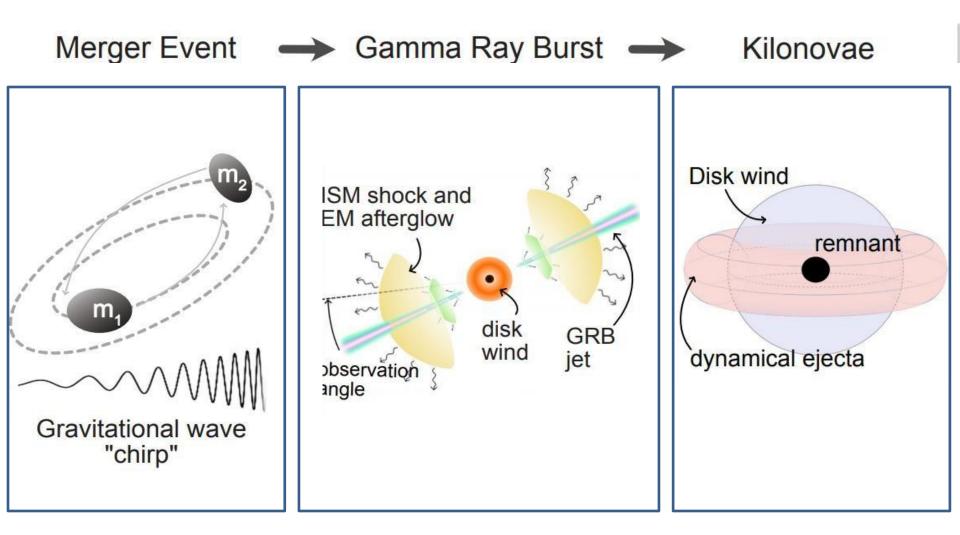




Enabling multi-messenger cosmology with the LIGO-Virgo alert system

S. Antier (CNRS/APC - CNES) antier@apc.in2p3.fr

Coupling the messengers of compact objects coalescences

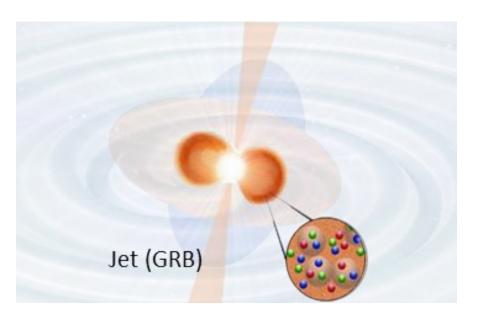


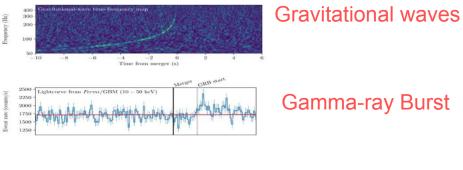
GW170817 – The only example

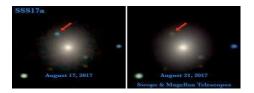
Matter behavior in extreme conditions ?

GW170817

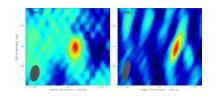
Uncertainties in the Equation of State of Utra-dense matter







