

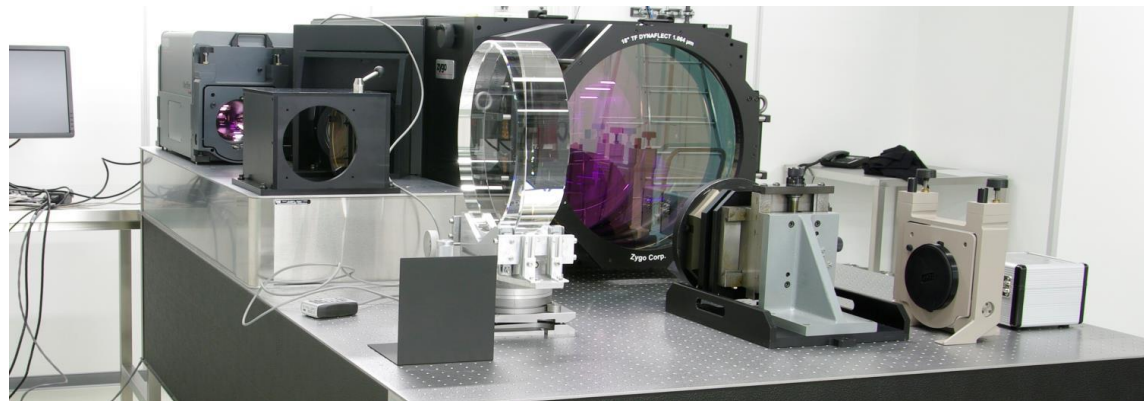


Fabrication des miroirs pour VIRGO

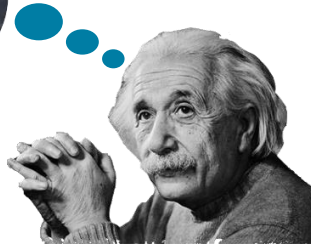
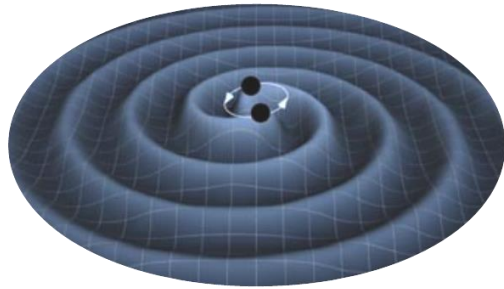
Journées Techniques Détecteurs IN2P3

1 juin 2021

Benoît Sassolas pour le LMA

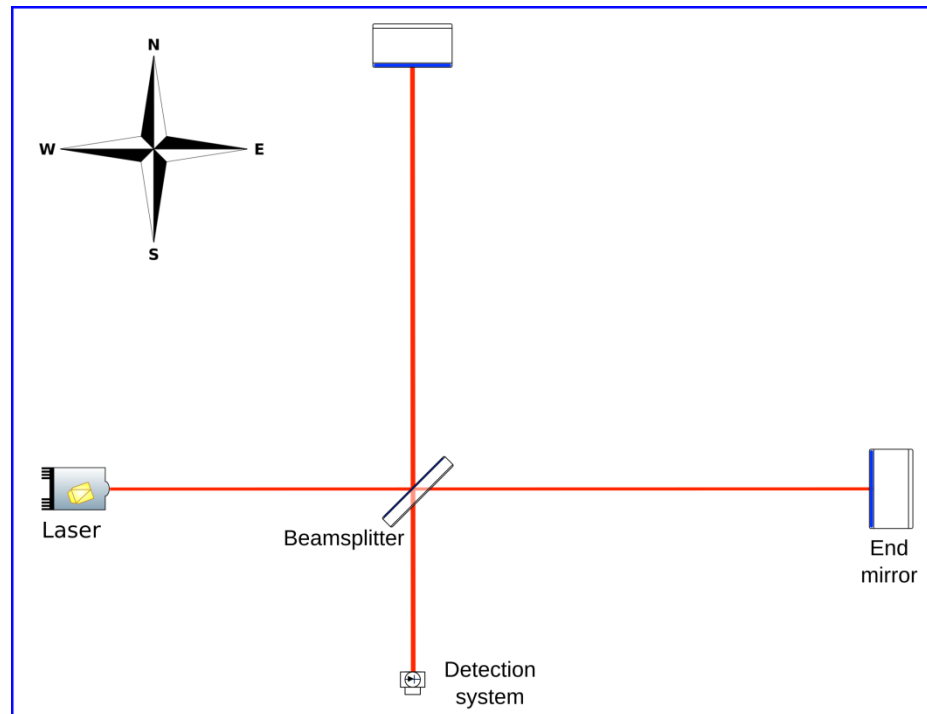


Gravitational waves



- General Relativity (1915)
- Gravitational waves = space-time perturbations
- Came from masses acceleration
- But very tiny signal : $\Delta L/L \sim 10^{-21}$

Example : massive objects merger



Giant detectors

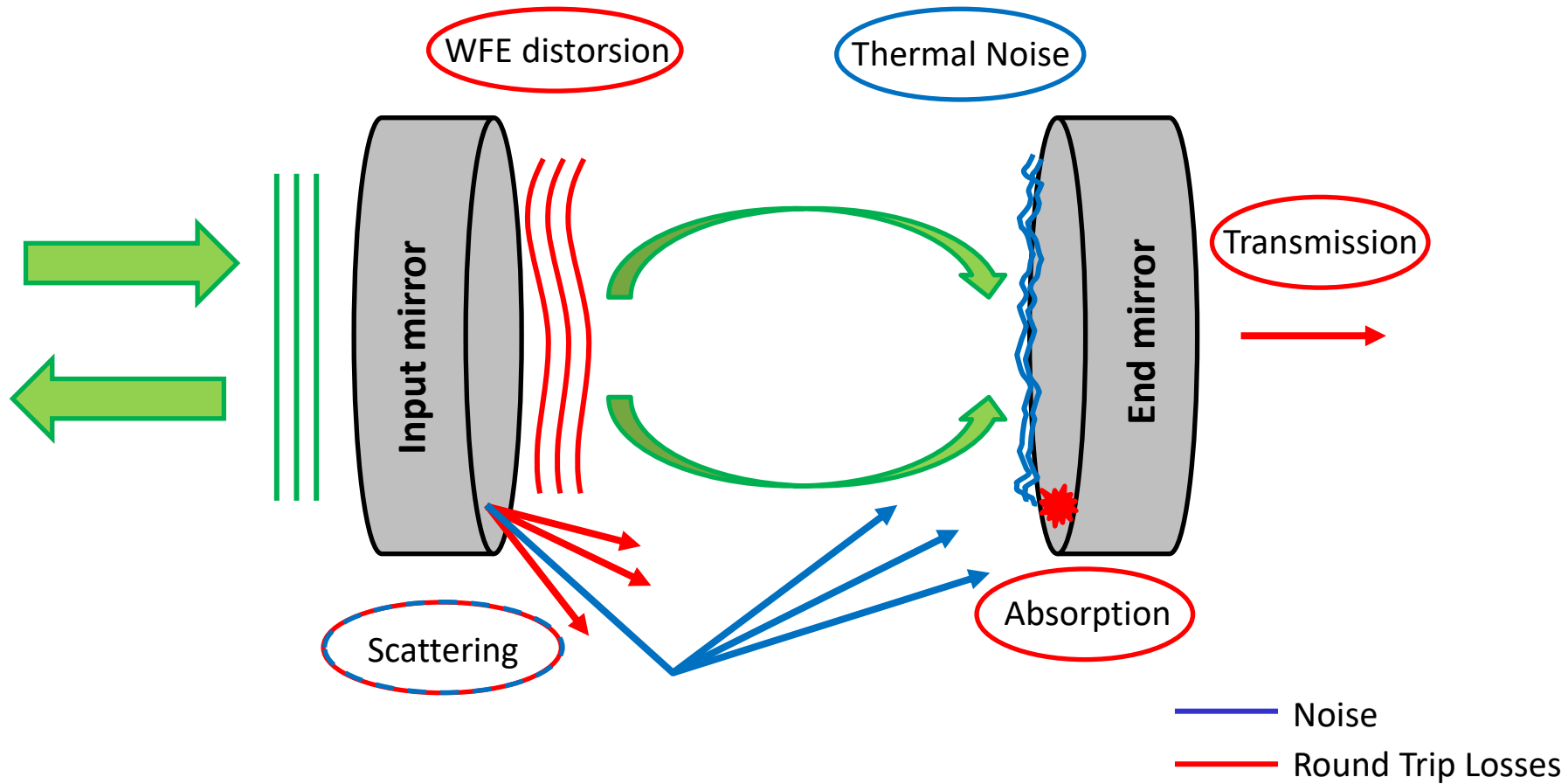


Michelson interferometer :

