Physicists Learning from Machines Learning

Smart but Interpretable Neural Networks for Physics at the LHC

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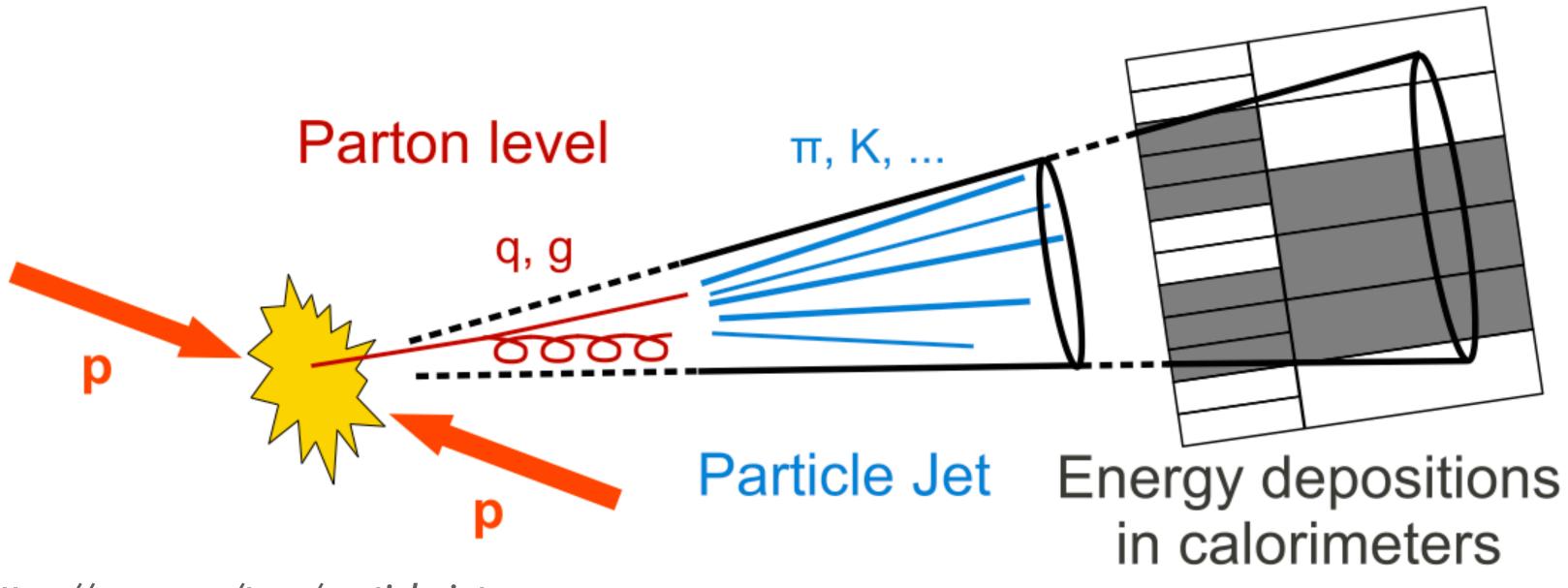
Mapping machine-learned physics into a human-readable space https://journals.aps.org/prd/abstract/10.1103/PhysRevD.103.036020

@ Clermont-Ferrand



Hadrons & Jets

- Particle collisions create particles with non-zero color charge (i.e. quarks & gluons)
- Free quarks/gluons hadronize to produce hadrons (e.g. mesons and baryons)
- "Jets" are collimated sprays of many hadrons in a cone.
- Identifying jets and different kinds of jets help distinguish high-energy processes.

