

Cold & Dense  
QCD : RGOPT  
improvement at  
NNLO

Loïc Fernandez,  
(PHD  
supervisor : J-L.  
Kneur)

QCD phase  
diagram

Thermal  
Quantum Field  
Theory (TQFT)

Renormalization  
Group Optimized  
Perturbation  
Theory

RGOPT :  $\lambda\phi^4$

Finite  
temperature  
QCD : RGOPT  
improvement

NNLO Cold &  
Dense QCD :  
RGOPT  
improvement

Back-up Slides

# QCD at high temperature and density : Renormalization group invariant resummation of perturbative expansion

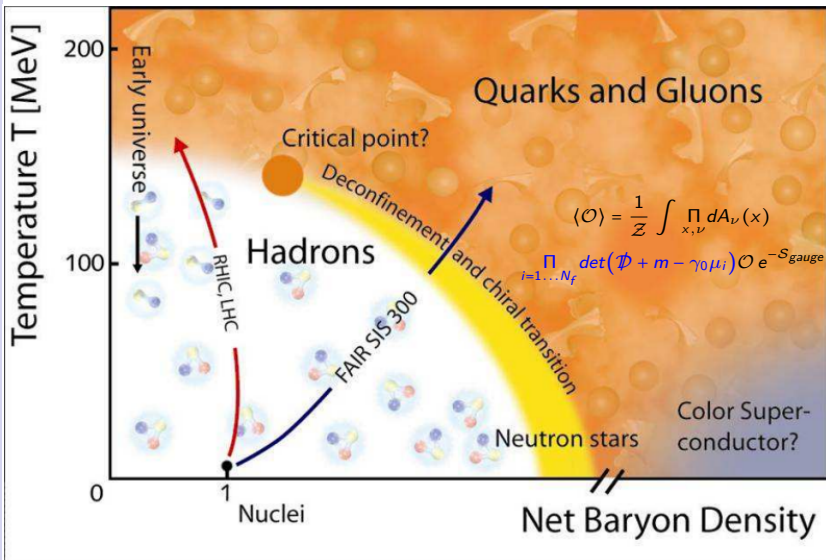
Loïc Fernandez (PHD supervisor : J-L. Kneur)

9 mars 2021



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- 2 Thermal Quantum Field Theory (TQFT)
- 3 Renormalization Group Optimized Perturbation Theory
- 4 RGOPT :  $\lambda\phi^4$
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- 6 NNLO Cold & Dense QCD : RGOPT improvement
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## QCD phase diagram



# Quantum Field Theory at finite temperature and density

Partition function of statistical physics :  $\mathcal{Z} = \text{Tr} e^{-\beta(\hat{H}-\mu\hat{N})}$ .  
Starting from QFT :  $\mathcal{Z}_{QFT}$  (Imaginary time formalism)

- 1  $\int_0^{+\infty} dt \xrightarrow{t=-i\tau} \int_0^{\frac{1}{T}} d\tau$  ,  $p_0 \rightarrow p_0 - i\mu$
- 2 Periodic/anti-periodic B.Cs for Bosons/Fermions
- 3 In Fourier Space :

$$\int_0^{\frac{1}{T}} d\tau \rightarrow \sum_{n=-\infty}^{\infty} , \quad p_0 \rightarrow \omega_n = \begin{cases} 2\pi n T & \text{Bosons} \\ (2n+1)\pi T - i\mu & \text{Fermions} \end{cases}$$

- 4  $T \sum_{\{\omega_n\}} \int d^3\vec{p} = \oint_{\{P\}}$  ,  $T \sum_{\omega_n} \int d^3\vec{p} = \oint_P$

$$\mathcal{Z}_{QFT} = \mathcal{Z}_{\text{free gas}} + \text{radiative corrections}$$

# Infrared divergences ( $\lambda\phi^4$ model)

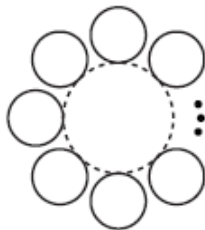
$\lambda\phi^4$  shares properties similar to QCD.

