



ESCAPE

European Science Cluster of Astronomy &
Particle physics ESFRI research Infrastructures

g- \mathcal{L} -like : likelihood maximisation
and profiling for astrophysical (and
beyond) applications

Cosimo Nigro, Javier Rico, IFAE

E-OSSR Onboarding Presentation

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What is gLike?

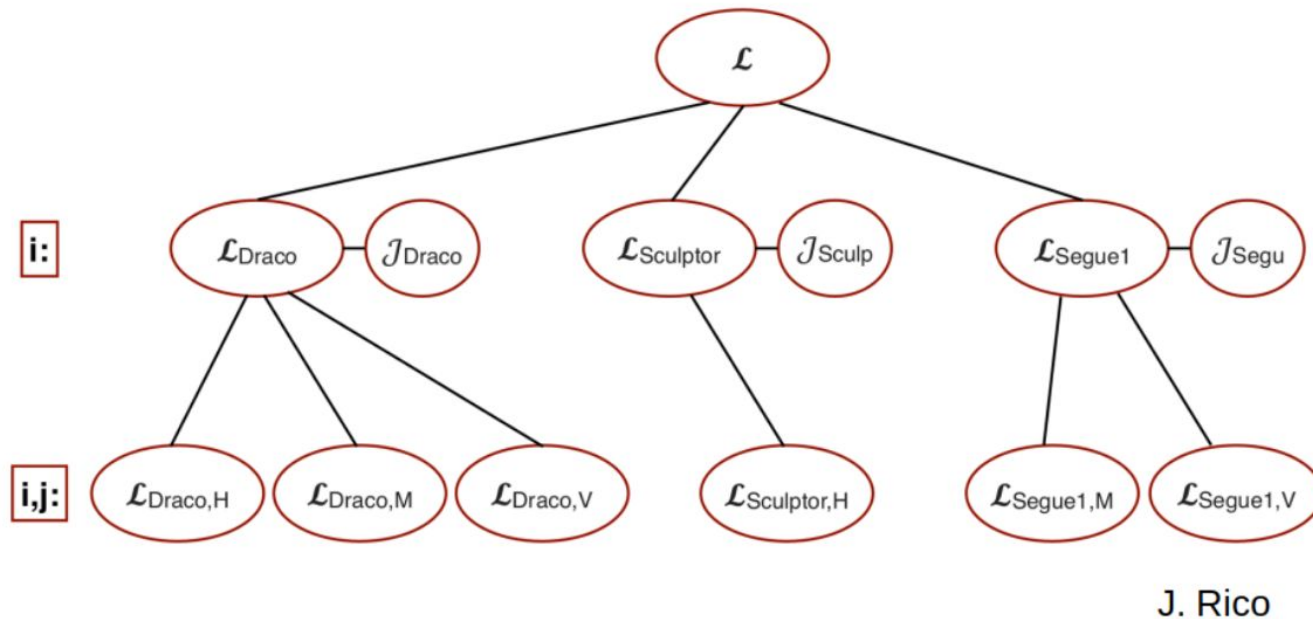
- **Framework for numerical maximisation of joint (e.g. multi-instrument) likelihood functions;**
- likelihood of **one free parameter (g)** and **as many nuisance parameters as wanted**, profiled over in the maximisation process;
- C++ code built on [ROOT](#);
- **Use cases** (in order of increasing complexity):
 - estimating the intensity of an astronomical source of signal particles in the presence of background particles from datasets obtained with different instrument and in different data formats;
 - estimating the dark matter annihilation cross-section combining observations of dwarf spheroidal gamma-ray and neutrino telescopes;
 - estimating the energy scale of quantum gravity by combining observations of recently measured GRBs (work in progress);
- already used for several publications, **highlights**:
 - [Fermi-LAT + MAGIC DM searches in dSphs](#),
 - [gloryduck project](#) (check [Tjark's presentation @Fermi Symposium](#)).



What is gLike?

