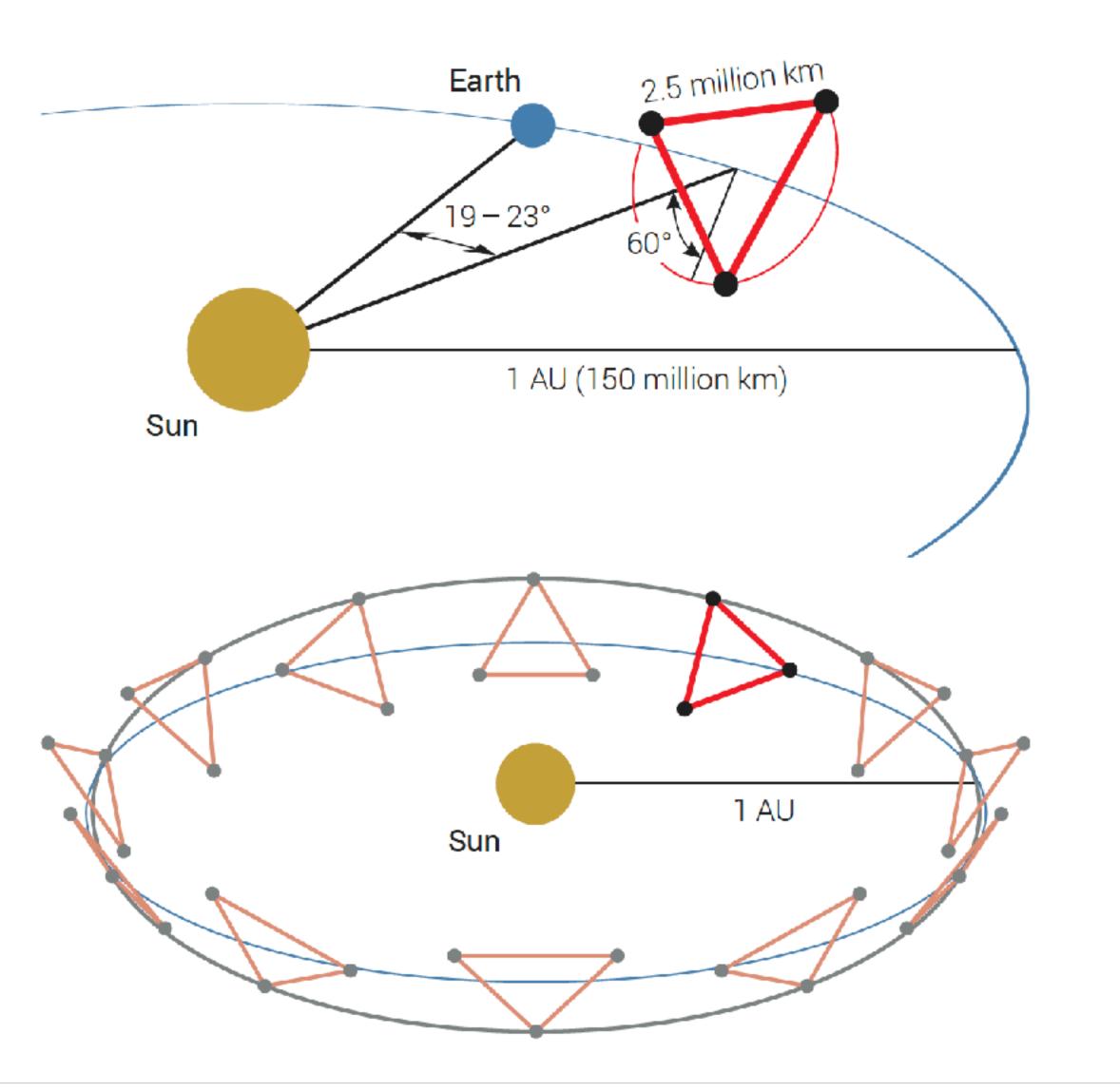


Image: LISA White paper



# Sources

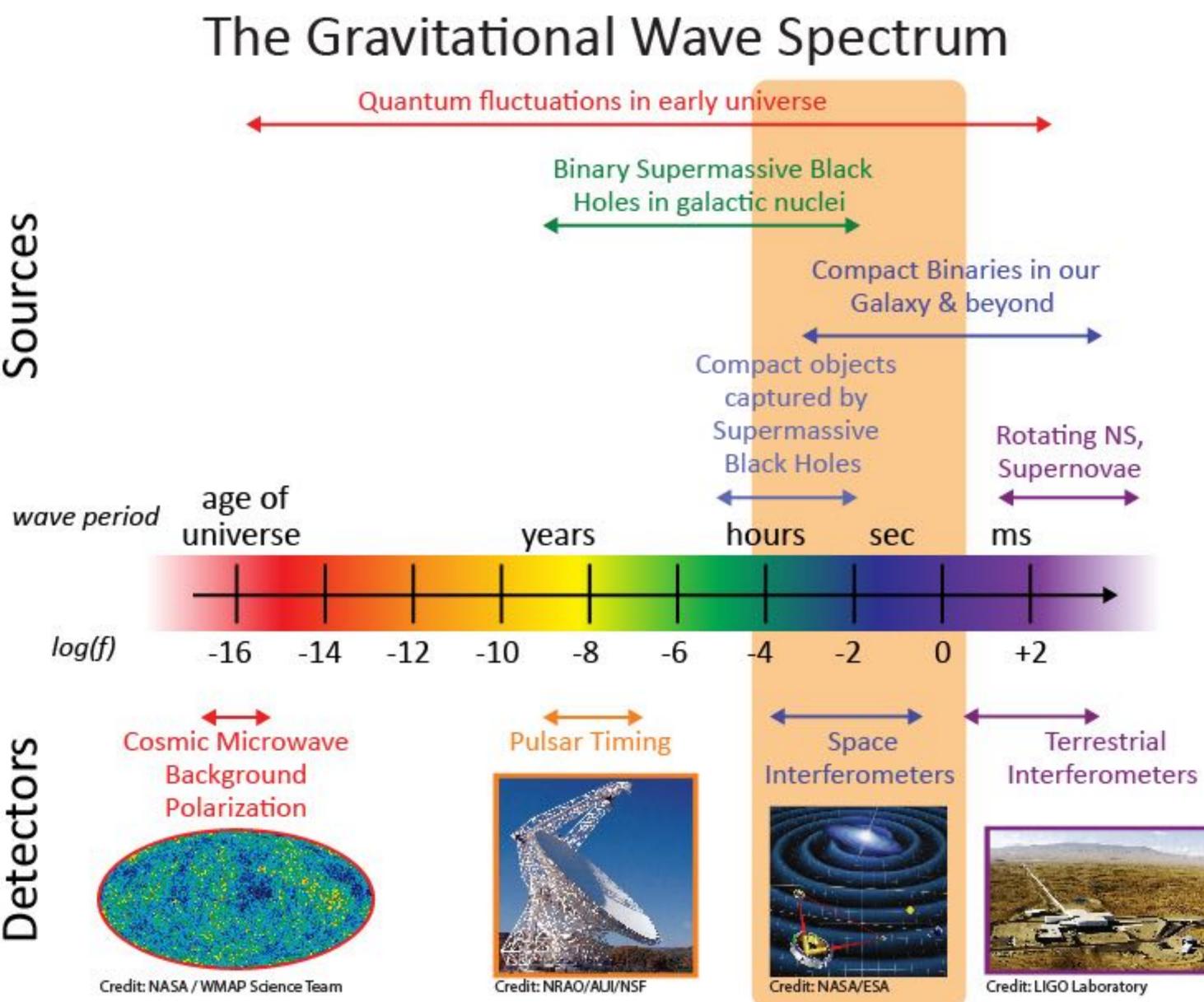


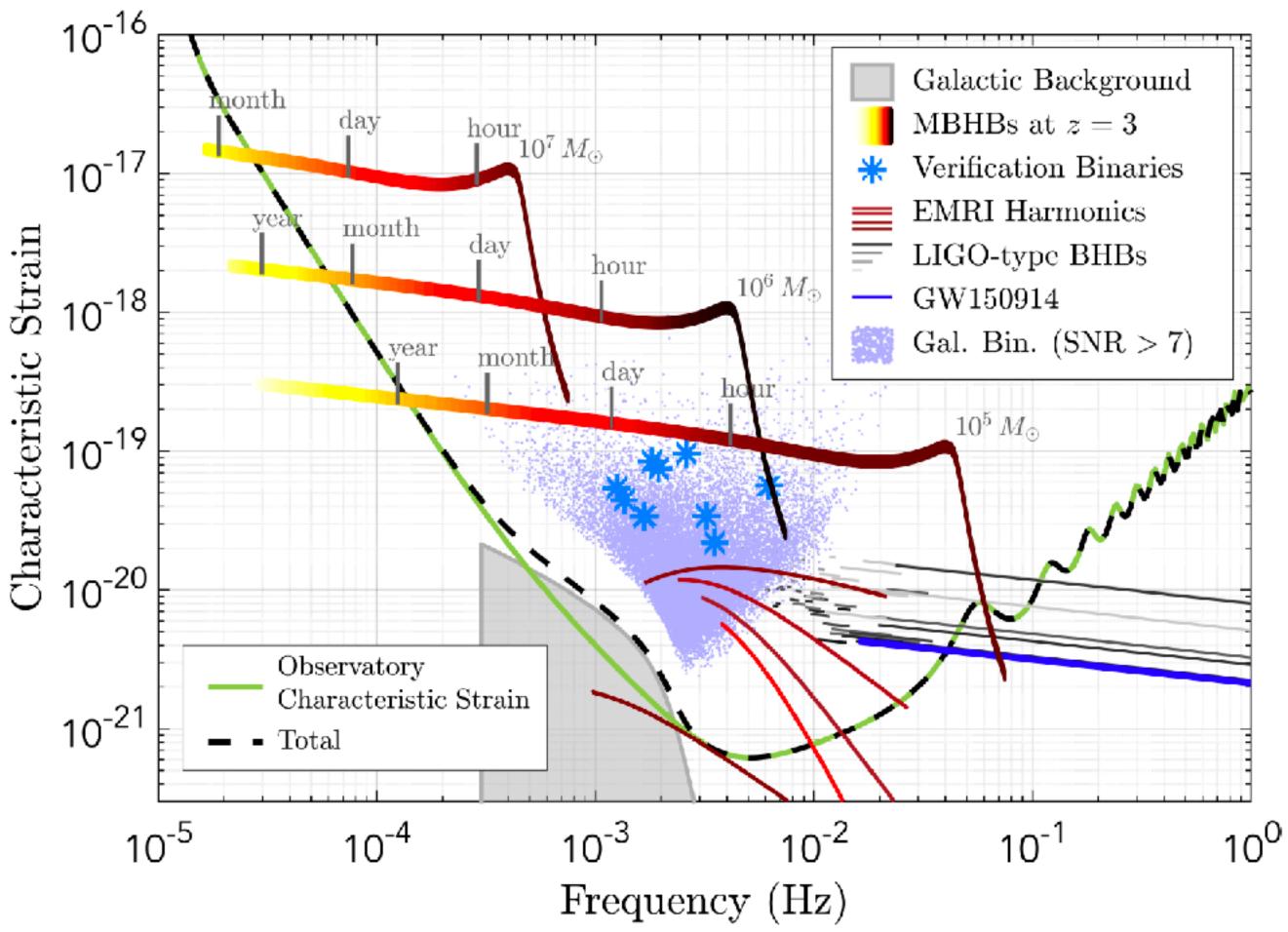
Image: NASA

#### LISA noise and sources

Signals observed:

- Massive Black Hole Binary -----
- Galactic Compact Binaries \_
- Extreme Mass Ratio Inspirals -
- Stellar Origin Black Hole Binaries
- Cosmological Background \_

— . . .

