

# Cryogenic Detectors for CEvNS and DM: Ideas and their Application

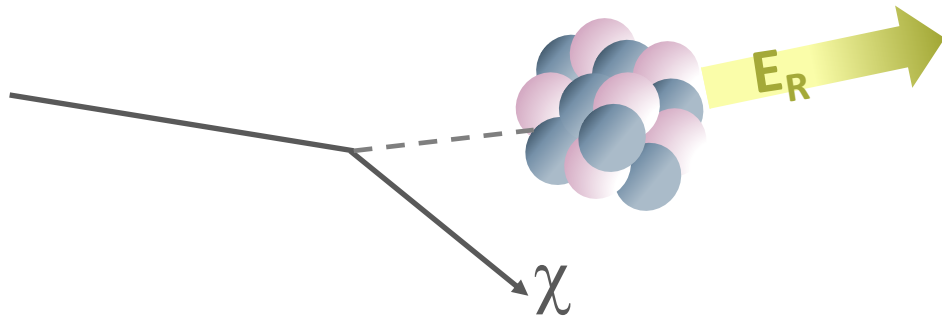


**V. Wagner**  
Technische Universität München  
BSM-nu Workshop,  
11. – 12.02.2012



# Direct Dark Matter Search

- For generic DM models: dark matter,  $\chi$ , interacts coherently with nucleus
- Experimental signature = nuclear recoil



- Featureless, smoothly decreasing spectrum

$$\frac{dR}{dE_R} = \frac{R_0}{E_0\mu} F^2(q) e^{-\frac{E_R}{E_0\mu}}$$

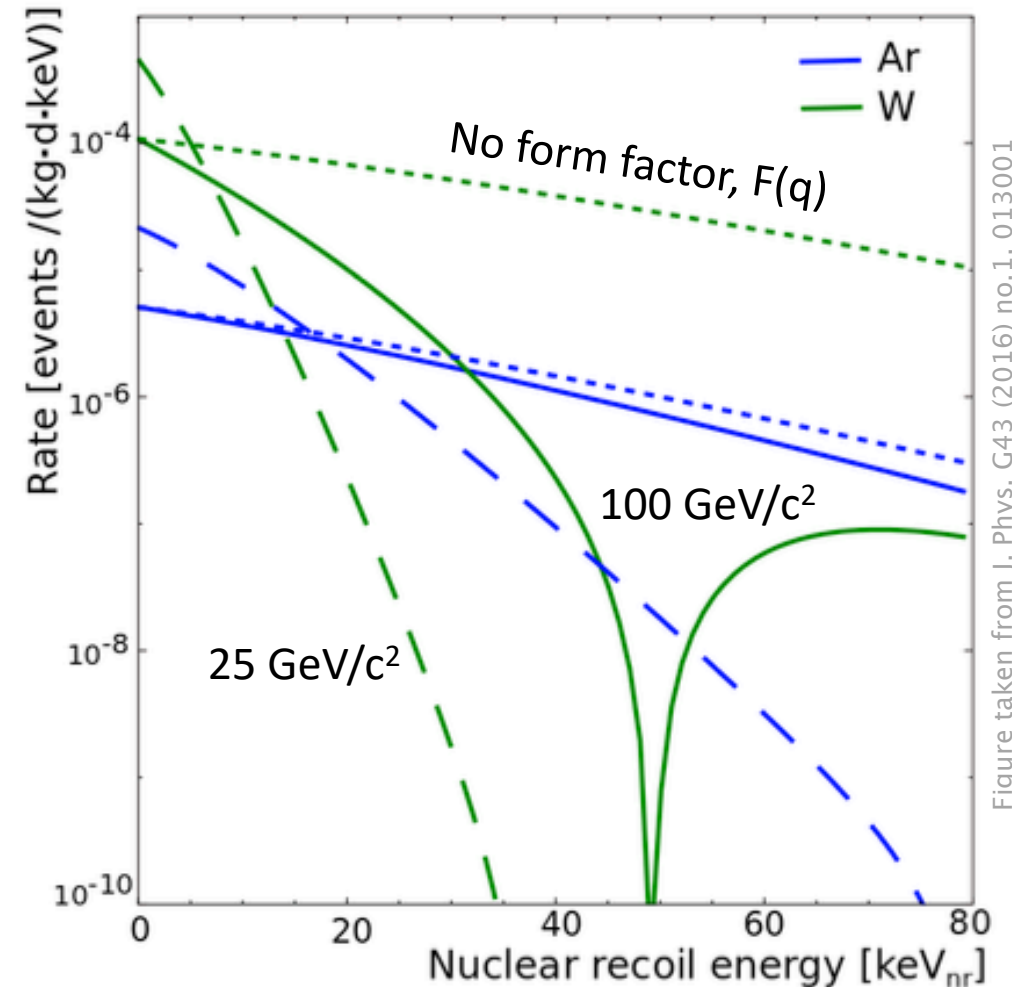
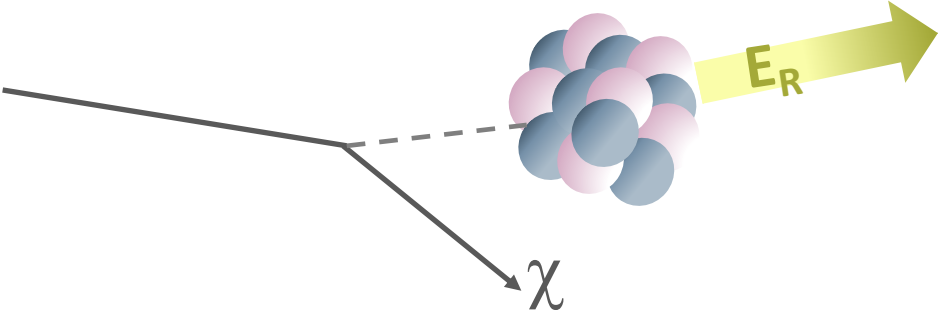


Figure taken from J. Phys. G43 (2016) no.1, 013001

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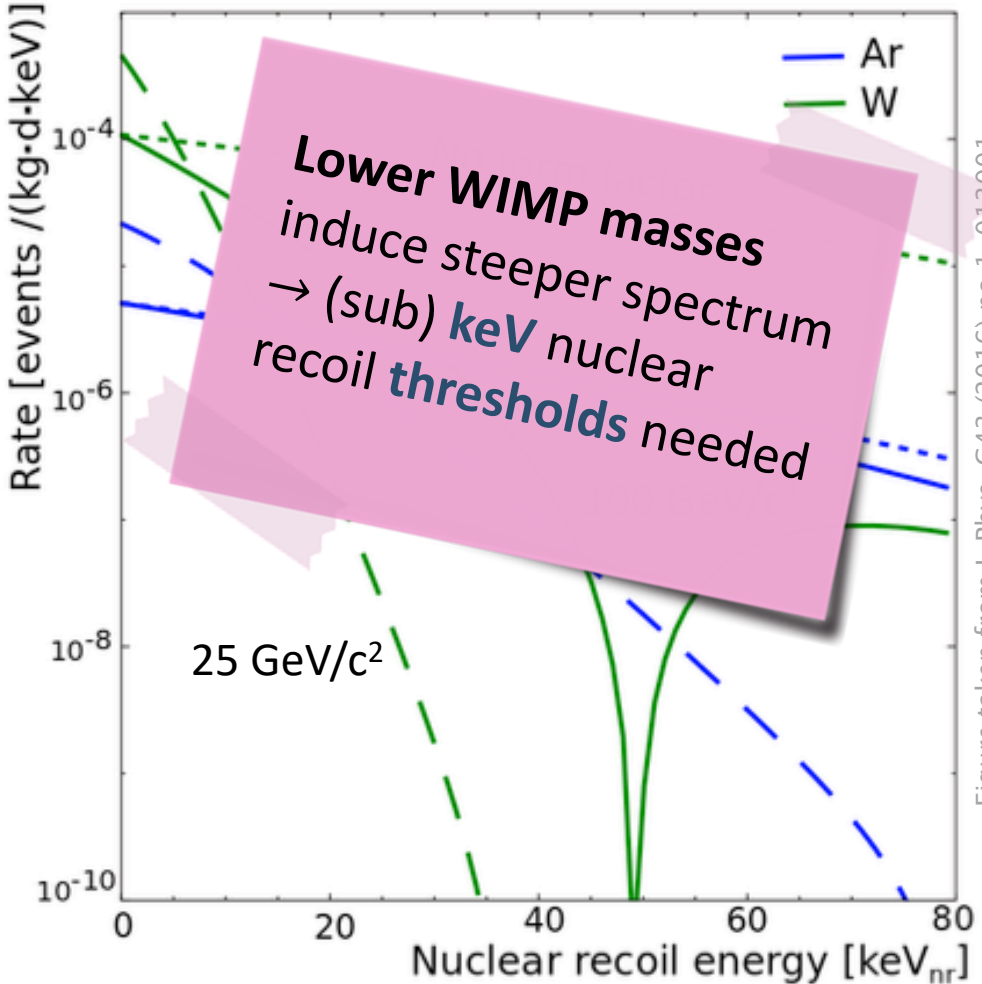
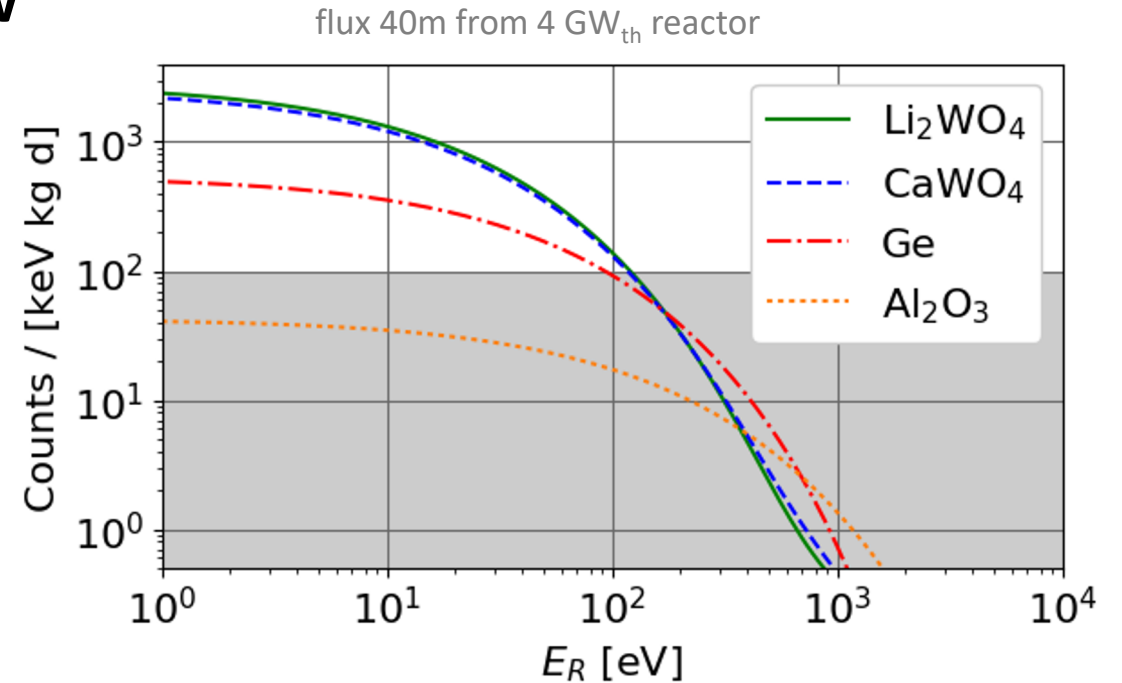
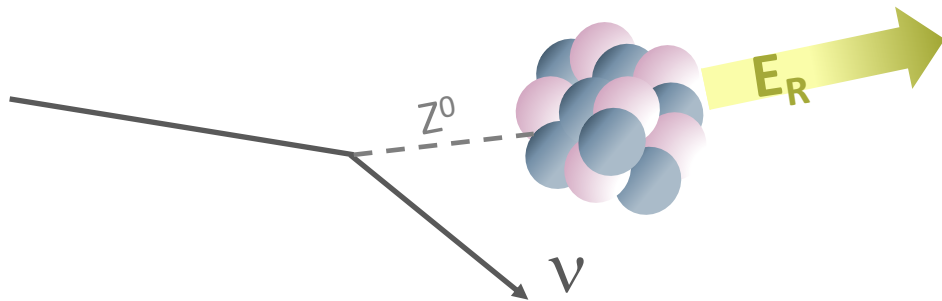


Figure taken from J. Phys. G43 (2016) no.1, 013001

# Coherent Elastic n-Nucleus Scattering (CEvNS)

- Probe of SM prediction & **search for new physics**
- Neutrino,  $\nu$ , scattering off elastically a nucleus
- Experimental signature = nuclear recoil