Likelihood formula:

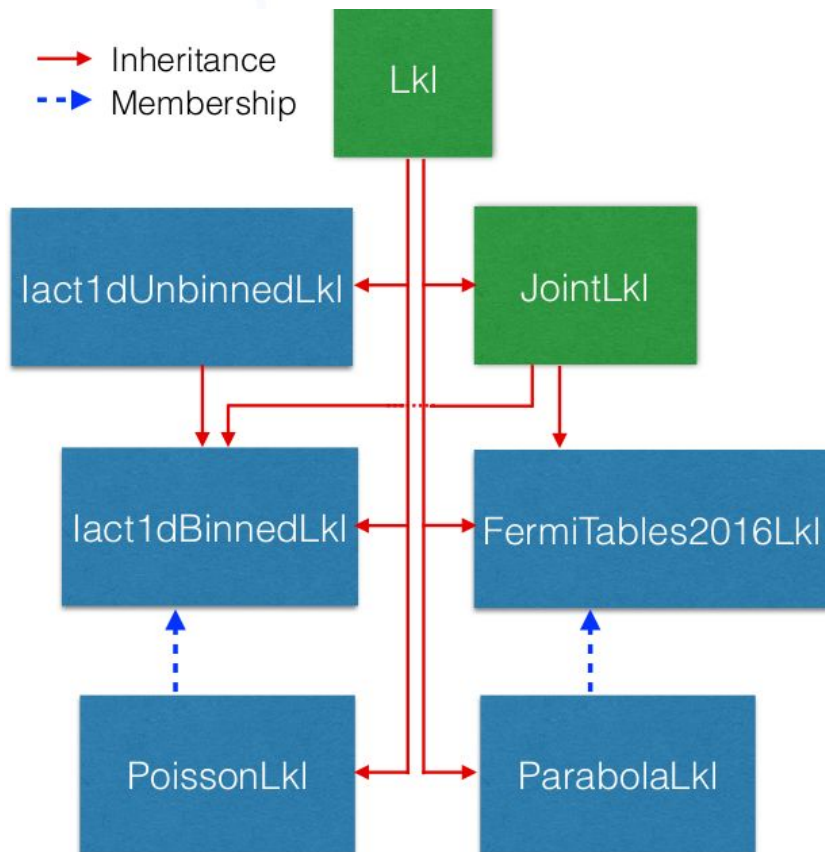
$$\mathcal{L}(\langle \sigma v \rangle; \nu \mid \mathcal{D}_{\text{dSphs}}) = \prod_{l=1}^{N_{\text{dSphs}}} \mathcal{L}_{\text{dSph},l}(\langle \sigma v \rangle; J_l, \nu_l \mid \mathcal{D}_{l,\text{measured}}) \times \mathcal{J}_l(J_l \mid J_{l,\text{obs}}, \sigma_{\log J_l})$$



- gLike can create complex likelihood functions from simpler likelihood terms;
- nuisance parameters **can selectively affect in the same way an arbitrary number of these terms.**



What is gLike?



CODE:

- **Lkl**: abstract linked to TMinuit, performing the minimisation of $-2 \log L$ vs g ;
- **JointLkl**: combines several **Lkl** objects in a joint likelihood;
- the remaining classes implement a particular likelihood function:
 - **lact1dUnbinnedLkl**: energy-dependent unbinned likelihood for energy spectra measured with IACTs,
 - **lact1dBinnedLkl**: binned version of **lact1dUnbinnedLkl**;
 - **FermiTables2016Lkl**: total L values according to Fermi tables (L vs flux for bins in Energy);
 - **PoissonLkl**: On/Off double Poisson likelihood (LiMa, Rolke,...);
 - **ParabolaLkl**: holds values of L vs g ;
 - **TemplateLkl**: template for new likelihood classes;
- **lactEventListIRF**: simple data class for IACT ON and OFF data and IRF components.

DATA:

- Data directory with gamma-ray spectra from DM annihilation and decay in different channels (from [Cirelli 2011](#)), used for $\langle \sigma v \rangle$ estimation;
- Test IACT unbinned / binned spectral data in ROOT and FITS format.