If  $m = 0$  (gluons) and  $n = 0 \Rightarrow \omega_n = 0$  then :

$$\int dp \frac{1}{(p^2)^k},$$

- IR-divergent !

**Daisy diagram** : Most divergent  
diagram at order N



# IR divergences : Resummation

- Structure of a Taylor Expansion
- Resum a subclass of **Daisy diagrams**
- 

$$\sum_{N=0}^{\infty} \frac{1}{N!} \left(\frac{\lambda_B T^2}{4}\right)^N \left(\frac{d}{dm_B^2}\right)^N \left(\frac{-m_B^3 T}{12\pi}\right) = \frac{-T}{12\pi} \left(m_B^2 + \frac{\lambda_B T^2}{4}\right)^{\frac{3}{2}}$$

- **Thermal mass** generated by the dynamics of the theory
- Expansion in massless theory  $\rightarrow$  IR-divergences
- "Equivalent" resummation :  $\mathcal{L} \rightarrow \mathcal{L} + m_{thermal}^2 \phi^2$
- Other approach of resummation : (OPT/SPT)

# Renormalization Group Optimized Perturbation Theory (RGOPT)

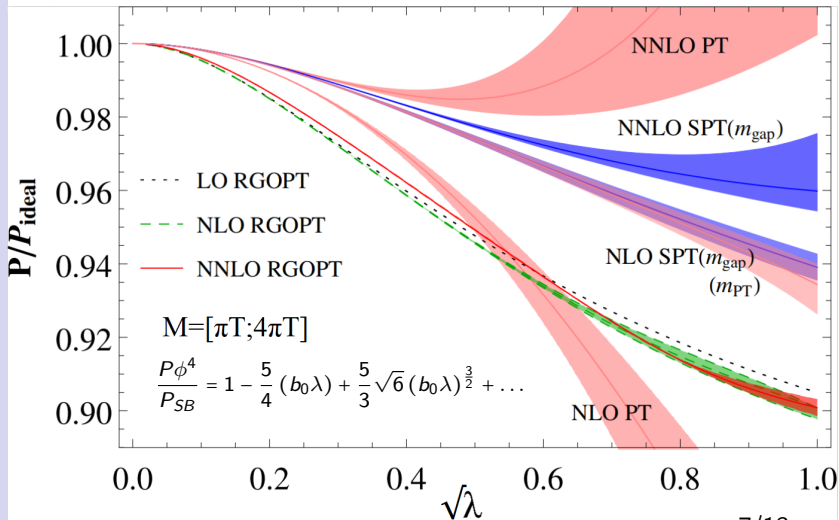
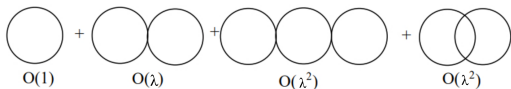
- J-L. Kneur, A. Neveu, QCD  $T=0$ , '13
  - J-L. Kneur, M. Pinto,  $\lambda\phi^4$   $T \neq 0$ , '16
- ①  $\mathcal{L}(\lambda, 0) \rightarrow \mathcal{L}(\delta\lambda, m(1-\delta)^a)$
  - ②  $m \rightarrow m(1-\delta)^a$  ;  $\lambda \rightarrow \delta\lambda$
  - ③ Expand in  $\delta$  at order  $\mathcal{O}(\lambda^k)$  then  $\delta \rightarrow 1$
  - ④ RG invariance requires : "a" and vacuum subtraction terms

$$\left. \frac{\partial}{\partial m} \mathcal{P}(\lambda, m) \right|_{m=\bar{m}} = 0 ; \left( M \frac{\partial}{\partial m} + \beta(g) \frac{\partial}{\partial g} - \gamma_m m \frac{\partial}{\partial m} \right) \mathcal{P}(\lambda, m) = 0$$

OPT  $\oplus$  RG = reduced RG :

$$\left( M \frac{\partial}{\partial M} + \beta(\lambda) \frac{\partial}{\partial \lambda} \right) \mathcal{P}(\lambda, m, \delta = 1) = 0$$

one-loop Order : RG  $\rightarrow a = \frac{\gamma_0}{b_0} = \frac{1}{6}$  & OPT  $\rightarrow \bar{m}^2 \sim \mathcal{O}(\lambda T^2)$





