



UNIVERSITÉ DE NANTES



GDR Groupement
de recherche
QCD Chromodynamique quantique

Study of 2nd order susceptibilities with
EPOS event generator

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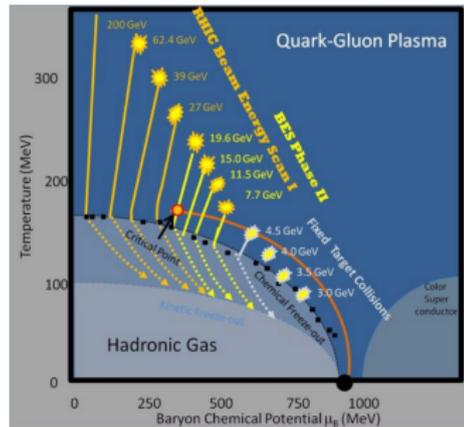
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QCD phase diagram and CEP

Since the QGP has been observed (indirectly), efforts has been made to learn about its properties, and to map the QCD phase diagram.

- **Theoretically** : use models & theories to make predictions (T_C , μ_{B_C}) or to extract information from measurements (T , μ_B , $\eta/s...$)
- **Experimentally** : exploration of QCD phase diagram thanks to the Beam Energy Scan (BES) program, measurements of observables of interest



D. Cebra, 2013

Question(s) of interest : is there a 1st order phase transition and a critical endpoint between QGP and hadronic gas phases ? If yes, **where** ?

Susceptibilities

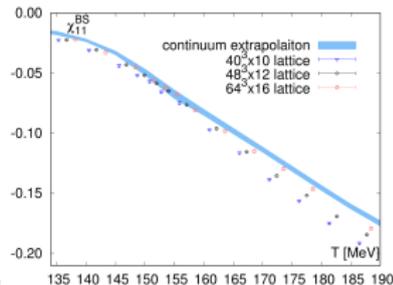
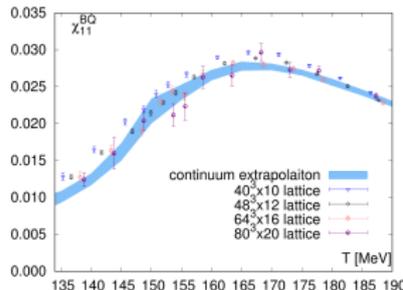
To answer this question, there are many tools that can be used, among which are the **susceptibilities**.

In a grand-canonical ensemble, to what a heavy-ion collision can be compared to, they are defined as derivatives of the partition function Z :

$$\chi_{i,j}^{X,Y} = \frac{1}{VT^3} \cdot \left[\frac{\partial^{i+j} Z(T, V, \mu)}{(\partial \hat{\mu}_X)^i (\partial \hat{\mu}_Y)^j} \right]_{\mu_{X,Y}=0}$$

$$(\hat{\mu} = \frac{\mu}{T})$$

As we are searching for **radical changes in the state of nuclear matter**, i.e. phase transition, these derivatives of Z should reveal them.



R. Bellwied et al., 2019

Susceptibilities

In a more convenient and understandable way, susceptibilities can be written as a function of the **net-charge cumulants** ($N_X = n_X - n_{\bar{X}}$).

They represent in fact **event-by-event fluctuations** of the considered net charges.

Also, in order to have observables **independent from volume or temperature**, which cannot be measured directly in experiments, **ratios** are often used.

2nd order susceptibilities for B, Q, S

Linked to the (co)variances of the considered charges :

$$\chi_{11}^{XY} = \frac{1}{VT^3} \sigma_{XY}^{11} = \frac{\langle N_X N_Y \rangle - \langle N_X \rangle \langle N_Y \rangle}{VT^3}$$

$$\chi_2^X = \frac{1}{VT^3} \sigma_X^2 = \frac{\langle N_X^2 \rangle - \langle N_X \rangle^2}{VT^3}$$