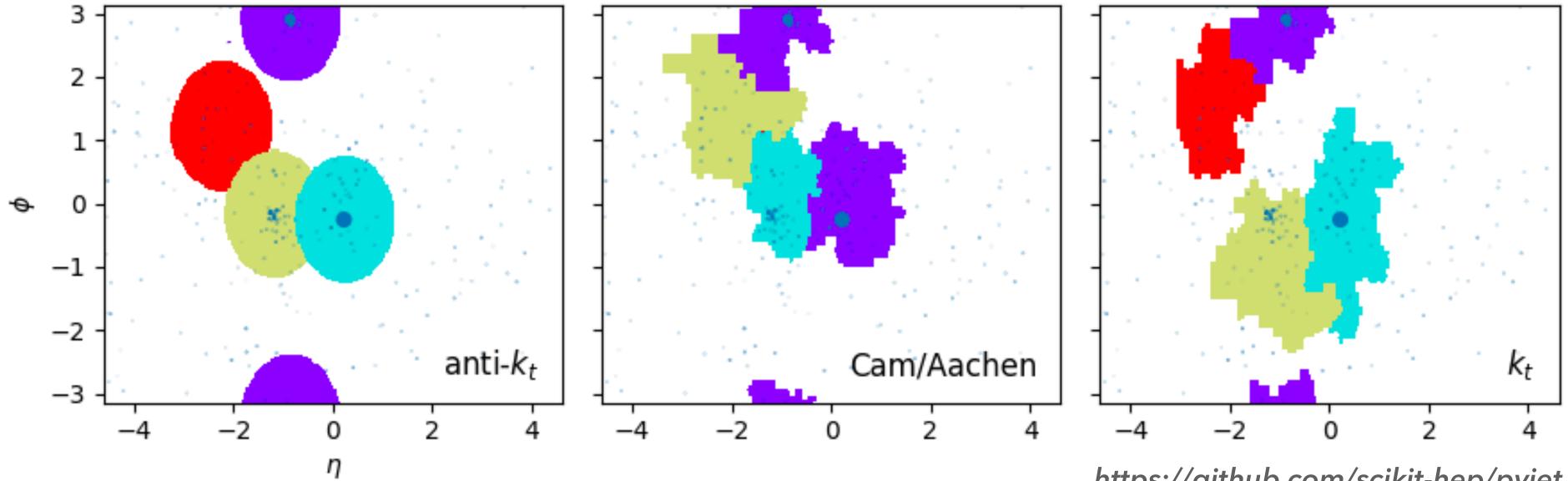


https://cms.cern/tags/particle-jet



Machine Learning Jet Substructure

- Jets are isolated in the detector using clustering algorithms
- How do we distinguish different types of jets from one another?



https://github.com/scikit-hep/pyjet



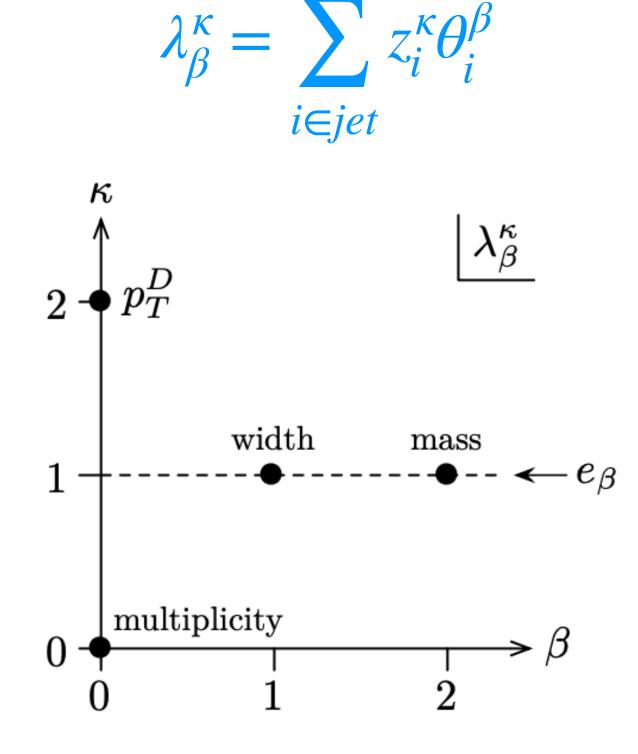
Jet Substructure (JSS)

What variables would be useful for training ML to classify different jets?

- Quark/Gluon discrimination?
 - Gluons produce more particles.
 - Look at "jet multiplicity" and similar proxies/weightings of momentum/ angular separation
 - From "generalized angularity": multiplicity, LHA, pTD...
- What about Jets with multiple sub-jets?
 - W/Z/h decay to 2 jets
 - We invent "N-subjettiness" to quantify separable jet substructure

$$\tau_N = \frac{1}{d_0} \sum_k p_{T,k} \min\left\{\Delta R_{0,k}, \Delta R_{1,k} \dots \Delta R_{1,k}\right\}$$