- Coherent typically for  $E_\nu < 30$  MeV
- Cross-section approximated by

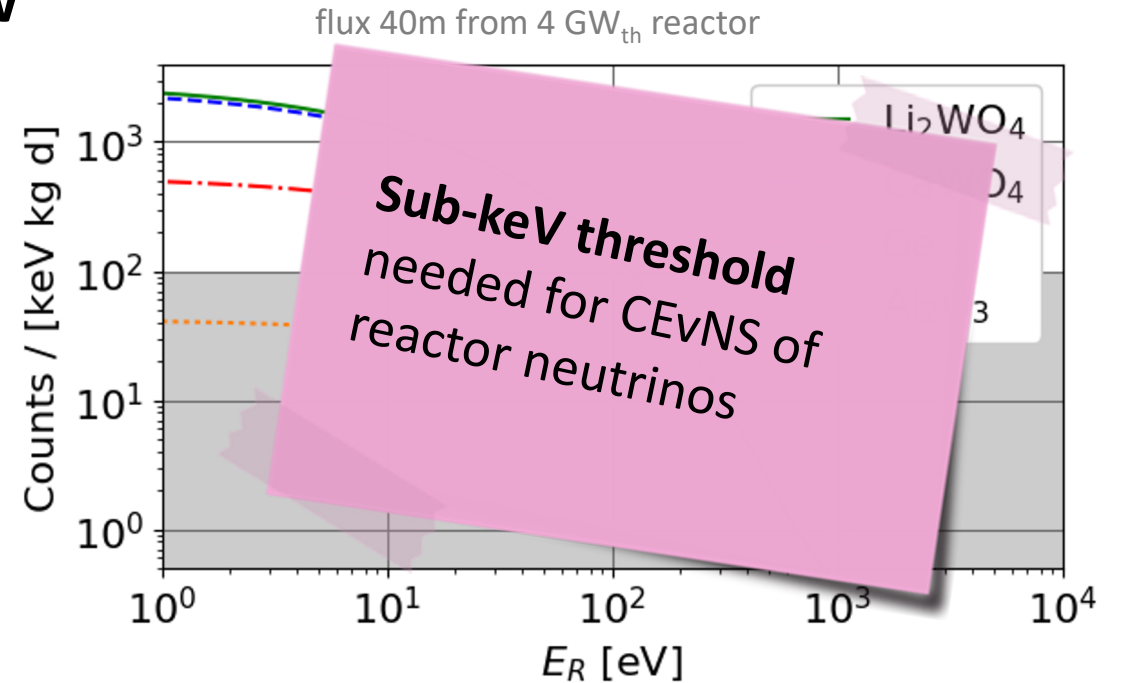
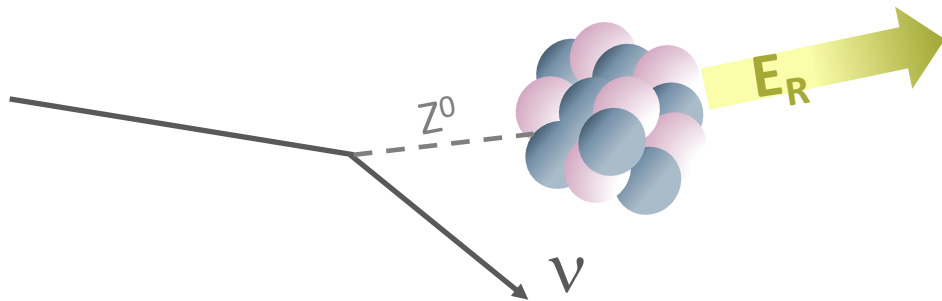
$$\sigma \approx \frac{G_F^2 N^2}{4\pi} E_\nu^2$$

$$E_{\text{nr}}^{\text{max}} = \frac{2E_\nu^2}{m_n + 2E_\nu}$$



# Coherent Elastic n-Nucleus Scattering (CEvNS)

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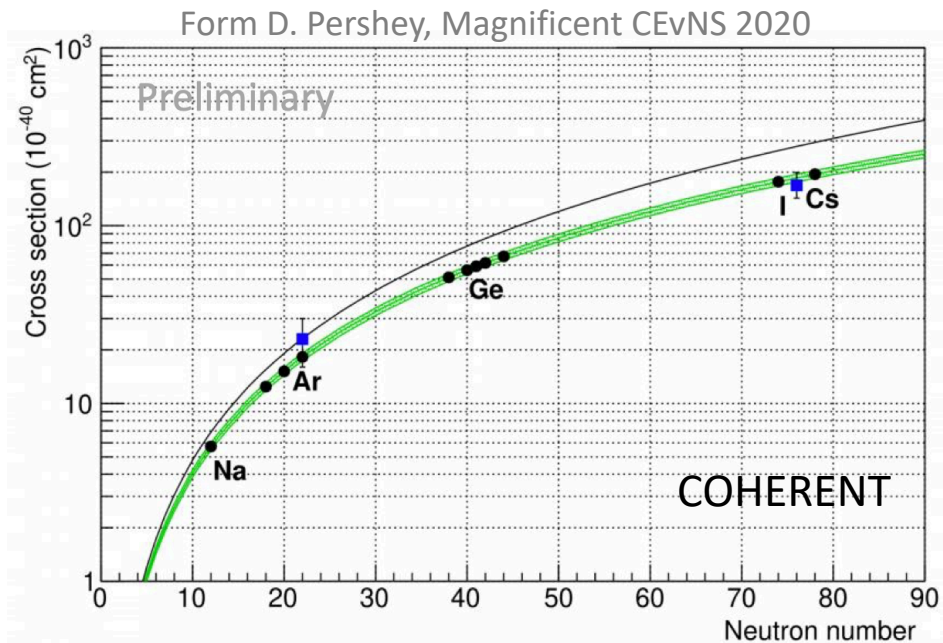
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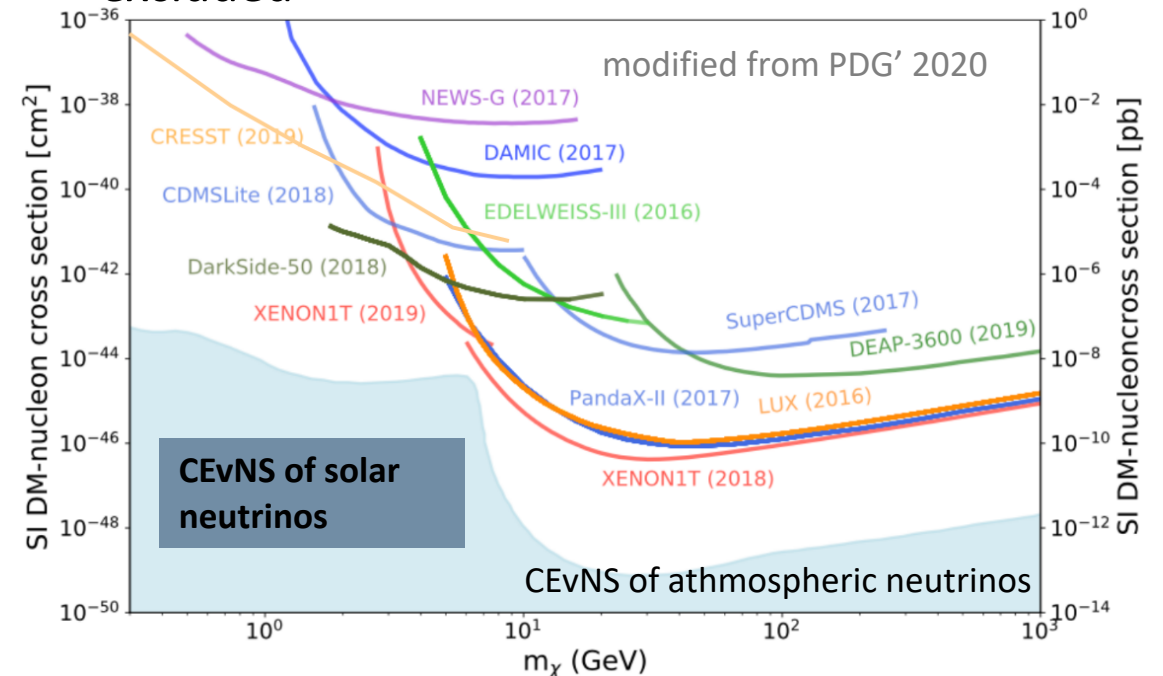
# Status CEvNS & Direct Dark Matter Search

- First observation at  $\pi$ -DAR source with multiple detector types and keV thresholds



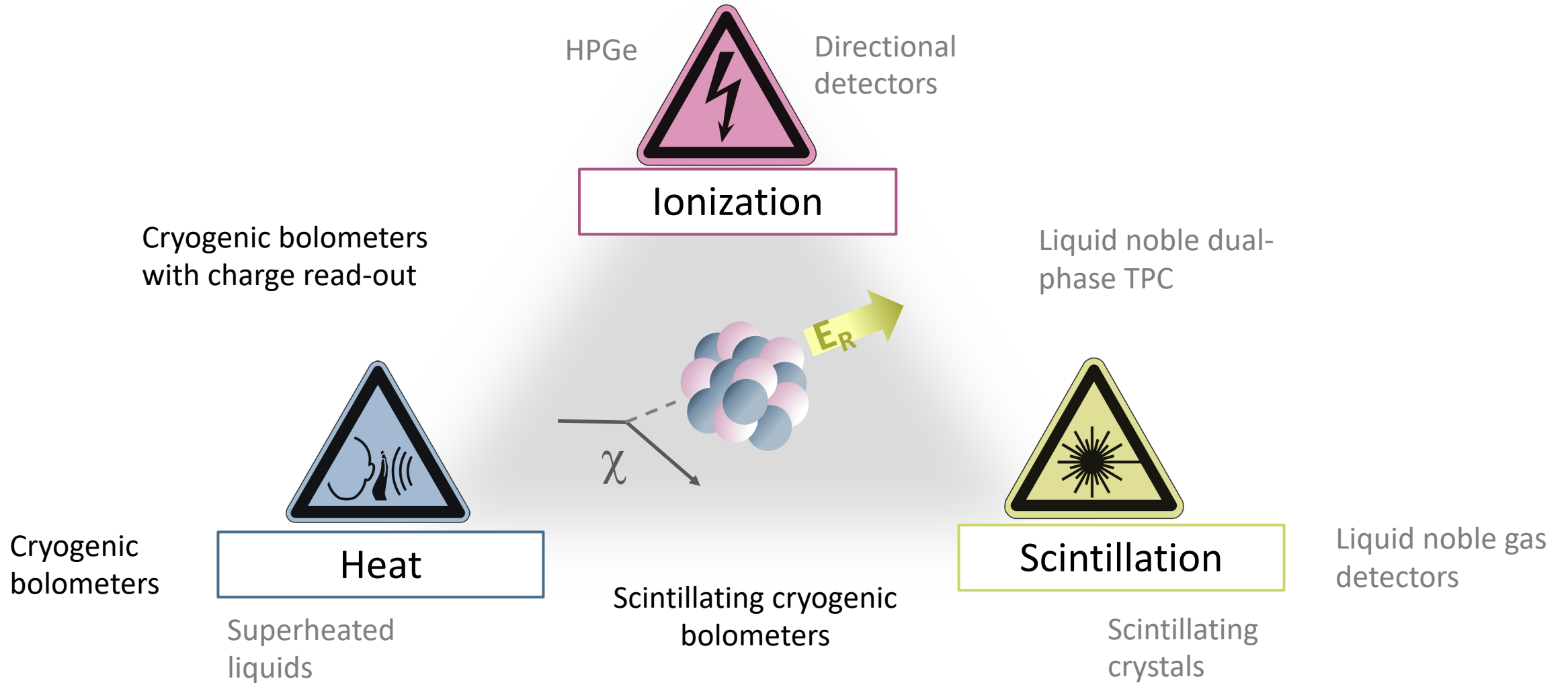
- CEvNS of reactor neutrinos unobserved so far
- Need of sub keV threshold