- 3 km length
- Fabry-Perot cavities
- Vacuum $<10^{-9}$ mbar
- 700 kW of light inside the cavity
- 40kg mirrors (\varnothing 35 cm; 20 cm thick)

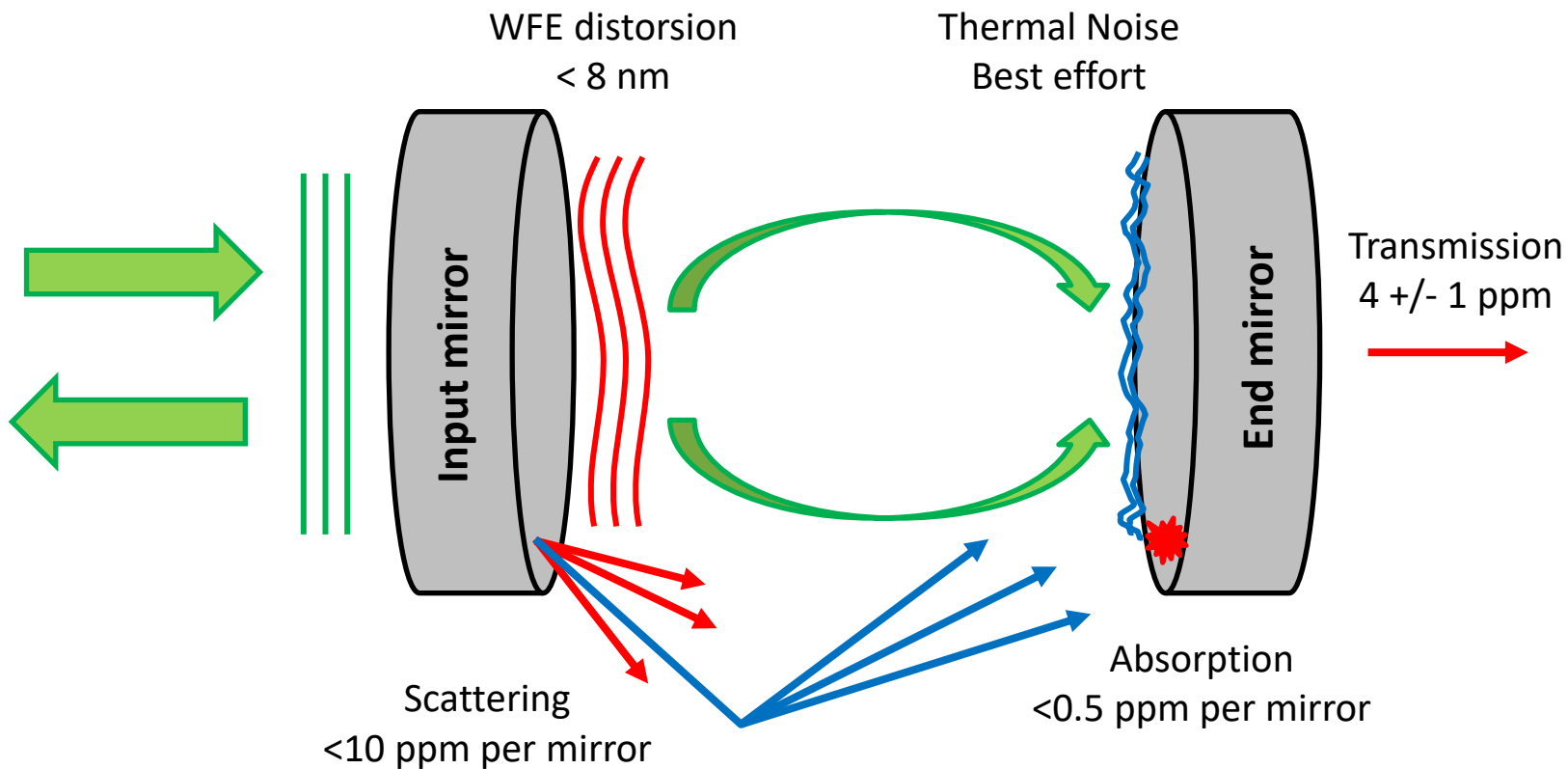
Major effects in a cavity

Goal : reflect the light properly as much as possible but ...

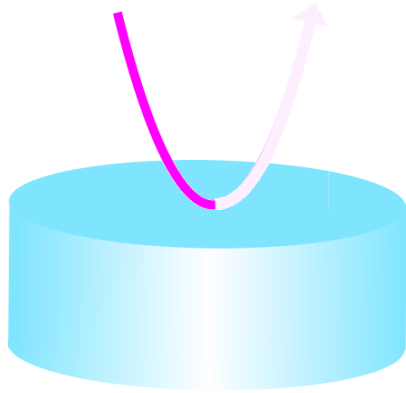


In order to limit these effects,
the mirror must be as performant as possible

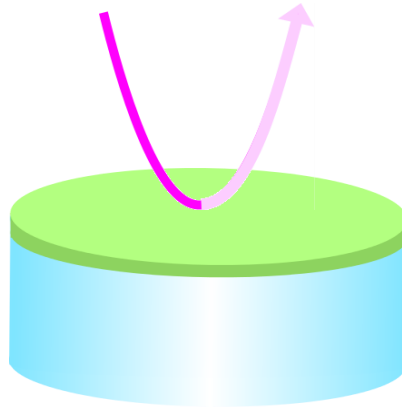
Requirements



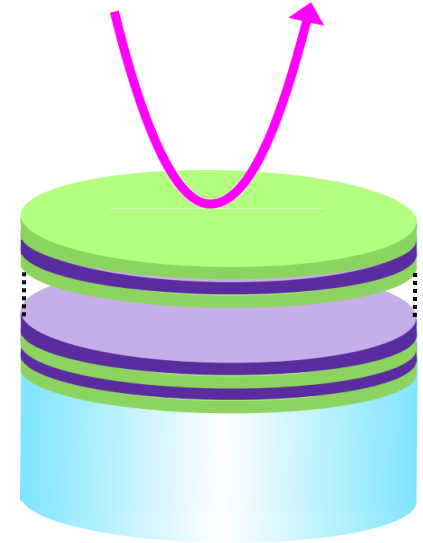
How does a mirror work ?