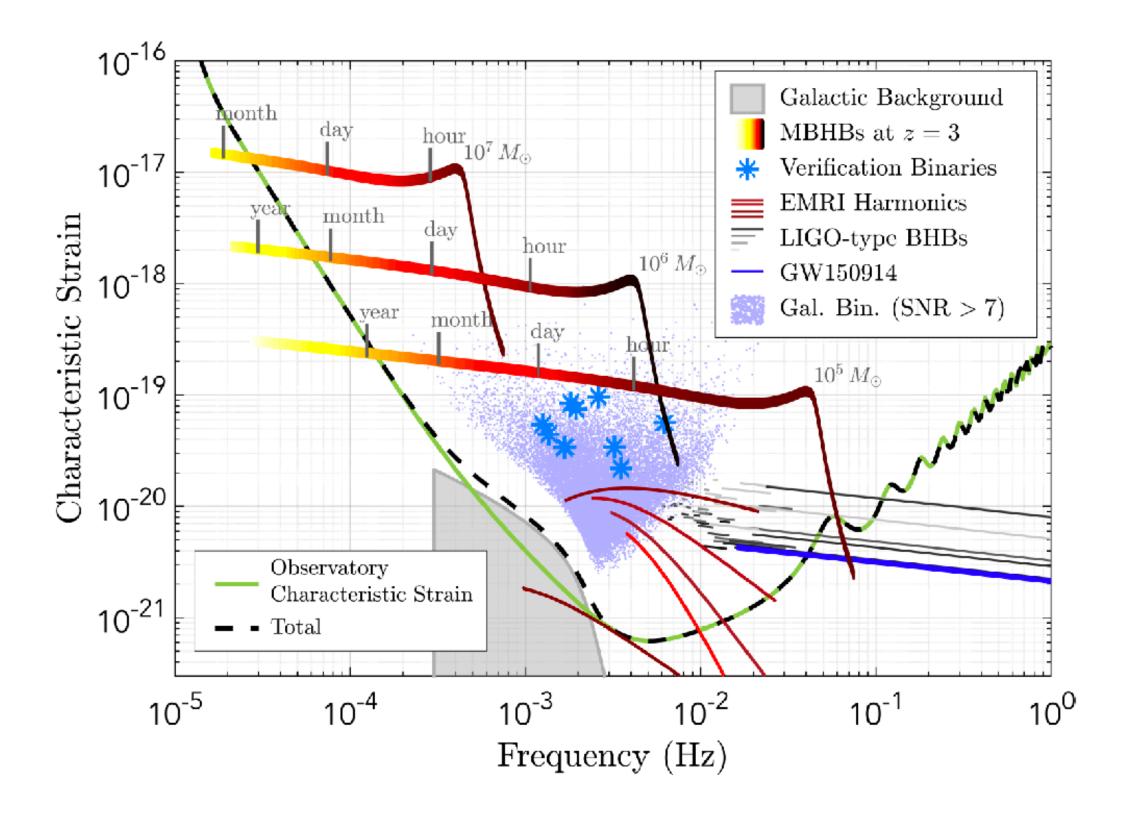


#### Massive Black Hole Binaries

Signals from MBHB mergers observed by LISA depend on

- assumptions regarding MBH formation,
- the recipes employed for the black hole mass growth via merger and gas accretion
- 10 to 100 sources / year





#### Possible electromagnetic counterparts

Multiple authors suggest that the electromagnetic counterparts will be observed as a transient during merger or also during inspiral and merger.

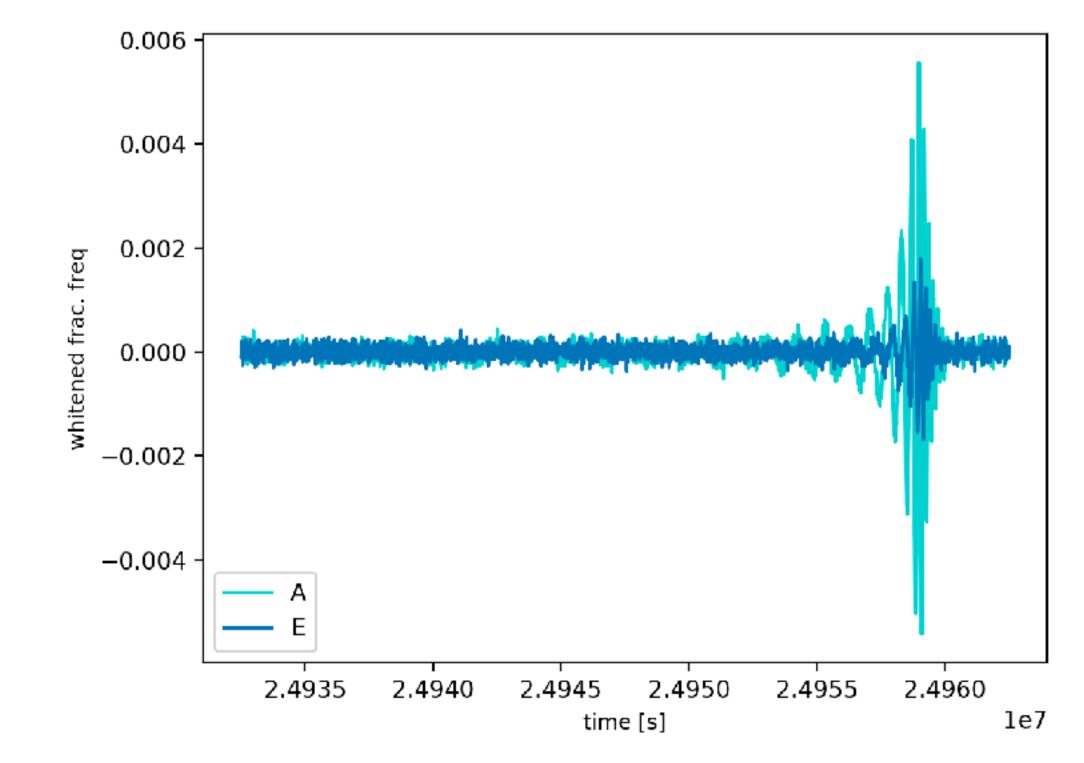
Electromagnetic counterparts will occur due to presence of

- matter or
- magnetic fields.

For example:

- Accretion during merger
- Jets produced by the external magnetic fields

. . .



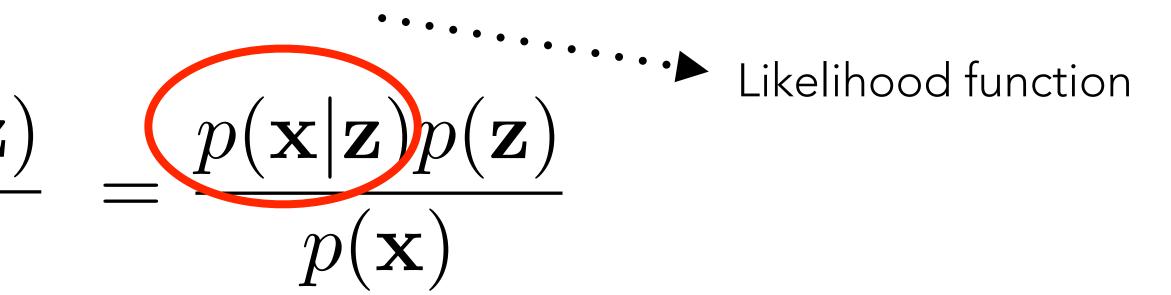
#### Inference

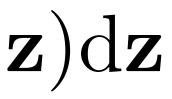
By inference we mean computing the **posterior distribution** of the parameter given the observed data:

$$p(\mathbf{z}|\mathbf{x}) = \frac{p(\mathbf{x}, \mathbf{z})}{p(\mathbf{x})}$$

The problem is that we have to compute marginal likelihood for the observation:

$$p(\mathbf{x}) = \int p(\mathbf{x},$$







#### **Parameter estimation**

It is not possible to perform exact inference for the general problem. We have to introduce some simplifications.