- Much of the WIMP parameter space already excluded



- To explore  $m_\chi < \text{GeV}/c^2$  need to lower threshold

# Observables



# Principle of Bolometer

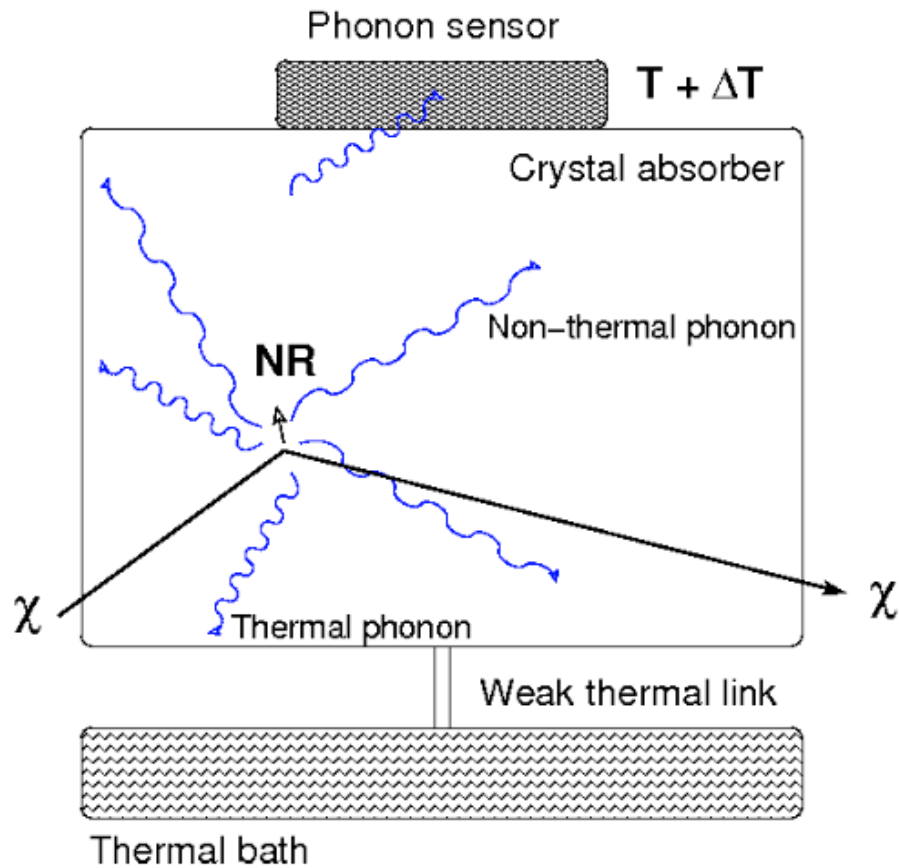
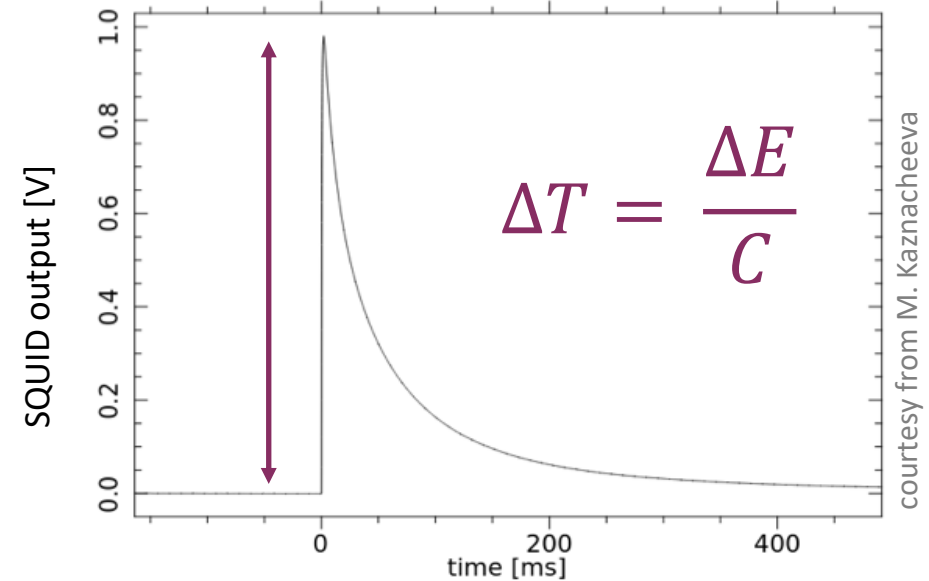


Figure taken from J. Phys. G43 (2016) no.1, 013001

Heat pulse measured with W-TES

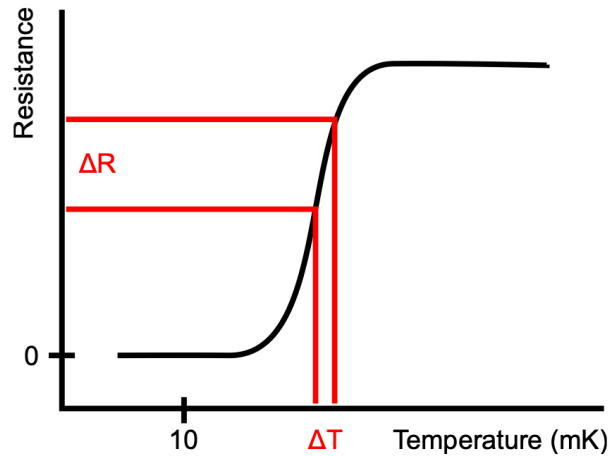


- ✓ Measurement of full recoil energy
- ✓ Achieve very low thresholds & excellent energy resolution
- ✓ Wide range of target materials
- ✗ Operation at (10–100) mK
- ✗ Limited scalability of detectors

Synergy of CEvNS and dark matter

# Thermometers

## Transition Edge Sensor (TES)

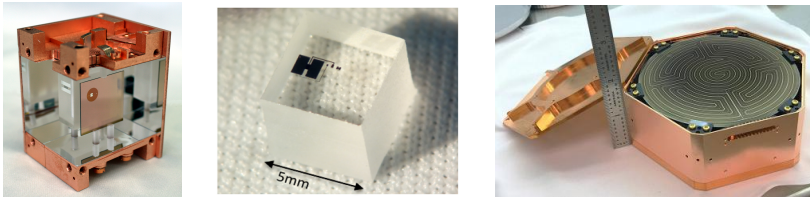


$$\Delta T = O(100 \mu\text{K})$$

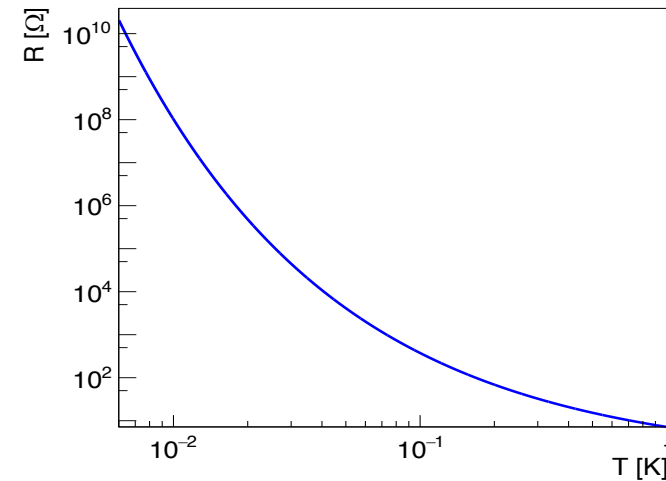
↓

$$\Delta R = O(100 \text{ m}\Omega)$$

- SQUID read-out
- Fast response  $O(\mu\text{s})$
- Narrow temperature range



## Neutron Transmutation Doped (NTD) Ge

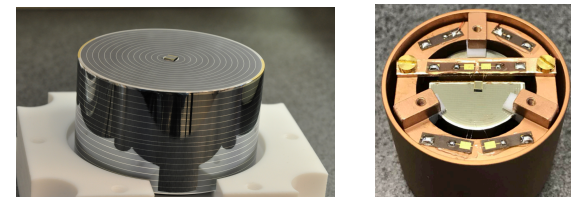


$$\Delta T = O(100 \mu\text{K})$$

↓

$$\Delta R = O(8 \text{ M}\Omega)$$

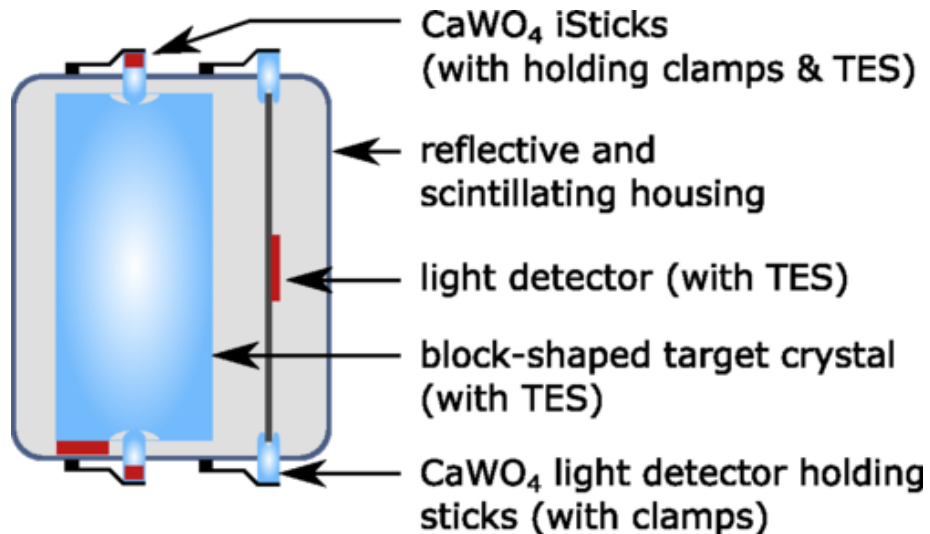
- Voltage amplifier read-out
- Slow response  $O(\text{ms})$
- Wider temperature range



# CRESST-III

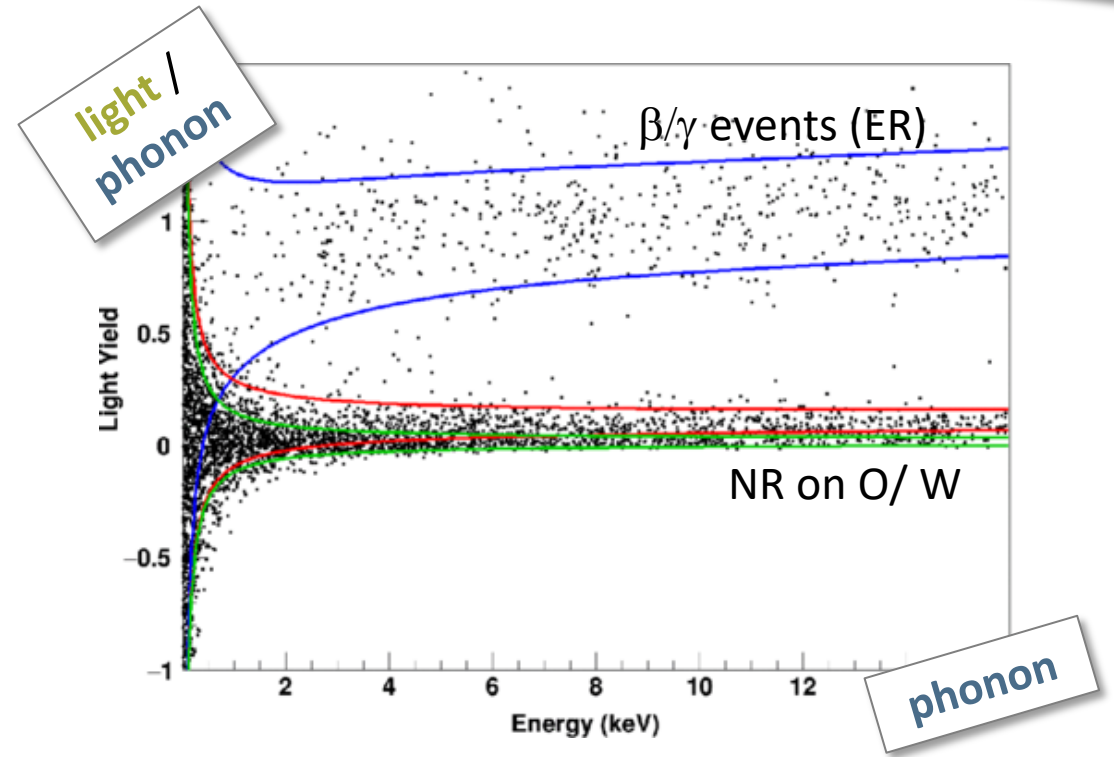
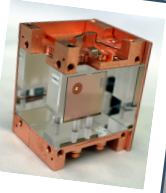
CRESST, PhysRevD.100.102002

- Optimized for low-mass dark matter
- Scintillating  $\text{CaWO}_4$  target crystals
- Dimensions:  $(20 \times 20 \times 10) \text{mm}^3 \approx 24 \text{g}$
- Achieved threshold:  $30.1 \text{ eV}_{\text{nr}}$



CRESST-III

Target:  $\text{CaWO}_4$   
Mass: 10x 24 g, 3x 220g  
Location: LNGS  
Status: running





# Towards 10 eV<sub>nr</sub> Thresholds

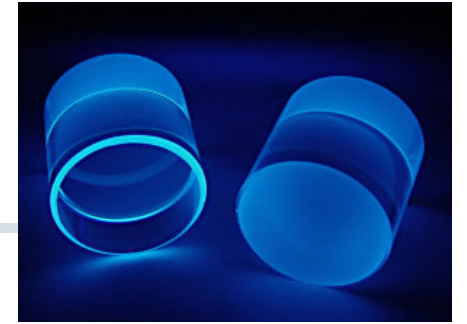
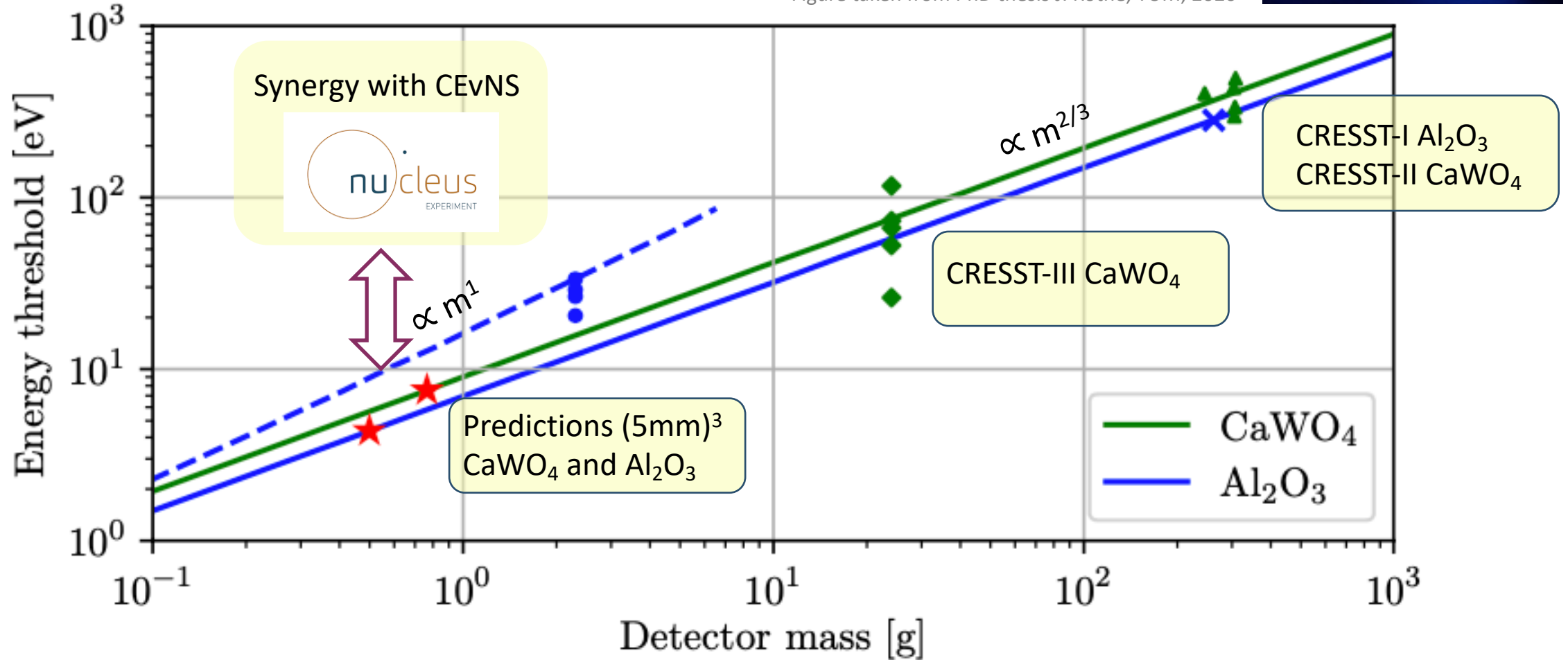


