

Probing primordial features with the SGWB

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with J. Fumagalli and L. Witkowski

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GEO**DESI**



Take-home message

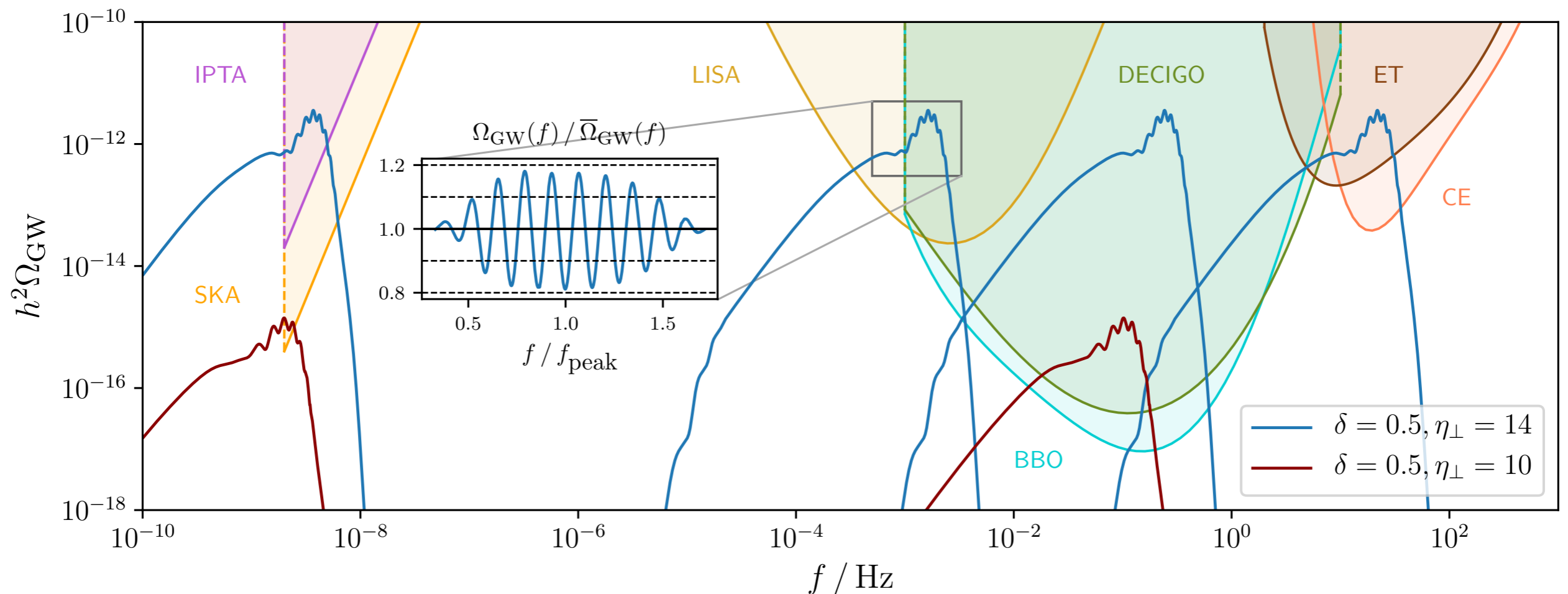
Primordial features invaluable probe of early universe

Oscillations in primordial scalar power spectrum



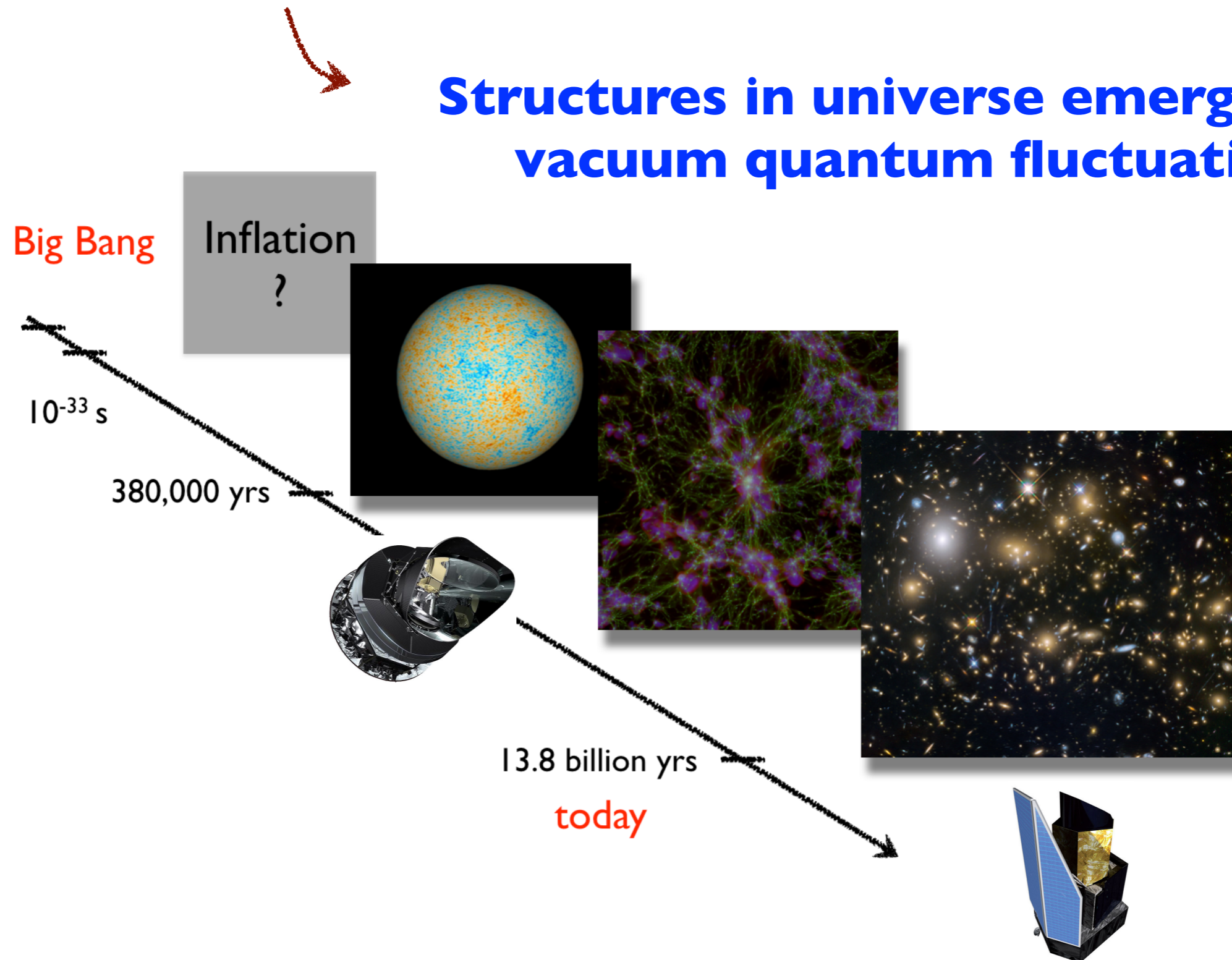
Oscillations in frequency profile of $\Omega_{\text{GW}}(f)$

Precious probe of inflation on small scales

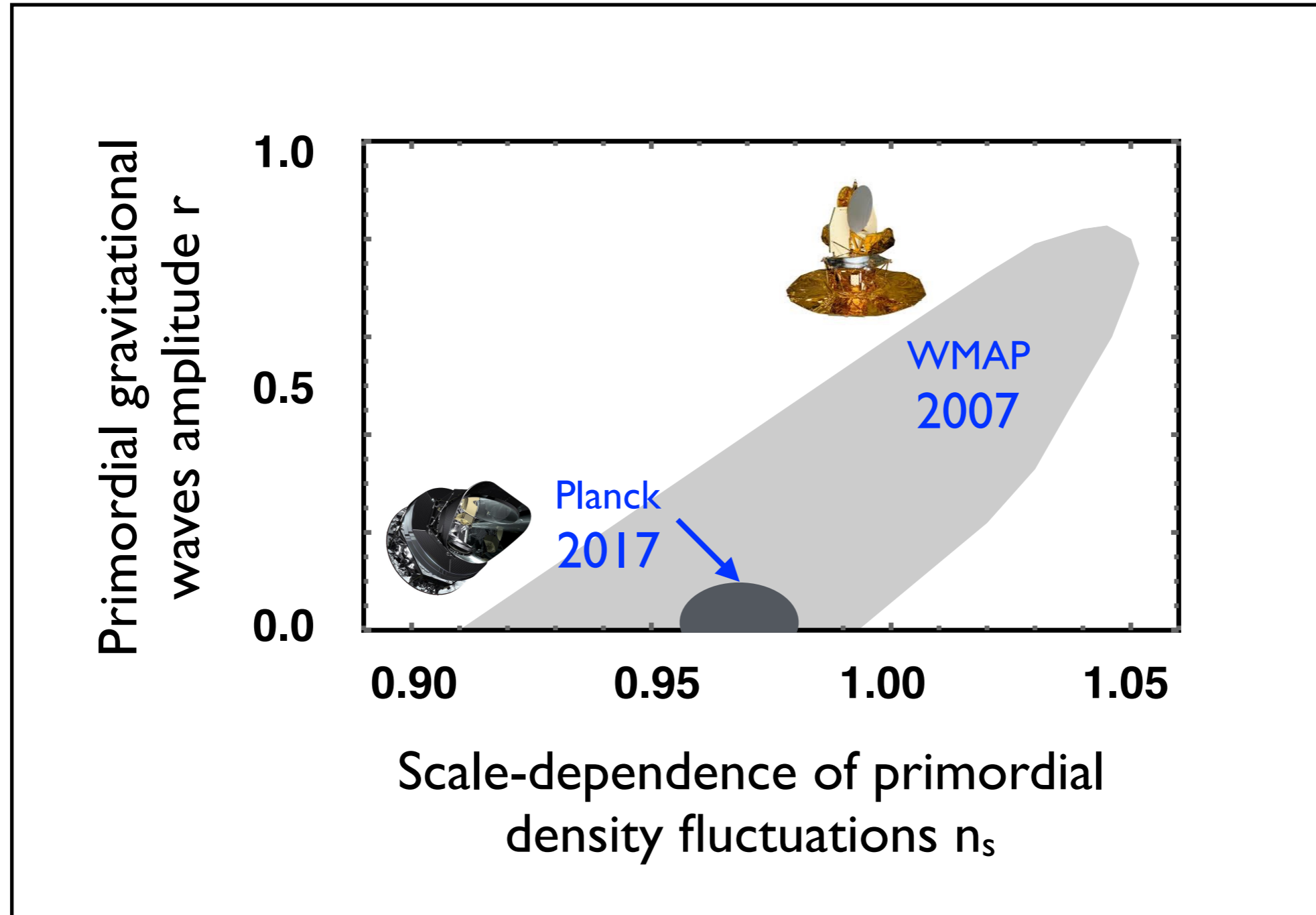


Inflation: a period of accelerated expansion before the radiation era that solves the problems of the Hot Big bang model

