

MiniBooNE Oscillation Results with Complete Dataset

Adrien Hourlier

on behalf of the MiniBooNE Collaboration

2020/11/24



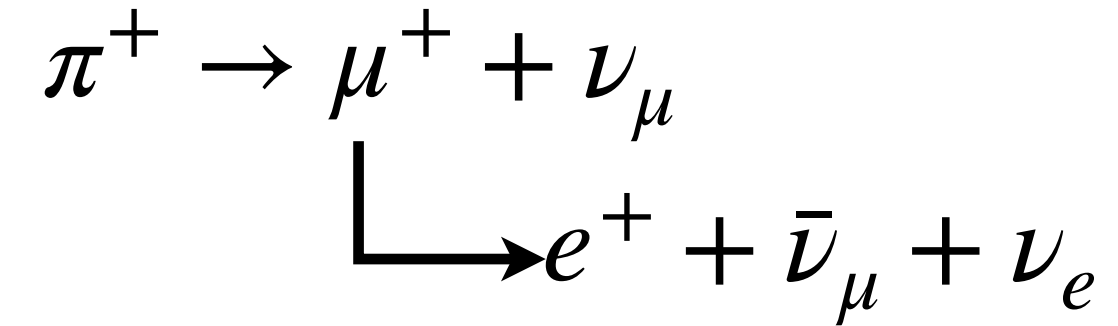
Massachusetts
Institute of
Technology

preprint available at [arxiv:2006.16883](https://arxiv.org/abs/2006.16883)

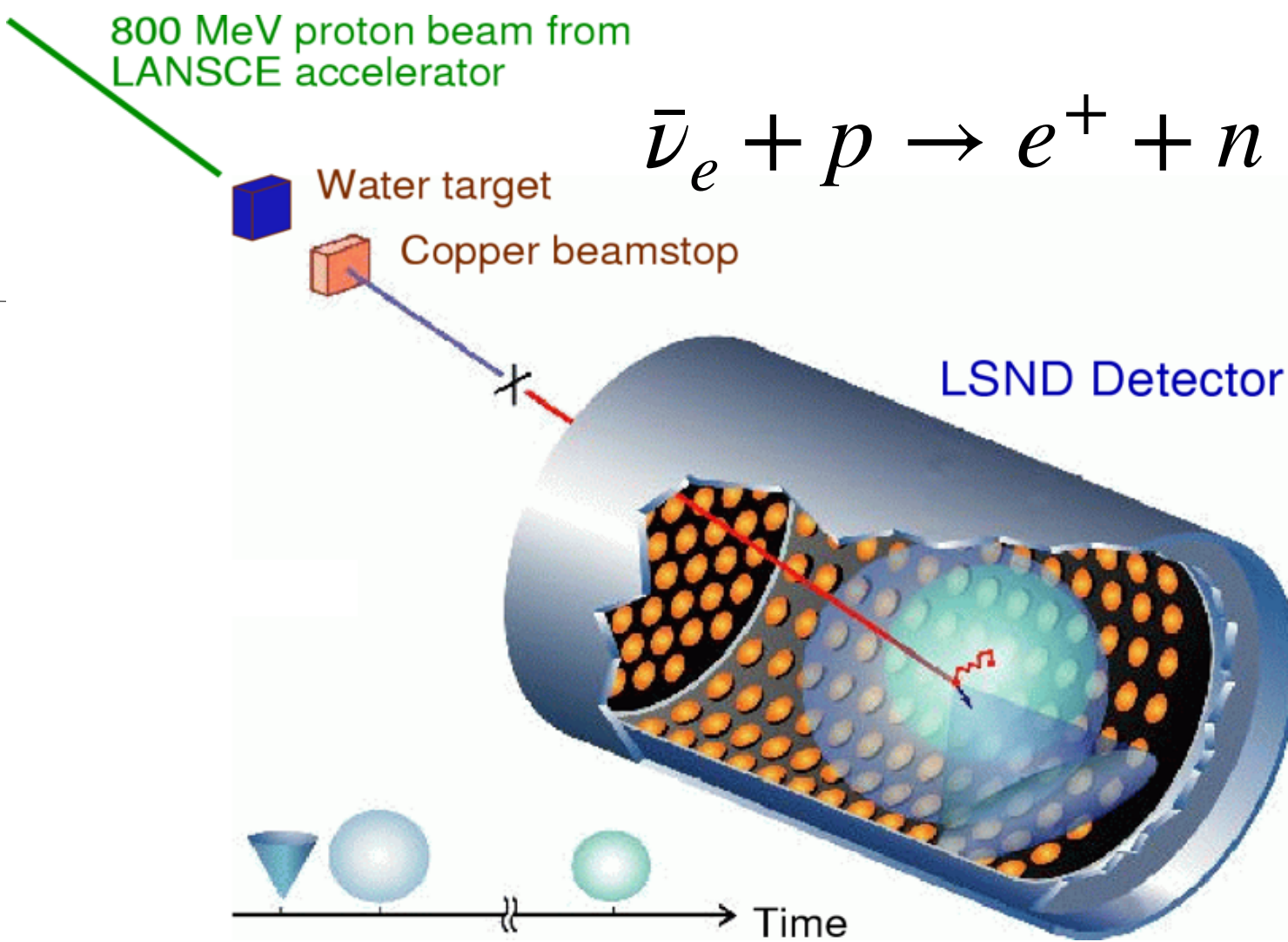
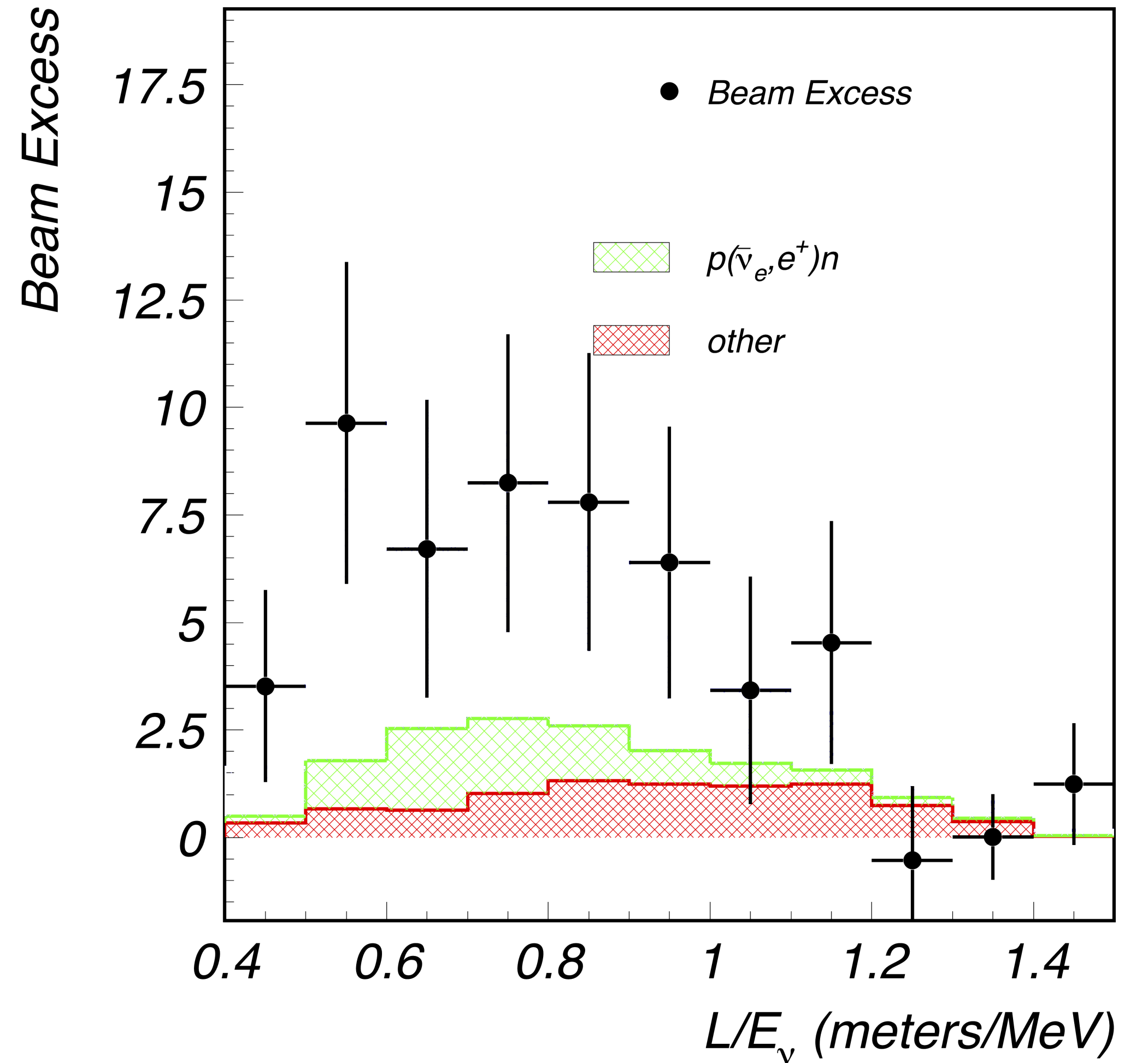
Before MiniBooNE : The LSND Anomaly

Phys.Rev.D 64.112007

- Stopped π beam at LANL

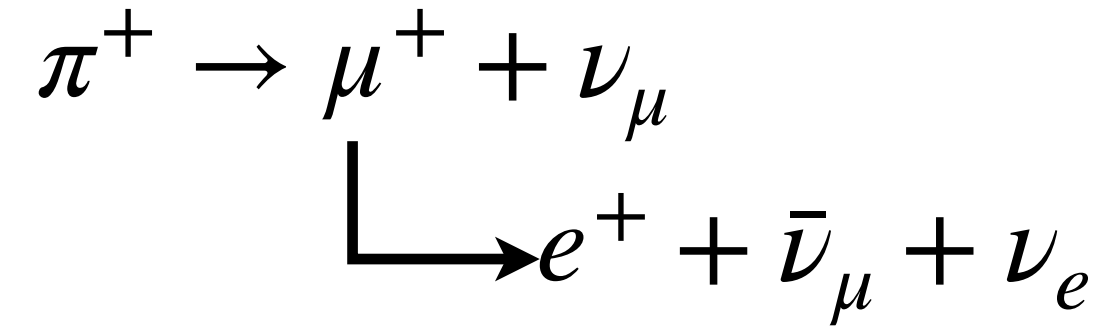


- appearance of $\bar{\nu}_e$ in a $\bar{\nu}_\mu$ beam
- $\bar{\nu}_e$ signature : Cherenkov light from e^+ with delayed n-capture
- Excess = $87.9 \pm 22.4 \pm 6$ (3.8σ)



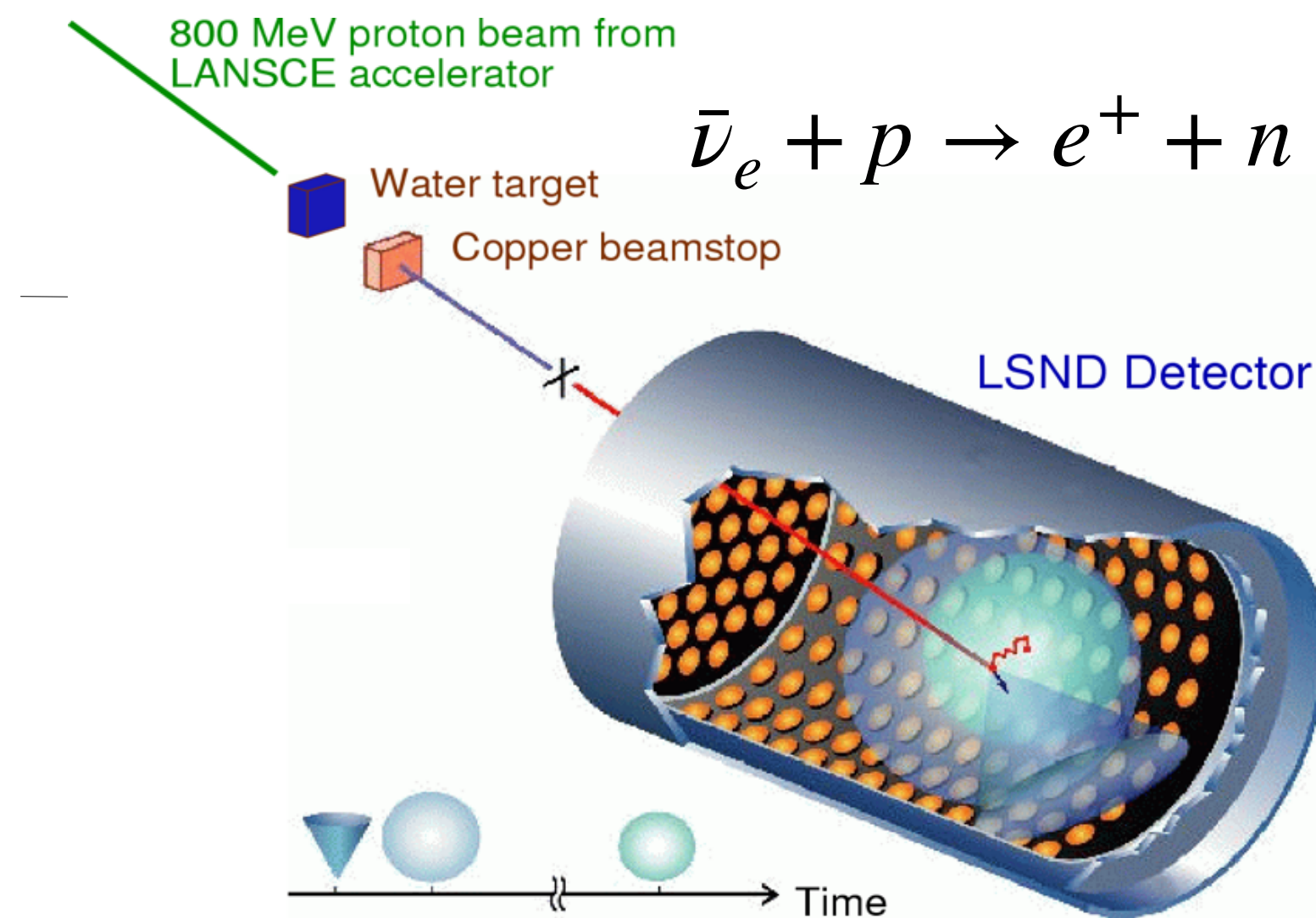
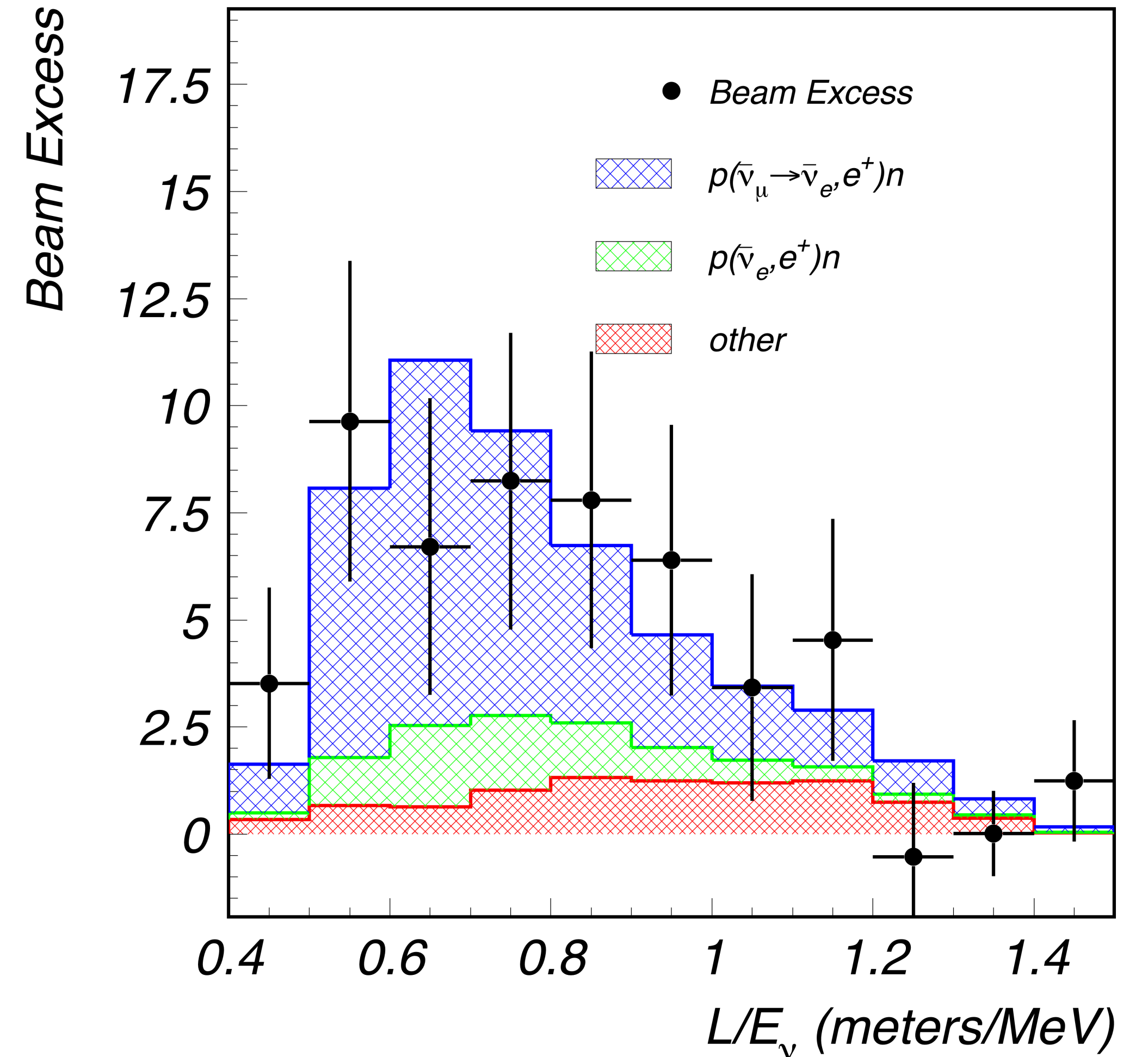
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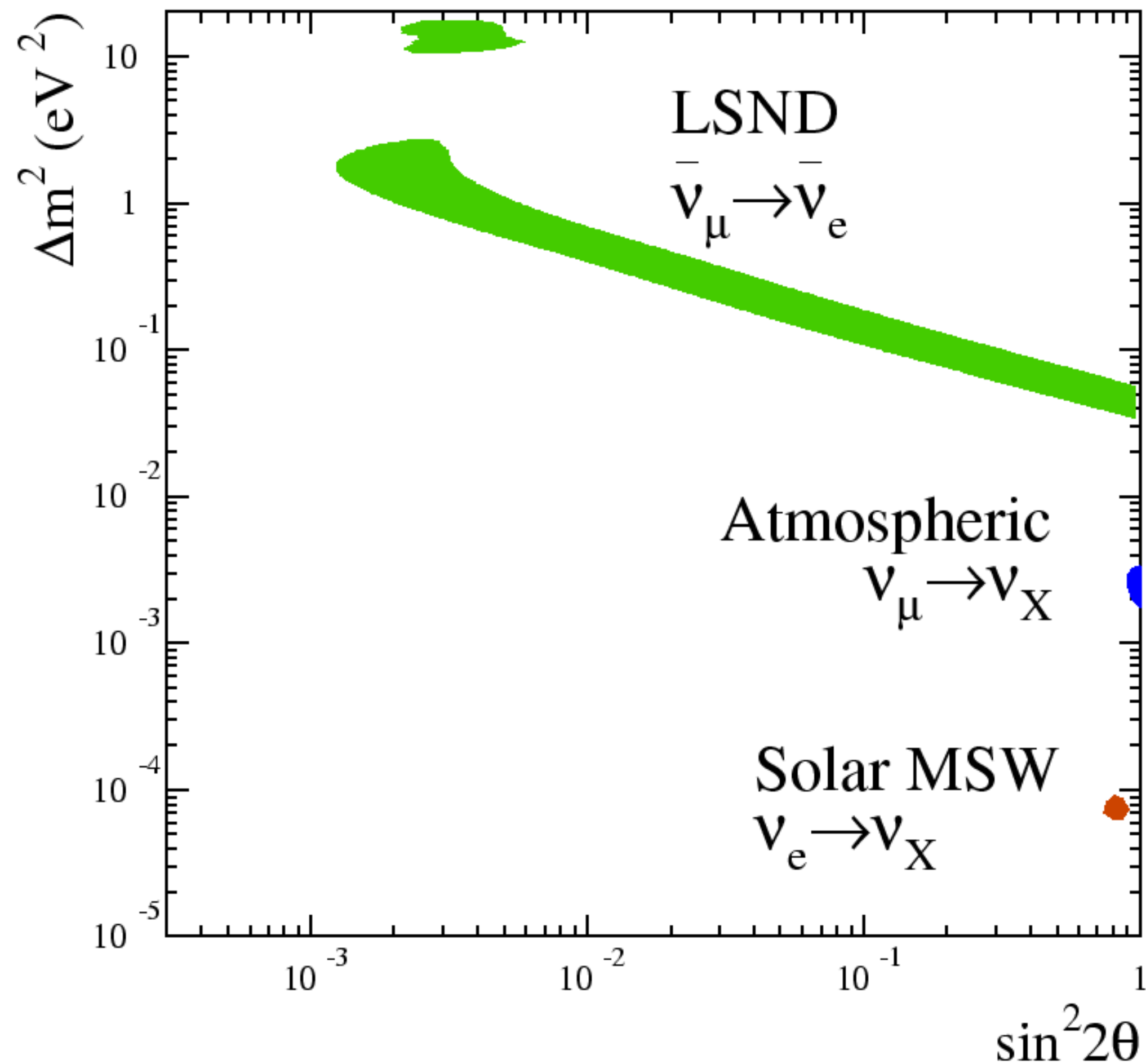


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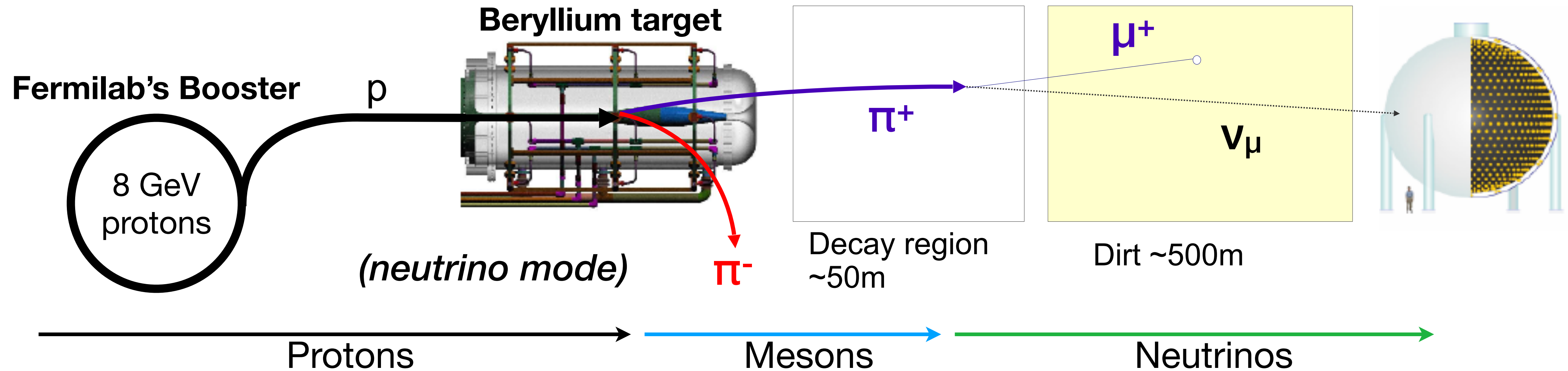