Figure taken from PhD thesis J. Rothe, TUM, 2020

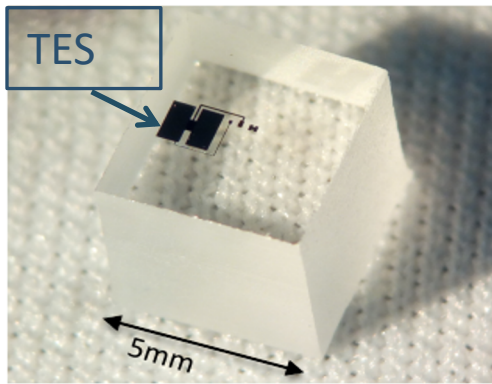



# NUCLEUS

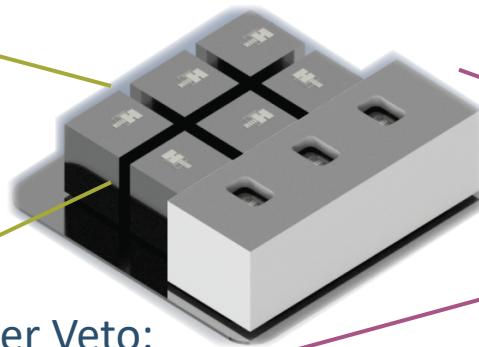
R. Strauss et al., Eur. Phys. J. C 77 (2017) 506

- Based on CRESST technology
- Demonstrated  $20 \text{ eV}_{\text{nr}}$  threshold
- 2 arrays of gram-scale cryogenic calorimeters
- Multitarget approach for  $N^2$  dependence of CEvNS
- Cryogenic veto system for background rejection

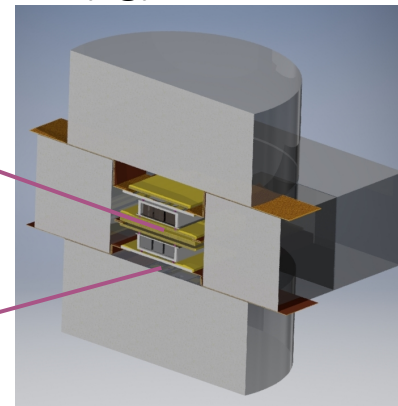
**NUCLEUS-10g**  
Target:  $\text{CaWO}_4, \text{Al}_2\text{O}_3$   
Mass: 6.8g + 4.4g  
Location: Chooz, France  
Status: Construction



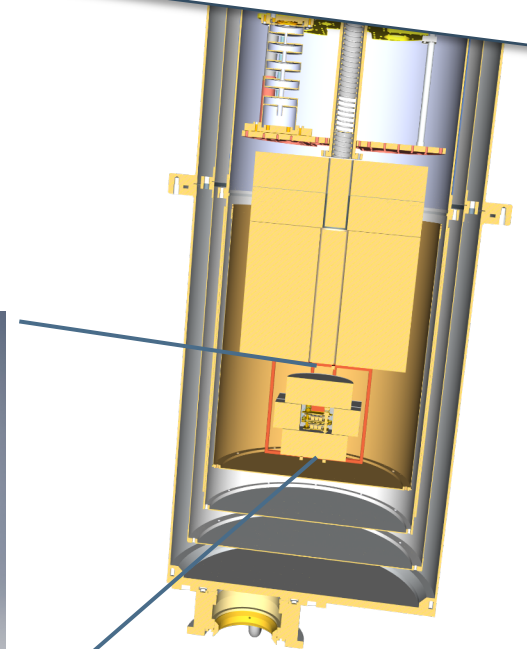
0.5g  $\text{Al}_2\text{O}_3$  prototype



Inner Veto:  
Instrumented holder



Outer Veto:  
O(kg) Ge detector

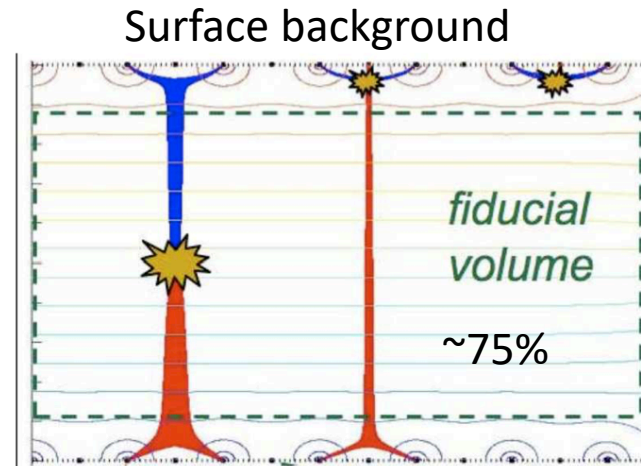
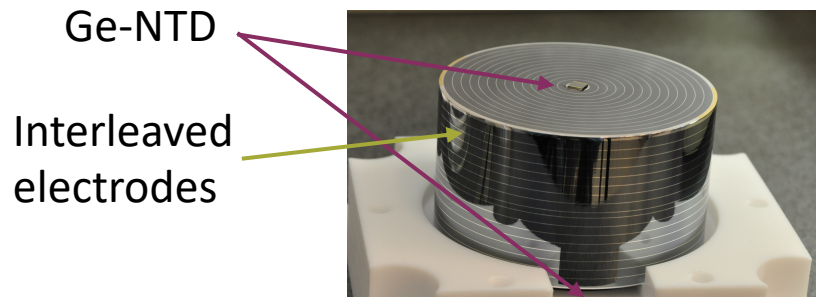




# EDELWEISS-III

E. Armengaud *et al* 2017 *JINST* 12 P08010

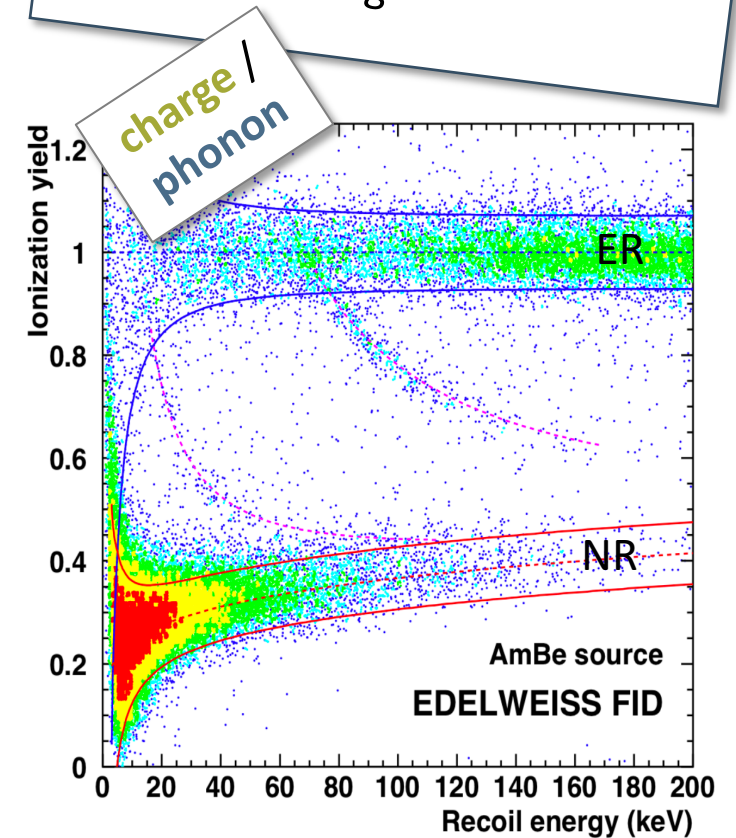
- 800g Ge detectors
- Heat & ionization readout
- Ionization energy resolution  $< 0.7 \text{ keV}_{ee}$  (FWHM)
- Heat energy resolution  $< 1.5 \text{ keV}$  (FWHM)



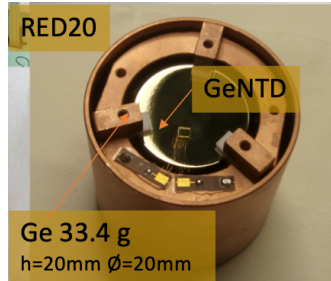
[http://dx.doi.org/10.3204/DESY-PROC-2010-03/gerbier\\_gilles](http://dx.doi.org/10.3204/DESY-PROC-2010-03/gerbier_gilles)

**EDELWEISS-III**

Target: Ge  
 Mass: 24 x 800g  
 Location: LSM  
 Status: Running



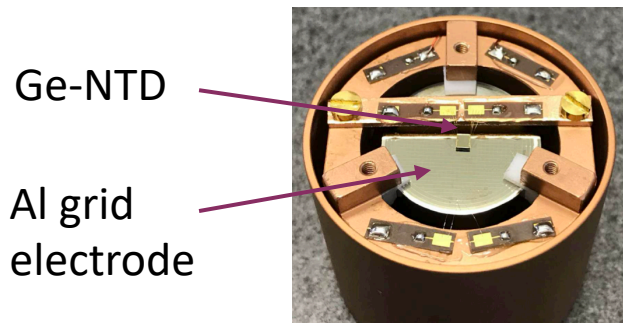
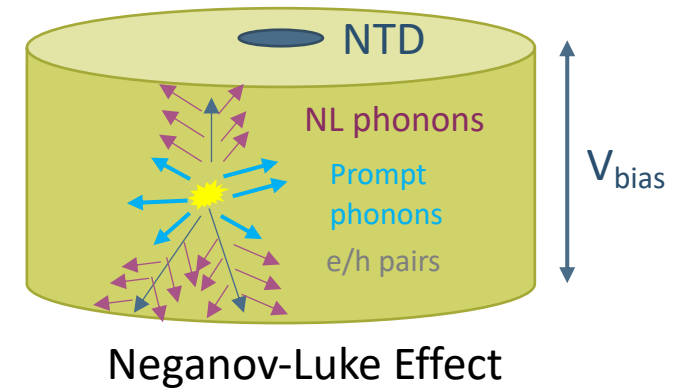
# Towards low Threshold with EDELWEISS-III



## EDELWEISS-Surf

*E. Armengaud et al., Phys. Rev. D 99, 082003 (2019)*

- Mass reduced to **33 g**
- Analysis threshold reduced to **55 eV<sub>nr</sub>**
- No ionization channel



## High voltage R&D

- Single e-/h sensitivity
- Use Neganov-Luke effect

## Low Voltage R&D:

- Goal: 10 eV (20 eV) heat (ionization) resolution
- R&D on front-end electronics with HEMT preamplifier

Synergy with CEvNS

**RICCOCHET**  
A Coherent Neutrino Scattering Program

# RICOCHET

J Billard et al 2017 *J. Phys. G: Nucl. Part. Phys.* **44** 105101

- O(kg) cryogenic detector array with 10 eV threshold to measure CEvNS
- Mixed array of Ge and Zn target detectors

**RICOCHET**  
A Coherent Neutrino Scattering Program

Target: Ge, Zn  
 Mass: 27 x 32g  
 Location: ILL, Grenoble  
 Status: Construction