# Mass term for the Gluons : Hard thermal Loops

$m^2 A_\mu A^\mu$  : explicitly breaks **gauge invariance**, (Curci-Ferrari model, cf. Van Egmond's talk)

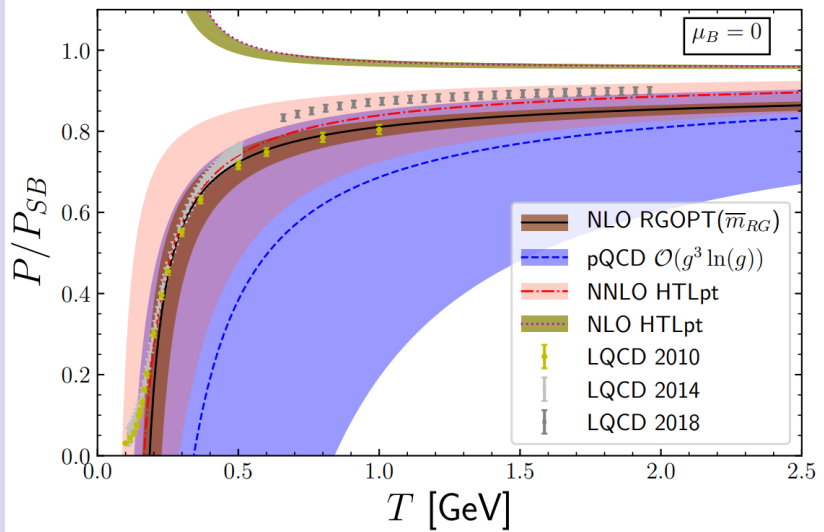
Hard Thermal Loop term (E.Braaten and D.Pisarski (1990)) :

$$-\frac{m^2}{2} \text{Tr} \left[ G_{\mu\alpha} \left\langle \frac{y^\alpha y^\beta}{(y \cdot \mathcal{D})^2} \right\rangle G^\mu{}_\beta \right] ; \quad y^\mu = (1, \vec{y})$$

**Gauge invariant but non-linear and add dressed vertices and propagators !**

Systematic high temperature expansion : HTLpt (Andersen et al. 2010-2014) State-of-the-art calculations : NNLO

# High Temperature QCD



# RGOPT LO Cold& Dense

RG invariance needs vacuum subtraction :

$$g = 4\pi\alpha_s \ ; \ \mathcal{E}_0 = \frac{-m^4}{g} \sum_{n=0}^k s_n g^n \ ; \ LO \rightarrow \ \mathcal{E}_0 = \frac{-m^4 s_0}{g}$$

(missing in HTL,HDLpt)

$$\mathcal{P}_{0,f}^{PT}(\mu, m) = N_c \frac{m^4}{(4\pi)^2 b_0 g} - N_c \frac{m^4}{8\pi^2} \left( \frac{3}{4} - \text{Log} \left( \frac{m}{M} \right) \right) \\ + \Theta(\mu^2 - m^2) \frac{N_c}{12\pi^2} \left[ \mu p_F \left( \mu^2 - \frac{5}{2} m^2 \right) + \frac{3}{2} m^4 \ln \left( \frac{\mu + p_F}{m} \right) \right]$$

$$OPT \rightarrow \bar{m} \sim \mathcal{O}(\sqrt{g}\mu) \ \& \ RG \rightarrow a = \frac{\gamma_0}{2b_0}$$

$$P^{PT} = \text{Diagram 1} + \text{Diagram 2} + \mathcal{O}(g^2)$$

The diagram shows the perturbative expansion of the vacuum energy  $P^{PT}$ . It consists of three terms: a one-loop vacuum bubble (a circle with two arrows), a two-loop vacuum bubble with a gluon exchange (two circles with arrows, connected by a wavy line), and higher-order terms  $\mathcal{O}(g^2)$ .

