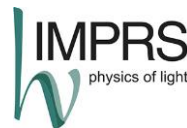


Event reconstruction for KM3NeT/ORCA using convolutional neural networks

Stefan Reck,
GDR meeting, 24.11.2020



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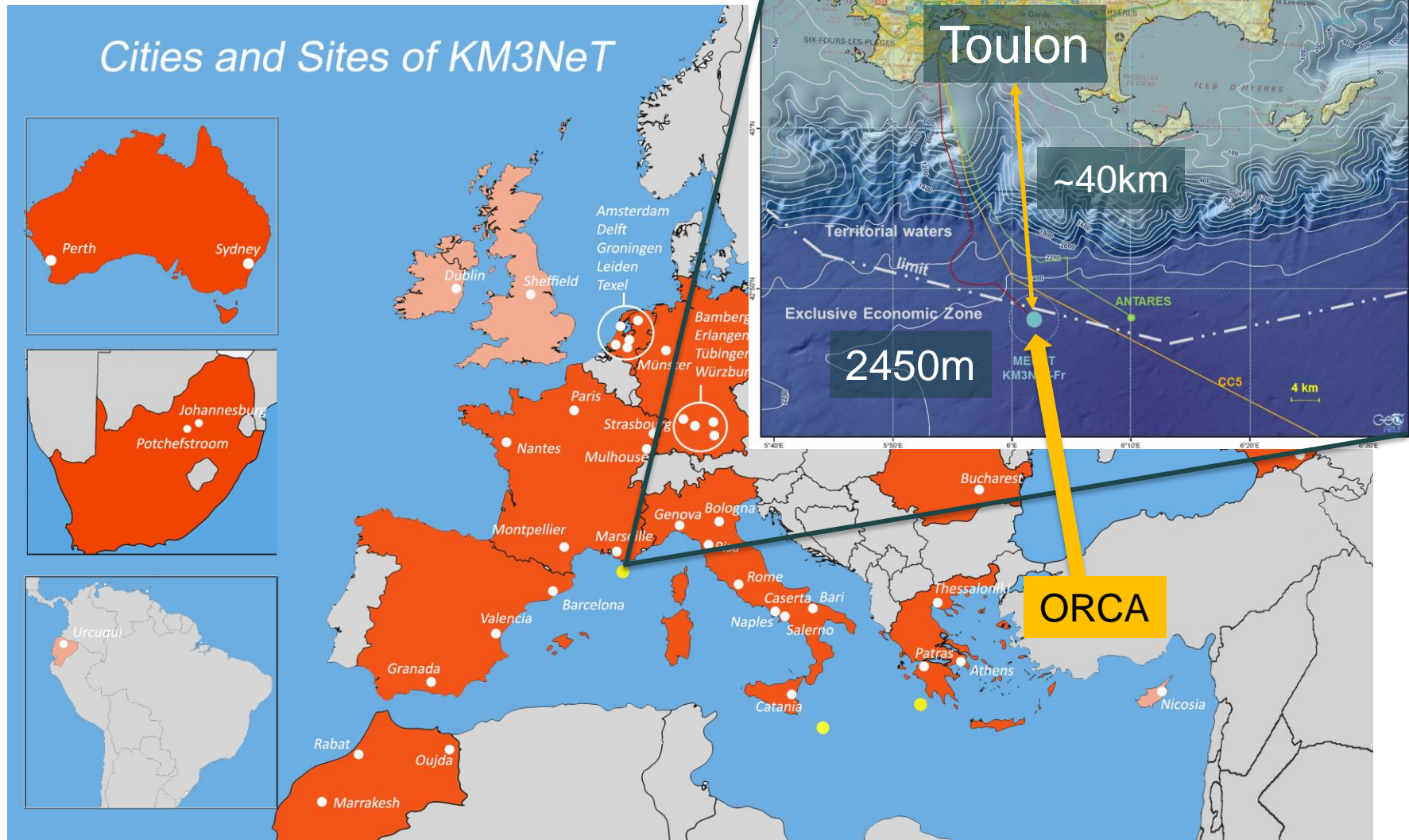
- This talk is based on the paper published in JINST:
<https://arxiv.org/abs/2004.08254>
by primary authors **Michael Moser** and **Thomas Eberl**

- First application of a deep convolutional network in water-Cherenkov detector

KM3NeT/ORCA: a neutrino detector



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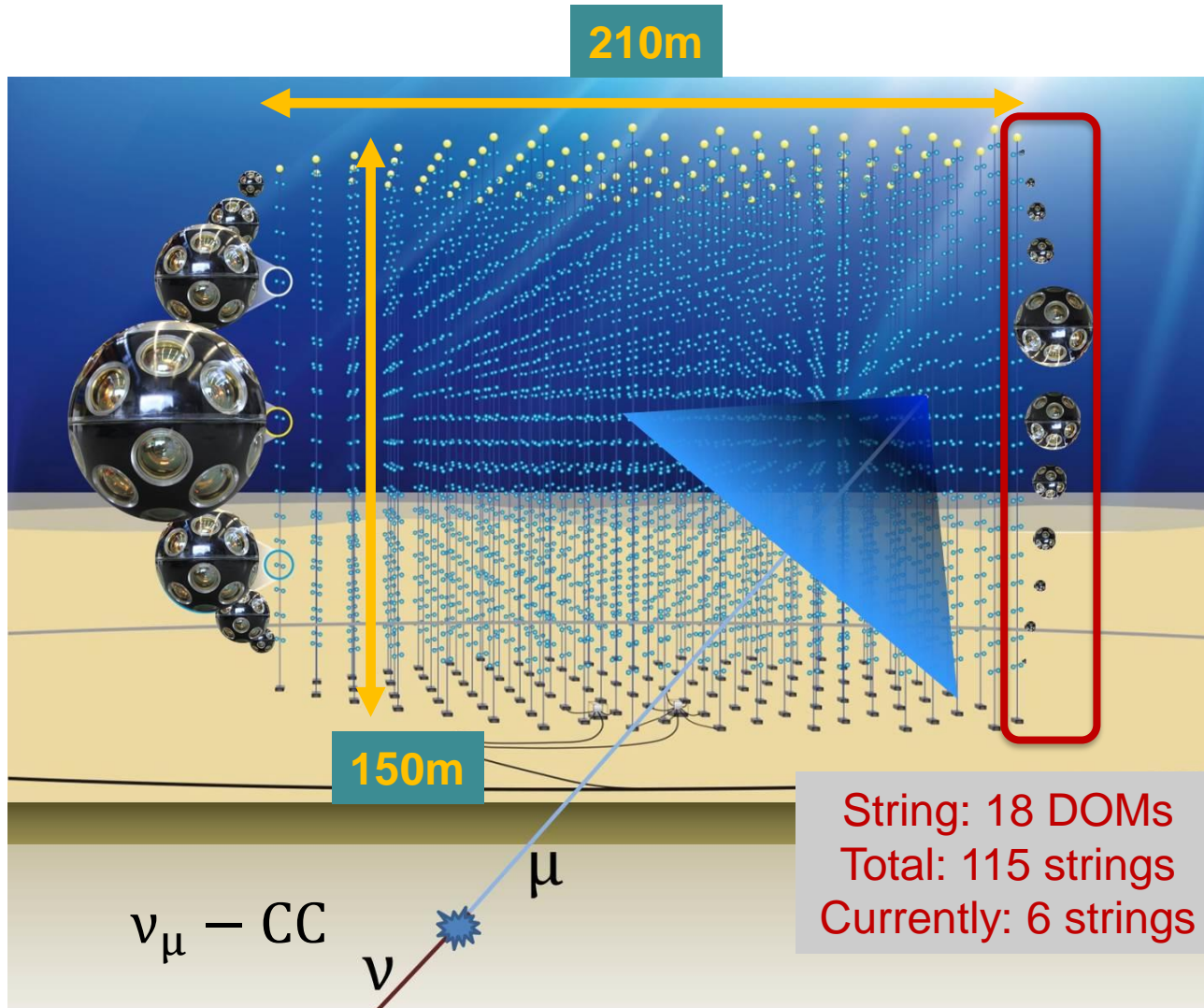


ORCA: a deep sea neutrino detector



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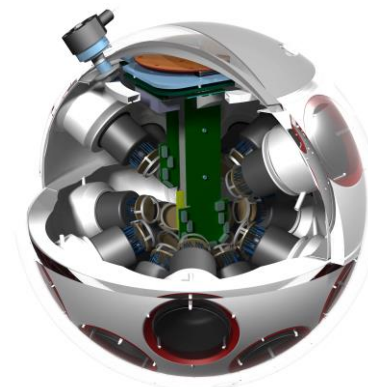
Aim: measure atmospheric neutrinos with energies from 1-100 GeV



ORCA DOM

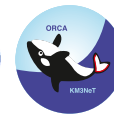


43 cm



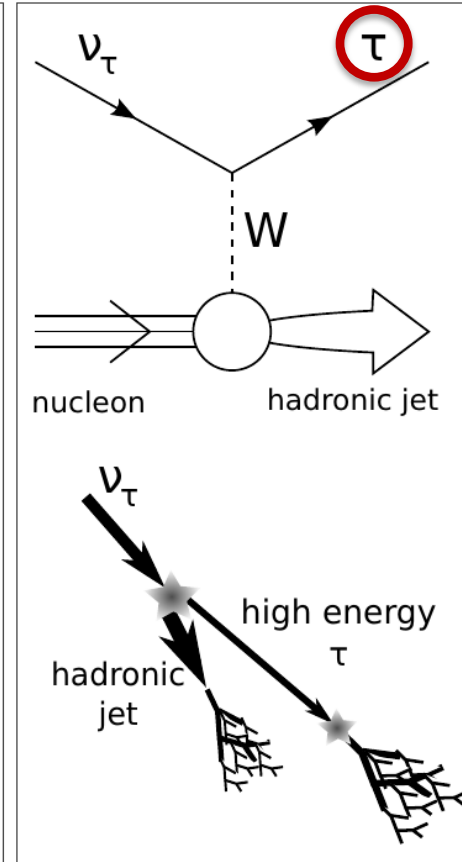
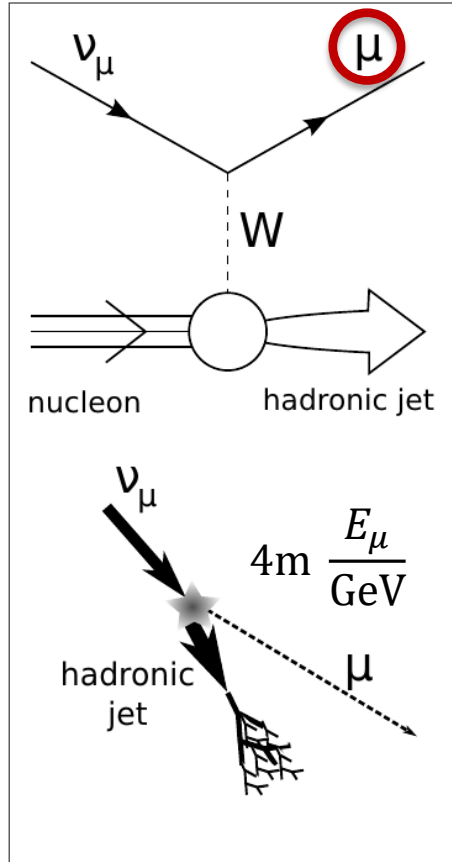
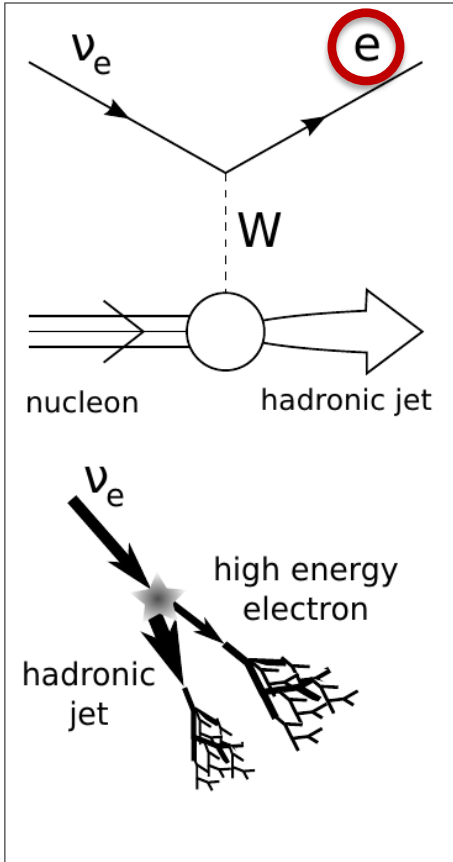
31 PMTs