We can use approximate inference:

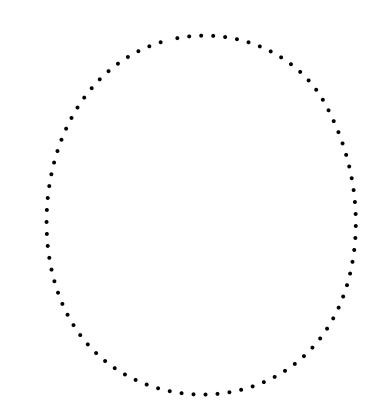
- Markov Chain Monte Carlo/Nested Sampling: sample from the exact posterior

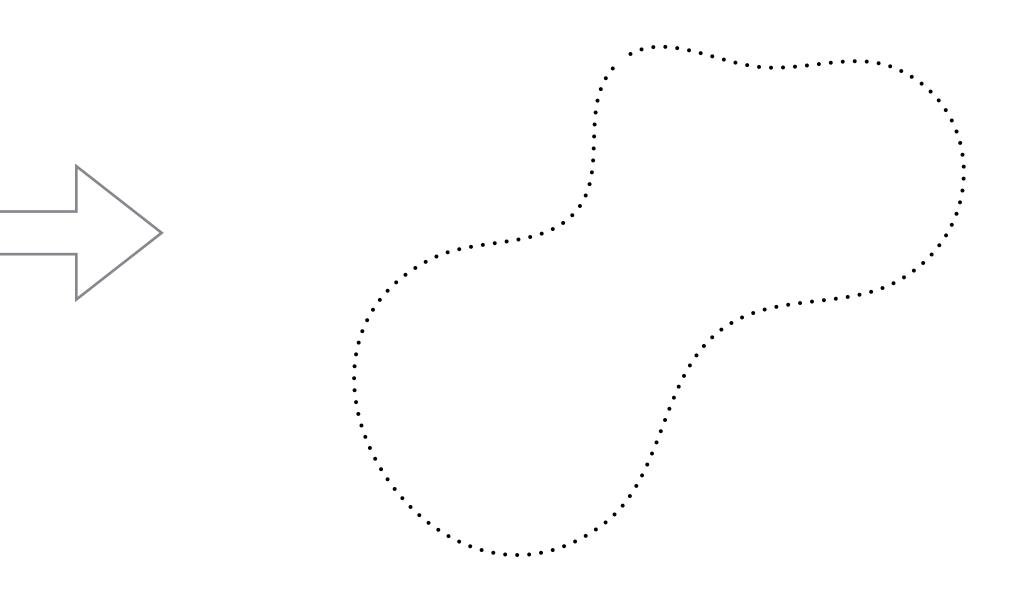
There are some exceptions for the models with some simplifications:

- Gaussian mixture models
- Invertible models

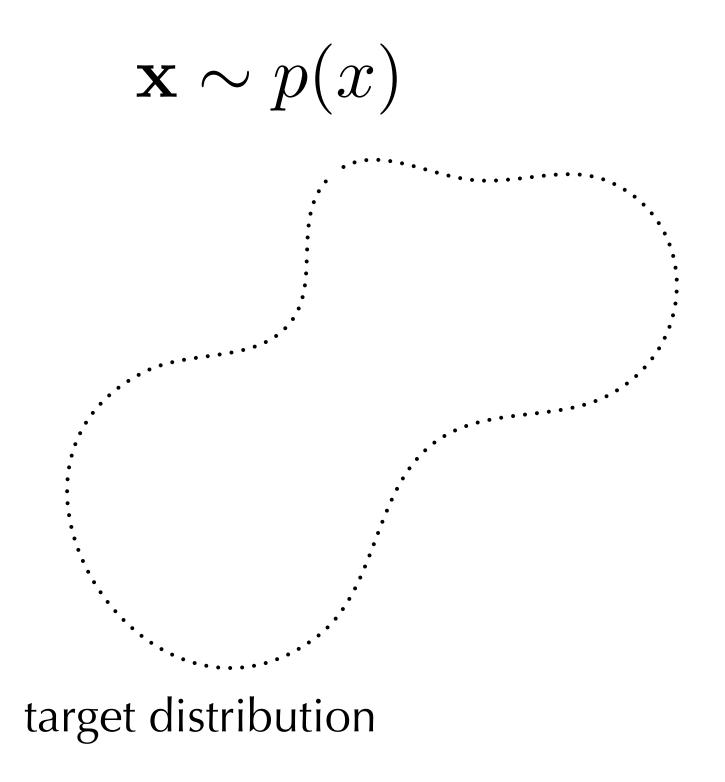
- Variational Inference: approximate the posterior distribution with a tractable distribution

We want to make a deterministic map from the simple and easy to sample distribution to a complex one