Kilonova



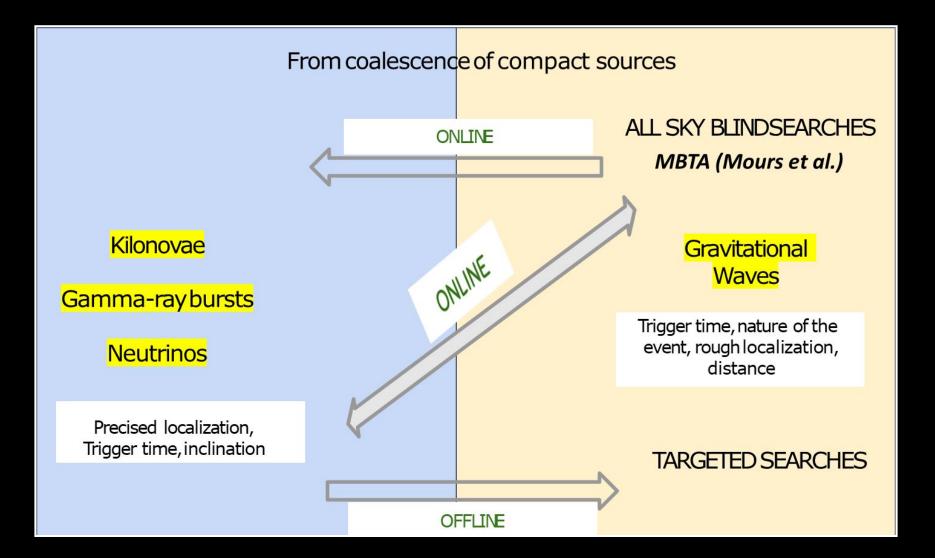
Afterglow

How are produced ejected matter in compact collision ?

Uncertainty of the central object Mechanisms for the jet emission



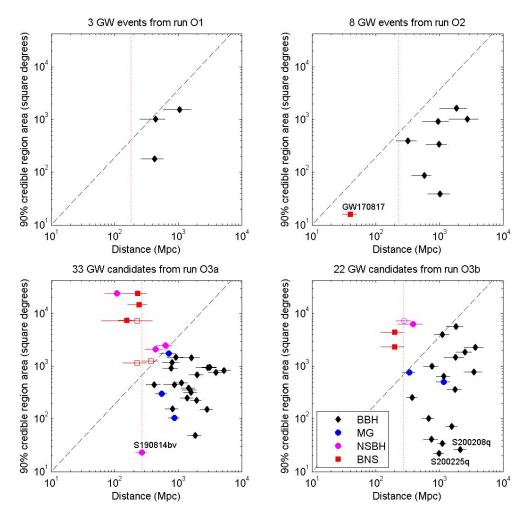
LIGO-Virgo Alert System



Enabling multi-messenger discoveries with



Alert statistics O1, O2 and O3



Antier et al., GRANDMA O3, 2020

Fom april 2019 to march 2020 ~ 330 days

80 +1 alerts with 56+1 "still alive" 24 retractations (30 % retracted)

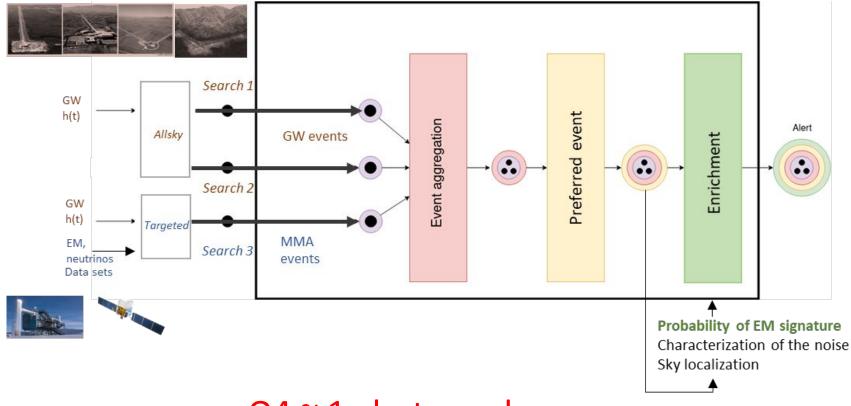
A traffic of alerts 4 times bigger than O2

41 alerts with updated sky localization areas (73% of the total of the alerts)

52 of the alerts with P (Terres) < 50 %

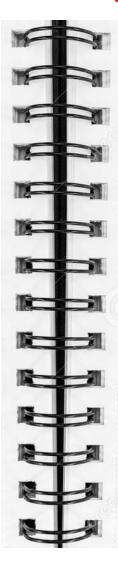
8 BNS candidates 6 NS-BH candidates 5 Mass-Gap candidates 36 BBH candidates 1 burst candidate 1 LIGO/Virgo - GBM-190816

Toward O4 - The LIGO-Virgo alert system



O4 ~ 1 alert per day

Work in progress in the alert system I am the key contact for Virgo-CNRS node



- Integrate the KAGRA detector
 Reduce the latency of the alert (< 5 min)
- Provide early-warning alert before the merger time
- Integrate new observatories (Km3NET, SVOM)
- Reduce the flow of false positive alerts (purity ~90 %)
- Provide updates with full paramters estimation within a day
- Provide from an update to another, how informations has changed
- Work on visibility and clearness

Predicted mass ejecta and HasEjecta on alerts Stachie et al.