Neutrino interactions

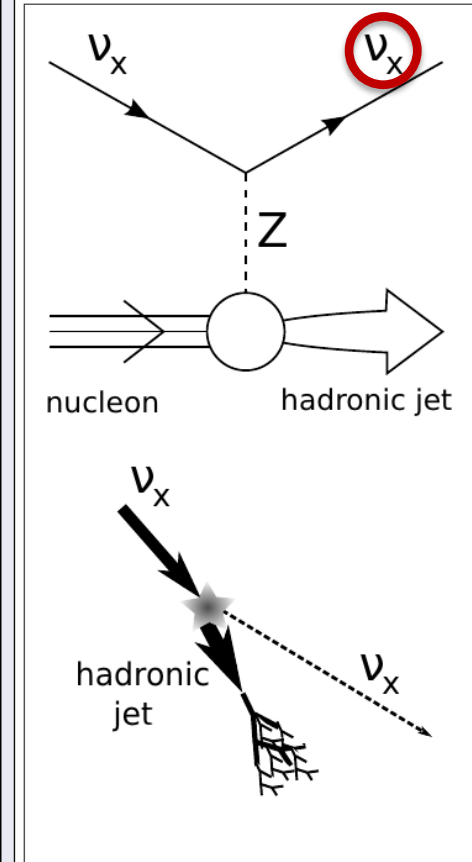


Charged Current

Credit: J. Tiffenberg, NUSKY11



Neutral Current



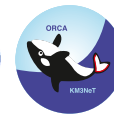
$\nu_e - CC$

$\nu_\mu - CC$

$\nu_\tau - CC$

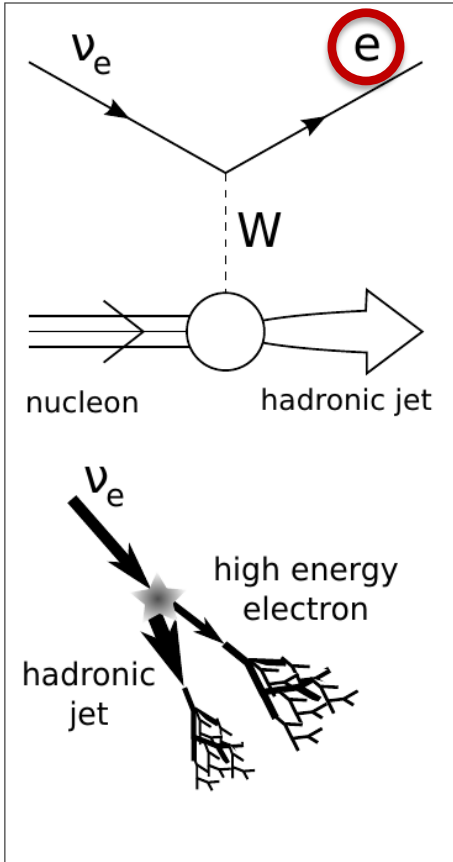
$\nu - NC$

Neutrino interactions

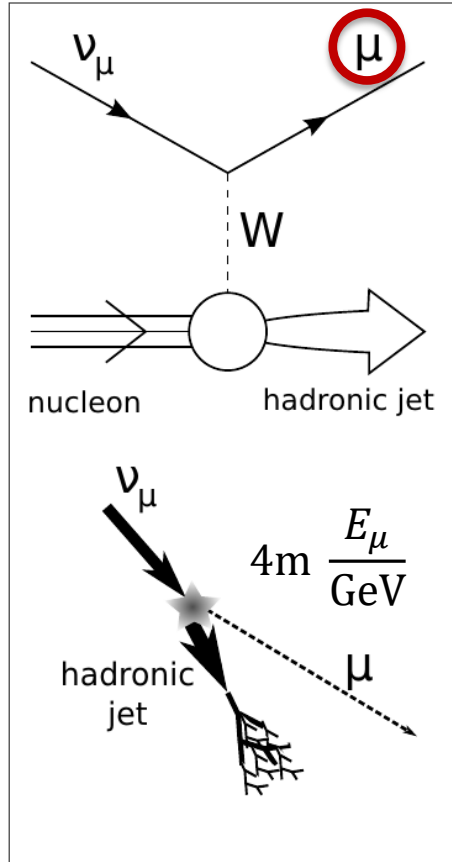


Charged Current

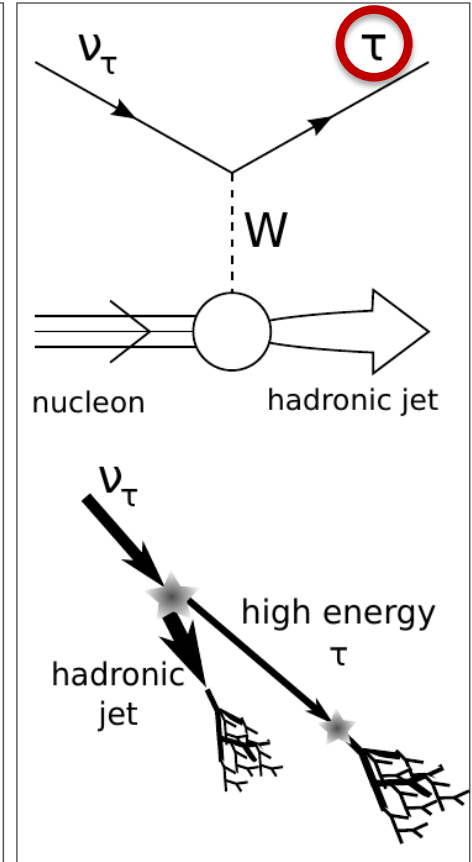
Credit: J. Tiffenberg, NUSKY11



shower-like

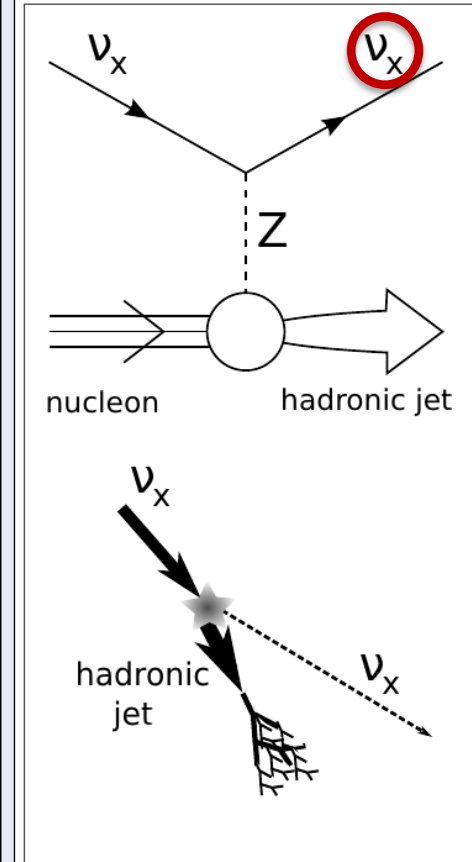


track-like



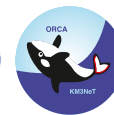
shower-like (83%)
track-like (17%)

Neutral Current

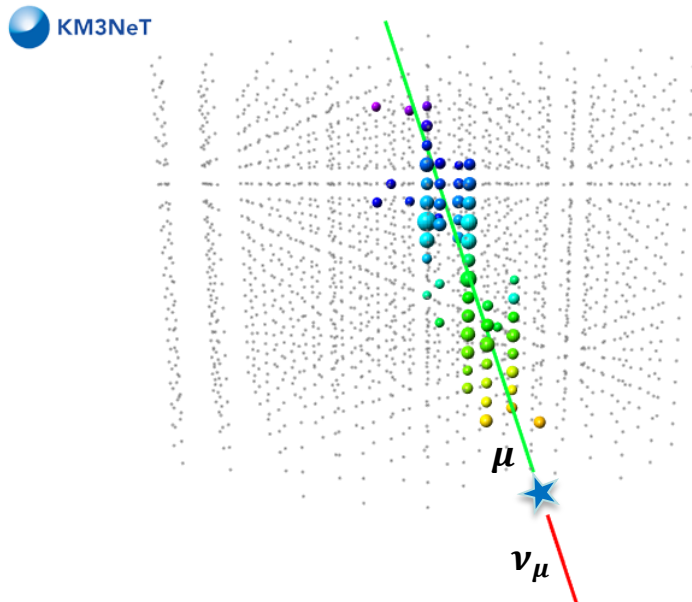


shower-like

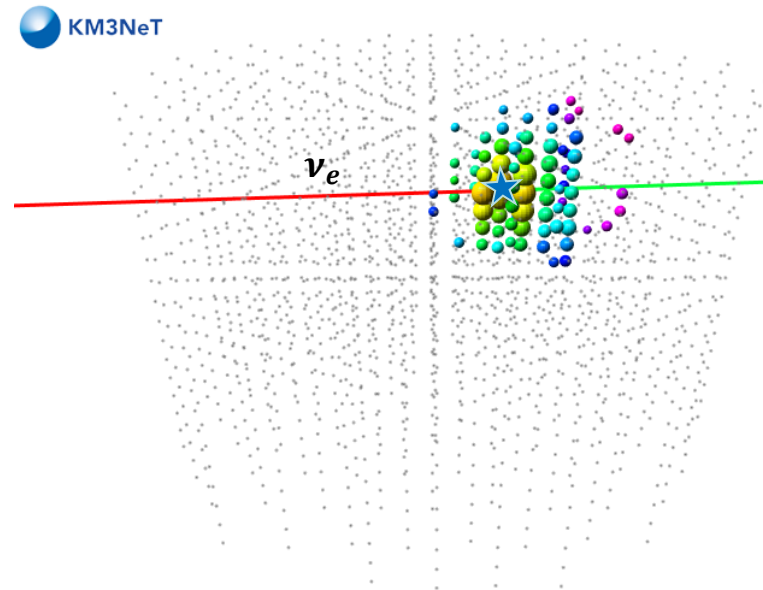
Event topologies

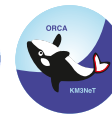


Up-going ν_μ – CC track-like event



ν_e – CC shower-like event



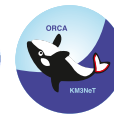


Reconstruction of neutrinos from low-level detector data

1. Discriminate **background** from neutrinos
2. Classify neutrino events into **track-** and **shower-**like events
3. Reconstruct neutrino properties like **energy** and **direction**

This work:

Perform this based on so-called **deep learning** techniques.