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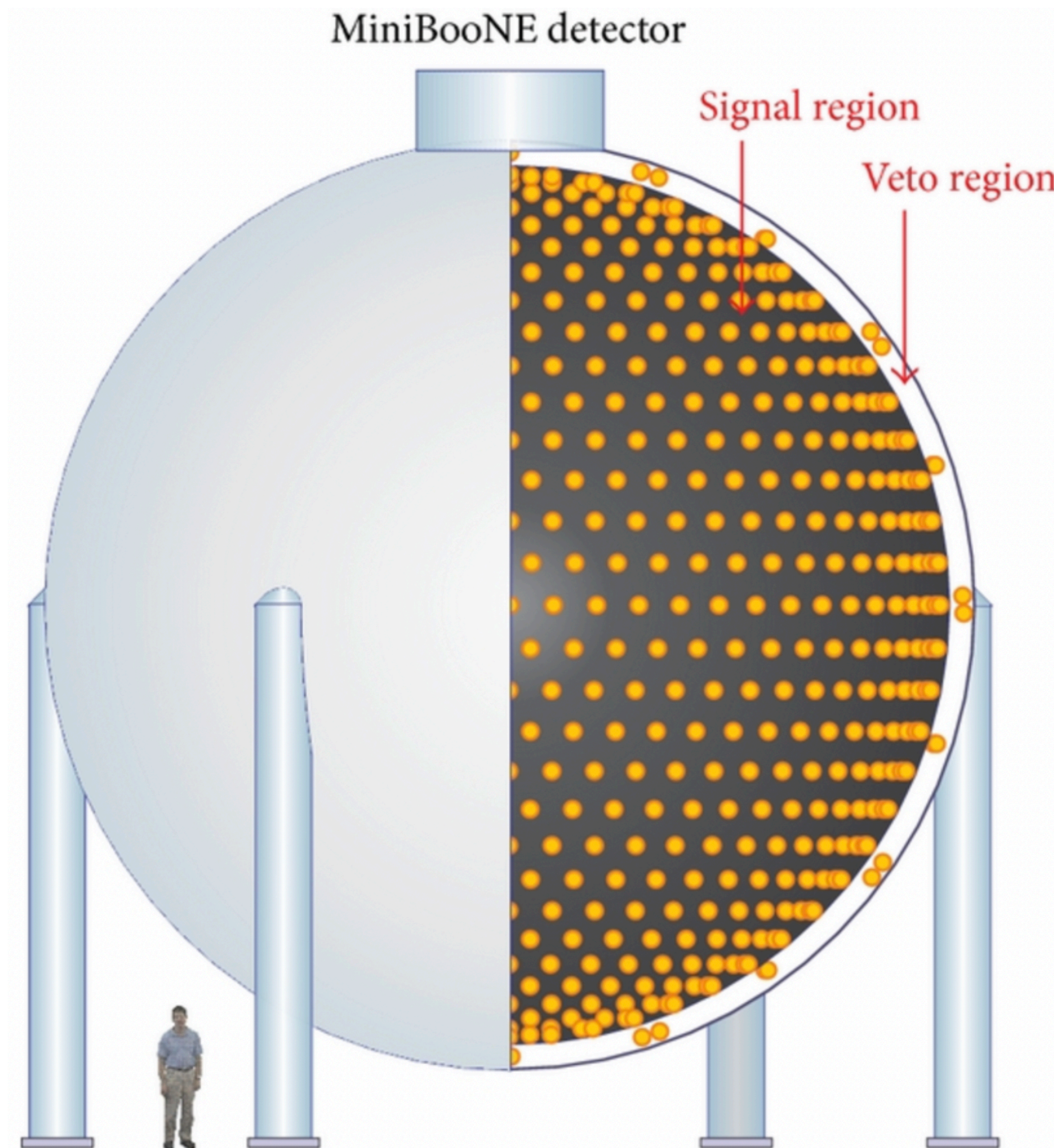
- LSND's oscillation fit yields parameter space that is not compatible with either atmospheric ($\Delta m^2 \sim 10^{-3}$ eV^2) or solar ($\Delta m^2 \sim 1-5$ eV^2) oscillations
- There cannot be three independent Δm^2 in a 3ν scheme
- LSND's result indicates a possible 4th generation of neutrino, but there are only 3 "active" flavors, that couple to the Z-boson.
- The additional neutrino must be "sterile" (not coupling to the Z and W bosons)

The MiniBooNE Experiment

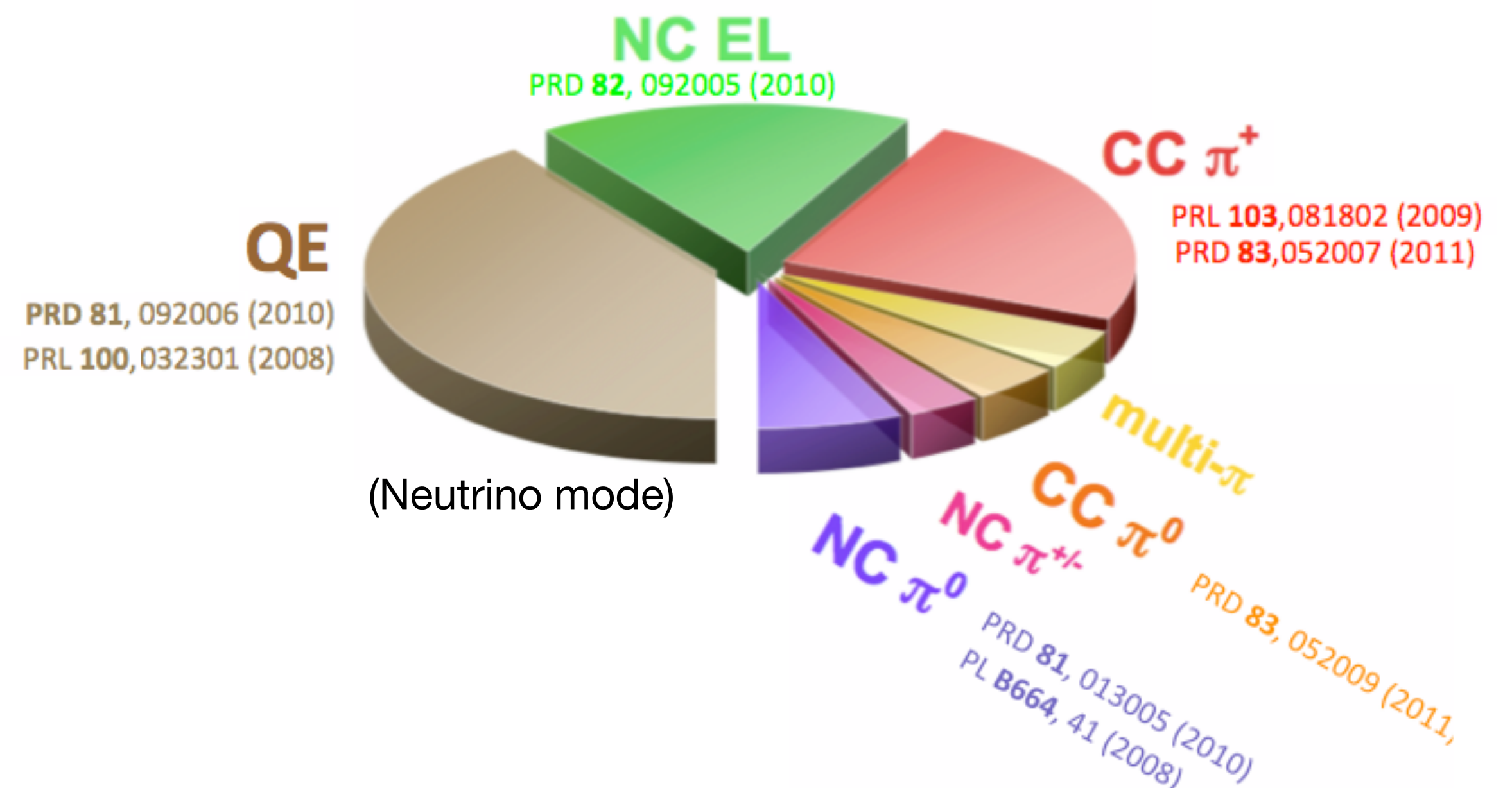


- Proposed to investigate the LSND anomaly, in search for sterile neutrinos
- Located on the Booster Neutrino Beam at Fermilab
- Single horn focused neutrino beam : Selection of neutrino/antineutrino modes
- Similar L/E as LSND :
 - MiniBooNE ~500 m / ~500 MeV
 - LSND ~30 m / ~ 30 MeV
- Different systematics due to different fluxes, event signatures and backgrounds

The MiniBooNE Detector

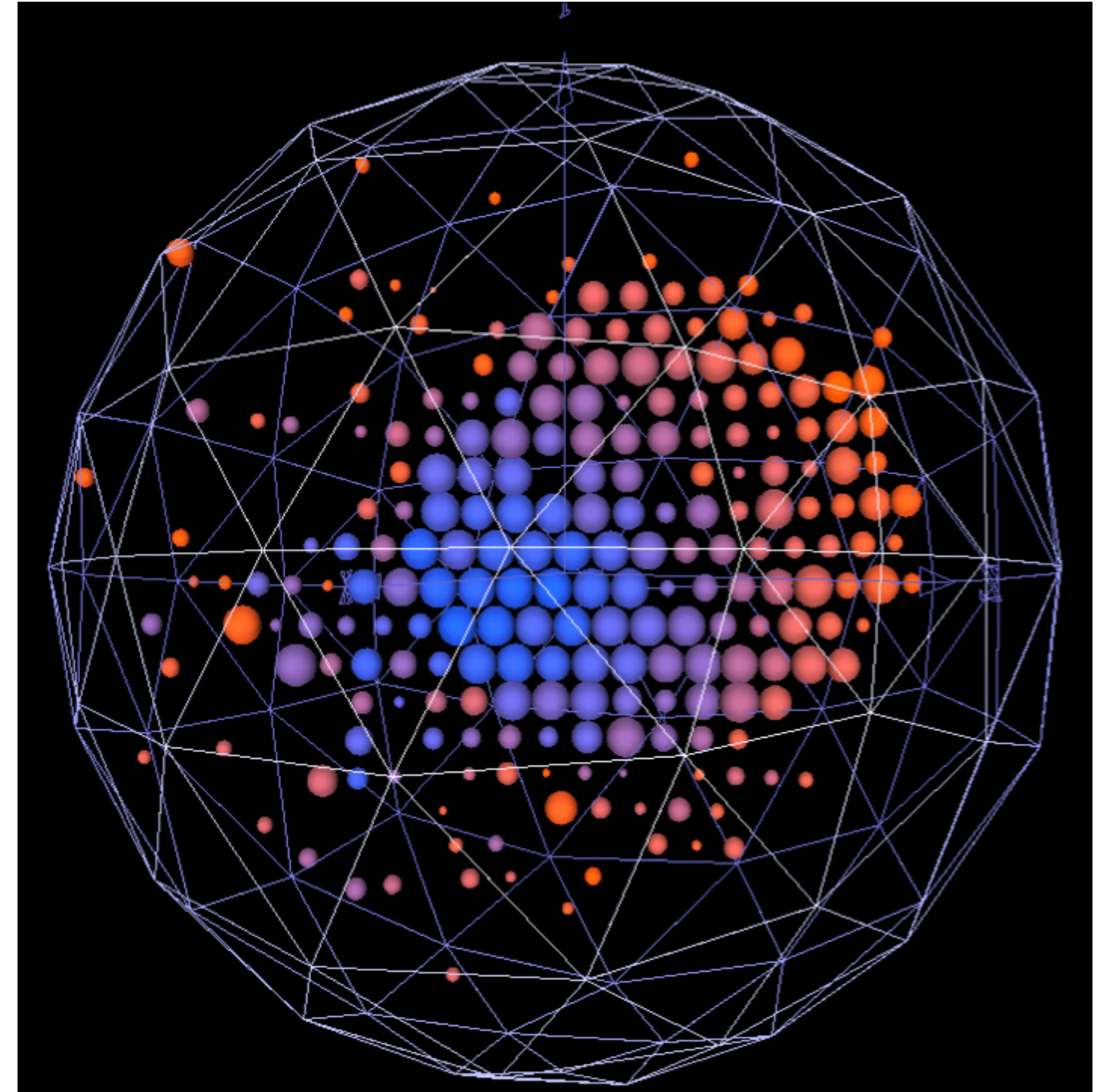
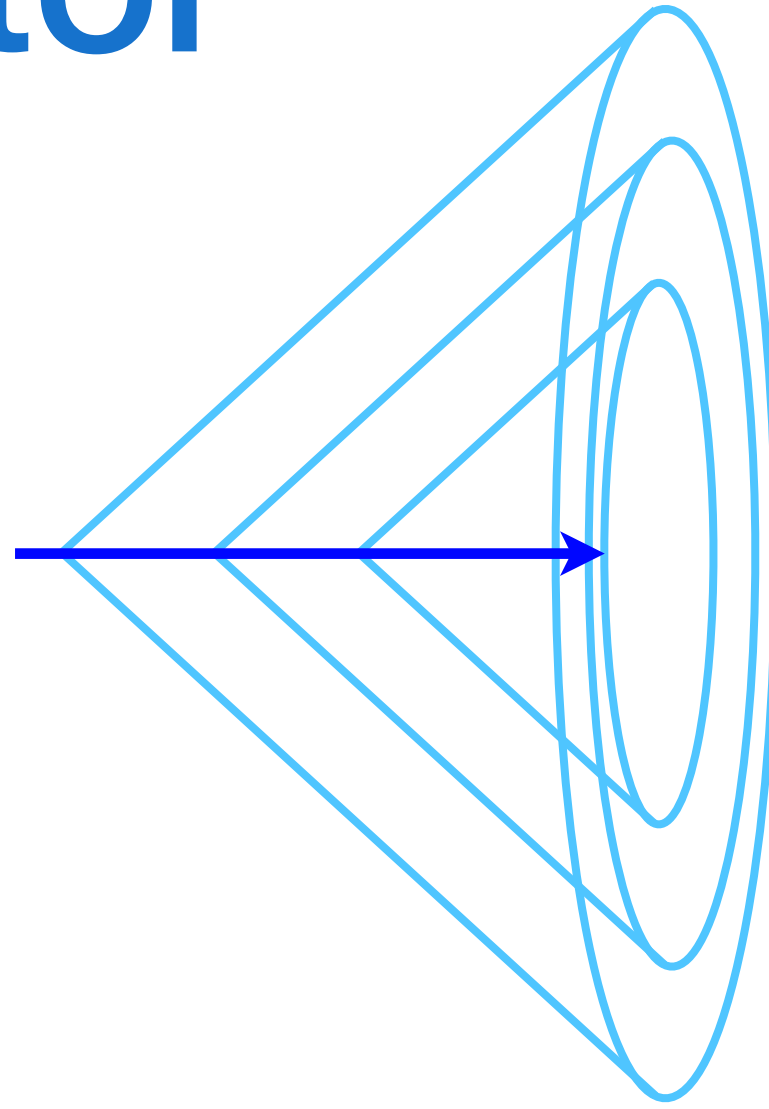


- $\varnothing 12.2$ m sphere, $\varnothing 10$ m fiducial volume
- 800 tons of mineral oil, 450 tons fiducial mass
- 2 optically isolated volumes
- 1280 inner PMTs, 240 veto PMTs
- Very well understood detector
 - 3% change of the energy scale over 17 years of running
 - Measurements of cross sections for $>90\%$ of the neutrino and anti-neutrino processes



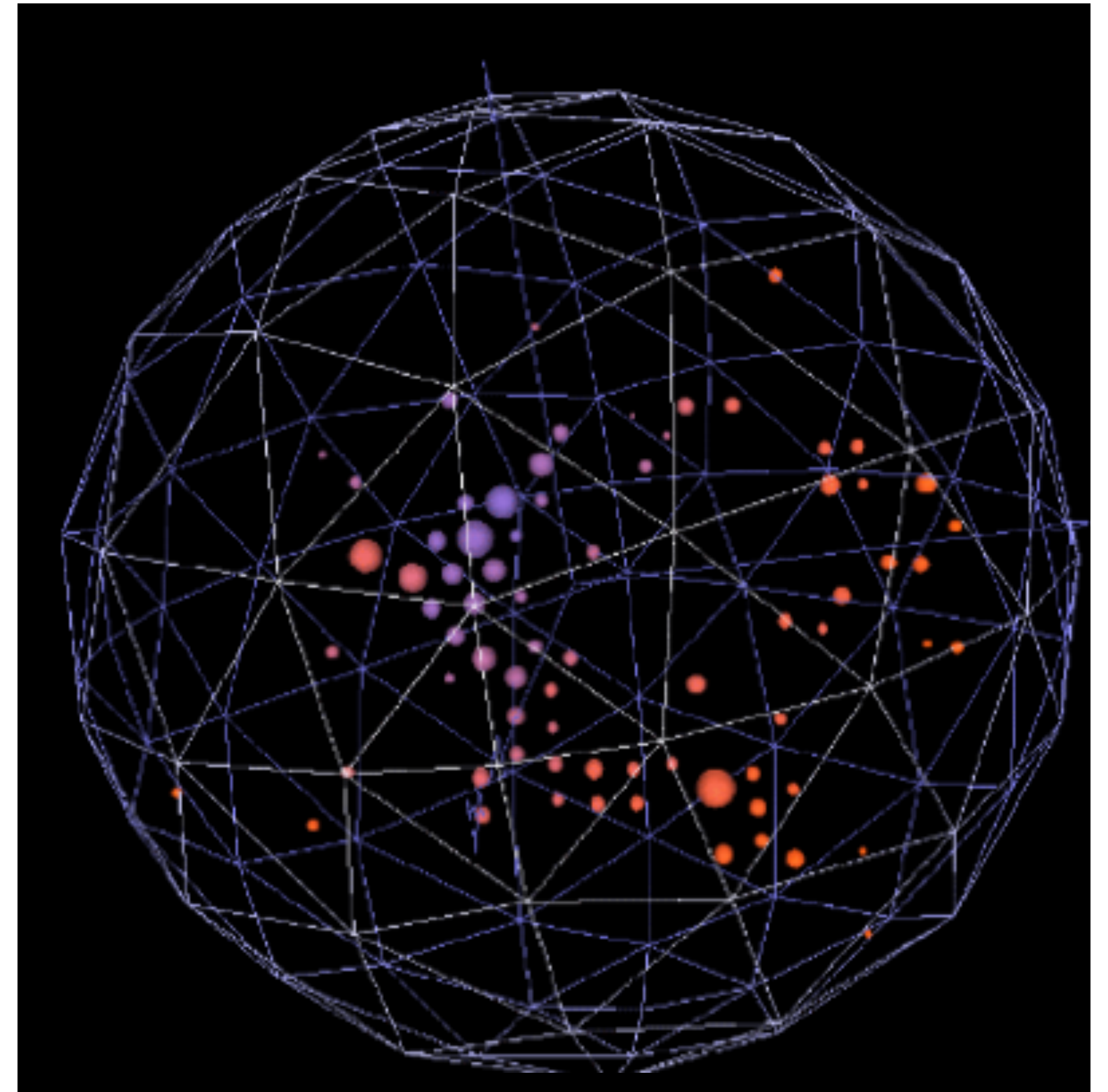
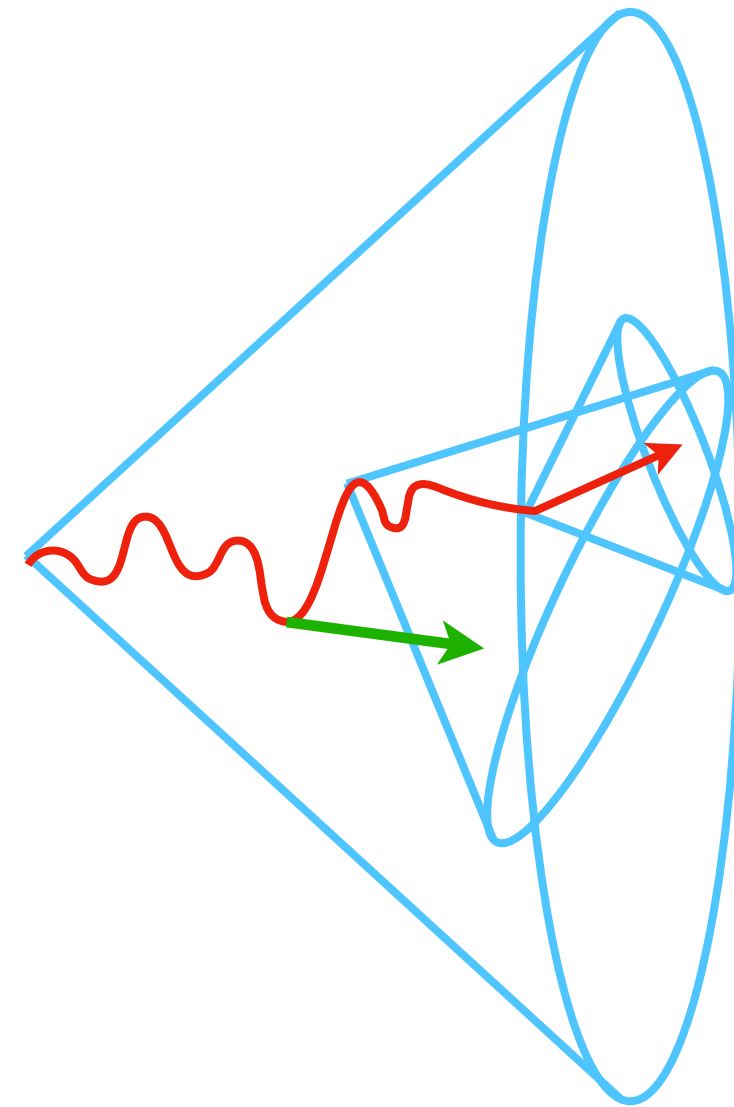
Events in the Detector

- Muons
 - Long straight tracks
 - Sharp clean rings or disks
- Electrons
 - Multiple scattering
 - Radiative processes
 - Scattered fussy rings
- Neutral pions
 - Decay to 2 photons
 - Double fuzzy rings
- NC elastic scattering
 - No Cherenkov radiation
 - Isotropic scintillation hits



