

# Studying dark-energy with the large-scale structures of our Universe

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How the structures of the Universe can inform us about dark energy?

Observing structures with eBOSS

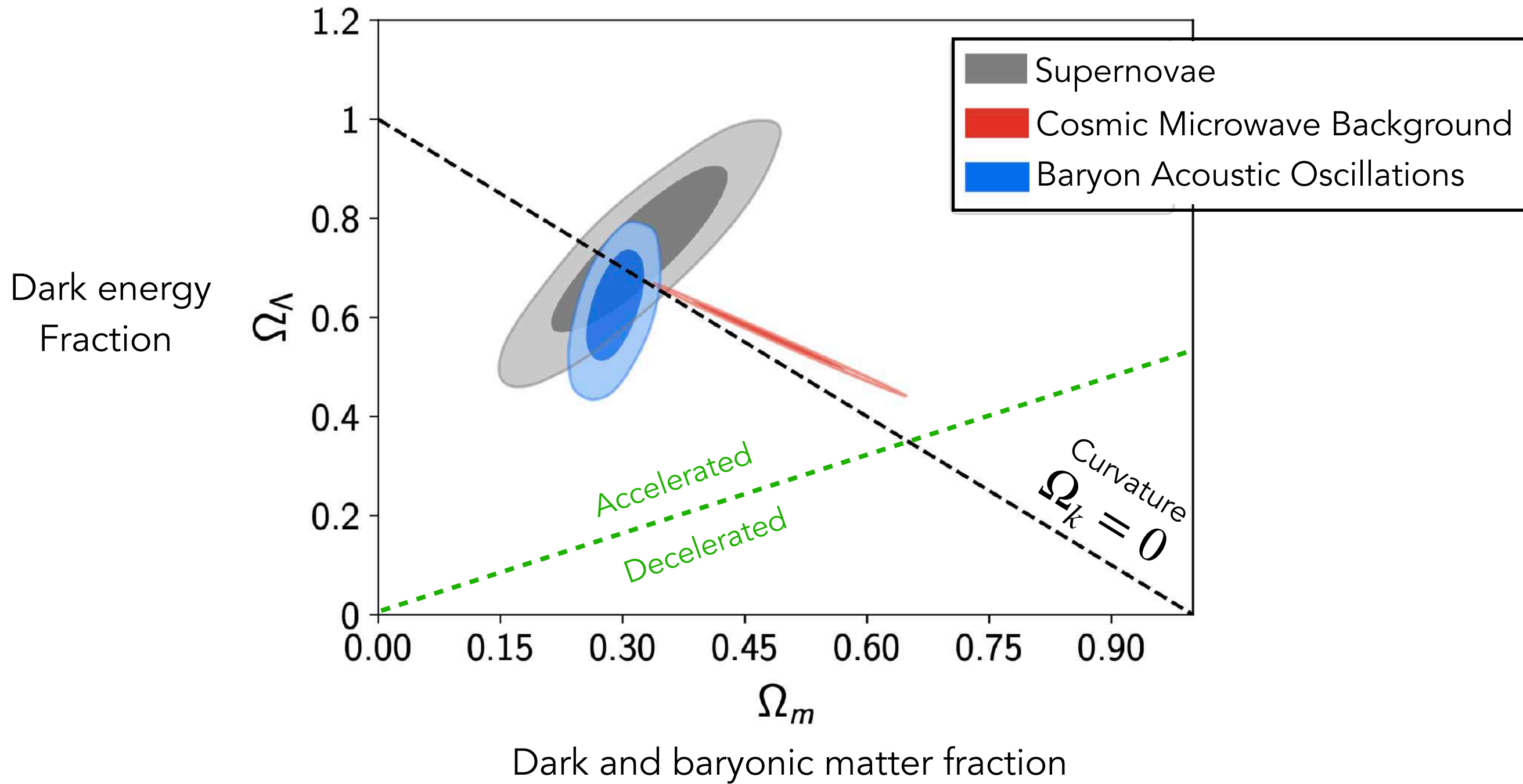
Analysis of the red galaxy sample from eBOSS

Future surveys

Cosmology with type-Ia supernovae

# Acceleration of the expansion of the Universe

eBOSS Collaboration 2020



**Independent probes confirm acceleration**

# Acceleration of the expansion of the Universe

General Relativity

+

$\Lambda$  : cosmological constant

or

$\Lambda, w_0, w_a$ : dark energy

or

Modifications or alternatives

to

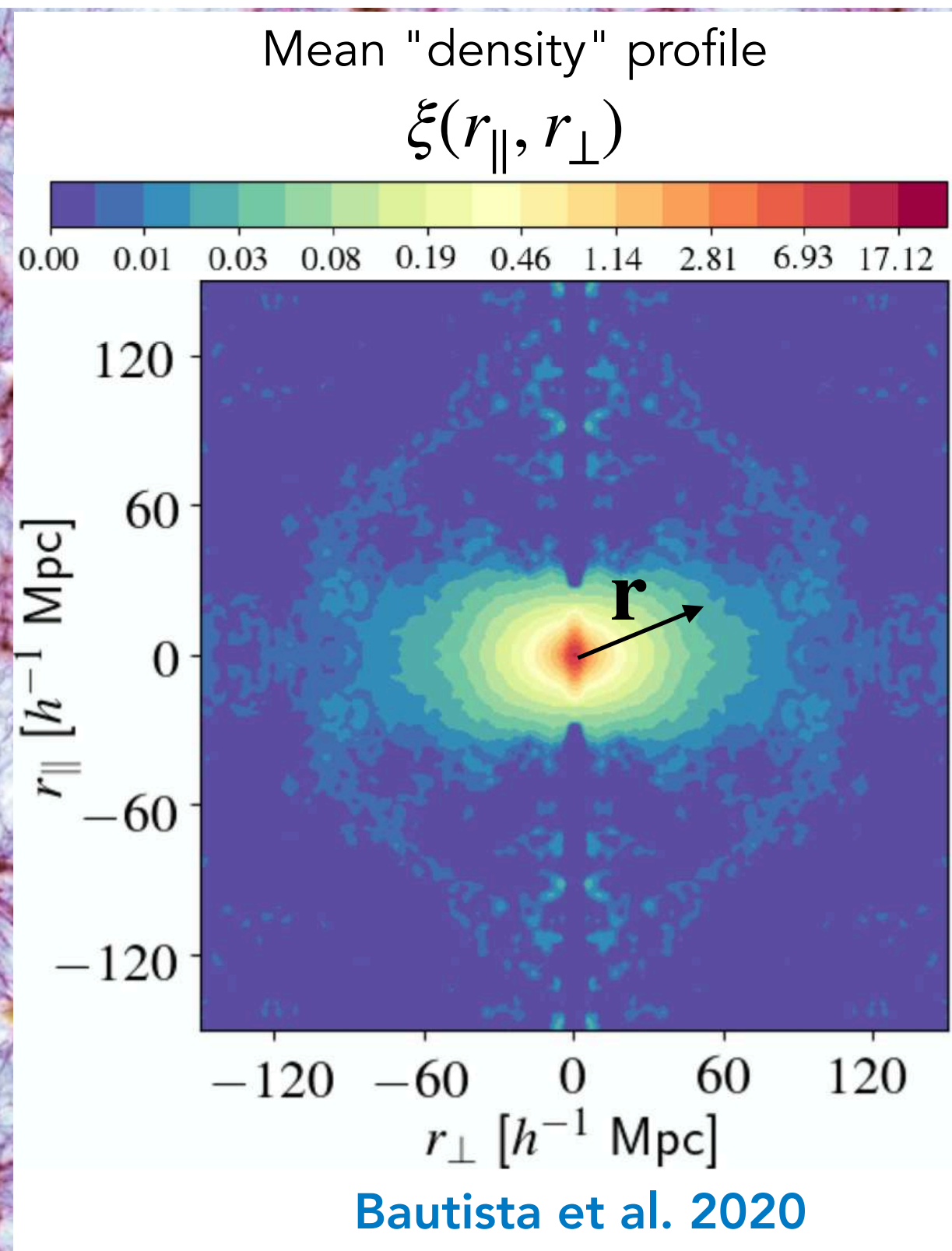
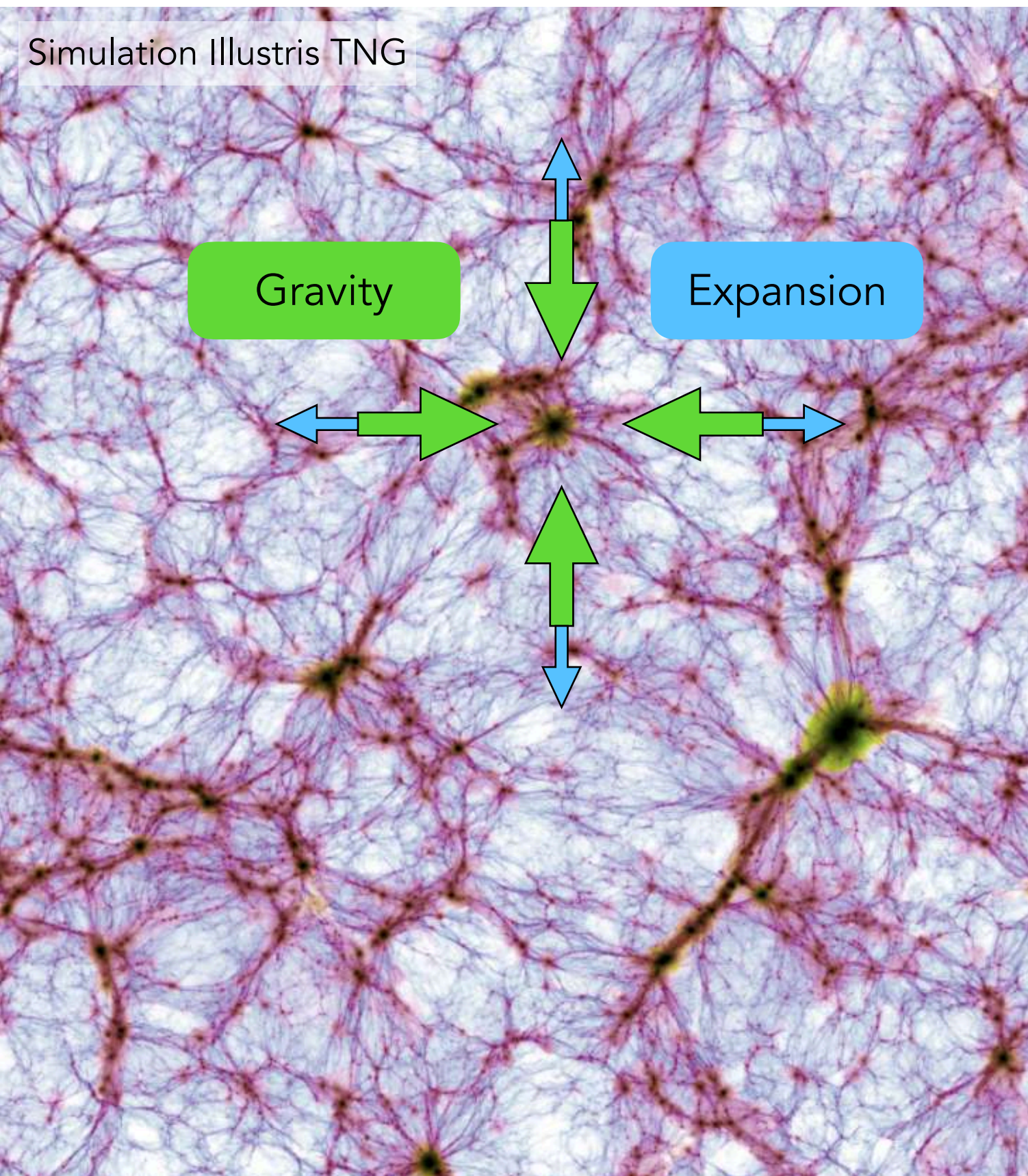
General Relativity

How to distinguish between these two ?