Ratios

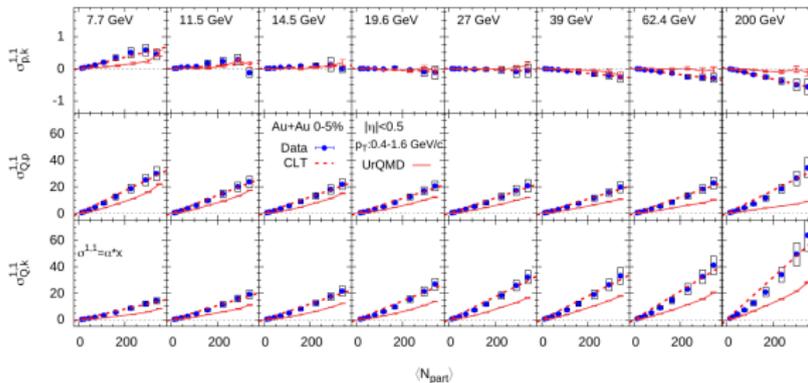
$$C_{BS} = \frac{\sigma_{BS}^{11}}{\sigma_S^2} \quad C_{QB} = \frac{\sigma_{QB}^{11}}{\sigma_B^2} \quad C_{QS} = \frac{\sigma_{QS}^{11}}{\sigma_S^2}$$

What has been done recently ?

Experimental results

STAR collaboration measured (for particles with $|\eta| < 0.5 + 0.4 < p_T < 1.6 \text{ GeV}/c$) :

- (co)variances $\sigma_{p,Q,k}^{11,2}$ (proxies for $\chi_{11,2}^{B,Q,S}$)
 - as a function of the η -window
 - as a function of $\langle N_{part} \rangle$

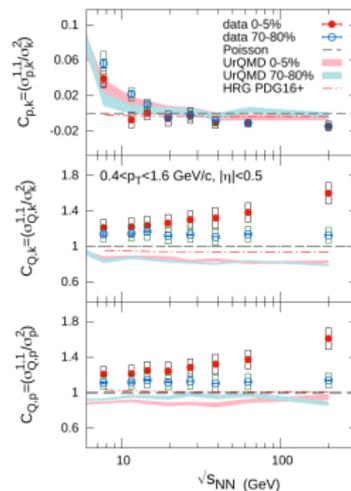
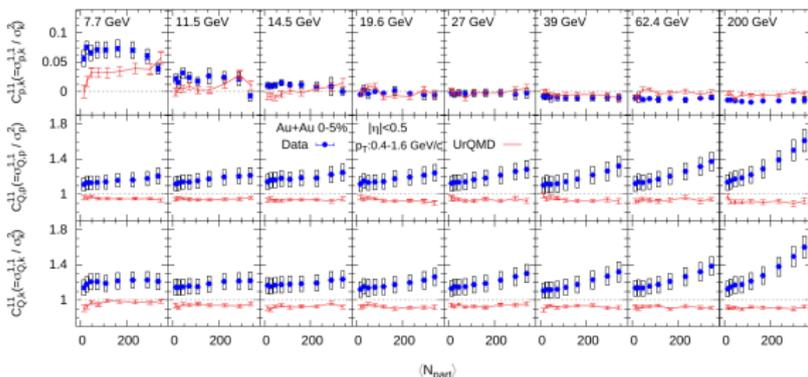


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STAR collaboration measured (for particles with $|\eta| < 0.5 + 0.4 < p_T < 1.6 \text{ GeV}/c$) :

- (co)variances $\sigma_{p,Q,k}^{11,2}$ (proxies for $\chi^{B,Q,S}_{11,2}$)
 - as a function of the η -window
 - as a function of $\langle N_{part} \rangle$
- Koch ratios $C_{Qp,Qk,pk}$ (proxies for $C_{QB,QS,BS}$)
 - as a function of $\langle N_{part} \rangle$
 - as a function of $\sqrt{S_{NN}}$



What has been done recently ?