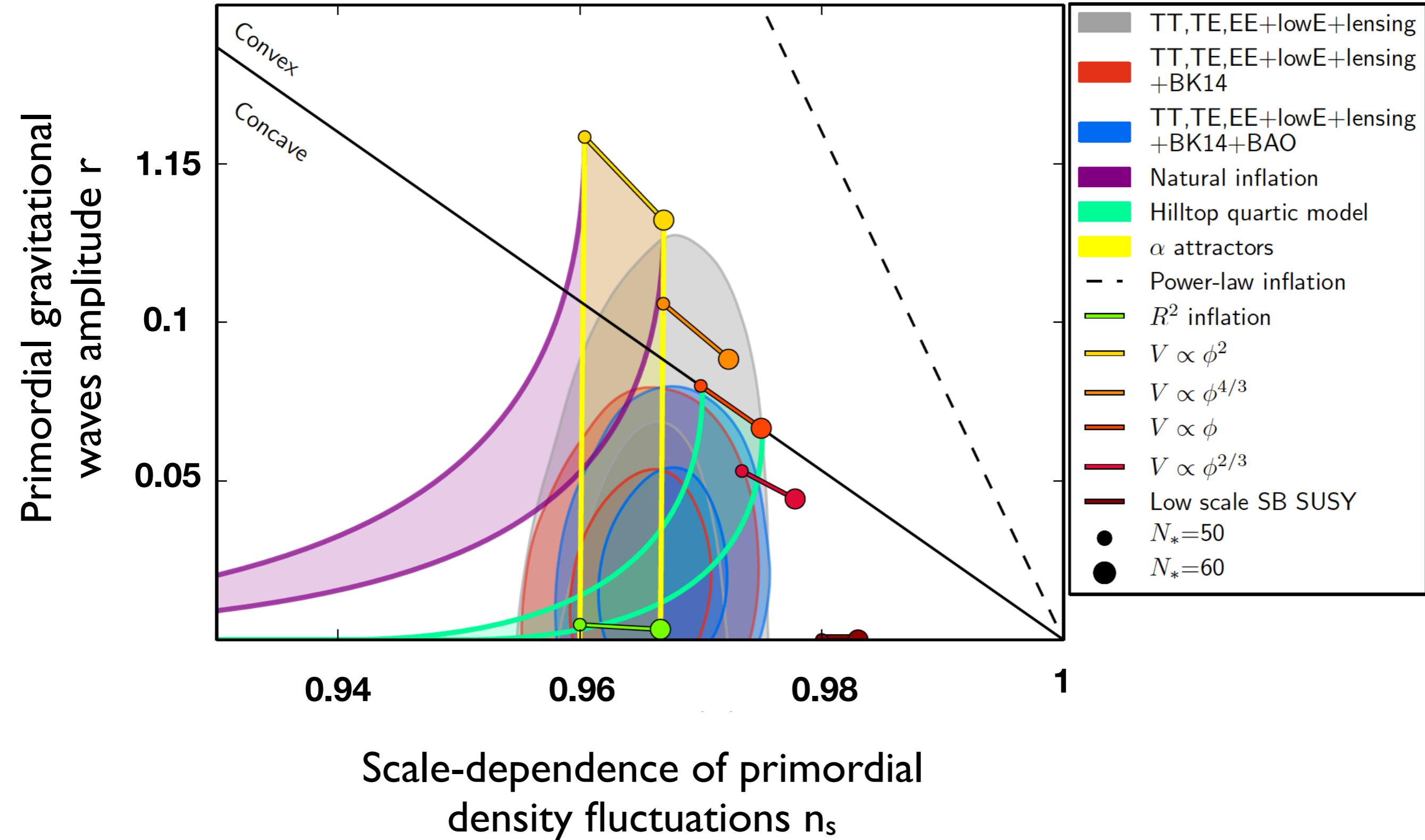
Structures in universe emerge from vacuum quantum fluctuations



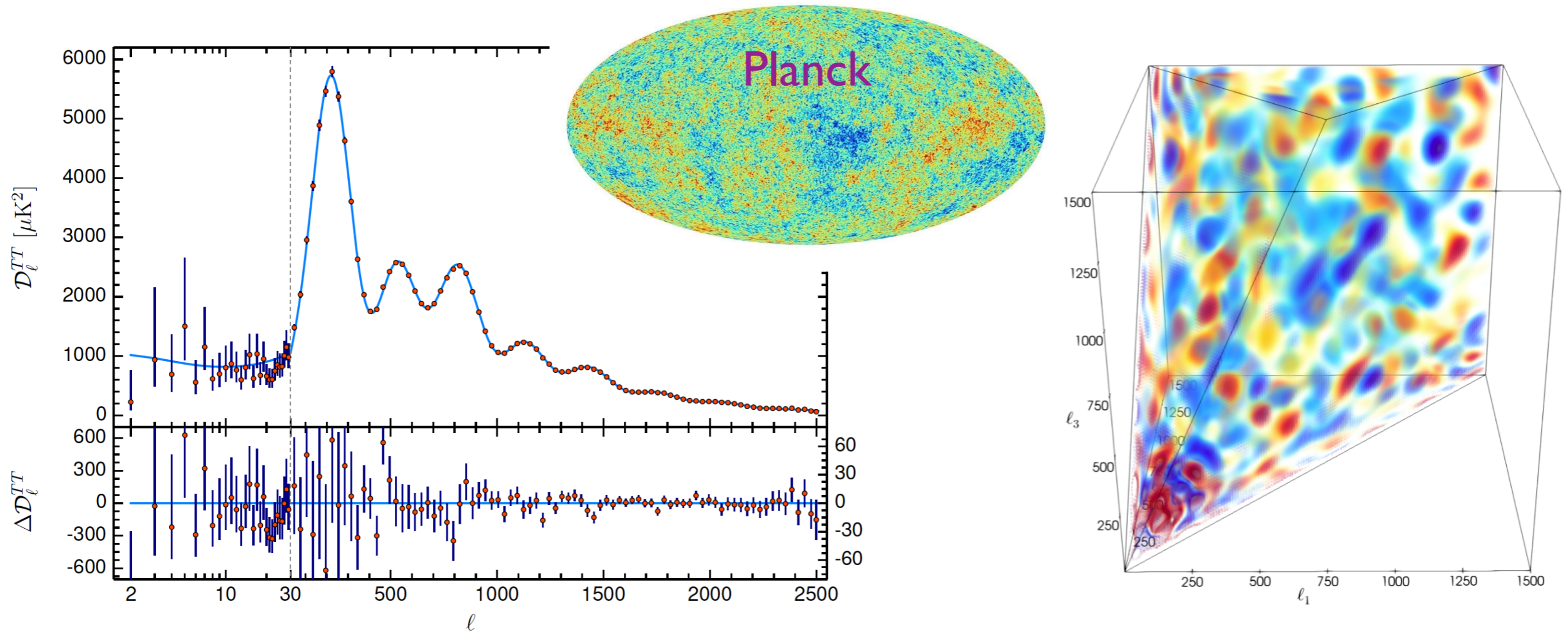
Observational progress



Observational progress



Observational progress



Density fluctuations:

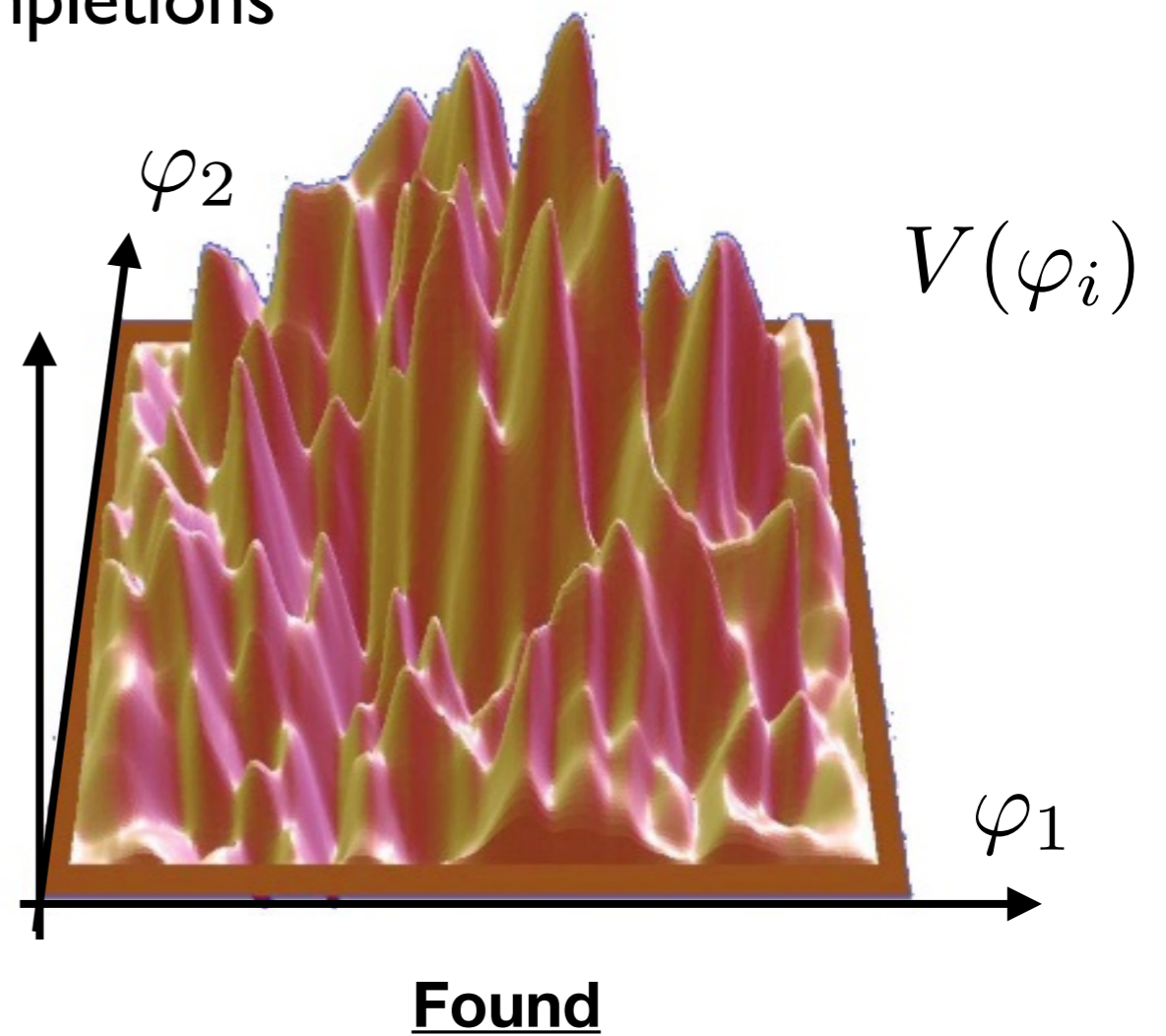
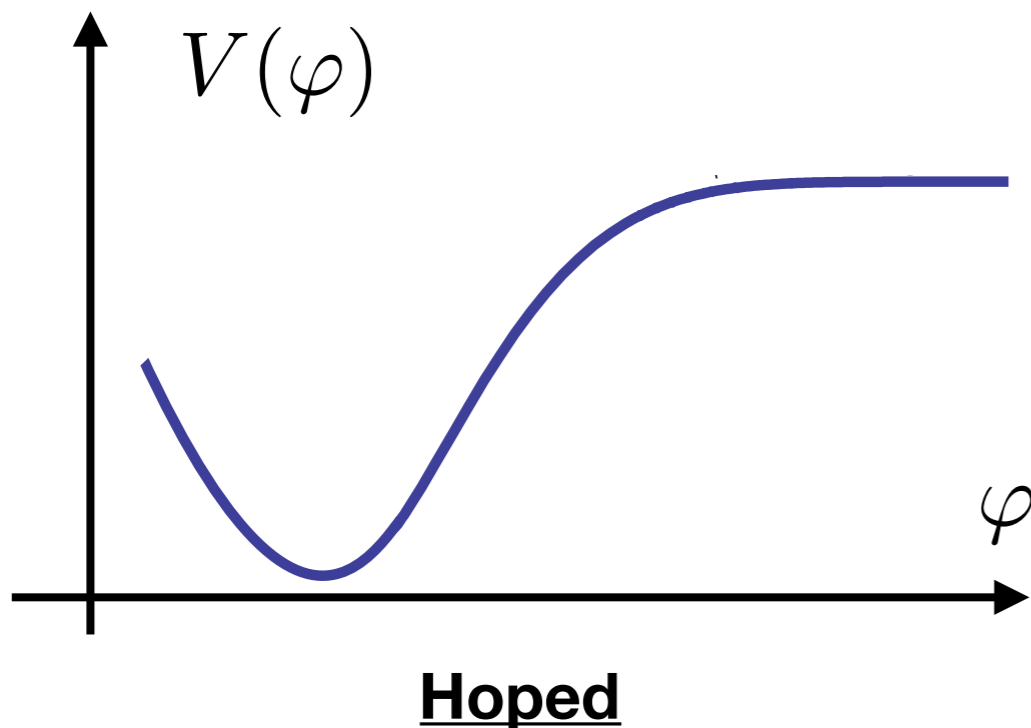
Superhorizon - adiabatic
almost scale-invariant - Gaussian

Simplest fit to data: single-field slow-roll inflation

Physics of inflation?

No-one believes **single-field slow-roll** models are more than **toy models**

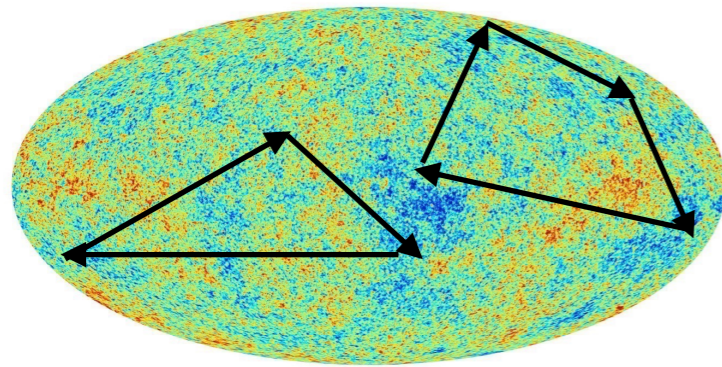
- decoupled from the rest of physics
- lack UV completions



Looking for new physics

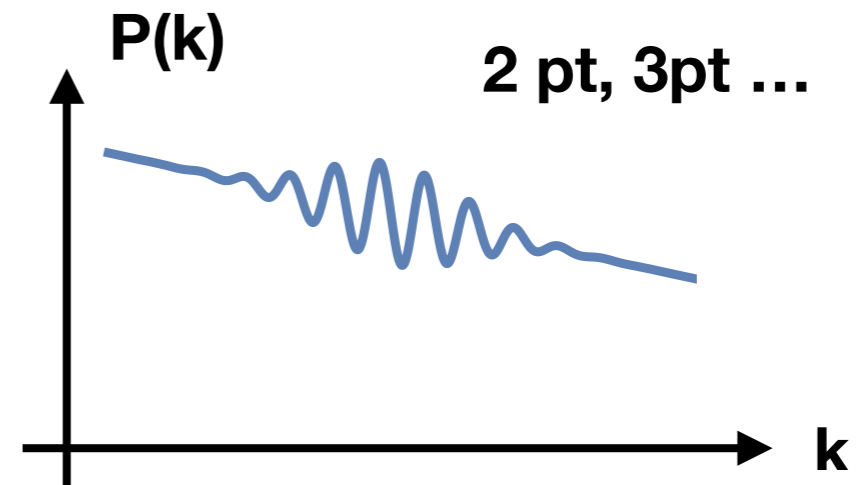
- Single-field slow-roll: at best emergent approximate description
- Cosmologists seek deviations to it in motivated manner

Primordial non-Gaussianities



3pt, 4 pt ...