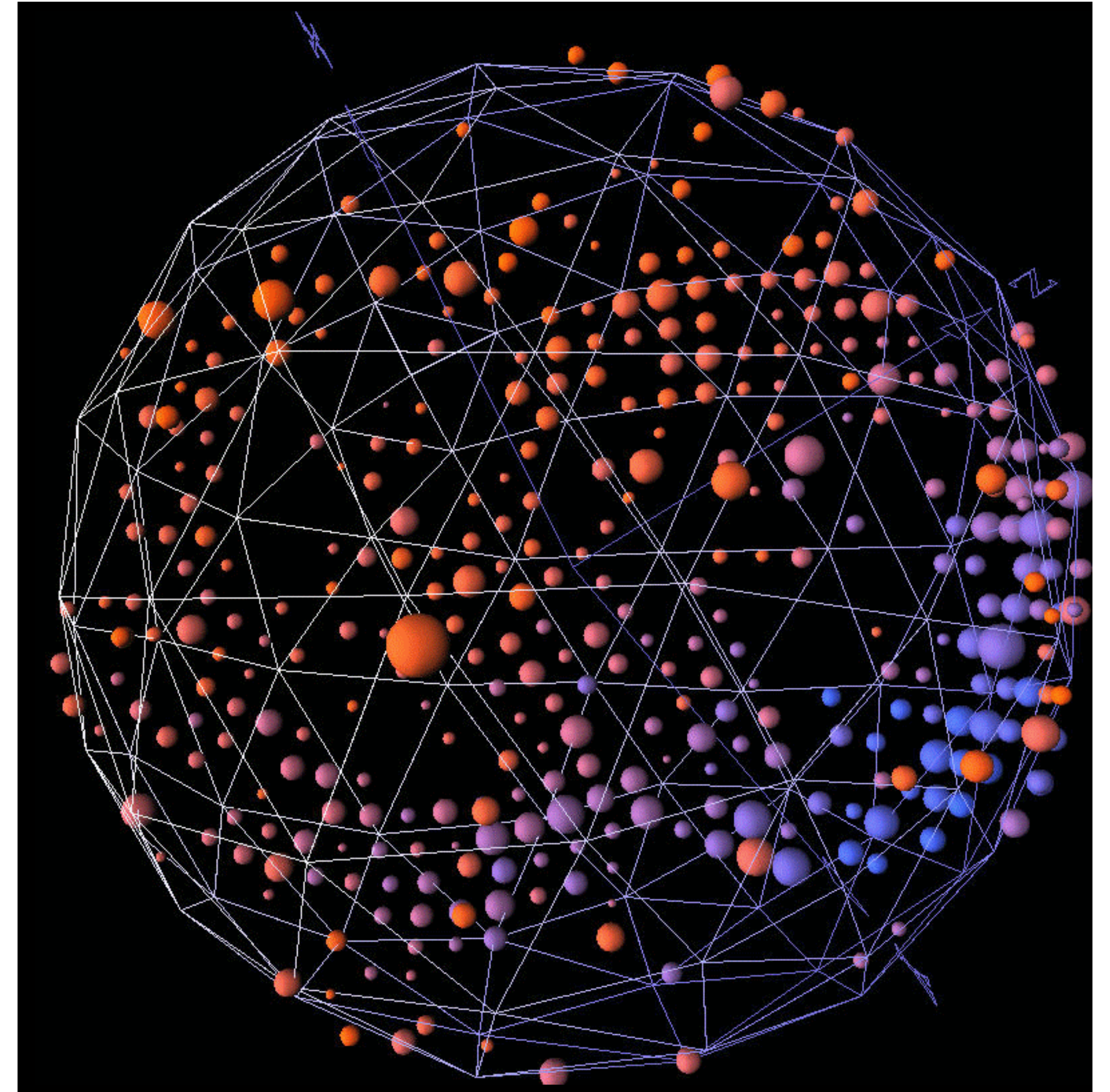
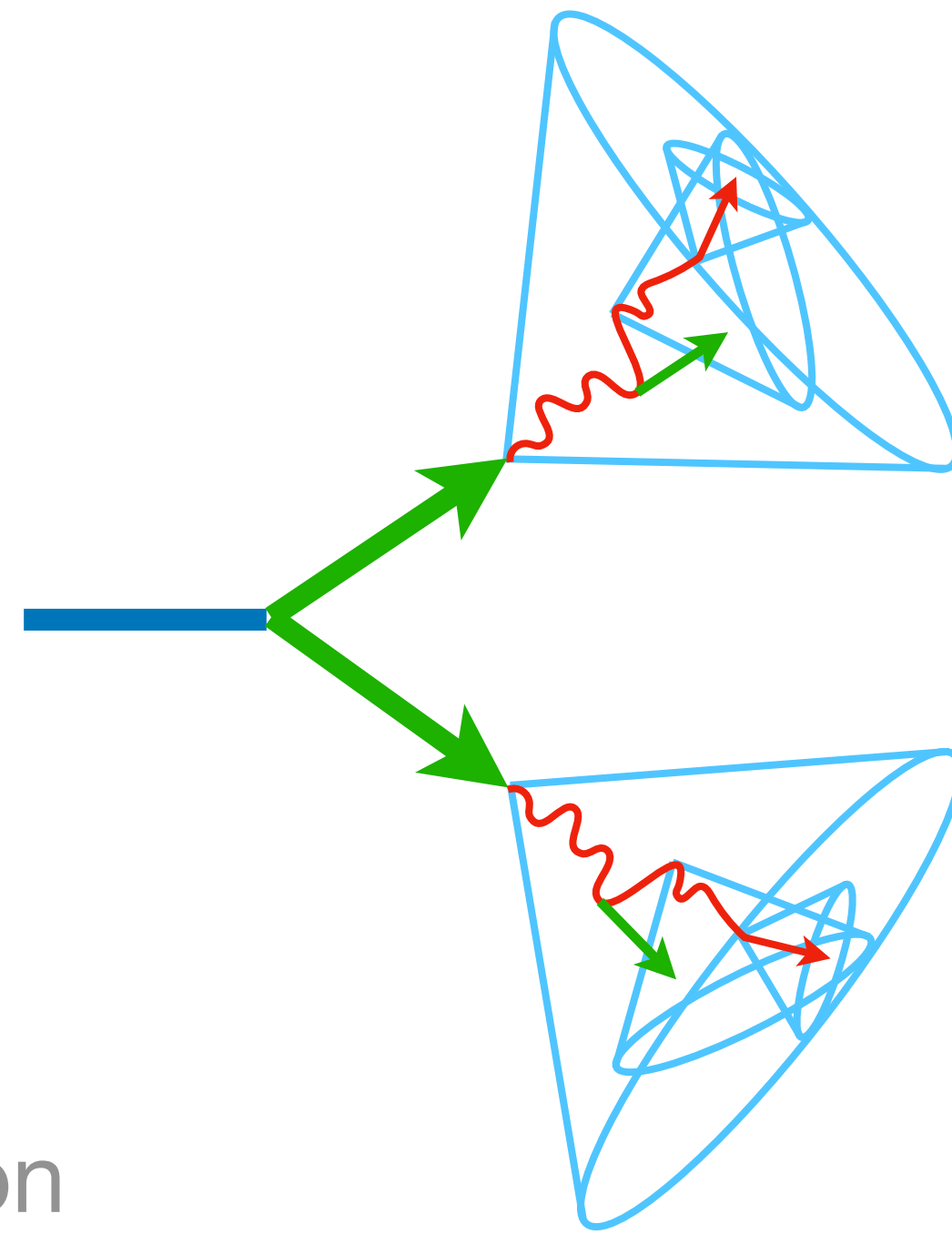
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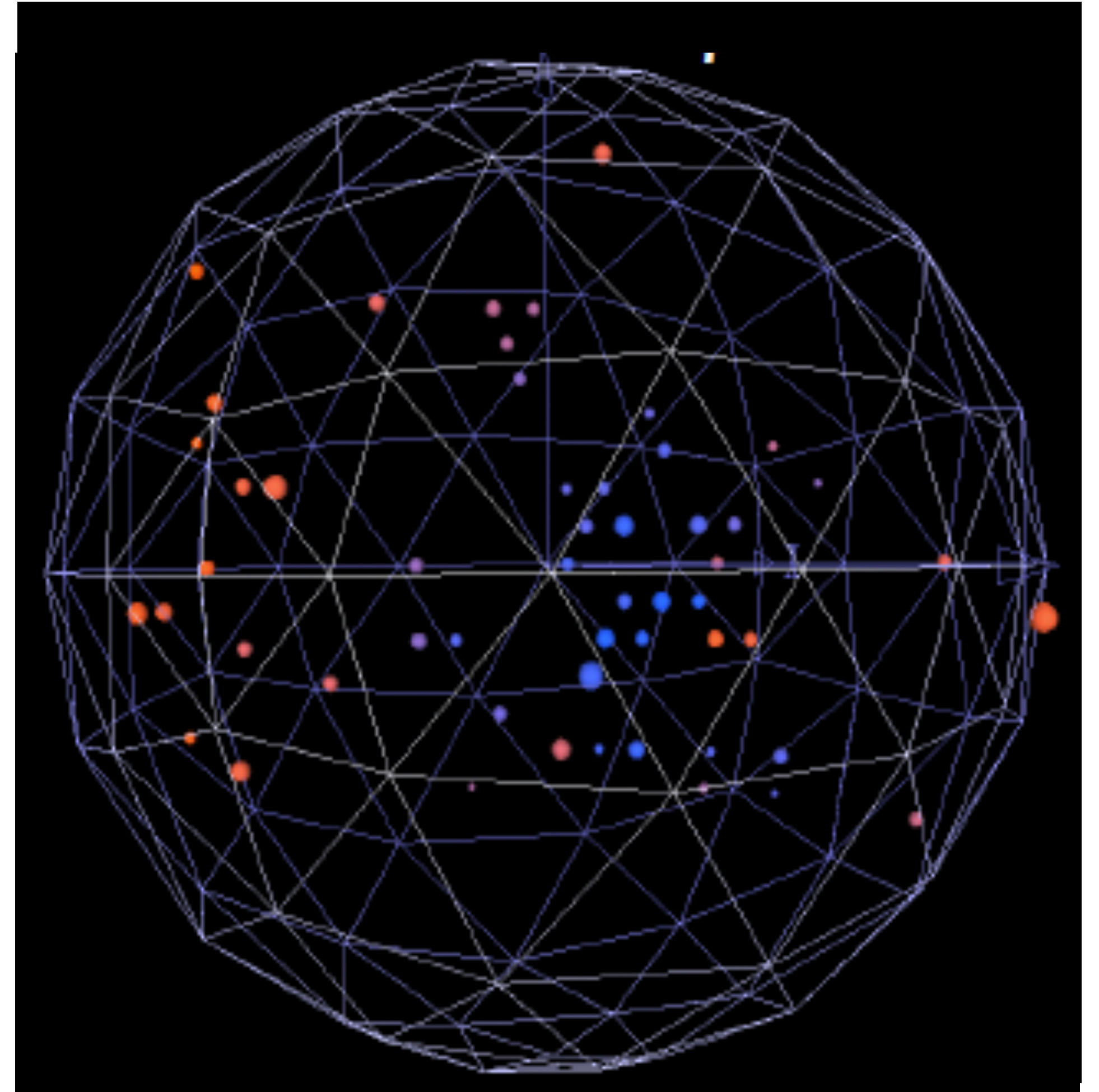
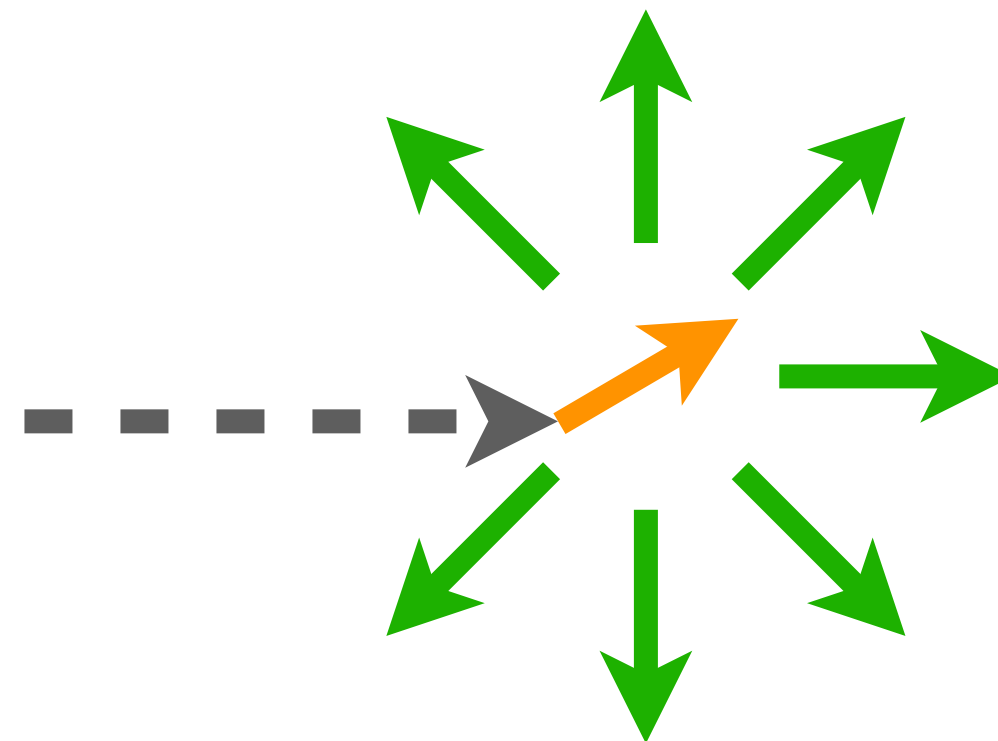
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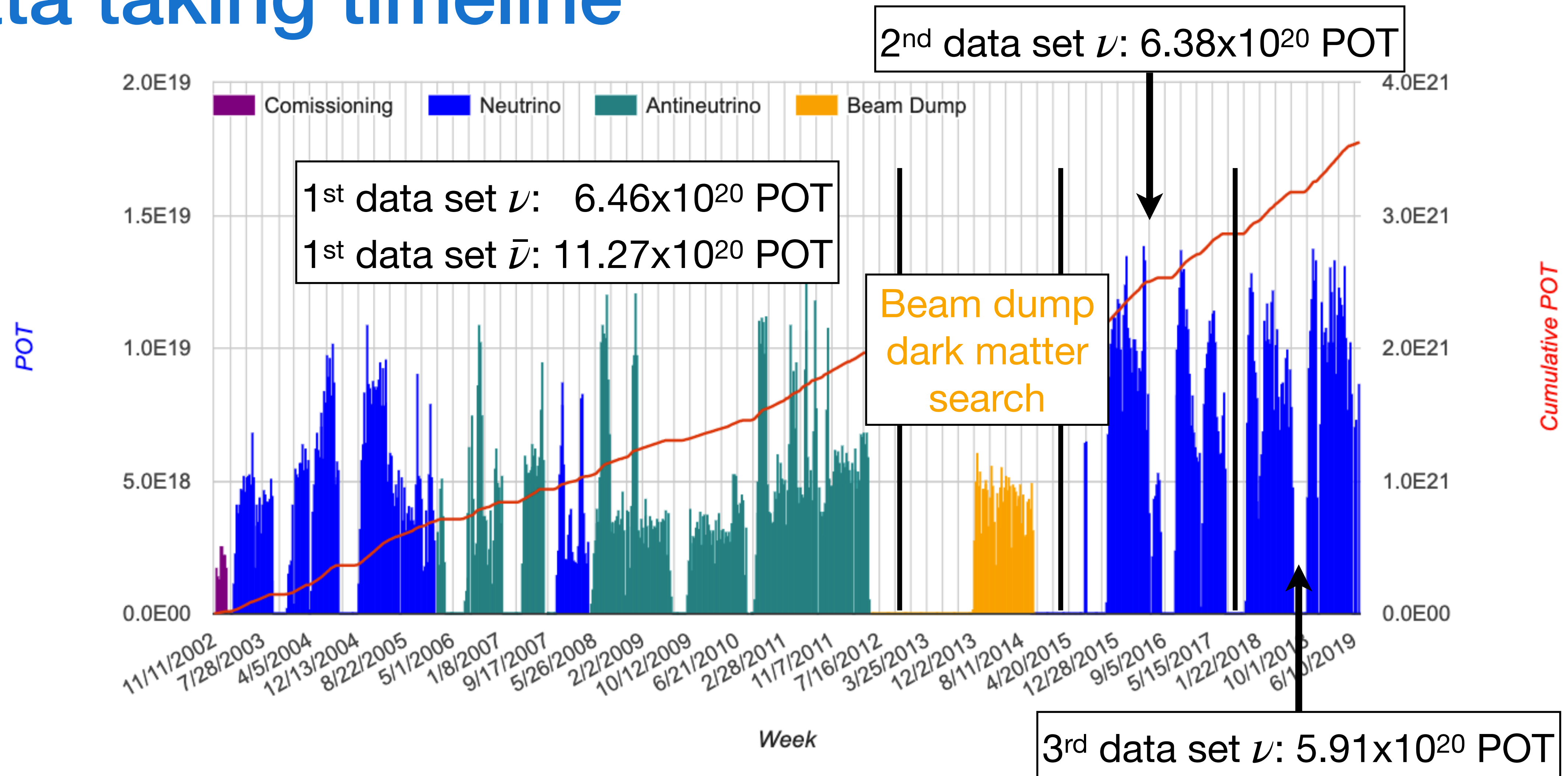


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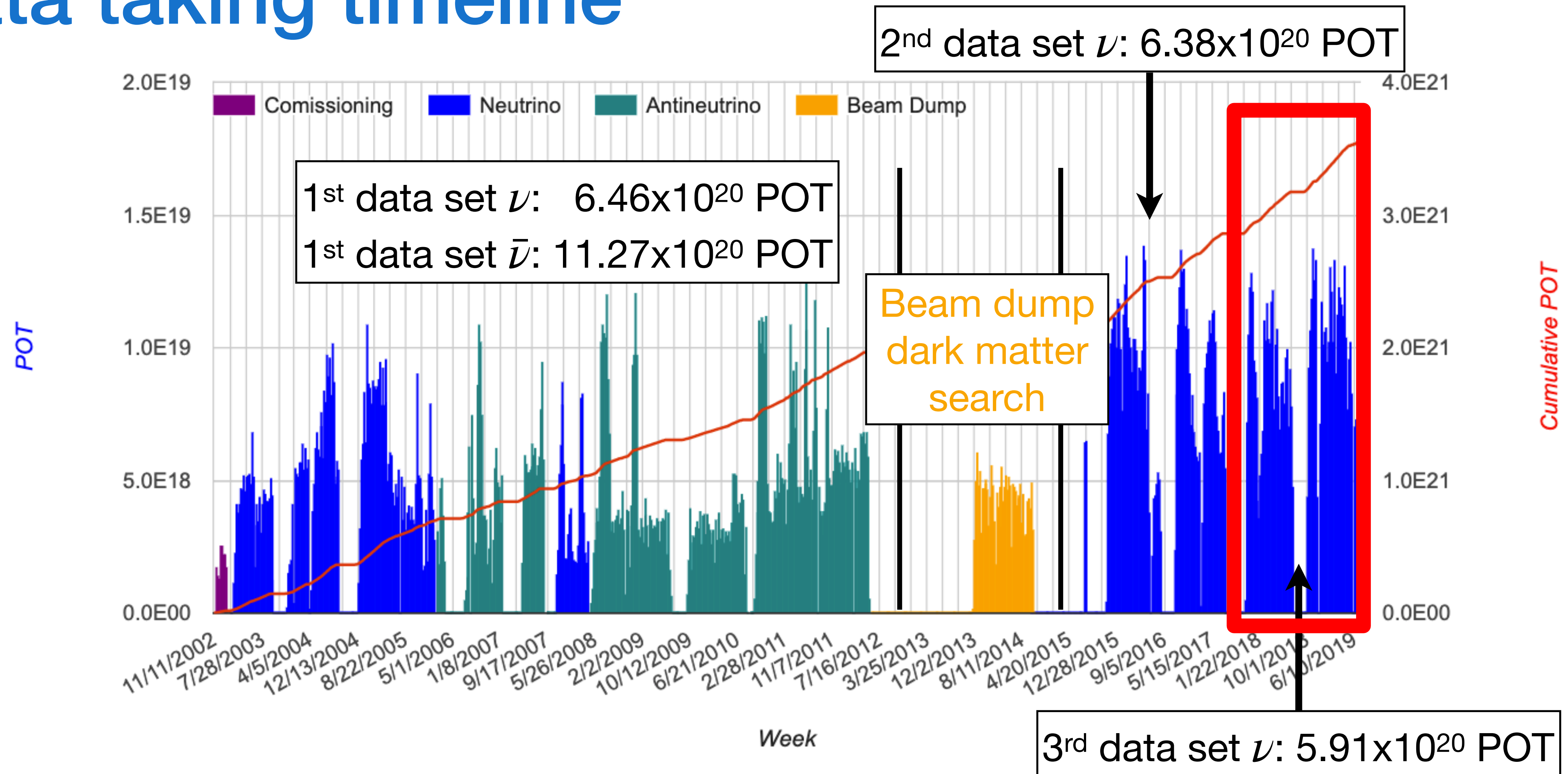