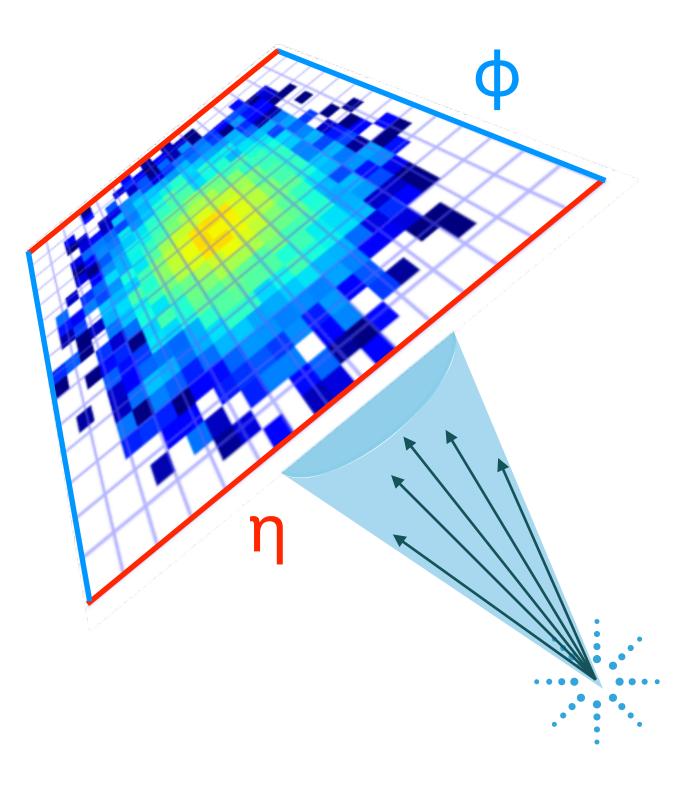






What's the alternative?

- Do we need to hand-pick observables in every study?
- What if we let a Deep Neural Network learn to solve the problem?
- Can a CNN to learn to classify directly from the calorimeter data?