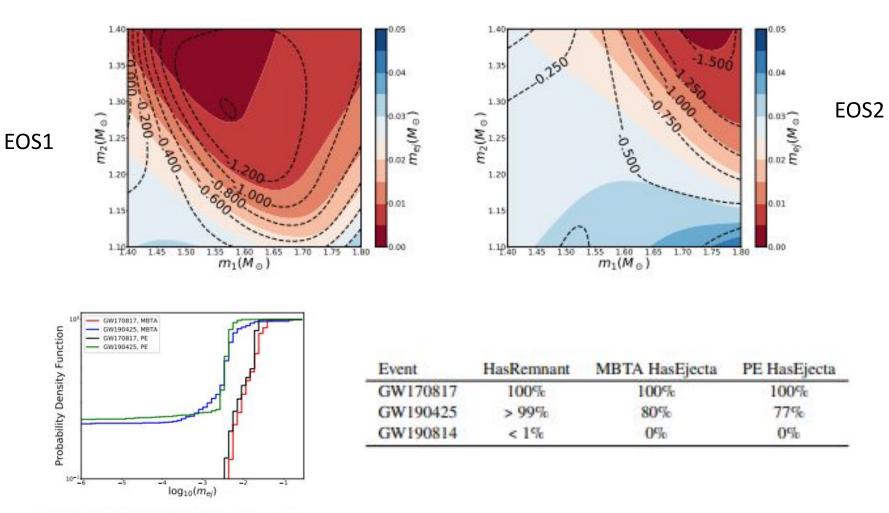
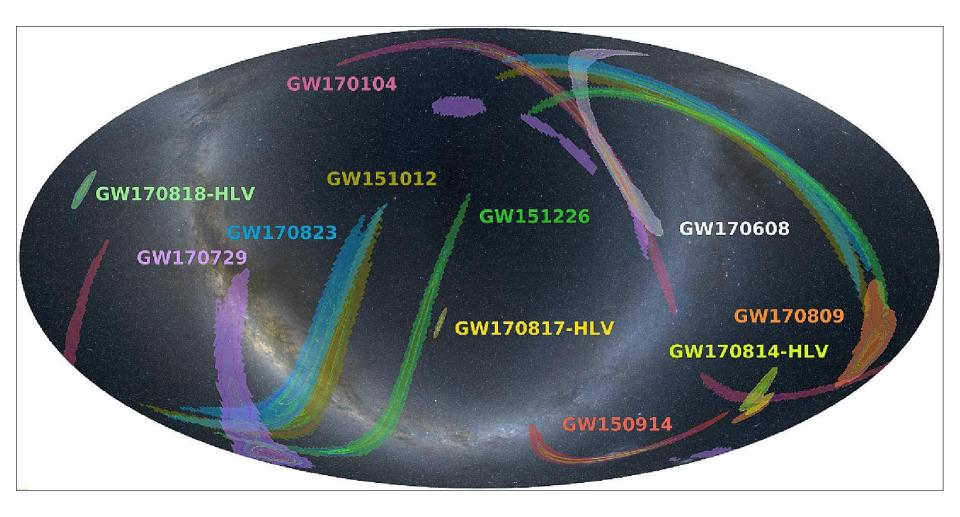


Figure 7. Cumulative distribution function of the ejected mass m_{ej} for the following GW triggers: GW170817, GW190425. There are both the low-latency results (the input data is represented by the MBTA weighted templates) and the PE results (the input data is represented by the offline PE (Veitch et al. 2015) posteriors).

Localizations



P-A Duverne, Singer et al.