# Large-scale structures of the Universe

Simulation Illustris TNG





# Cosmological probes in the large-scale structures

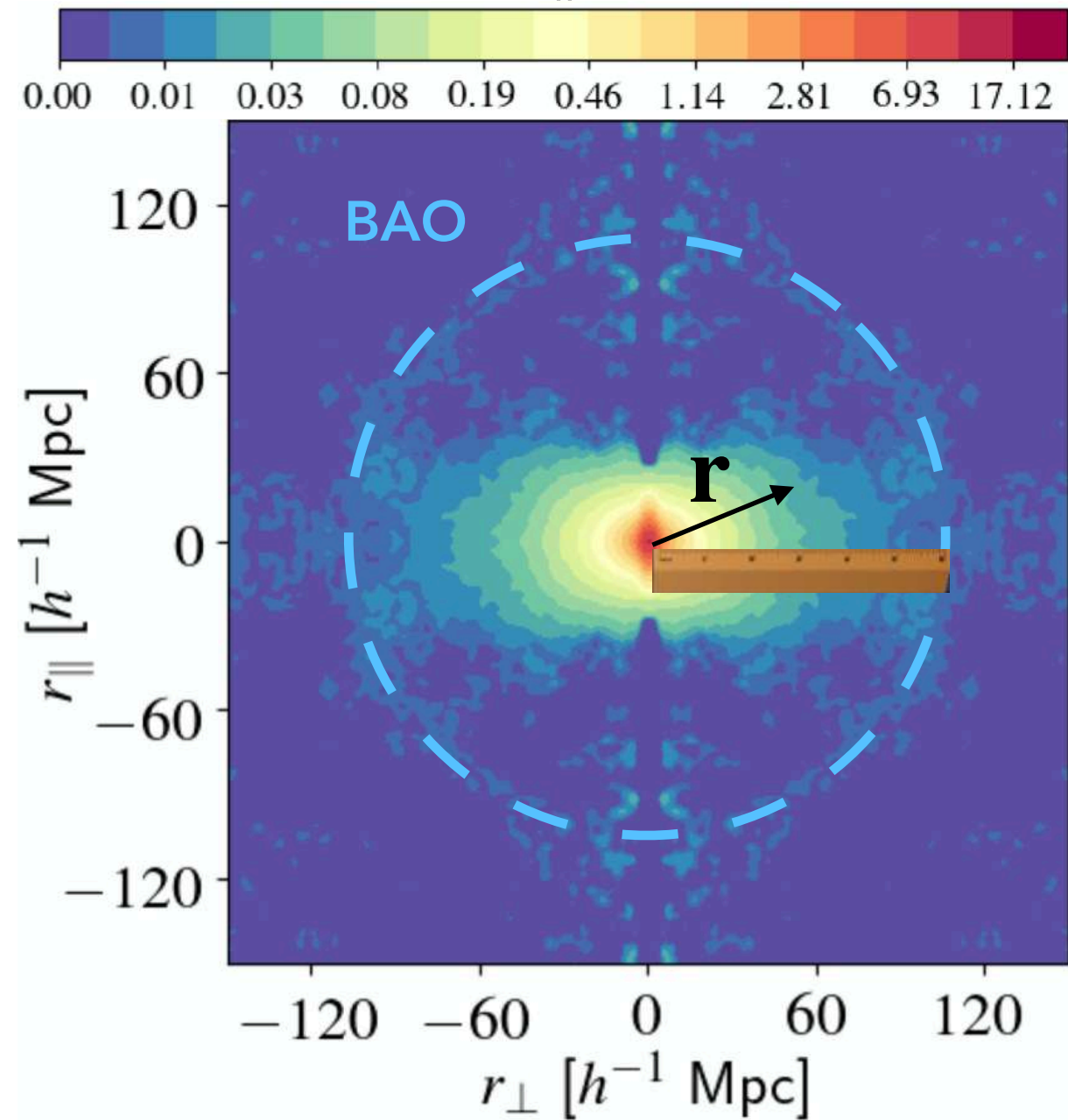
**BAO** - Baryon Acoustic Oscillations

Expansion rate  $H(z)$

General Relativity  
+  
 $\Lambda, w_0, w_a$ : dark energy

Mean "density" profile

$$\xi(r_{\parallel}, r_{\perp})$$



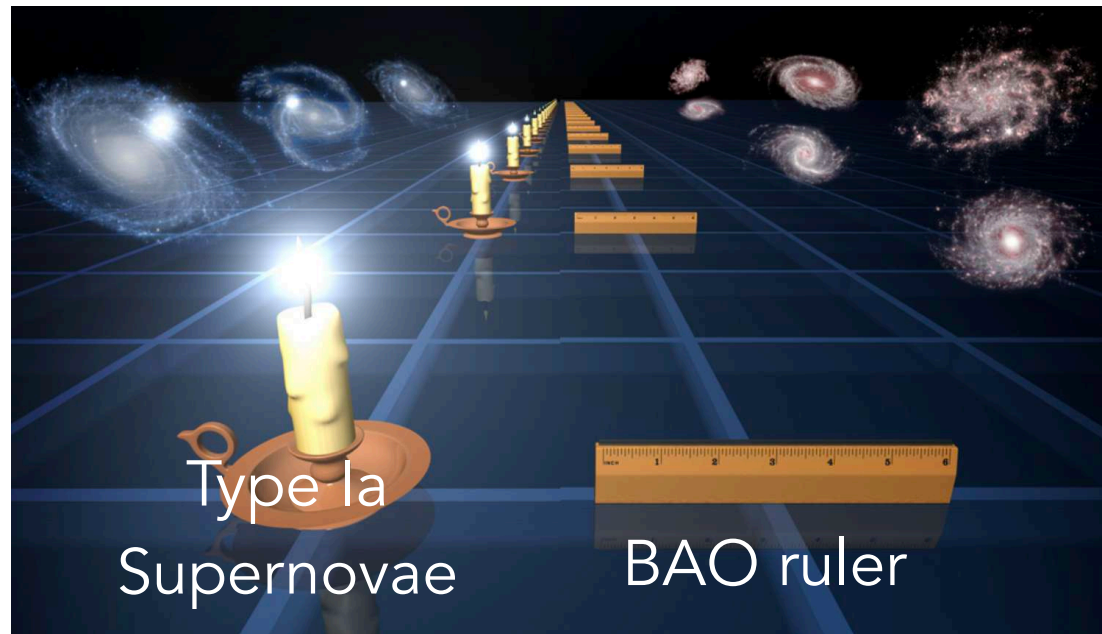
Bautista et al. 2020

# Cosmological probes in the large-scale structures

**BAO** - Baryon Acoustic Oscillations

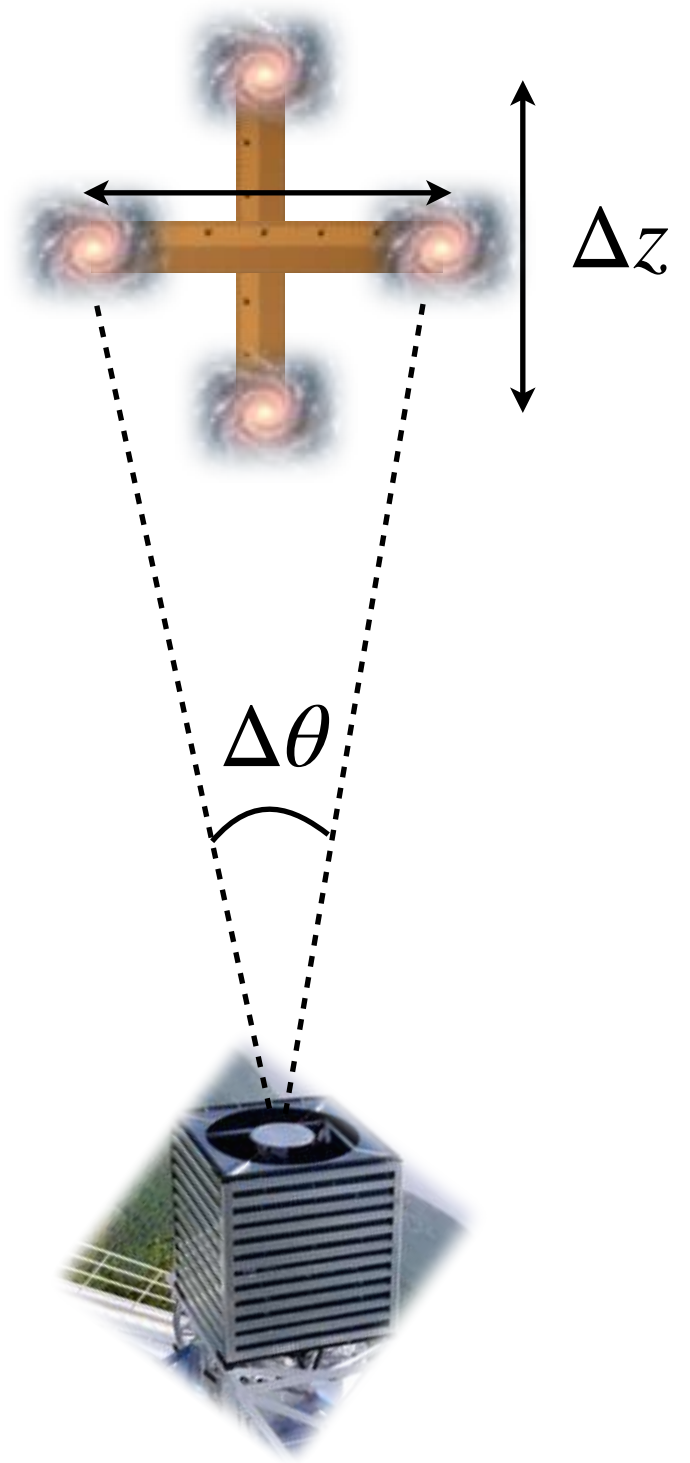
Expansion rate  $H(z)$

General Relativity  
+  
 $\Lambda, w_0, w_a$ : dark energy



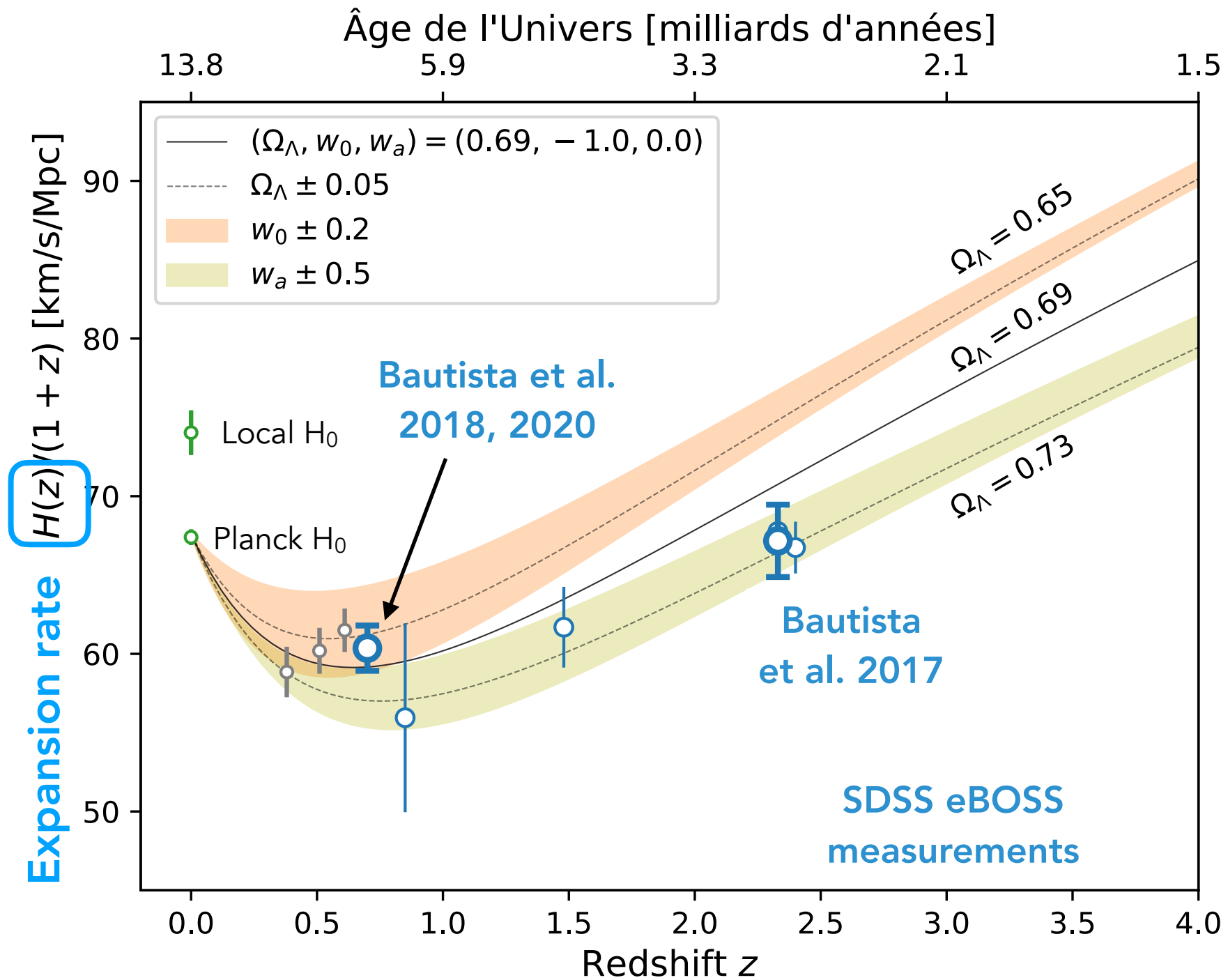
$$\Delta z = \frac{r_{\text{BAO}}}{c \overline{H(z)}}$$

$$\Delta \theta = \frac{r_{\text{BAO}}}{c \int \frac{dz}{H(z)}}$$



# Cosmological probes in the large-scale structures

## BAO - Baryon Acoustic Oscillations





# Cosmological probes in the large-scale structures

**BAO** - Baryon Acoustic Oscillations

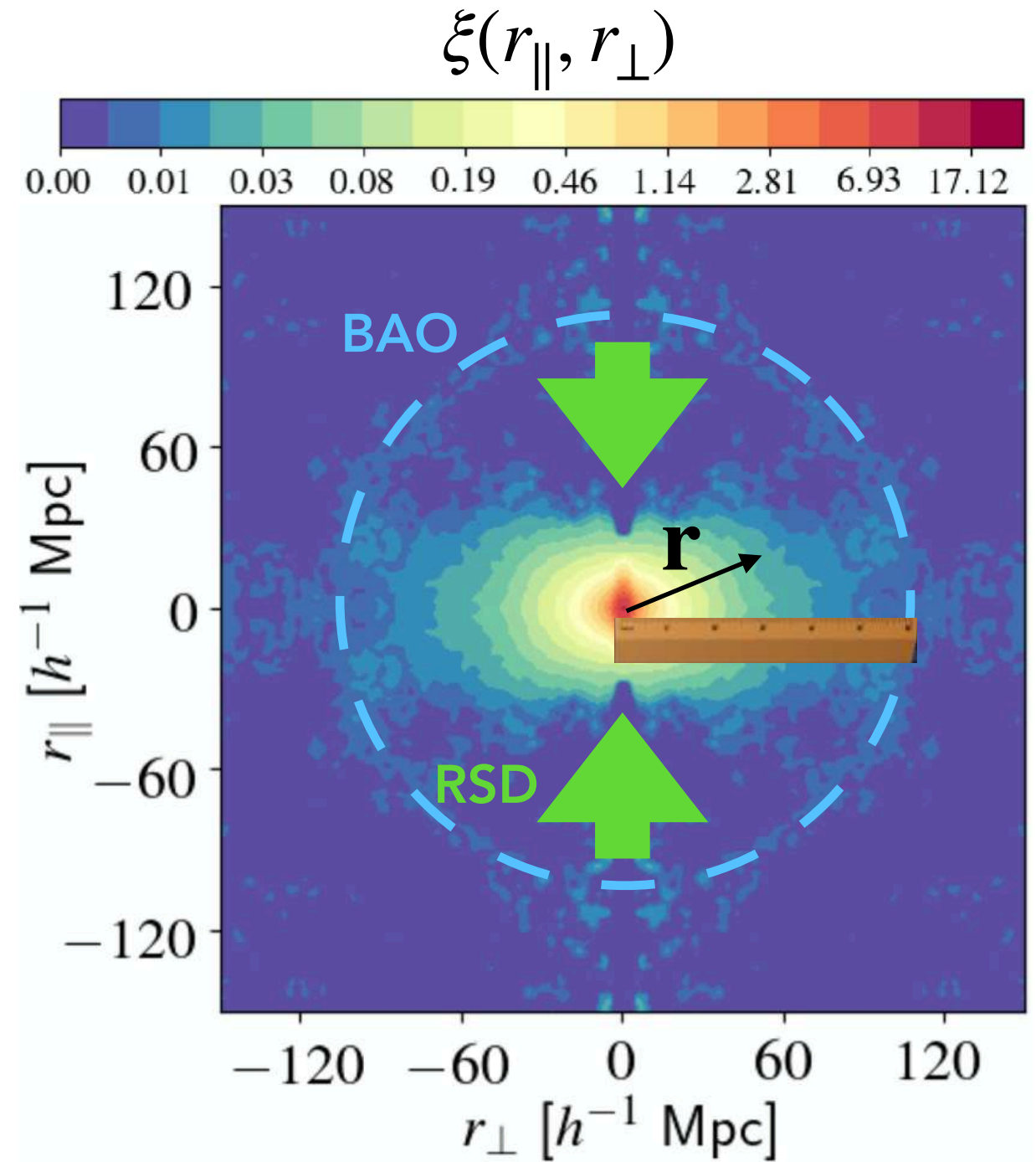
Expansion rate  $H(z)$

General Relativity  
+  
 $\Lambda, w_0, w_a$ : dark energy

**RSD** - Redshift-space distortions

Growth rate of structures  $f(z)$

Modifications or  
alternatives to  
General Relativity



Bautista et al. 2020

# Cosmological probes in the large-scale structures

## RSD - Redshift-space distortions

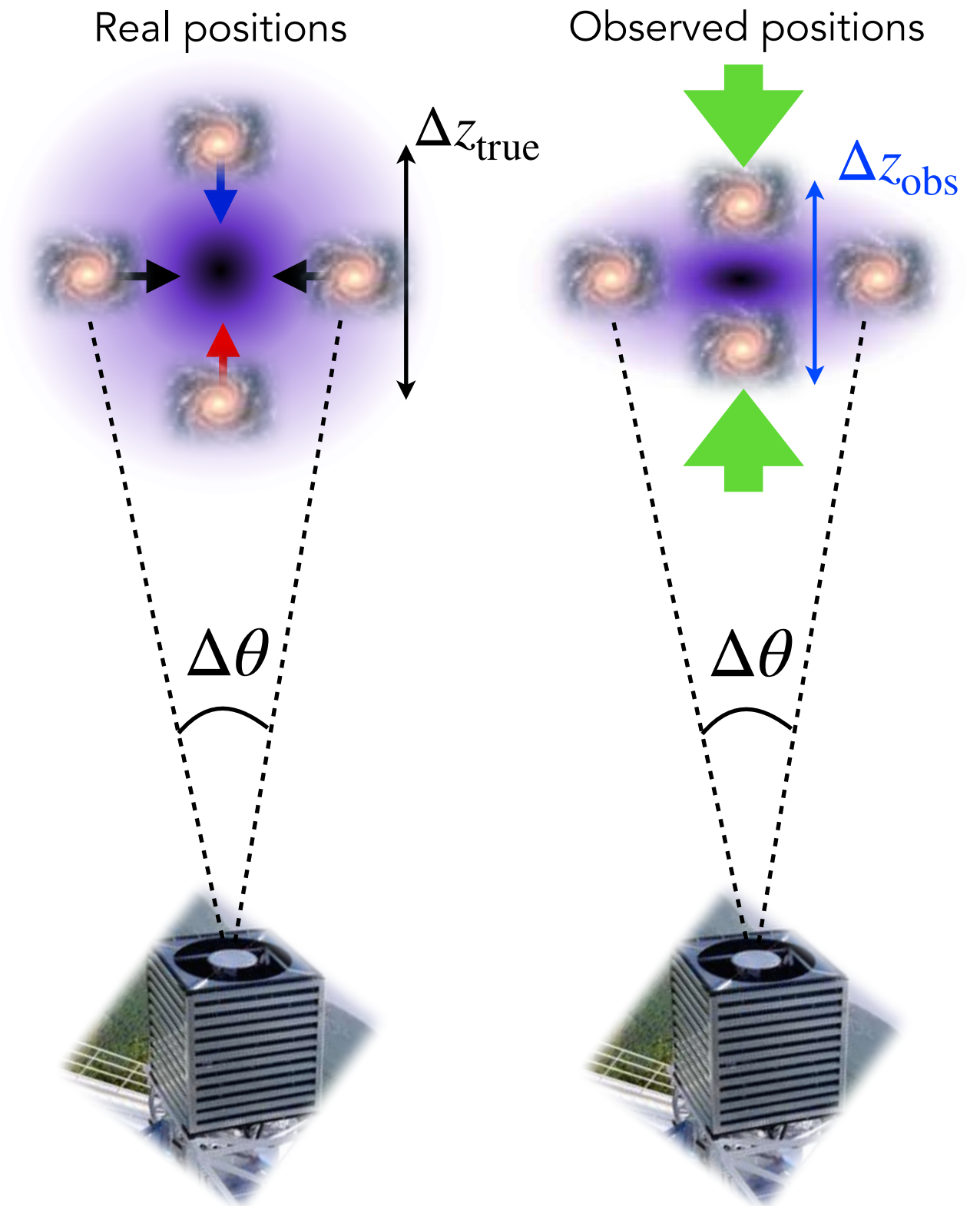
Velocities "flatten" the structures radially