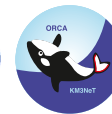
Reconstruction of neutrinos from low-level detector data

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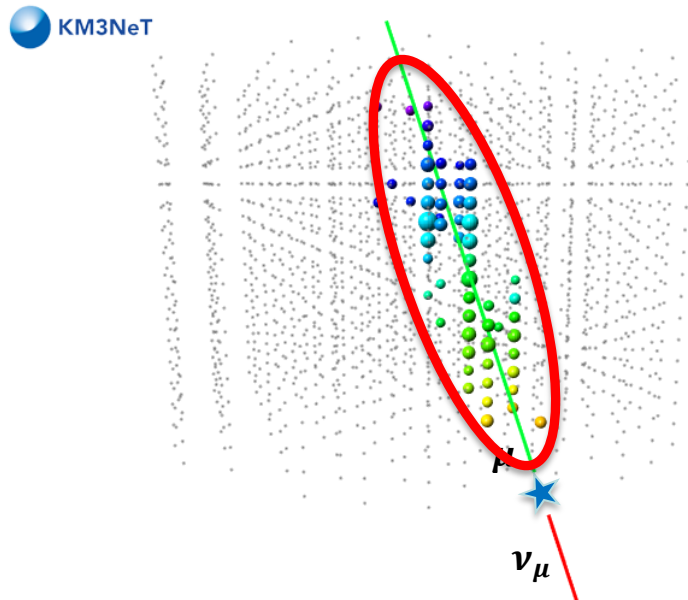
Perform this based on so-called **deep learning** techniques.

Event topology classification

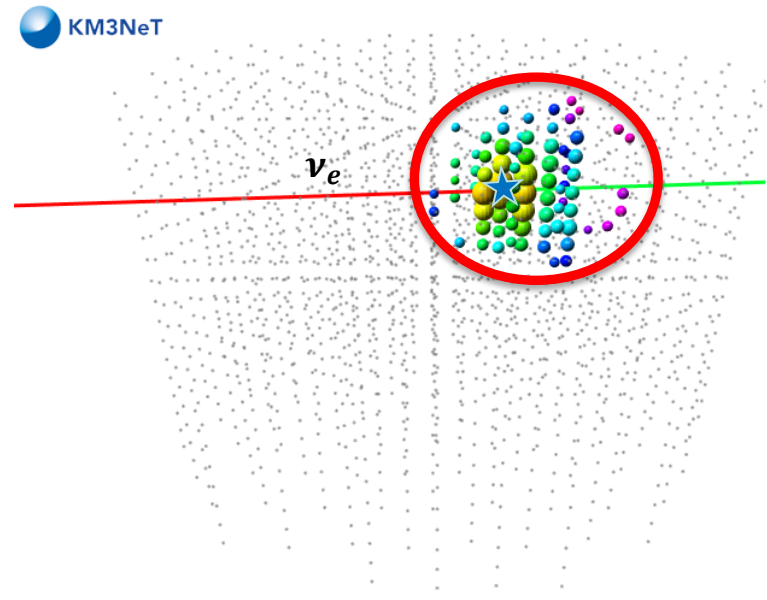


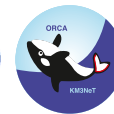
- How to distinguish tracks and showers?
→ Classical way: come up with characteristic features, e.g. **Sphericity**

Up-going ν_μ – CC **track**-like event

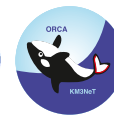


ν_e – CC **shower**-like event





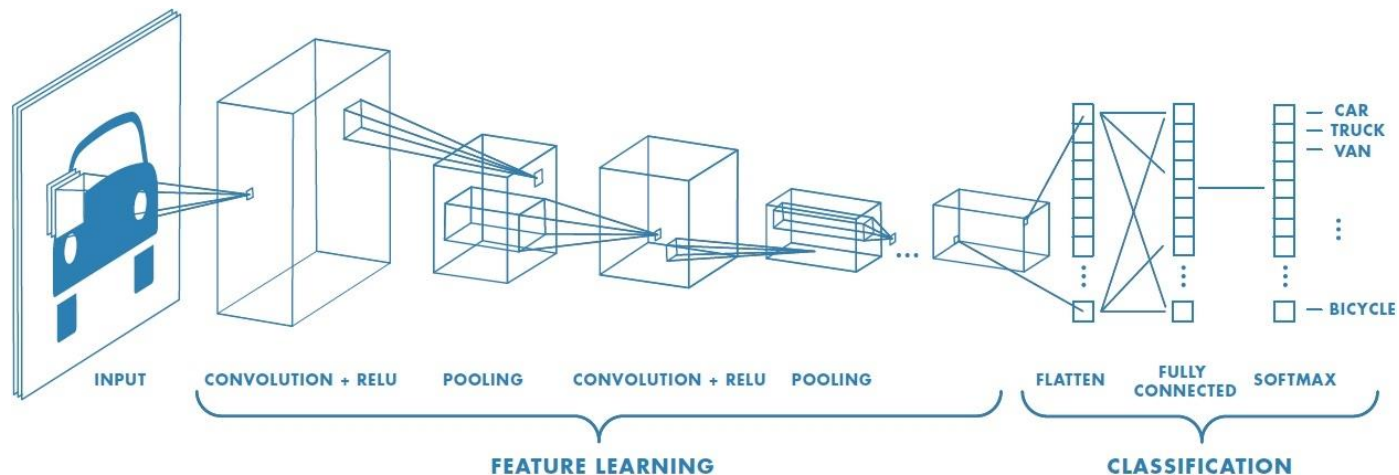
- Hand features to a machine learning based classifier
- Standard ORCA background classifier uses a Random Forest (RF) to combine the features
- Machine learning algorithm is „trained“ with simulations
- Problem: **feature design is not easy** and maybe we missed some good features?



- Solution: let an algorithm learn the features by itself based on simulations of low-level detector data
- Possible with recently emerging machine learning algorithms like deep neural networks, also called ***deep learning*** techniques

- How can we apply deep learning methods to ORCA data?

Successful model architecture in image recognition:
Convolutional neural networks (CNNs)



Simplified working principle of a CNN

Source: https://www.mathworks.com/solutions/deep-learning/convolutional-neural-network/_jcr_content/mainParsys/band_copy_copy_14735_1026954091/mainParsys/columns_1606542234_c/2/image.adapt.full.high.jpg/1575485682772.jpg

- ORCA data can be interpreted to be 5D (XYZ, T, 31 PMT channels)

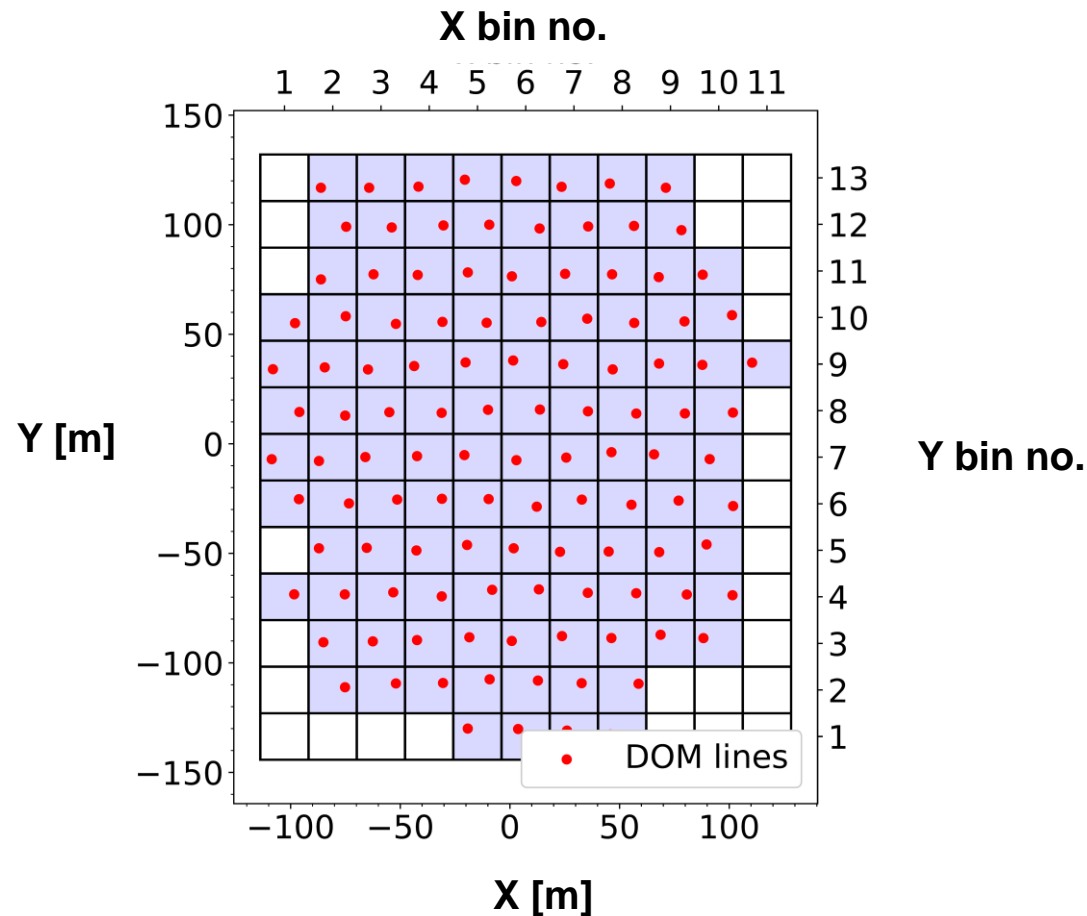


Most Deep Learning frameworks, like Google's Tensorflow, only support 4D input (colored videos)

- CNN expects image-like, pixelated input
 - Bin XYZT dimensions to get pixelated **event images**