1 surface, $R \sim 4\%$



2 surfaces, $R \sim 20\%$



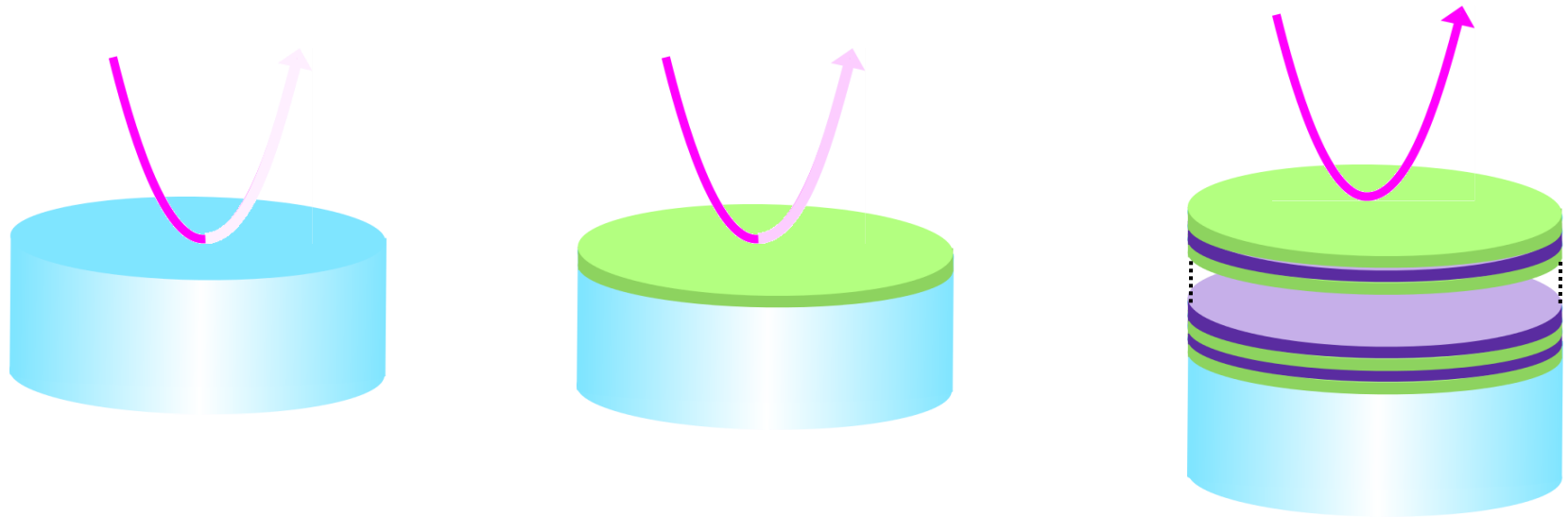
40 surfaces, $R \sim 100\%$

$$R_{max} \cong 1 - 4n_S \left(\frac{n_L}{n_H} \right)^{2p}$$

Stack of thin films :

- $2p$ layers
- High index material : titania-doped tantala ($\text{Ti:Ta}_2\text{O}_5$)
- Low index material : silica (SiO_2)
- Constructive interferences in reflection $\varphi_1 = \varphi_2 = \varphi_3 = \pi [2\pi]$
- Thicknesses of the layers $n_H e_H = n_B e_B = \lambda/4$; i.e. Quarter Wave Layer (QWL)

How does a mirror work ?



Developped at LMA for 10 years.

The best optical and mechanical material so far !

A huge topic in itself !

titania-doped tantala ($\text{Ti:Ta}_2\text{O}_5$)



Harry et al 2006 Class. Quantum Grav. 24 405

Flaminio et al 2010 Class. Quantum Grav. 27 084030

Amato et al 2019 J. Phys. Mater. 2 035004

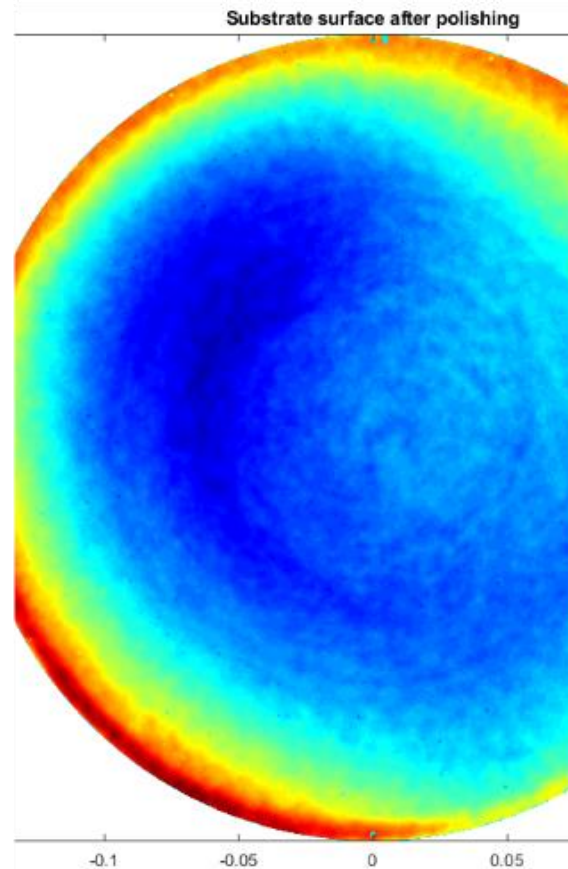
Granata et al 2020 Class. Quantum Grav. 37 095004 ...

High performance means high quality

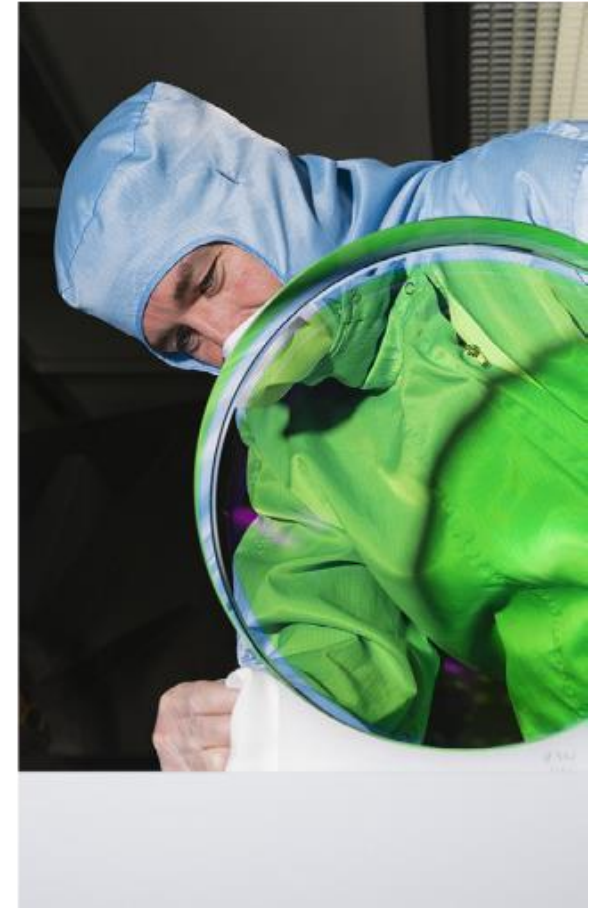
Substrate



Polishing



Coating



All 3 ingredients need to be exceptional !

The most perfect piece of glass !

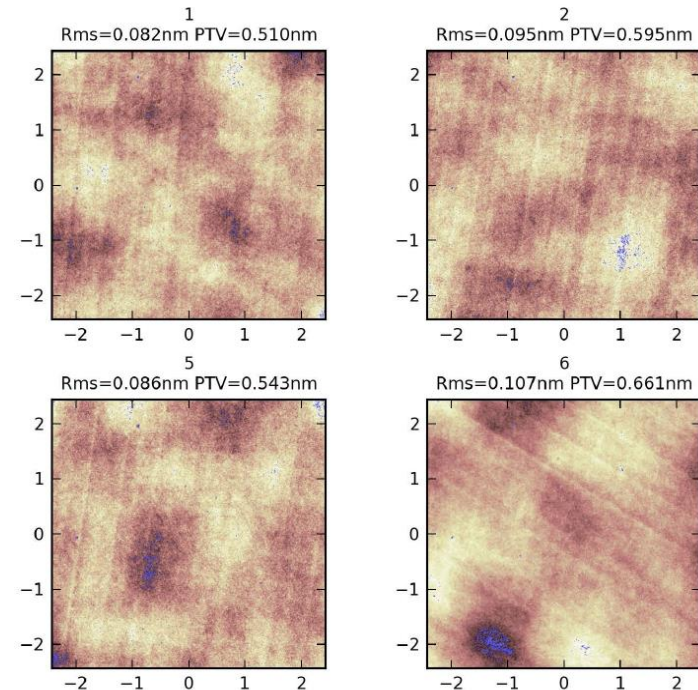
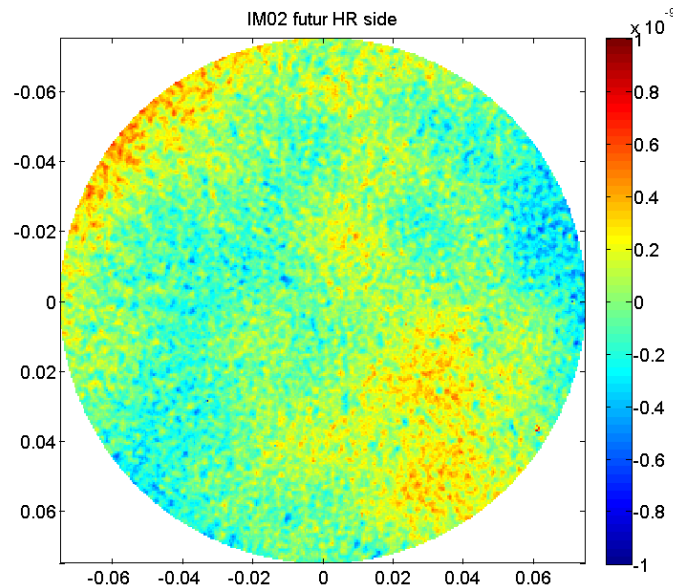