Growth rate of structures in general relativity

$$f(z) \sim [\Omega_m(z)]^{\gamma=0.55}$$

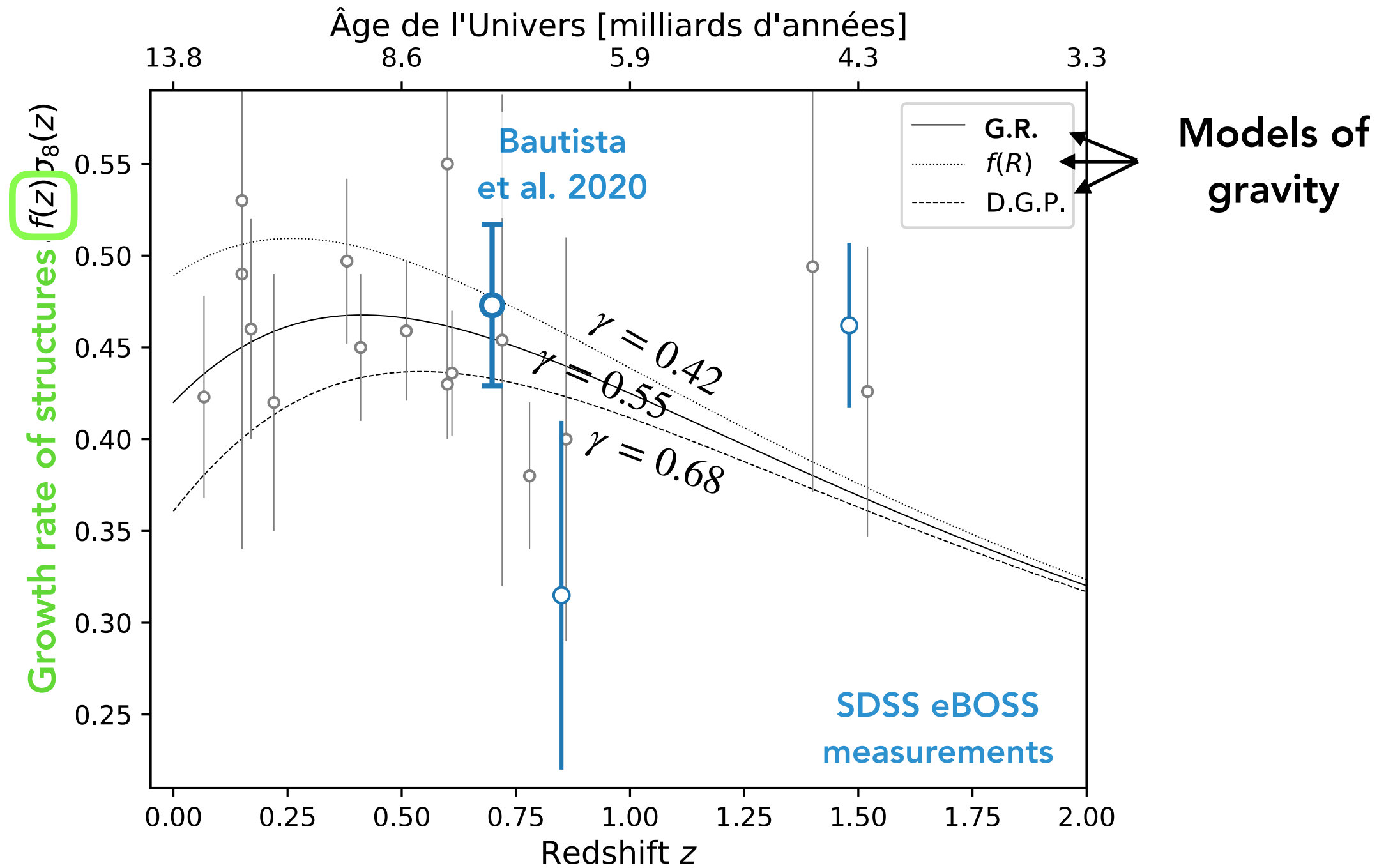
Else:  $\gamma \neq 0.55$

Modifications or alternatives to General Relativity



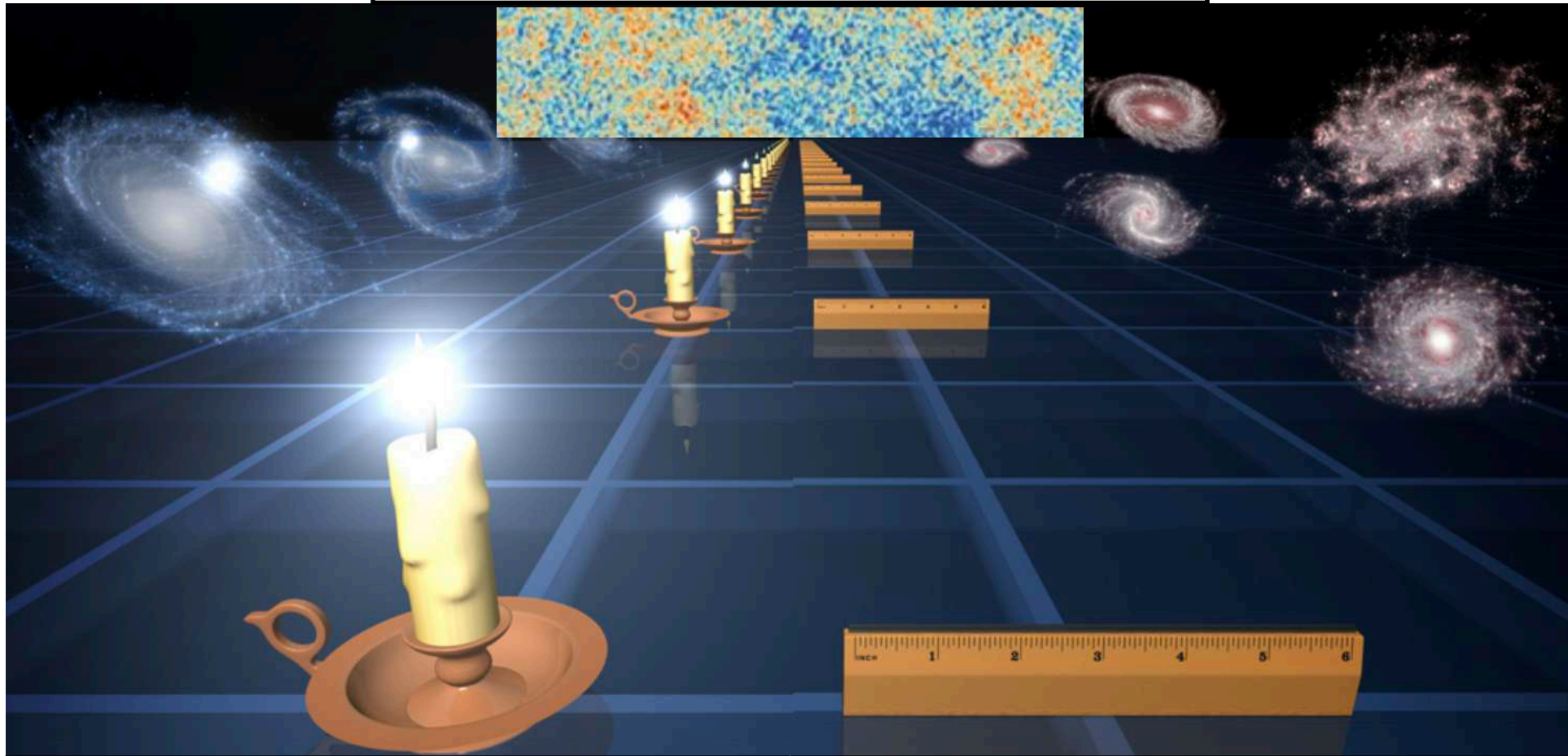
# Cosmological probes in the large-scale structures

## RSD - Redshift-space distortions



## Expansion rate $H(z)$

Cosmic microwave background (CMB)  
 $z \sim 1100$  or  $t \sim 380\,000$  years



Type-Ia Supernovae (SNIa)  
as standard candles  
 $0 < z < 1.5$   
 $5 \text{ Gy} < t < 13.8 \text{ Gy}$

Baryon Acoustic Oscillations (BAO)  
as standard ruler  
 $0.1 < z < 2.5$   
 $3 \text{ Gy} < t < 13 \text{ Gy}$

$$F = \frac{L_{\text{candle}}}{4\pi D_L^2(z)}$$

$$\Delta\theta = \frac{r_{\text{ruler}}}{D_M(z)}$$

$$\Delta z = \frac{r_{\text{ruler}}}{D_H(z)}$$



## Growth rate of structures $f(z)$

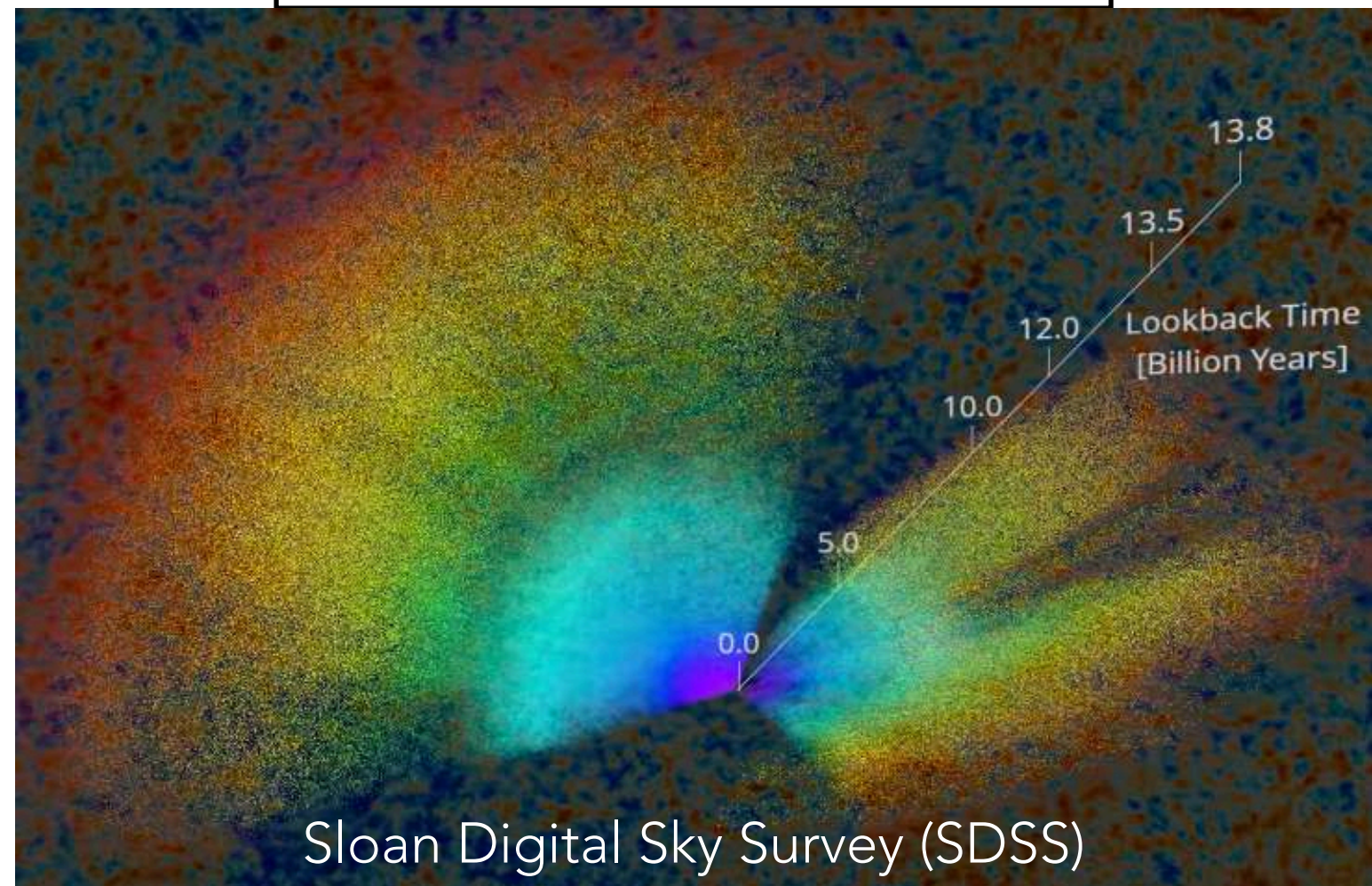
Cosmic microwave background (CMB)  
 $z \sim 1100$  or  $t \sim 380\,000$  years



Gravitational lensing  
 $0 < z < 1$   
 $6\text{ Gy} < z < 13.8\text{ Gy}$



RSD of  
galaxies, quasars, clusters, voids



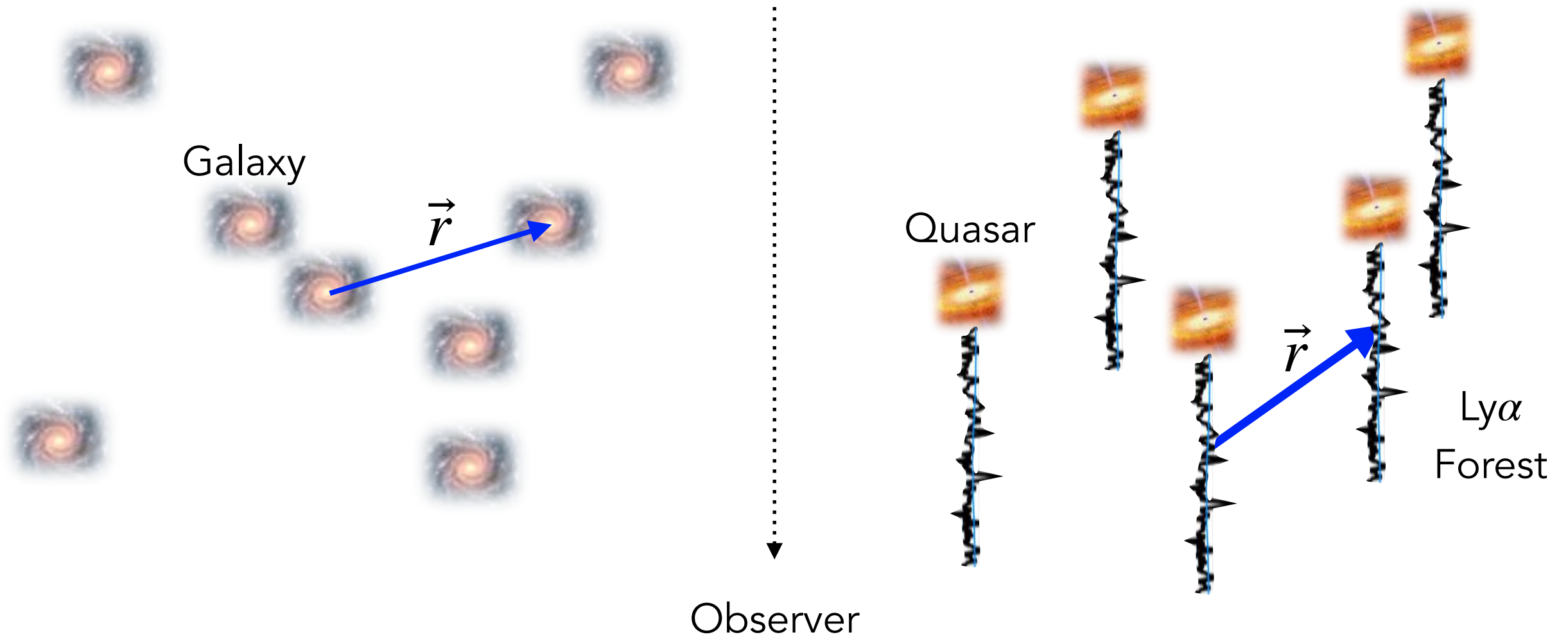
# Observing the structures

## Galaxies

- low redshift,  $z < 2$
- trace high density regions
- traditional method

## Lyman- $\alpha$ forests

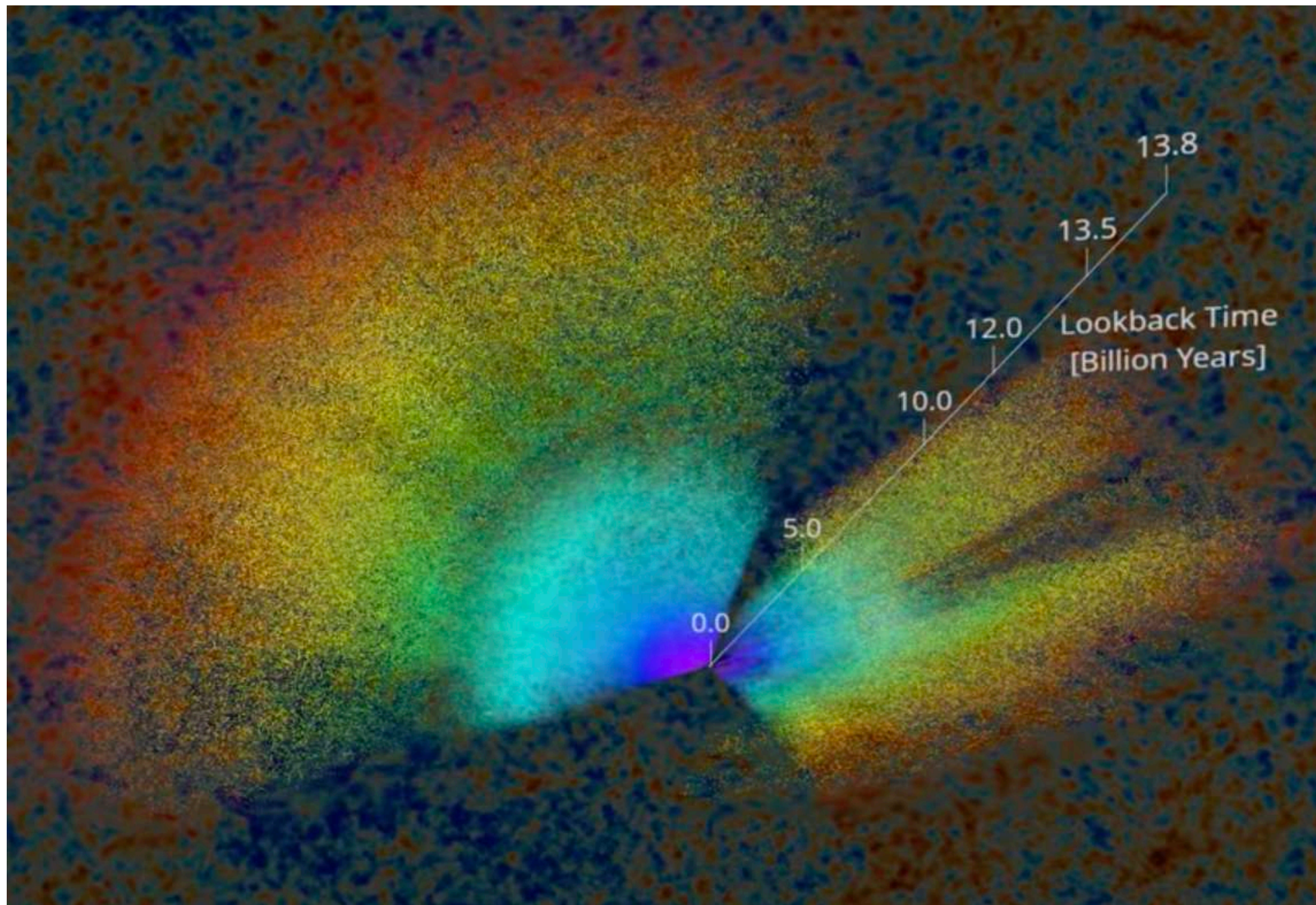
- unique\* access to high redshift,  $z > 2$
- trace low density in the line-of-sight
- recent method



# eBOSS

and the state-of-the-art map of the Universe's structures





## 20 years of redshift surveys with SDSS

<https://www.youtube.com/watch?v=KJJXbcf8kxA> (by EPFL.ch)