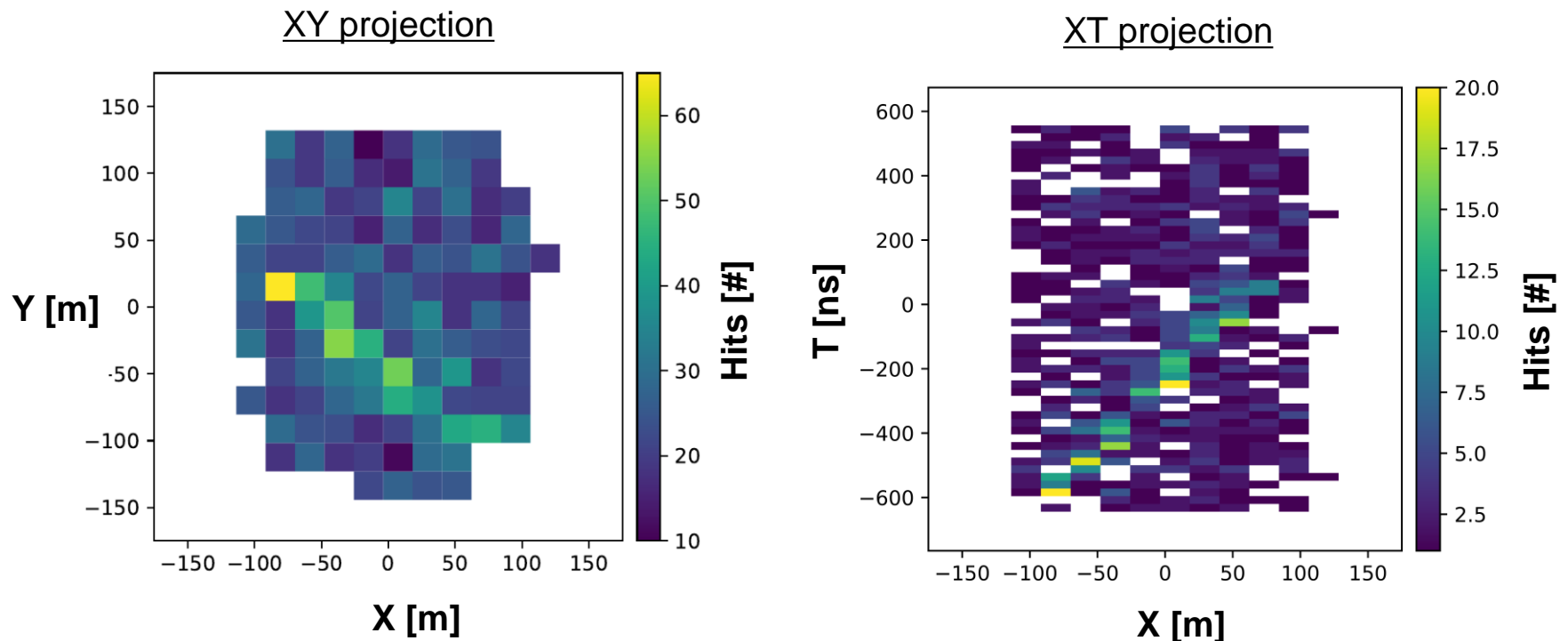
- Spatial binning (XYZ): 1 DOM / 3D pixel

Top view (XY) of the ORCA detector

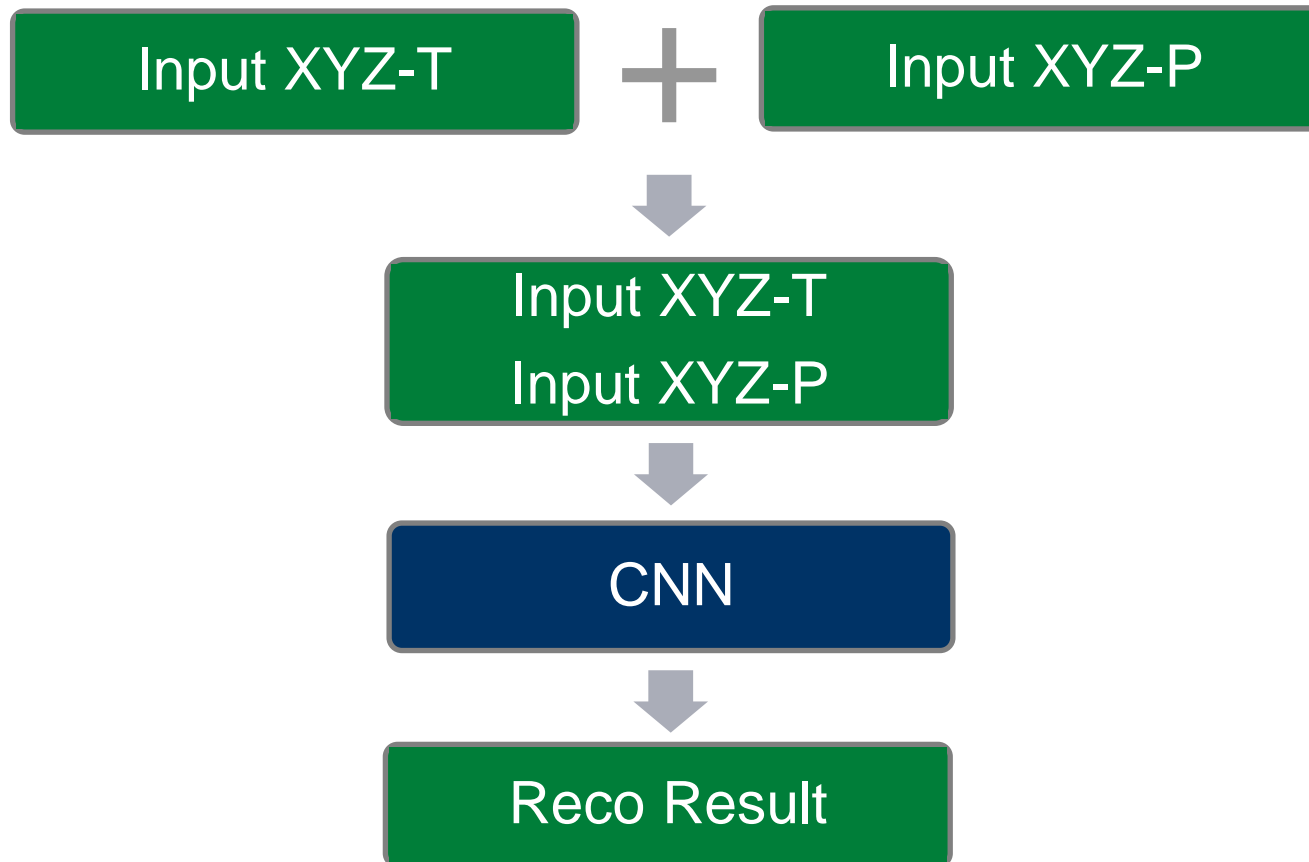


- Time binning:
 - 100 bins $\rightarrow \sim 9.5\text{ns} / \text{bin}$
 - Number of bins limited by computational cost

XY & XT projection of the 4D XYZT „image“ for an 80 GeV ν_μ - CC event



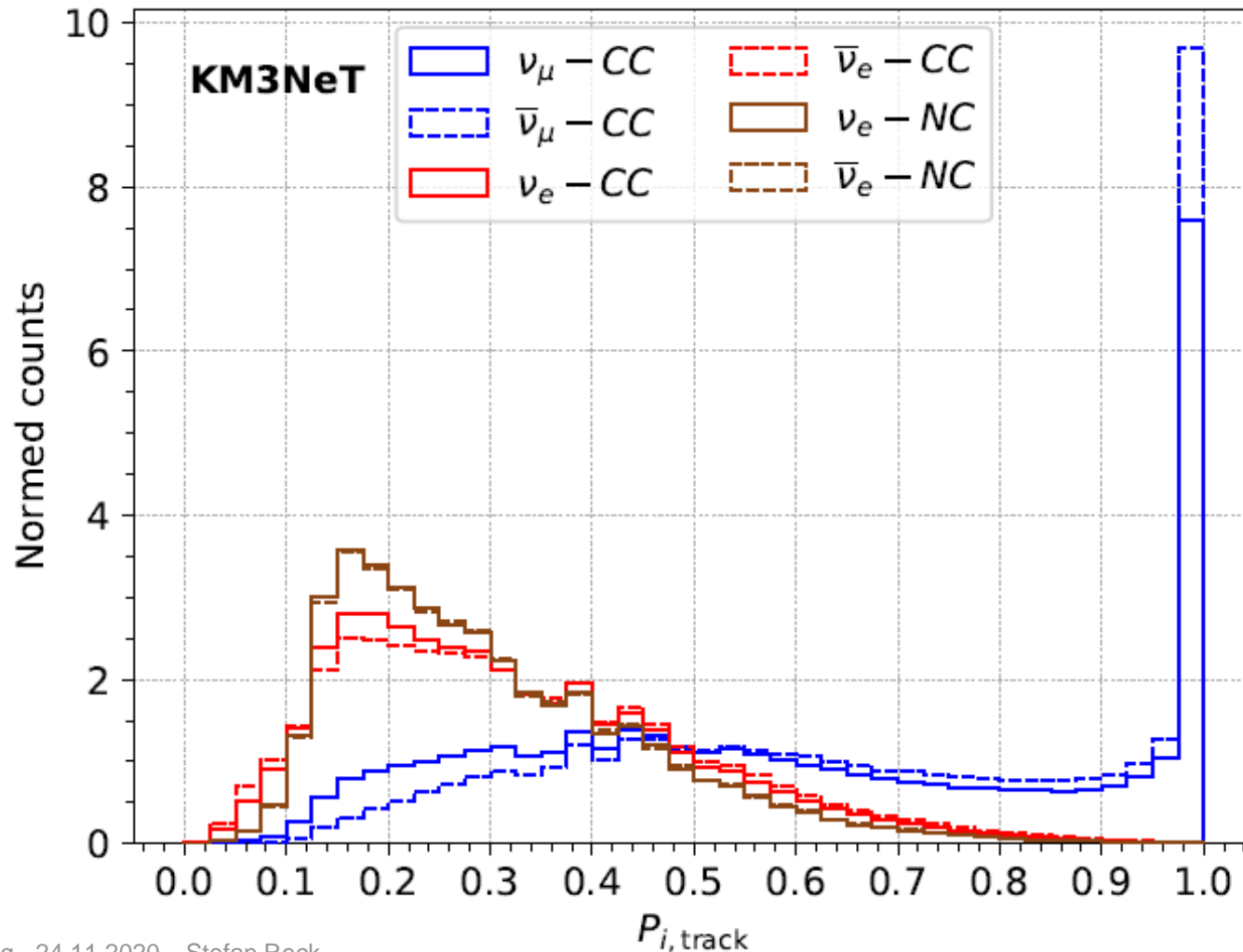
- How to use convolutions on our 5D XYZT-P input?
→ Supply network with two **4D projections**