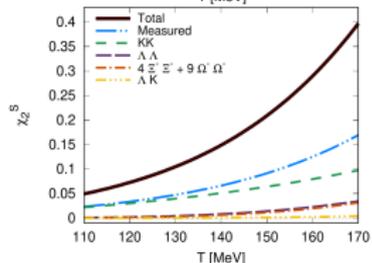
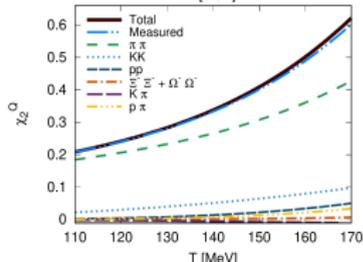
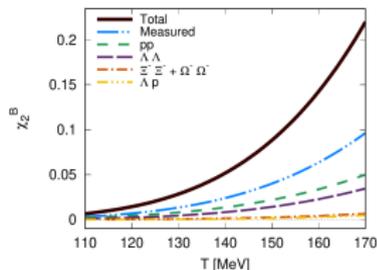
Lattice QCD + HRG model

R. Bellwied *et al.* :

- breakdown of hadronic species contributions to susceptibilities, using Hadron Resonance Gas model
 \Rightarrow **best proxies for ratios**
 (so potentially the most sensitive ones)

$$C_{QS} = \frac{\chi_{11}^{QS}}{\chi_2^S} = \frac{1}{2} \cdot \frac{\sigma_K^2}{\sigma_K^2 + \sigma_\Lambda^2}$$

$$C_{BS} = \frac{\chi_{11}^{BS}}{\chi_2^S} = \frac{\sigma_\Lambda^2}{\sigma_K^2 + \sigma_\Lambda^2}$$



What has been done recently ?

Lattice QCD + HRG model

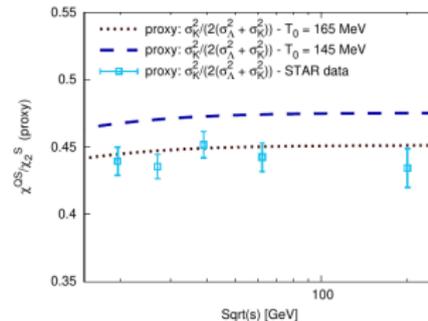
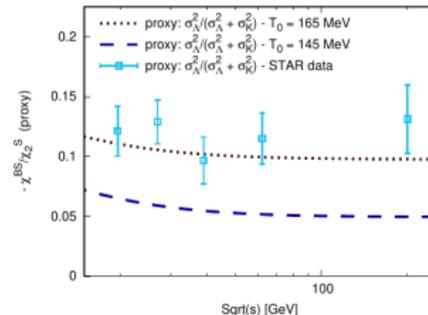
R. Bellwied *et al.* :

- breakdown of hadronic species contributions to susceptibilities, using Hadron Resonance Gas model
- ⇒ **best proxies for ratios**
(so potentially the most sensitive ones)
- ⇒ **results depending on \sqrt{s} + kinematic cuts compared with STAR data**

$$C_{QS} = \frac{\chi_{11}^{QS}}{\chi_2^S} = \frac{1}{2} \cdot \frac{\sigma_K^2}{\sigma_K^2 + \sigma_\Lambda^2}$$

$$C_{BS} = \frac{\chi_{11}^{BS}}{\chi_2^S} = \frac{\sigma_\Lambda^2}{\sigma_K^2 + \sigma_\Lambda^2}$$

... and what about event generators ?



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EPOS

Event generators are programs made to compute models and simulate every steps of a collision (e.g. [EPOS](#), [PYTHIA](#), [HIJING++...](#)).

Advantage : perfect detector, as final-state particles are all listed (no uncertainties)

Energy conserving quantum mechanical approach, based on

Partons, parton ladders, strings,

Off-shell remnants, and

Saturation of parton ladders

Event generator based on [Parton-Based Gribov-Regge Theory](#), which unifies [Parton model](#) and [Gribov-Regge theory](#) by [solving inconsistencies](#) of both models.