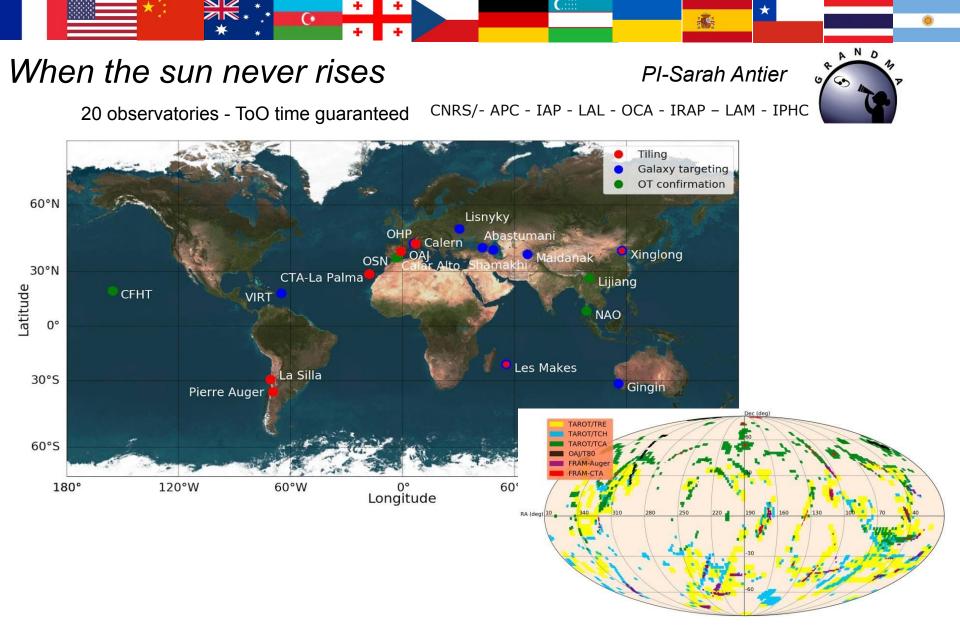
https://git.ligo.org/lscsoft/lalsuite-archive/blob/5fc7b8eede53e3aee325dd16f0916e5d77636510/lalinference/python/lalinference/ba

Once the alert is received

Lots of observations are running



Everyone is looking at the same region of the search area to find the counterpart of GW events Our proposition - Coordination



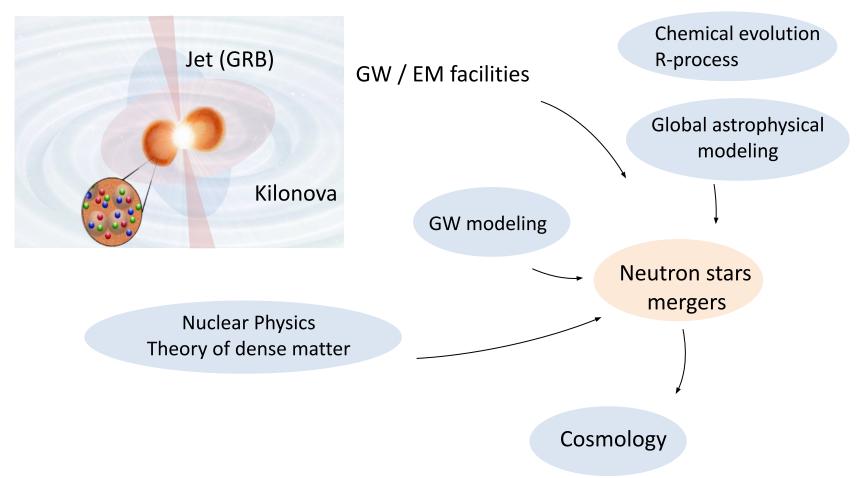
O3b and global summary of O3: <u>GRANDMA Observations of Advanced LIGO's and Advanced Virgo's Third</u> <u>Observational Campaign</u> <u>O3a and presentation of the collaboration</u>: <u>The first six months of the Advanced LIGO's and</u> <u>Advanced Virgo's third observing run with GRANDMA, 2020, MNRAS, 492, 3904</u>