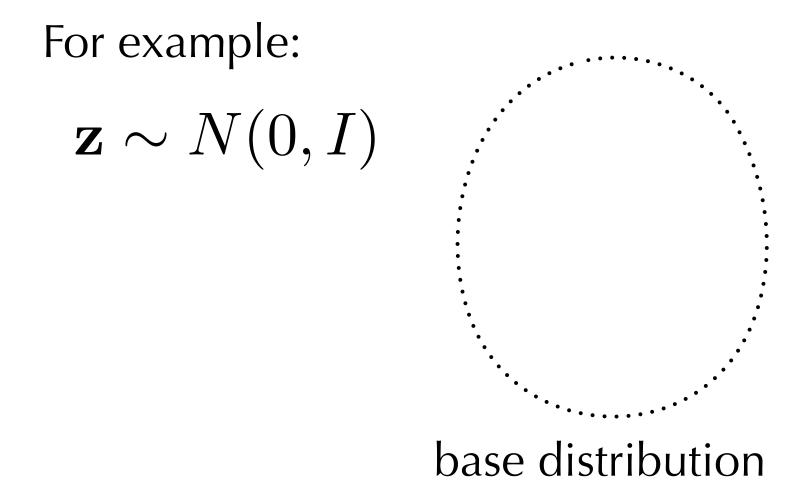


#### The variable transformed with the mapping

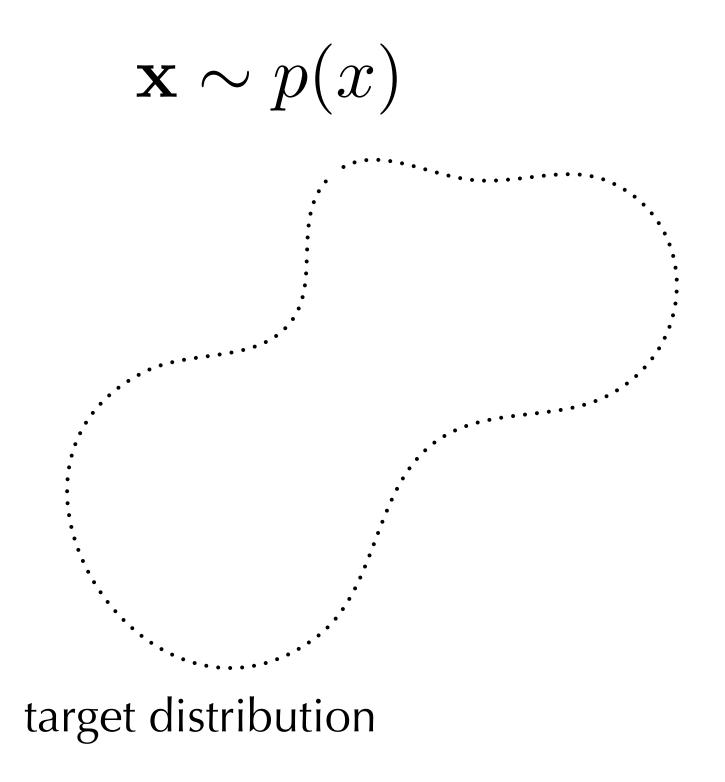


We take a random variable

$$\mathbf{z} \sim q(z)$$

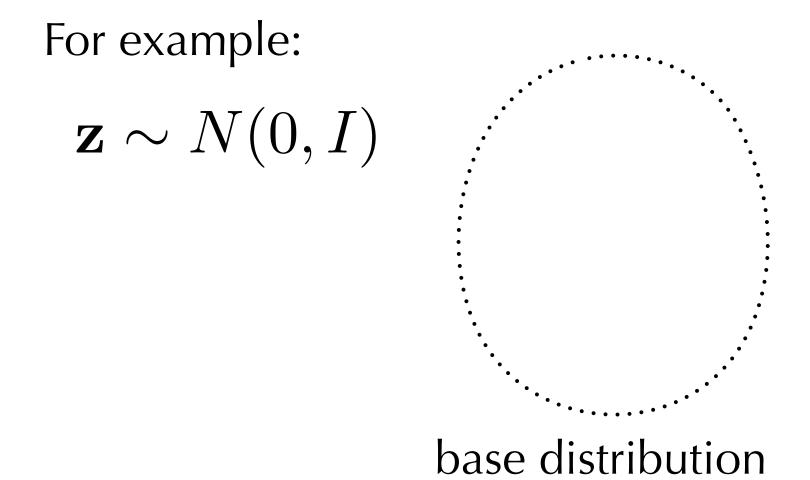


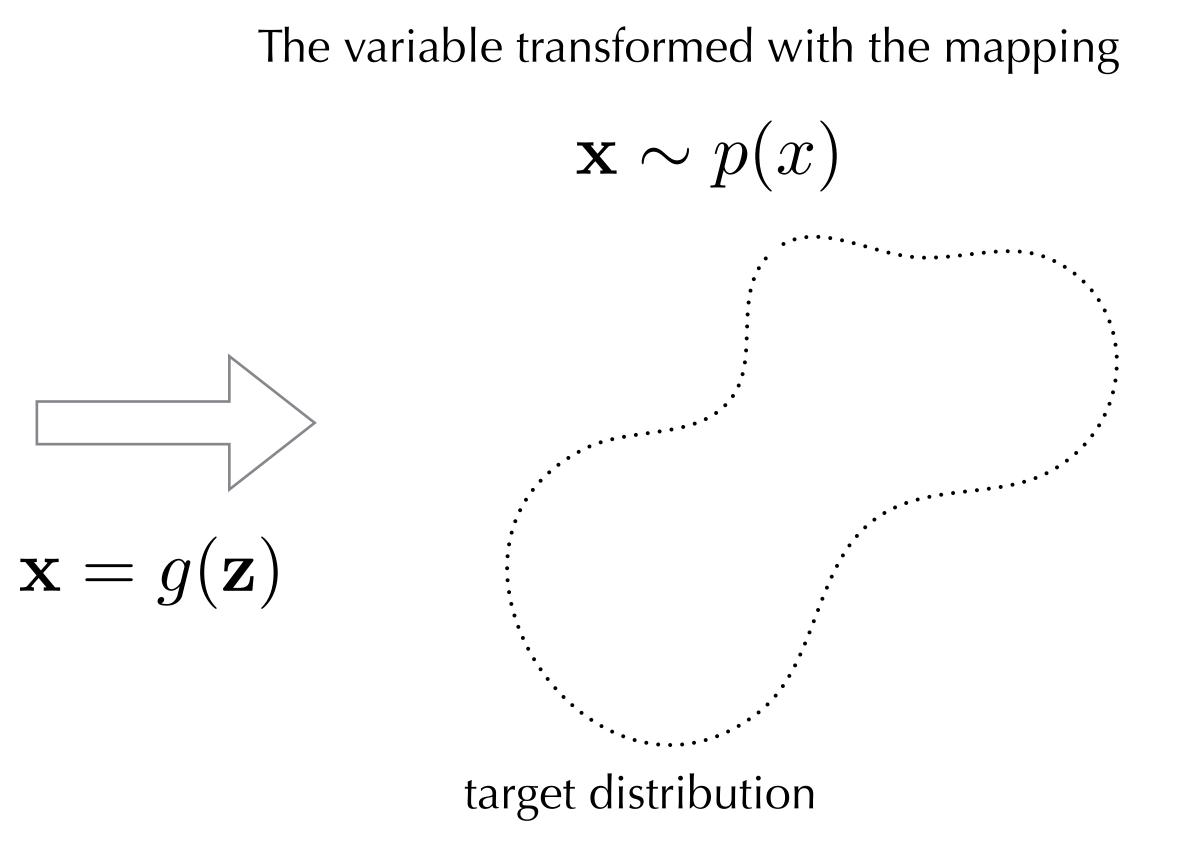
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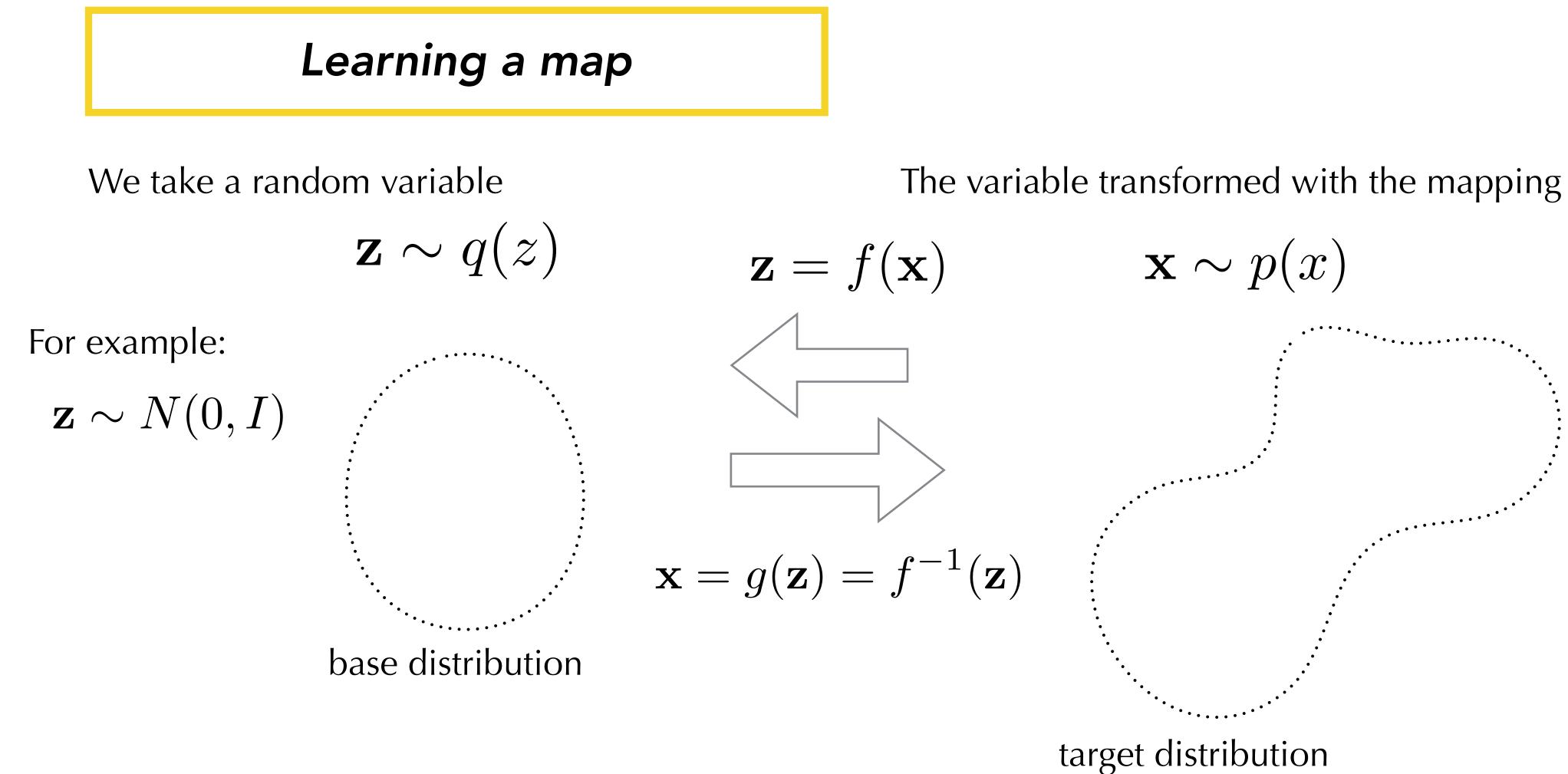