# eBOSS

extended Baryon Oscillation Spectroscopic Survey

Dawson et al. 2016

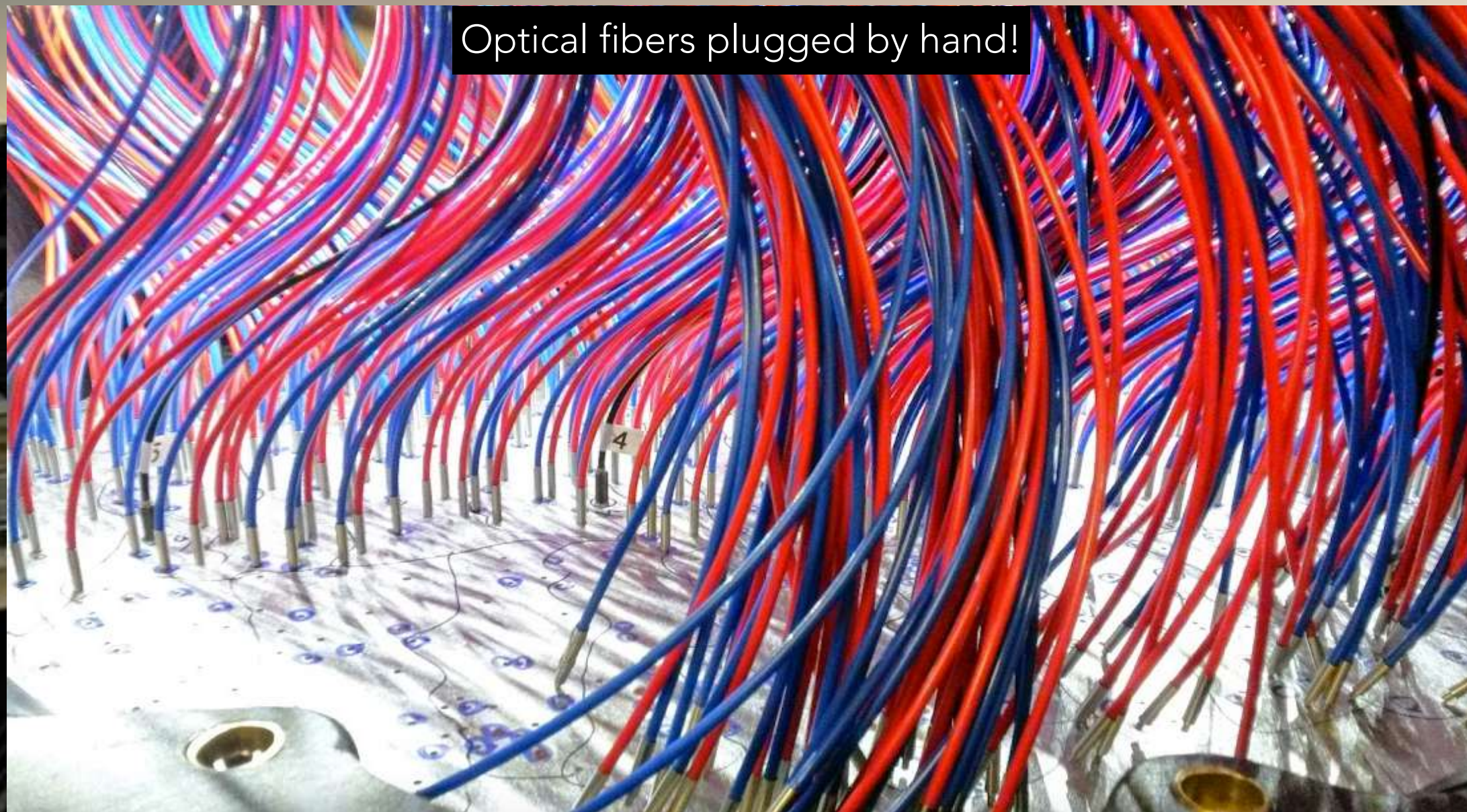
2.5-meter  
mirror

Sloan Digital Sky Survey Telescope  
Apache Point Observatory, New Mexico, USA



# eBOSS

extended Baryon Oscillation Spectroscopic Survey



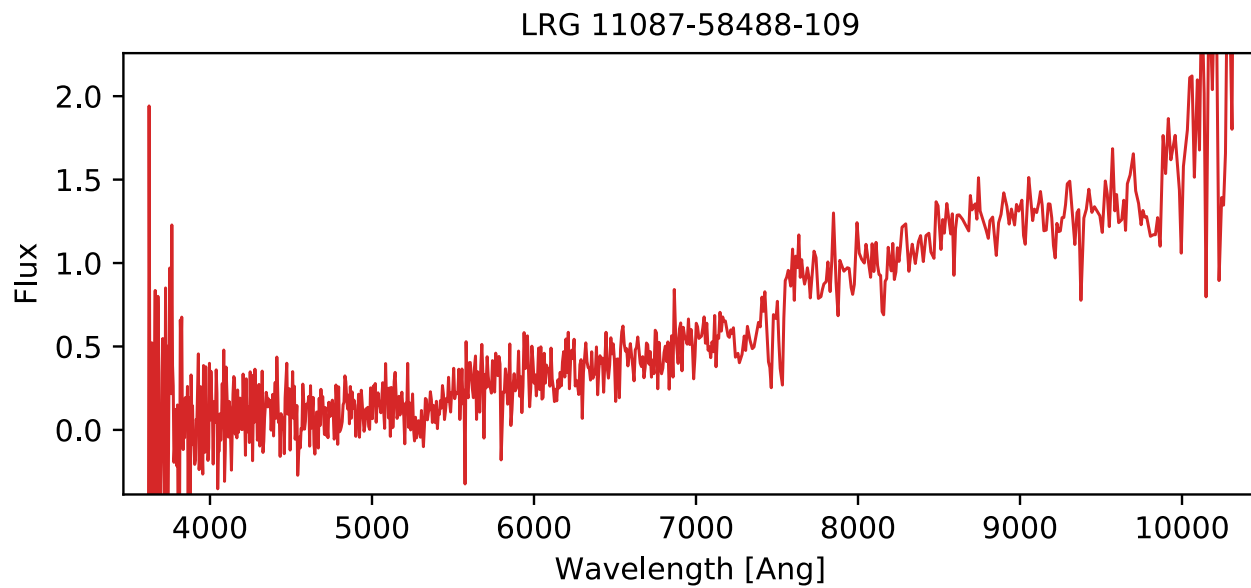
Optical fibers plugged by hand!

1-meter focal plane

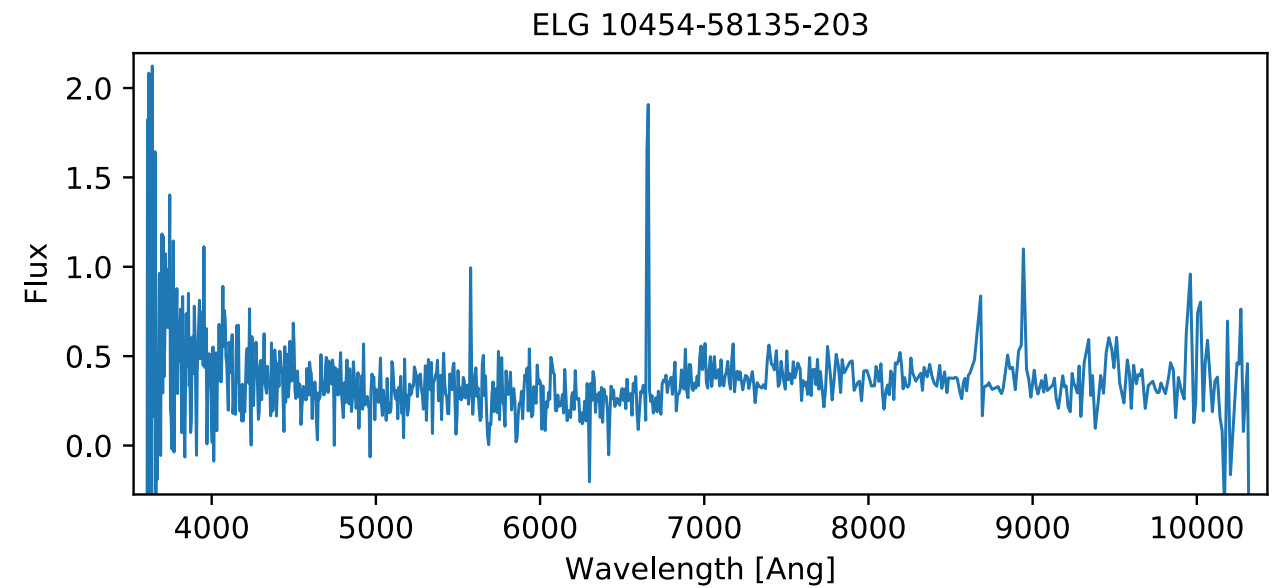


# eBOSS Spectra

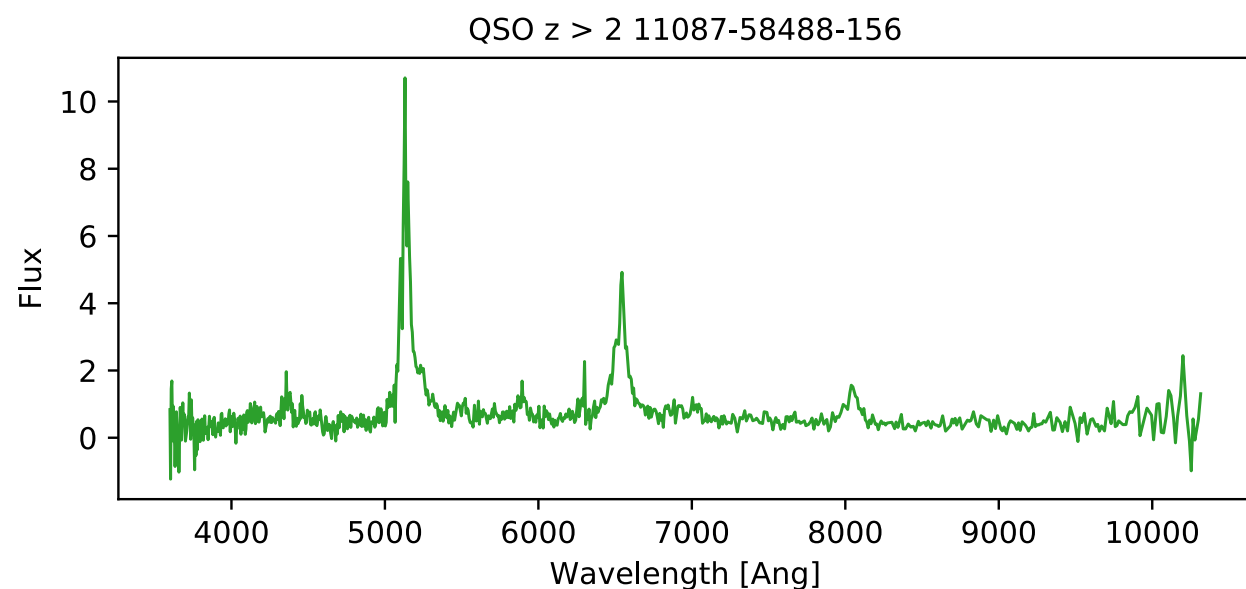
Luminous Red Galaxies ( $0.6 < z < 1.0$ )



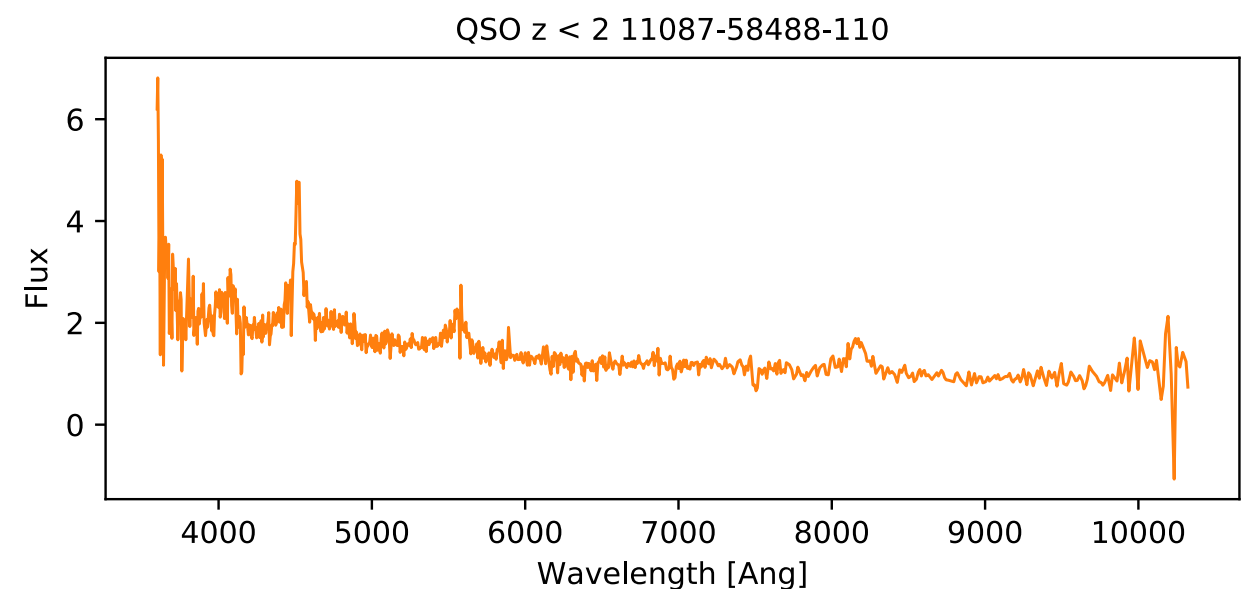
Emission Line Galaxies ( $0.7 < z < 1.1$ )



Quasars ( $z > 2$ ) for Lyman-alpha forest



Quasars for clustering ( $0.8 < z < 2.2$ )




Redshifts obtained with PCA templates, neural networks  
and some visual inspection (for QSOs)

# Extracting cosmological overdensities

Ross, **Bautista**, Tojeiro et al. 2020

Galaxy overdensity:  $\delta_g(\vec{x}) = \frac{n_g(\vec{x})}{\bar{n}_g} - 1$

Random catalog



Features or systematics to take into account:

Survey footprint

Observational completeness

Fake overdensities caused by photometry

"Collisions" of fibers

Reconstruction of linear density field for BAO

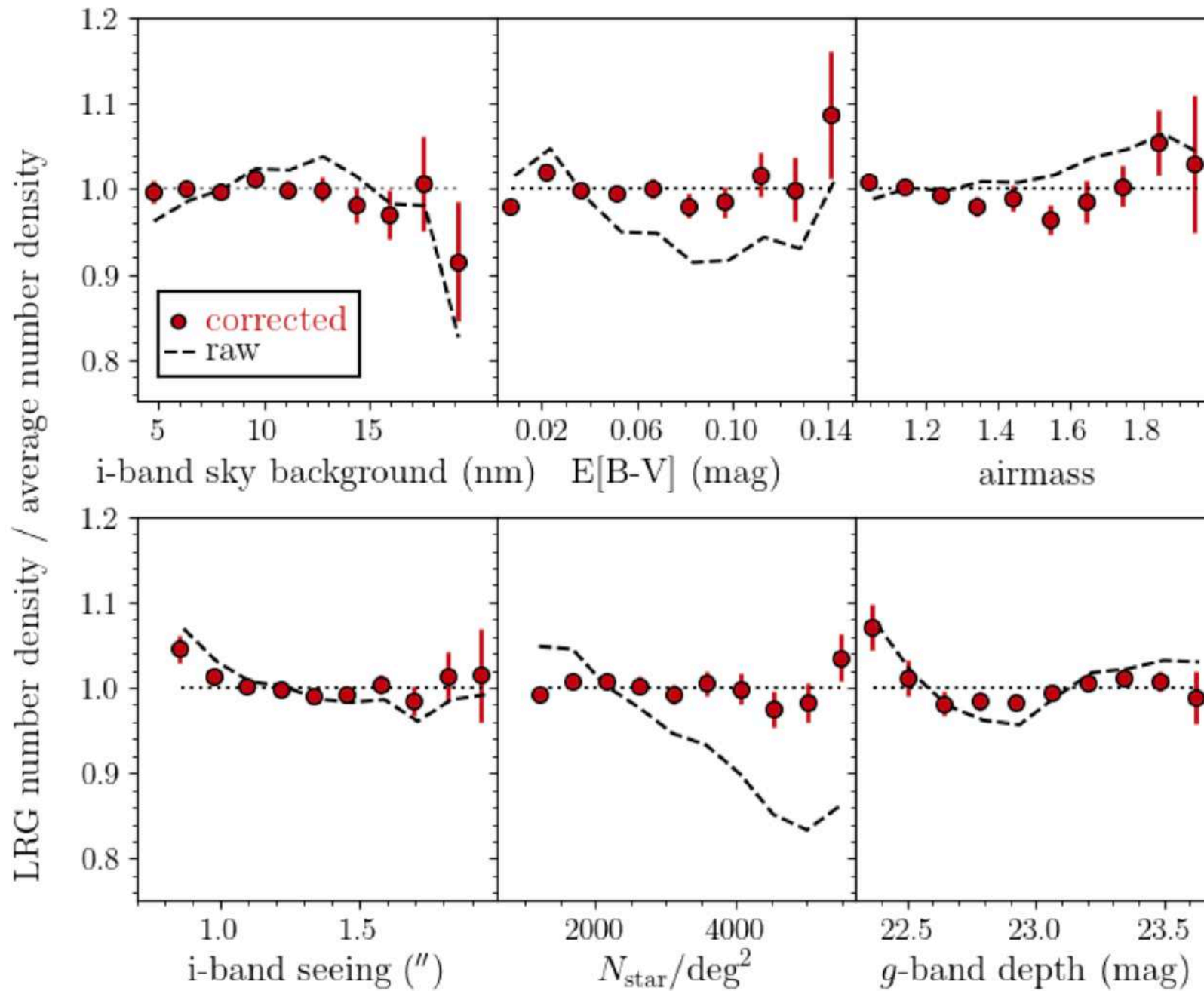
Spectra without confident redshift measurement

1000 simulated surveys used to test methods, covariance, systematic errors  
(Zhao, Chuang, **Bautista**, et al. 2020)