Can simulate any type of collision :

$$e^{+/-} + e^{+/-}$$

$$e^{+/-} + p$$

$$p + p$$

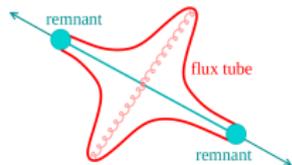
$$p + A$$

$$A + A$$

Initial conditions

Primary interactions treated with PBGRT exchange of multiple Pomerons in parallel

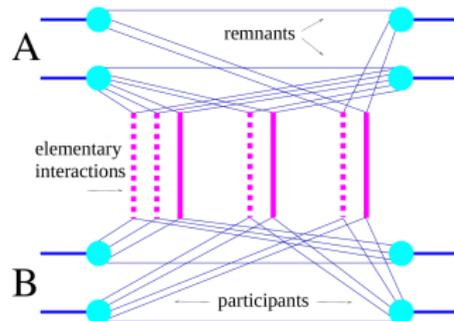
→ can be seen as cut (particle production) or
uncut (σ calculation) parton ladders



$\times N_{int} =$



K. Werner, 2018



K. Werner et al., 2000

Core-corona separation

- **Core** = high string density region
- **Corona** = escaping segments (with high p_T)

Medium evolution, hadronisation and re-scattering

Core evolution

Viscous 3D+1 hydrodynamics expansion
based on a cross-over transition EoS

+

Hadronisation of the medium via
Cooper-Frye procedure

Corona evolution

Strings evolution following dynamics of
gauge invariant Lagrangian

+

String fragmentation to produce hadrons



Re-scatterings between formed hadrons with **UrQMD model** until
chemical freeze-out (no more inelastic scatterings)
kinetic freeze-out (no more elastic scatterings)



Final state particle

What we can (not) study with EPOS

The **hydrodynamic** evolution of the core in EPOS **does not include fluctuations** :
susceptibilities are **NOT sensitive** to any possible **CEP**

⇒ search for signatures of CEP **impossible with EPOS** by construction !

In fact, in EPOS, all the **fluctuations** are coming from **initial conditions**,
hadronisation process and/or **hadronic cascades**.

Hence, what we plan to do is to

study the impact of hadronisation and hadronic cascades on the susceptibilities

by comparing susceptibilities :

- from **micro-canonical** & **grand-canonical** decays
- from **before** & **after** UrQMD

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Koch ratios vs energy

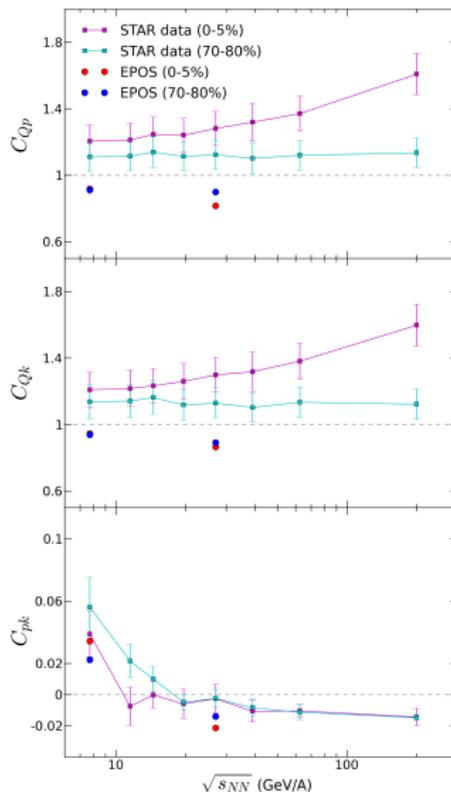
Results from **EPOS 4** with **BEST EoS**, with
 $T_C = 155$ MeV and $\mu_{B_C} = 350$ MeV
 (cf. recent *presentation from M. Stefaniak*
for more details)

!!! CAUTION : β version of EPOS 4 !!!
 → just started parametrisation for
 LHC data