Material

- $\varnothing 35$ cm ; 20 cm thick
- High purity fused silica
- No bubble or inclusion

Polishing

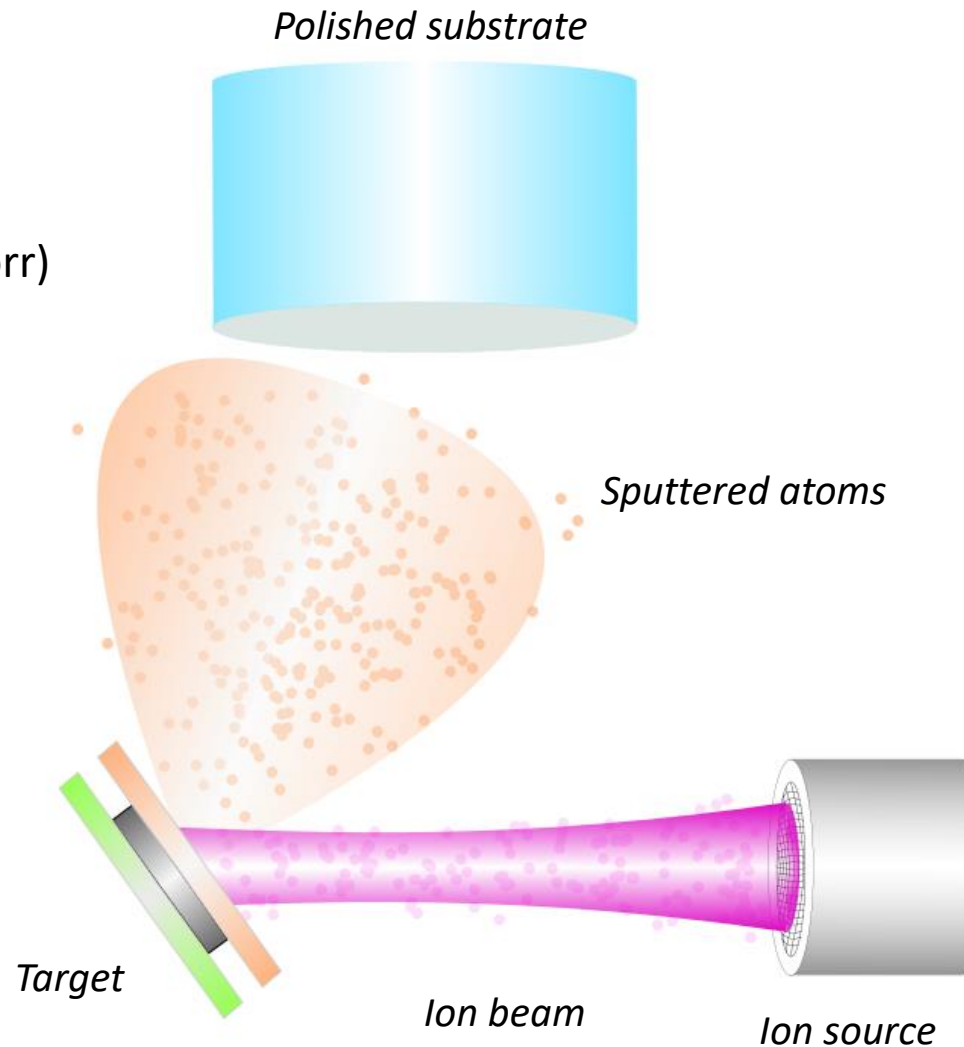
- RoC = 1.5km +/- 2m
- Flatness $PV = \lambda/300$
- Roughness < 0.1 nm rms



How to make the « HR coating » ?

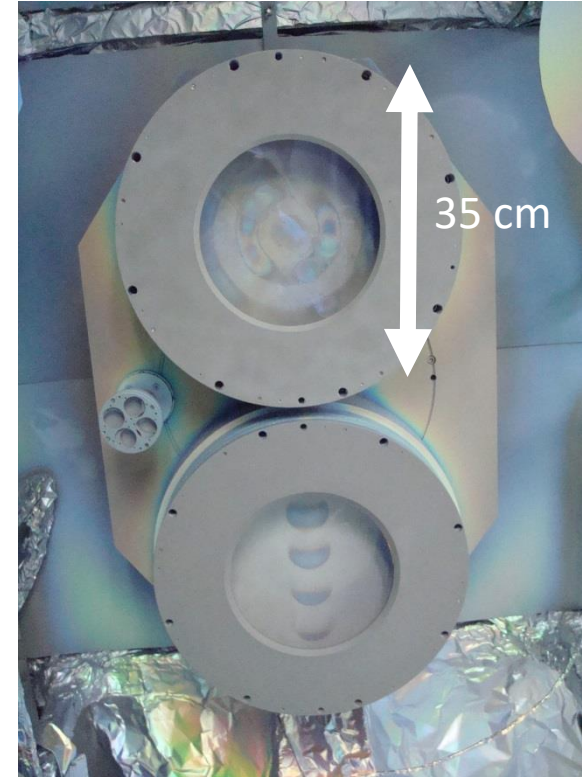
Ion Beam Sputtering (IBS)

- High vacuum
(base pressure $\sim 10^{-7}$ torr)
- Energetic deposition
- High density
- Low scattering
- Low absorption