# Extracting cosmological overdensities

Ross, **Bautista**, Tojeiro et al. 2020

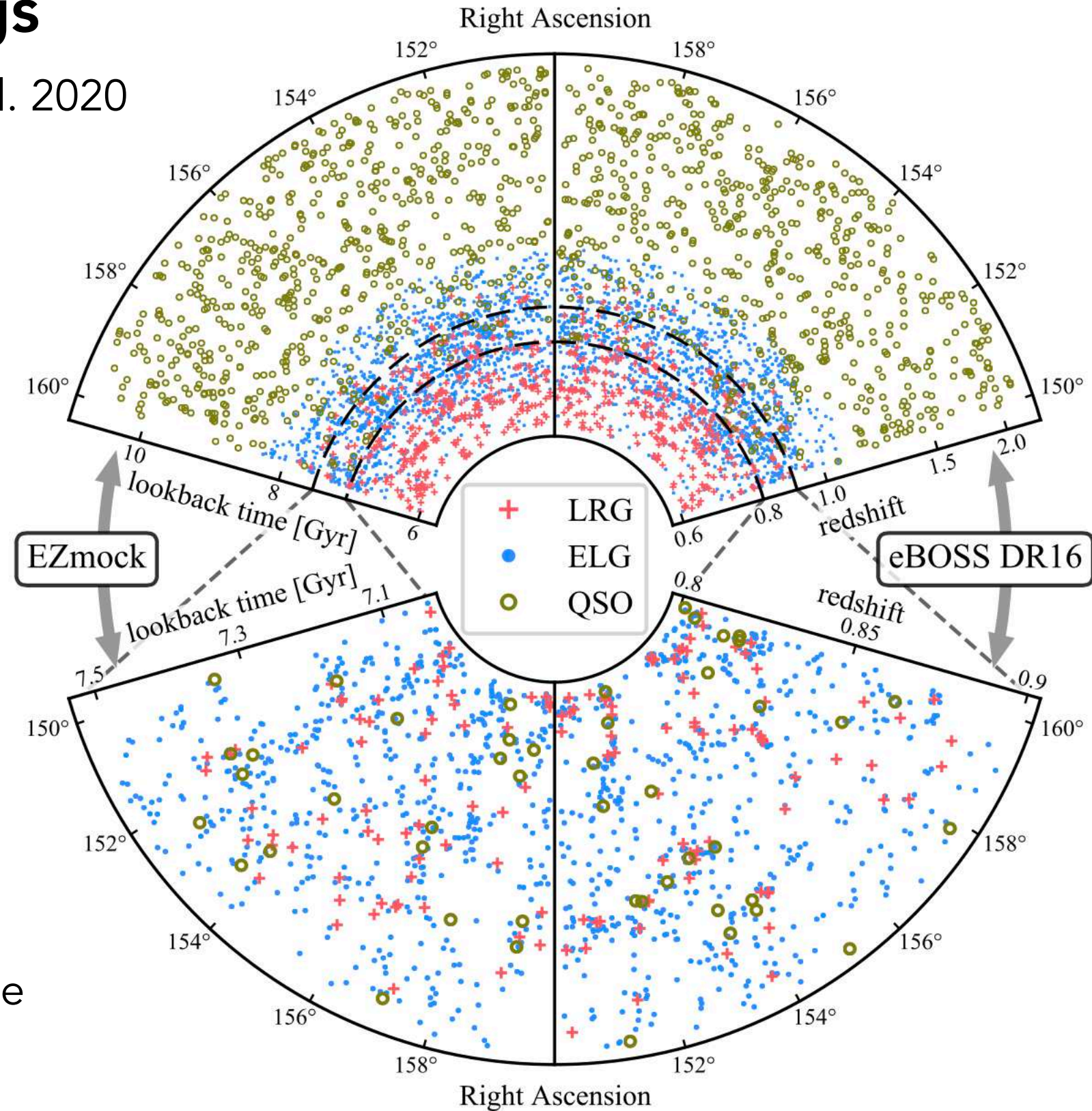


Fake overdensities  
caused by photometry

# EZmock catalogs

Zhao, Chuang, **Bautista**, et al. 2020

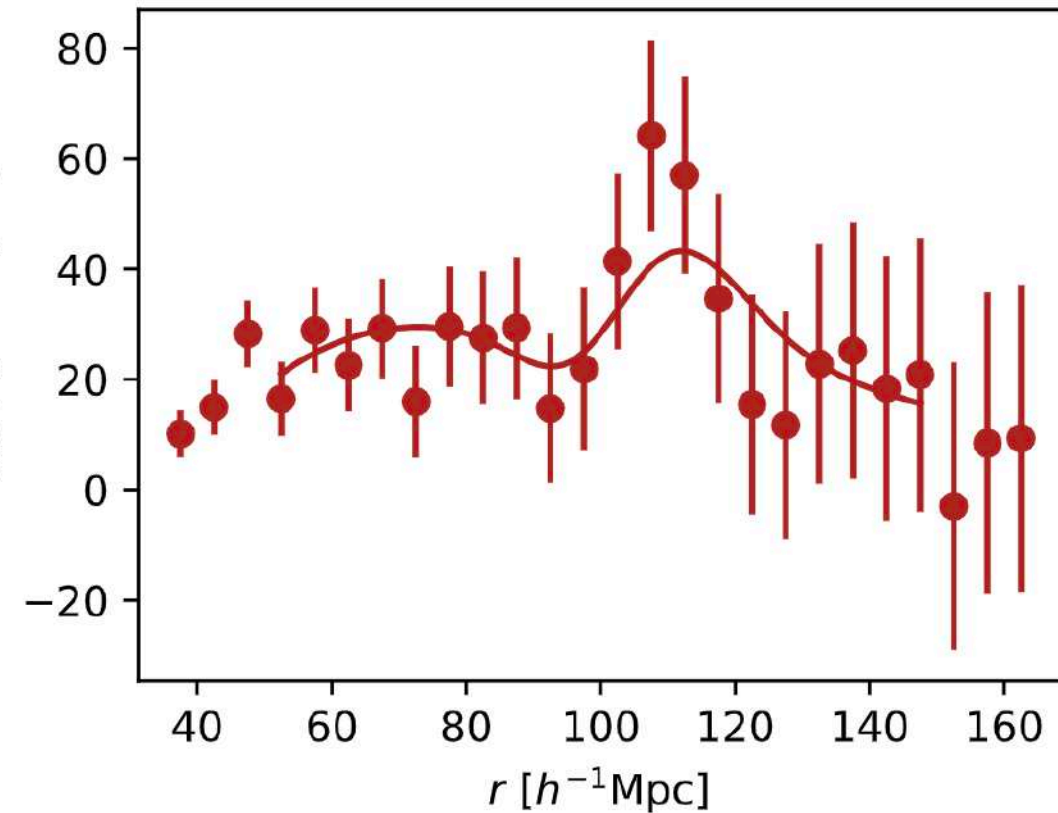
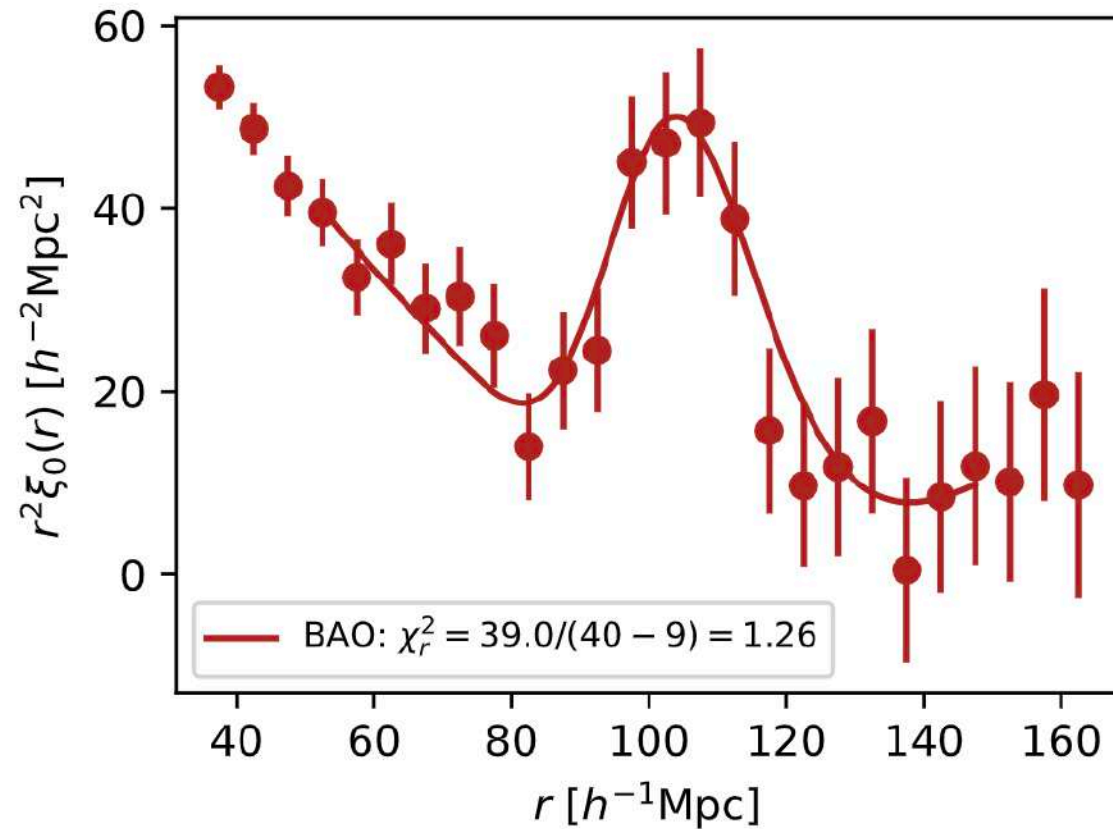
- Zel'dovich approximations to rapidly construct density field
- 1000 realizations of the survey
- includes redshift evolution
- includes observational effects
- includes cross-correlations between tracers
- used to test our methods, estimate systematic errors and compute covariance matrices



# BAO analysis

eBOSS LRG sample

Post-reconstruction multipoles



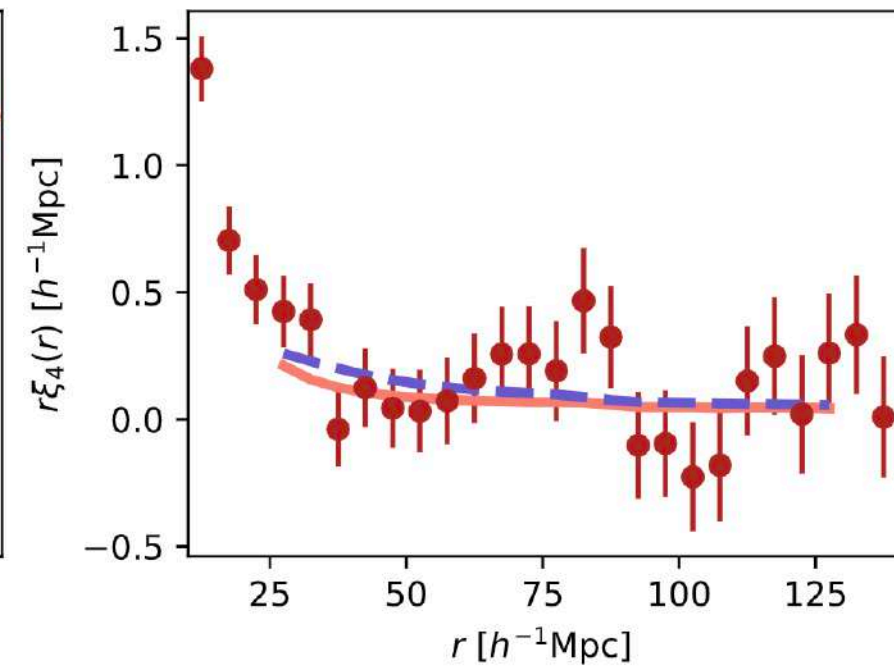
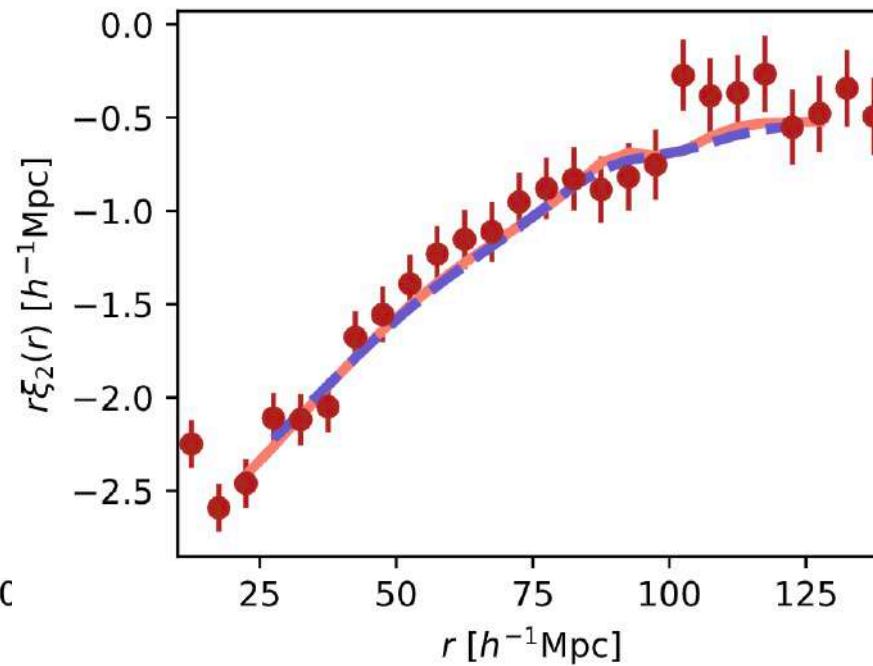
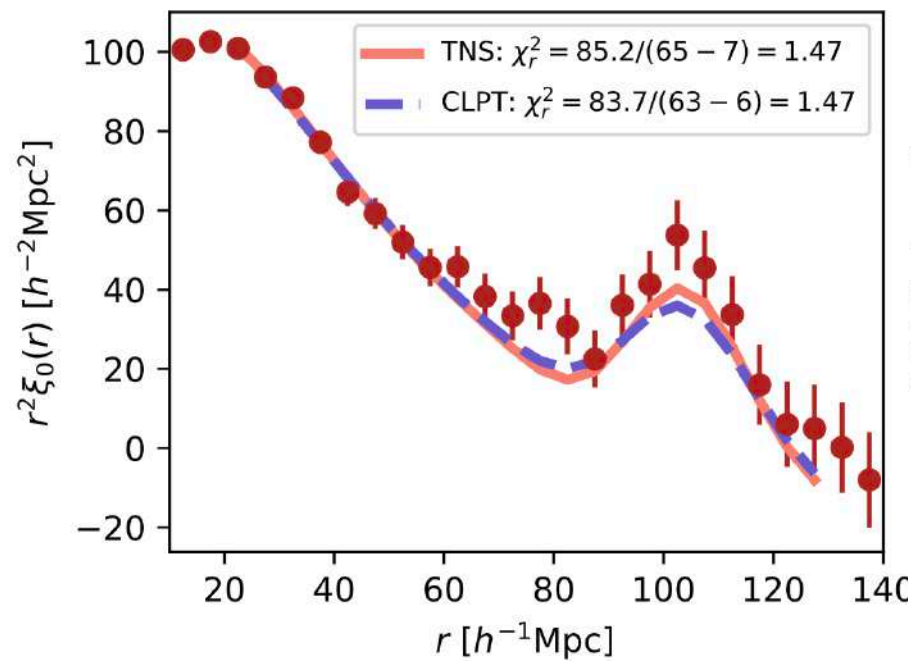
Estimate  $D_M/r_d$  and  $D_H/r_d$  from BAO peak position (and nothing else)  
Model from Bautista et al. 2018



# RSD analysis

eBOSS LRG sample

Pre-reconstruction multipoles



## Growth rate of structures

Estimate  $D_M/r_d$ ,  $D_H/r_d$  and  $f\sigma_8$  from the full-shape of the correlation function

Two models:

TNS (Taruya et al. 2010) with non-linear bias

CLPT (Carlson et al. 2013) with Gaussian streaming (Reid & White 2011)



