

**GT4: Quel est l'apport de la physique nucléaire
à la compréhension de l'astrophysique**

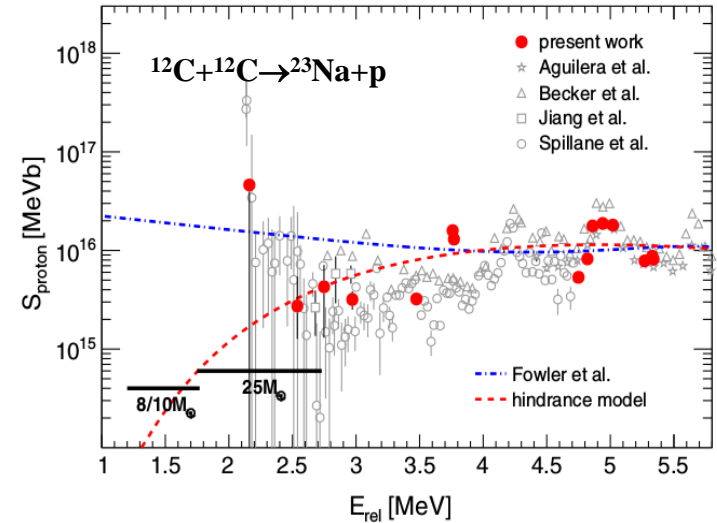
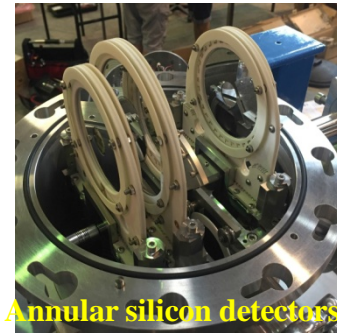
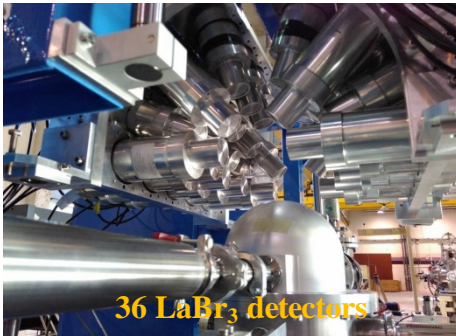
**Short review (Nucleosynthesis)
and perspectives**

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Some of the open questions that were addressed in the 1st joint meeting GT4-GDR RESANET & OG (Nucleosynthesis) and the achievements since

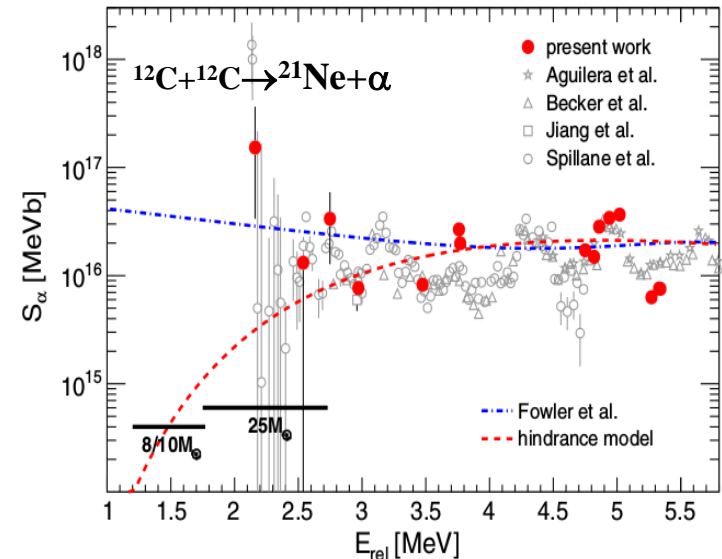
- **Carbon burning in massive stars:** Measurements of $^{12}\text{C}+^{12}\text{C}$ cross-sections using STELLA setup @ Andromède

Fruet+ PRL2020



→ The present results support strongly the fusion hindrance model @ deep sub-barrier energies for the $^{23}\text{Na}+p$ exit channel

→ The present results support the fusion hindrance model @ deep sub-barrier energies for the $^{21}\text{Ne}+\alpha$ & exit channel & the presence of a resonance @ 2.14 MeV for the $^{21}\text{Ne}+\alpha$ exit channel