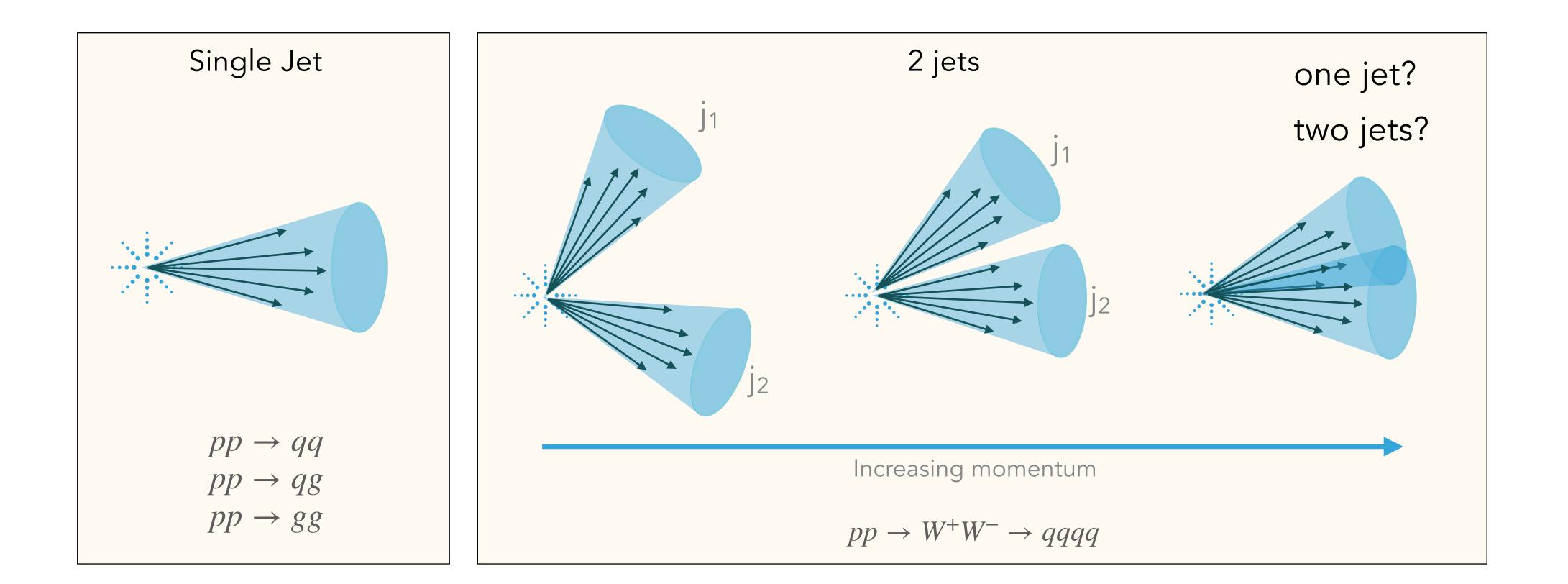


- $-E_T = Transverse Energy$
- Position (η, ϕ)
- $-\eta = -\ln(\tan(\theta/2))$



Test Case: Boosted W vs QCD

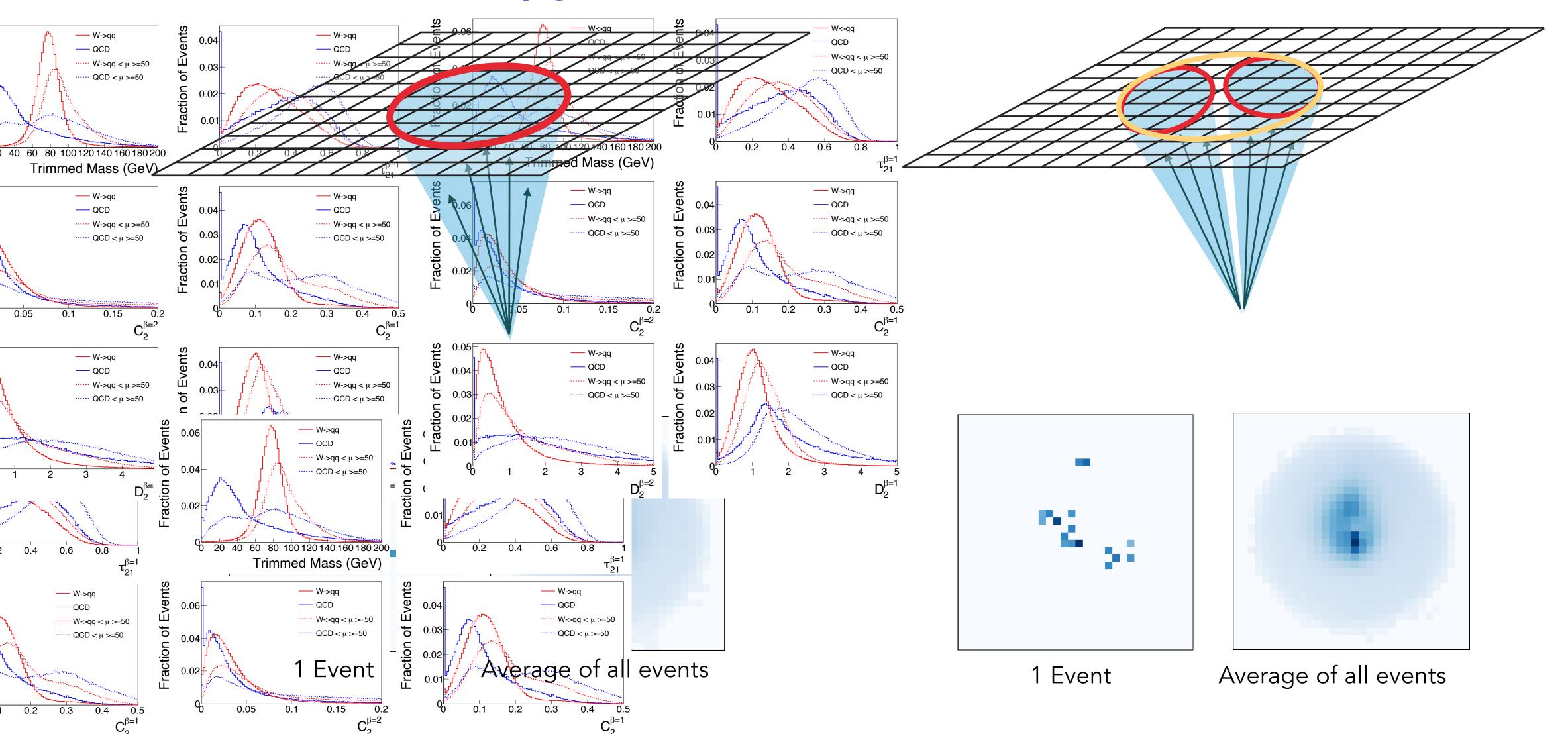
Boosted W bosons (W \rightarrow qq') create highly collimated di-jets. Can a CNN separate boosted W jets from QCD jets?