Primordial features



Rule out all simplest models,
detection of heavy particles, etc

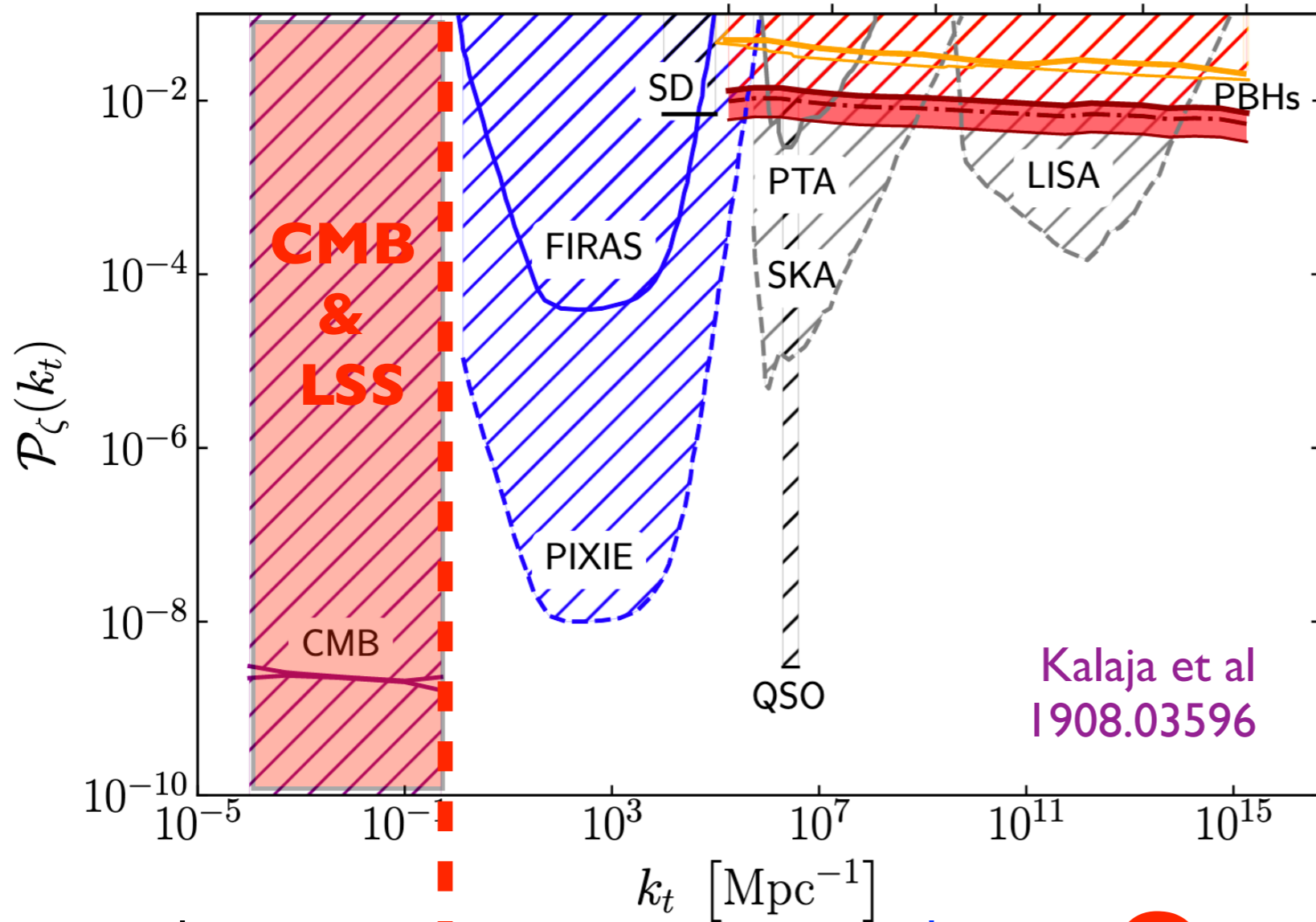
Studied for CMB, LSS, 21 cm



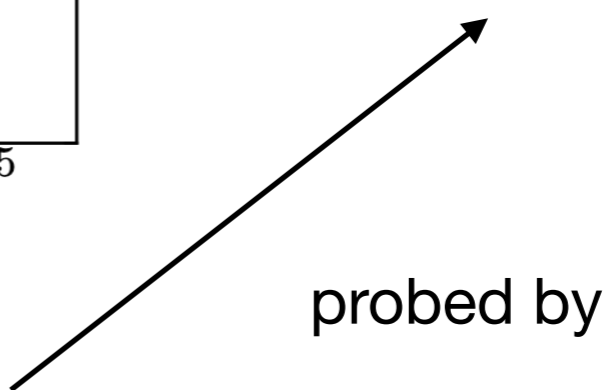
Euclid (202?)

SKA (2020 +)

Inflation on small scales?



Spectral distortions
 Primordial Black Holes
Stochastic GW background



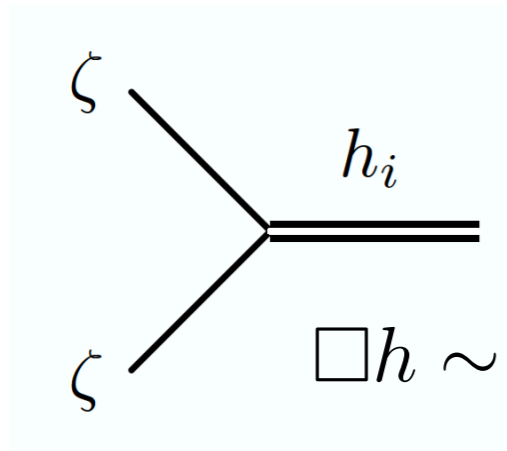
early stage
 of inflation
 \approx
Slow Roll
 +
Corrections

later
 stage **?**

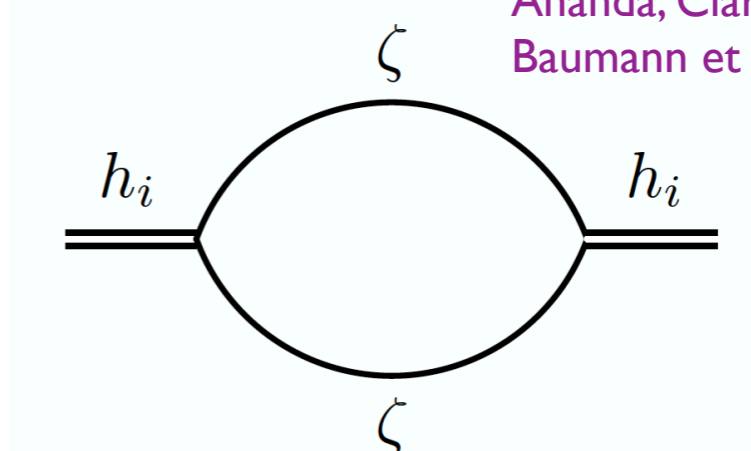
Drastically different?
**High-Energy effects
 to be discovered?**

Scalar-induced GWs

Acquaviva et al. '02
 Mollerach, Harari, Matarrese '03
 Ananda, Clarkson, Wands '06
 Baumann et al. '07 ...

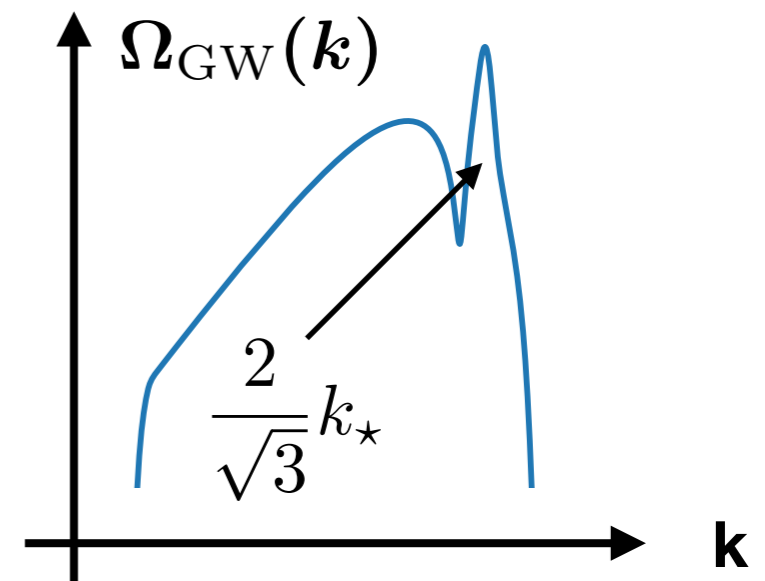
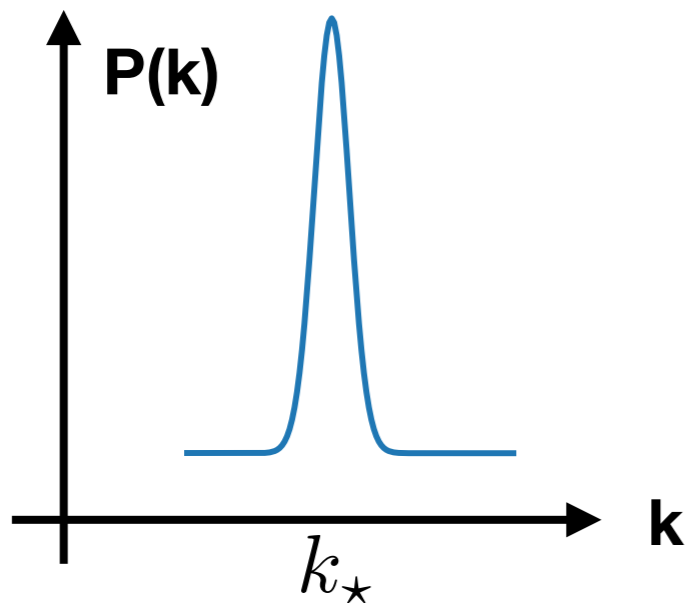


Enhanced $\delta\rho$



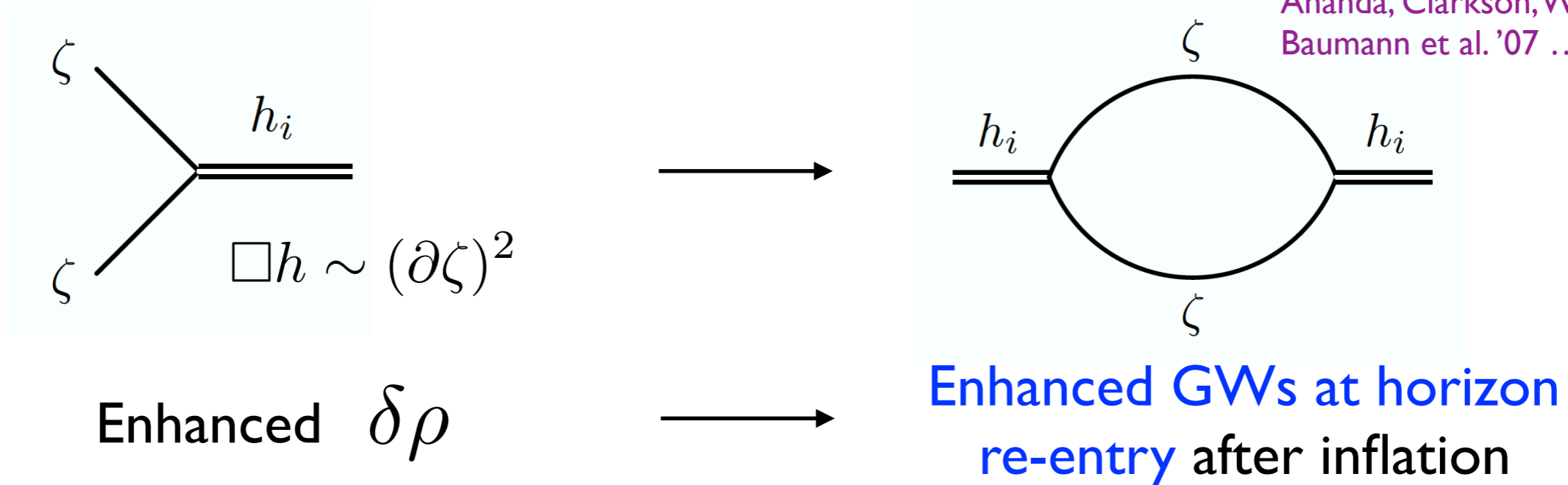
Enhanced GWs at horizon re-entry after inflation

$$\Omega_{\text{GW}}(k) = \int \int T(u, v) \mathcal{P}_\zeta(ku) \mathcal{P}_\zeta(kv) \sim 10^{-5} \mathcal{P}_\zeta^2$$