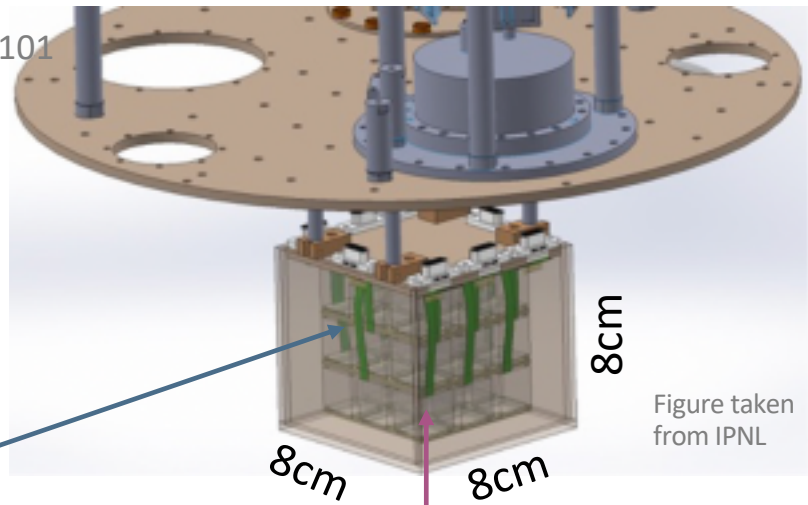
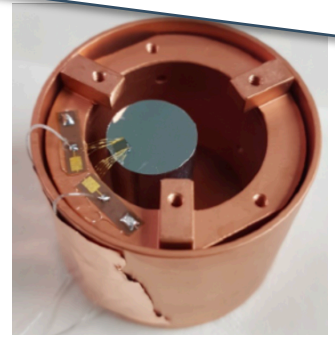
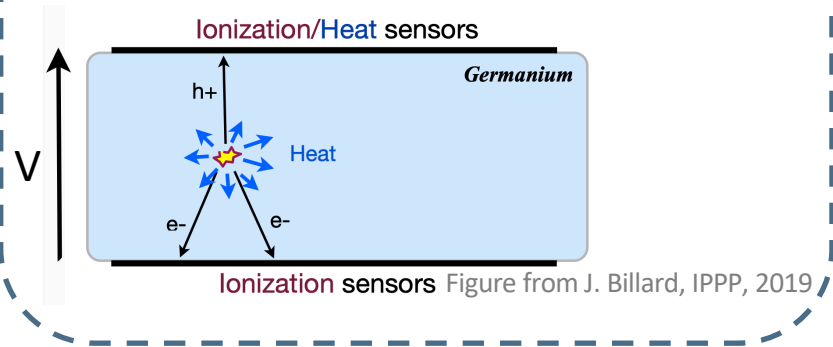


Figure taken from IPNL



## Ge semiconductors

- Well known technique with proven 55 eV threshold



## Zn superconducting metal

- NEW technology
- Prompt vs delayed heat signal depend on particle type

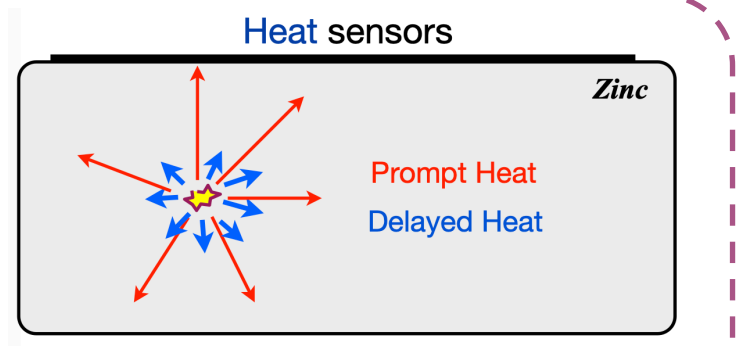


Figure from J. Billard, IPPP, 2019

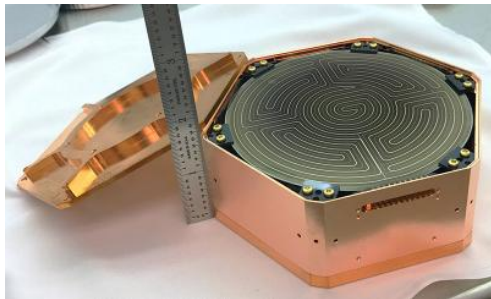
# SuperCDMS

SuperCDMS Collaboration, Phys. Rev. D **95**, 082002

1.4 kg Ge and 0.6 kg Si detectors with two designs

## iZIP Detector

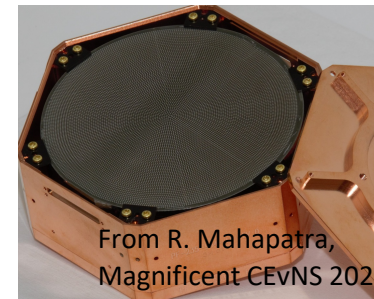
- Focus on background,  $m_\chi \sim 5 \text{ GeV}/c^2$
- Interleaved phonon and ionization sensors
- Surface event discrimination
- ER/NR discrimination down to  $2 \text{ keV}_{nr}$



## HV detector

- Focus on low threshold to explore  $m_\chi < 5 \text{ GeV}/c^2$
- $\sim 100\text{V}$  bias to increase of phonon signal:
 
$$E_{total} = E_{recoil} + n_{eh}e\Delta V = E_{recoil} \left( 1 + \frac{Ye\Delta V}{\langle E_{eh} \rangle} \right)$$
- Only (TES-)based phonon sensors


| Anticipated                | iZIP |     | HV  |     |
|----------------------------|------|-----|-----|-----|
|                            | Ge   | Si  | Ge  | Si  |
| Number of detectors        | 10   | 2   | 8   | 4   |
| Total exposure (kg·yr)     | 56   | 4.8 | 44  | 9.6 |
| Phonon resolution (eV)     | 50   | 25  | 10  | 5   |
| Ionization resolution (eV) | 100  | 110 | –   | –   |
| Voltage Bias (V)           | 6    | 8   | 100 | 100 |



Synergy with CEvNS



**SuperCDMS**



Target: Ge, Si  
 Mass: 10kg  
 Location: SNOLAB  
 Status: construction



# MIvER

Phys. Rev. D 94, 093002 (2016)

<http://people.tamu.edu/~mahapatra/miner.html>

MIvER



Target: Ge, Si  
Mass: 10kg

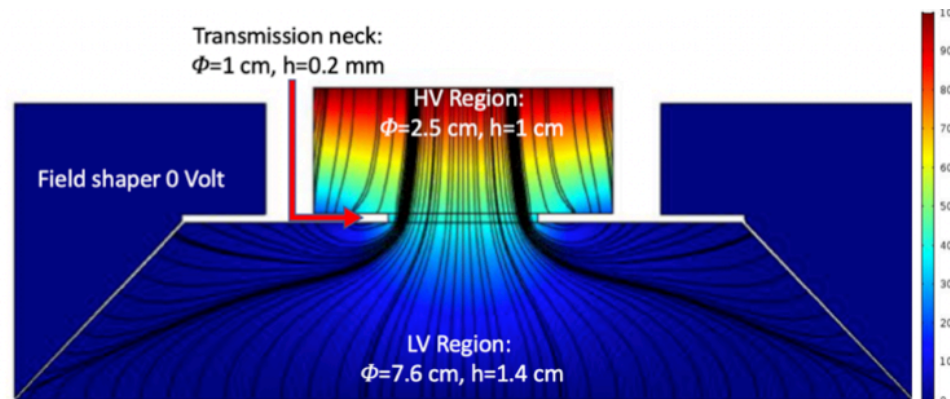
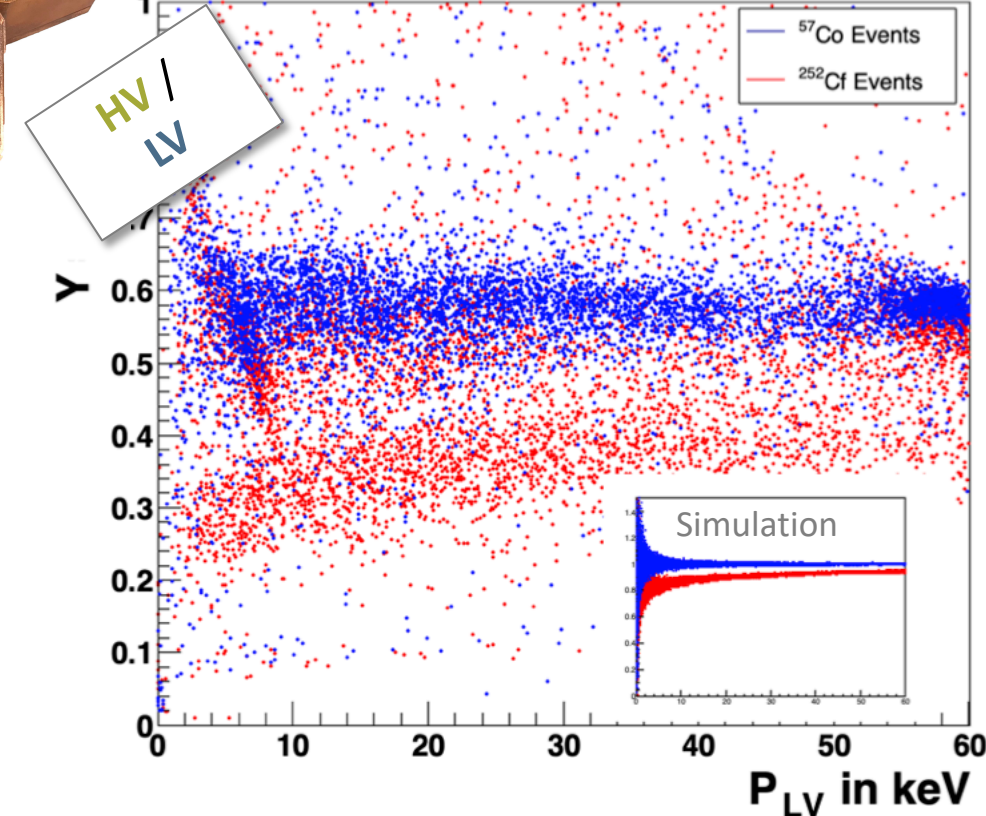
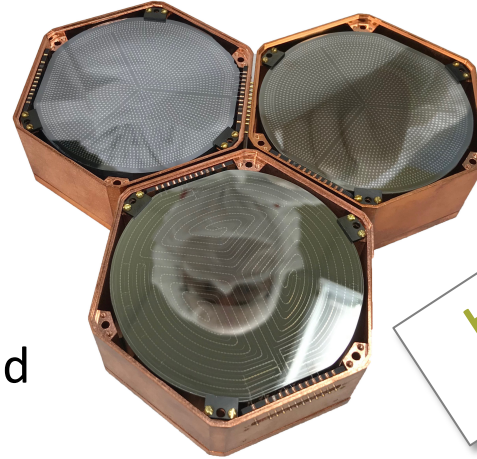
Location: TAMU, Texas

Status: running

- Phase 1:
  - 3 Si HV detectors (single e/h sensitive) for CEvNS
  - 1 new Si Hybrid in-situ background measurement

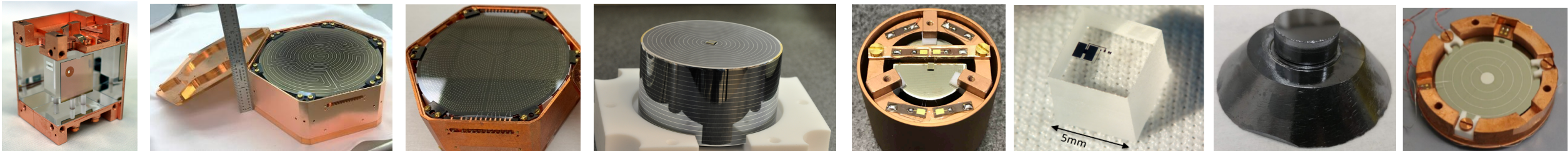
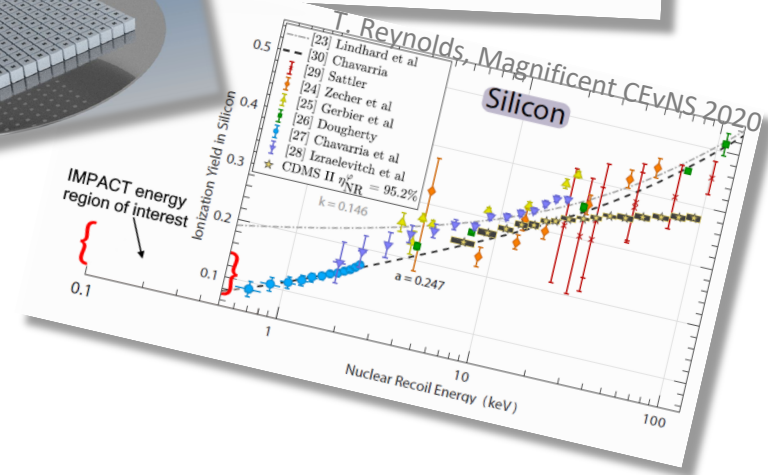
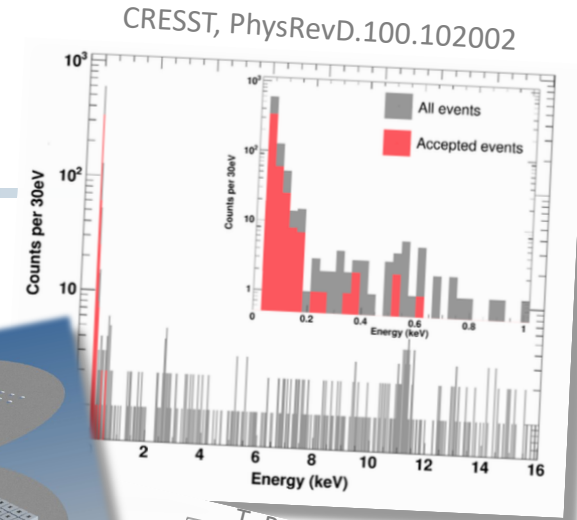
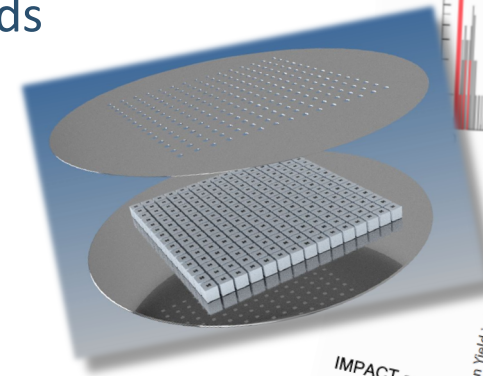
## Hybrid HV Detector [arXiv:2006.13139v1](https://arxiv.org/abs/2006.13139v1)