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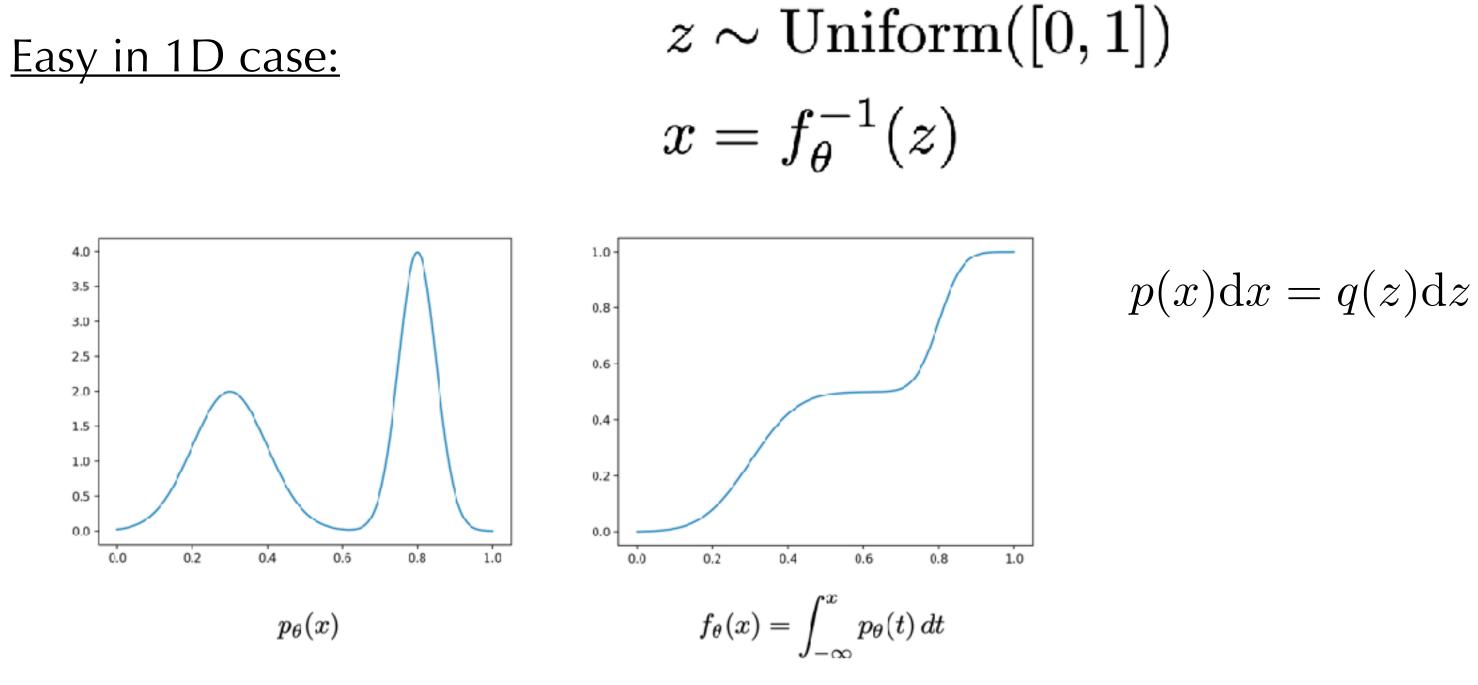


How to estimate the map?

$$\mathbf{x} = f^{-1}(\mathbf{z})$$

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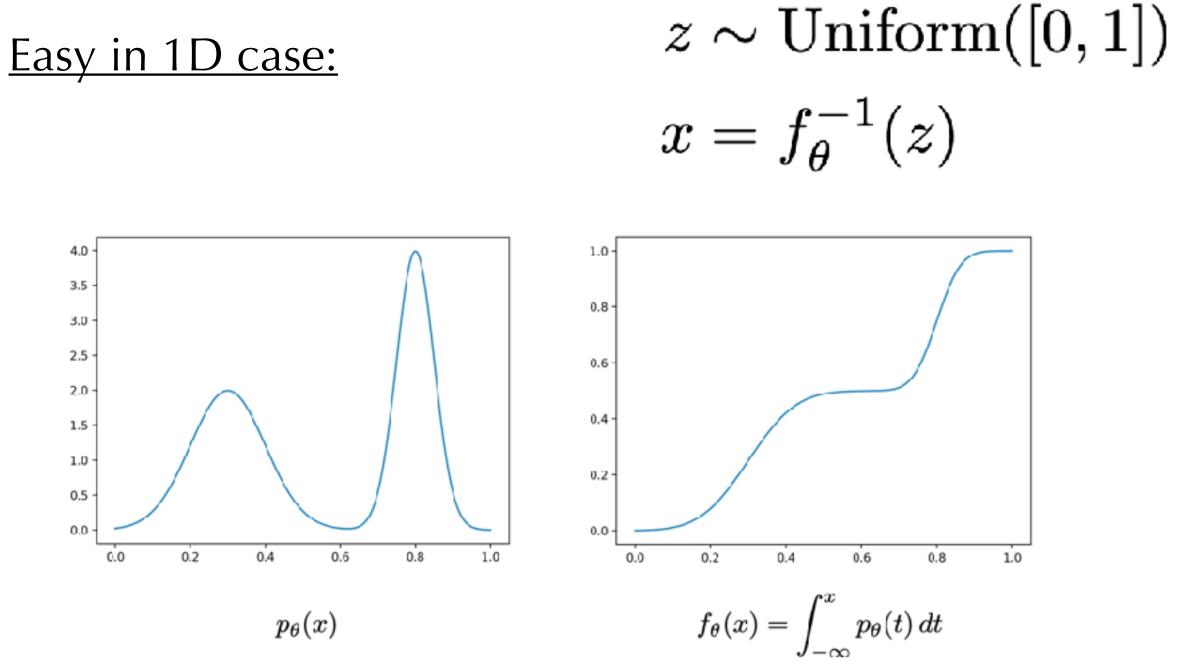