Event topology classification



- Network output: Probability of event being a track

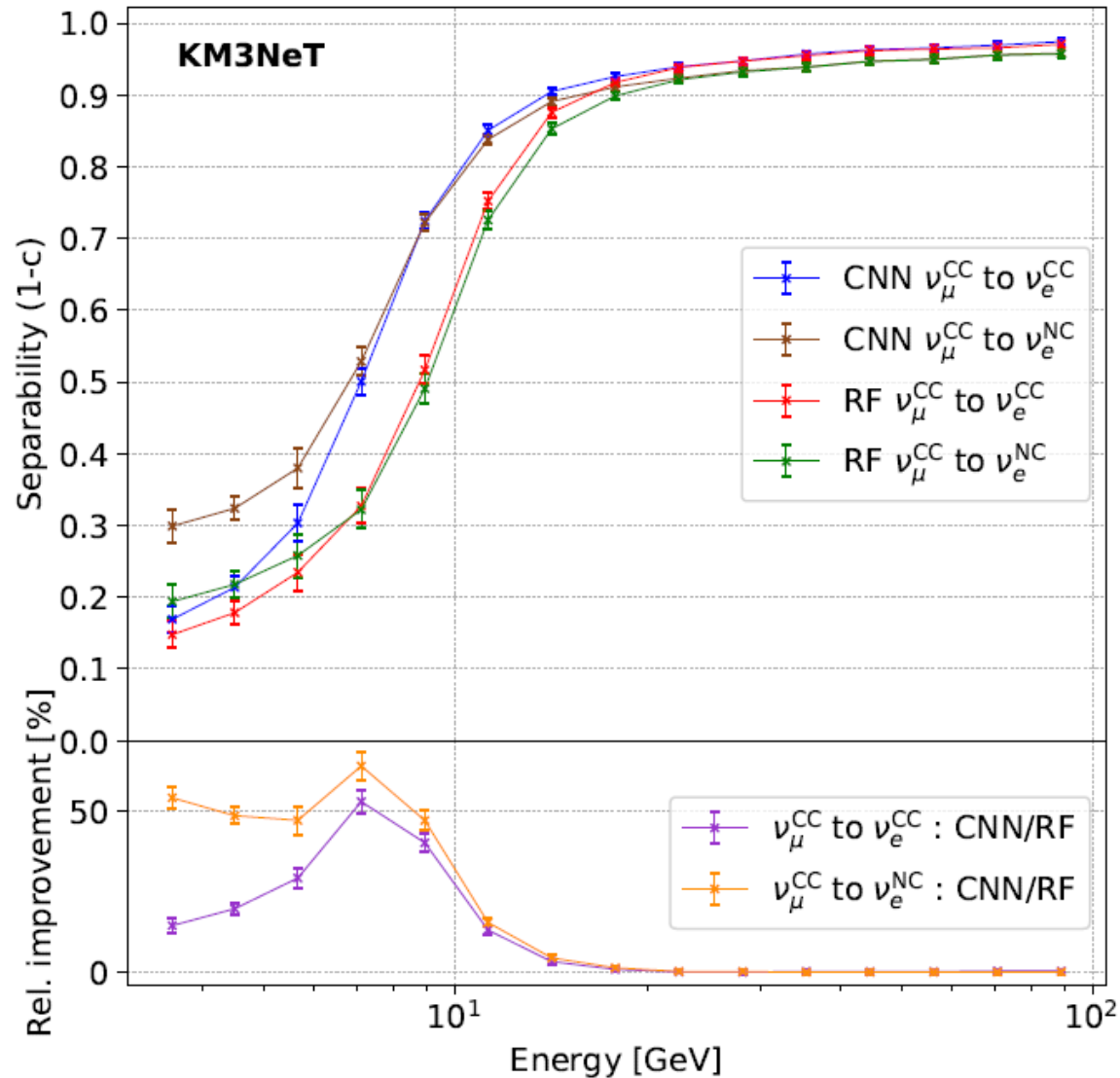


Event topology classification



- Separability between track and shower

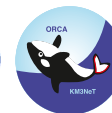
Neural network
(this work)
VS
random forest
(standard method)



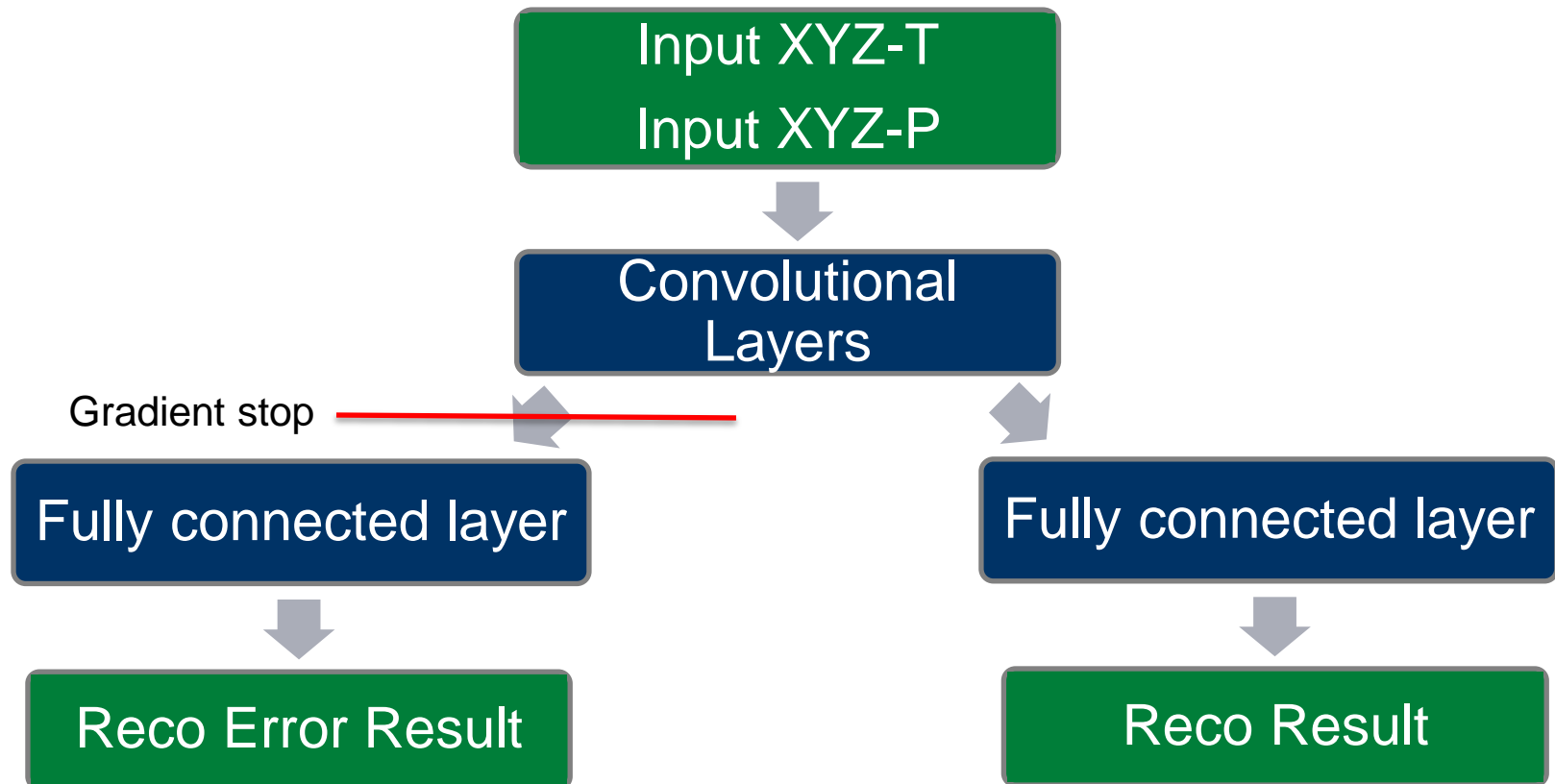


direction reconstruction

Regression results – direction



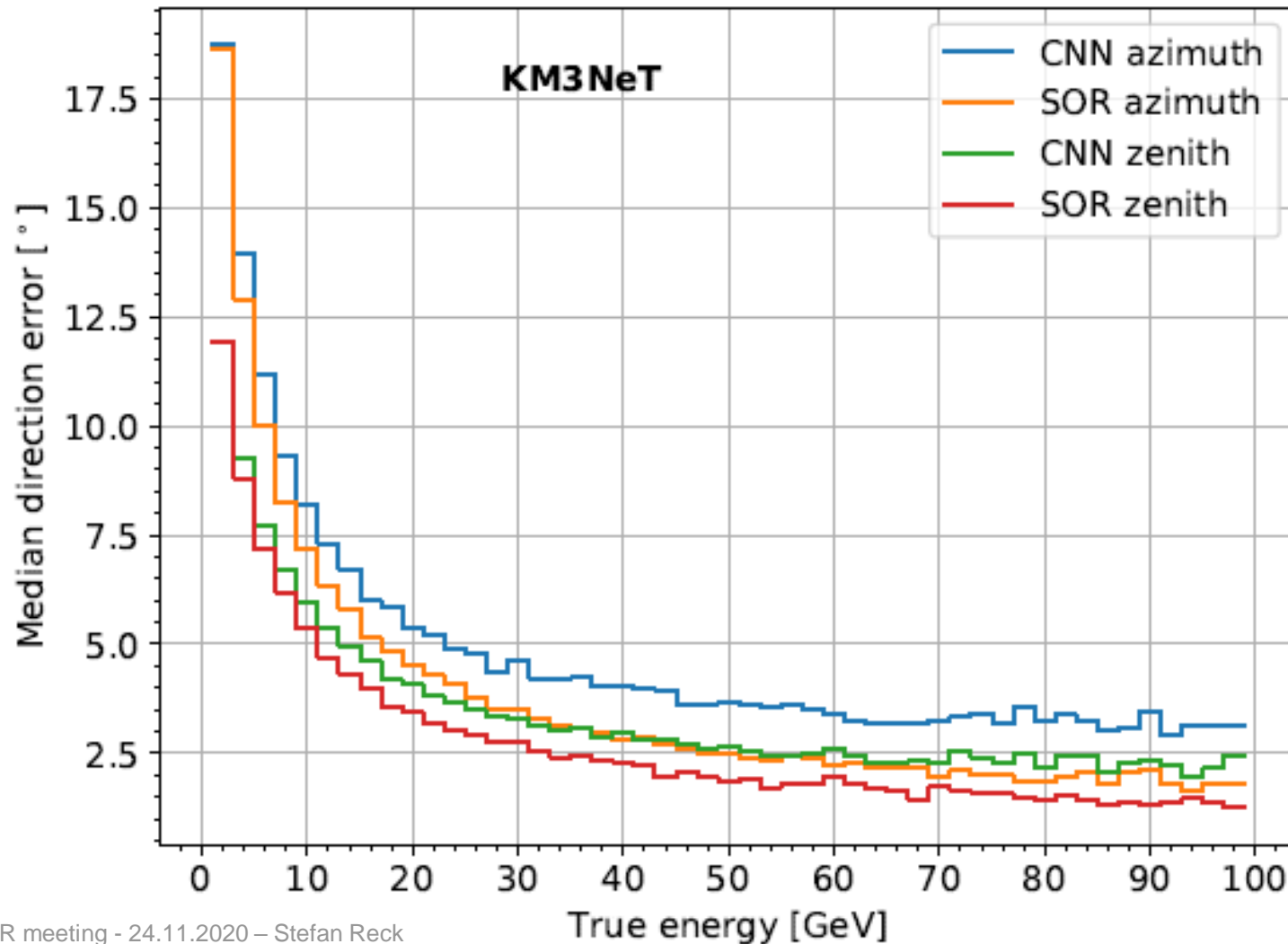
- Also predicts the standard deviation of any regression variable
- Done by adding a second dense network at the end



