### ENIGMASS – WP 2 GRAVITATIONAL WAVES AND MULTI-MESSENGER ASTROPHYSICS (Virgo, HESS-CTA, AMS-02, Auger, Theory)



Executive Summary

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### Astronomy revolutions in the XXth & XXlth century

#### Looking at the sky beyond the limitation of the visible photon band

- ►~1930's : Radio astronomy
- > | 9 | 1: Extraterrestrial nature of cosmic rays
- > > 1934: Indirect (air shower) techniques
- >1957: Space Era (essential for atmospheric opacity)
- >1960's MeV Neutrinos as astrophysical messengers (Sun, SN 1987A)
- >1970's Cherenkov technique for gamma-ray showers (Whipple...HEGRA...)
- >2000 "Mature" CR direct & indirect probes, with high statistics & redundancy
- Beginning of high-energy neutrino astrophysics (IceCube >2012)
- Beginning of gravitational wave astrophysics (>2015)

#### Multiwawel.-messenger: added value of combining different channels



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#### Enigmass WP 2's activities fit in this broader context and concern:

- Virgo @ LAPP
- Pierre Auger Observatory @ LSPC
- HESS/CTA @ LAPP
- AMS-02 @ LPSC
- Some interdisciplinary topics (e.g. transients)
- Theoretical and phenomenological analyses (LAPTh, LAPP, LPSC)



## LIGO/Virgo:O3 News

- steep increase in event rate in O3a due to the hard work between O2 and O3 to upgrade the detectors.
- detection of 39 candidates in O3a consistent with merger rate inferred from OI+O2.
- Second GW catalog (OI+O2+O3a): 50 events
  - OI (4 months): 3 BBHs
  - O2 (8 m.): 7 BBHs and I BNS
  - O3a (6 m.): 36 BBHs, I BNS, 2 BH + lighter object
  - BBHs with large mass asymmetry (radiation beyond main quadrupolar order, tests of General Relativity)
  - Object of ~2.5-3 M<sub>•</sub> : either heaviest NS or lightest BH!
  - Black hole of 85  $M_{\odot}$  : « forbidden » mass (pair instability)
  - $M_{fin} = 142 M_{\odot}$  : intermediate mass black hole





**Cumulative Count of Events** O1 = 3, O2 =8, O3a =39, Total =50 50 02 01 03a \_ 40 Cumulative #Events 10 200 300 400 500 100 Time (Days) LIGO-G2001862 Credit: LIGO-Virgo Collaboration

# Synoptic view of LAPP activities in Virgo: towards better sensitivity and precision for GW analyses



## New gamma-ray observatory: The CTA Project

2 sites for full & complementary sky coverage 3 telescope sizes to cover a large energy range













## Strong links btw technical & scientific activities





## Ultra-high-energy cosmic rays (UHECRs)



#### Extensive atmospheric air shower

One measures features of the particle shower induced by primaries interacting in the atmosphere They develop differently for different primaries, offering identification capability.

• Understand their origin and nature

• Astrophysics: sources, acceleration, propagation processes. • New fundamental physics at UHE?





### Pierre Auger Observatory & YUHE searches

#### Key role of hybrid detection



#### Interest for $\gamma_{\text{UHE}}$

Unavoidably produced in association to the GZK "propagation-induced"

## Activities $\gamma_{UHE}$ (*D*) LPSC

### Current analyses

- Discriminating variables to identify  $\gamma_{@}$  SD
- Multivariate analysis (PhD thesis defended)
- Study of irreducible background noise

### **R&D** related to AugerPrime

- Shower dependence from key parameters (E, X<sub>max</sub>, direction,  $\# \mu$ )
- Identified universal components enter the MICRO project

### $\gamma$ shower reconstruction via universality principle

- Feasibility and validity for  $\gamma$  searches
- Using complementary measurements from AugerPrime
- On line alert performing analysis (PhD thesis Labex 2020-23: Zoé Torrès)
- Multimessenger exploration of UHECR origin (ANR project)

### 2 staff, I postdoc, I PhD

Signal total = Signal e.m. (pur)



## Galactic Cosmic Rays: AMS-02

High statistics (large acceptance, long exposure time) and high precision data:

- Elemental fluxes from H to Ni
- Isotopic fluxes for light nuclei
- Low energy time dependance

LPSC Involvement in data analysis:

- Elements from H to O
- He, Li, and Be Isotopes



- A longstanding one on CR data analysis
- gamma-ray background based on Fermi-LAT and IACT observations...)
- cross-correlation studies GW / cosmological surveys)

### Theory & pheno studies

No 'dedicated-WP2' theory group at any of the labs, yet regular activity, notably with inter-Lab collaborations:

Some examples:

On multi-wavelength aspects of gamma-rays (source variability, models of diffuse)

• Exploratory studies in multimessenger astrophysics, within a Idex UGA grant, a junior ANR... (e.g.: role in the investigation of gamma-ray excess at Galactic center,



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### Thank you for your attention!

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