Once EM counterpart is

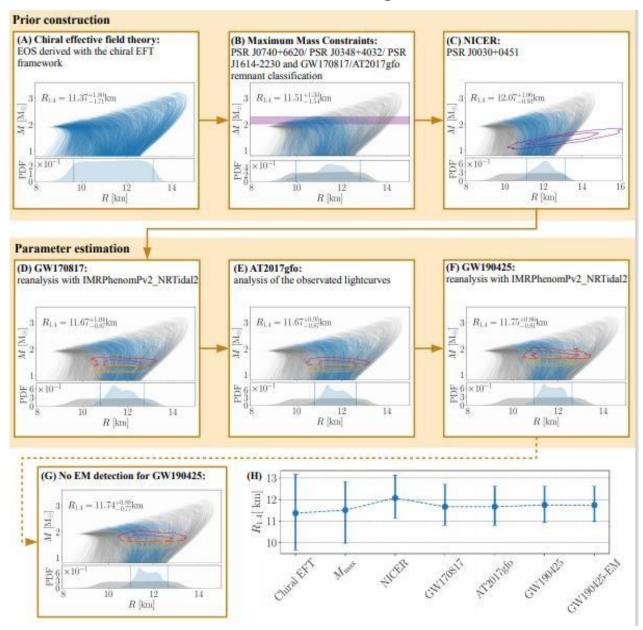
found

Multi-physics

Multi-physics framework



The multi-messenger framework

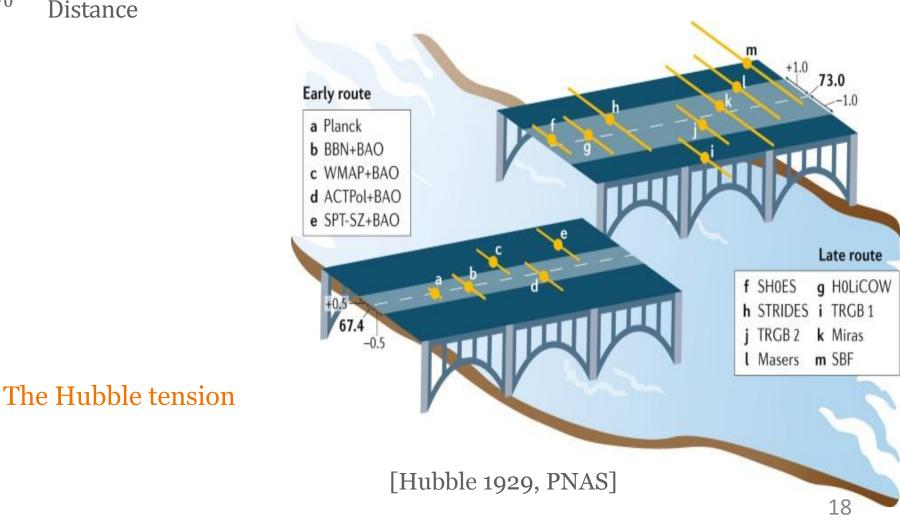


Dietrich T et al., New Constraints on the Supranuclear Equation of State and the Hubble Constant from Nuclear 16 Physics – MMA