Regression results – direction



- CNN reco is comparable to classical likelihood based reco

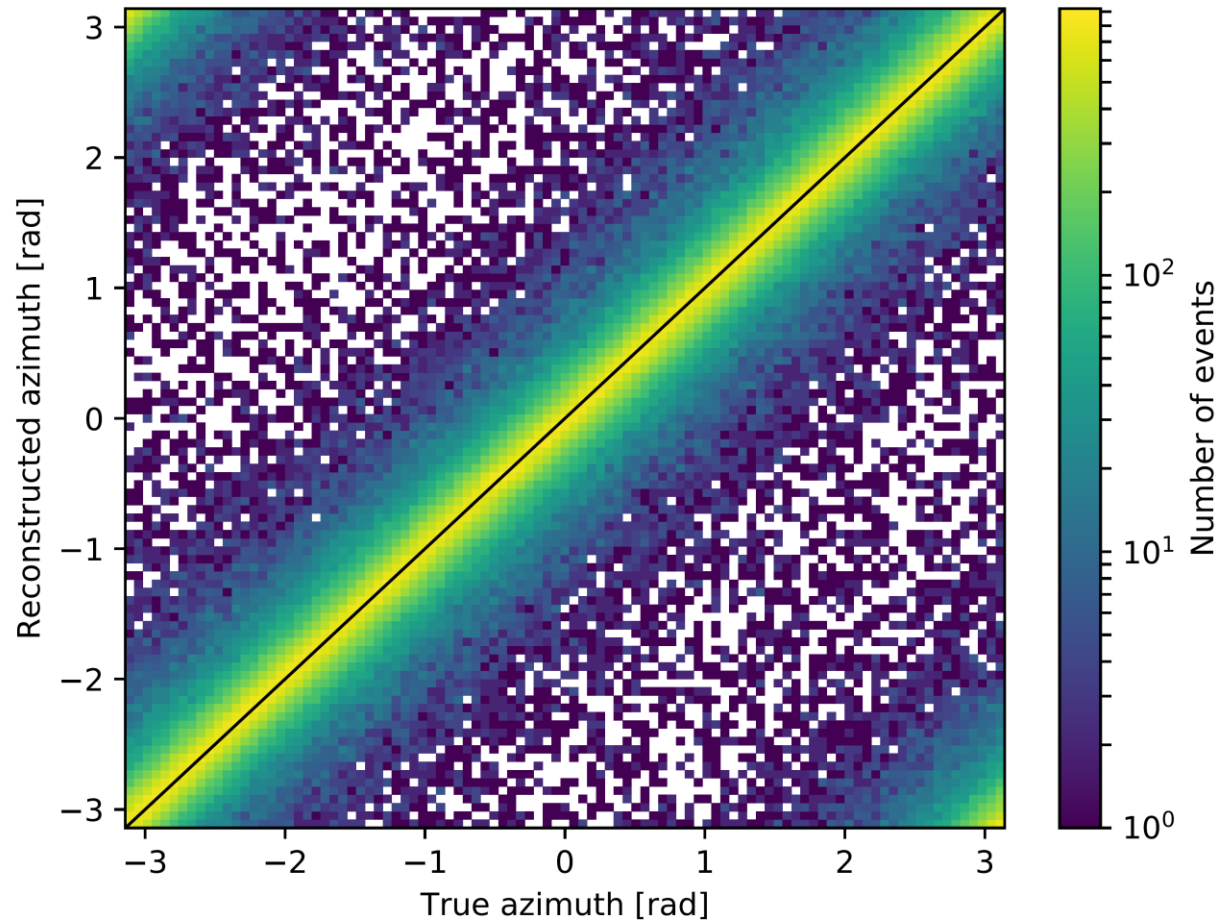


Regression results – direction



- can do cuts using network's error estimation

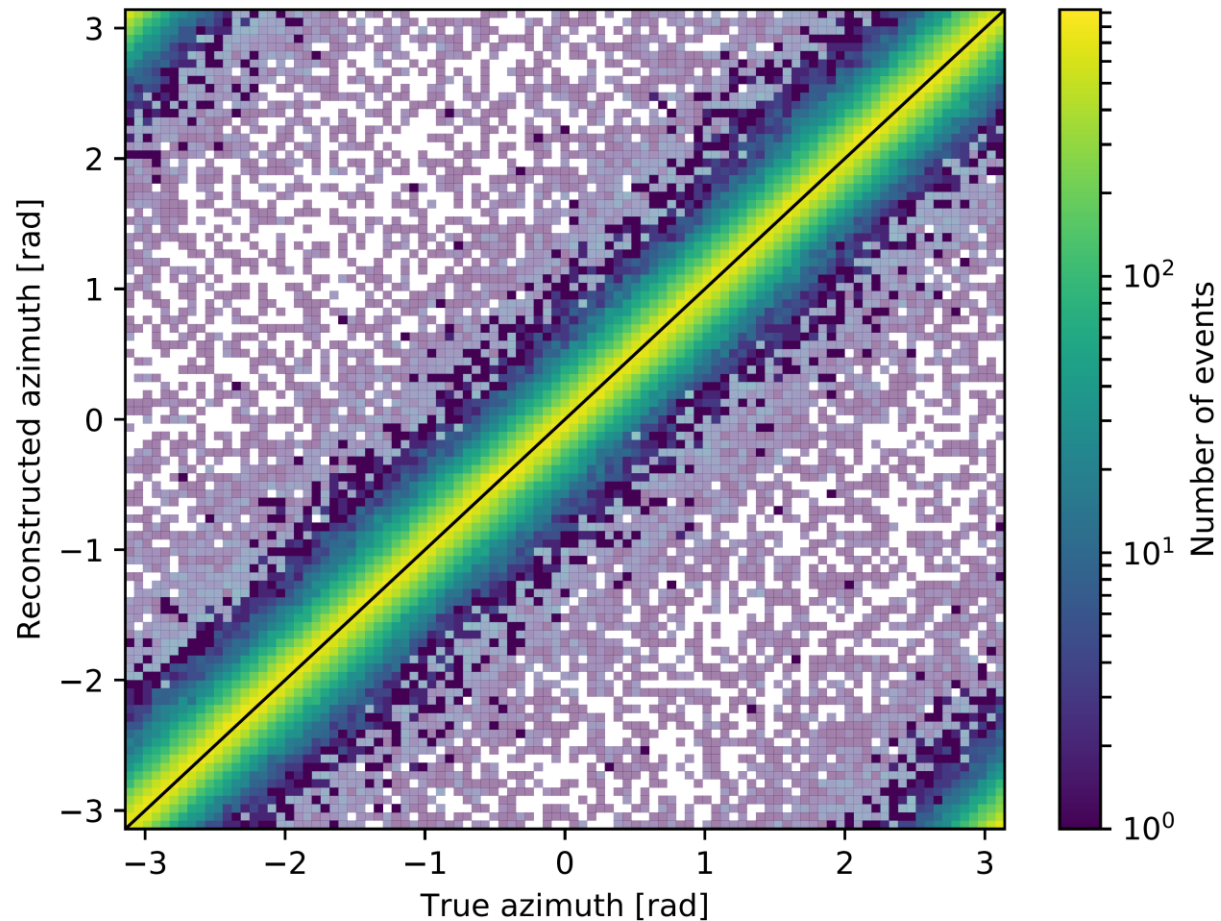
OrcaNet: Track like ($\nu_\mu - CC$), 100% of total events



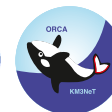
Regression results – direction



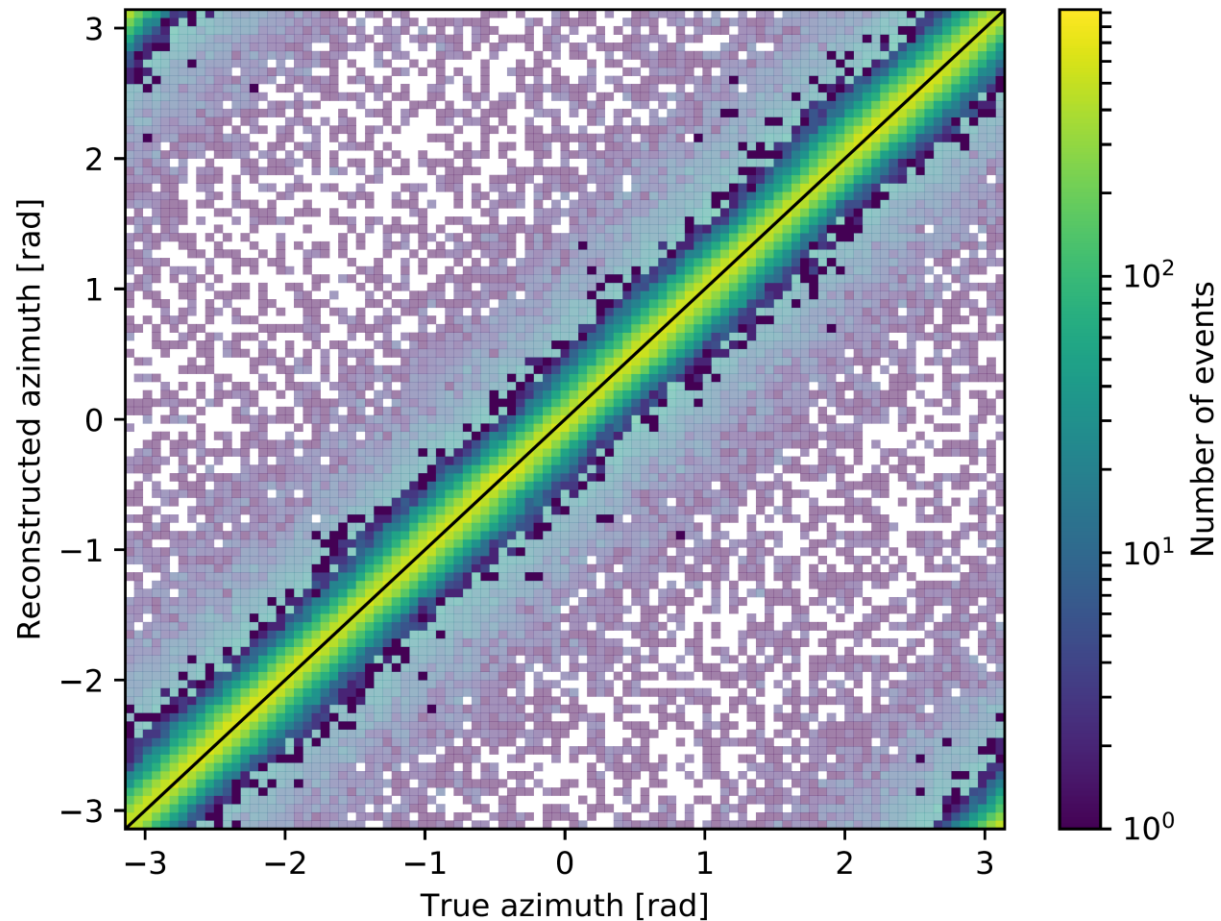
OrcaNet: Track like ($\nu_\mu - CC$), 80% of total events