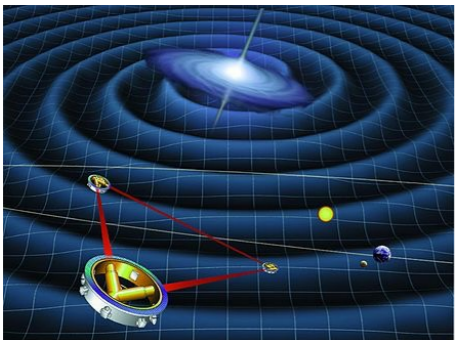
Scalar-induced GWs

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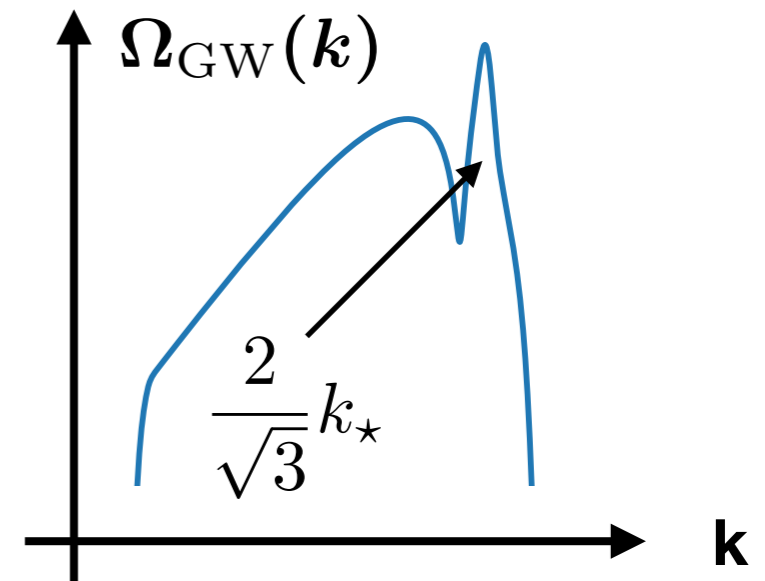


$$\Omega_{\text{GW}}(k) = \int \int T(u, v) \mathcal{P}_\zeta(ku) \mathcal{P}_\zeta(kv) \sim 10^{-5} \mathcal{P}_\zeta^2$$

$$\log\left(\frac{f}{10^{-3}\text{Hz}}\right) \simeq \log\left(\frac{k}{10^{12}\text{Mpc}^{-1}}\right) \simeq N_{\text{after CMB}} - 30$$



GW observatories probe inflation on small scales



Primordial features

Sharp features

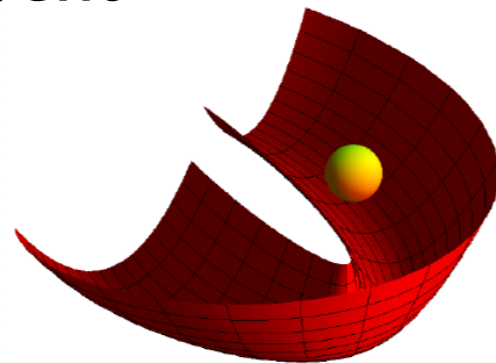
Resonant features

$$\frac{\mathcal{P}_\zeta(k)}{\mathcal{P}_{\text{env}}(k)} = \left[1 + A_{\text{lin}} \cos(\omega_{\text{lin}} k + \varphi_{\text{lin}}) \right]$$

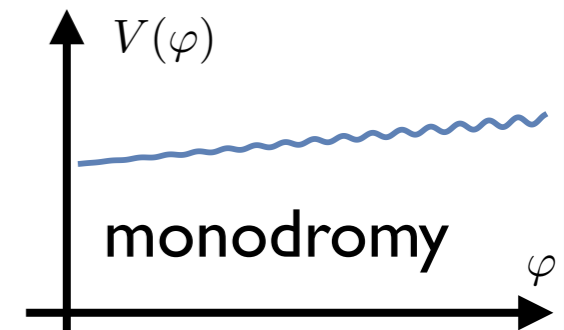
$$\left[1 + A_{\text{log}} \cos(\omega_{\text{log}} \log(k/k_\star) + \varphi_{\text{log}}) \right]$$

Localized event

step in potential
turn in field space ...



Resonance btw
background
oscillations and
quantum modes
oscillations



$$e^{-ik\tau_f} = e^{ik/k_f}$$

selection of
preferred time/scale

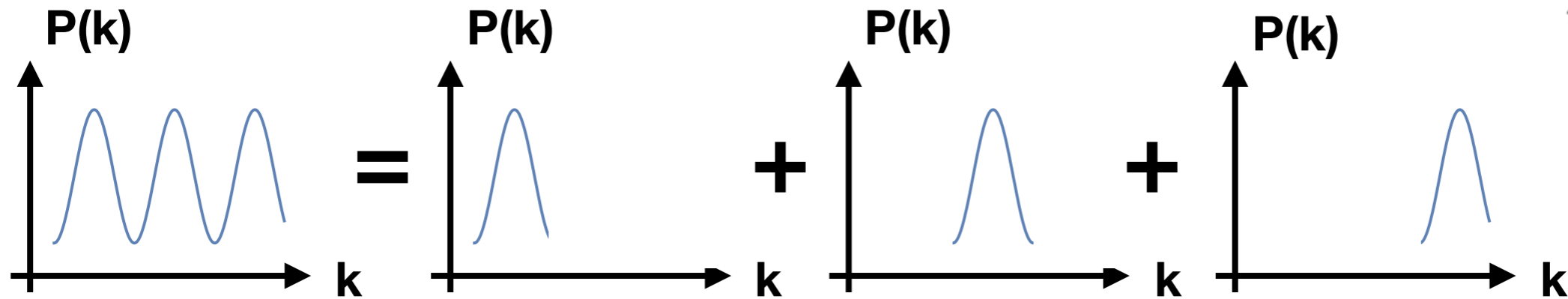
$$e^{iMt} = e^{i\frac{M}{H}N} \rightarrow e^{i\omega_{\text{log}} \log(k/k_\star)}$$

$$k \propto e^{N_k}$$

Intuition

1) Decompose as **sum of individual peaks**

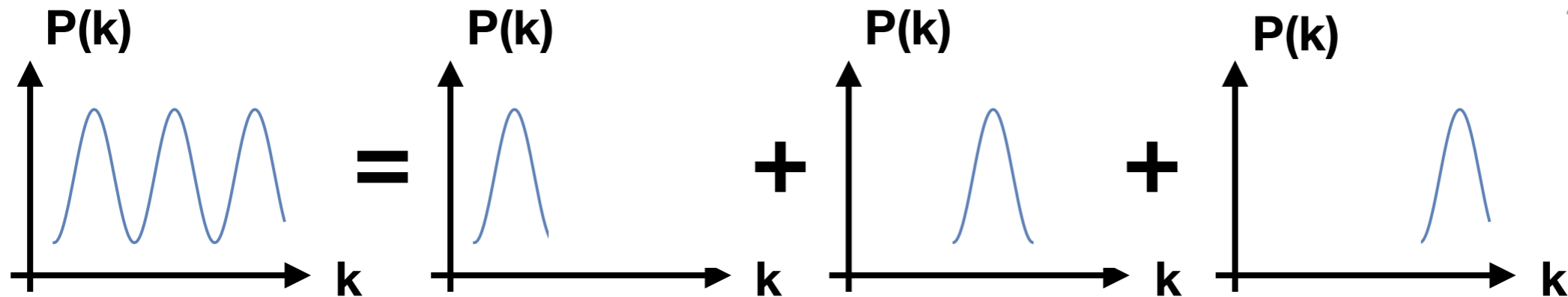
$$\mathcal{P}_\zeta(k) = \sum_{i=1}^{n_p} \mathcal{P}_{k_{*i}}(k)$$



Intuition

1) Decompose as **sum of individual peaks**

$$\mathcal{P}_\zeta(k) = \sum_{i=1}^{n_p} \mathcal{P}_{k_{\star i}}(k)$$



2) $\Omega_{\text{GW}} = \sum_{ij} \Omega_{\text{GW}}(\text{peak } i, \text{peak } j)$