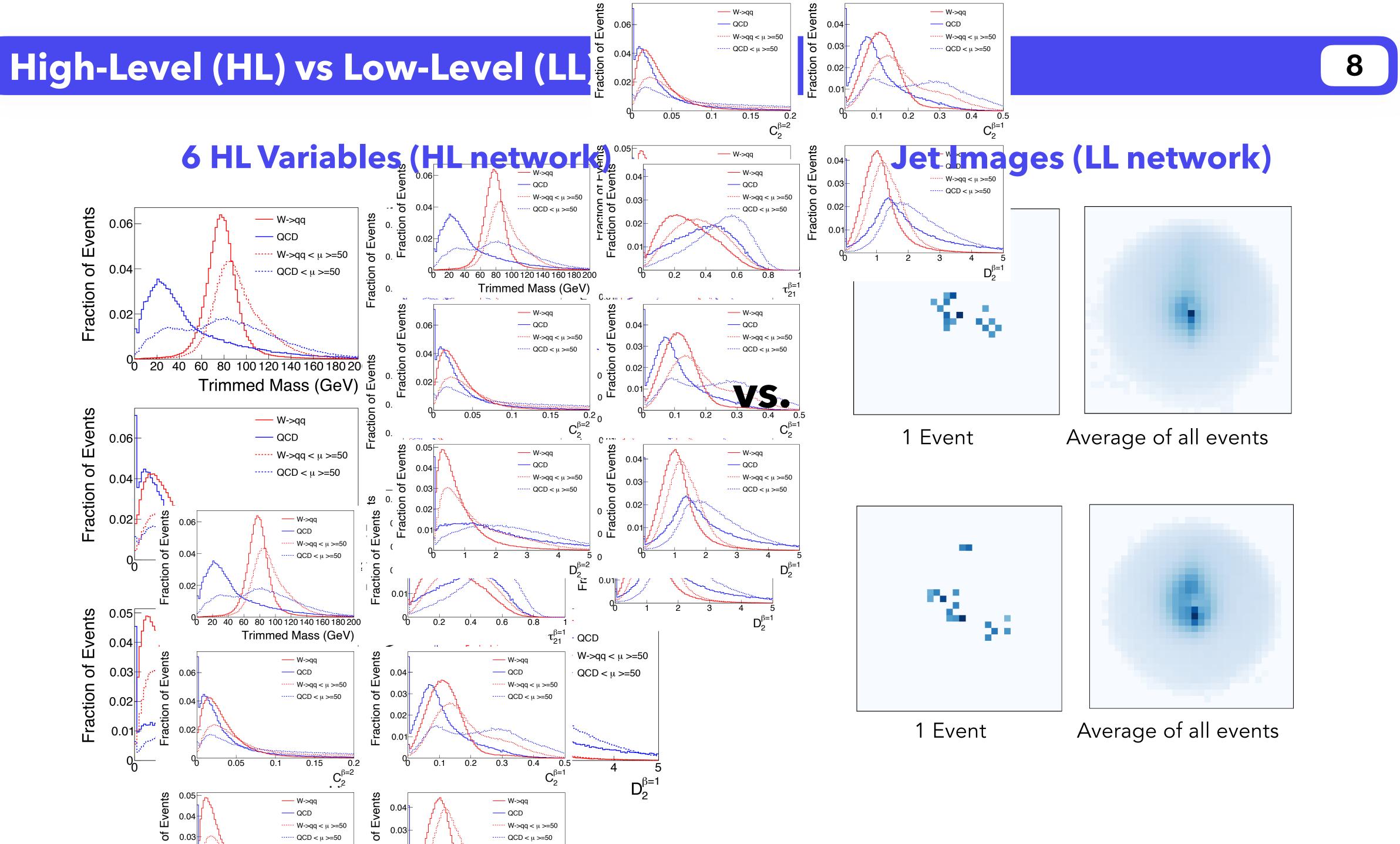
QCD mono-jet vs Boosted W di-jet

QCD Jet (q, g)