Regression results – direction



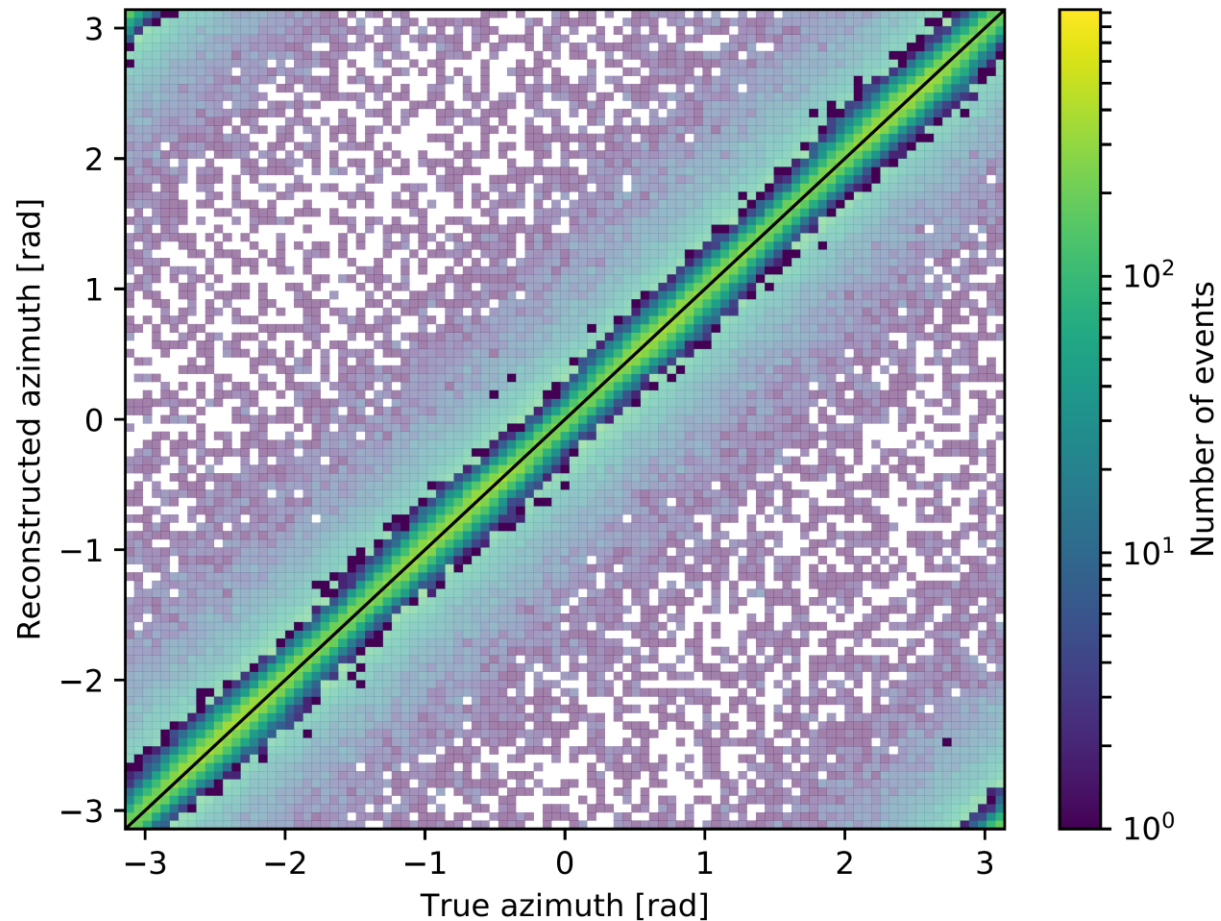
OrcaNet: Track like ($\nu_\mu - CC$), 50% of total events



Regression results – direction



OrcaNet: Track like ($\nu_\mu - CC$), 20% of total events

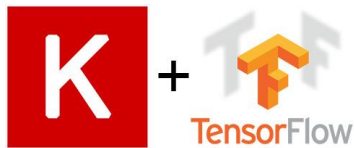


- **Background classification**
significant improvement on rejection of atmospheric muons, same performance for random noise
- **Track-shower classification**
significant improvement on the classification accuracy
- **Regression of energy, direction and neutrino interaction point**
 - Significant improvement for the energy reconstruction of track-like events
 - Competitive performance for other variables
 - error estimation for each reconstruction



backup

- Network needs to be trained in order to distinguish the background
 - 43 Mio. simulated events used for the training
 - Consists of neutrinos, atmospheric muons, random noise events
 - CNN architecture with 10 layers
 - 3 million free parameters
 - Trained for about 1.5 weeks



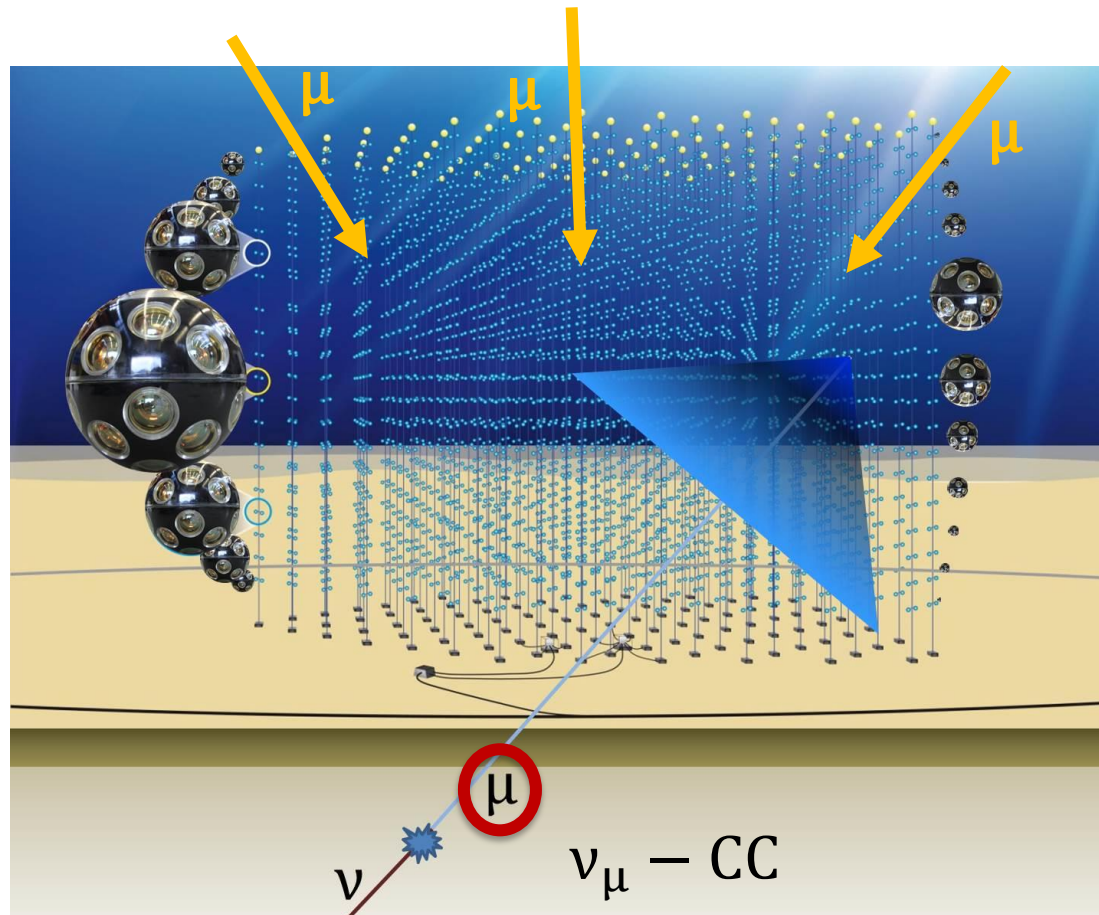


background rejection

ORCA backgrounds



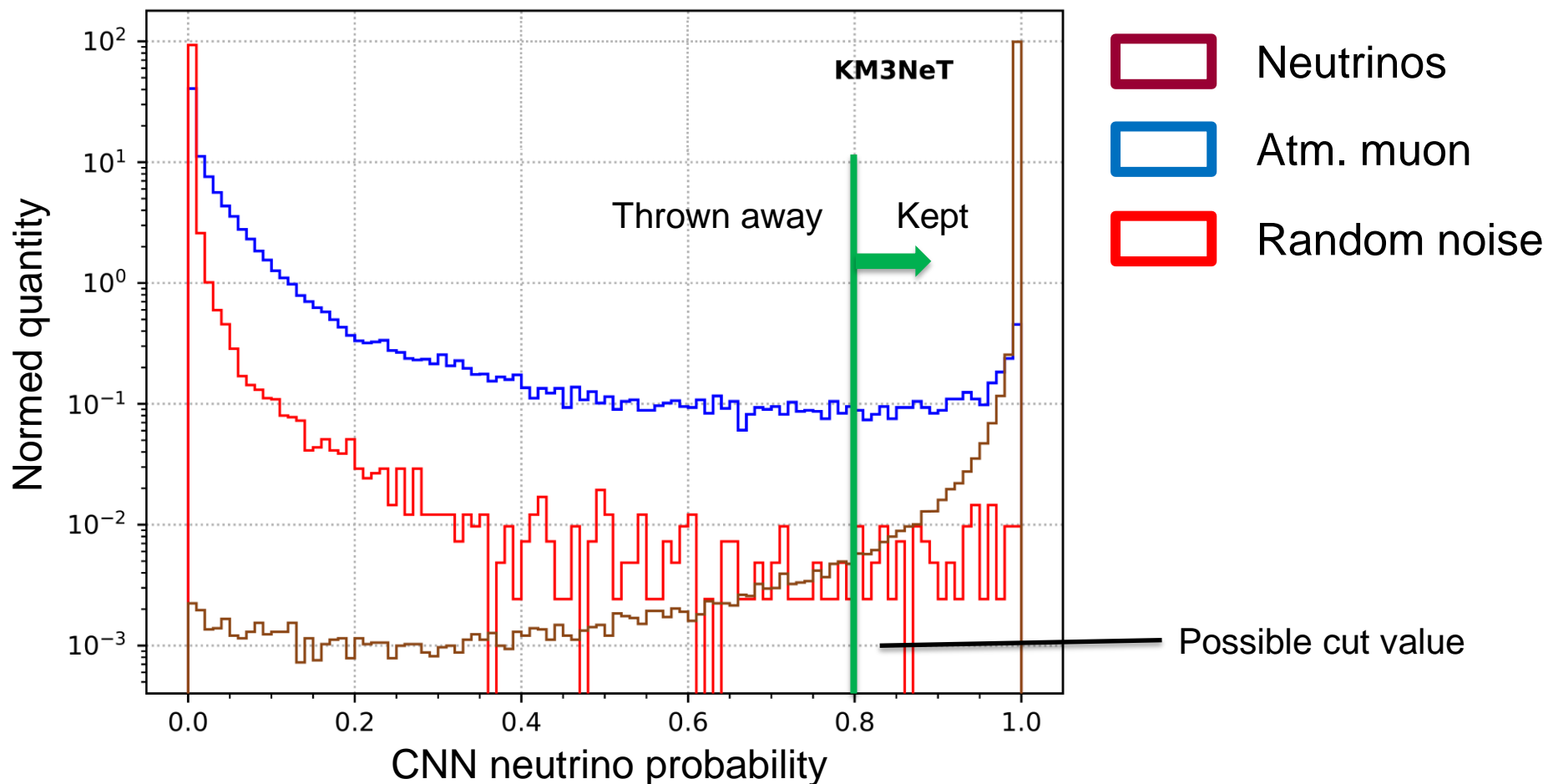
- Two types of backgrounds producing photons in the deep sea:
 1. **Atmospheric muons** passing the detector from above
 2. **Random noise**, by K-40 beta decays and bioluminescent animals



Background separation performance



- Output of the CNN on never-before-seen test data
- Reject backgrounds by cutting on the predicted neutrino probability



Background separation performance



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- Assuming an oscillated neutrino and atmospheric muon flux