# Renormalization Scheme Change at NNLO

$$m \rightarrow m(1 + B3g^3)$$

RSC was not needed for  $\lambda\phi^4$

Theoretical constraint on  $\bar{m}$  :

- Asymptotic Freedom (AF) matching

Expected properties of  $\bar{m}$  :

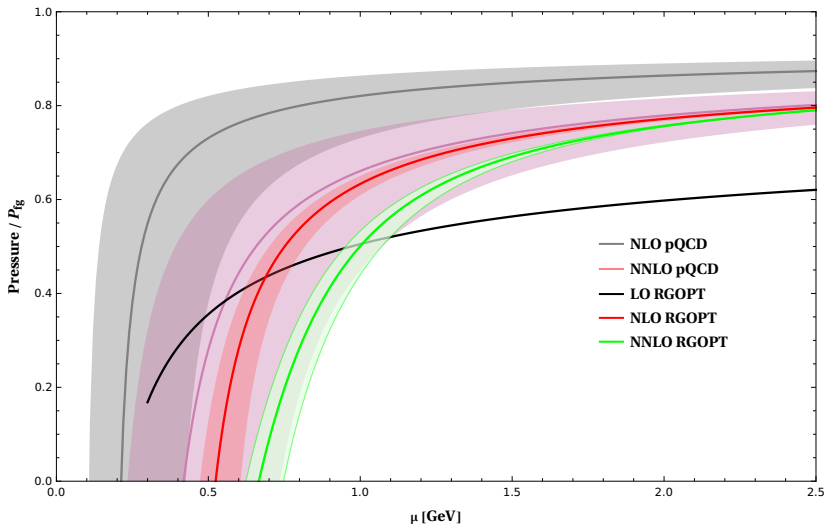
- Pressure should not exceed  $P_{fg}$
- $P(M = 4\mu) > P(M = \mu)$  for large enough  $\mu$
- Small deviation from  $\bar{MS}$  scheme

Prescription :

*Fix  $B3(g)$  such as to restore reality of the solution*

NB : In HTLpt,  $m_{PT}$  is used, so no problem of imaginary solutions.  
But loose resummation properties. Start-of-the-art : NNLO

# (Preliminary !!) Results



# Future of RGOPT

- Equation of State at NNLO for Neutron Star
- Include (HTL) variational mass for gluons
- NLO, and ultimately, NNLO for full QCD at  $(T, \mu)$

# NLO Cold&Dense QCD

- No real solutions for every  $\mu$  ( $\text{Im}(m) \neq 0$ )

One way to recover realness of the solutions : Renormalization Scheme Change (RSC)

$$m \rightarrow m (1 + B2g^2)$$

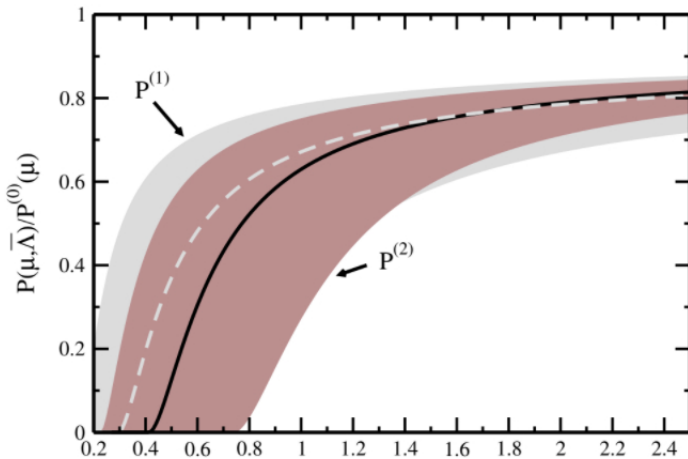
- LO&NLO :  $\text{Log}\left(\frac{m}{M}\right)$  and  $\text{Log}\left(\frac{\mu+p_F}{m}\right)$  recombine as  $\text{Log}\left(\frac{\mu+p_F}{M}\right) = L_\mu$
- RG equation  $\leftrightarrow c1(m,g,M) L_\mu^2 + c2(m,g,M) L_\mu + c3(m,g,M)$
- Quadratic equation in  $L_\mu$
- Discriminant suggest :

$$B2 = -\frac{7}{81\pi^2 g}$$

# NNLO Cold&Dense QCD

Starting from  $\mathcal{P}_2^{PT}(g, \mu, m)$  calculated by :