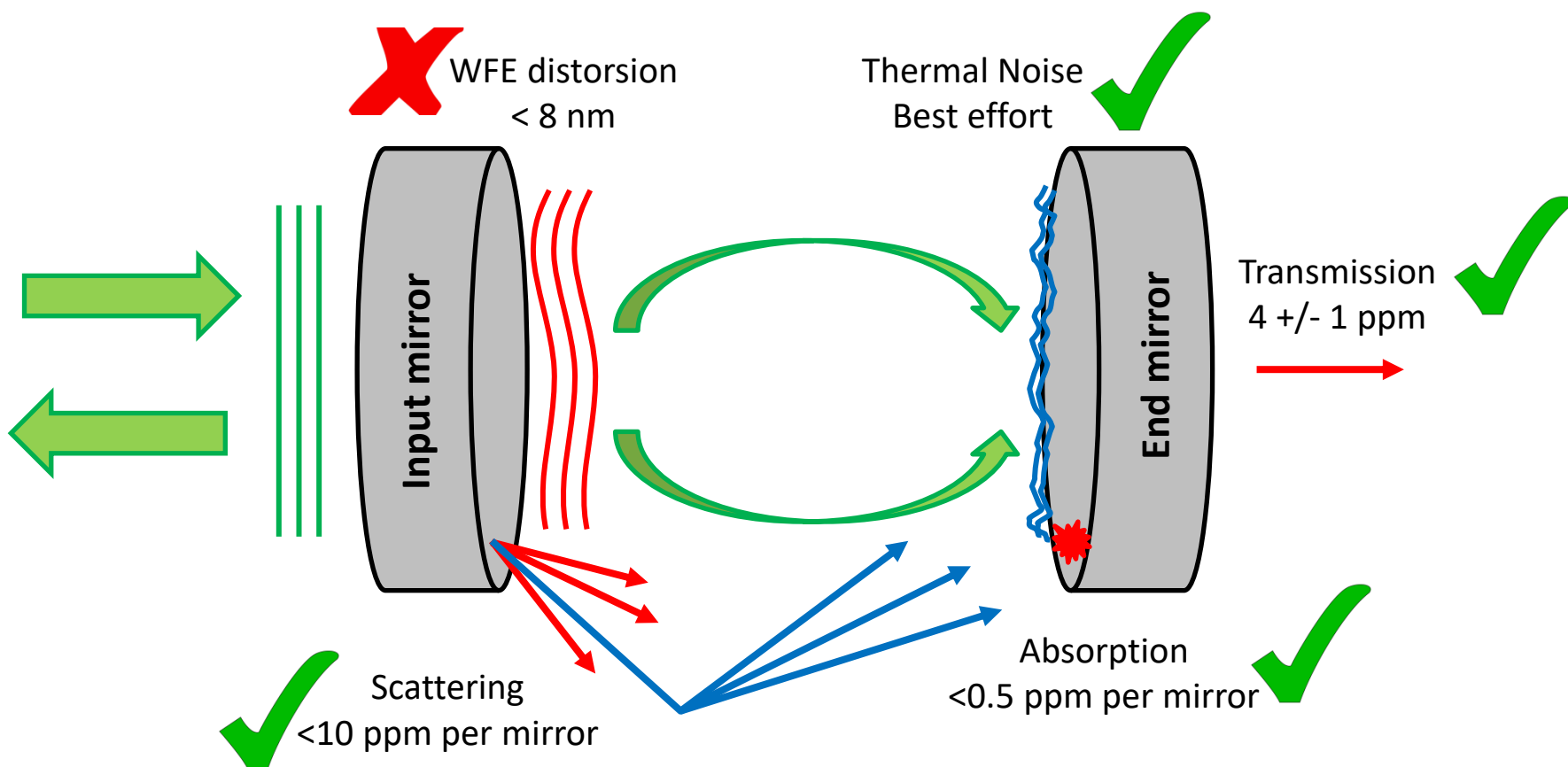
High optical quality

An home made machine



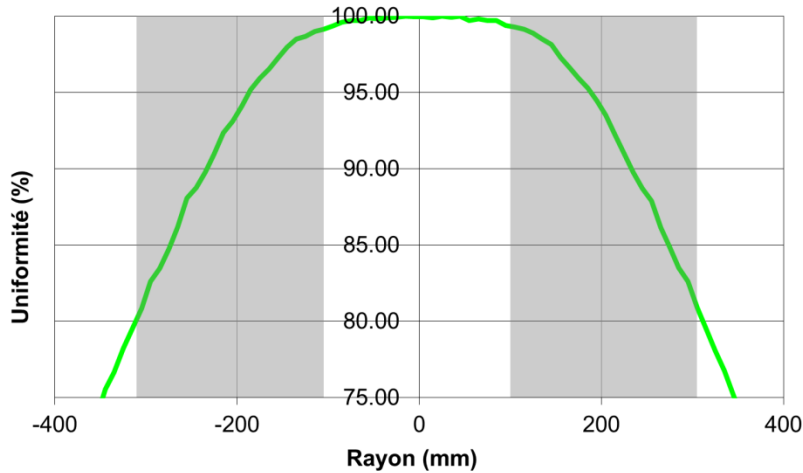
- 10 m³ vacuum chamber
- Able to coat up to \varnothing 1m or 2 Virgo mirrors at a time
- Install in a ISO3 cleanroom

Requirements



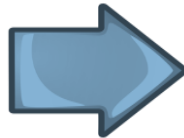
IBS technique allows to produce low-loss coating.

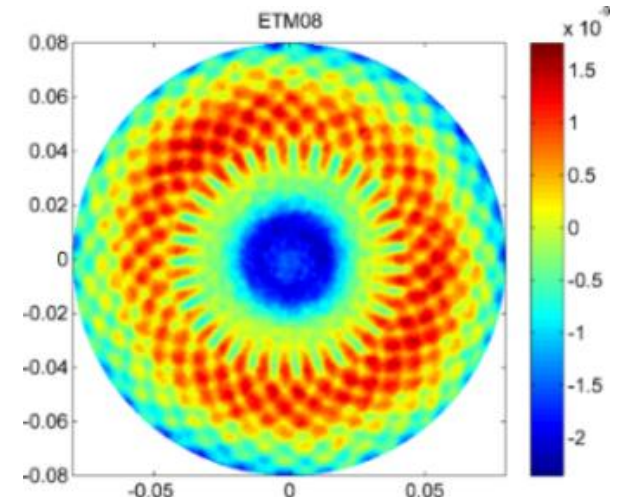
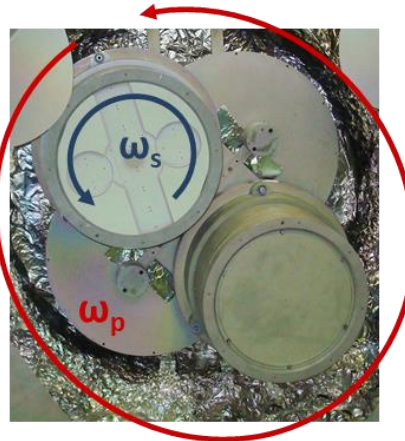
WFE improvement



Quite large thickness variation

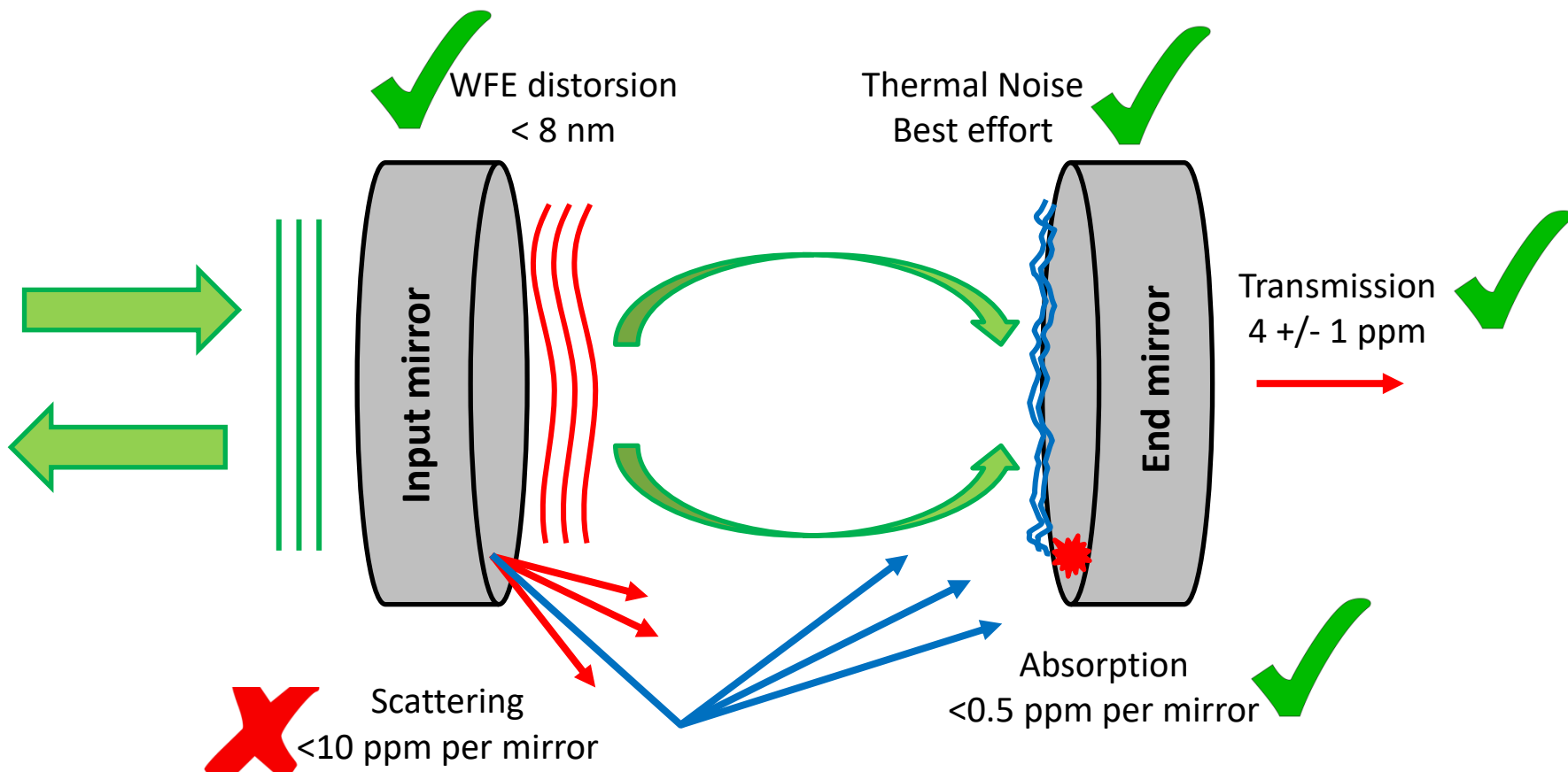
- Strong impact on the optical properties
- Large WFE