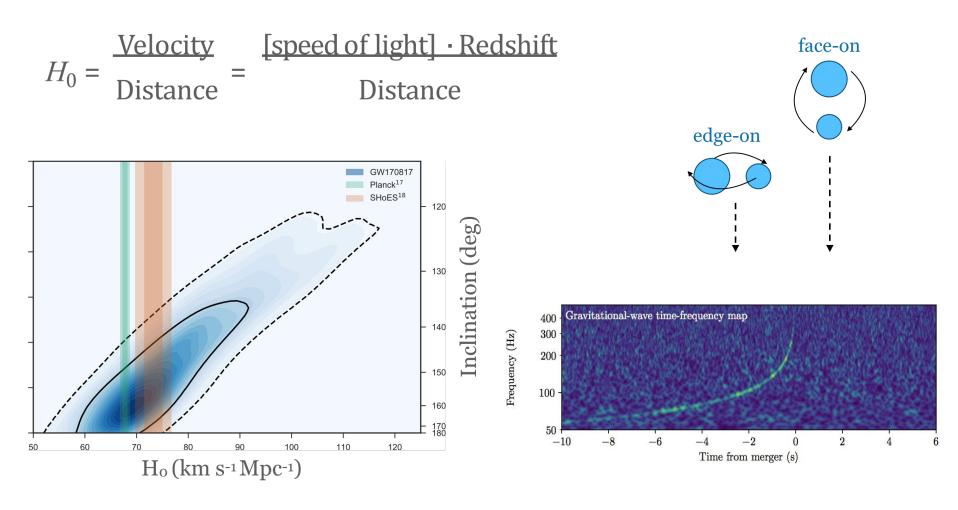
The Ultimate -Cosmology

The local expansion rate of the Universe

 $H_0 = \frac{\text{Velocity}}{\text{Distance}}$

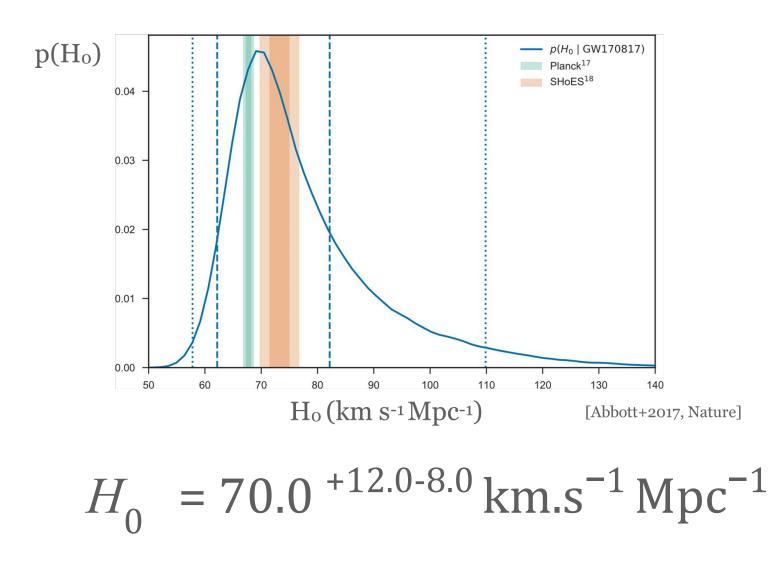


[Schutz 1986, Nature; Holz & Hughes 2005, ApJ]

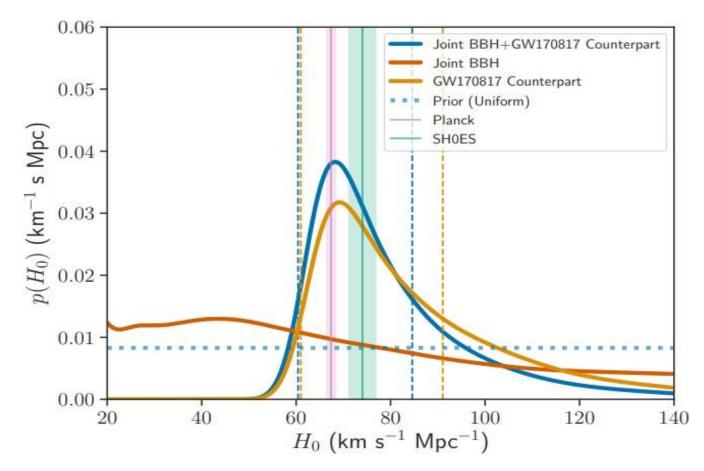


19

[Schutz 1986, Nature; Holz & Hughes 2005, ApJ]



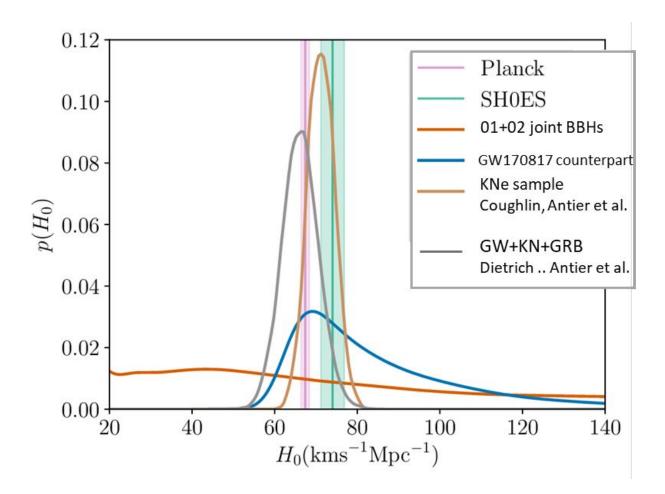
[Schutz 1986, Nature; Holz & Hughes 2005, ApJ]



A gravitational-wave measurement of the Hubble constant following the second observing run of Advanced LIGO and Virgo, O2 run, LVC

- Method 1 : GW + KN
- Method 2 : Statistical approchs with BBH (prob loca and catalogs)

[Schutz 1986, Nature; Holz & Hughes 2005, ApJ]



- Method 1 : GW + KN + help the degenary of the distance inclination
- Method 2 : Statistical approchs with BBH (prob loca and catalogs)
- Method 3 : KNe as standard canddles