Non-linear:
individual peaks
+ interactions between peaks

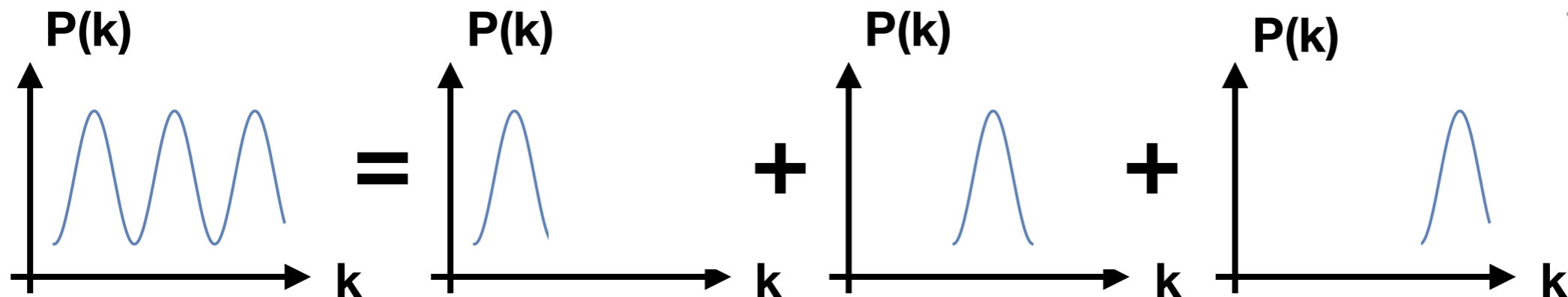
$$k_{\text{max},ij} = \frac{1}{\sqrt{3}}(k_{\star i} + k_{\star j}), \quad \text{with } k_{\text{max},ij} > |k_{\star i} - k_{\star j}|$$

Cai et al, 1901.10152

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Cai et al, 1901.10152

3)

Periodicity Δk

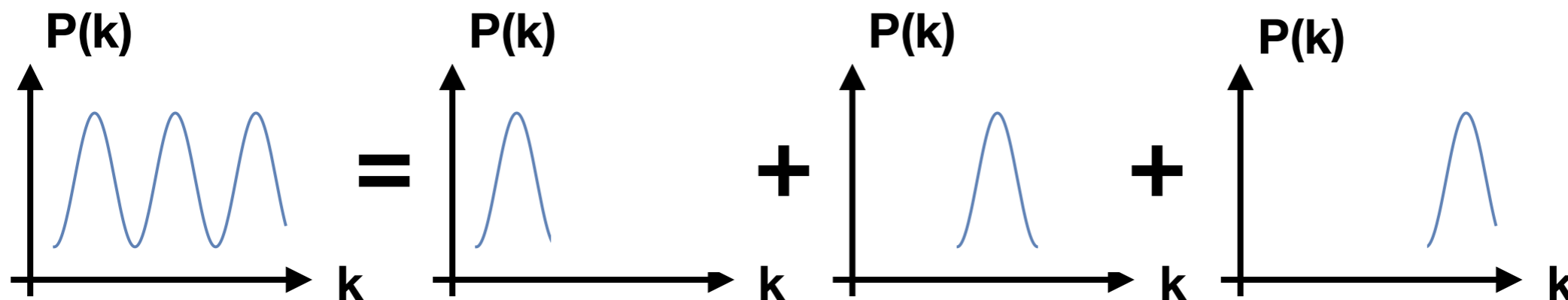


Periodicity $\frac{\Delta k}{\sqrt{3}}$

Intuition

1) Decompose as **sum of individual peaks**

$$\mathcal{P}_\zeta(k) = \sum_{i=1}^{n_p} \mathcal{P}_{k_{*i}}(k)$$



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Cai et al, 1901.10152

3 bis)

Periodicity $\Delta \log k \longrightarrow$ Periodicity $\Delta \log k$

in preparation




+ caveat

SGWB signature of sharp features

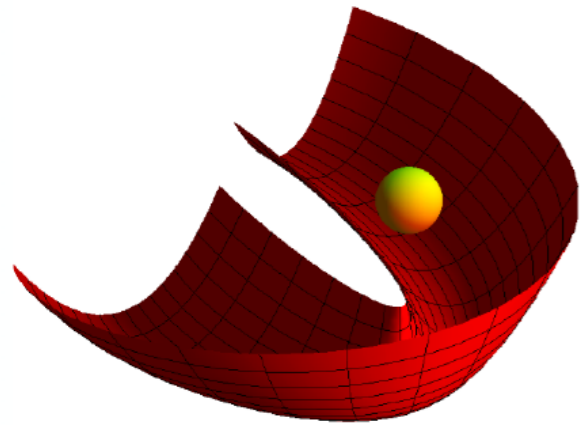
$$\mathcal{P}_\zeta(k) = \mathcal{P}_{\text{env}}(k) \times \left[1 + A_{\text{lin}} \cos(\omega_{\text{lin}} k + \varphi_{\text{lin}}) \right]$$

$$\omega_{\text{lin}}^{\text{GW}} = \sqrt{3} \omega_{\text{lin}}$$

$$\Omega_{\text{GW}}(k) \sim \bar{\Omega}_{\text{GW}}(k) \left[1 + A \cos(\omega_{\text{lin}}^{\text{GW}} k + \phi) \right]$$

-  Overall shape determined by envelope of power spectrum
-  Periodic structure in k \longrightarrow Periodic structure in k
-  Averaging-out effect: e.g. from 100% modulation to 10%

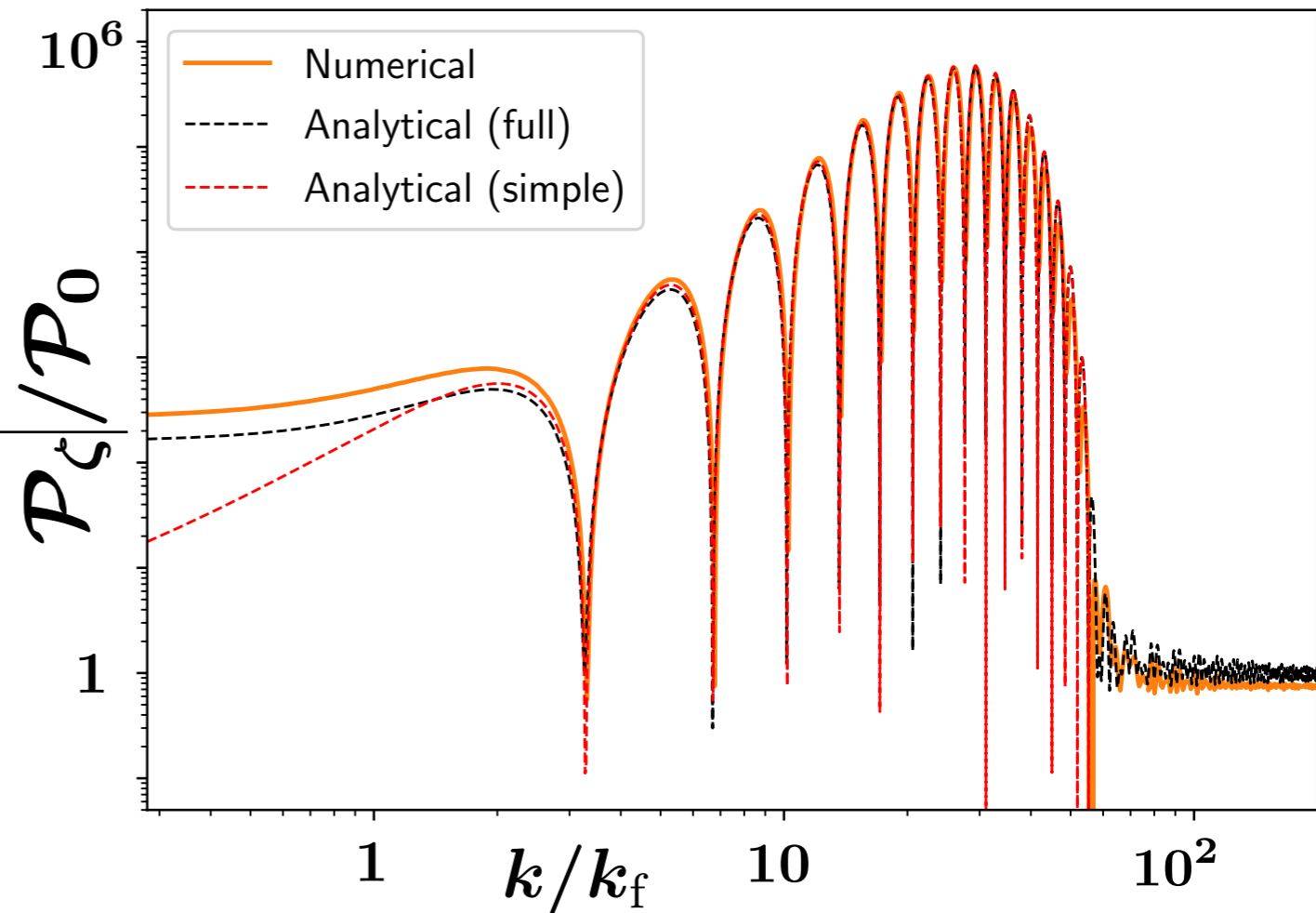
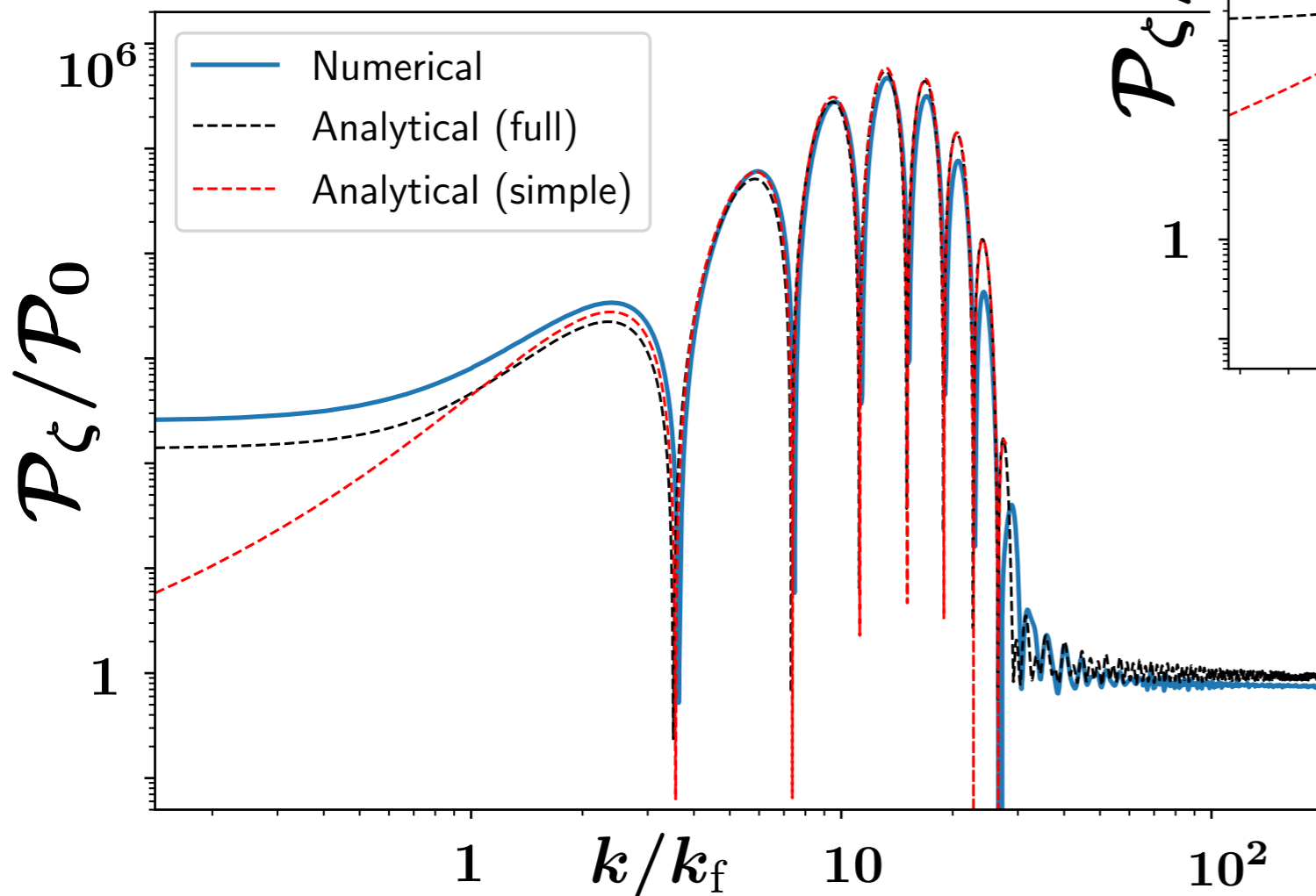
An explicit example