Data taking timeline



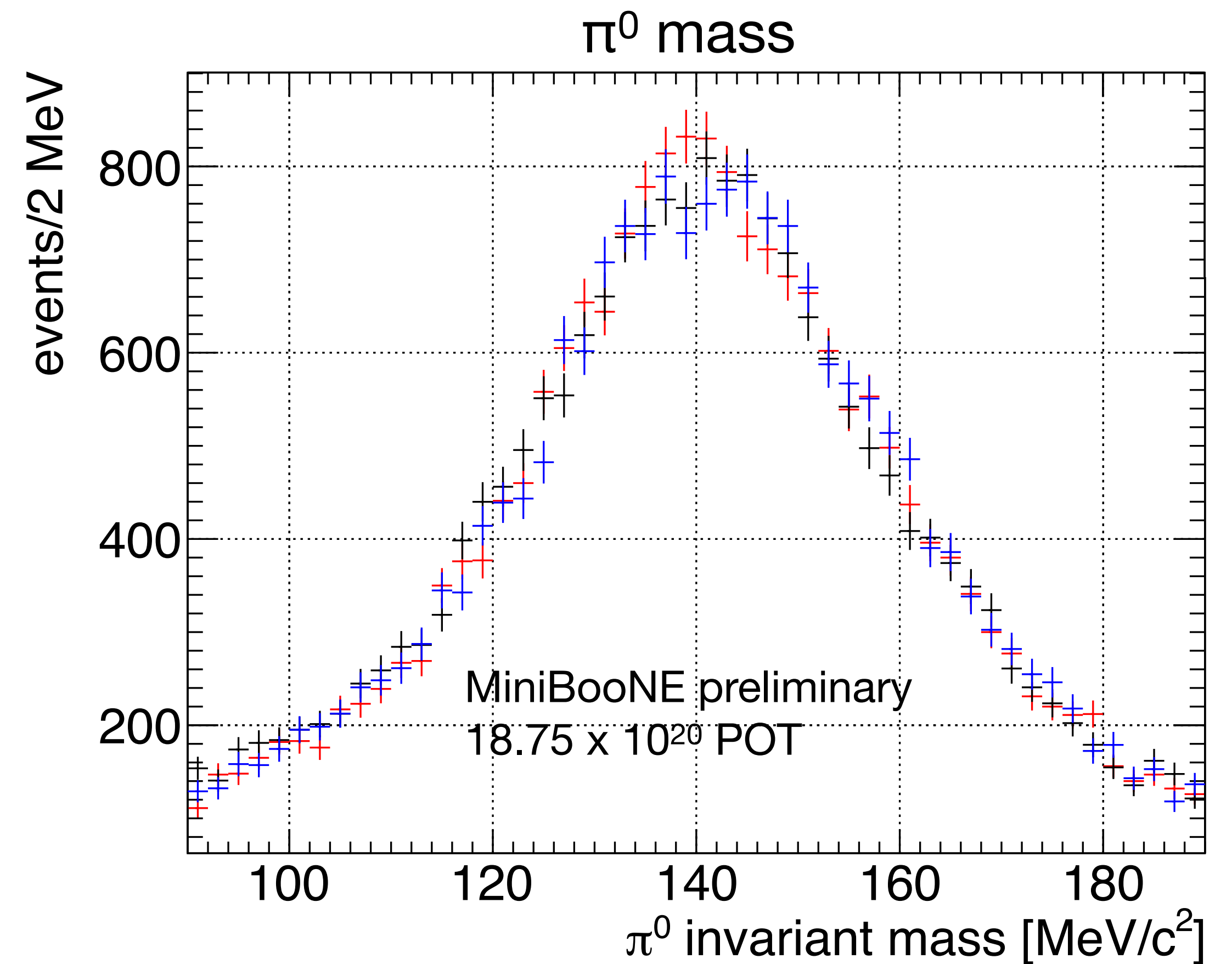
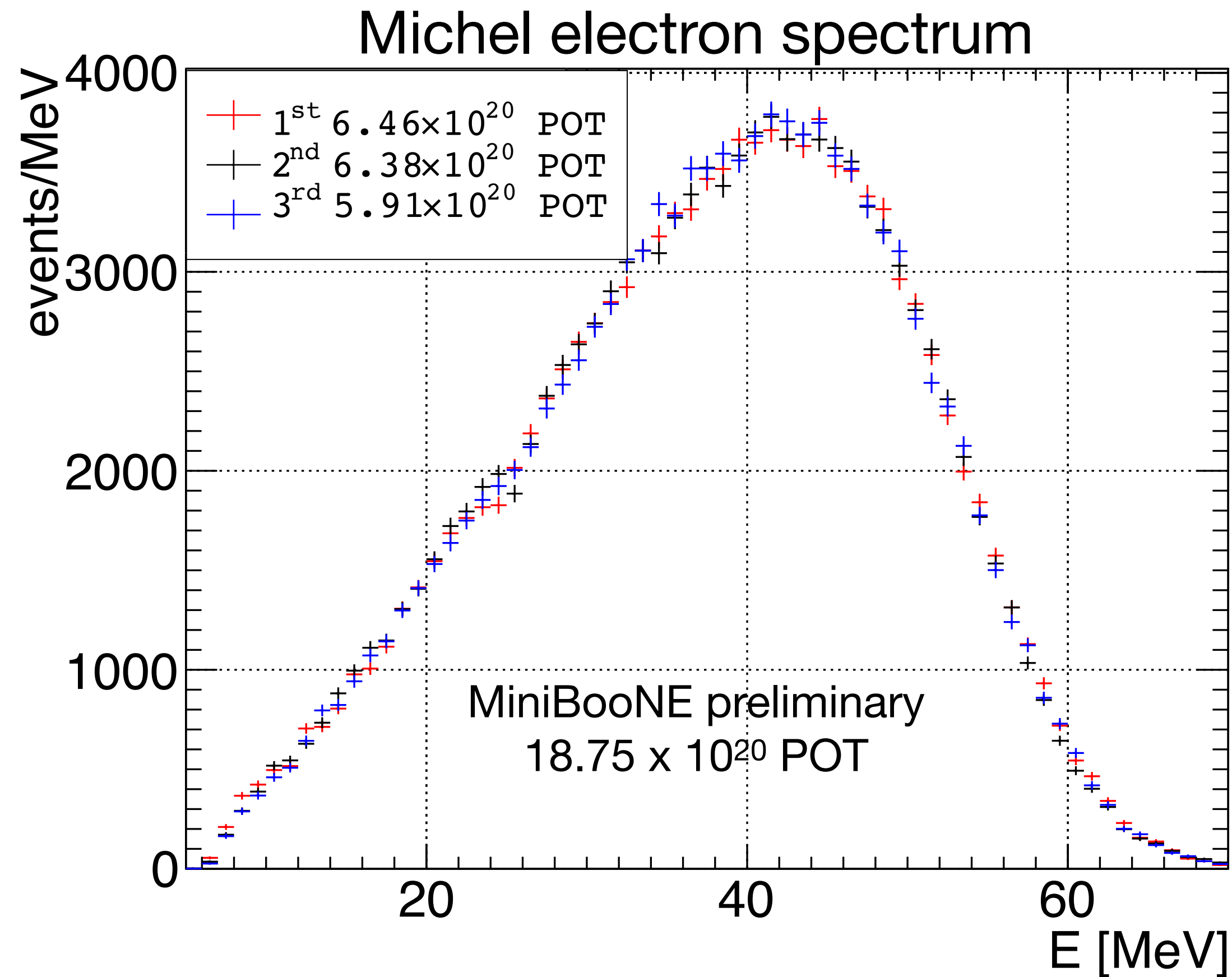
- We have added another $\sim 6 \times 10^{20}$ POT to the neutrino dataset since the previous data release.
- The detector was turned off at the end of summer 2019, mothballed and waiting for future use...
- Almost 17 years of running, or as much as 5 army ants worth of protons!

Data taking timeline



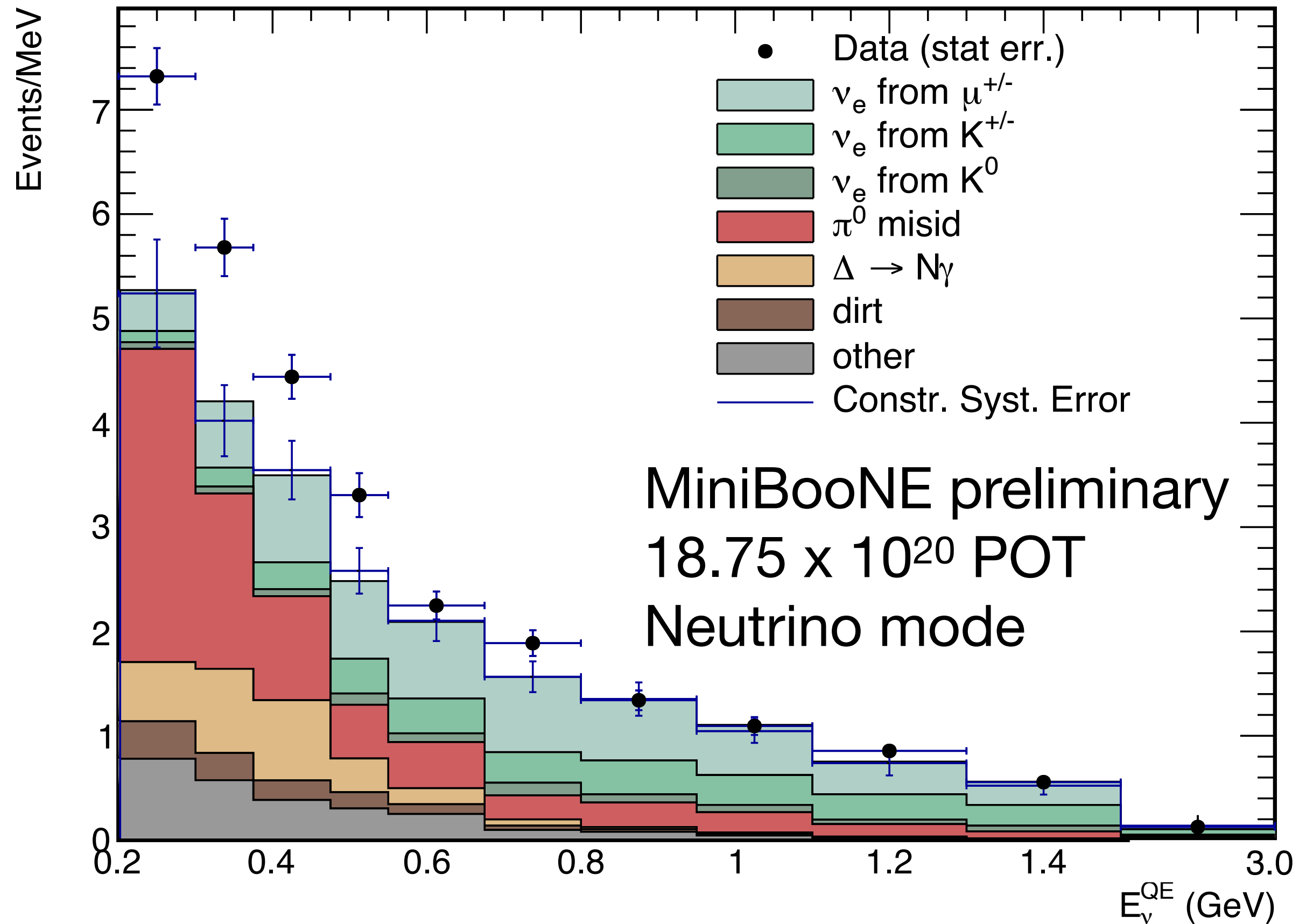
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Excellent long term detector stability over 3 run periods



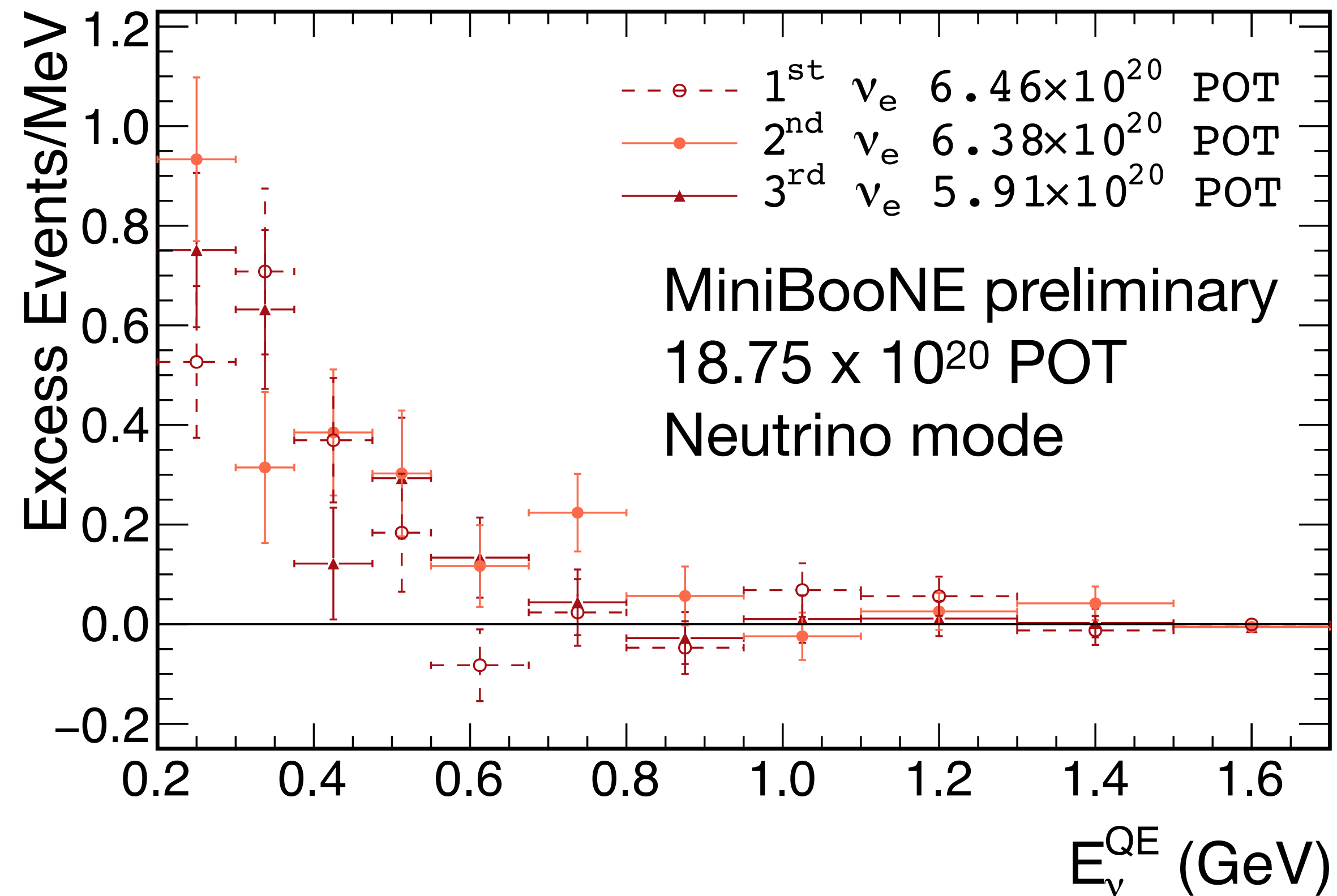
- We can use two standard candles to calibrate the energy scale over the different data sets
- Scale up the energy response to match the original 2007 data release:
 - 2015-2017 neutrino data => 2% energy scaling
 - 2017-2019 neutrino data => 3% energy scaling

Neutrino energy and 3ν prediction



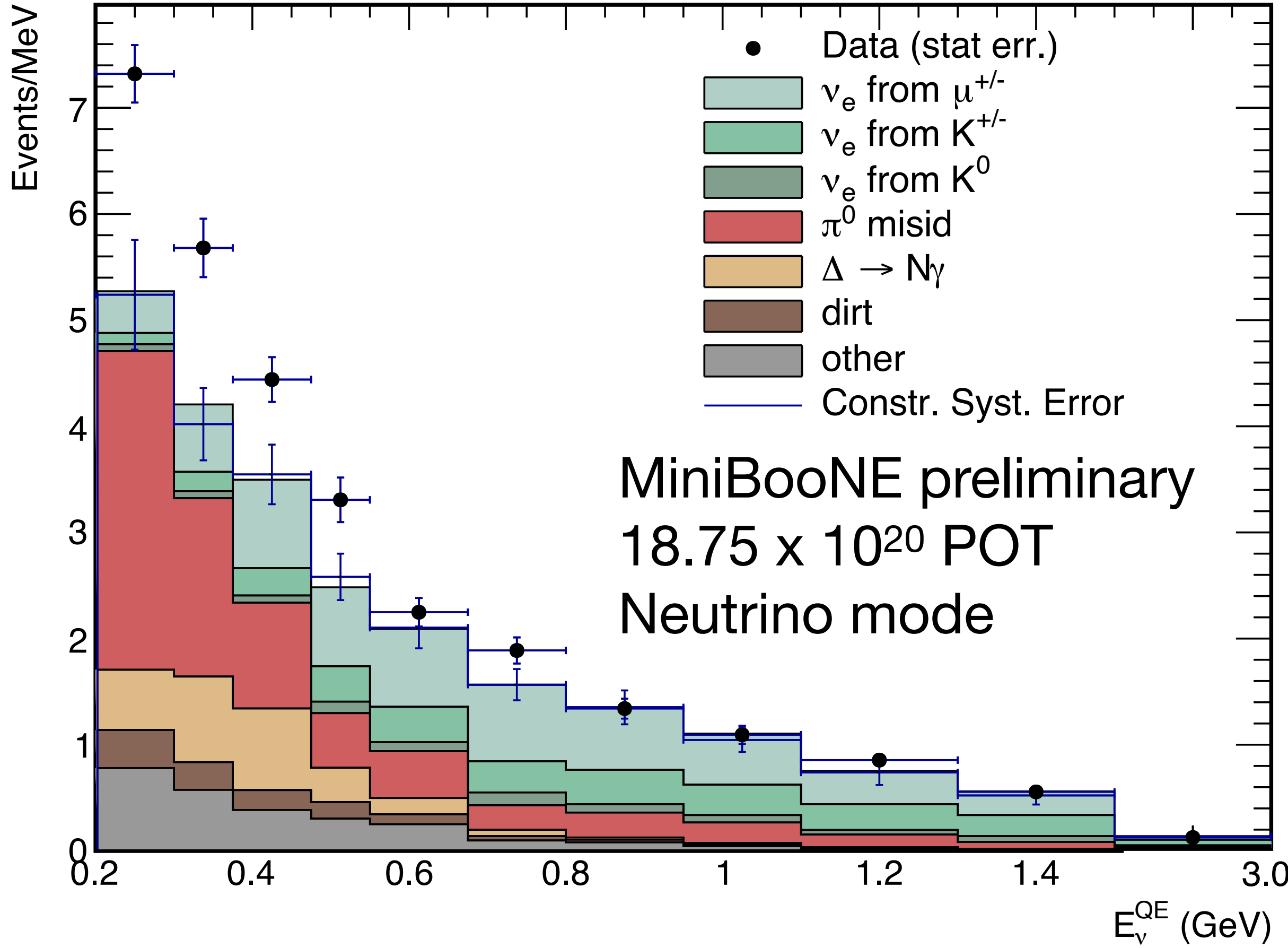
- Excess of data events with respect to our background prediction
- We report an excess of 560.6 ± 119.6 electron-like events (neutrino mode)
- **Significance : 4.7 σ in neutrino mode only**

ν_e -like excess stable across 3 runs

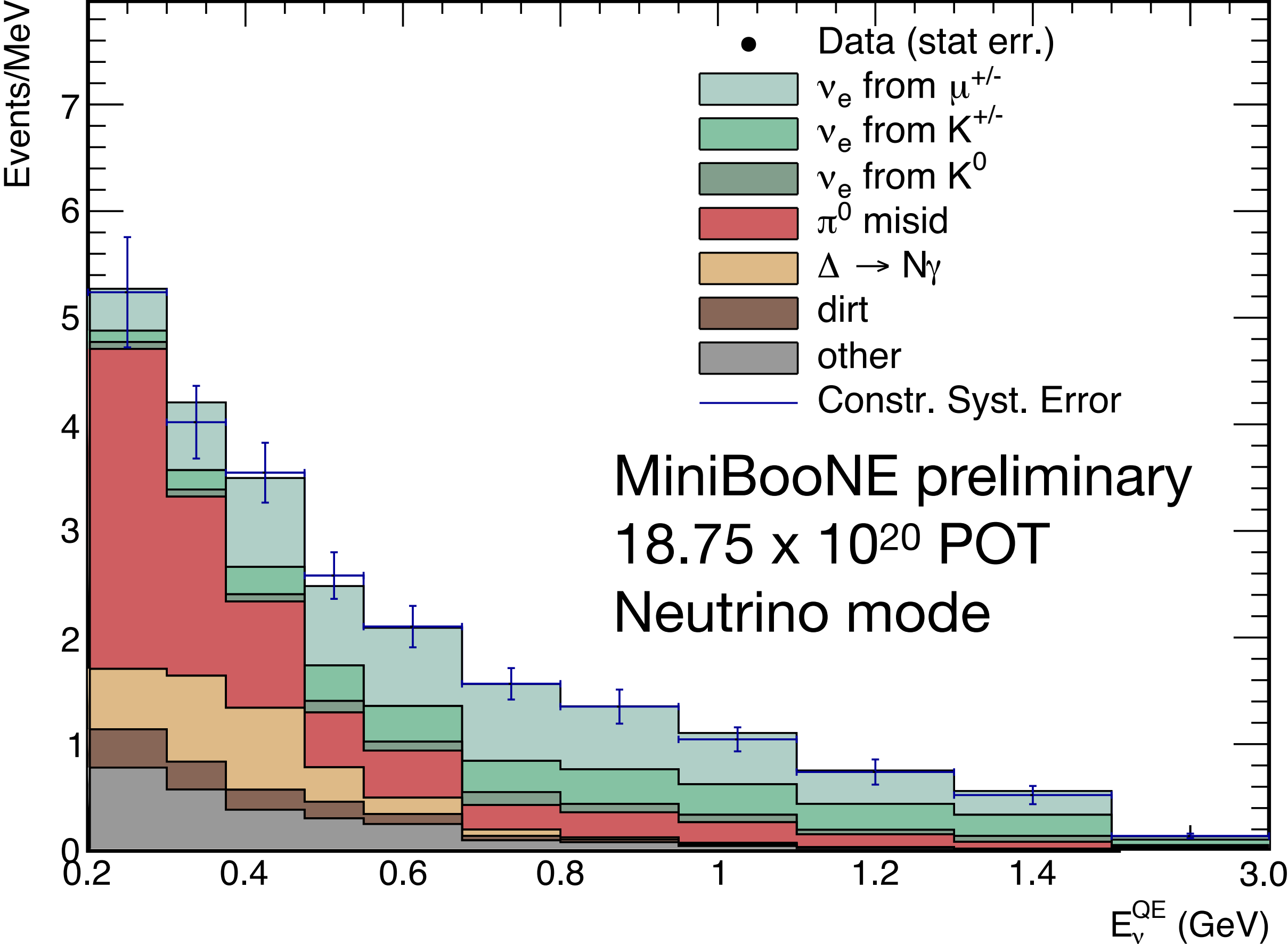


- Comparing the data-prediction excess for three data taking periods in neutrino mode
- Comparable statistics between the three data releases
- **The observed excess remains well compatible between the three data sets**

Constraining the backgrounds



Constraining the backgrounds

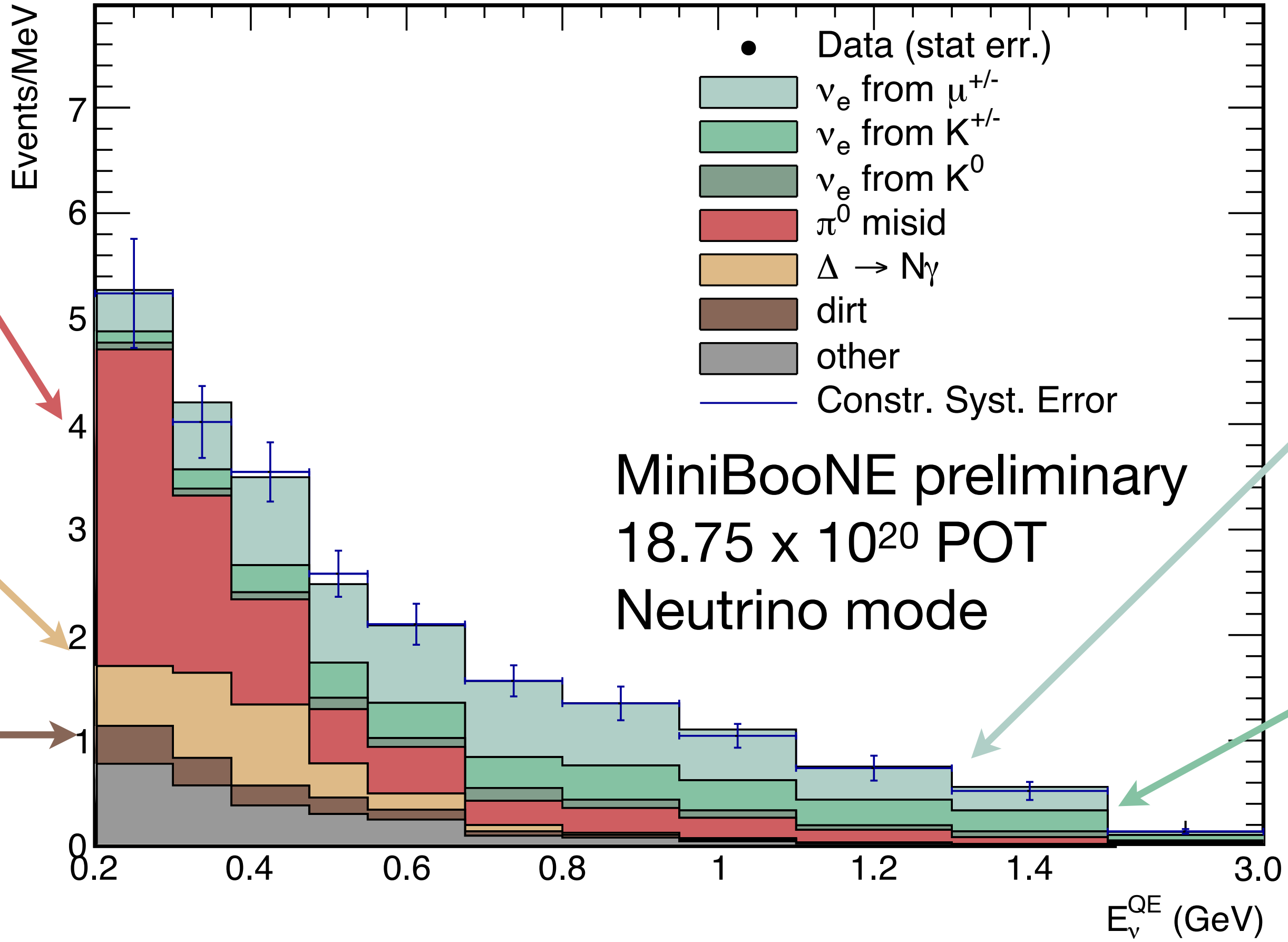


Constraining the backgrounds

π^0 MisID
 constrained from *in situ* measurement of NC π^0 rate

$\Delta \rightarrow N\gamma$ resonance
 constrained from *in situ* measured NC π^0 rate and theoretical prediction