# Final results from the LRG sample at $z = 0.7$

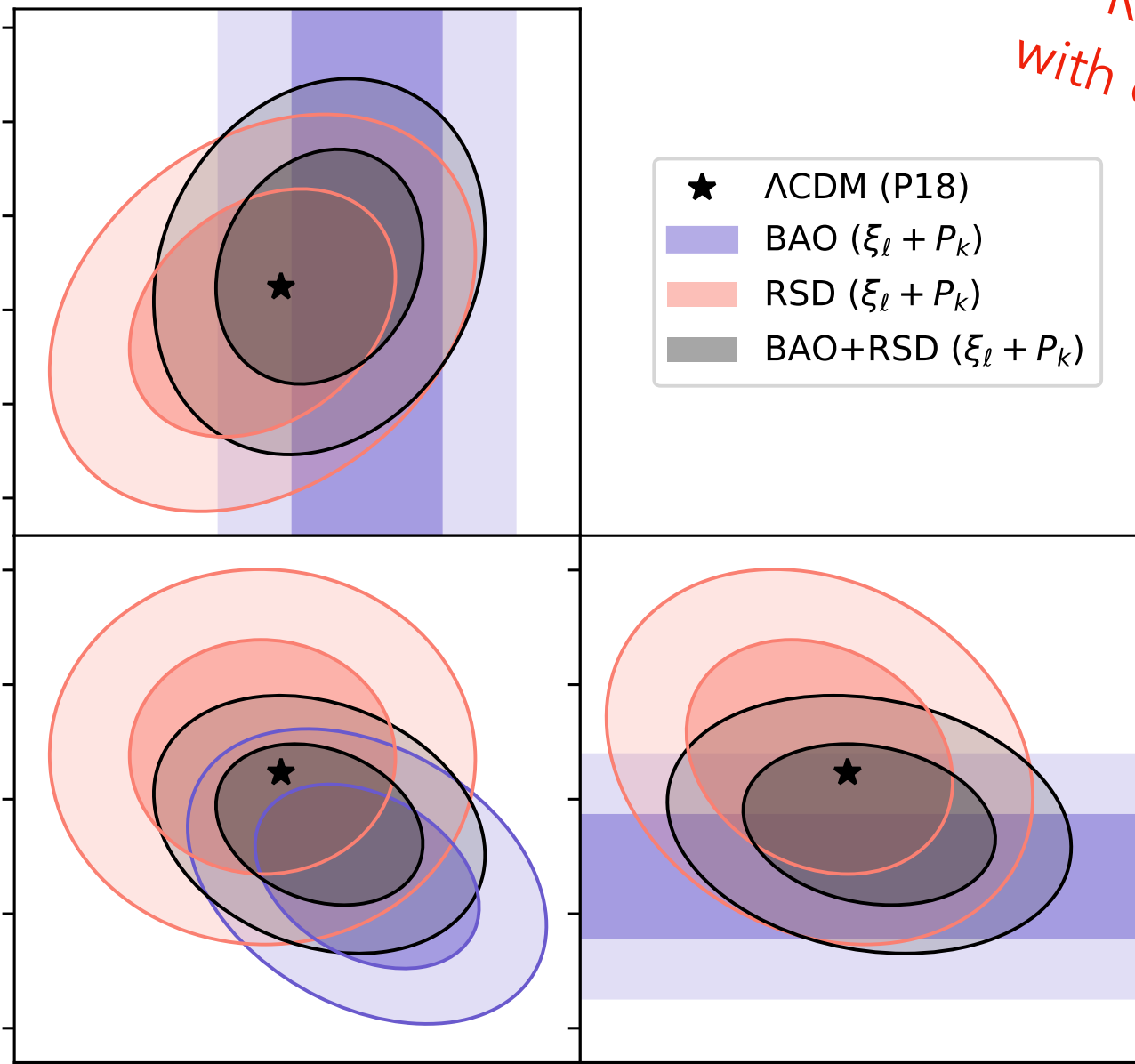
BAO + RSD  $\xi_\ell + P_\ell$

Growth rate  
of structures

$$f\sigma_8$$

Radial  
BAO scale

$$D_H(z)/r_{\text{drag}}$$



Transverse BAO scale

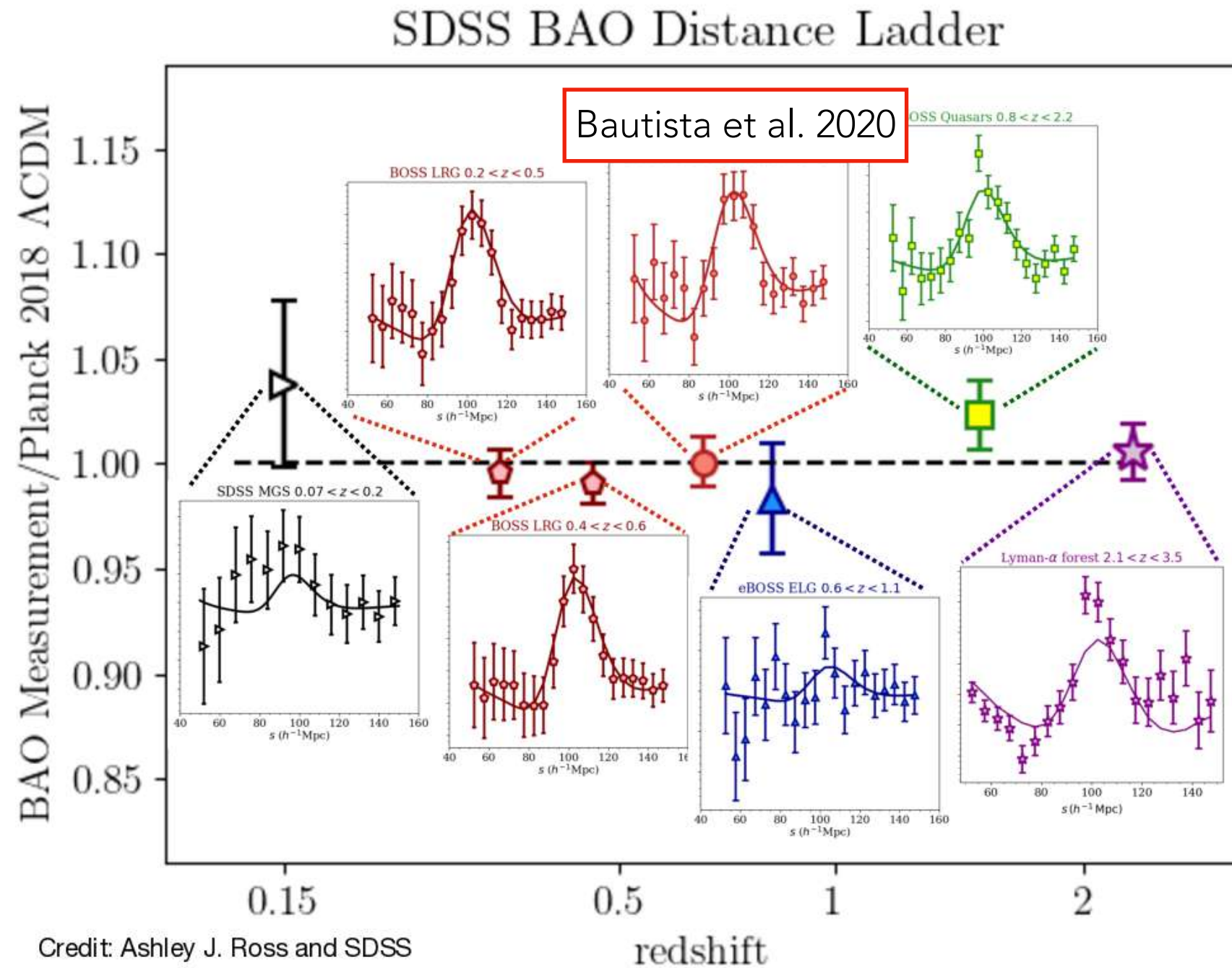
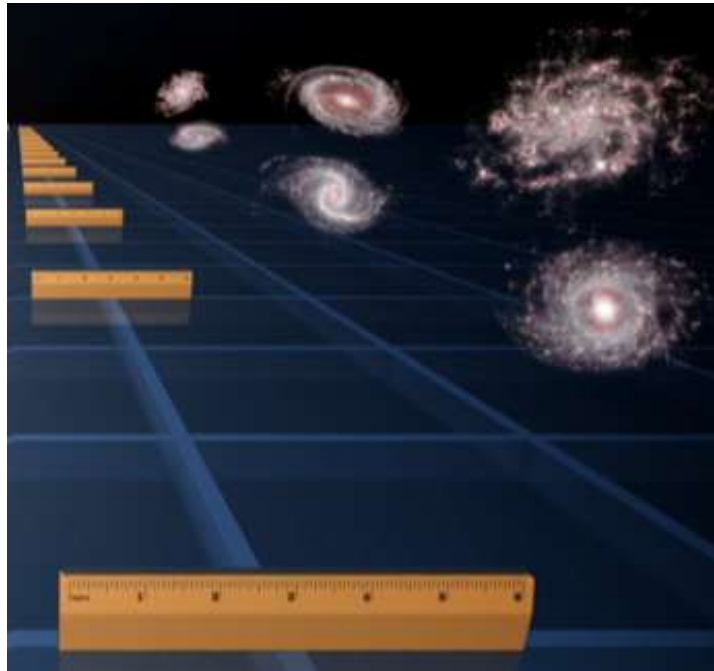
$$D_M(z)/r_{\text{drag}}$$

Growth rate  
of structures

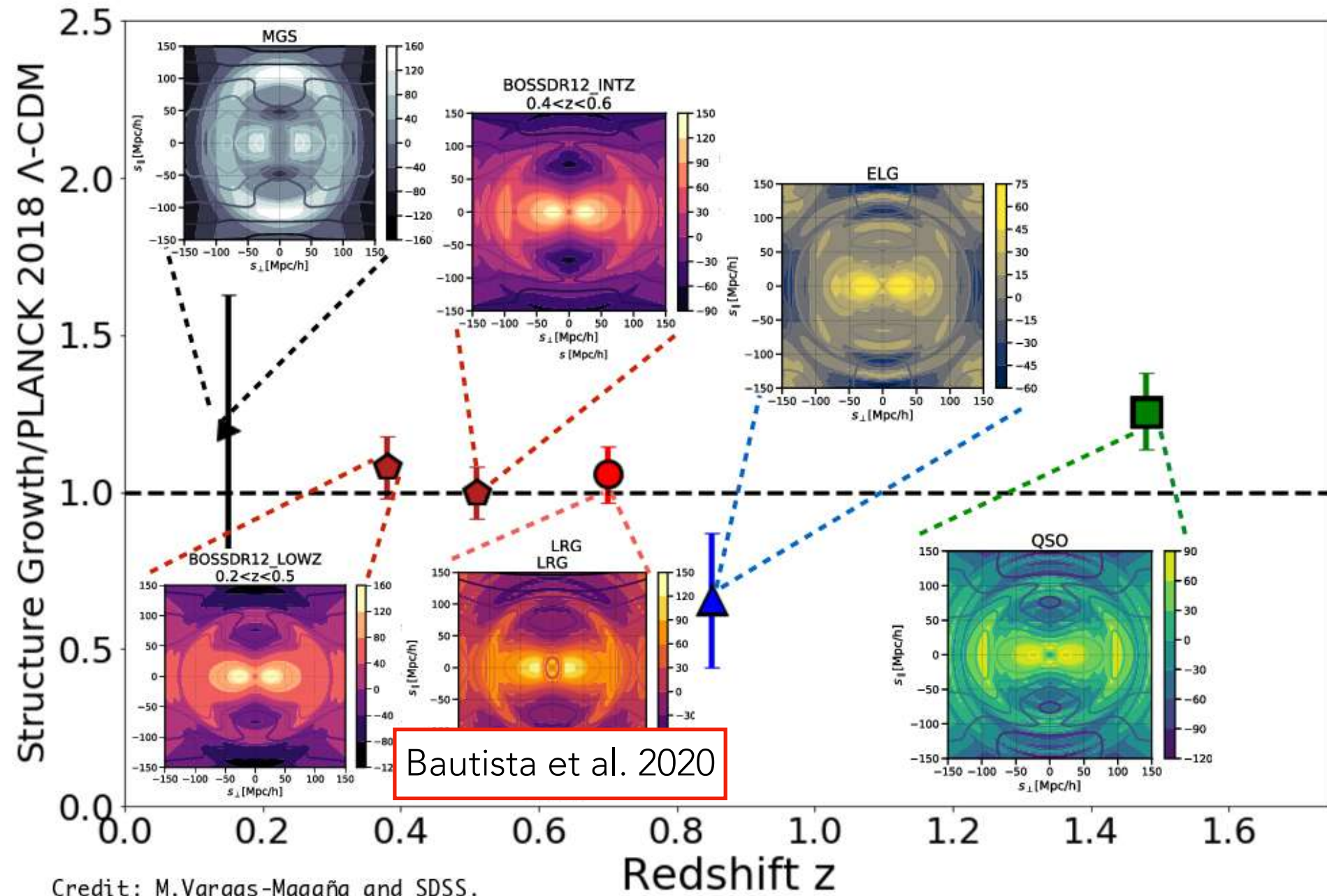
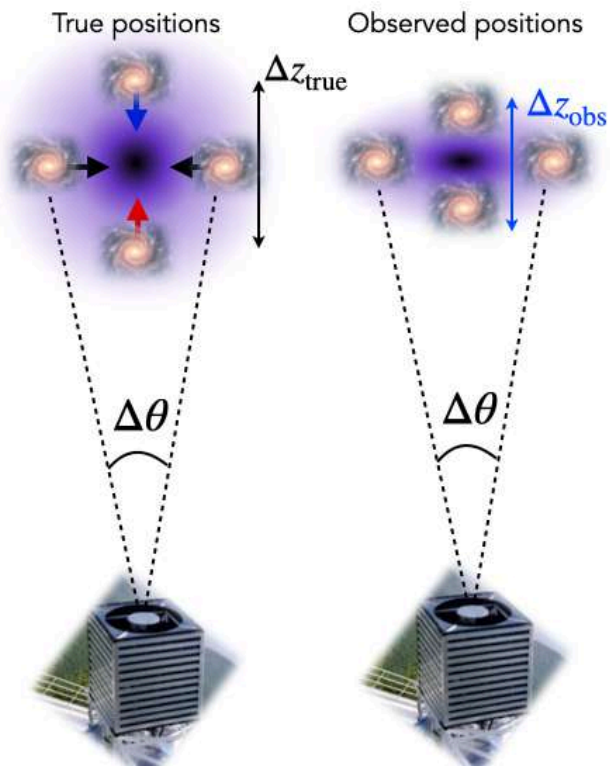
$$f\sigma_8$$

Results don't change  
with order of combination!

# Expansion-rate with Baryon Acoustic Oscillations (BAO)



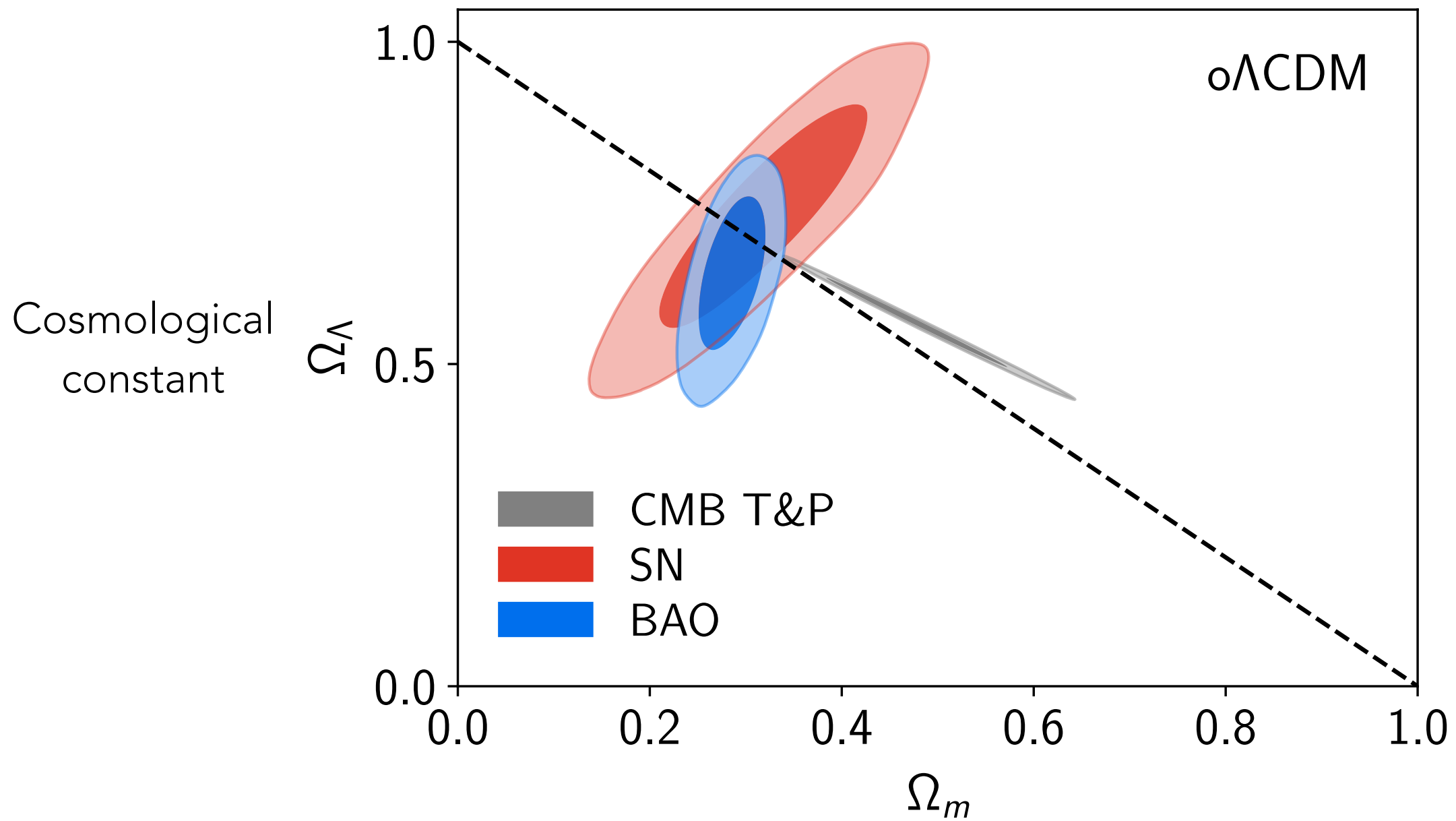
# Growth-rate of structures with redshift-space distortions (RSD)



Also using cosmic-voids: Aubert, Cousinou, Escoffier, et al. 2020

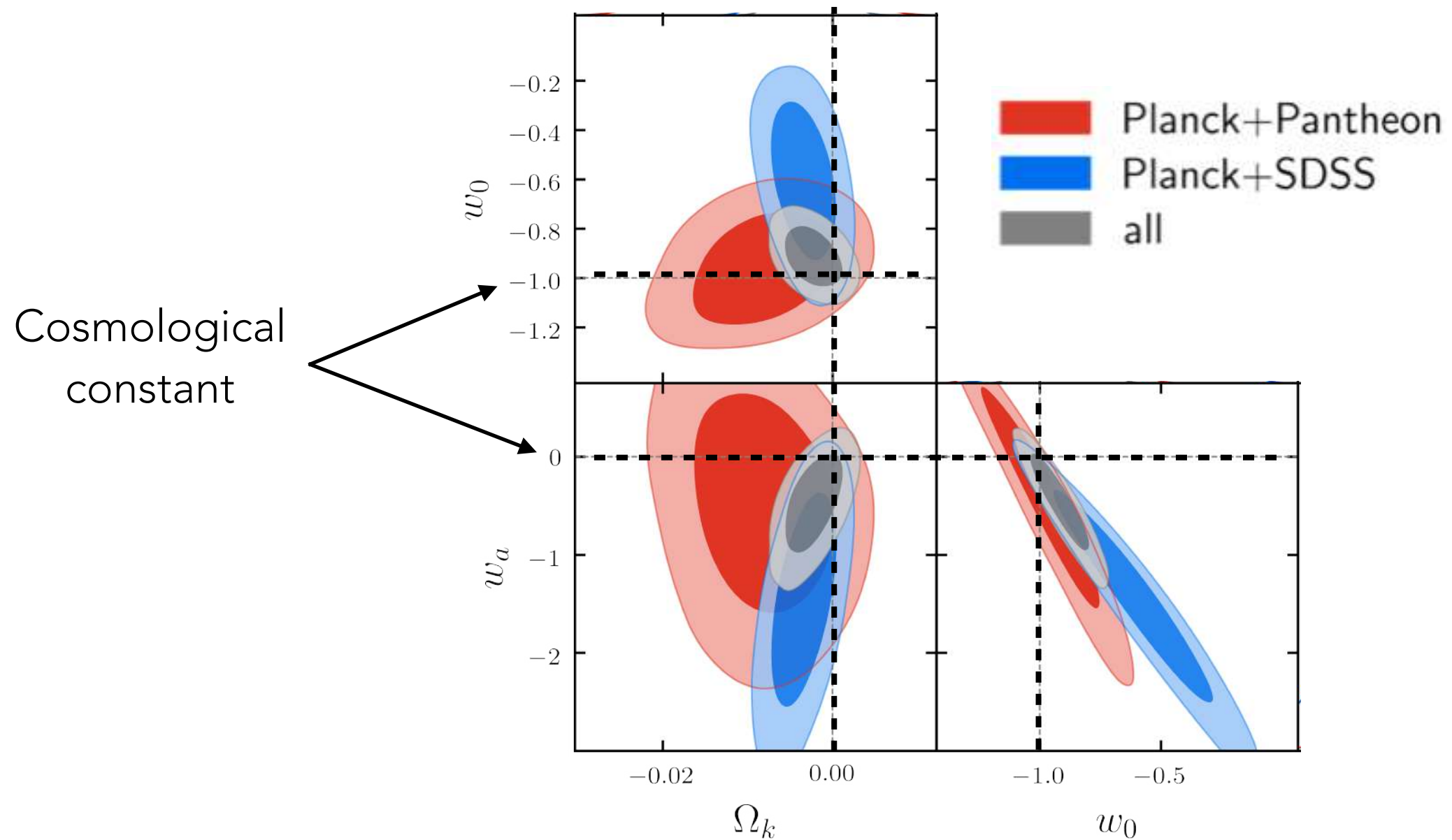
# Cosmological implications

(eBOSS Collaboration, 2020)