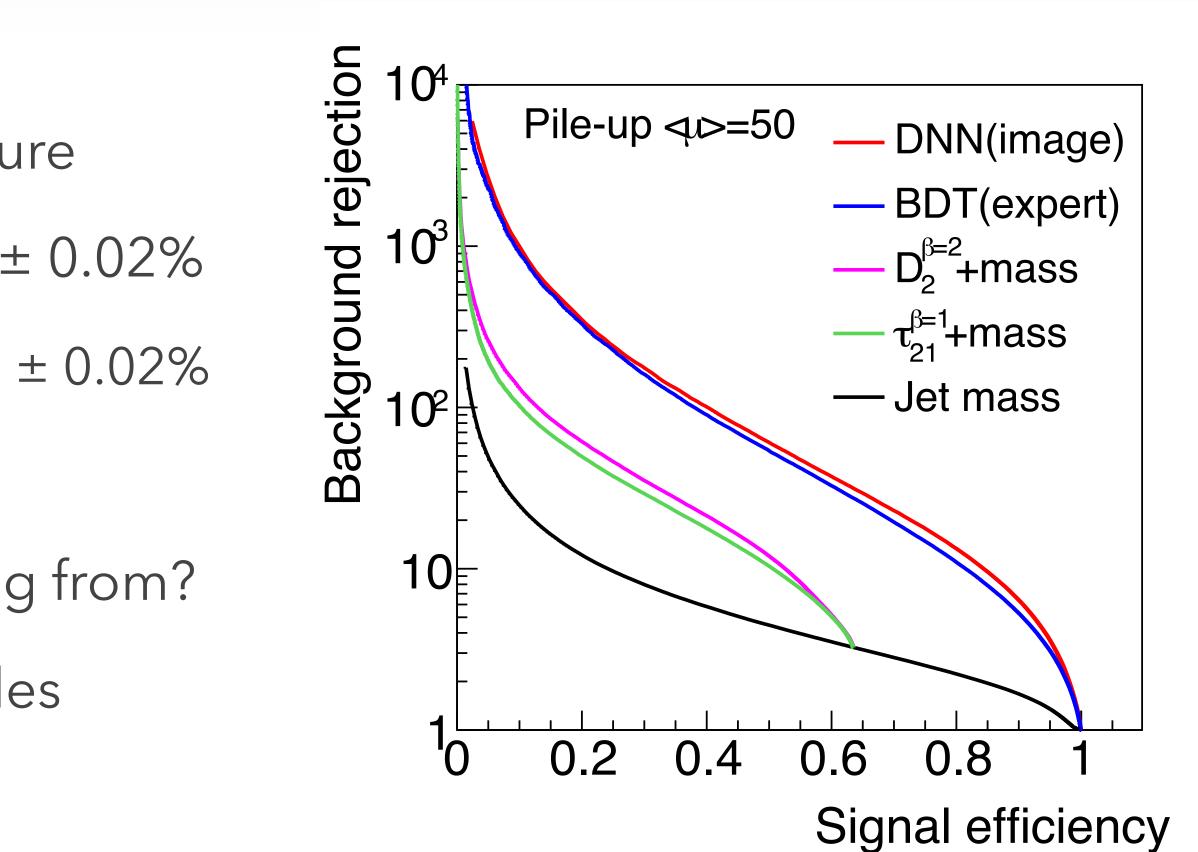
W jet





Jet Images out-perform JSS

- CNN performs better than Jet Substructure
 - LL network (red line): AUC = 95.30% ± 0.02%
 - ► HL network (blue line): AUC = 95.00% ± 0.02%
- But wait!
 - Where is that extra information coming from?
 - Why don't Jet Substructure observables contain this information?
- We've used a black box, so how can we investigate?



Baldi, P., Bauer, K., Eng, C., Sadowski, P., & Whiteson, D. (2016, March 30). Jet Substructure Classification in High-Energy Physics with Deep Neural Networks. arXiv.org. http://doi.org/ 10.1103/PhysRevD.93.094034

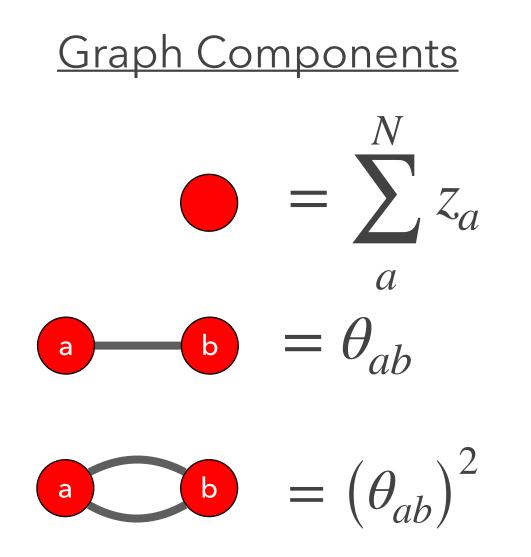


Another approach to generalizing JSS

- We weren't the only ones thinking about a generalized approach to understanding JSS!
 - Jesse Thaler, Patrick Komiske, Eric Metodiev

https://arxiv.org/abs/1712.07124

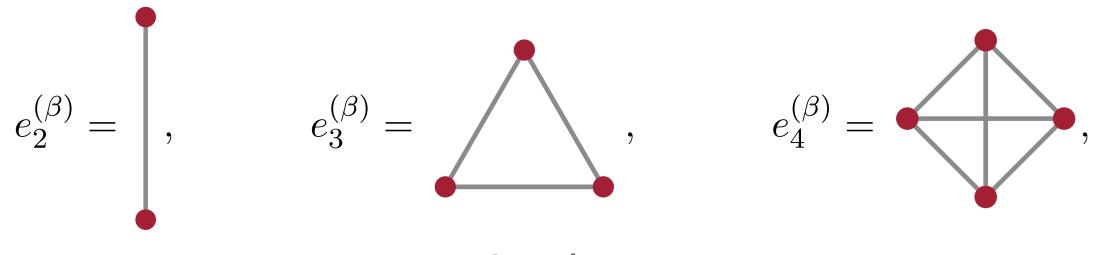
Energy Flow Polynomials (EFP): A complete linear basis set for jet substructure.





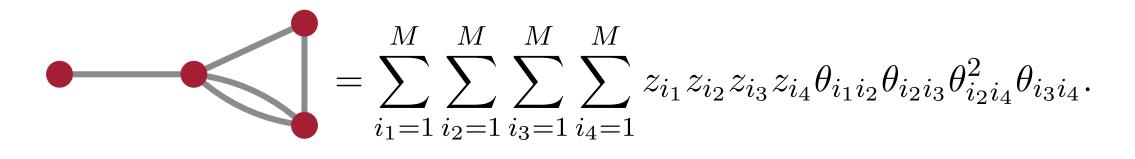


Existing JSS exists in the EFP space



Energy Correlation

<u>We can also explore more exotic observables.</u>



<u>With variables and parameters (κ , β)</u>

$$z_{i}^{\kappa} = \left(\frac{p_{T,i}}{\sum_{i} p_{T_{j}}}\right)^{\kappa} \qquad \theta_{ij}^{\beta} = \left(\Delta \eta_{ij}^{2} + \Delta \phi_{ij}^{2}\right)^{\beta/2}$$

Marry the two methods

- 1. Black-box Learning
 - **Benefits:** Powerful Performance and no need to hand-pick observables
 - **Drawbacks:** Not interpretable. What are we learning about the problem?
- 2. Energy Flow Polynomials (EFP)
 - Benefits:
 - Physics motivated
 - Modeling can be verified
 - Uncertainties can be defined
 - Compact and efficient

Combine them! The DNN has learned how to solve the problem. Let the DNN tell us what EFPs to choose!