20 %  0.075%



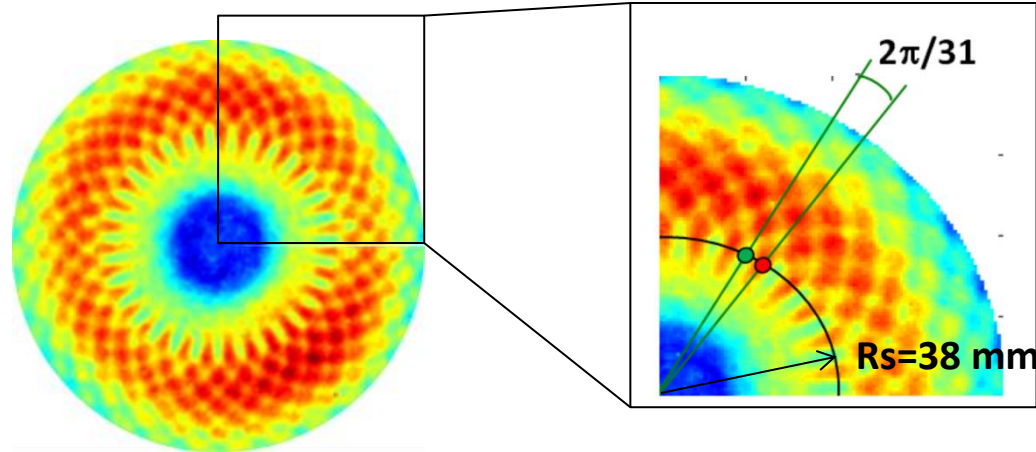
WFE in reflection
over $\varnothing 160\text{mm}$ (nm)

Requirements

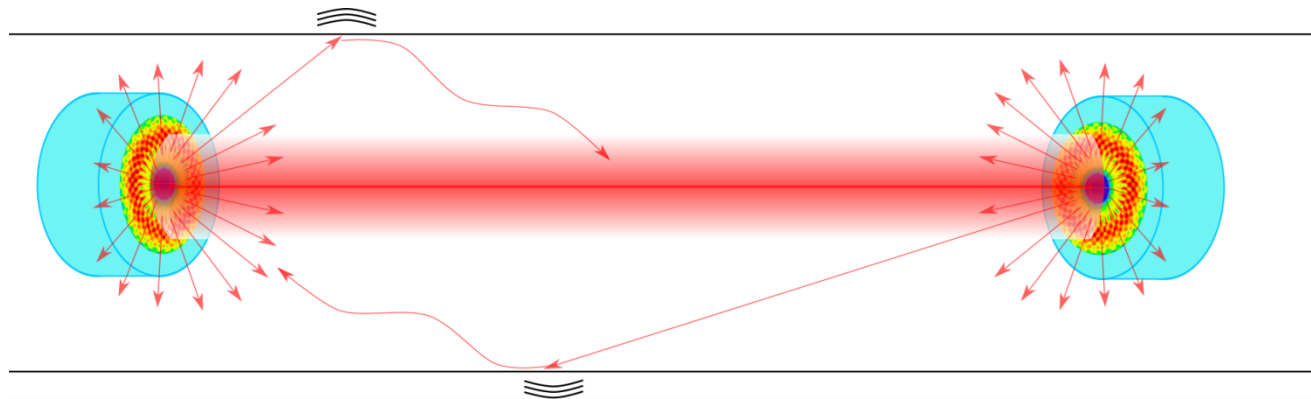


What's happened ?

The spiral pattern



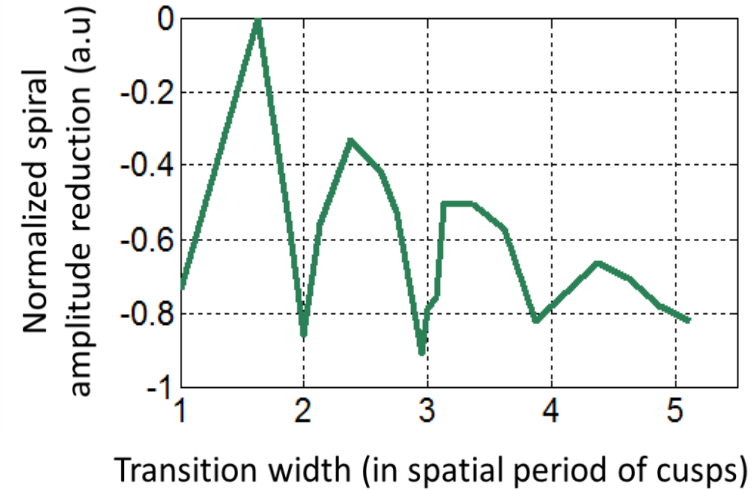
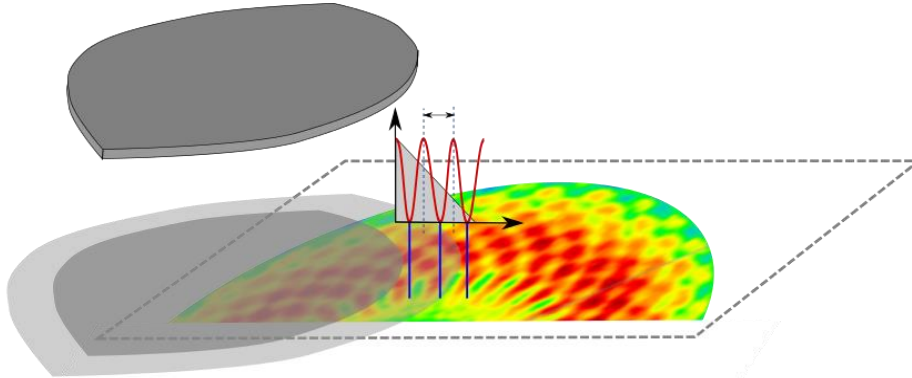
- Maximal spiral amplitude on radius $R_s = 38$ mm and angular period of $2\pi/31$
- Generated by the combination of the planetary motion and the shadowing mask
- Small defect : only 1.5 nm for a 6 μm coating total thickness



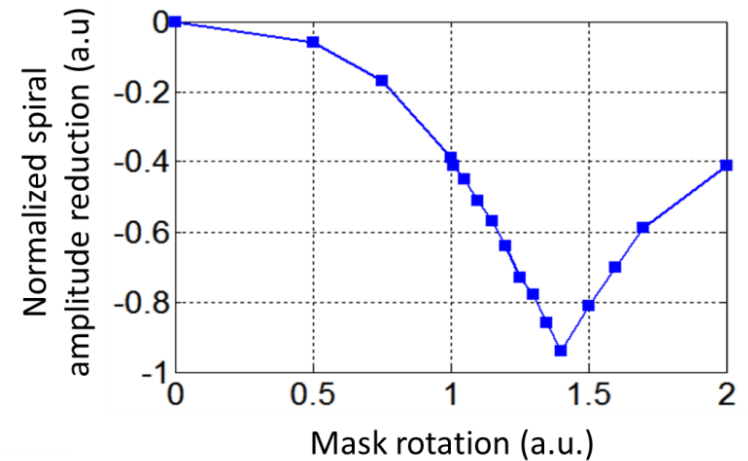
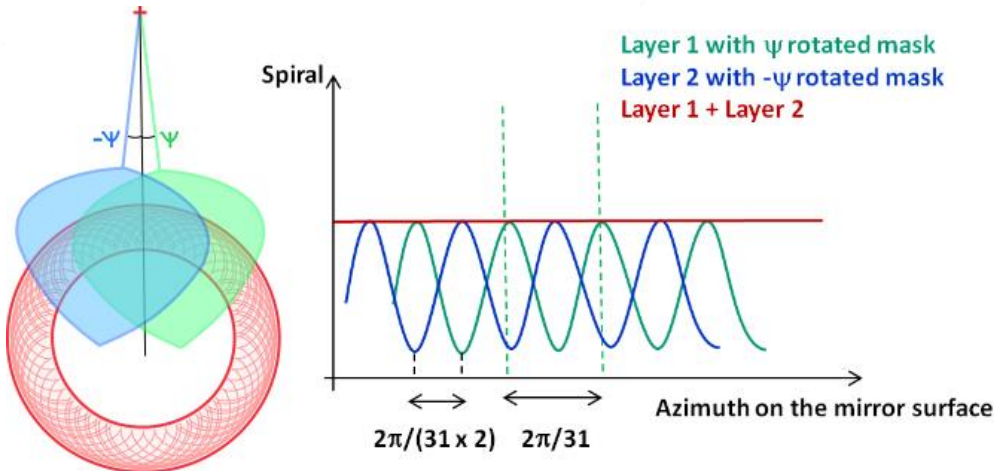
BUT this tiny defect limits the sensitivity of the GW detectors