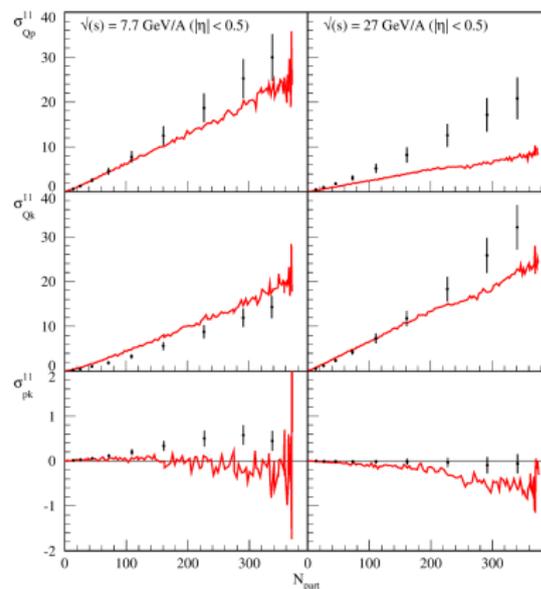
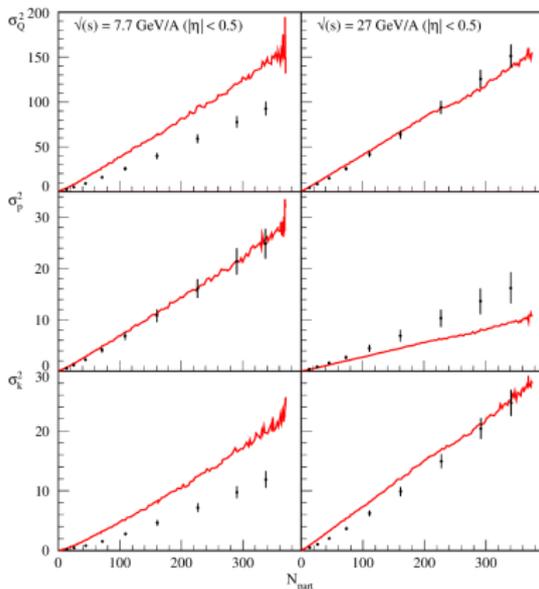
(+ only 2 energies tested yet)

⇒ EPOS underestimate ratios

Let's check in details...



(Co)variances vs Npart



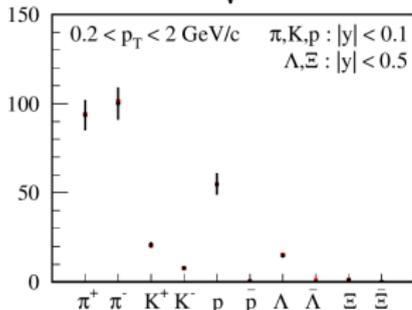
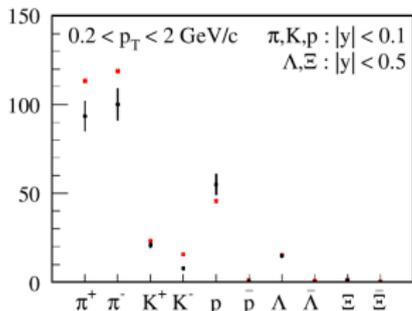
Qualitatively ✓

- EPOS general behaviour \equiv STAR data
- globally same variation trends w energy

Quantitatively ✗

- few EPOS results match with STAR data
- too much / not enough variation

Hadronic species multiplicity

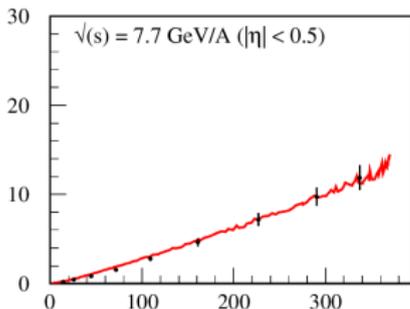
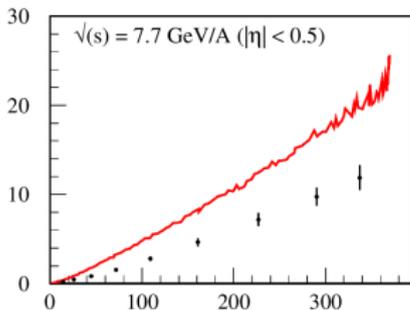


In fact, **discrepancies** for these 2nd order cumulants are **directly linked to the multiplicity** of the **considered species**.