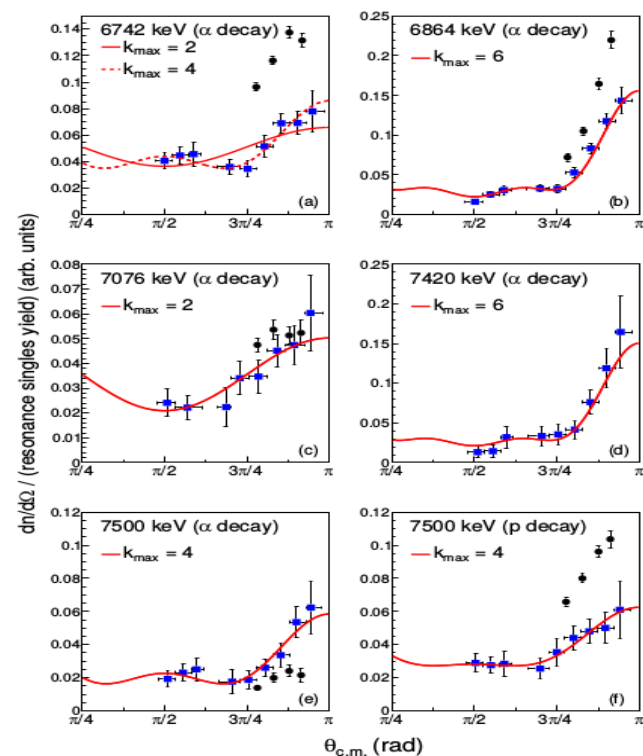
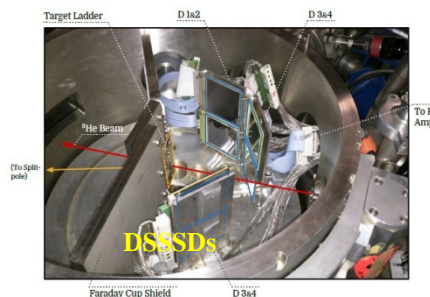
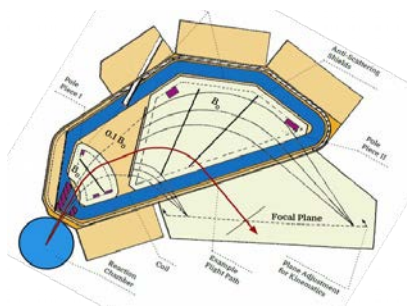


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➤ **Classical novae** → The 4 **key reactions** with significant uncertainties:

- ◆ $^{25}\text{Al}(p,g)^{26}\text{Si}$
- ◆ $^{22}\text{Na}(p,\gamma)^{23}\text{Mg}$ (Fougères+Thesis, analysis nearly finalized, AGATA/SPIDER/VAMOS/GANIL)
- ◆ $^{30}\text{P}(p,\gamma)^{31}\text{S}$ (Meyer+Thesis 2020, SPLIT-POLE/ALTO)
- ◆ $^{18}\text{F}(p,\alpha)^{15}\text{O}$: γ -ray emission ≤ 511 keV Boulay+Thesis2016(GANIL) & Riley+ PRC 2020 (ALTO)

↳ Studied via $^{19}\text{F}(^3\text{He},t)^{19}\text{Ne}(p|\alpha)^{18}\text{F}|^{15}\text{O}$ using SPLIT-POLE/DSSD/ALTO



Riley+ PRC 2020

→ Confirmation of **spin-parity** and **α -branching ratios** in a single measurement (angular correlation analysis)

→ The **role of the sub-threshold resonances** to $^{18}\text{F}(p,\alpha)^{15}\text{O}$ reaction rate is **clarified**