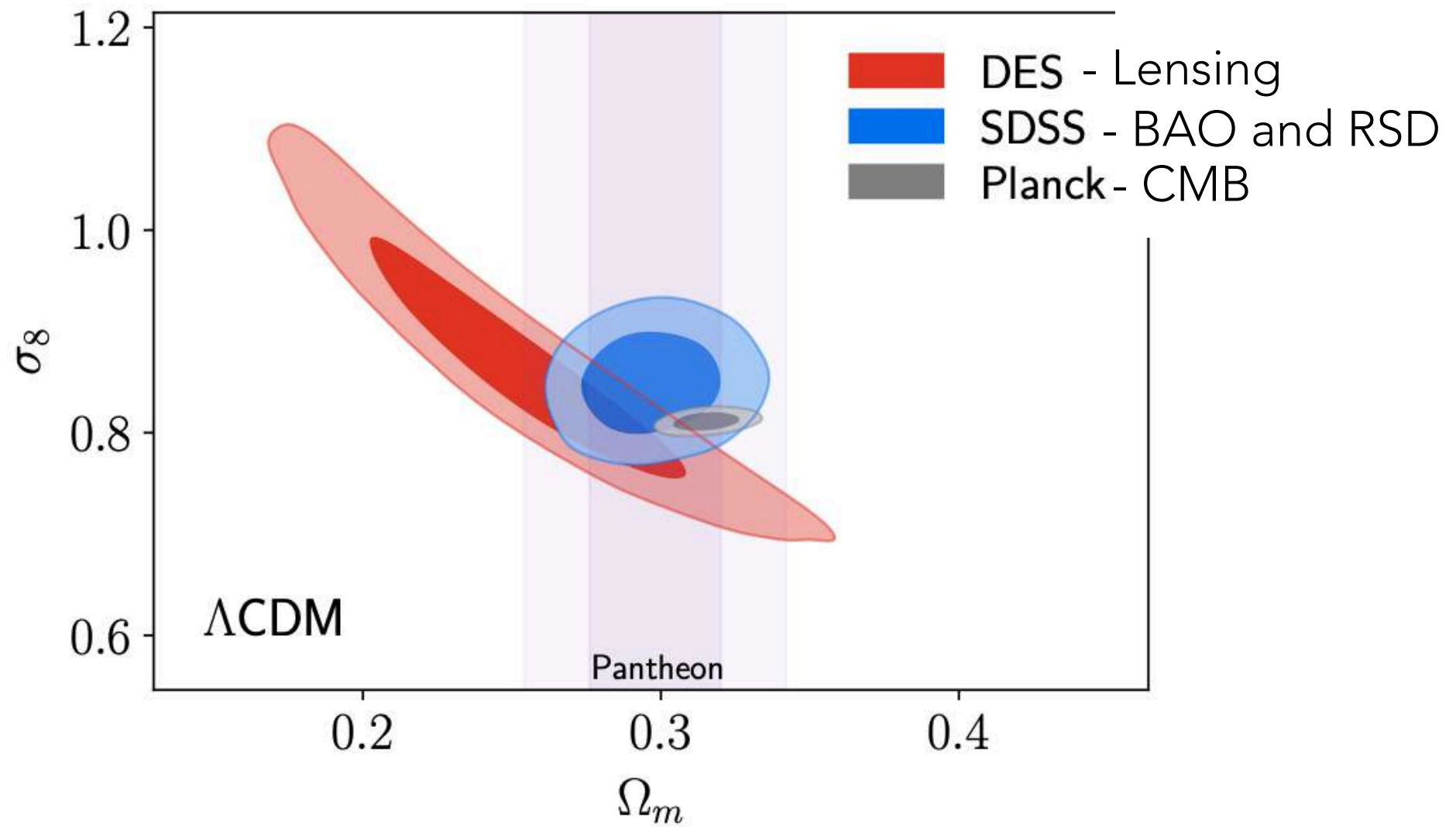
# Cosmological implications



(eBOSS Collaboration, 2020)

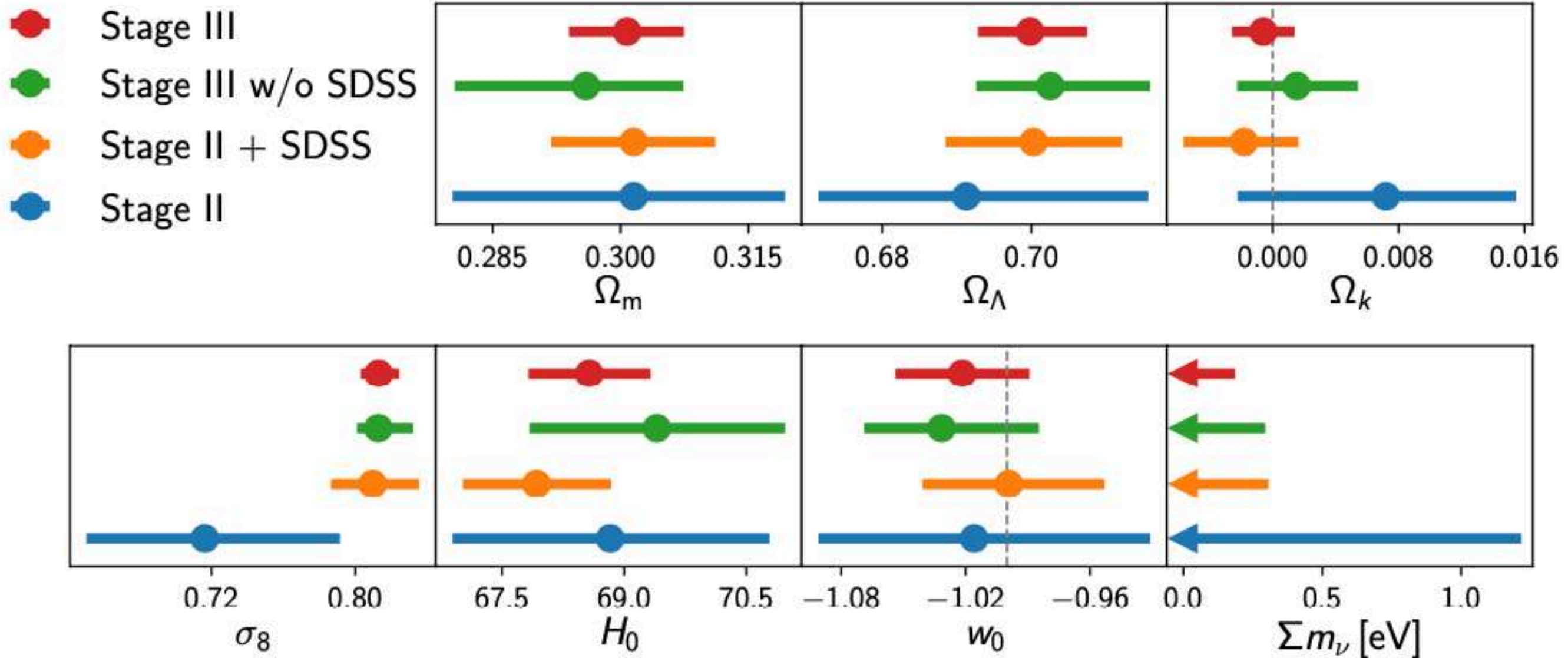
# Cosmological implications

Variance of  $\delta$   
(with 8 Mpc/h smoothing)



(eBOSS Collaboration, 2020)

# Cosmological implications



(eBOSS Collaboration, 2020)

Stage III: SDSS final, Planck CMB, Pantheon SN Ia, and DES 3x2pt  
 Stage II: SDSS DR7, WMAP CMB, JLA SN Ia



# Future

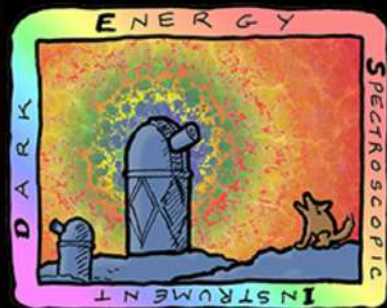
and making high-resolution maps

# Next-generation surveys of the structures

Satellite with 2m mirror  
~30 million galaxies  
2022 - 2028

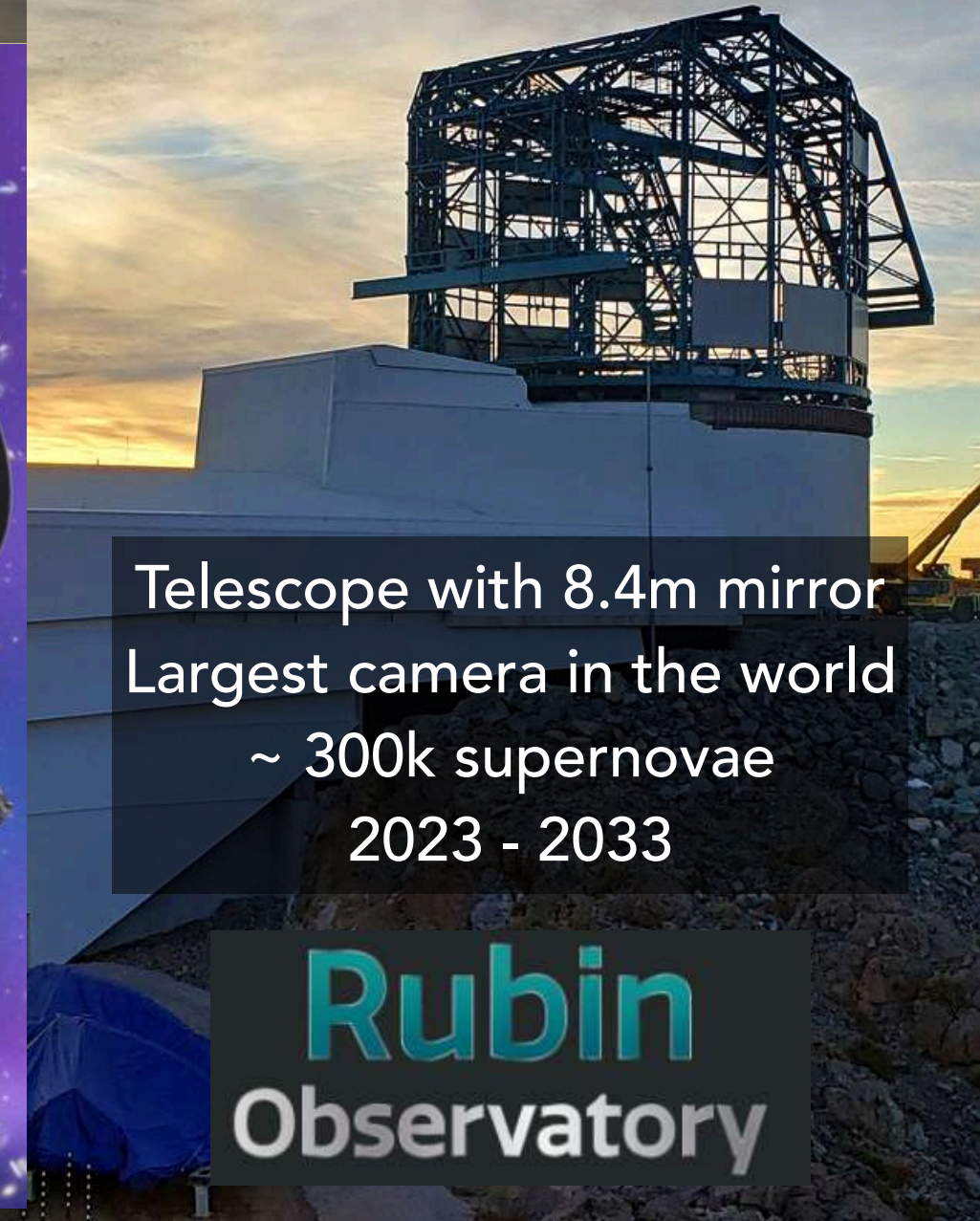


Telescope with 4m mirror  
5000 spectra at a time  
~ 20 million galaxies  
2021-2026



DARK ENERGY  
SPECTROSCOPIC  
INSTRUMENT

Telescope with 8.4m mirror  
Largest camera in the world  
~ 300k supernovae  
2023 - 2033