- A. Kurkela, P. Romatschke, and A. Vuorinen. Cold Quark Matter. *Phys. Rev. D*, 81:105021, 2010





# NNLO Cold&Dense QCD

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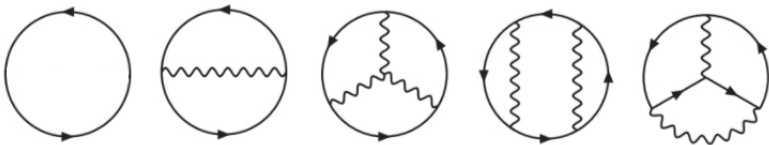
NNLO Cold &  
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Back-up Slides



Model :  $N_f = 3$  ;  $m_u = m_d = m_s = 0$  ;  $m_{\text{variational}}$   
 (Investigating  $N_f = 2 + 1$  ;  $m_u = m_d = 0$  ;  $m_s$  ;  $m_{v,ud} \neq m_{v,s}$ )  
 Starting from  $\mathcal{P}_2^{PT}(g, \mu, m)$  without the Gluons.

- ①  $\mathcal{P}_{2,f}^{PT}(g, \mu, m, \delta = 1)$  (for one flavor)
- ② Reduced RG equation gives a real solution  $\bar{m}$  for  $\mu \in [1.2, 2.5]$  GeV



## Strange quark current mass :

$$\mathcal{N}_F = 2 + 1$$

$m_s$  considered as a perturbation to the (all-order) original massless theory :

$$\mathcal{N}_F \mathcal{P}_{2,f}(g, m) \rightarrow (\mathcal{N}_F - 1) \mathcal{P}_{2,f}(g, m) + \mathcal{P}_{2,f}(g, m + m_s)$$

$$M \frac{d}{dM} = M \frac{\partial}{\partial M} + \beta(g) \frac{\partial}{\partial g} - \gamma_m(g) m_s \frac{\partial}{\partial m_s}$$

# Renormalization Group

Renormalization  $\rightarrow$  Scale dependence :  $M$

$$g \rightarrow g(M) ; m \rightarrow m(M)$$

$$\alpha_s(M = 1 \text{ GeV}) \simeq 0.42 \rightarrow \Lambda_{QCD}$$

Observables are independent of the choice of  $M$

Perturbative RG equation :

$$M \frac{d}{dM} \mathcal{O} = \left( M \frac{\partial}{\partial M} + \beta(g) \frac{\partial}{\partial g} - \gamma_m m \frac{\partial}{\partial m} \right) \mathcal{O} = 0$$

OPT **breaks** the RG invariance

$$\begin{aligned}
 \mathcal{P}_{2,f}^{PT}(\mu, m) = & -N_c \frac{m^4}{8\pi^2} \left( \frac{3}{4} - L_m \right) + \Theta(\mu^2 - m^2) \frac{N_c}{12\pi^2} \left[ \mu p_F \left( \mu^2 - \frac{5}{2} m^2 \right) + \frac{3}{2} m^4 \ln \left( \frac{\mu + p_F}{m} \right) \right] \\
 & - \frac{d_A g}{4(2\pi)^4} m^4 \left( 3L_m^2 - 4L_m + \frac{9}{4} \right) - \Theta(\mu^2 - m^2) \frac{d_A g}{4(2\pi)^4} \left\{ 3 \left[ m^2 \ln \left( \frac{\mu + p_F}{m} \right) - \mu p_F \right]^2 - 2p_F^4 \right\} \\
 & - \Theta(\mu^2 - m^2) \frac{d_A g}{4(2\pi)^4} m^2 (4 - 6L_m) \left[ \mu p_F - m^2 \ln \left( \frac{\mu + p_F}{m} \right) \right] \\
 & + \frac{g^2 m^4}{135(4\pi)^6} \left( \alpha_{0,2} + \alpha_{1,2} L_m + \alpha_{2,2} L_m^2 + \alpha_{3,2} L_m^3 \right) \\
 & + \frac{g^2 d_A \mu^4}{(4\pi)^4 2\pi^2} \Theta(\mu^2 - m^2) \left\{ -\hat{m}^2 [(11C_A - 2N_f)z + 18C_F(2z - \hat{u})](L_m)^2 - \frac{1}{3} \left[ C_A \left( 22\hat{u}^4 - \frac{185}{2} z\hat{m}^2 - 33z^2 \right) \right. \right. \\
 & \left. \left. + \frac{9C_F}{2} (16\hat{m}^2 \hat{u}(1 - \hat{u}) - 3(7\hat{m}^2 - 8\hat{u})z - 24z^2) - N_f(4\hat{u}^4 - 13z\hat{m}^2 - 6z^2) \right] L_m \right. \\
 & \left. + C_A \left( -\frac{11}{3} \ln \frac{\hat{m}}{2} - \frac{71}{9} + G_1(\hat{m}) \right) + C_F \left( \frac{17}{4} + G_2(\hat{m}) \right) + N_f \left( \frac{2}{3} \ln \frac{\hat{m}}{2} + \frac{11}{9} + G_3(\hat{m}) \right) + G_4(\hat{m}) \right\}, \tag{1}
 \end{aligned}$$