Fumagalli et al
2004.08369, 2012.02761
Palma et al 2004.06106

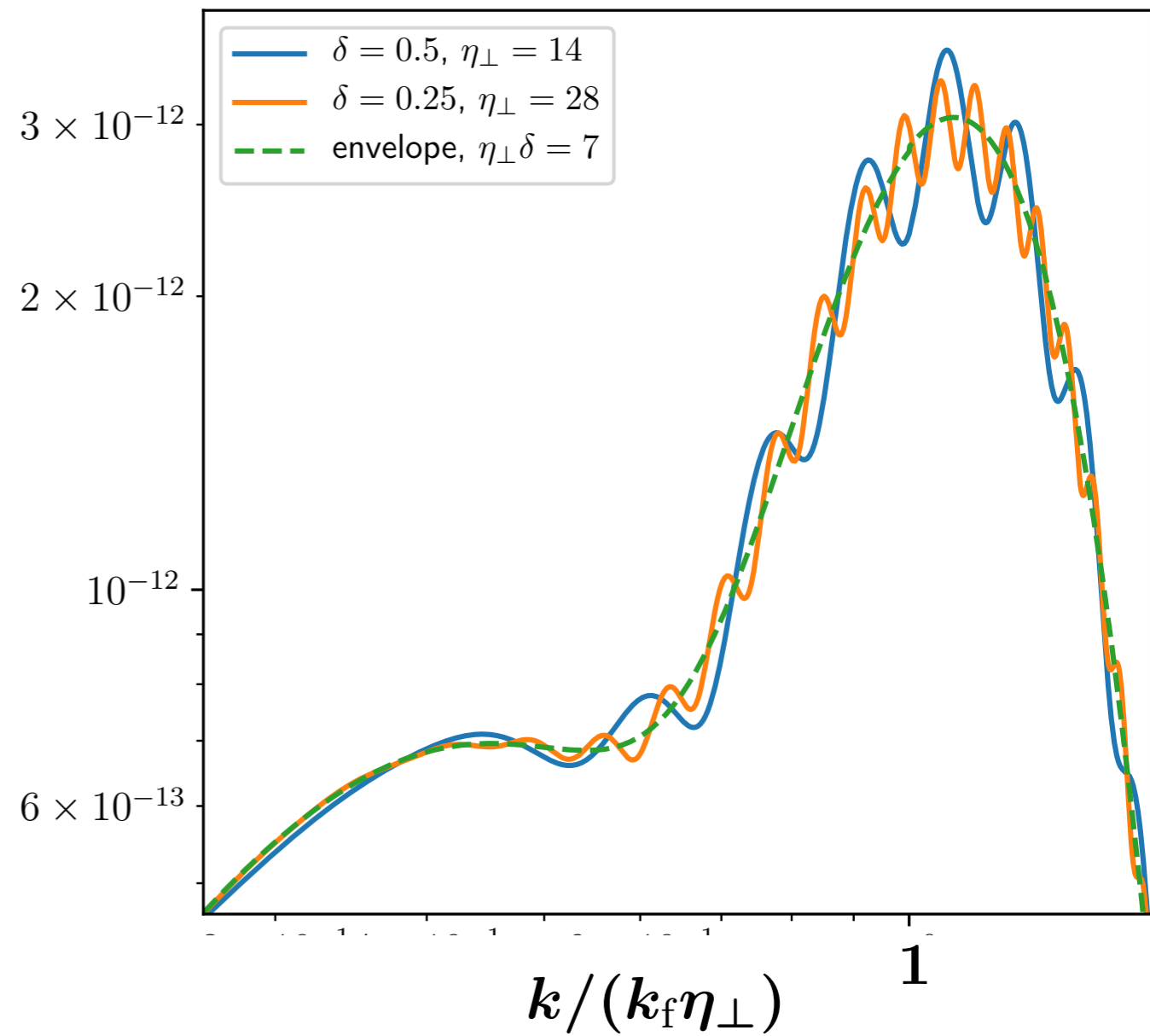
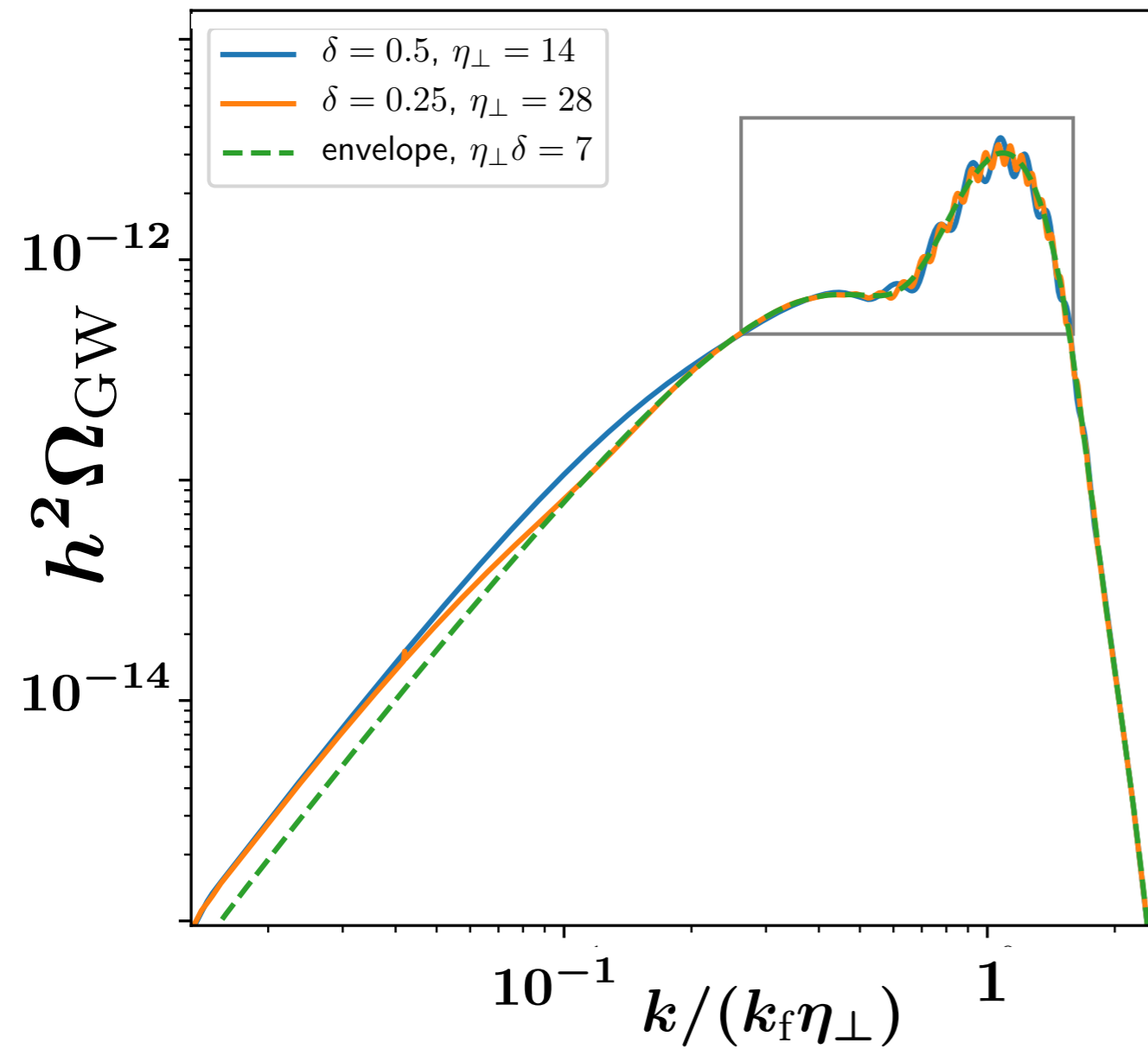
$\delta = 0.25, \eta_{\perp} = 28$