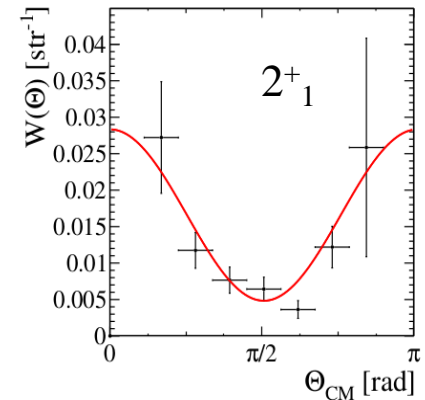
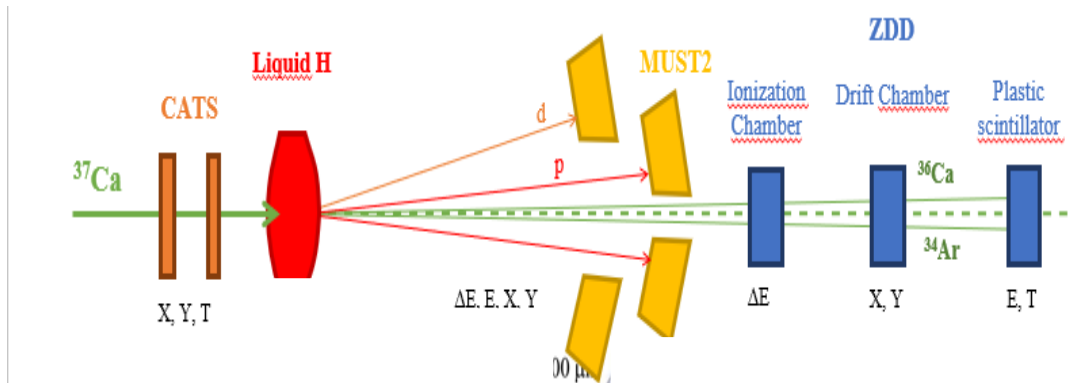
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➤ **X-ray bursts Type I** → Sensitivity studies have shown that only few tens of reactions play an important role.

◆ $^{15}\text{O}(\alpha,\gamma)^{19}\text{Ne}$ Sanchez+ (Ongoing analysis, MUGAST/AGATA/VAMOS/GANIL 2019)

◆ $^{35}\text{K}(p,\gamma)^{36}\text{Ca}$ Lalanne+ PRC2020 to be submitted (Hupin-GT1 talk)

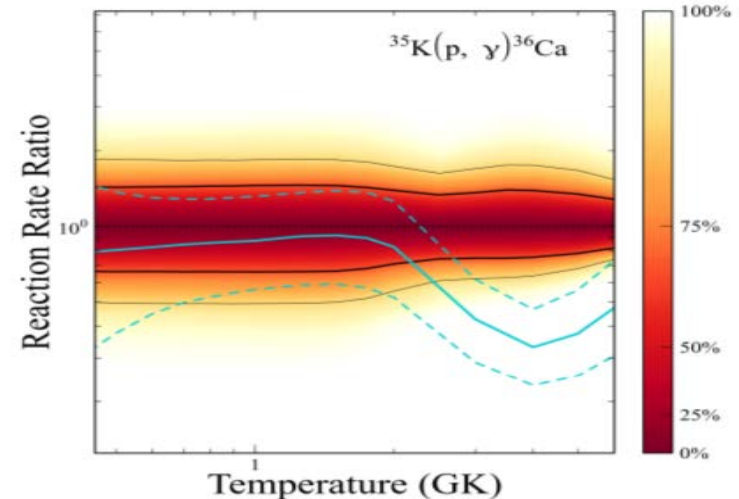
↳ Studied via $^1\text{H}(^{37}\text{Ca},d)^{36}\text{Ca}$ (p) ^{35}K using MUST2@LISE/GANIL



→ New resonances were measured and proton decay branching ratios were determined from the proton angular correlation measurements for the first time

→ The $^{35}\text{K}(p,\gamma)^{36}\text{Ca}$ reaction rate is now experimentally constrained.

→ $^{35}\text{K}(p,\gamma)^{36}\text{Ca}$ will not affect the shape of the X-ray bursts' light curve.