- Combine single e/h sensitivity & background discrimination
- LV: measurement of primary phonons (iZIP-like)
- HV: measurement of Neganov-Luke phonons



# Summary Conclusion

- Low mass dark matter search & CEvNS are very active fields
- Many new ideas under way
- Challenges
  - Low energy background
  - Calibration of nuclear recoil
  - Measurement of quenching factors
  - Scalability of read-out



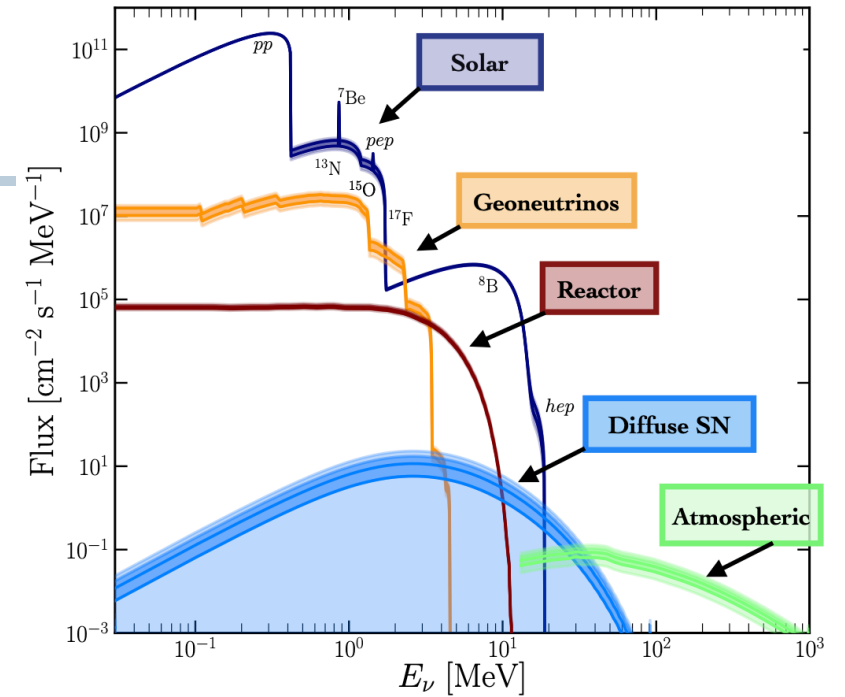
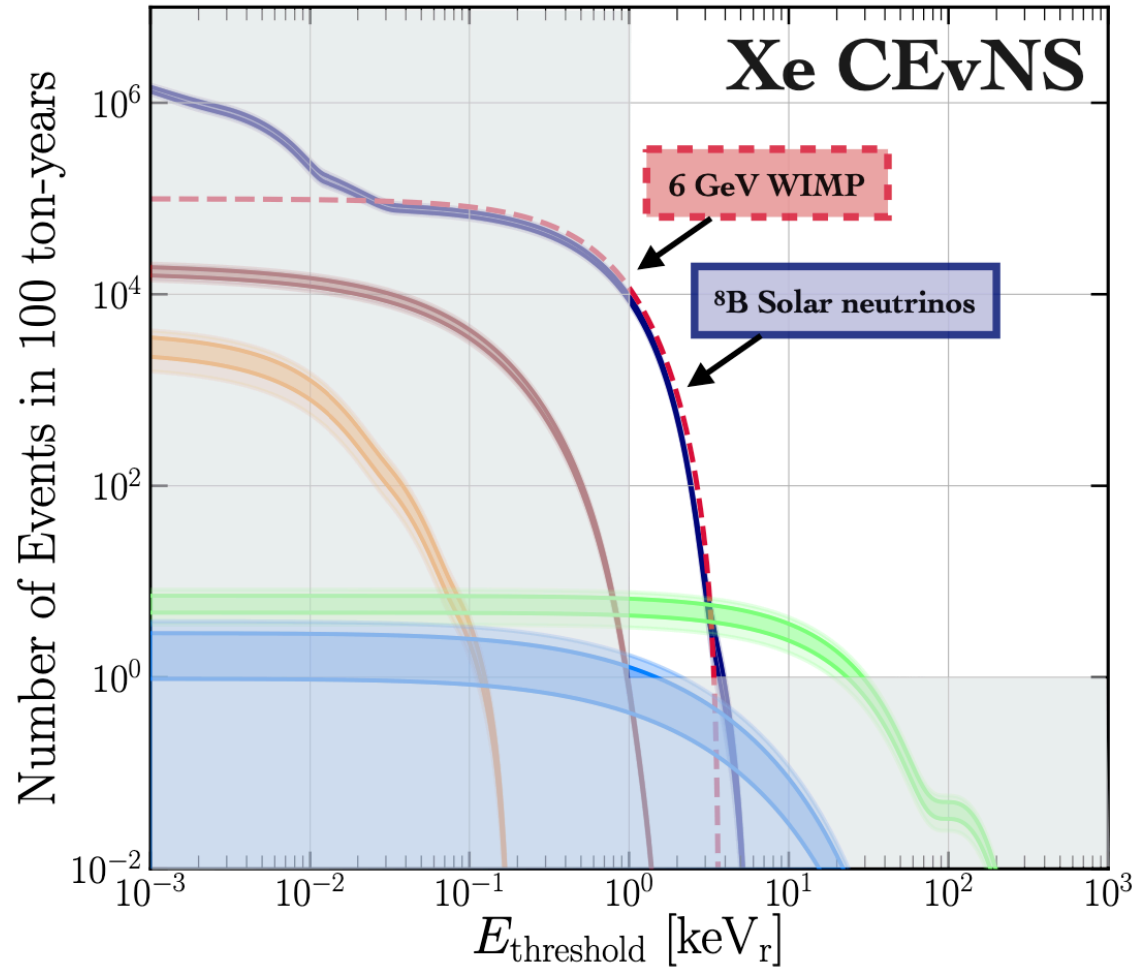


MUNICH  
Oct 5-7 2021

# Thank you for your attention

# Neutrino Floor

C. O'Hare, Magnificent CEvNS 2020

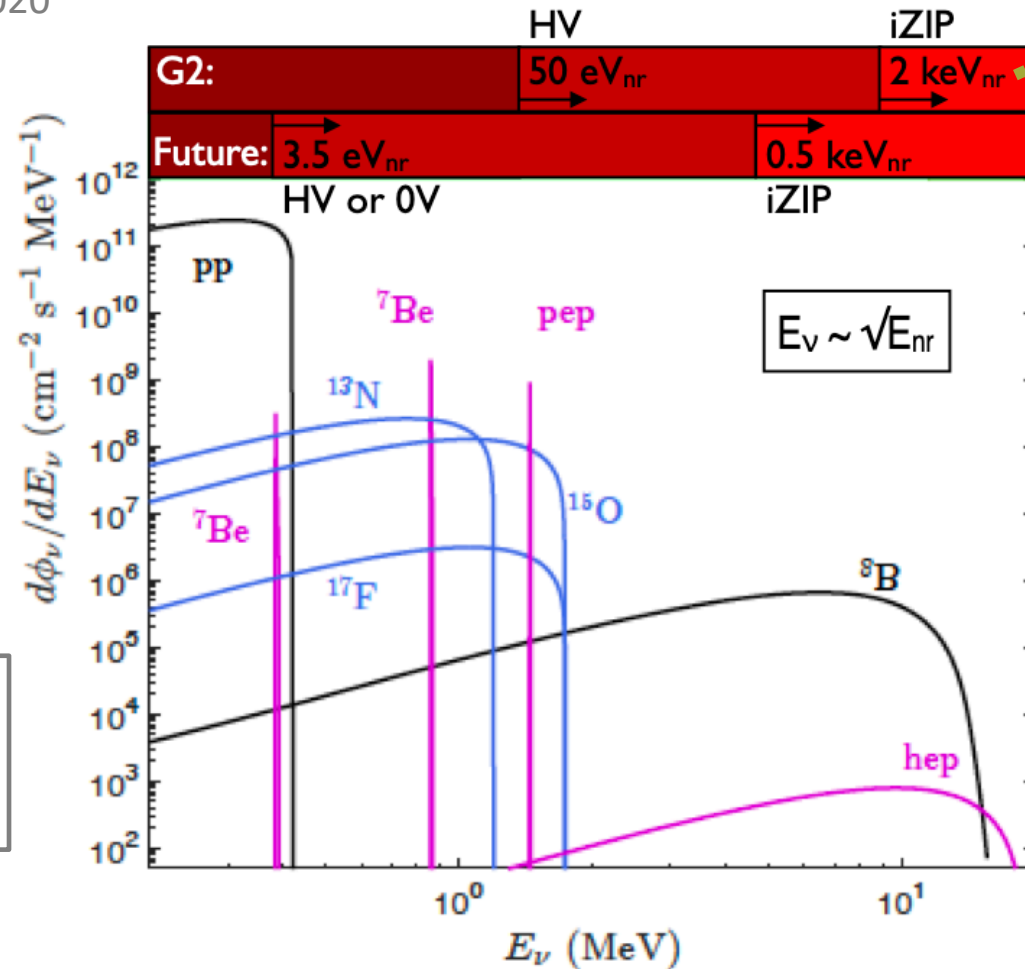


- Discrimination by
  - Annual modulation
  - Complementary targets
  - Directional detectors
  - Improved energy resolution



# CEvNS with SuperCDMS

S. Golwala, Magnificent CEvNS 2020



HV detectors:

- No discrimination
- neutrinos  $\ll$  background

## iZIP detectors:

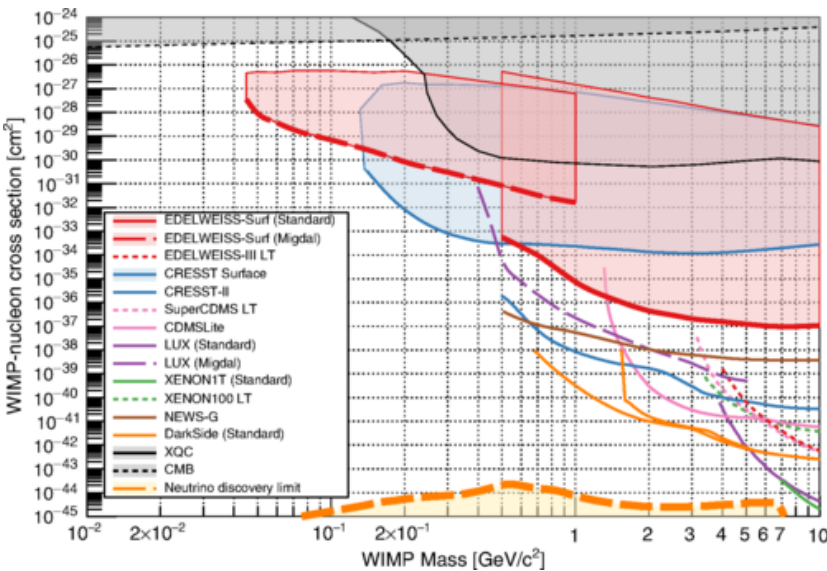
Few  $^8\text{B}$  w/ discrimination, O(50) down to 350 eV<sub>nr</sub> w/o discrimination

Discrimination down to 0.5 keV<sub>nr</sub>

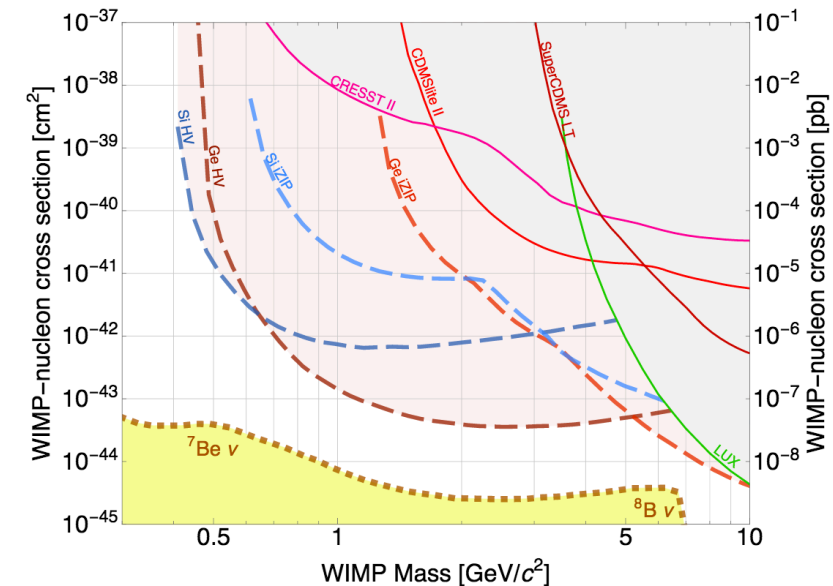
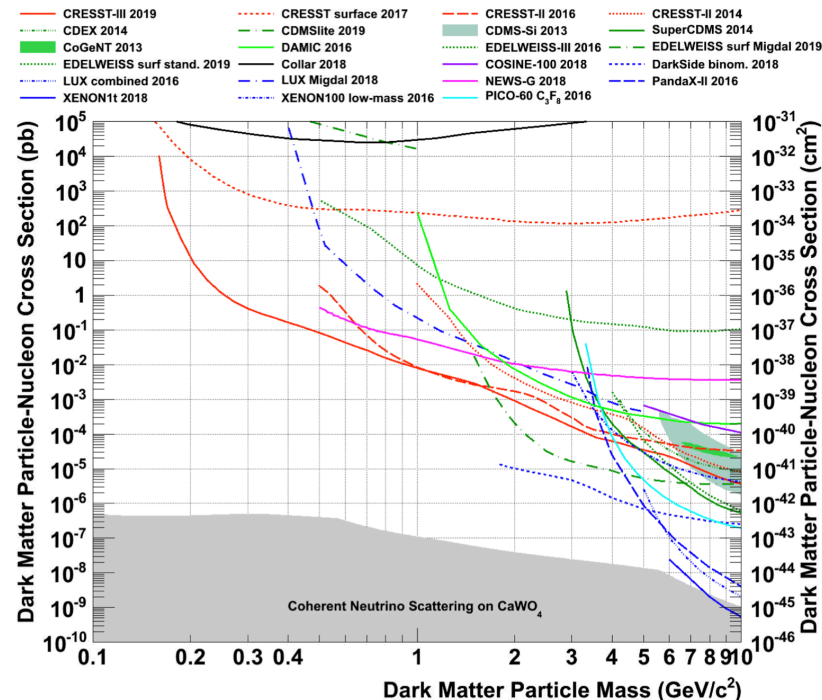
Few more  $^8\text{B}$  w/ discrimination