$\delta = 0.5, \eta_{\perp} = 14$



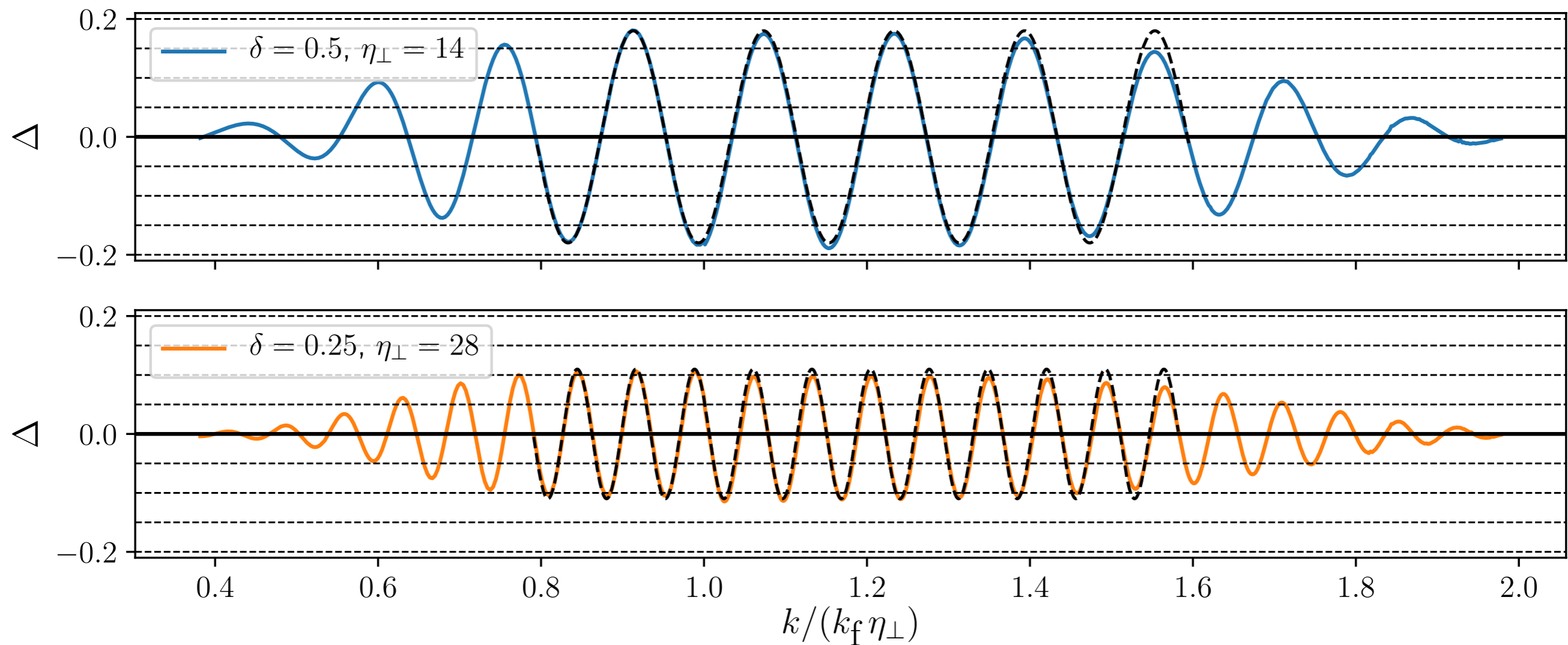
δ duration of turn in e-fold
 $\eta_{\perp} = \frac{\Omega}{H}$ importance of turn

and its SGWB signature



Feature template

$$\Delta(k) \equiv \frac{\Omega_{\text{GW}}(k) - \bar{\Omega}_{\text{GW}}(k)}{\bar{\Omega}_{\text{GW}}(k)} = A \cos(\omega_{\text{lin}}^{\text{GW}} k + \phi)$$

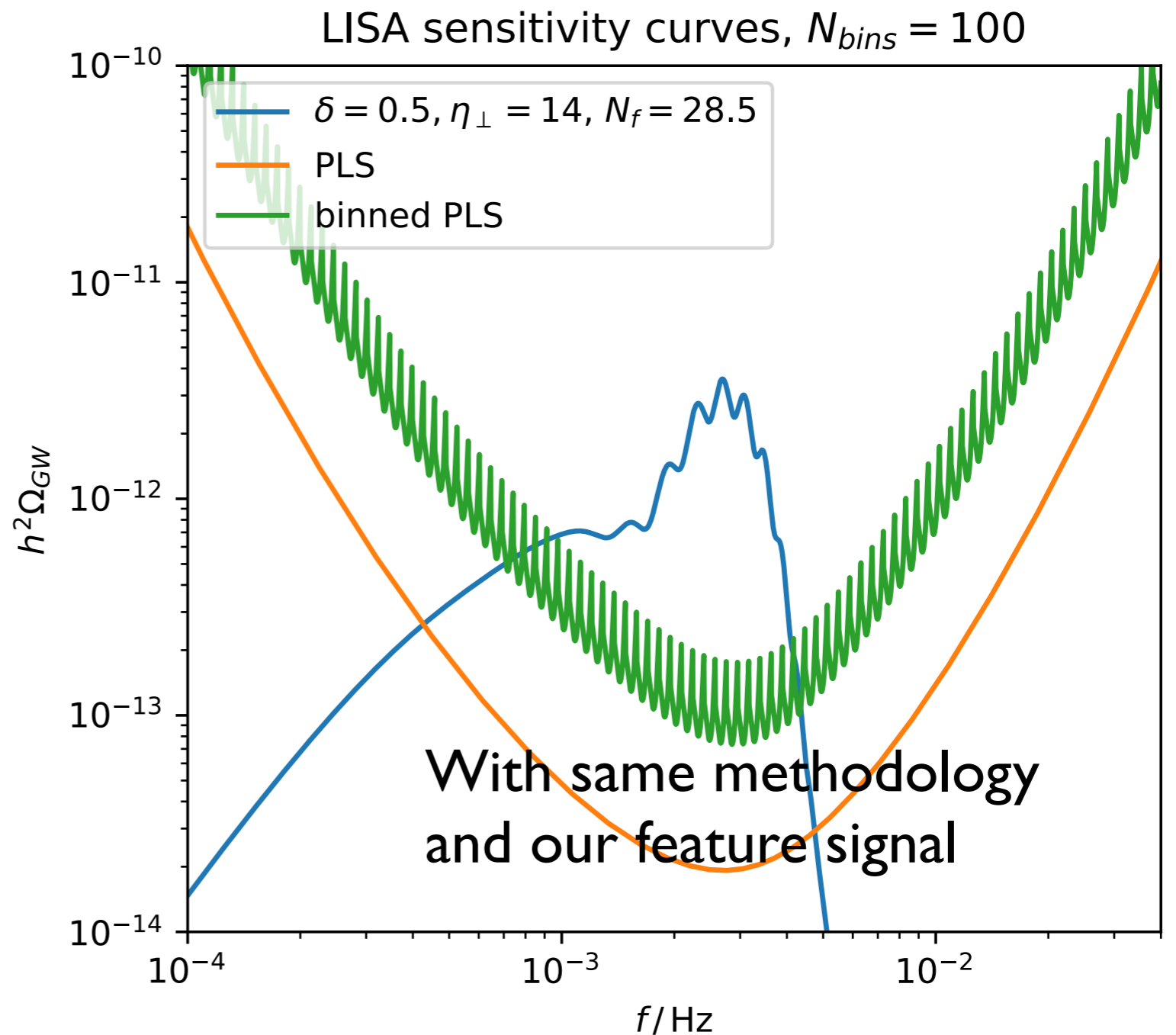


Detectability

Reconstruction of spectral shape with LISA

Caprini et al, LISA CosWG
1906.09244

A wiggly signal can be reconstructed (without particular motivation there)



Oscillations resolvable by LISA T=3 years

see also Braglia et al 2012.05821

Theoretical consistency

Enhanced $\zeta \simeq \frac{\delta\rho}{\rho}$

model
independent

Scalar-induced GWs
after inflation

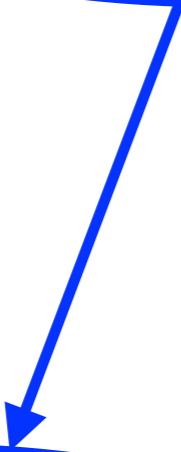
Theoretical consistency

Physical mechanism

Enhanced $\zeta \simeq \frac{\delta\rho}{\rho}$



mechanism
dependent



model
independent



Scalar-induced GWs
during inflation

Scalar-induced GWs
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in preparation

Theoretical consistency

Physical mechanism

Enhanced $\zeta \simeq \frac{\delta\rho}{\rho}$



Backreaction
Perturbativity

mechanism
dependent

model
independent

Scalar-induced GWs
during inflation

Scalar-induced GWs
after inflation

in preparation

Take-home message

Primordial features invaluable probe of early universe

Oscillations in primordial scalar power spectrum



Oscillations in frequency profile of $\Omega_{\text{GW}}(f)$

Precious probe of inflation on small scales