$$\alpha_{0,2} = (357315 + 176\pi^4 + 960\pi^2(\log 2)^2 - 960(\log 2)^4 - 23040 Li_4(1/2) + 12960 \zeta(3))$$

$$+ 90N_f(-393 + 224 \zeta(3))$$

$$\alpha_{1,2} = 180(-3817 + 286N_f + 48 \zeta(3))$$

$$\alpha_{2,2} = -720(-807 + 26N_f)$$

$$\alpha_{3,2} = 2880(-81 + 2N_f)$$

$$\hat{m} = \frac{m}{\mu}, \quad L_m = \ln \frac{m}{M}, \quad \hat{u} = \frac{u}{\mu} = \frac{\sqrt{\mu^2 - m^2}}{\mu}, \quad z = \hat{u} - \hat{m}^2 \ln \frac{1 + \hat{u}}{\hat{m}}, \quad L = \ln \hat{m}$$

$$G_1(\hat{m}) = 32\pi^4 \hat{m}^2 (-0.01863 + 0.02038\hat{m}^2 - 0.039\hat{m}^2 L + 0.02581\hat{m}^2 L^2 - 0.03153\hat{m}^2 L^3 + 0.01151\hat{m}^2 L^4)$$

$$G_2(\hat{m}) = 32\pi^4 \hat{m}^2 (-0.1998 - 0.04797L + 0.1988\hat{m}^2 - 0.3569\hat{m}^2 L + 0.3043\hat{m}^2 L^2 - 0.1611\hat{m}^2 L^3 + 0.09791\hat{m}^2 L^4)$$

$$G_3(\hat{m}) = 32\pi^4 \hat{m}^2 (-0.05741 - 0.02679L - 0.002828L^2 + 0.05716\hat{m}^2 - 0.08777\hat{m}^2 L + 0.0666\hat{m}^2 L^2 - 0.02381\hat{m}^2 L^3 + 0.01384\hat{m}^2 L^4)$$

$$G_4(\hat{m}) = 32\pi^4 \hat{m}^2 (0.07823 + 0.0388L + 0.004873L^2 - 0.07822\hat{m}^2 + 0.1183\hat{m}^2 L - 0.08755\hat{m}^2 L^2 + 0.03293\hat{m}^2 L^3 - 0.01644\hat{m}^2 L^4).$$

# Sign Problem

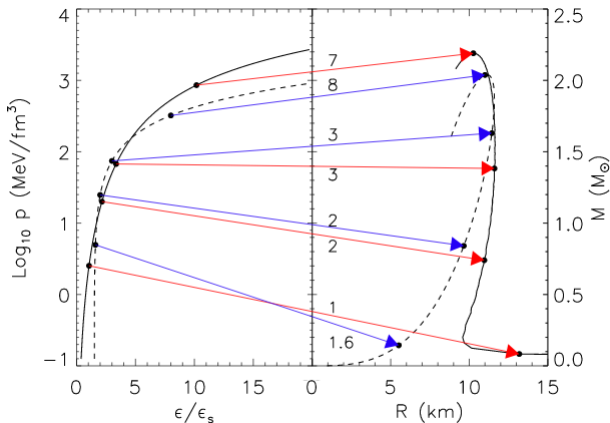
$$\mathcal{Z} = \int \mathcal{D}A_a^\mu \mathcal{D}\bar{\psi} \mathcal{D}\psi \mathcal{D}\bar{C} \mathcal{D}C e^{-\int_0^\beta d\tau \int d^3x (\mathcal{L}_E - \psi \mu \bar{\gamma}^0 \psi)} \quad (3)$$