Software Development

- **Software development on github:**
<https://github.com/javierrico/gLike>
- **Documentation:**
 - github wiki;
 - .html documentation can be generated with one of the scripts or at compilation with cmake;
- **License:** GNU General Public License v3.0.
- **Test and CI / CD:**
 - only a small test unit available at the moment, not integrated in a workflow;
 - **no workflow for CI / CD set up at the moment.**



Software Development

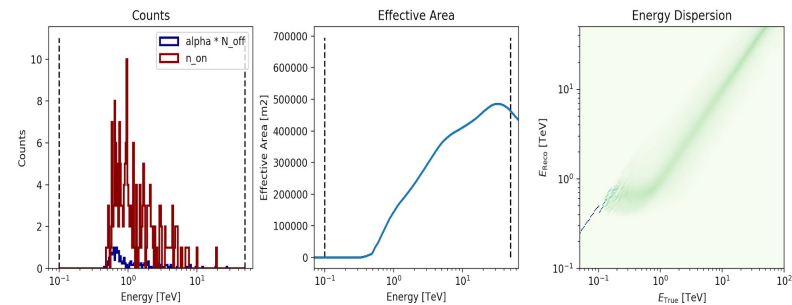
- **Operating System, compilation environment:**
 - ROOT (sole dependency);
 - cmake build;
- **hardware requirements:**
 - no particular hardware requirement;
- **containerisation and portability requirements:**
 - containers available for two root versions;
- **workflow/interface requirements to other software/services:**
 - compatibility with gammapy, interface at the data level.



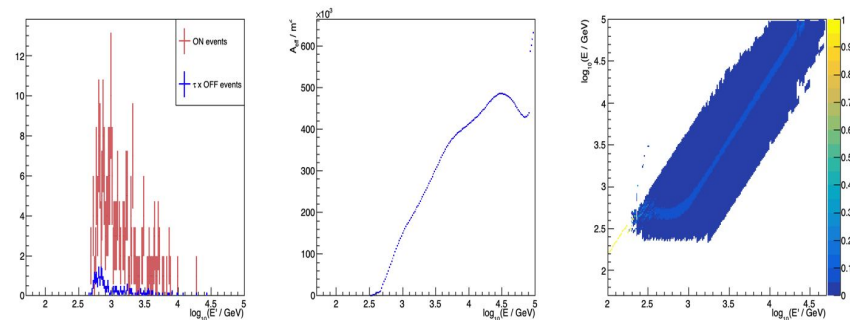
Software Development

- Decided to make gLike **interfaceable with gammapy at the data level**;
- user reduces data with gammapy, **DL3 -> unbinned spectral data (list of ON - OFF events + IRF)**;
- the **unbinned spectral data** are then fed to gLike for likelihood minimisation / profiling;
- script to produce unbinned spectral data with gammapy in [gammapy-extra](#);
- added interface to FITS data in gLike (via ROOT);
- to read the OGIP format (in which the spectral data are stored) a patch in ROOT was needed. The ROOT version including this patch is > v6.20;
- we maintained compatibility with old ROOT versions (used by most of the analysers).

Gammapy output (from data reduction)



gLike input (to likelihood analysis)



OSSR Integration

- **What is available?**
 - source code;
 - singularity containers, two root versions:
 - v5-34-36, used by analysers,
 - v6-22-06, ROOT patched to read IRF in .fits data.
- **What will be onboarded (source code, container, test workflow incl. data)?**
 - source code;
 - containers;
- **Are there open points and requirements?**
 - expand the test unit (already working on this);
 - include the tests and the container generation in a git automatic workflow;
 - add interface to binned spectral data ([OGIP](#)) obtainable with gammapy - and produced by many high-energy instruments (X-ray telescopes...).



OSSR Integration

- **What is the “user story” of a EOSC user taking on the software?**
- a user can reduce gamma-ray data with gammapy and analyse them with gLike;
- a user can combine **data from different instruments** (gamma-ray and neutrino observatories) **and in different formats** (ROOT and FITS) in the same likelihood framework with gLike;
- it will be extremely relevant for the **Dark Matter Test Science Project:**
 - flexible likelihood template can be used to add likelihood terms for particle physics data.



Time for a short demo (~ 10 min)