• Drawbacks: It's an infinite space. How do we begin to choose observables and their parameters?



Decision Ordering: A New Metric for Decision Similarity

- Consider boosted W vs QCD jet binary classification
- CNN has learned where to draw an ideal decision surface in its feature space.
- We want a HL feature space that makes equally good decisions.
 How do we compare the decision surface for the CNN to the HL
- How do we compare the decision features?



Decision Ordering

Comparing pair orderings

- Take a pair of signal (x) and background (x') features,

- Predictions for 2 NN (f(x) and g(x)) of the features will increase/decrease relative to input

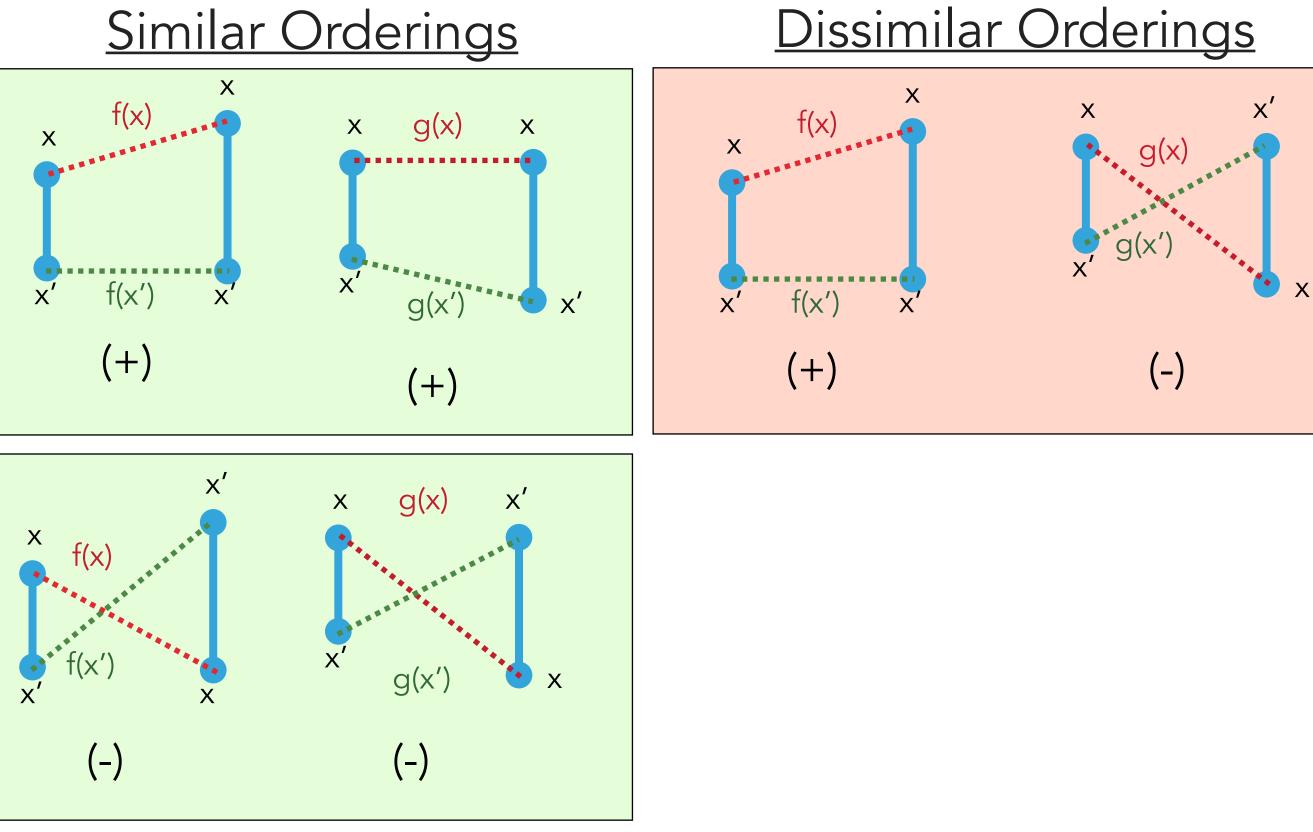
$$\mathrm{DO}(x, x') = \Theta\left[(f(x) - f(x')) \cdot (g(x) - g(x')) \right]$$

Heaviside step function (Θ) sets DO=1 for similar order, DO=0 for dissimilar order.