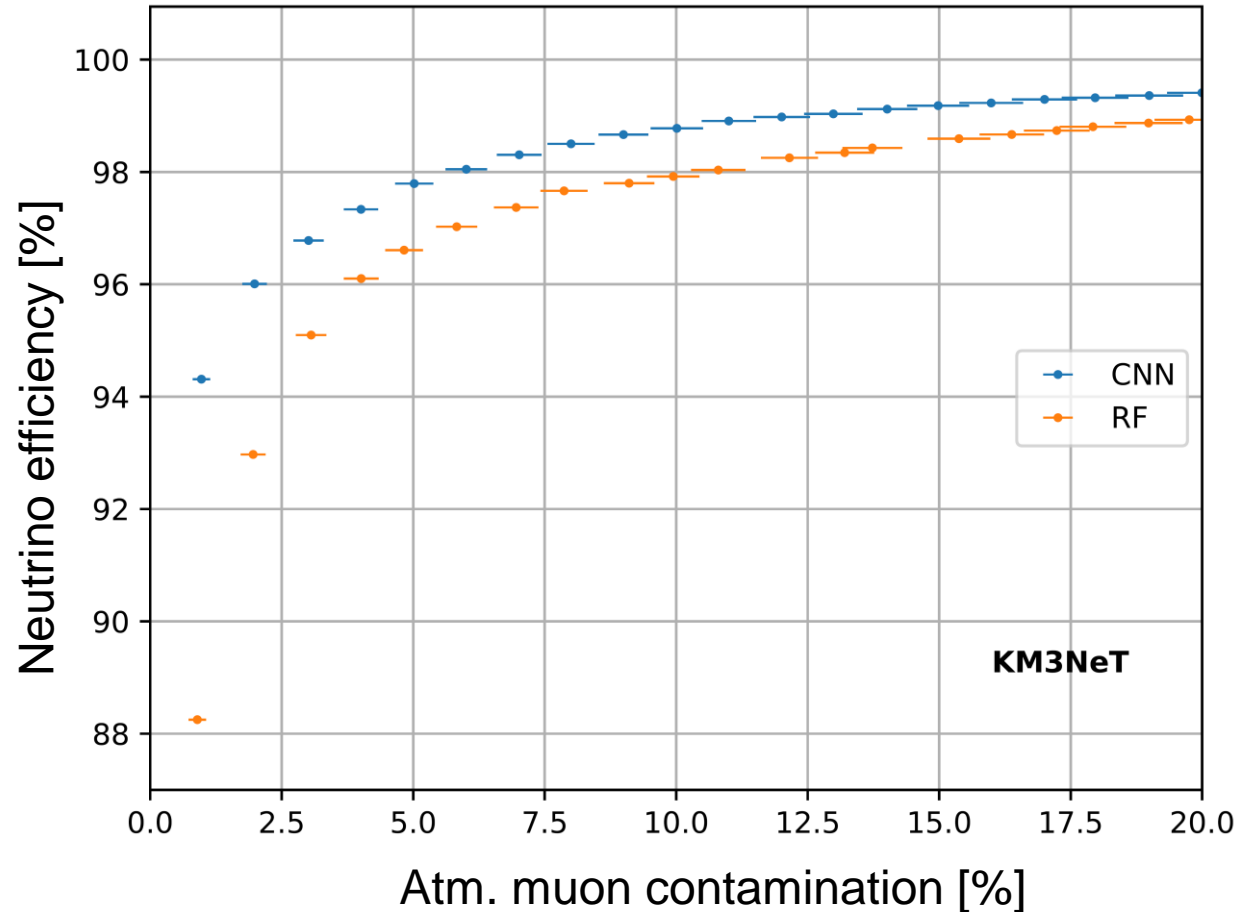
- CNN (this work)
- RF (standard method)

Muon contamination:

$$C_{\mu} = \frac{N_{\mu}(p)}{N_{total}(p)}$$

Neutrino efficiency:

$$v_{eff} = \frac{N_{\nu}(p)}{N_{\nu,total}}$$



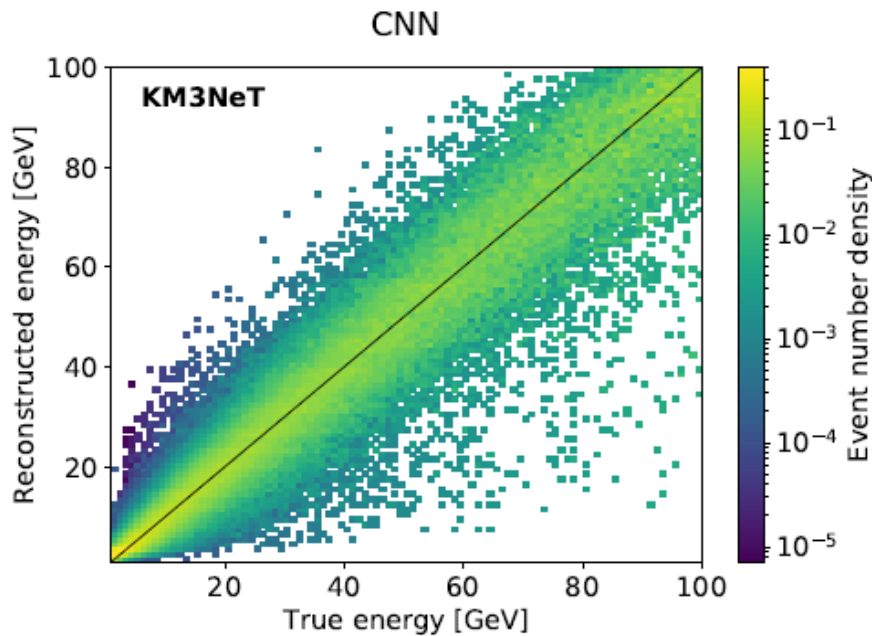


energy reconstruction

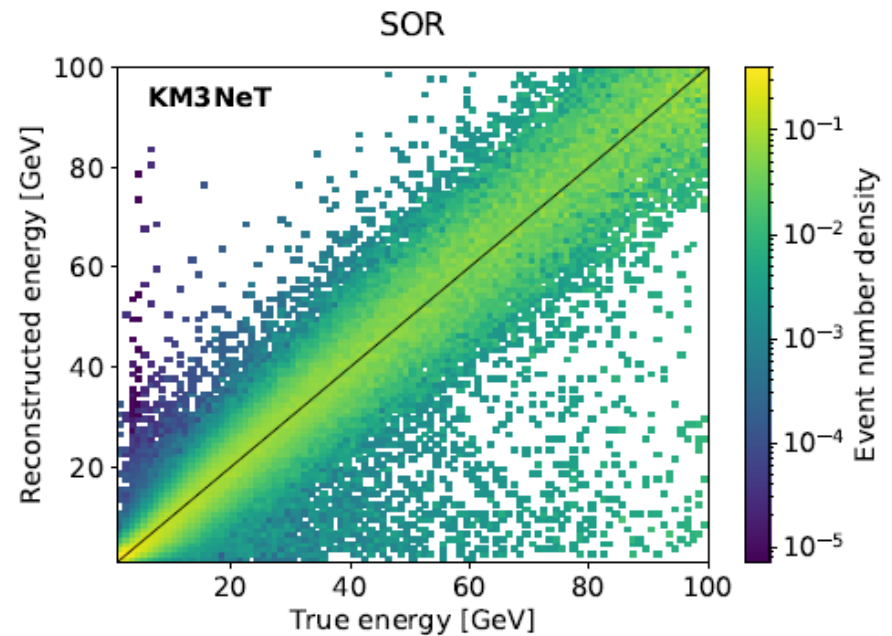
Regression results – shower energy



- OrcaNet shower energy reco is comparable to max. lkl. based reco



Deep Learning

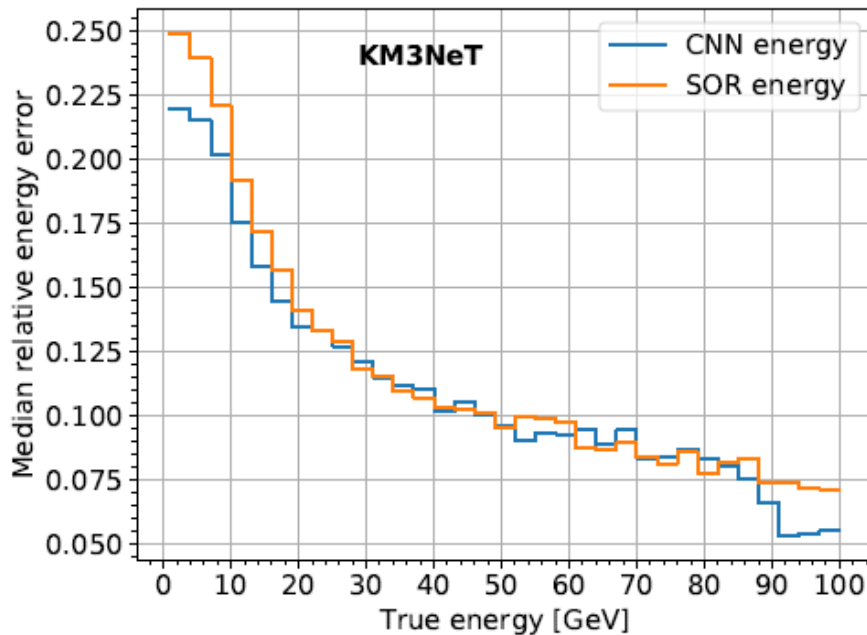


standard reco

Regression results – shower energy

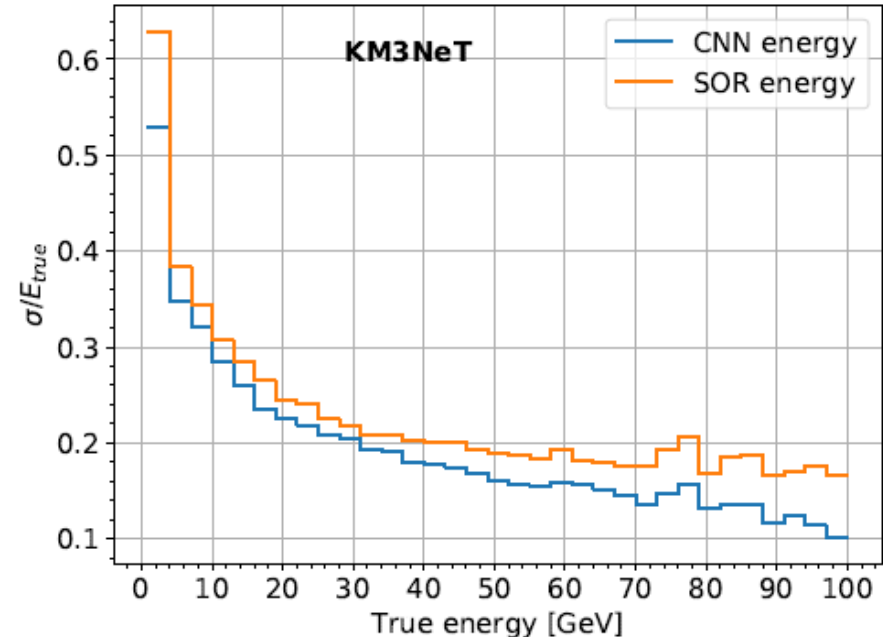


- CNN energy reco is comparable to standard reco



Median relative error

$$\text{Median}\left(\left|\frac{E_{\text{reco}} - E_{\text{true}}}{E_{\text{true}}}\right|\right)$$



Relative standard deviation

$$\frac{\sigma}{E_{\text{true}}}$$