Spiral mitigation techniques

The « blurring » technique

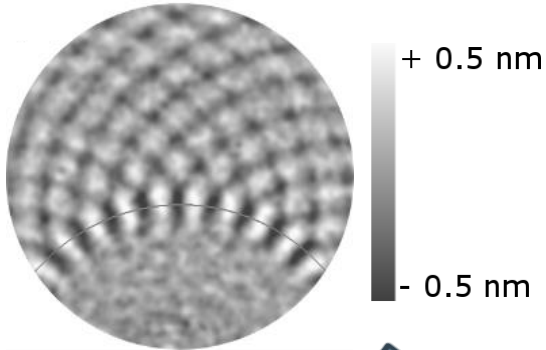


The « interference » technique



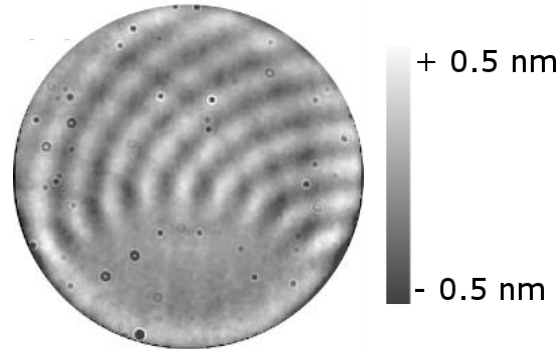
Final results

Before mitigation



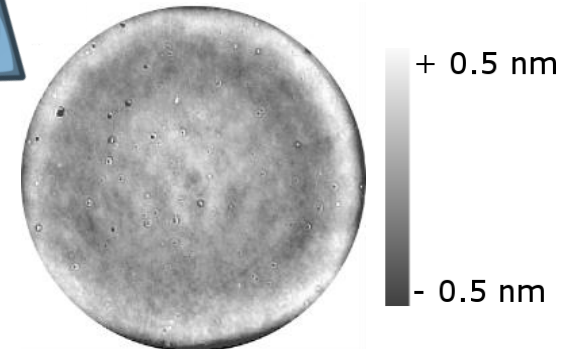
x 1/2

With the blurring technique

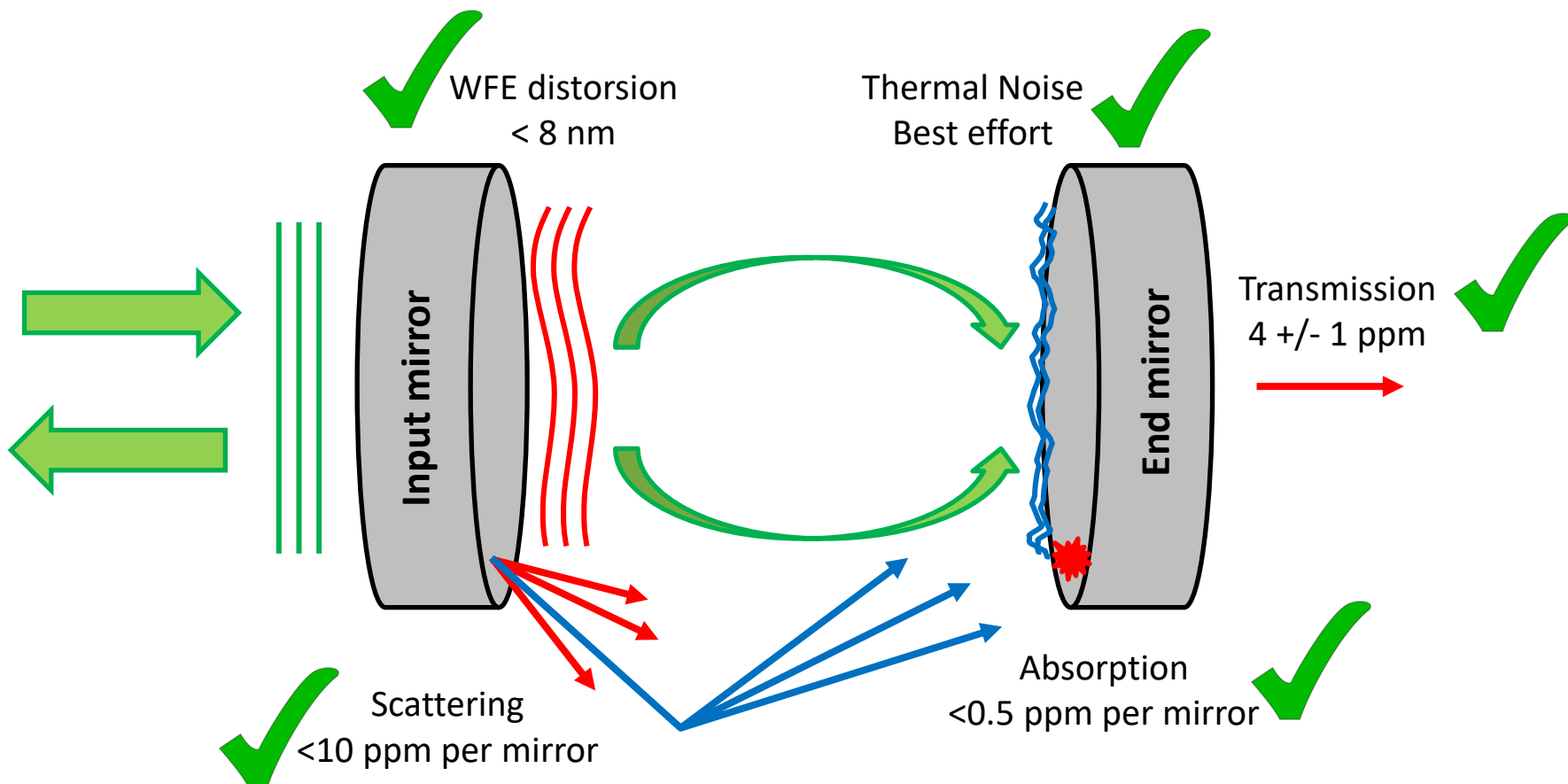


x 1/3

With the blurring and the interference technique



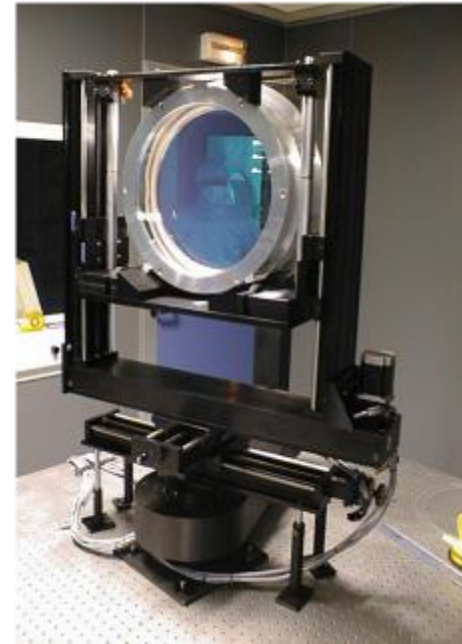
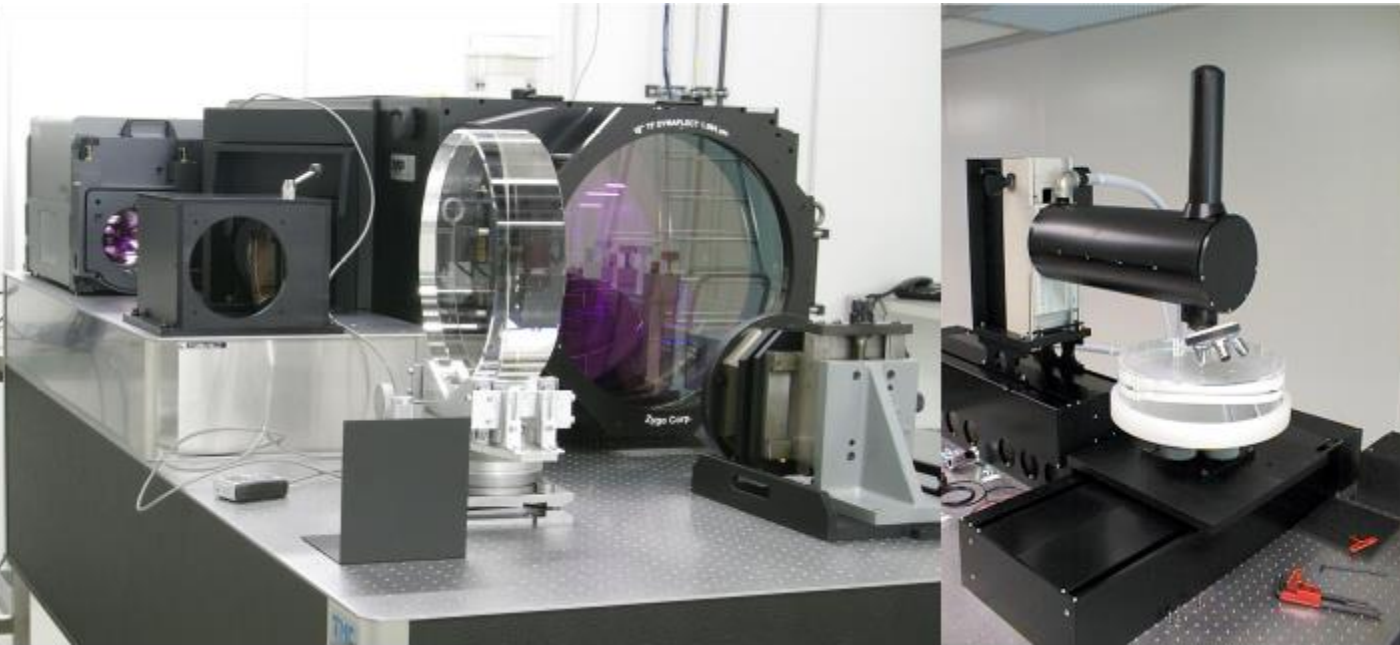
Requirements



Complete optical metrology

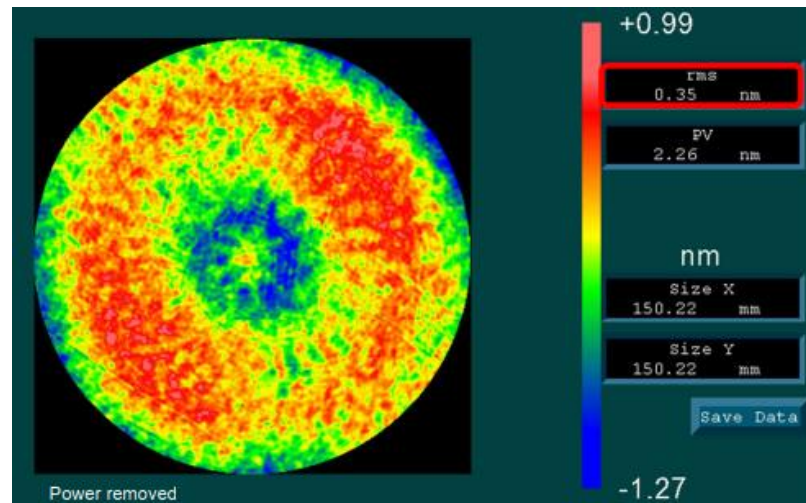
Essential step in order to check the actual performances of the optics :

- Transmission
- Optical absorption
- Waveront error (WFE)