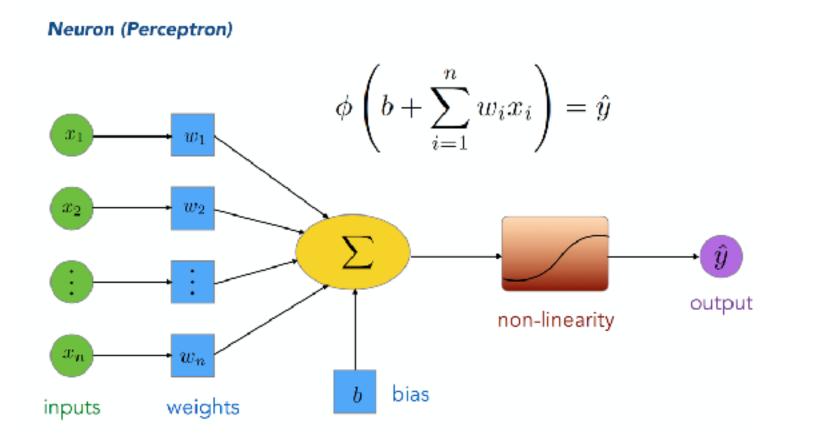
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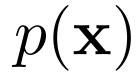
$$\mathbf{x} = f^{-1}(\mathbf{z})$$

#### Multidimensional case:

Parameterise a map by the Neural Network

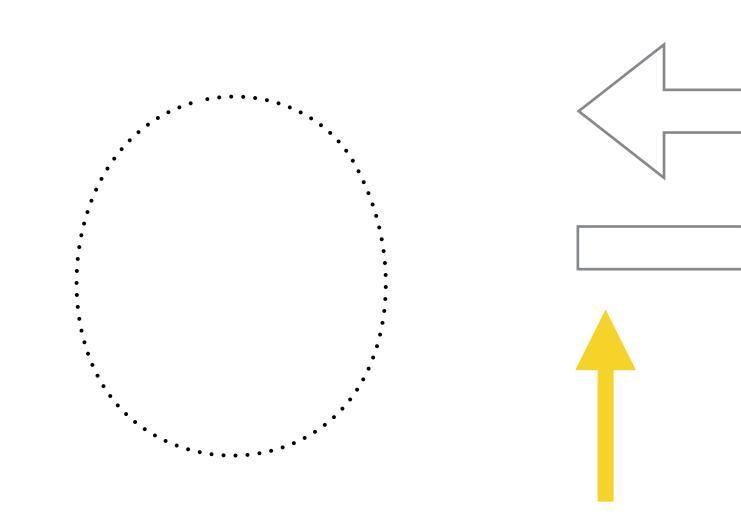






$$= q(f(\mathbf{x})) \left| \det \frac{\partial f(\mathbf{x})}{\partial \mathbf{x}} \right|$$

#### **Condition on the Waveform**

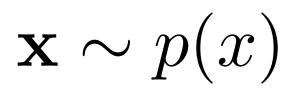


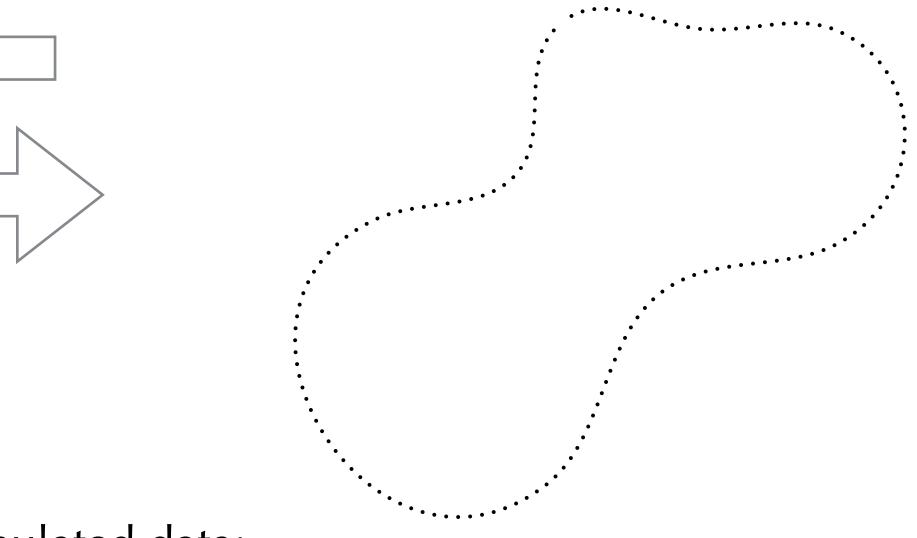
Condition map on the simulated data:

$$\mathbf{d} = h(\mathbf{x}) + \mathbf{n}$$

Therefore we have access to the joint sample:

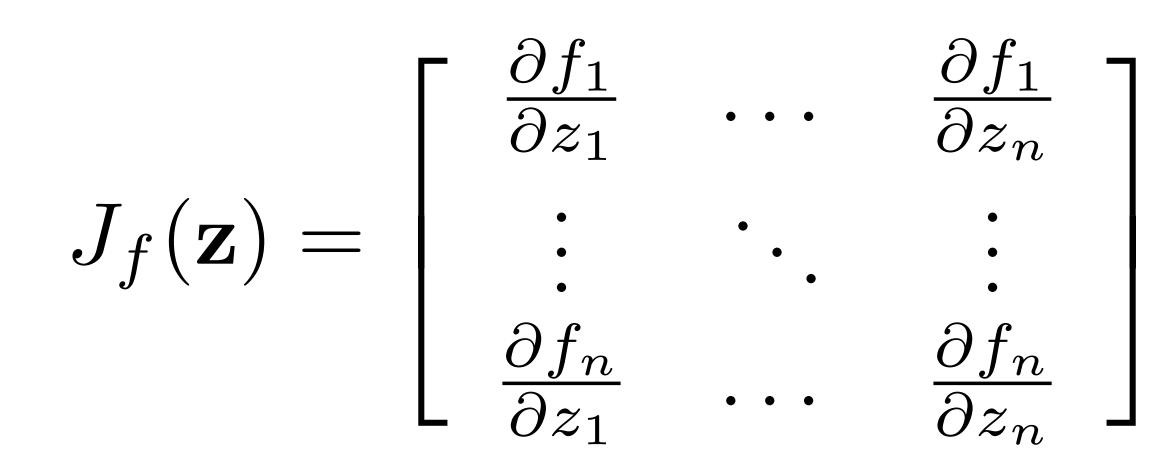
Samples from a prior of a physical parameter