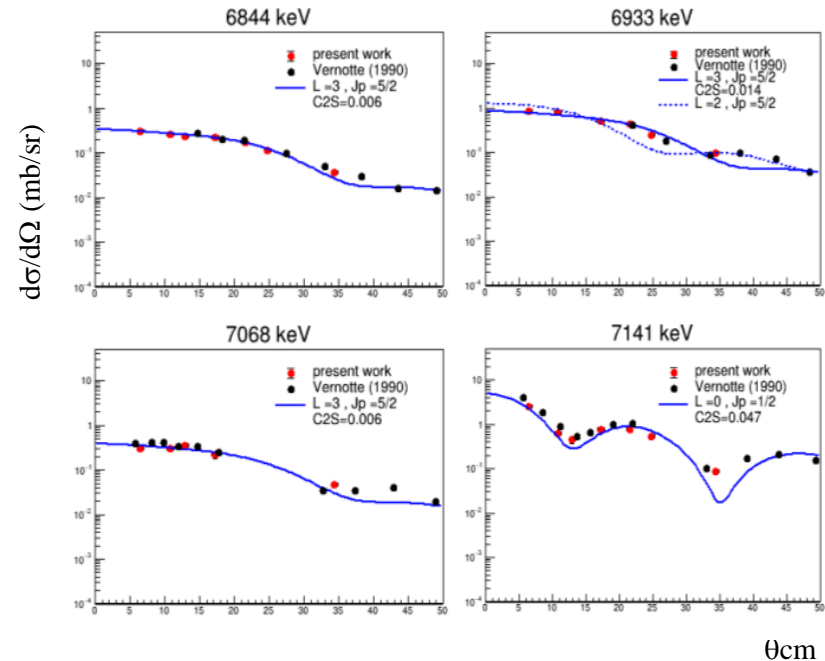
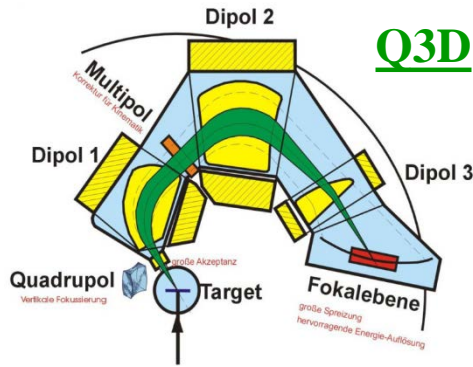
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➤ **Globular clusters** → One of the main question in this field is what is the nature of the 1st generation of stars? need to improve the uncertainty of the reaction rate of 4 reactions.

◆ $^{39}\text{K}(p,\gamma)^{40}\text{Ca}$ Adsley+(ongoing analysis DRAGON/TRIUMF 2019)

◆ $^{30}\text{Si}(p,\gamma)^{31}\text{P}$ Harrouz+PRC paper in progress

↪ Studied via $^{30}\text{Si}(^3\text{He},d)^{31}\text{P}$ using Q3D@MLL



→ Determination of Γ_p & resonance strengths of key resonances for the first time → constrain the $^{30}\text{Si}(p,\gamma)^{31}\text{P}$ reaction rate

Workshop organized with F. Hannachi, Orsay 12-13 june 2019 : What are the possibilities of experiments in nuclear astrophysics with lasers?

- **1 day dedicated to talks :**
 - Direct measurements of cross-sections of astrophysical interest: characteristics and challenges (**M. Heine -IPHC**)
 - Les plasmas lasers, description et diagnostic (**J. santos-CELIA**)
 - Accélérateurs d'ions dans l'interaction laser plasma: état de l'art (**F. Hannachi-CENBG**)
 - Détection d'observables nucléaires en milieu laser (**M. Tariesien-CENBG**)
- **1 day dedicated to discussions :**
 - For which kind of measurements the lasers/plasma could be used?
 - The **electron screening effects**
 - Nuclear reaction on **isomers with short lifetime** (>ns) produced by lasers and multi-incident particle reactions with the ultra intense beams accelerated with lasers
 - Nuclear Properties in plasma environments
 - The technical difficulties
 - Perspectives: - Calculations of XS, production & detection to evaluate the feasibility
 - propose a **“already feasible”** experiment for the next APOLLON call

Some ideas for GT4 workshops/meetings for the next two years

- A 1 day meeting on “[globular clusters nucleosynthesis](#)”
- A 1 day common GT1-GT4 workshop on “[how our nuclear physics knowledge at the dripline impact our understanding of nucleosynthesis processes and supernova neutron star matter](#)”
 - A 2 days workshop on “[X-ray bursts Type I](#):
 1. Probing neutron star physics using thermonuclear X-ray bursts
 2. Nucleosynthesis: is an experimental program at GANIL/SPIRAL1 really possible?”
- A 1 day workshop on [r-process](#) (experiments, theory & modeling)
- A 1 day workshop on [p-process](#) (experiments, theory & modeling)
 - A joint RESANET-OG-GDRs workshop

**Any other ideas of actions are
welcome**