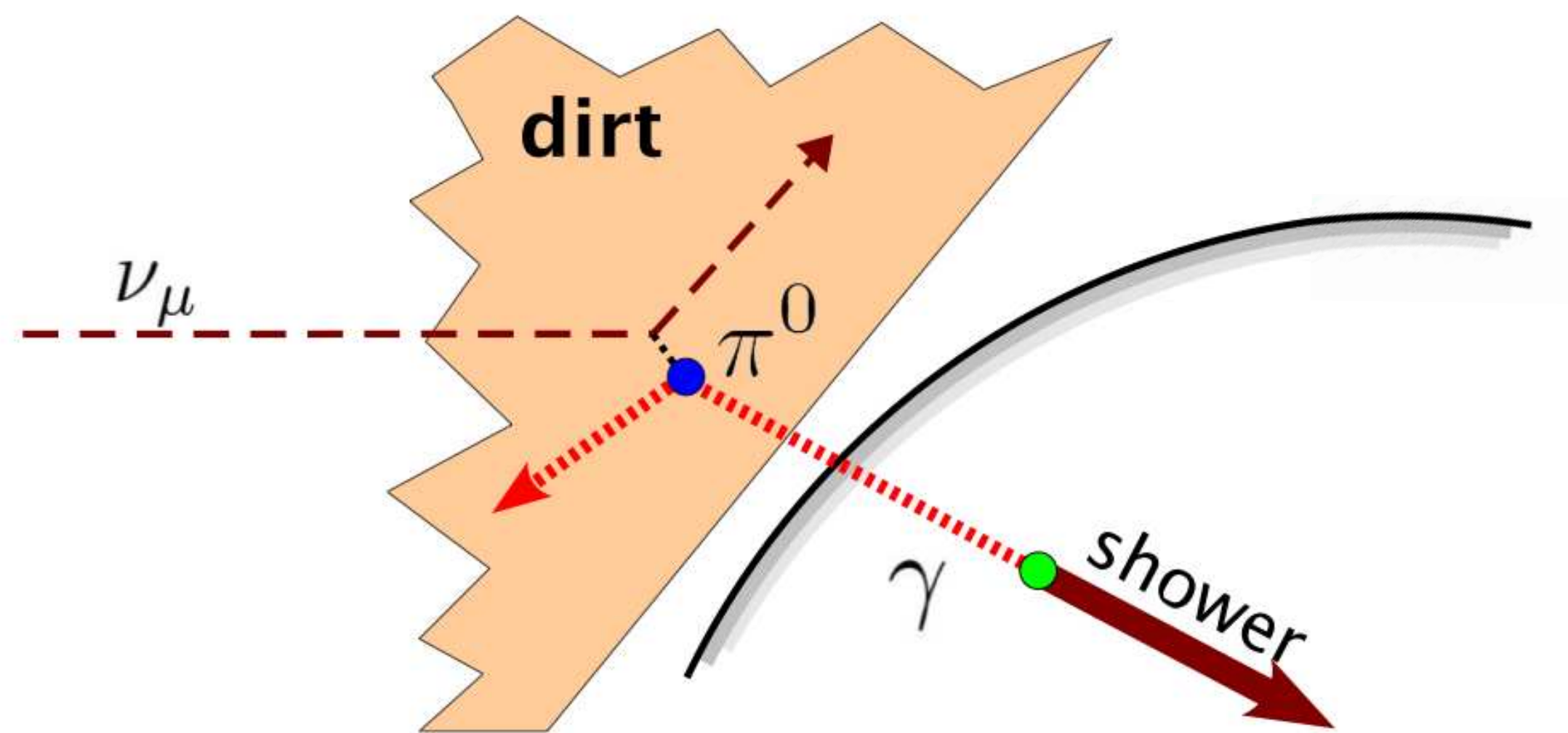
Dirt
 constrained from *in situ* dirt data sample



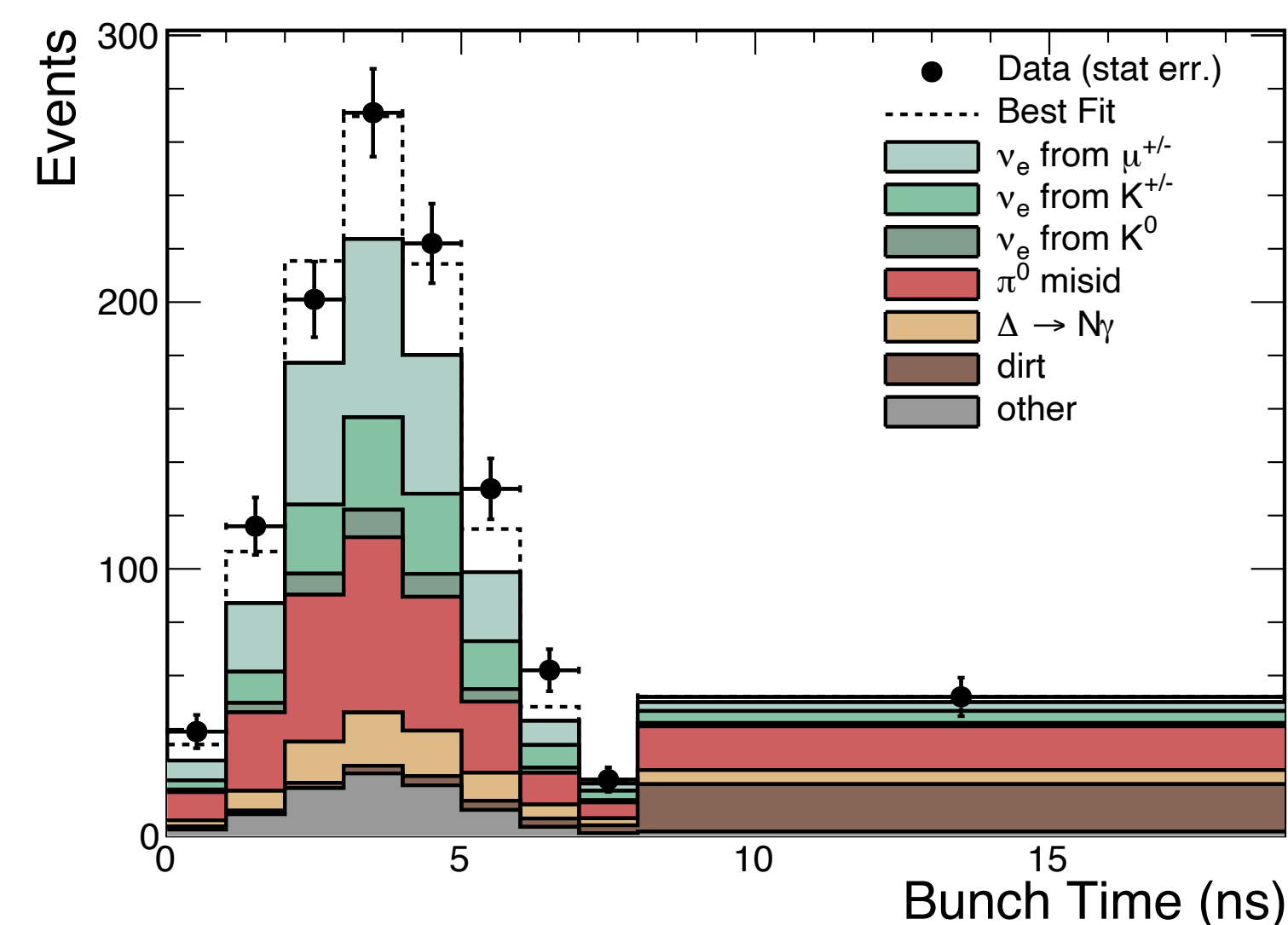
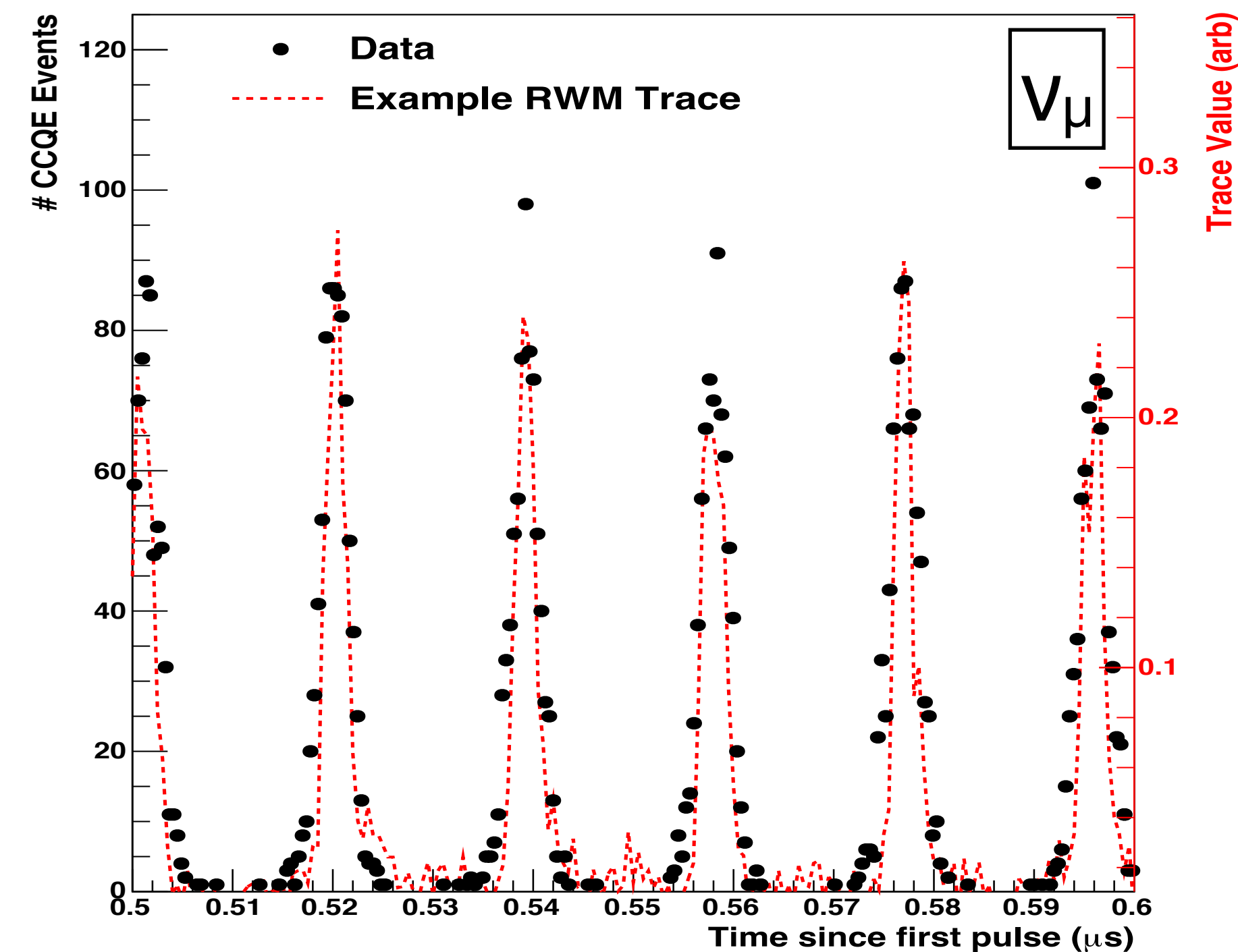
ν_e from μ decay
 is constrained by *in situ* ν_μ CCQE measurement

ν_e from K decay
 constrained from *in situ* high energy events + SciBooNE high energy ν_μ event rate

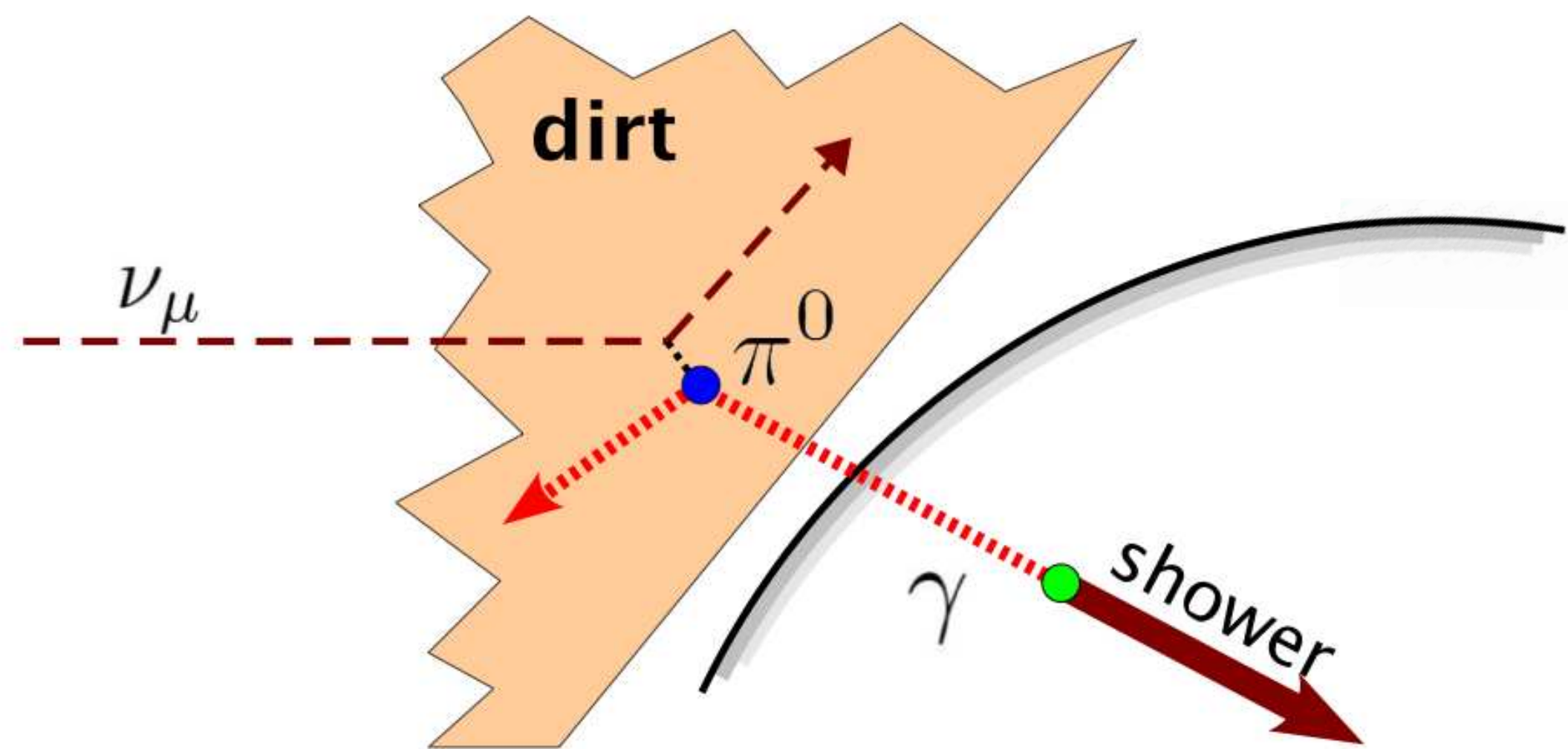
Dirt constraint with timing



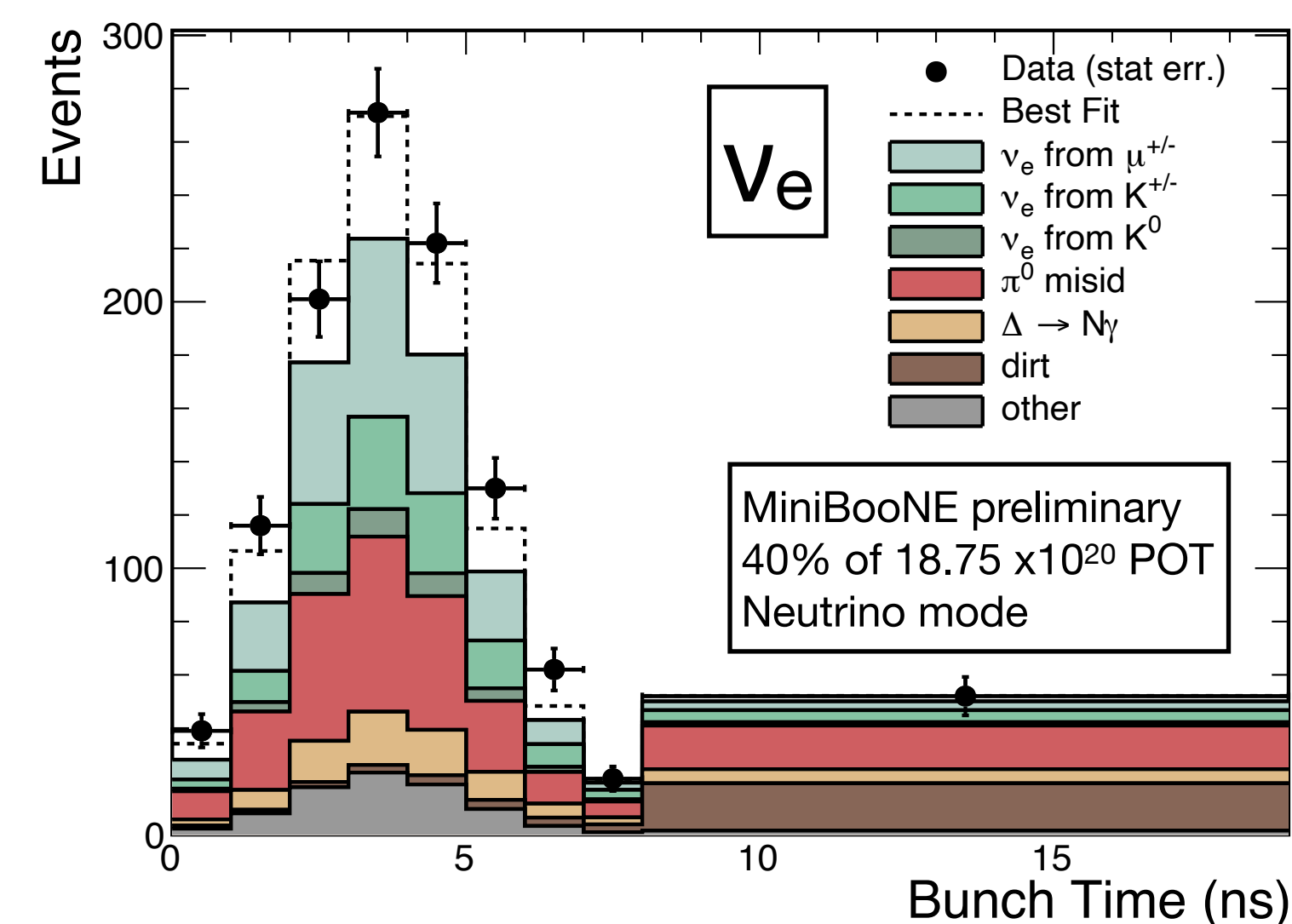
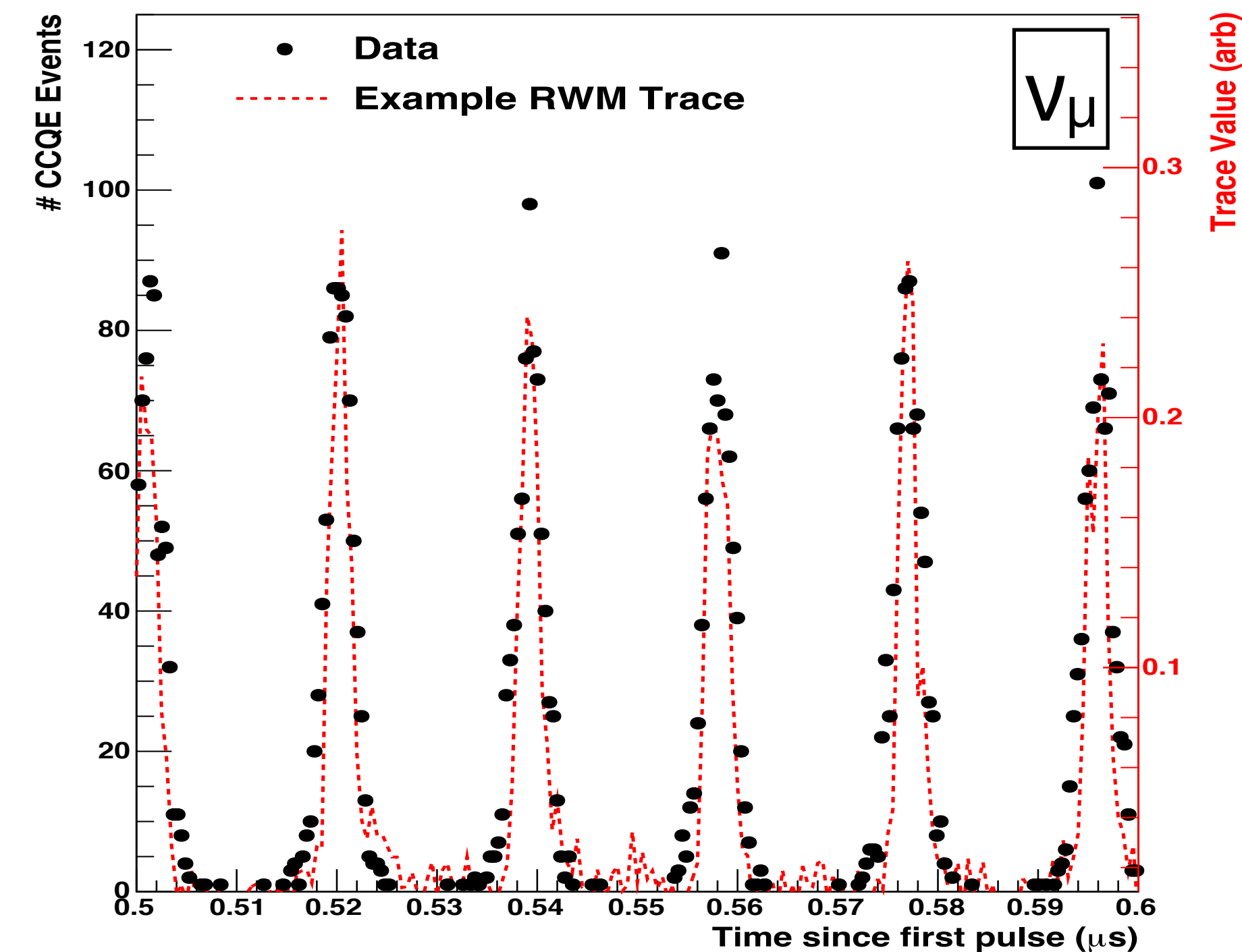
- Dirt events:
 - ➔ beam-related neutrino interactions in the rocks surrounding the detector
 - ➔ time shift due to extra flight path before particles enter the detector
- No cut on the event timing within the beam spill (RF cavity structure of 52.81 MHz)
- Event timing shows no significant excess of off-bunch data
 - ➔ **dirt constrained to better than 5σ**