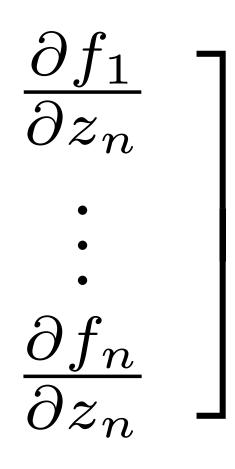


 $p(\mathbf{d}, \mathbf{x}) = p(\mathbf{x})p(\mathbf{d}|\mathbf{x})$ 

#### **Evaluation of Jacobian**



The calculation of determinant Jacobian will take  $O(n^3)$ To make it faster we have to ensure that the Jacobian is triangular Because the determinant of the triangular matrix is just a product of the diagonal elements



#### Affine transformations

location-scale transformation

 $\tau(z_i; \mathbf{h}_i) = \alpha_i z_i + \beta_i \qquad \mathbf{h}_i = -$ 

Invertibility for 
$$\alpha_i 
eq 0$$

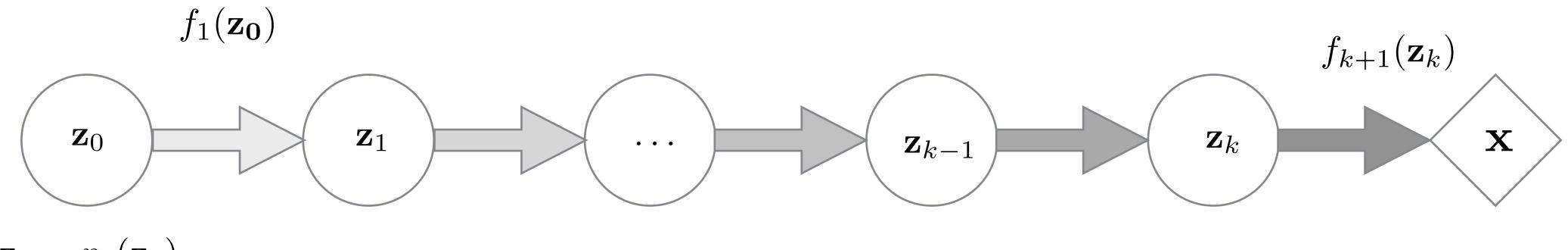
log-Jacobian becomes

$$\log |\det J_f(\mathbf{z})| = \sum_{i=1}^N \log$$

 $\mathbf{h}_i = \{\alpha_i, \beta_i\}$ 

# $|lpha_i|$

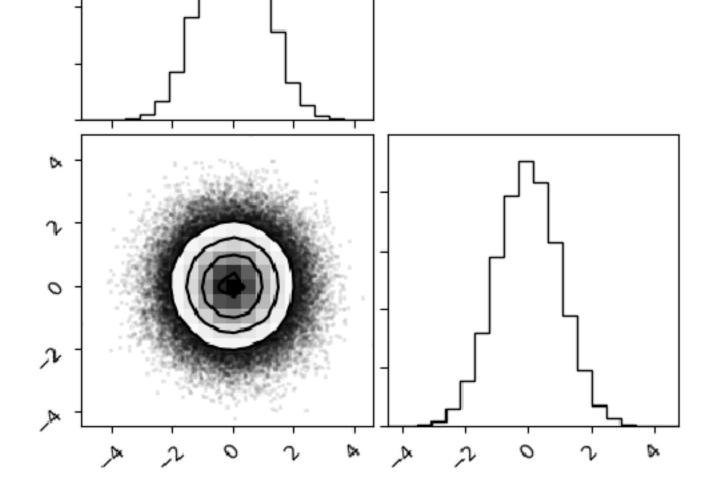
# **Combining transformations**

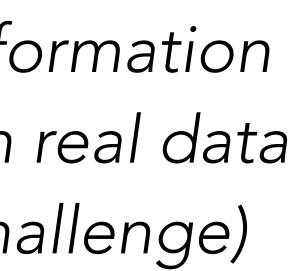


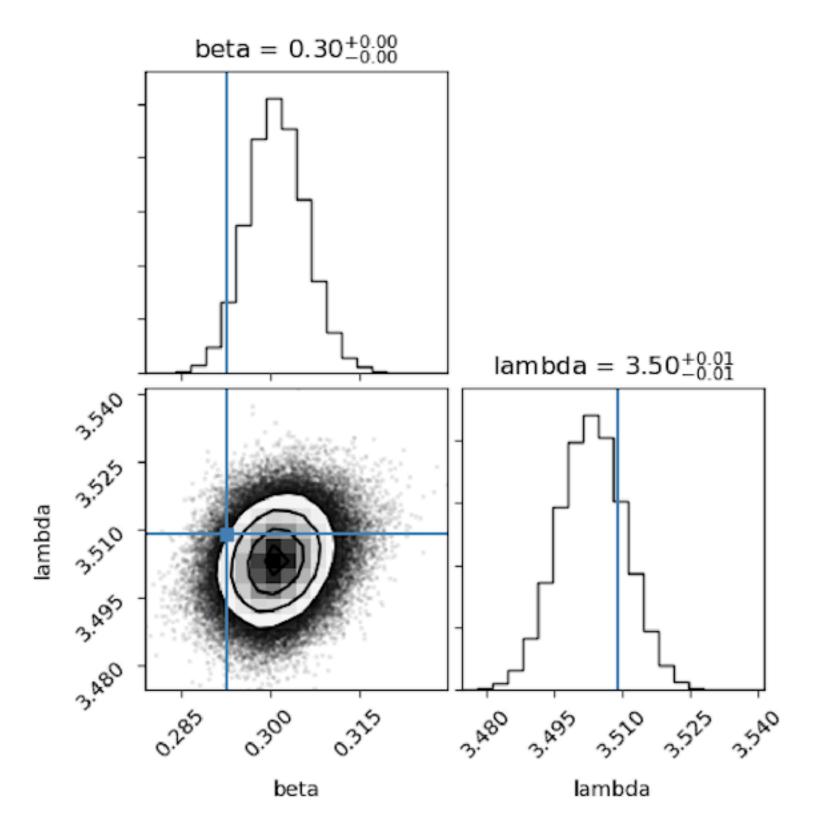
 $\mathbf{z}_0 \sim p_0(\mathbf{z}_0)$ 

# Example

# Learned transformation conditioned on real data (LISA Data Challenge)









New way to do Bayesian Inference for the Gravitational Wave data analysis

Time consuming calculations are done at the training time