# Perspective Dark Matter

- Cryogenic experiments optimized for low-mass DM
- Can probe masses down to 0.2 GeV/c<sup>2</sup>
- Future phases will probe cross sections down to  $10^{-43} - 10^{-44} \text{ cm}^2$



Phys. Rev. D **99**, 082003

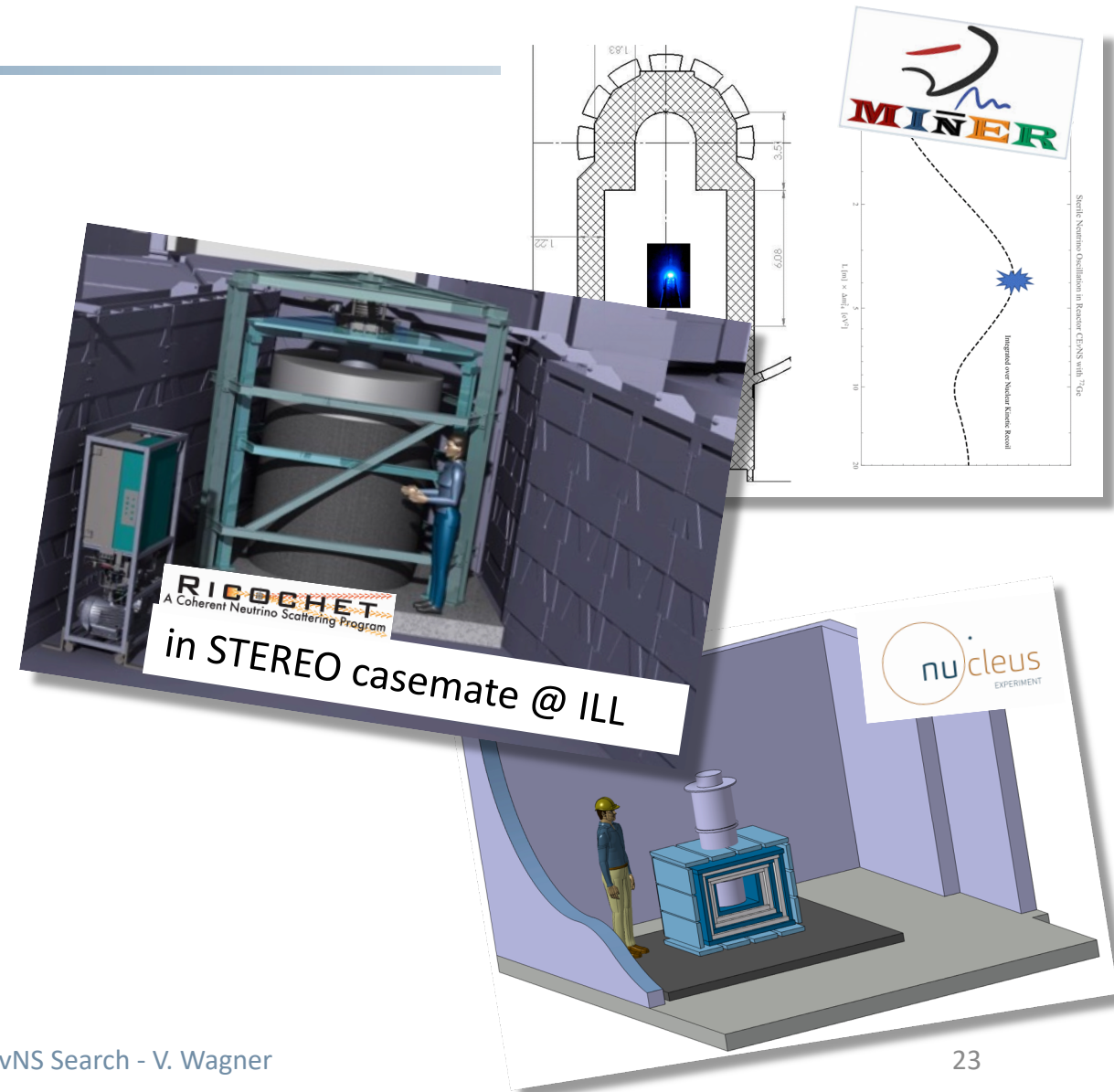


R. Agnese et al., Phys. Rev. D 95, 082002 (2017)

R. Agnese et al., arXiv:1707.01632v1

# Reactor CEvNS Experiments with Cryogenic Detectors

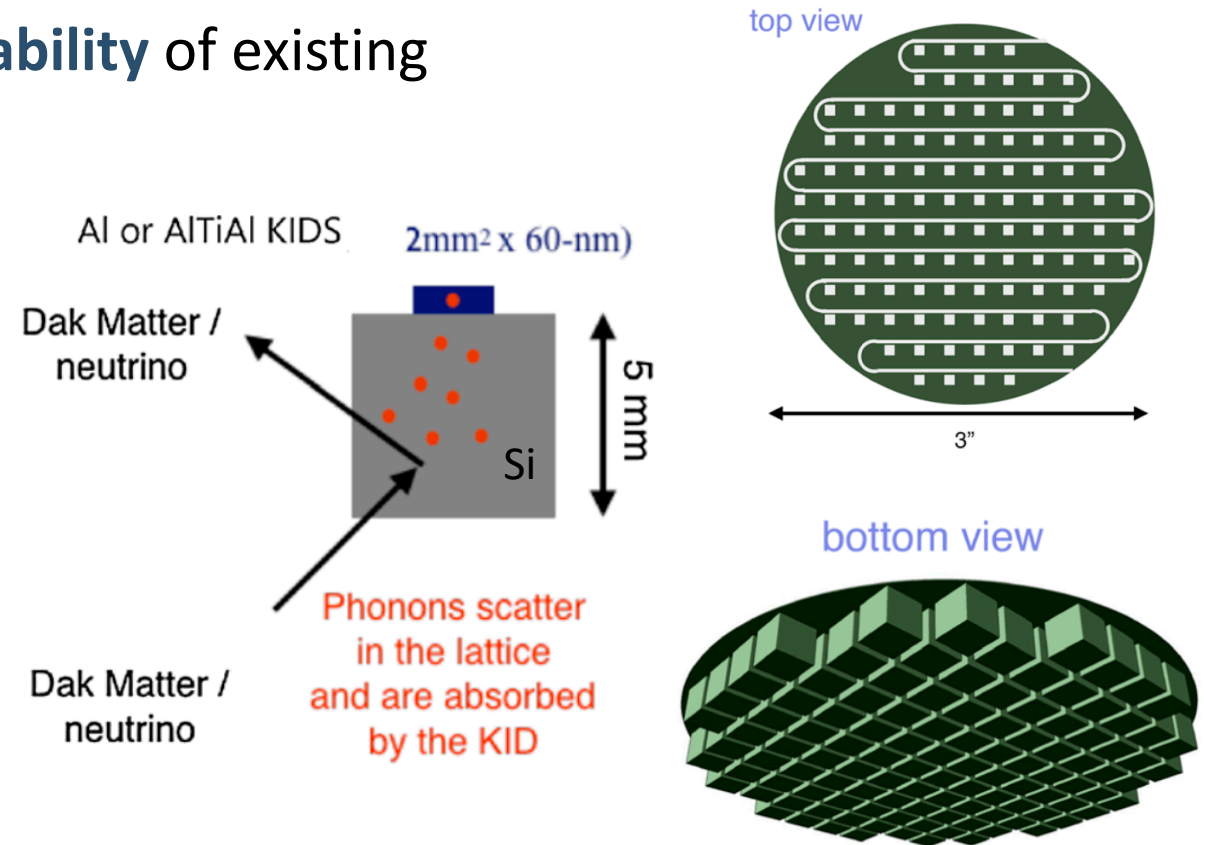
| Experiment | Target   | Threshold                | $\nu$ -flux [cm <sup>-2</sup> s <sup>-1</sup> ] |
|------------|--|--------------------------|---|
| MIvER      | Ge, Si   | O(100 eV <sub>nr</sub> ) | 4x10 <sup>11</sup>                              |
| RICOCHET   | Ge, Zn   | 55 eV <sub>nr</sub>      | 9x10 <sup>11</sup>                              |
| NUCLEUS    | CaWO <sub>4</sub> , Al <sub>2</sub> O <sub>3</sub> | 20 eV <sub>nr</sub>      | 2x10 <sup>12</sup>                              |



# BULLKID

Journal of Low Temperature Physics (2020) 199:593–597

- New detector concept addressing **scalability** of existing cryogenic technologies
- Goal:
  - energy threshold  $< 100 \text{ eV}_{\text{nr}}$  and
  - **O(1kg)** target mass
- Low phonon efficiency O(10%)
- Different KID materials under study:
  - Al, AlTiAl (25 eV resolution with CALDER<sup>1</sup>)
  - W + Al and Hf



<sup>1</sup>Casali et al., EPJ C 79 724

# MKID

Energy deposition



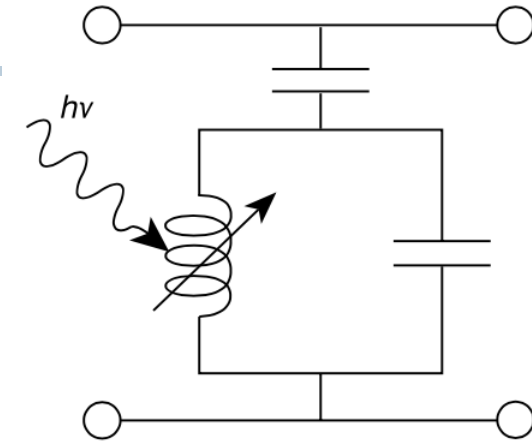
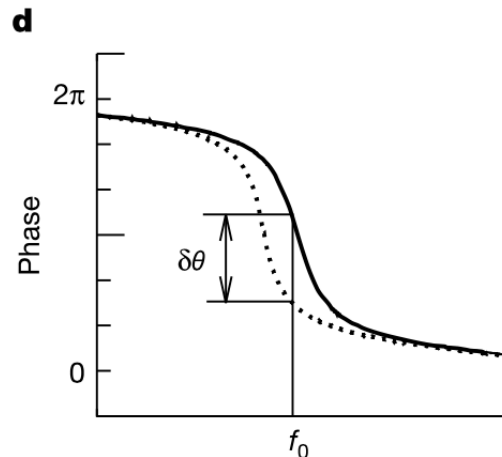
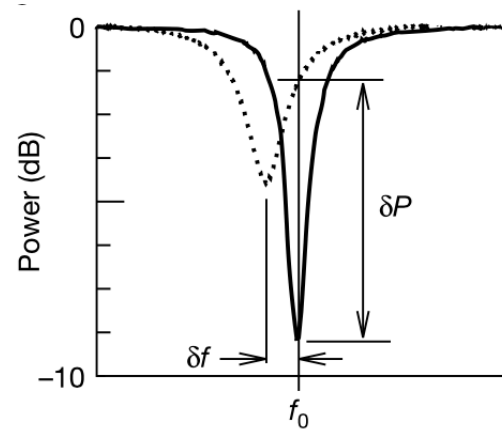
Breaking of Cooper pairs