Expectation value of an operator  $\mathcal{O}$  :

$$\langle \mathcal{O} \rangle = \frac{1}{\mathcal{Z}} \int \prod_{x,\nu} dA_\nu(x) \prod_{i=1 \dots N_f} \det(\not{D} + m - \gamma_0 \mu_i) \mathcal{O} e^{-S_{gauge}}$$

With  $\mu_i \neq 0$  the determinant is **no longer positive definite** which is needed for Monte-Carlo method used for evaluation.

# Neutron Star



$M_{MAX} \lesssim 2.1 M_{\odot}$   
 $M > 0.9 M_{\odot}$   
 $C_s^2 < 1$  : Causal  
 limit  
 $C_s^2 < 1/3$  :  
 Conformal limit

J. M. Lattimer, Annu. Rev. Nucl. Part. Sci. 62, 485 (2012), URL  
<https://doi.org/10.1146/>

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# state-of-the-art of NS Equation of State (EoS)

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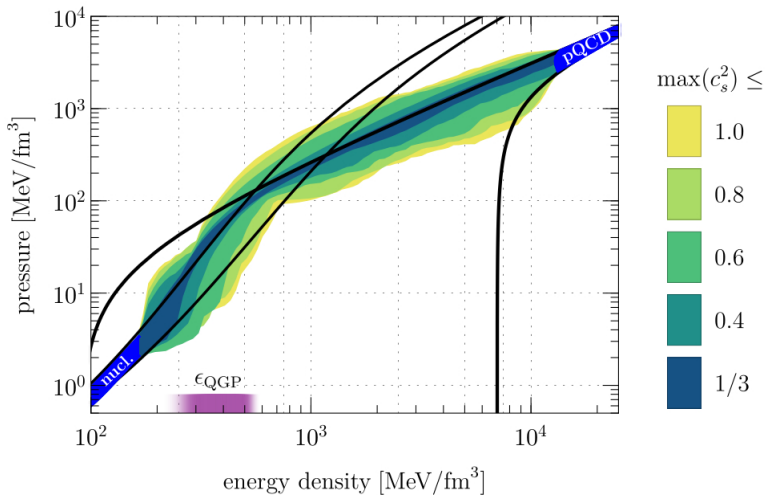
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# Tolman Oppenheimer Volkoff equation (TOV)

Assume a static, spherically symmetric perfect fluid.

$$g_{\mu\nu} dx^\mu dx^\nu = e^\nu c^2 dt^2 - \left(1 - \frac{2G m}{r c^2}\right)^{-1} dr^2 - r^2 (d\theta^2 + \sin^2\theta d\phi^2)$$

$$\text{EinsteinEquations} \rightarrow \frac{d\nu}{dr} = - \left(1 - \frac{2G m}{r c^2}\right) \frac{dP}{dr}$$

$$\left\{ \begin{array}{l} \frac{dp}{dr} = - \frac{G\epsilon(r)\mathcal{M}(r)}{c^2 r^2} \left[1 + \frac{p(r)}{\epsilon(r)}\right] \left[1 + \frac{4\pi r^3 p(r)}{\mathcal{M}(r)c^2}\right] \left[1 - \frac{2G\mathcal{M}(r)}{c^2 r}\right]^{-1} \\ \frac{d\mathcal{M}}{dr} = 4\pi r^2 \rho(r) = \frac{4\pi r^2 \epsilon(r)}{c^2} \\ \mathcal{M}(r) = 4\pi \int_0^r dr' r'^2 \rho(r') = 4\pi \int_0^r dr' r'^2 \frac{\epsilon(r')}{c^2} \end{array} \right.$$



## RGOPT LO : Mass VS Radius