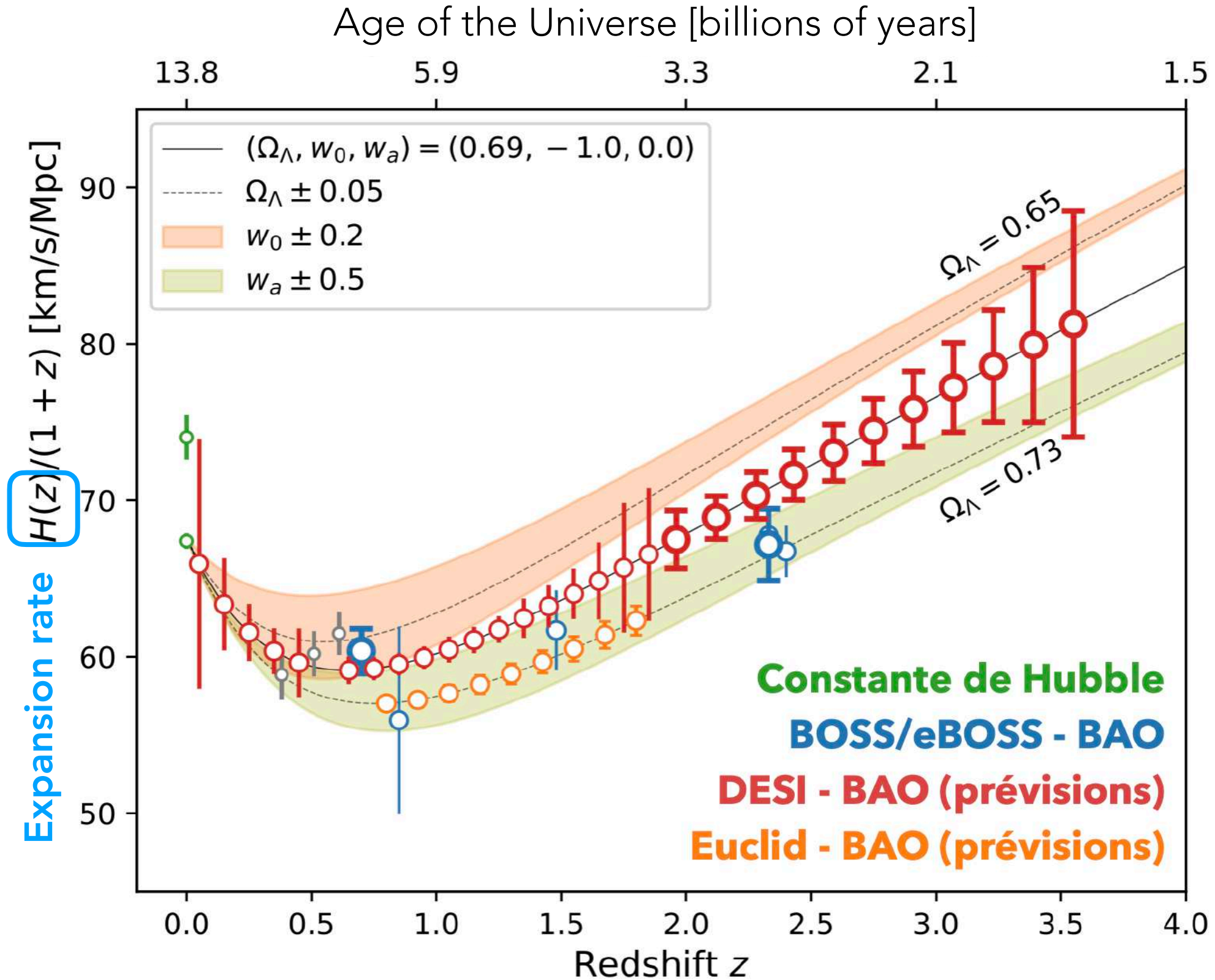


Rubin  
Observatory

Key participation of CPPM in these projects

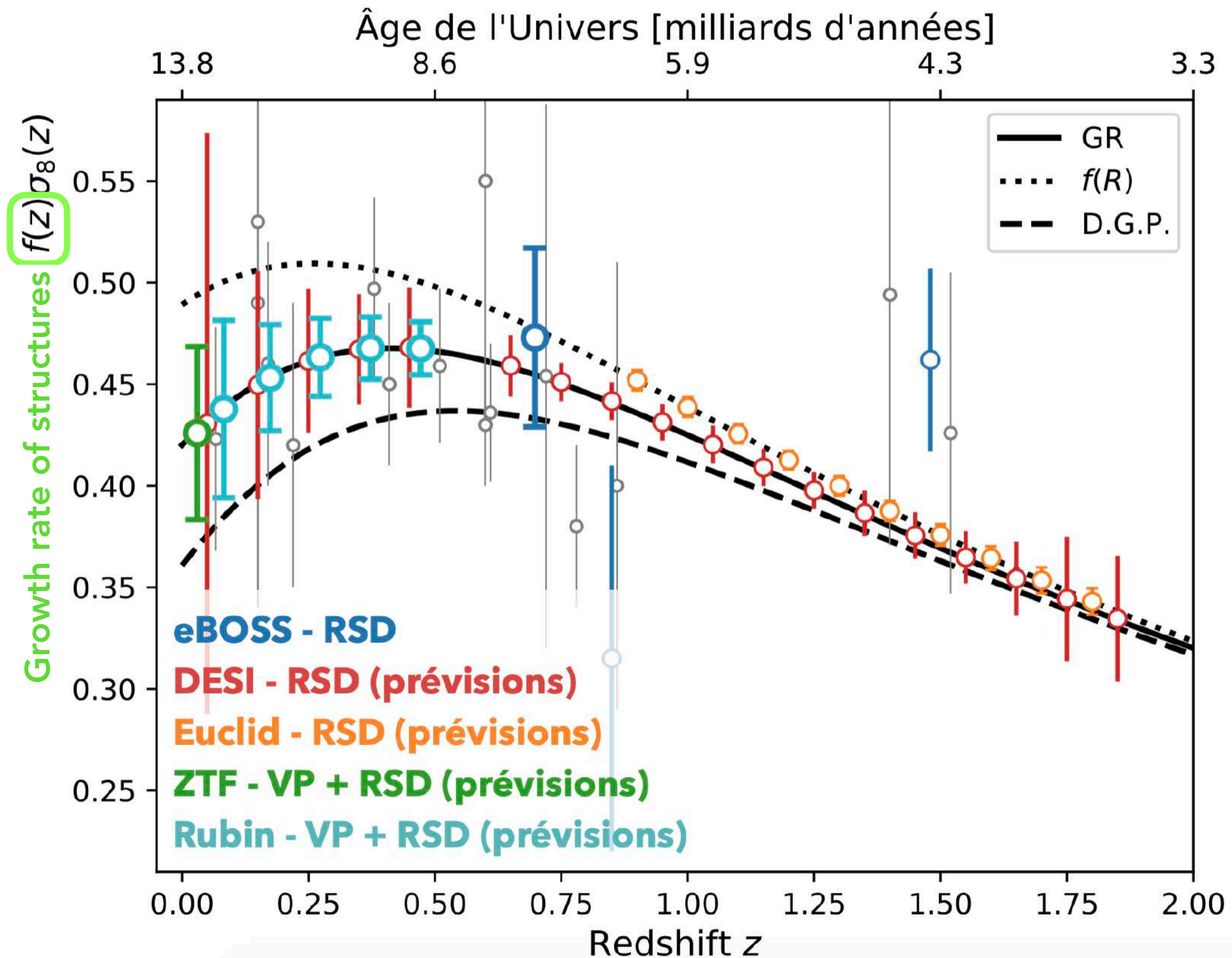


# Expansion-rate



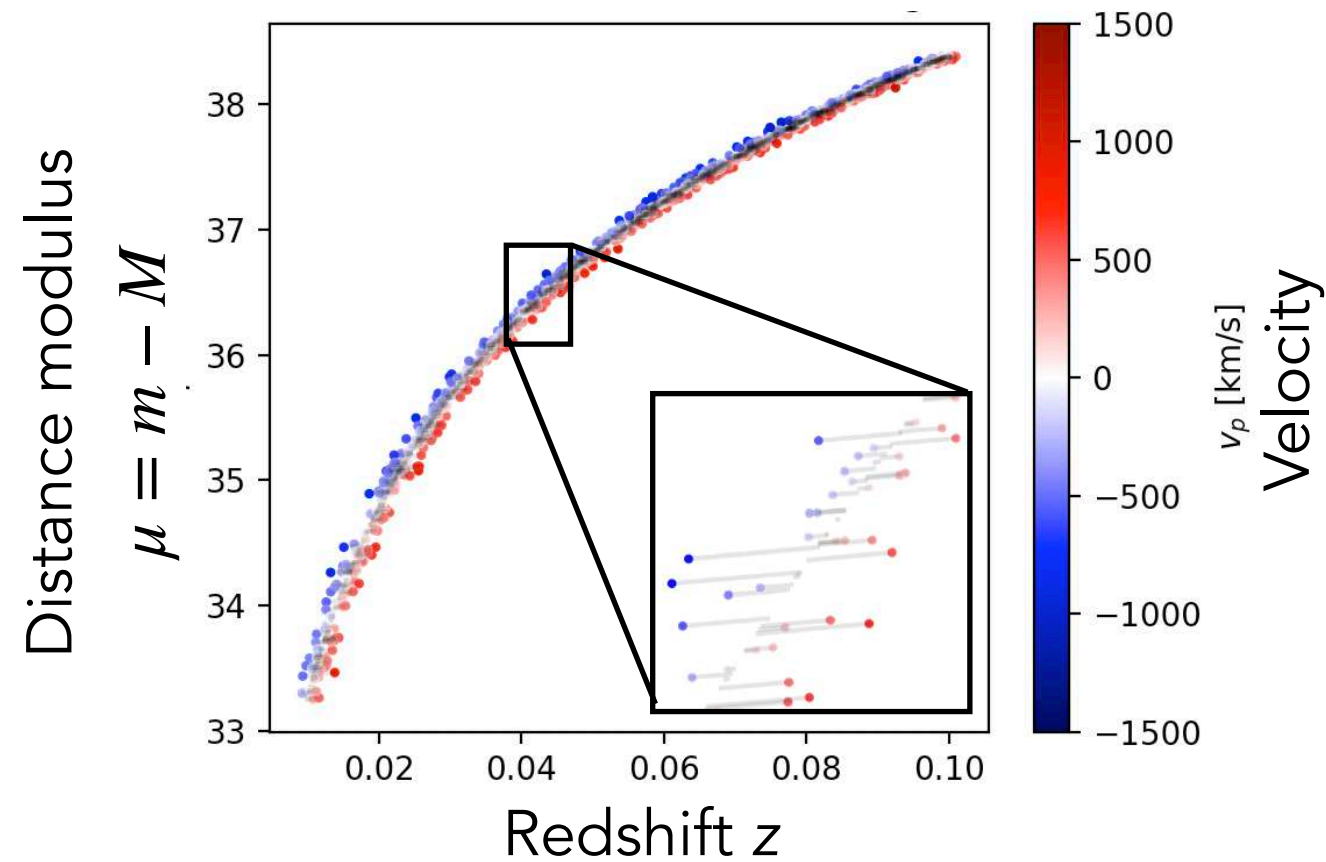


# Growth-rate

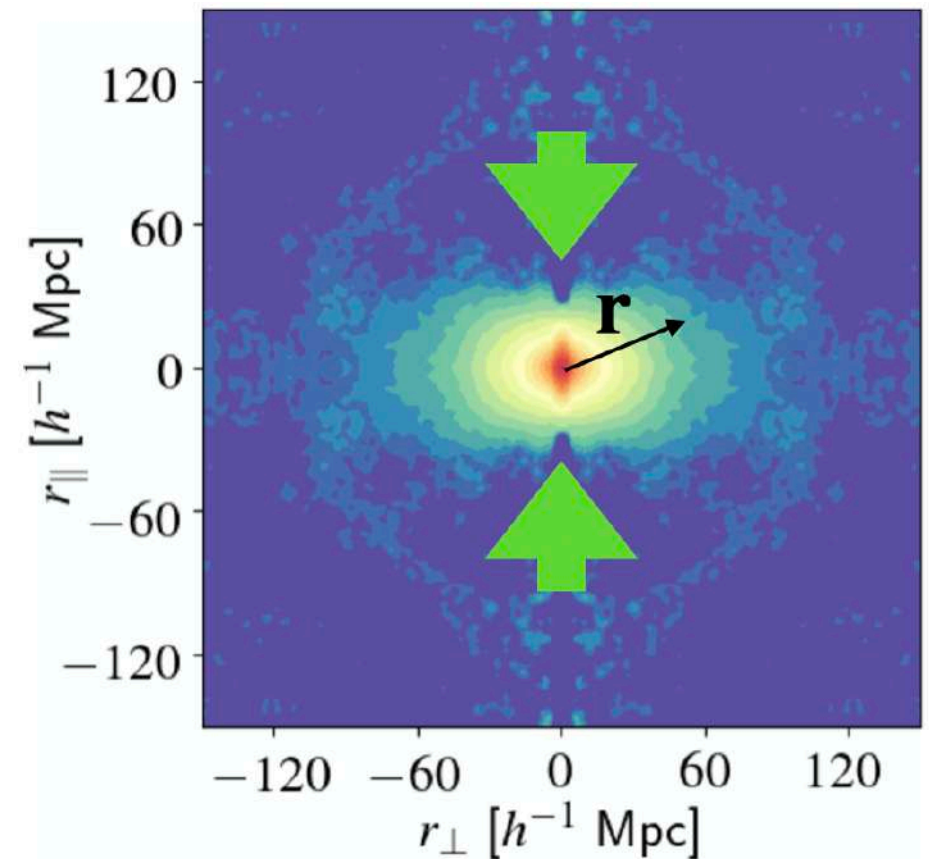


# Testing GR with type-Ia supernovae peculiar velocities

Peculiar velocities measured from SNIa



RSD of low redshift galaxies



Photometric surveys



ZTF



Rubin-LSST

Spectroscopic surveys



DESI

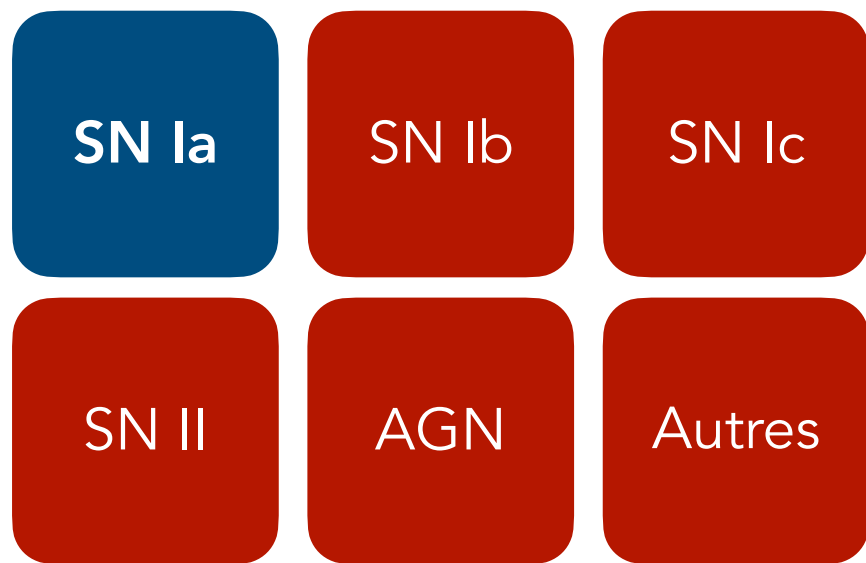


4MOST

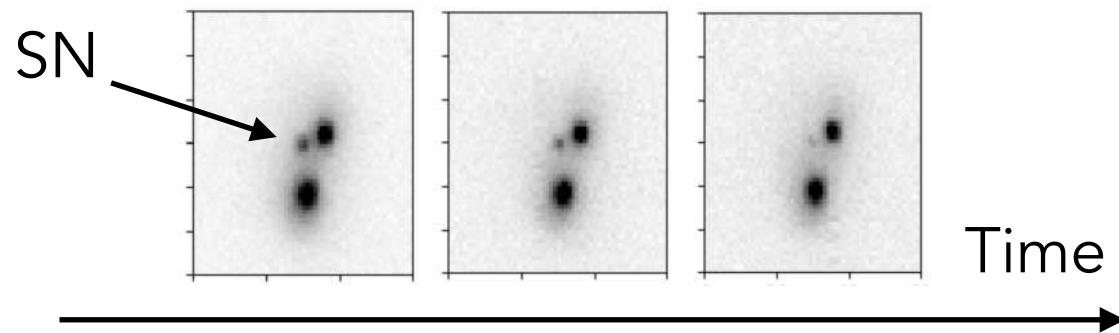
This new type of measurement will be competitive for the first time (with SNIa)

# Cosmology with type Ia supernovae

Photometric classification of supernovae with *deep learning*

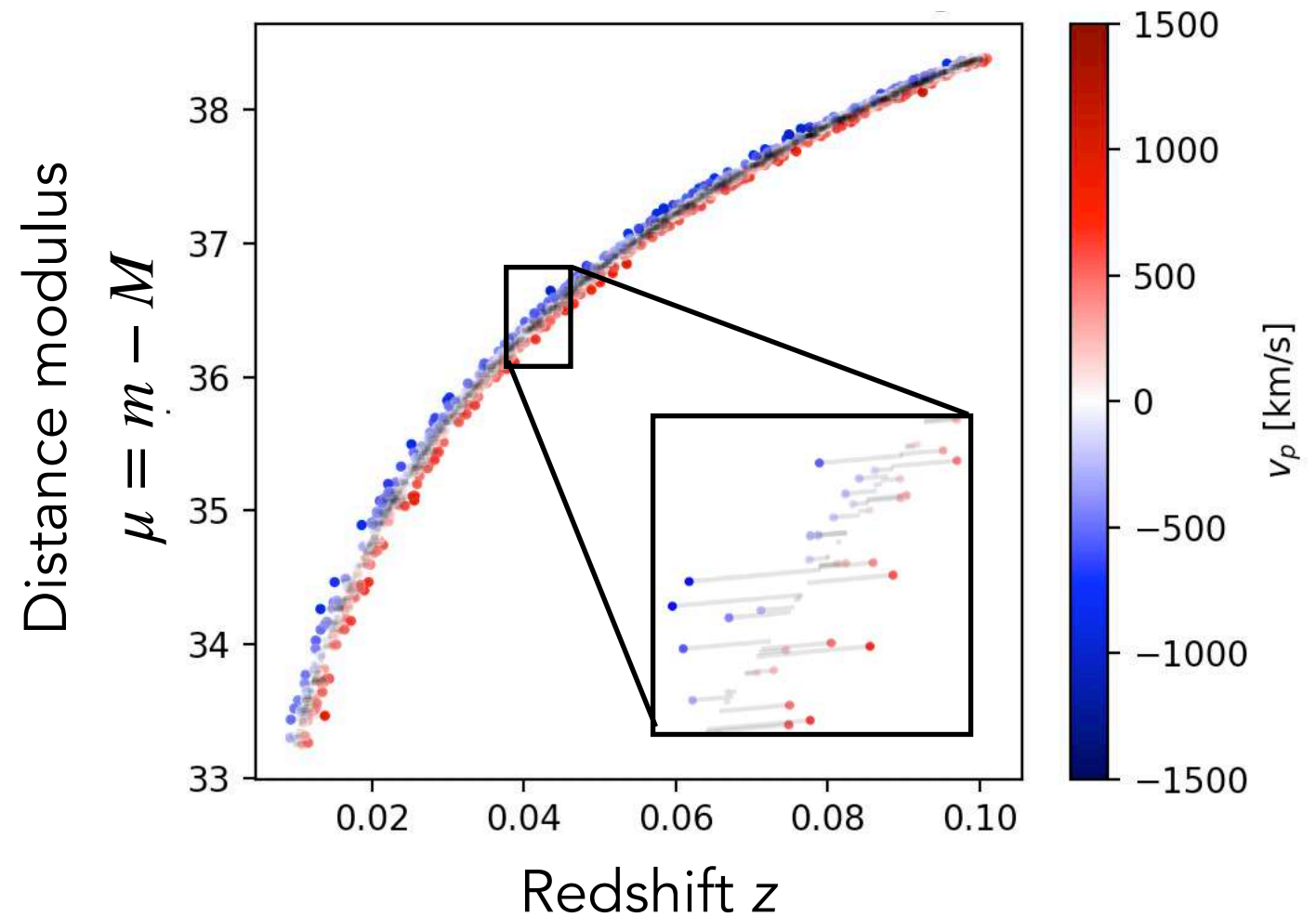


Use images as input for learning



Collaboration with LIRMM et TETIS

Peculiar velocities from SNIa



Project in collaboration with:  
Bastien Carreres, doctorant, CPPM  
Mariam Sabalbal, master 2, AMU  
Nattapon Preedasak, master 1, AMU

**Préparation à l'analyse cosmologique de la prochaine génération**



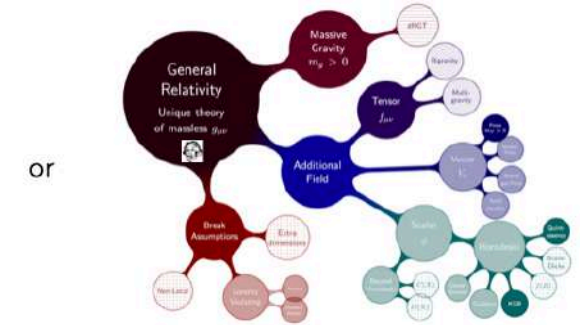
# Conclusion

What is dark energy? We don't know yet

$$\Omega_{\Lambda}[a(t)]^{-3(1+w_0+w_a)} e^{3w_a[1-a(t)]}$$

dark energy  
(quintessence, phantom force)

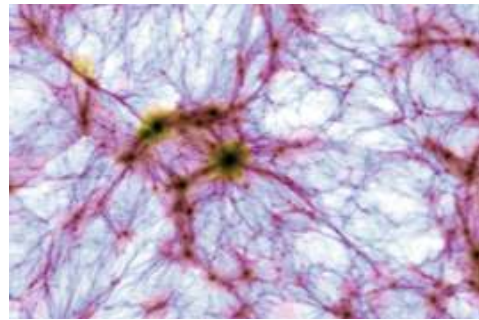
70%



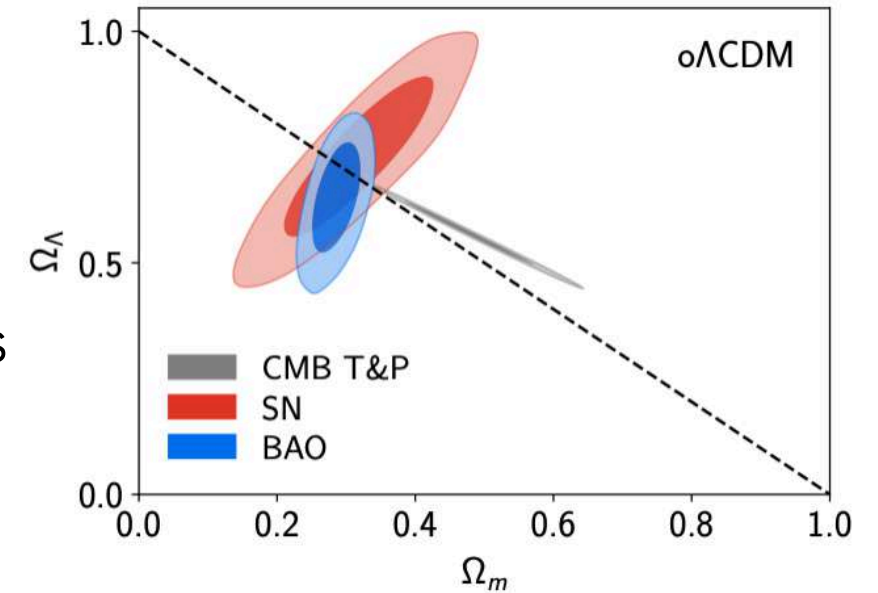
Empirical dark-energy?

Alternative to GR?

Statistics of the structures of our Universe can inform us



SDSS produced the state-of-the-art map of the structures measuring expansion-rates and growth-rates



Future is promising with next-gen surveys



euclid



Ongoing project to test GR with SNIa peculiar velocities