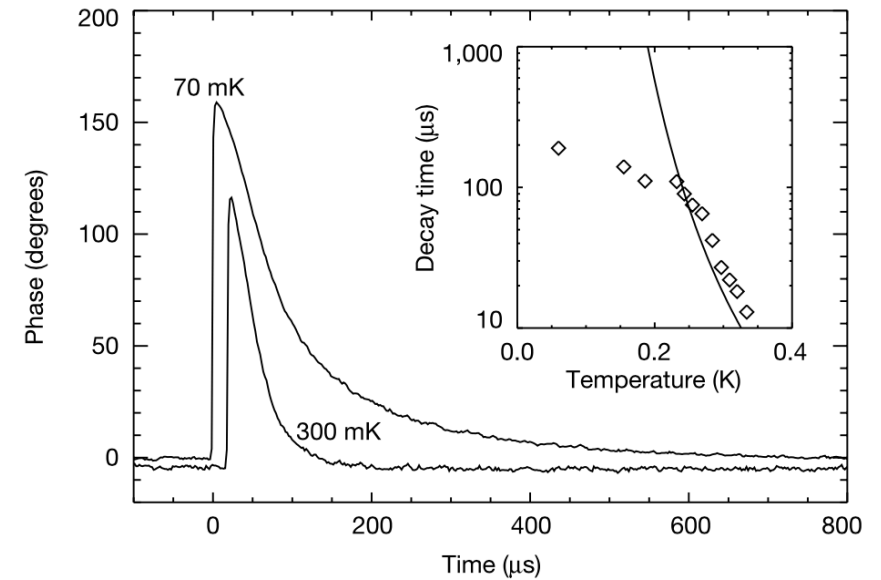
Change in kinetic inductance



Change of resonance frequency  $f_0$



Fine-tune  $f_0$   
for  
multiplexing in  
GHz range

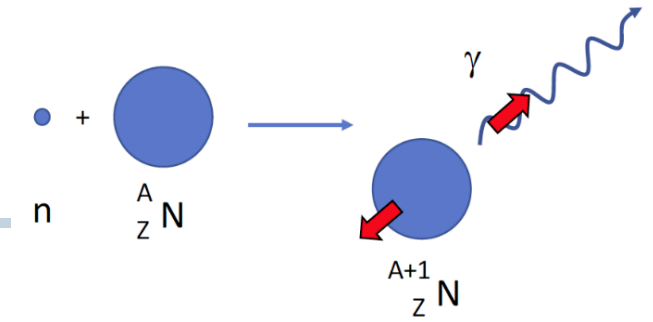
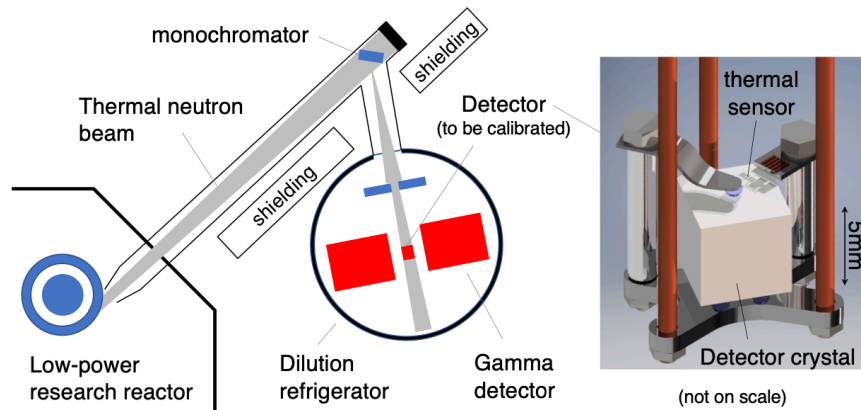


From nature02037

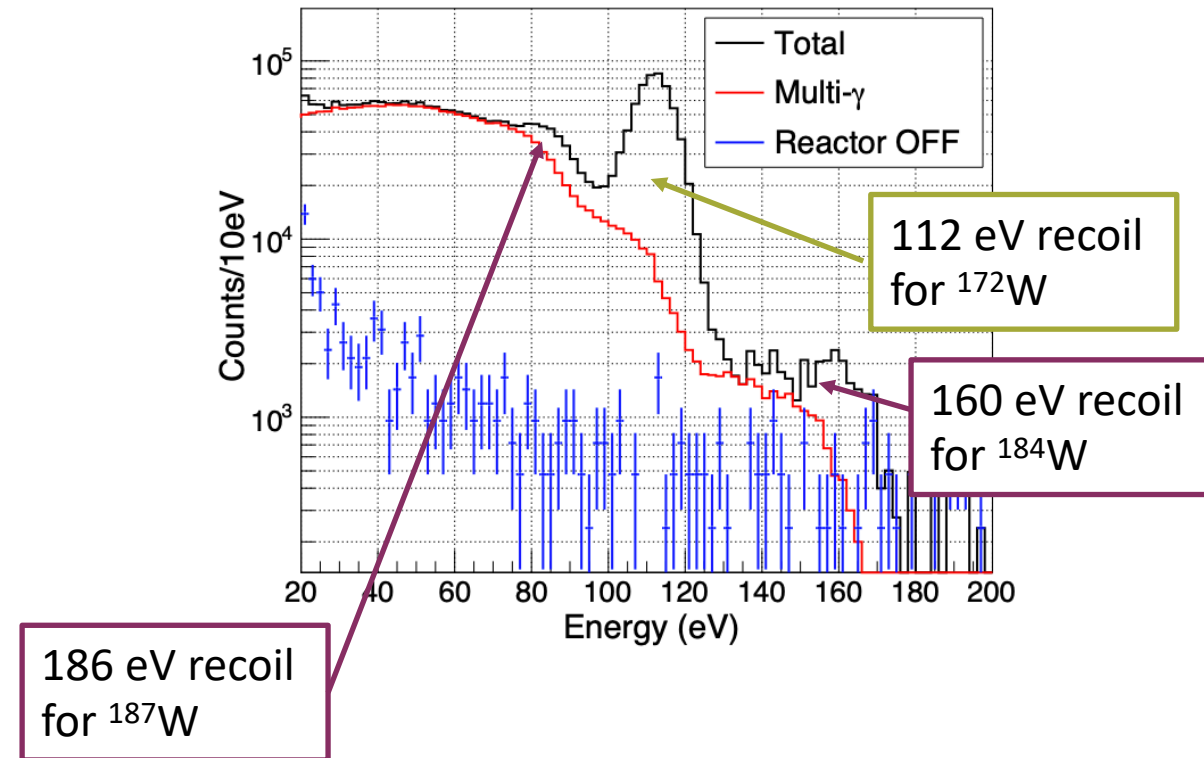
# CRAB

L. Thuillier arXiv: 2011.13803

- Calibration at the 100 eV scale using neutron-induced nuclear recoils
- **Idea:** nuclear recoils from gamma emission
- $\gamma$ -emission can be used for background reduction
- Suggested experiment at low power TRIGA reactor



Simulation for 0,76g CaWO<sub>4</sub> crystal and 5 eV energy resolution



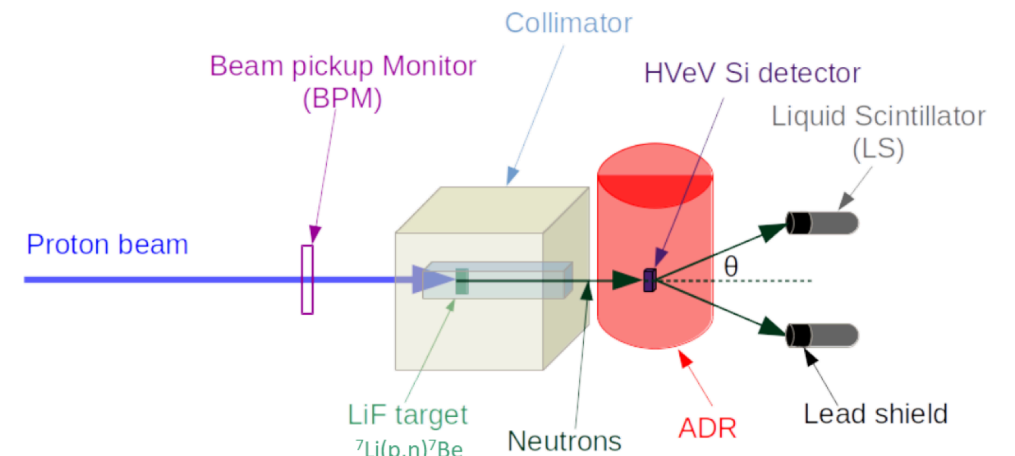
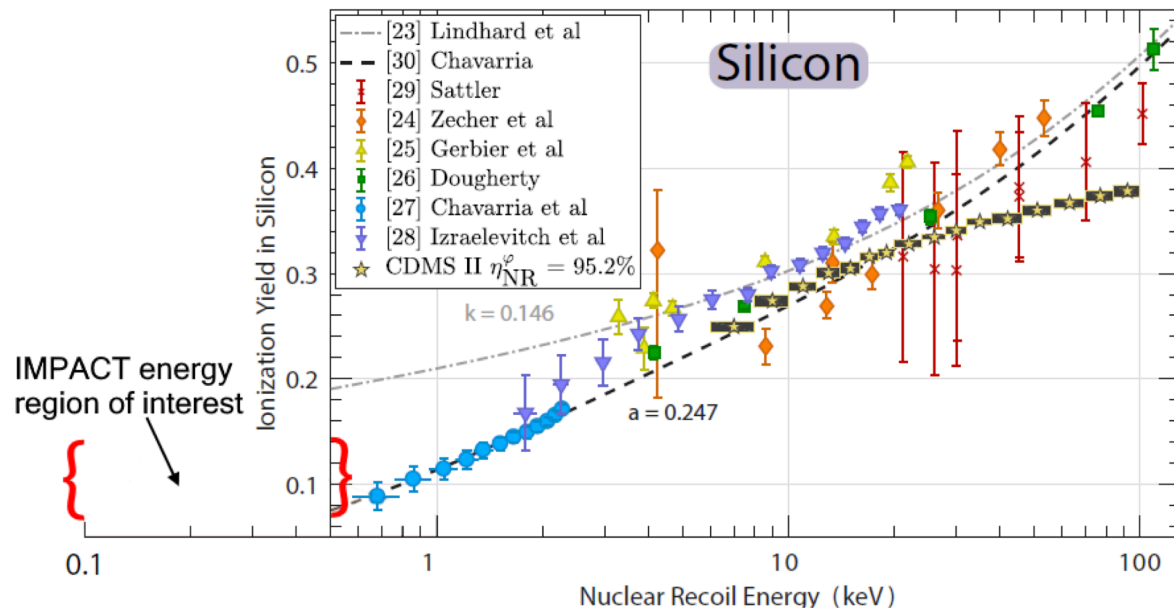


# SuperCDMS Ionization Yield Measurement

T. Reynolds, Magnificent CEvNS 2020

$$E_{total} = E_{recoil} + n_{eh}e\Delta V = E_{recoil} \left( 1 + \frac{Ye\Delta V}{\langle E_{eh} \rangle} \right)$$

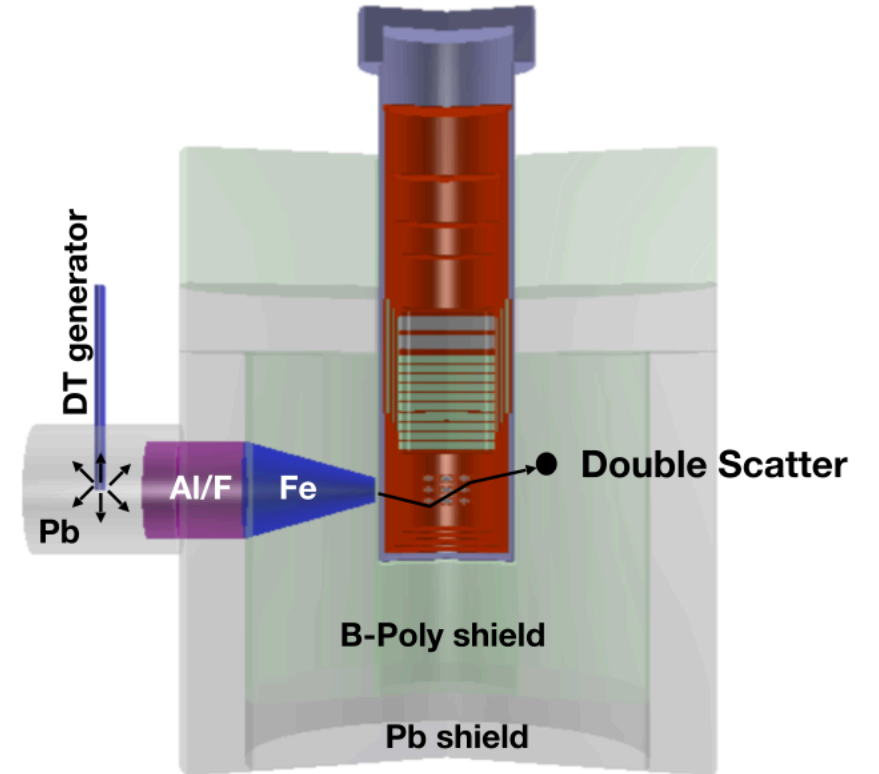
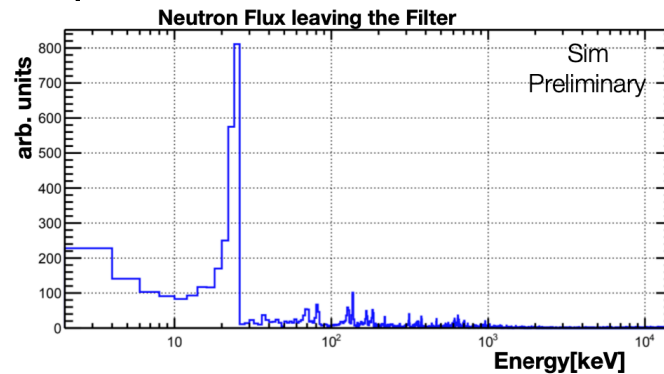
IMPACT@TUNL, results expected soon



# keV Neutron Source

- For  $< 100$  eV nuclear recoil need keV neutron source
- Ideally portable, to measure at experimental side
- Fe-filtered DT generator for RICOCHET:

P. Patel, Magnificent CEvNS 2020



T. Salanac, Magnificent CEvNS 2020