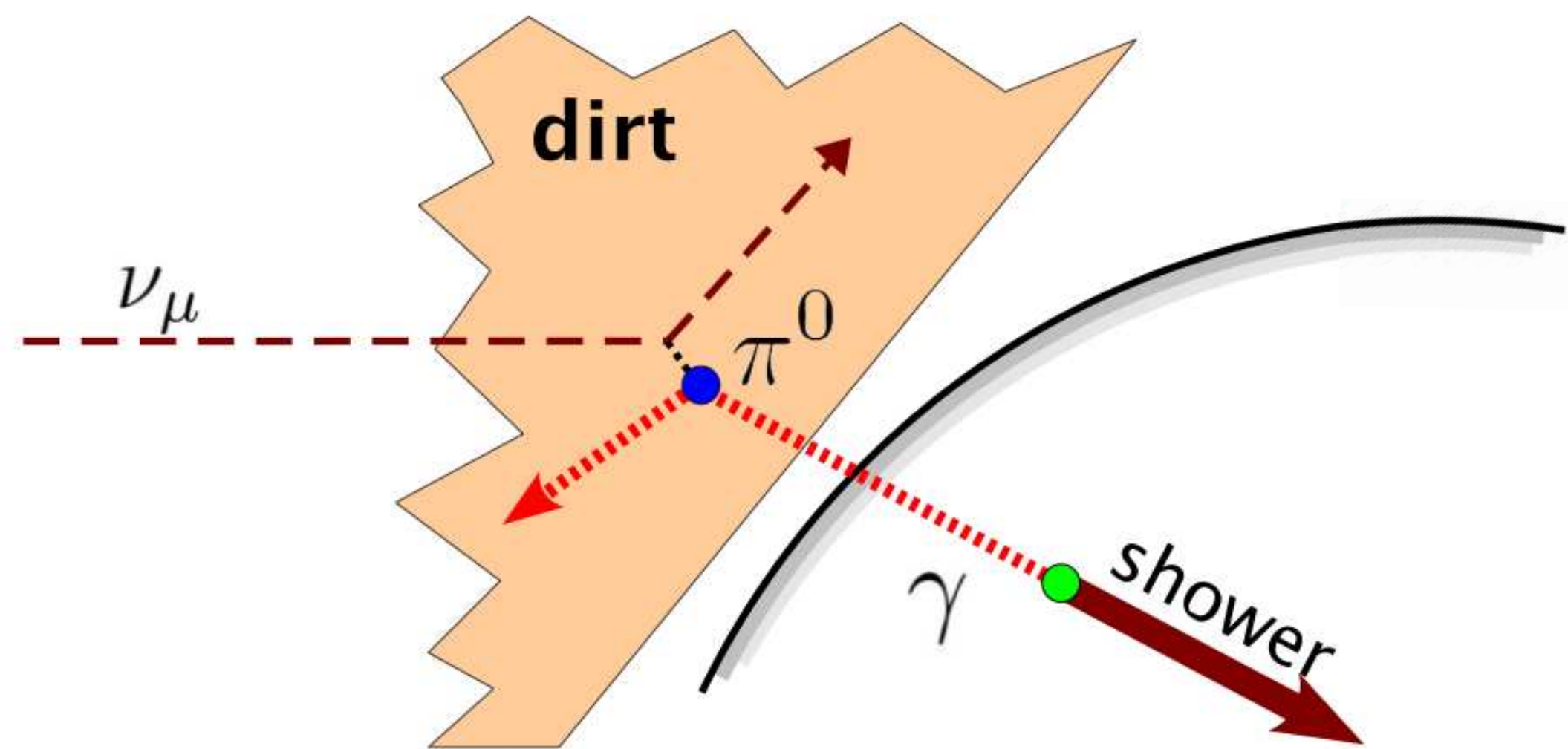
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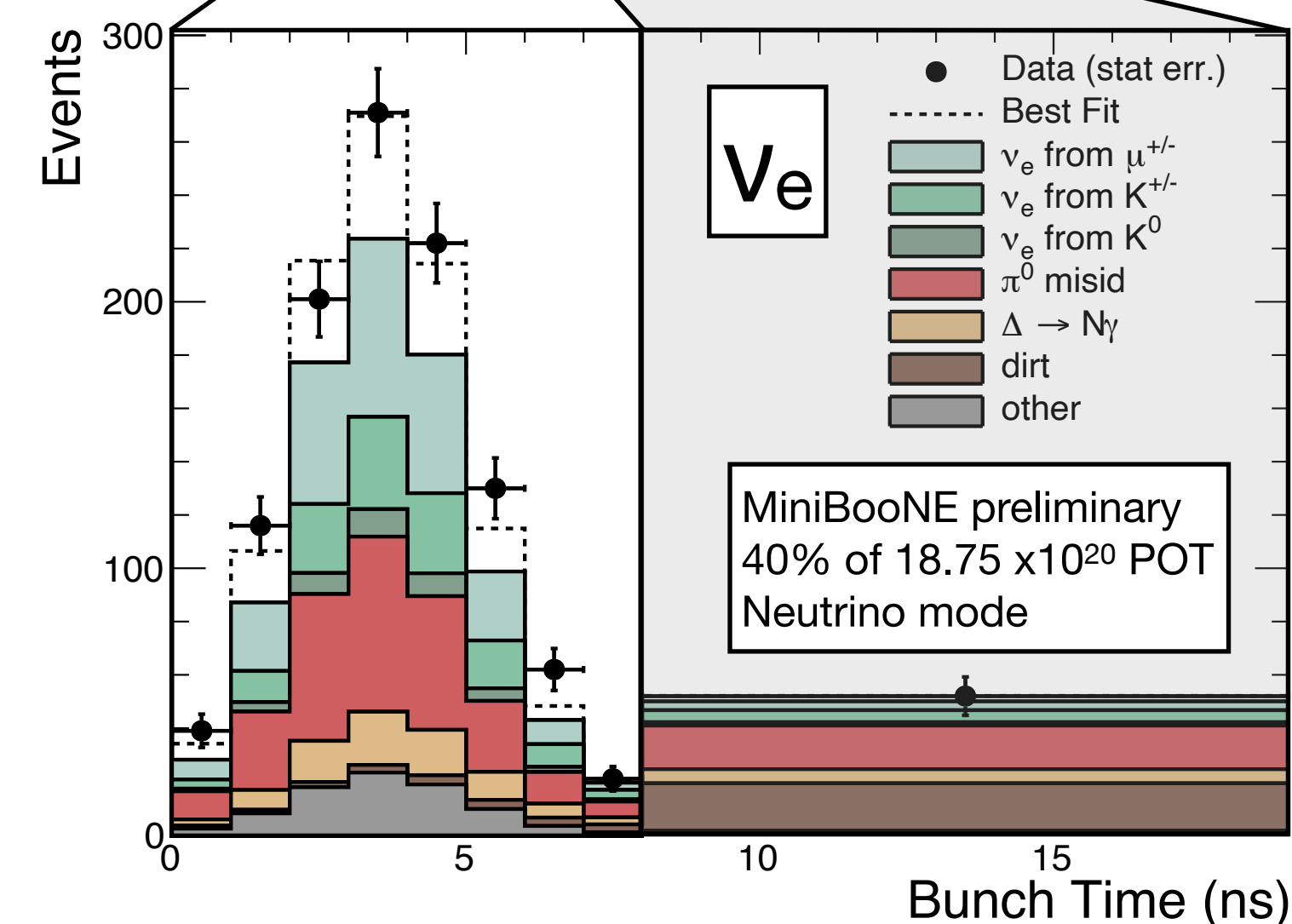
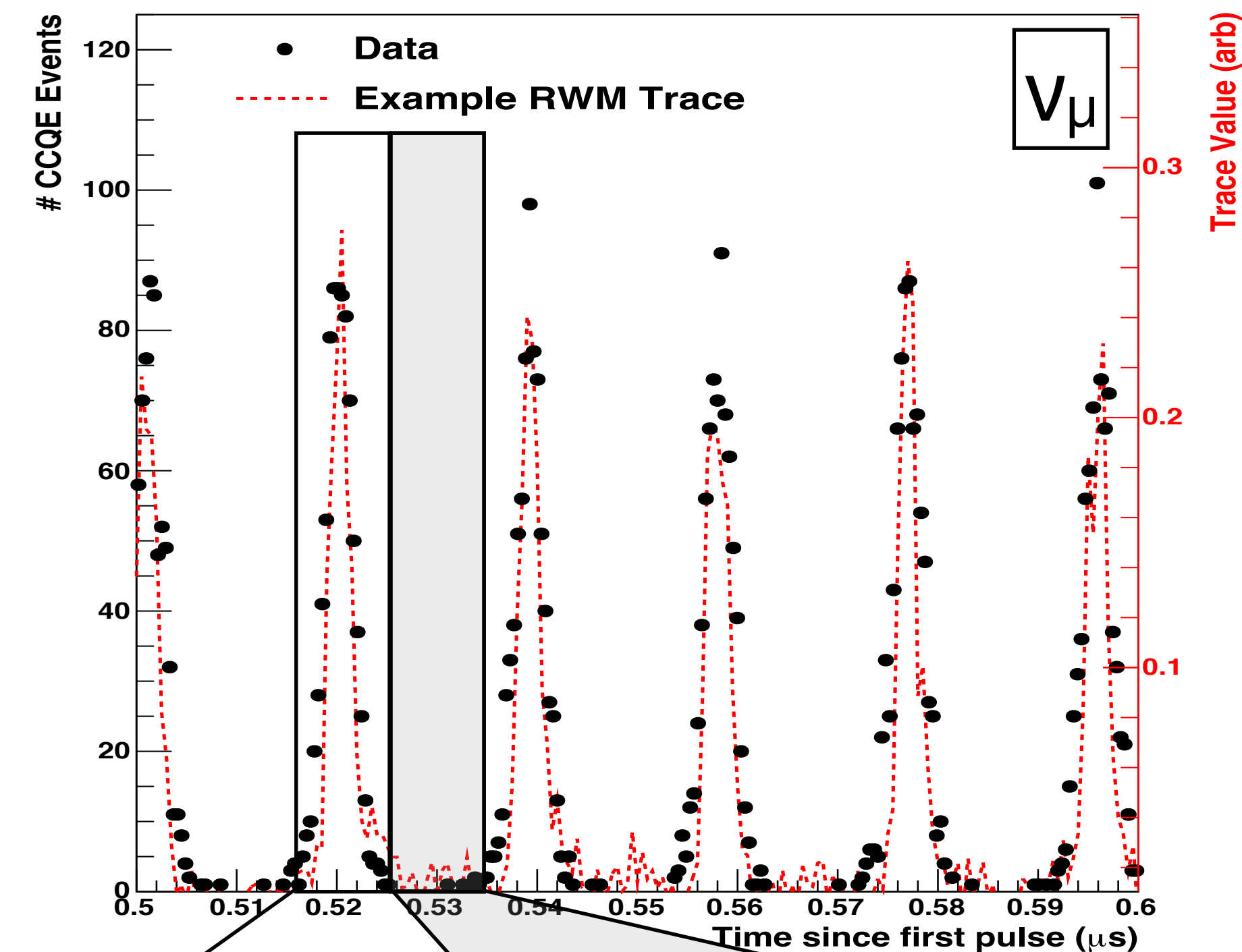
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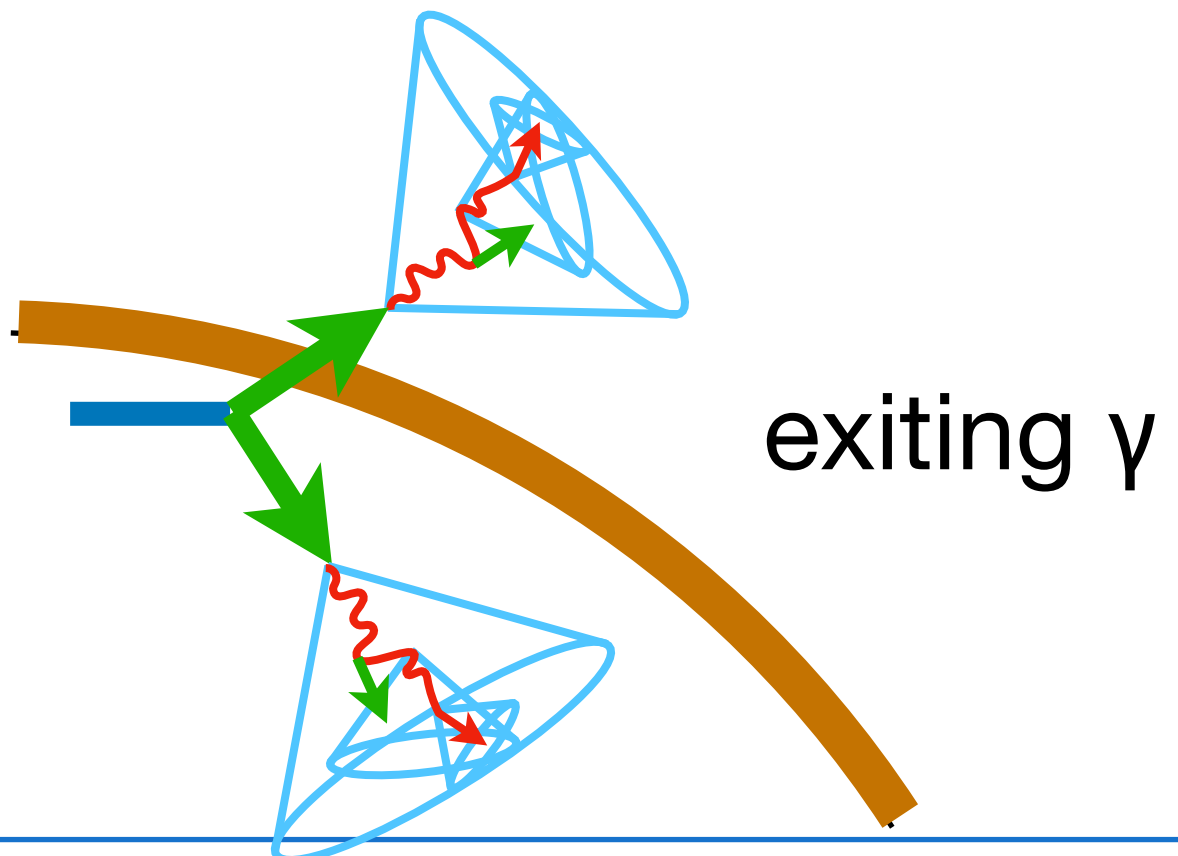
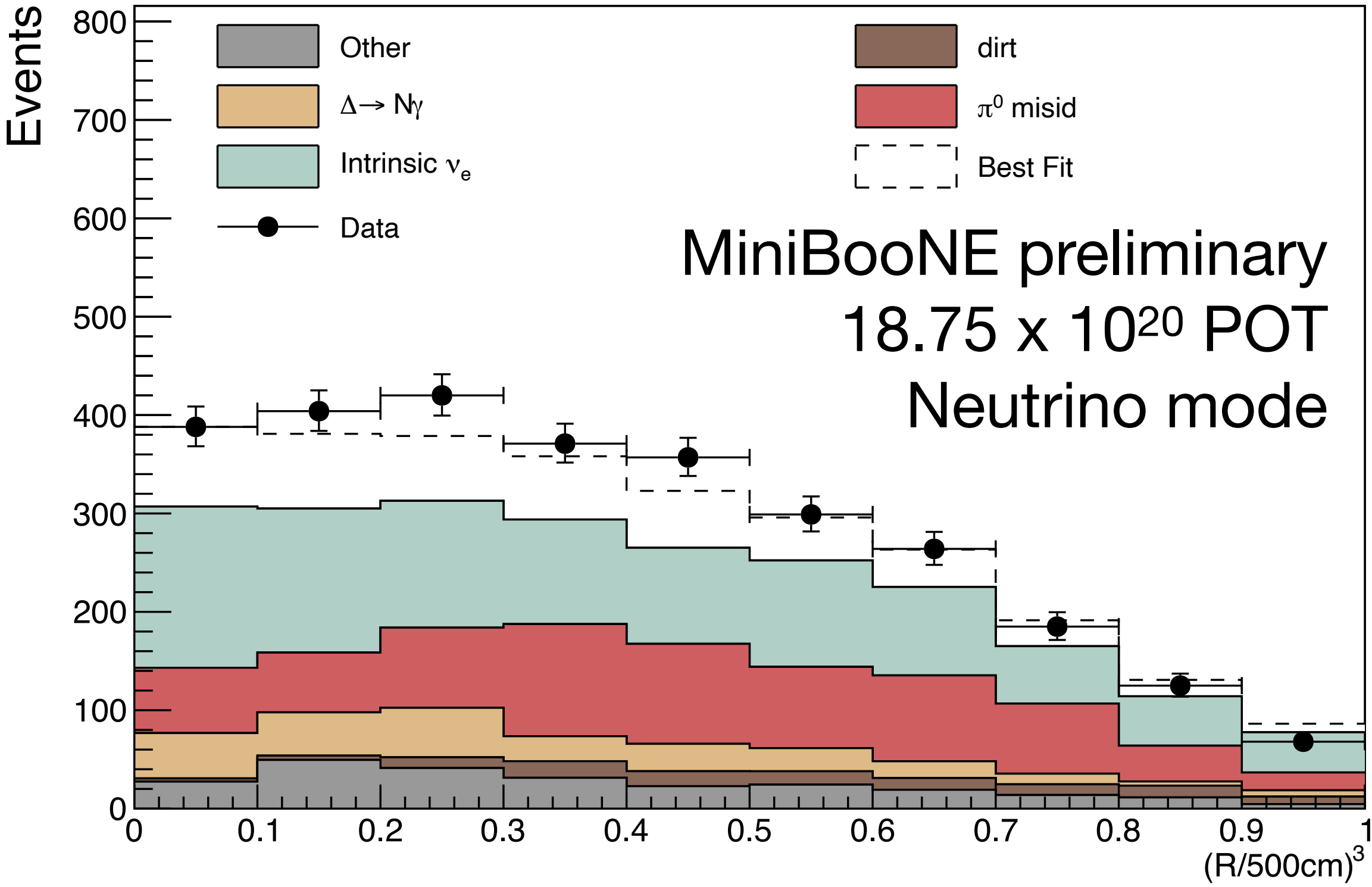
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Radial distribution disfavors dirt and π^0 background



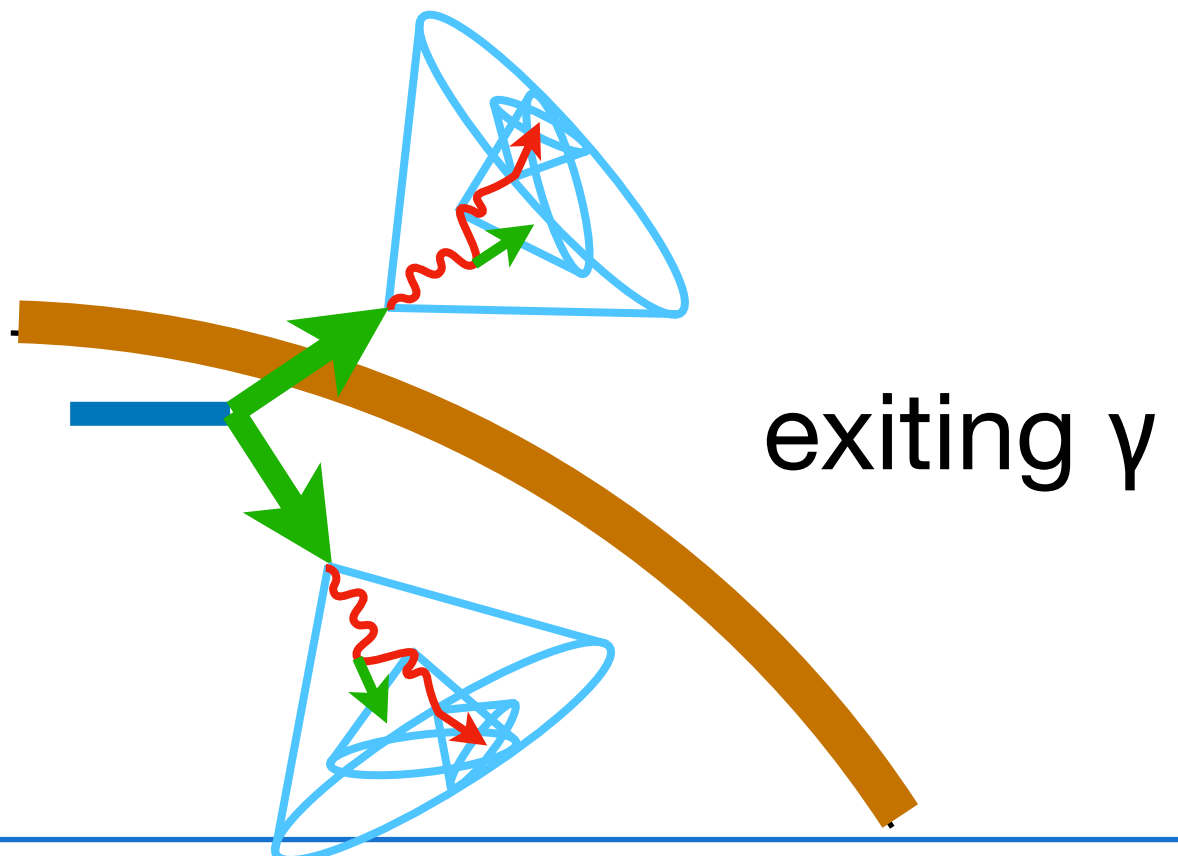
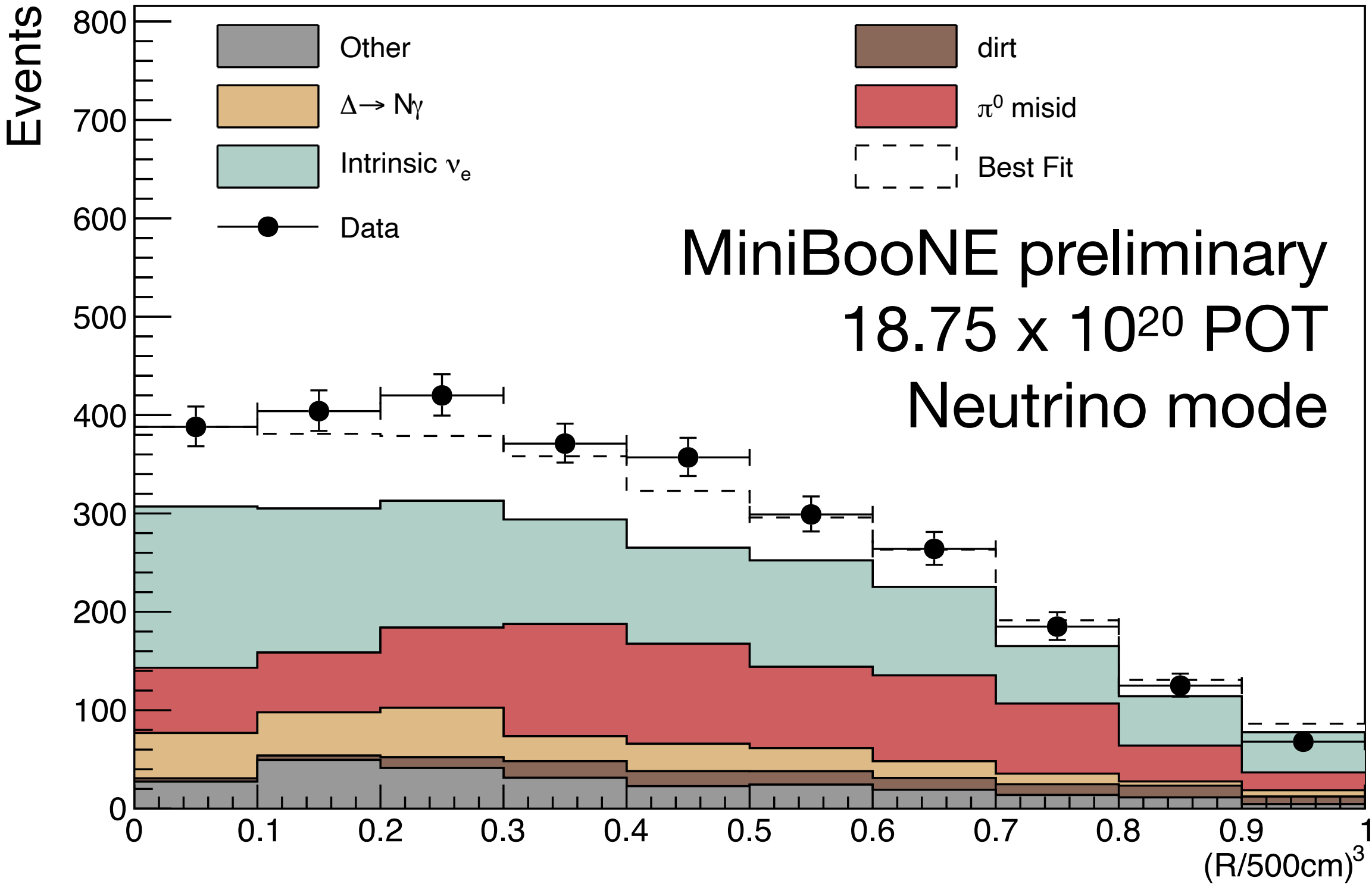
- Improved statistics allow for more distributions to be investigated
- Radial distribution shows that the excess is spread evenly within the volume of the detector
- An excess of π^0 background would have peaked near the edge (higher probability of missing one of the γ)
- Similar approach to SNO's CC/NC constraint
- Second best candidate : NC γ background