- Microroughness
- Scattering

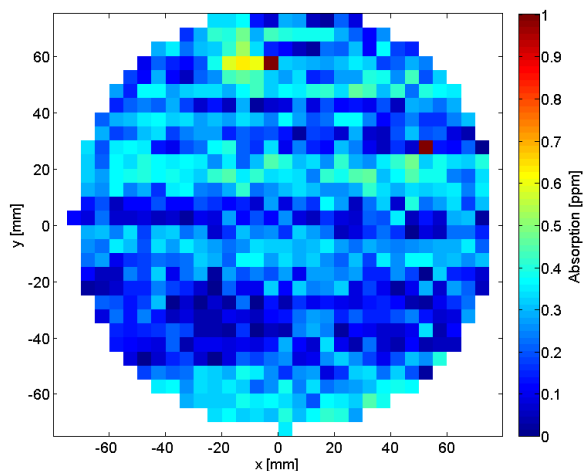


Final performances @ 1064 nm

- Flatness : 0.35 nm RMS (w/o curvature)
- Transmission matching : difference <0.002%
- End mirror transmission 4 ppm

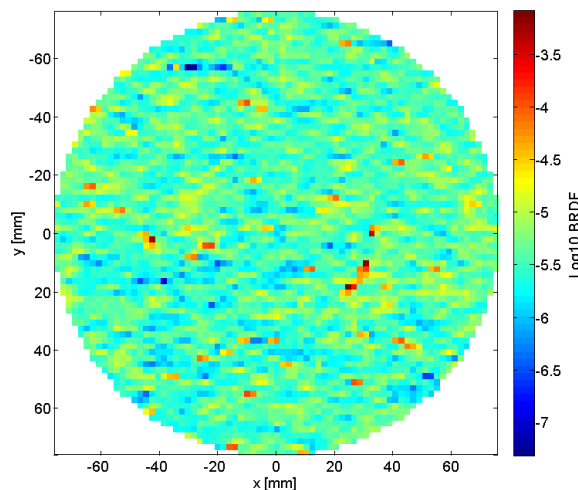


Absorption



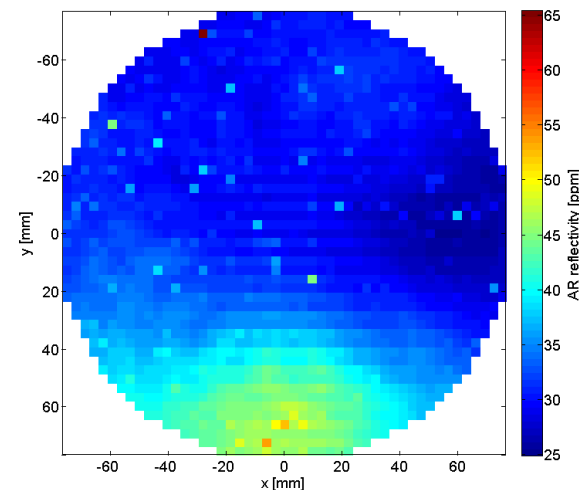
0.2 ppm

Scattering



4 ppm

AR reflectivity

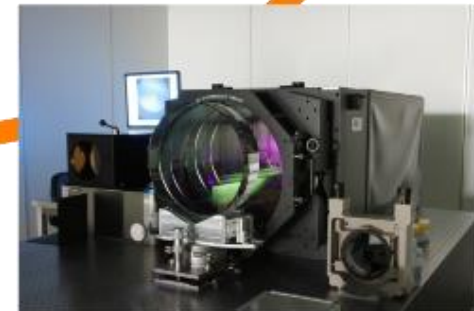


30 ppm

Some numbers



- 10 years R&D on the materials
- 4 years R&D on the uniformity
- 24 substrates of 40 kg coated
- 480 h of coating (including 15 nights)
- 0.1 mm total thickness coating
- 240 days of metrology



Thank you

Interferometric measurements