Indeed, if we multiply net particle numbers N_X and N_Y by factors c_X and c_Y , we get :

$$\begin{aligned}\sigma_{XY}^{\prime 11} &= \langle N'_X \cdot N'_Y \rangle - \langle N'_X \rangle \langle N'_Y \rangle \\ &= \langle c_X N_X \cdot c_Y N_Y \rangle - \langle c_X N_X \rangle \langle c_Y N_Y \rangle \\ &= c_X c_Y \langle N_X \cdot N_Y \rangle - c_X \langle N_X \rangle c_Y \langle N_Y \rangle\end{aligned}$$

Hadronic species multiplicity



In fact, **discrepancies** for these 2nd order cumulants are **directly linked to the multiplicity** of the considered species.

Indeed, if we multiply net particle numbers N_X and N_Y by factors c_X and c_Y , we get :

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Then

$$\sigma_{XY}^{\prime 11} = (c_X \cdot c_Y) \times \sigma_{XY}^{11}$$

and similarly

$$\sigma_X^{\prime 2} = (c_X)^2 \times \sigma_X^{11}$$

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Summary & outlook

Plan : use last version of **EPOS 4** study the **impact** of **hadronisation** and **hadronic cascades** on **2nd order susceptibilities** of **B, Q, S** .

In particular : compare values of **STAR proxies** and best proxies proposed by **R. Bellwied *et al.*** before and after hadronic cascades
+ compare them for different hadronisation processes (grand-canonical / microcanonical)

Status :

1. compare EPOS results with STAR measured proxies :

OK qualitatively, NOT OK quantitatively

BUT "works" technically, even without any RHIC data matching test yet

⇒ finish EPOS 4 parametrisation + add proxies from R. Bellwied *et al.*

→ investigate particle production (yields, dN/dy vs $N_{part}...$)

→ check results for other energies to see better the global tendency

2. compare results before and after UrQMD

3. compare results from different hadronisation processes

4. take a look at higher order cumulants and ratios (skewness, kurtosis...) ?

Thanks for your attention !

Time for questions



Every comments or suggestions are welcome 😊

A bit more about EPOS...

More references about EPOS :

- primary interactions & hydrodynamics in EPOS
- hydrodynamics in EPOS
- heavy flavors in EPOS
- jet-fluid interaction in EPOS

Recent developments for EPOS 4 :

- parton saturation (see also [here](#))
- factorisation
- BEST equation of state inclusion

Stay tuned ! More papers to come...

Hadron Resonance Gas Model (summarised from R. Bellwied *et al.*)

It assumes that a gas of interacting hadrons in ground states can be described by a gas of non-interacting hadrons and resonances.

One can then re-write partition function, allowing to consider kinematic cuts simply by changing the phase space integration :

$$\ln(\mathcal{Z}_R) = \eta_R \frac{V \cdot d_R}{2\pi^2 T^3} \int_0^\infty p^2 \cdot dp \cdot \ln \left(1 - \eta_R \cdot z_R \cdot e^{-\varepsilon_R/T} \right)$$

Hence, with such assumption, one can decompose susceptibilities as a function of hadronic species :

$$\chi_{ijk}^{BQS}(T, \hat{\mu}_B, \hat{\mu}_Q, \hat{\mu}_S) = \sum_R \sum_{i \in \text{stable}} (P_{R \rightarrow p})^l \times B_p^i Q_p^j S_p^k \times I_l^R(T, \hat{\mu}_B, \hat{\mu}_Q, \hat{\mu}_S)$$

with :

- $l = i + j + k$
- $P_{R \rightarrow p} = \sum_\alpha N_{R \rightarrow p}^\alpha \times n_{p,\alpha}^R$: $\langle n_p \rangle$ produced in process α by each resonance R
- B_p^i, Q_p^j, S_p^k : quantum numbers of particle specie p
- $I_l^R(T, \hat{\mu}_B, \hat{\mu}_Q, \hat{\mu}_S) = \frac{\partial^l}{\partial \hat{\mu}_R^l} \left[\frac{1}{V T^3} \sum_R \ln(\mathcal{Z}_R) \right]$ ($\hat{\mu}_R = \hat{\mu}_B \cdot B_R + \hat{\mu}_Q \cdot Q_R + \hat{\mu}_S \cdot S_R$)