Average many examples = Average Decision Ordering

$$ADO' = \sum DO(x, x')$$

х′



- ADO = 1 : Identical decisions
- ADO = 0.5 : Random similarity

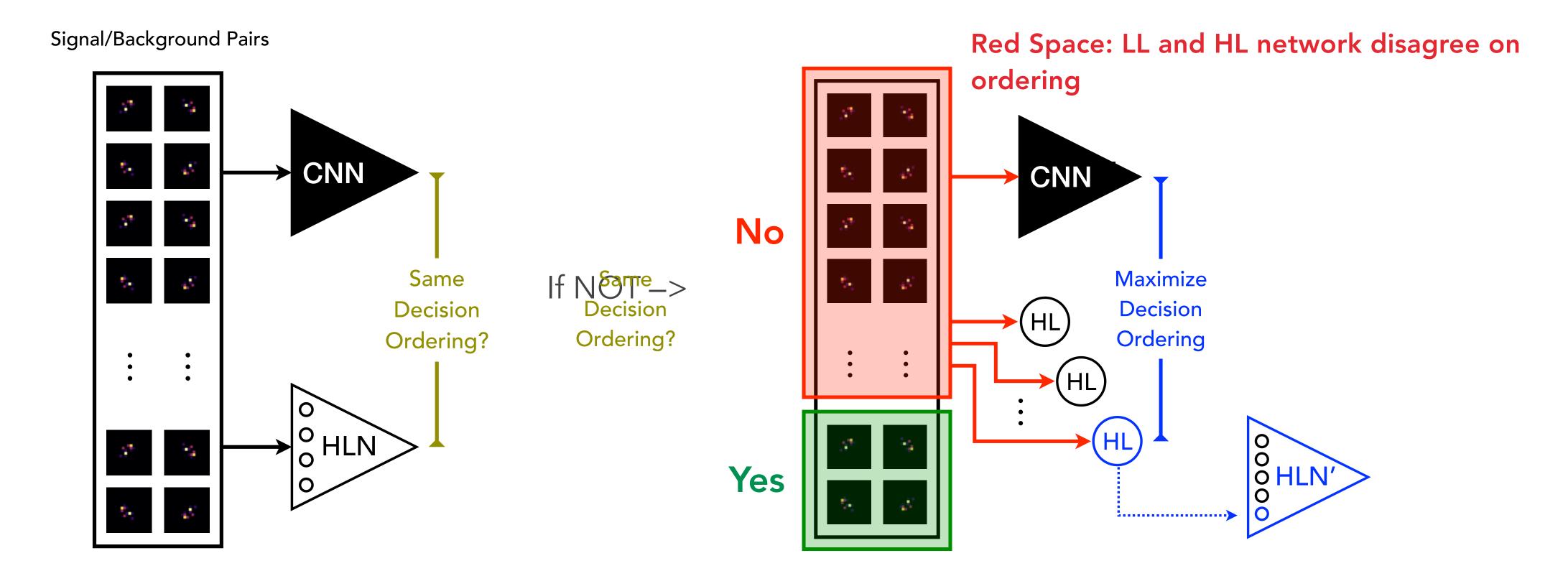




Guided Search

Why is ADO useful?

We can compare NN decision making. Where does the HL network and LL network disagree?



Use ADO to choose EFP that makes similar choices to the LL network in the "differently ordered" red space



Finding a new EFP with Guided Search

We only need 1 new observable to achieve equal performance with the CNN!

Observable	AUC	ADO[CNN, Obs.]
$M_{ m jet}$	0.898 ± 0.004	0.807
$C_2^{\beta=1}$	0.660 ± 0.006	0.584
$C_2^{\beta=2}$	0.604 ± 0.007	0.548
$D_2^{\beta=1}$	0.790 ± 0.005	0.743
$D_2^{\beta=2}$	0.807 ± 0.005	0.762
$ au_2^{\beta=1}$	0.662 ± 0.006	0.600
6HL	0.9504 ± 0.0002	0.971 Original
CNN	0.9531 ± 0.0002	1.000
$7 HL_{black-box}$	0.9528 ± 0.0003	0.971

Original HL + 1 EFP





Which EFP did we pick? $(\kappa=2, \beta=1/2) = \sum_{a,b,c,d=1}^{N} z_a^2 z_b^2 z_c^2 z_d^2 \sqrt{\theta_{ab} \theta_{bc} \theta_{ac} \theta_{ad}}$

Noteworthy details

- EFP is not Infrared-safe ($k \neq 1$)
- $\beta = 1/2$ is probing small-angle behaviour
- Chromatic #3 graph (probing deviations) from 2-prong substructure)
- Chromatic Number = Minimum number of prongs to not vanish

HL

Deep networks can identify gaps where low-level data contains unused information



ML Mapping strategies can capture and translate that information into understandable physics



Questions ?

