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LDM and Hadron Collider Predictions

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Outline

- Light Dark Matter
- LDM & p-p collisions
- Background vs. Signal
 - Conclusions

Light Dark Matter

LDM - Basics

C. Boehm et al., arXiv:0305261 [hep-ph]

- Aussume Mirror-parity (Z₂ symmetry)
- Non- chiral Yukawa couplings between SM fermions (f), their mirror partners (F) and a scalar field (S)
- Lowest mass particle: scalar S, m_s~ (MeV-GeV)
- S is stable thanks to Mirror-parity
 → S is a Dark Matter candidate



LDM - Cosmology

- Coupled to SM primordial thermal bath
- Chemical decoupling: Relic Density $\Omega_m h^2$
- Structure formation: LDM ≡ CDM (Boehm et al., arXiv:0012504 & arXiv:0410591 [astro-ph])



LDM – Relic Density

Follow DM distribution function throughout universal history:

- Boltzmann Equation L[f(x, p)] = C[f(x, p)]
- Number Density

$$\frac{g}{\left(2\pi\right)^{3}}\int f\left(x,p\right)d^{3}p\equiv n$$



$$LDM - Boltzmann Code$$

$$y(T) = \frac{n(T)}{T^3} \propto N(T) \quad \longleftrightarrow \quad \frac{dy}{dT} = \frac{\langle \sigma v \rangle_{ann}}{H_0} (y^2 - y_0^2)$$

- Start from thermal and chemical equilibrium at T > m_{DM}
- Follow exponential suppression with an implicit integrator
- By imposing y_0 , get $\langle \sigma v \rangle_{ann} \& T_{fo}$



$$\langle \sigma v \rangle_{ann} \sim 10^{-26} cm^3 s^{-1}$$



LDM – Charged Lepton Side

Unhealthy Implications

• Gamma-ray production in galaxies: $\sigma_{SS \rightarrow l^+l^-} v_{rel} < O(10^{-[31-28]} cm^3 s^{-1})$ (for m_{DM} ε [1-100] MeV)

•Anomalous magnetic moments (g-2) constraints for electrons and muons:

$$\delta a_{\mu,e}^{m_F \gg m_{\mu,e}} \simeq \frac{m_{\mu,e}}{16\pi^2} \frac{c_l c_r}{m_F}$$

Direct constraint for annihilation cross section:

$$\sigma_{SS \to l^+ l^-} v_{rel} \propto \frac{c_l c_r}{m_E} \longrightarrow \sigma_{SS \to l^+ l^-} v_{rel} < O(10^{-29} cm^3 s^{-1})$$

Cannot achieve relic density through annihilation into e nor into µ

LDM – Charged Lepton Side

Vigorous Application

• 511 keV line detected from the galactic center (Johnson et al., ApJ, 1972) \rightarrow presence of low energy positrons

• Morphology: distributed on the galactic bulge, suggests correlation with a typical DM halo profile

• Best fit to INTEGRAL signal: DM halo profile + annihilating LDM into e⁺e⁻ through F_e exchange (Ascasibar et al., arXiv:0507142 [astro-ph])

LDM – Neutrinos

• No particle physics nor astrophysical contraints yet \rightarrow can produce relic density through S-neutrino interactions: $\sigma_{SS \rightarrow vv} v_{rel} \sim O(10^{-26} cm^3 s^{-1})$

• As a bonus: if neutrinos are Majorana fermions, then expect



and if apply similar couplings and masses, get $m_v \le 1 \text{ eV}$. However, this would require F_v to be a singlet of (SU(2) X U(1)) (Boehm et al., arXiv:0612228 [hep-ph]).

$$LDM - Flavor Side$$

$$L \supset \left[s_{DM} \left(c_{l} \overline{q_{l}} F_{q,r} + c_{r} \overline{q_{r}} F_{q,l} \right) + h.c. \right] \xrightarrow{F_{q}}$$

- Assume same mass for all flavors
- No flavor mixing in the mirror side
- F_q colored particles, therefore strong coupling to gluons
- If m_{dm} < 100 MeV, no annihilation into quarks (nor heavy leptons)

q

LDM – Parameter Space

- New charged particle searches: $m_F > 100 \text{ GeV}$
- Big Bang Nucleosynthesis: $T_{fo} > T_{BBN} = 100 \text{ keV}$ => $m_{DM} > 1 \text{ MeV}$
- Relic Density in neutrino sector only: $m_{DM} < m_{\pi}$
- c_{l,r} can be large! (still want to be perturbative) : [0.1 ; 3]

LDM & p-p collisions

p-p collisions



 $protons \supset \{u, d, s, c, b, t, g\}$



di-F_q production (≈ di-F_q)



Exchange of a very light particle!

 $F_q - \overline{F_q}$ production



F_q production (≈ F_q)



Cross Sections computed using:

- cs22 function implemented in MicrOmegas (G. Bélangeret al., arXiv:0803.2360 [hep-ph])
- Vegas routine in Calchep-Comphep (A. Pukhov et al., arXiv:0908288 [hep-ph])
- Checked values analytically by comparing symbolic calculations in Calchep-Comphep with our own expressions
- Chose $m_F = 300 \text{ GeV}, m_{DM} = 2 \text{ MeV}$
- Implement only the case $q \equiv u$







p-p collisions – ATLAS

- Starting physics program soon
- Data will be taken at $\sqrt{s} = 7 \ TeV$
- Compute proton-proton collisions using hCollider function implemented in MicrOmegas
- Chose mF = 300 GeV, mDM = 2 MeV











p-p collisions – signature



p-p collisions

Expected events:

• From u-u, u- \overline{u} and g-g collisions:

 $p-p \rightarrow F + F(F) \rightarrow 2$ jets + missing energy

From u-g collisions:

 $p-p \rightarrow F + DM \rightarrow 1$ jet + missing energy

Background vs. Signal

Background vs. Signal

• Largest "2 jet + missing energy" events expected:

- Z + 2 jets
- $t,\overline{t} \rightarrow b,W \rightarrow b,I,v$ (ATLAS has b-tagging!)
- WW, WZ
- Generation of events using Comphep: p-p → F-F, F-F

 $c_{l,r} = 1$ $m_{DM} = 2 \text{ MeV}$ $m_F = 300 \text{ GeV}$ pcm = 3.5 TeV

- Choose a rather conservative point
- Include Z + 1, 2, 3 jet events and compare Z to one F





Background vs. Signal

• Perspective: a cut based analysis eliminates background down to a few fb (The ATLAS collaboration, arXiv:0901.0512 [hep-ex])

 Carry on a full simulation including final sates using well established MC simulators

• Wait for LHC to look for LDM!

 TEVATRON: background has been studied, but cross section behavior, kinematical cuts and conclusions are model dependent

Conclusions

- Ensuring the Relic Density in neutrino sector allows the Hadronic sector to be more reactive
- Light Dark Matter is a DM candidate that provides predictions for colliders
- Cross sections for p-p collisions can be (amazingly) large!
- LHC (ATLAS) may have the potential to explore a good deal of parameter space, namely masses and couplings
- Complete kinematic analysis is in preparation
- Expand the number of events to take into account:
 - Other flavors (i.e. include d, s, c, b, t)
 - Leptonic production

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