Excess shape tests

Hypothesis	Multiplicative factor	$\chi^2/9ndf$
NC $\Delta \rightarrow N\gamma$ Background	3.18	10.0
External Event Background	5.98	44.9
ν_e & $\bar{\nu}_e$ from K_L^0 Decay Background	7.85	14.8
ν_e & $\bar{\nu}_e$ from K^\pm Decay Background	2.95	16.3
ν_e & $\bar{\nu}_e$ from μ^\pm Decay Background	1.88	16.1
Other ν_e & $\bar{\nu}_e$ Background	3.21	12.5
NC π^0 Background	1.75	17.2
Best Fit Oscillations	1.24	8.4

Statistics only χ^2

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Statistics only χ^2

$NC \Delta \rightarrow N\gamma$ resonance

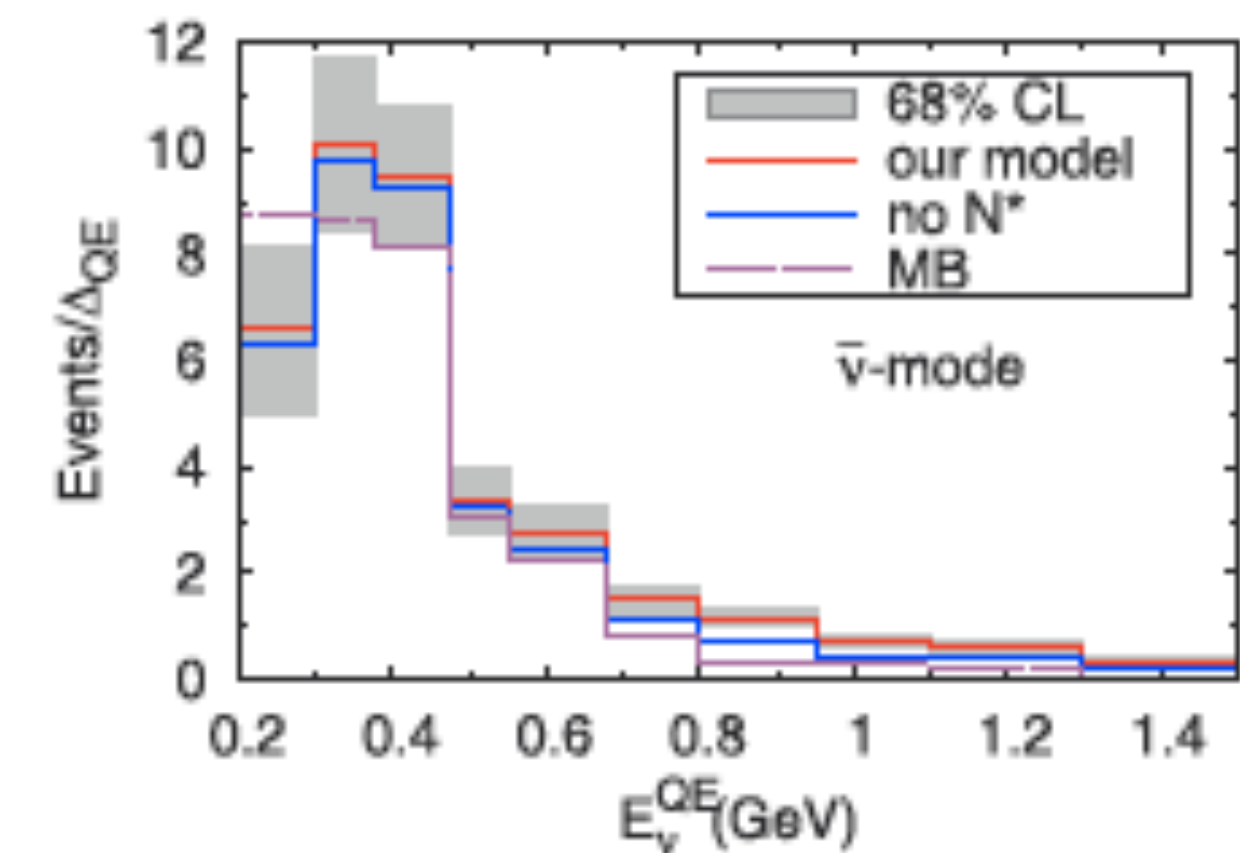
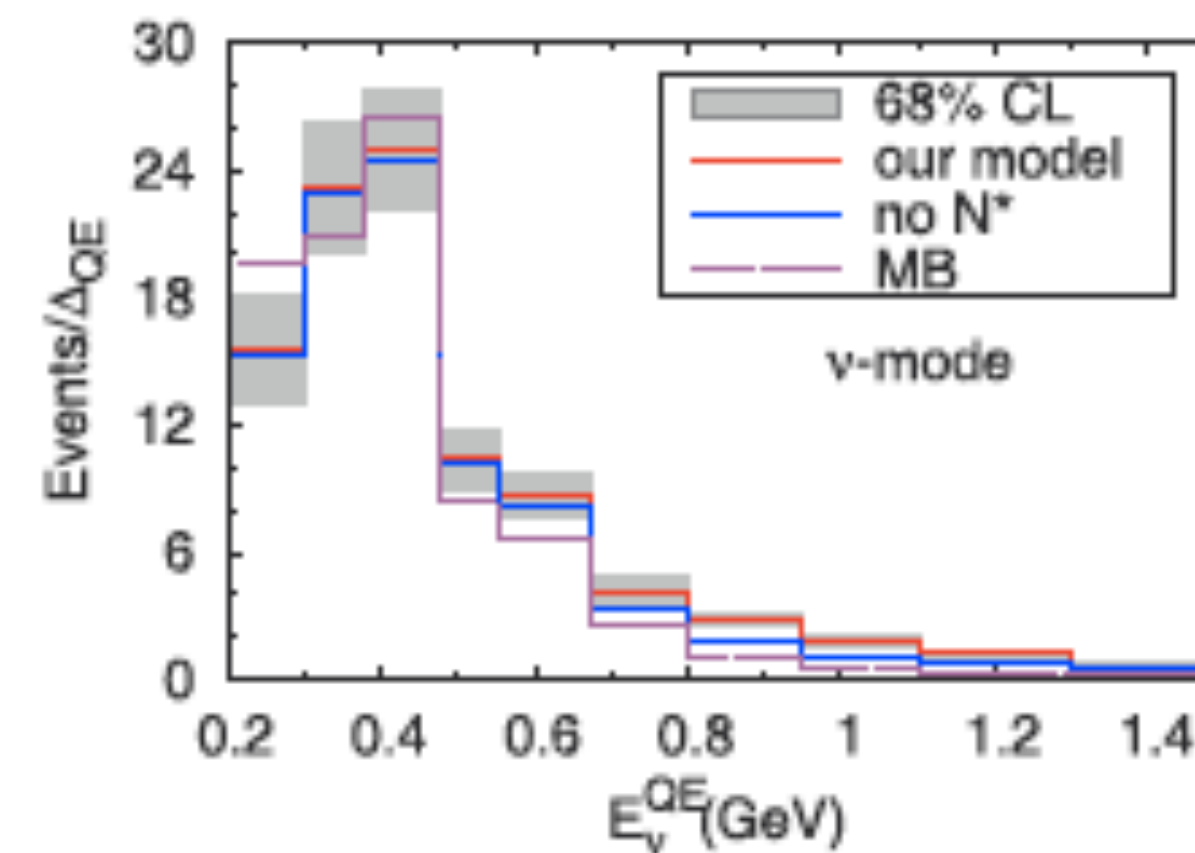
- The production of $NC \Delta \rightarrow N\gamma$ is highly correlated to the measurement of $NC \pi^0$
- Same probability of a NC interaction, the difference in final state is the relative rate of resonant production.
 - ➔ Our predicted single γ/π^0 ratio is $\sim 0.9\%$, which takes into account pion absorption in the nucleus, higher mass resonances, coherent scattering, and non-resonant processes
- Apply the same correction and fractional uncertainties to $NC \Delta \rightarrow N\gamma$ as $NC \pi^0$
- Additional uncertainty to account for final state interactions (FSI)
- The single gamma estimate agrees with theory

Single photon events from neutral current interactions at MiniBooNE

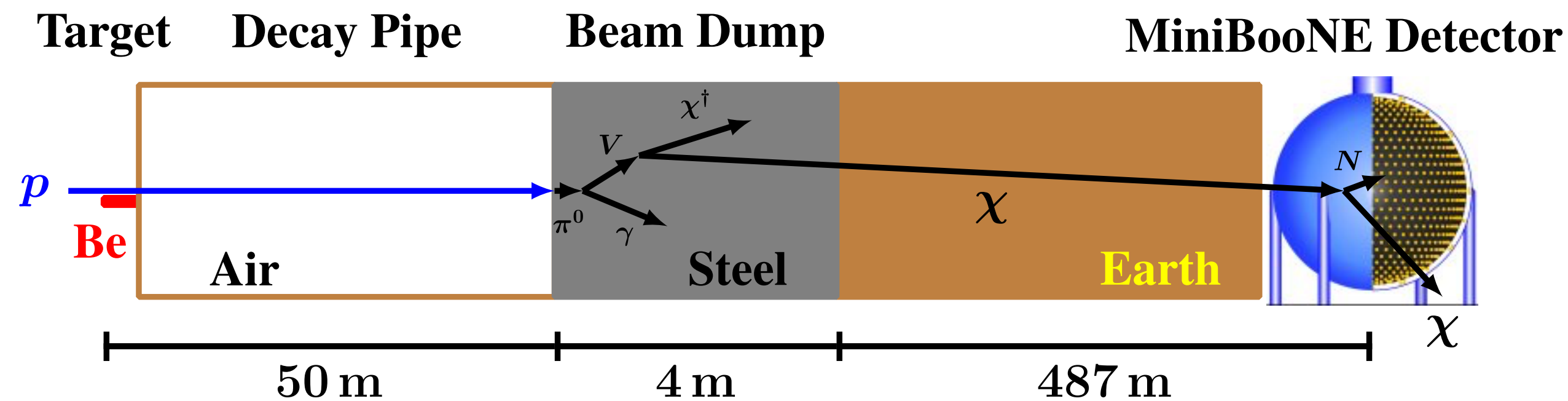
En Wang, Luis Alvarez-Ruso*, Juan Nieves

Instituto de Física Corpuscular (IFIC), Centro Mixto CSIC-Universidad de Valencia, Institutos de Investigación de Paterna, Apartado 22085, E-46071 Valencia, Spain

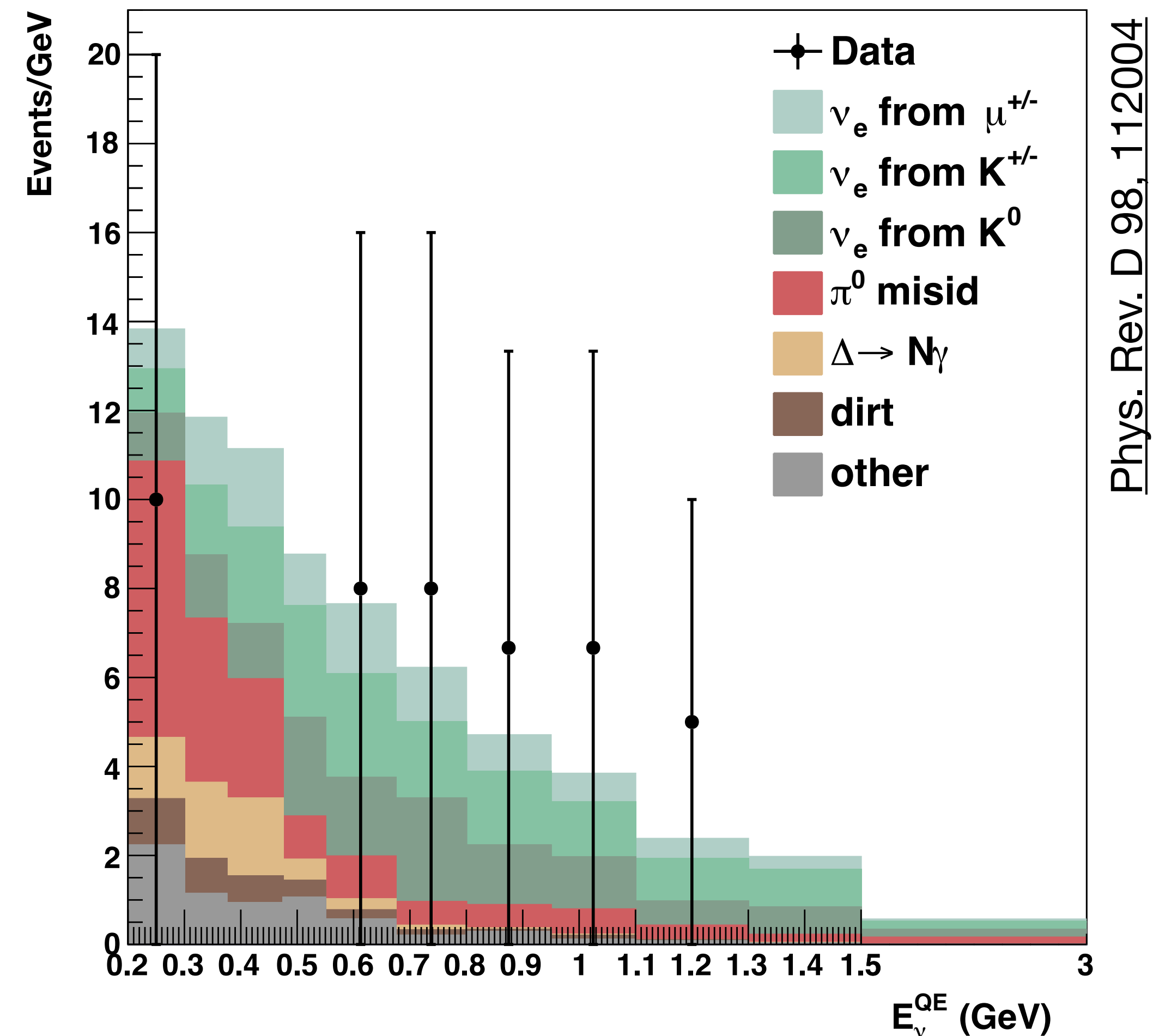
Phys. Lett. B 740 (2015) 16-22



Constraints from the beam dump run

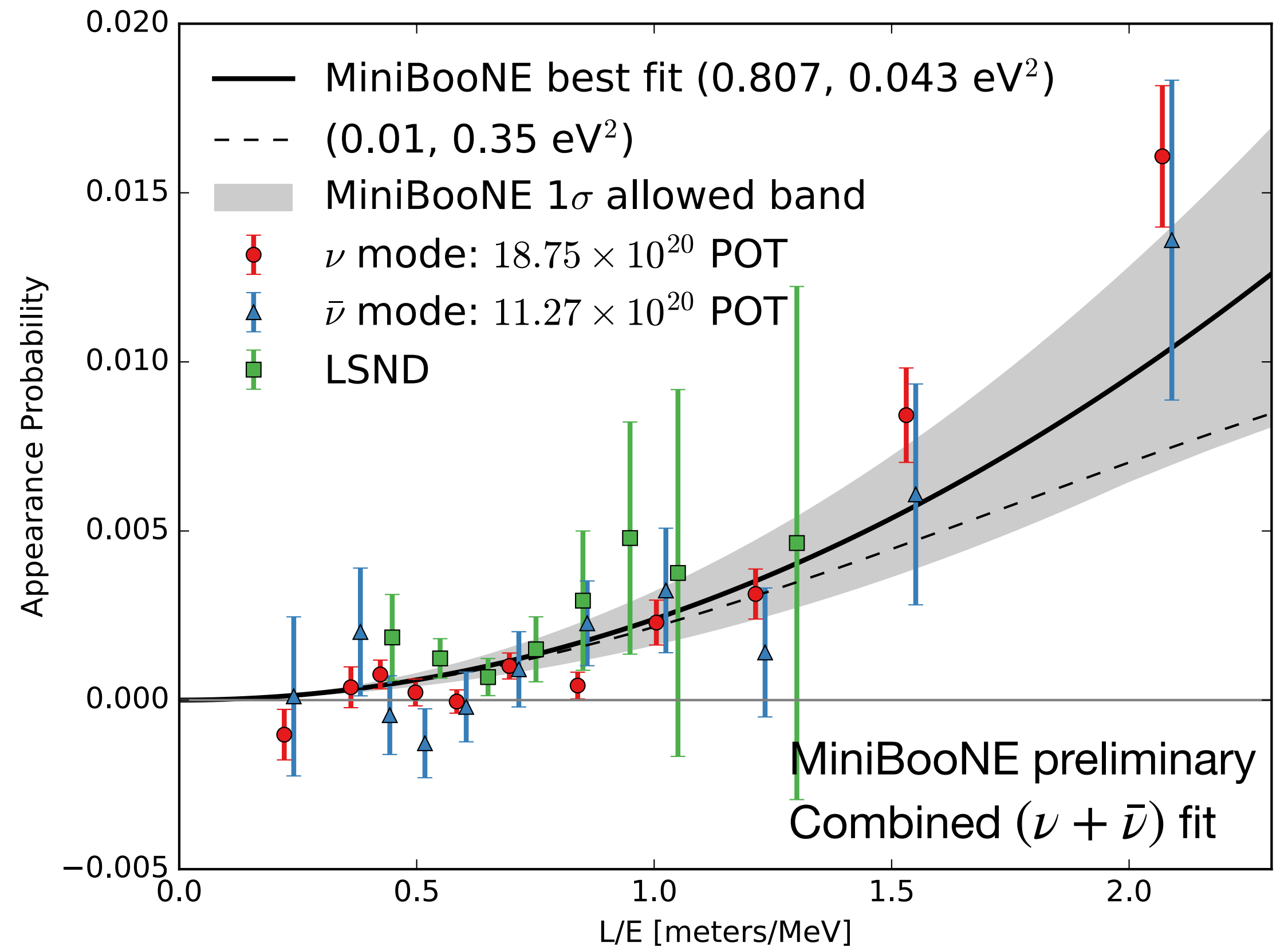
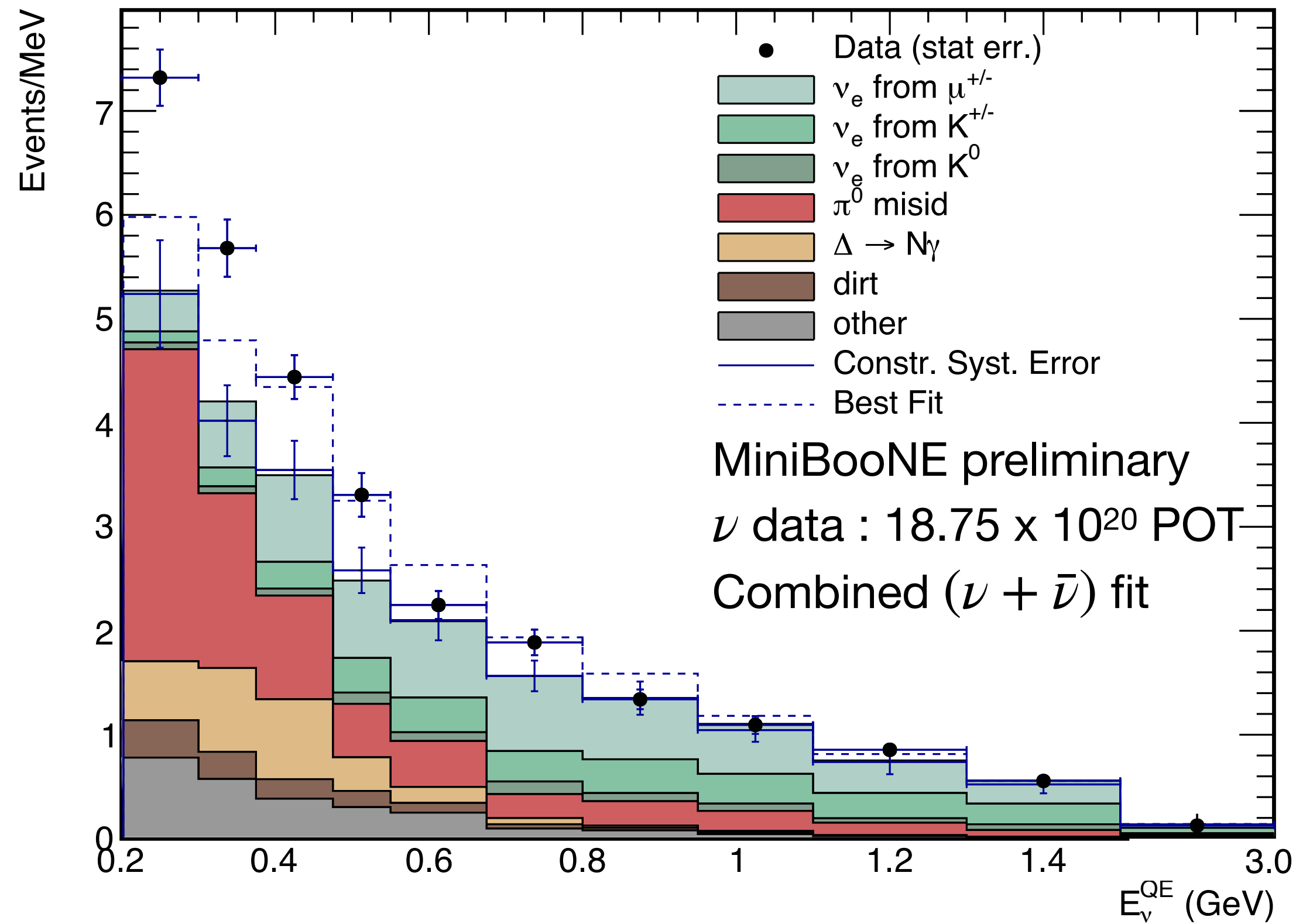


- Reduction in neutrino production by a factor of ~ 50
- No change to neutral meson production and proton bremsstrahlung to first order
 - ➔ Can directly test models that predict the oscillation excess does not scale as neutrino scales (e.g. vector portal, inelastic dark matter, ...)
- **Expected** : 35.5 ± 7.4 excess events in $[200, 1250]$ MeV for a POT-scaling excess
- **Measured** : 6 events, 8.8 backgrounds expected
 - ➔ -2.8 excess events
 - ➔ Explanation that scale only by POT instead of neutrino production are **disfavoured out at 4.6σ**



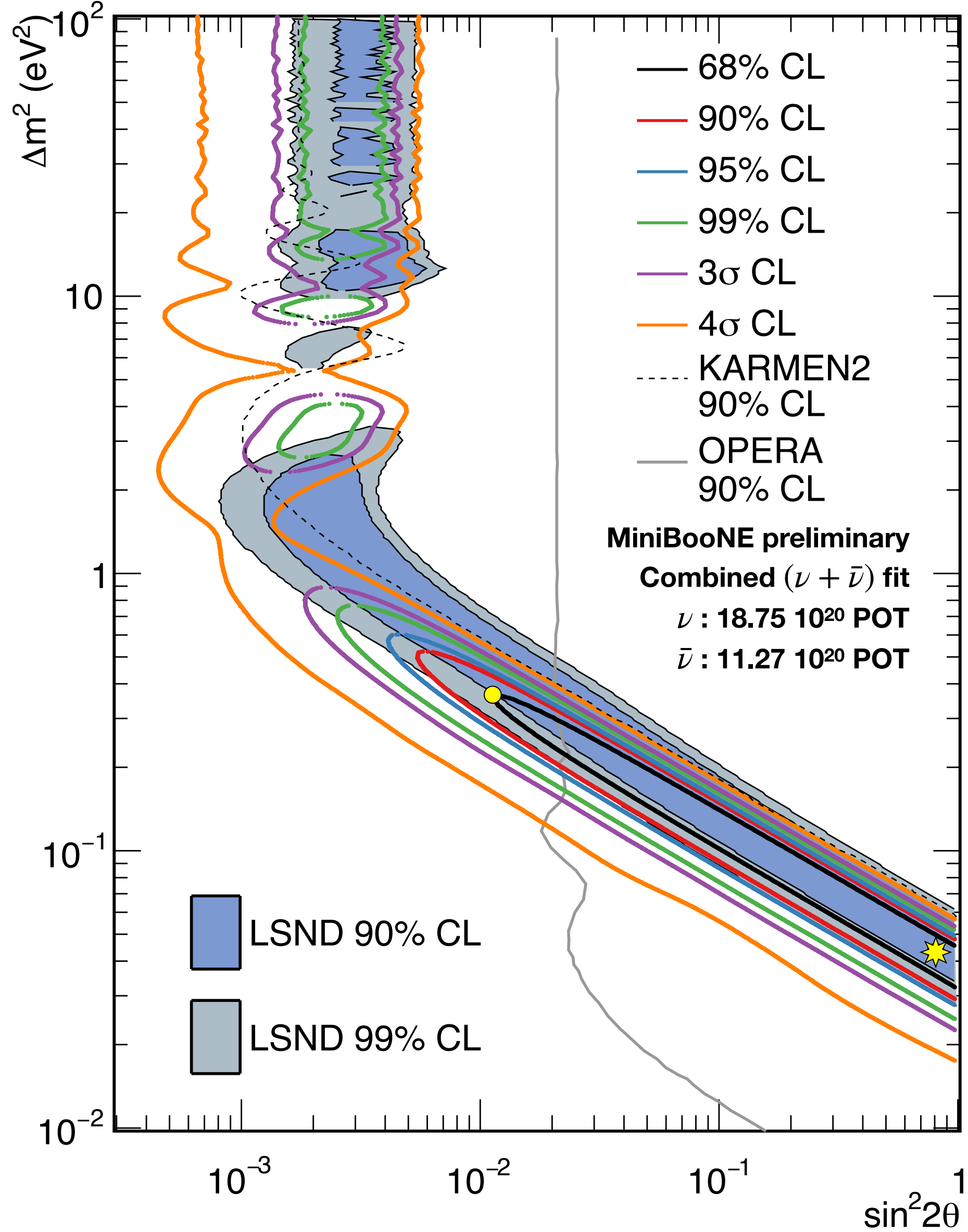
See Jordan et al. Phys. Rev. Lett. 122,081801 (2019) for constraints on different models

Excess interpretation in a sterile neutrino hypothesis

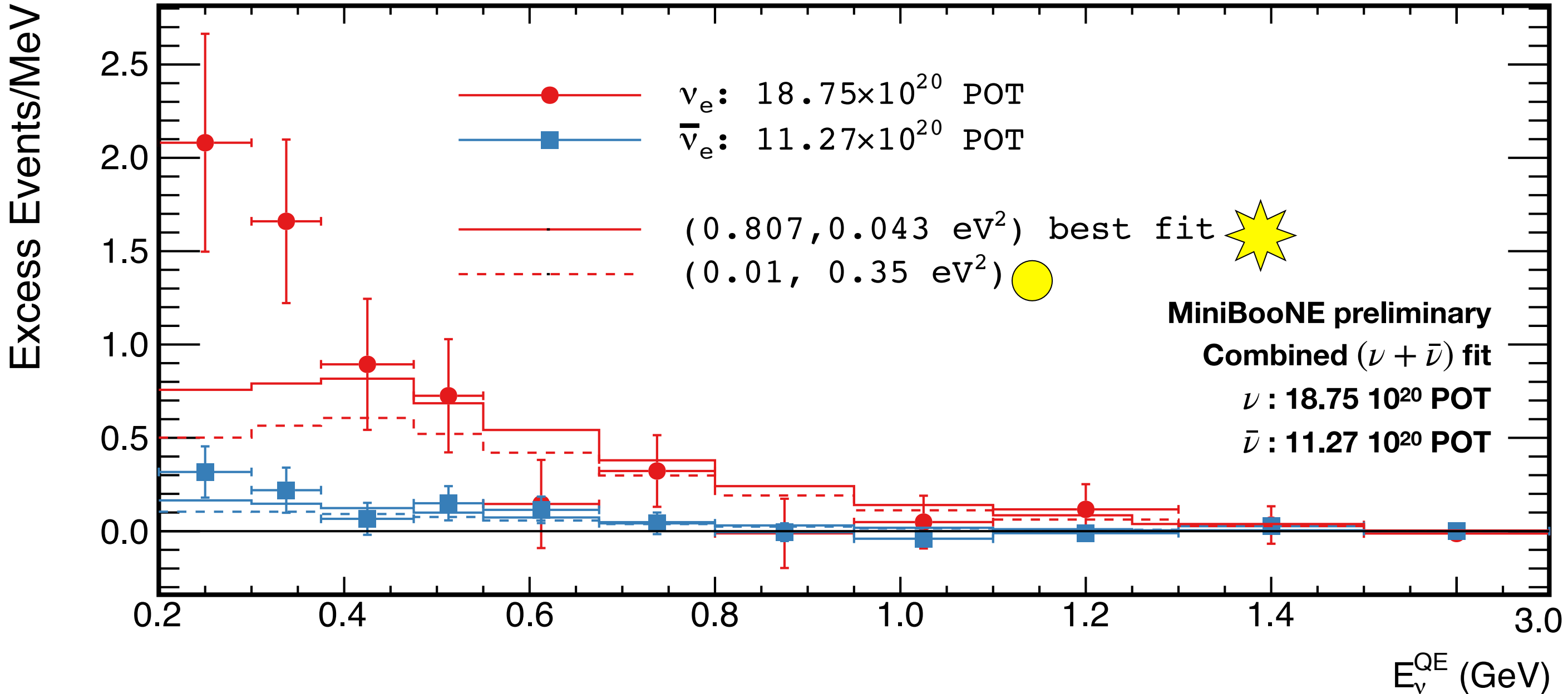


- Combined ($\nu + \bar{\nu}$) fit
- LSND and MiniBooNE points follow the same best fit 2ν oscillation interpretation

Preferred regions in sterile neutrino hypothesis



- Neutrino mode excess 4.7 σ ,
- **Neutrino+Anti-neutrino modes excess : 4.8 σ**

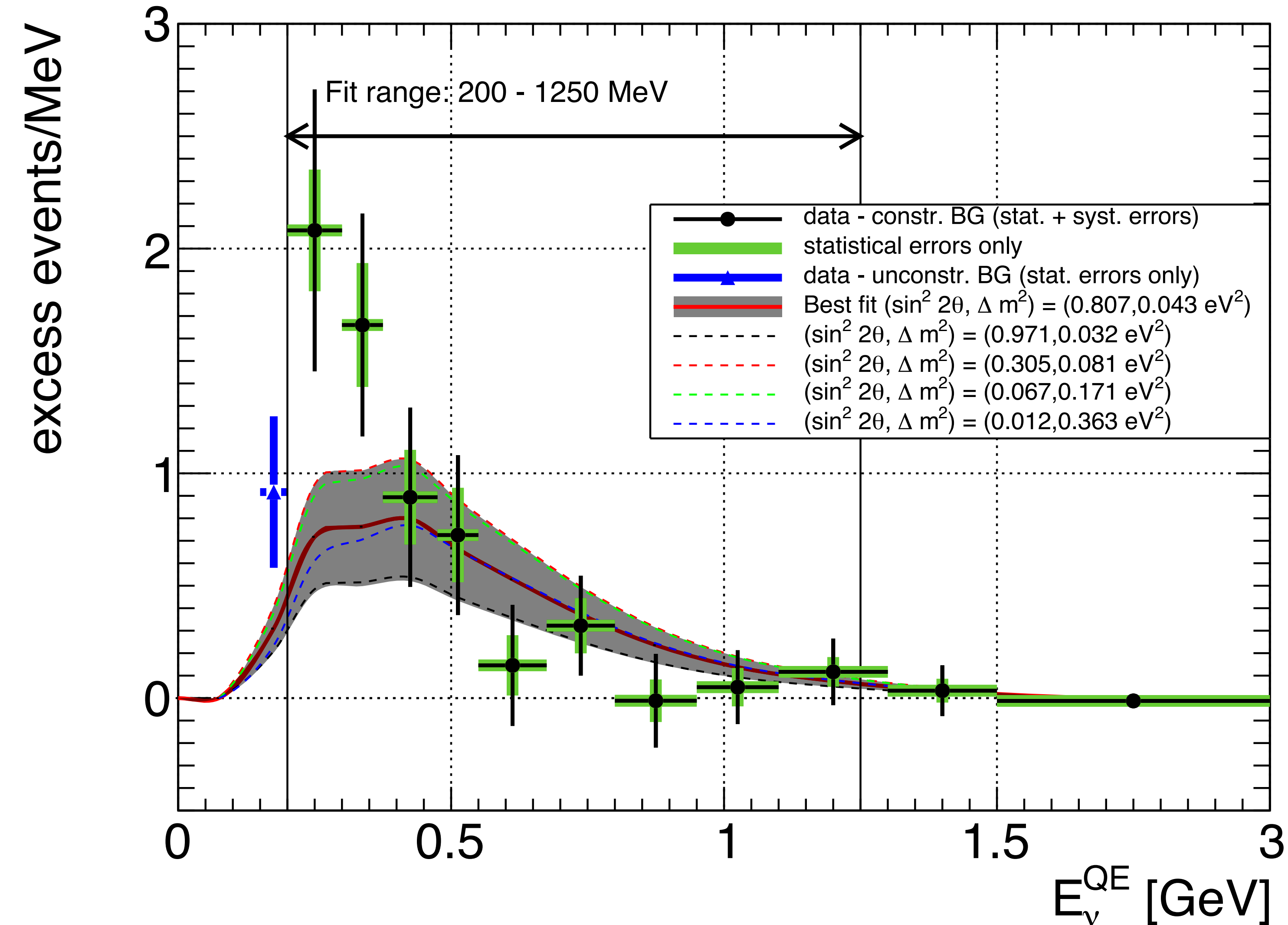


Neutrino + Anti-Neutrino Mode

$(\Delta m^2, \sin^2 2\theta) = (0.043 \text{ eV}^2, 0.807)$

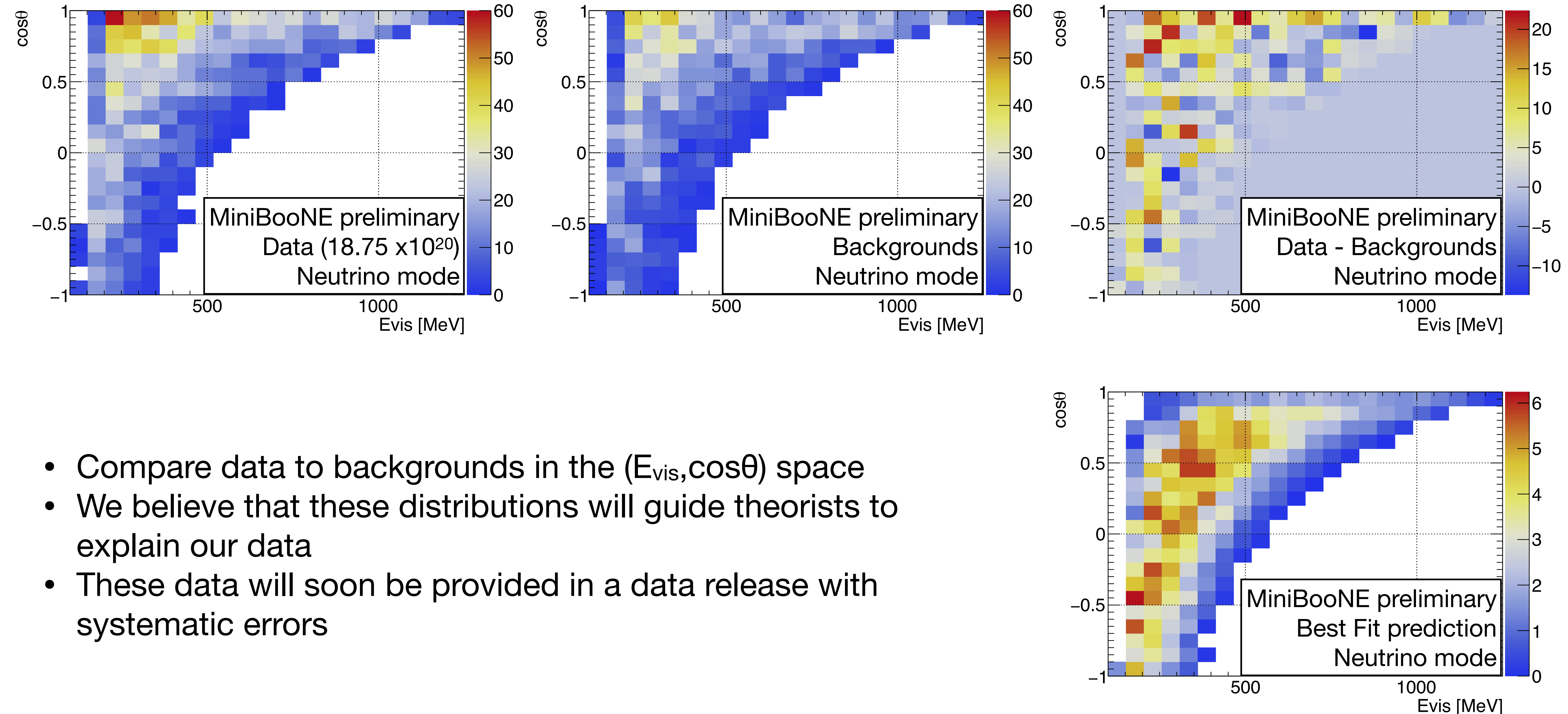
$\chi^2/ndf = 21.7/15.5$ (prob = 12.3%)

Full ν_e statistics excess plot



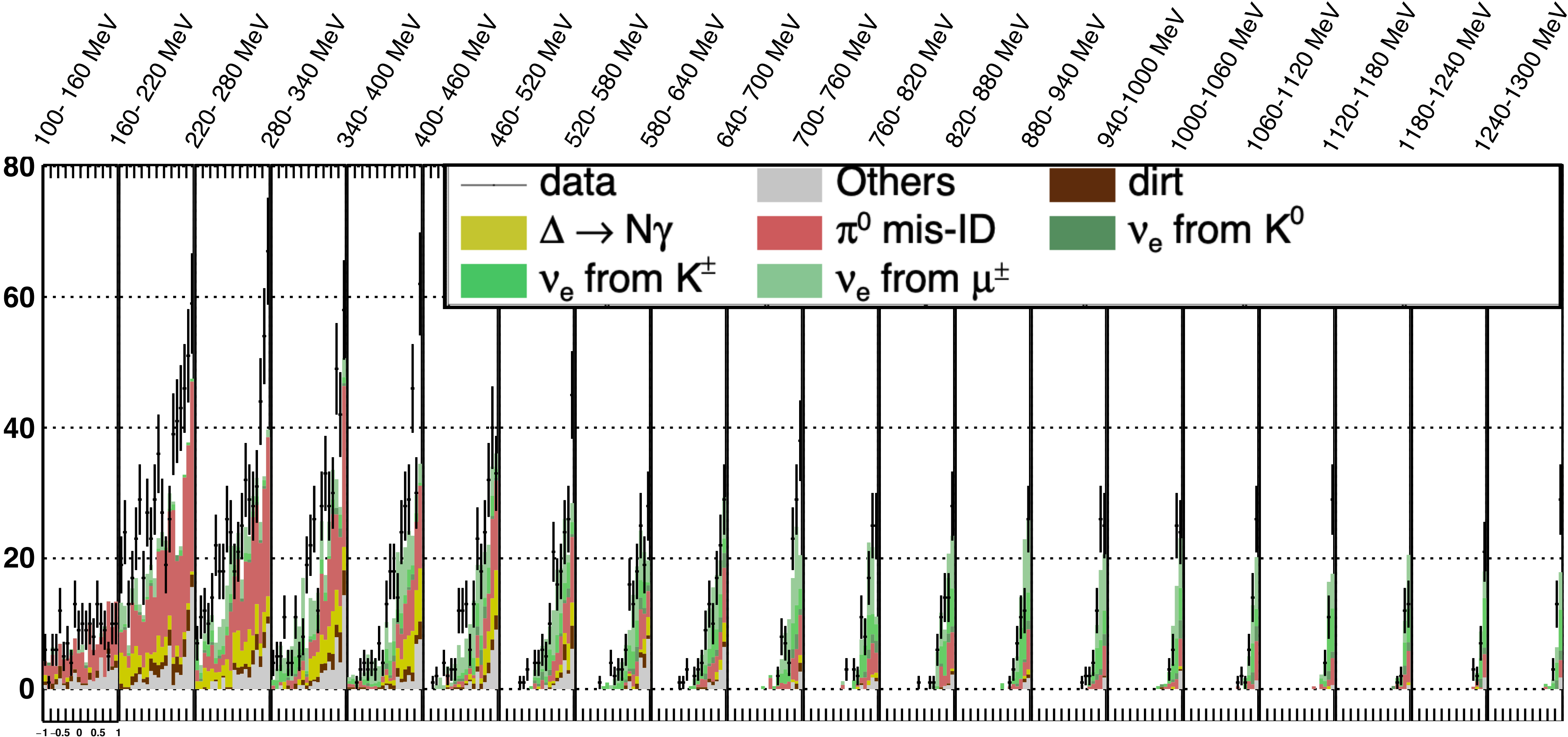
- The grey area corresponds to the allowed 1σ region of the oscillation fit
- The excess is well explained at high E (> 300 MeV) by a simple oscillation model
- Need an additional explanation for the first two bins
- Value in the 150-200 MeV bin is with statistical error only, and unconstrained backgrounds

$\cos(\theta)$ VS E_{vis} Distributions for ν_e samples



- Compare data to backgrounds in the $(E_{\text{vis}}, \cos\theta)$ space
- We believe that these distributions will guide theorists to explain our data
- These data will soon be provided in a data release with systematic errors

cos(θ) VS E_{vis} Distributions for ν_e samples



Data compared to stacked backgrounds
 cos θ for slices of E_{vis} [MeV]

- The excess at low energy occurs across a wide range of cos θ

Summary

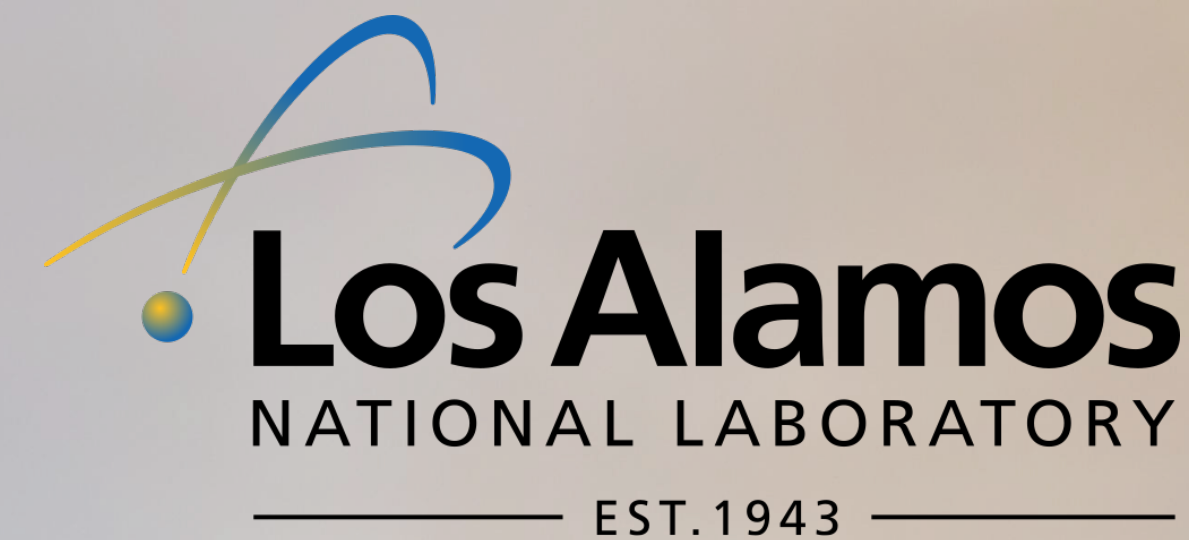
- MiniBooNE presented a full analysis of **17 years of data taking**
- The **event excess has remained stable** in shape and magnitude over the different data releases
- We now have a 4.7σ significance in neutrino mode only, and a **4.8σ significance** in a combined ($\nu + \bar{\nu}$) analysis for our **nominal cut of $R < 500$ cm.**
- Several explanations for the excess are disfavored
 - NC π^0
 - Dirt event
 - Dark matter run ruled out non-neutrino-related beam backgrounds
- In the spirit of responding to requests for more information, we are now in a position to provide additional information on timing, visible energy VS angle of lepton scatter, and radius.
- Further studies under way to better understand the excess, including investigating Meson Exchange Currents, stay tuned!

preprint available at [arxiv:2006.16883](https://arxiv.org/abs/2006.16883)



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Thank you!

