

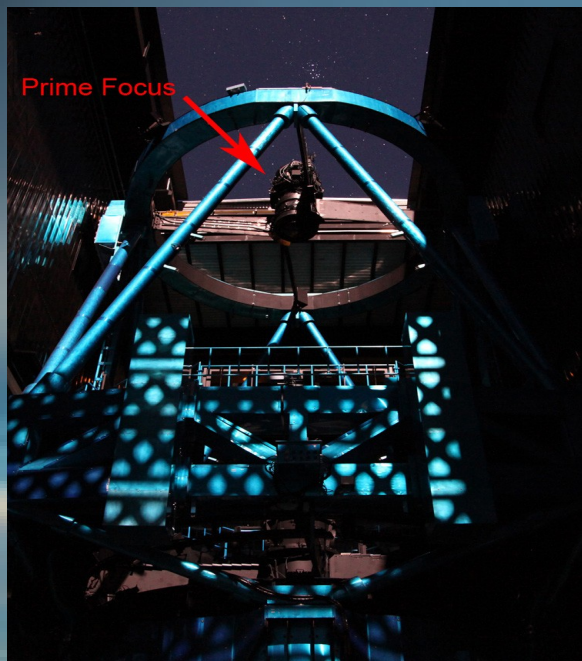


Transient detection in the HSC data

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Subaru telescope

Hyper Suprime-Cam



mauna kea

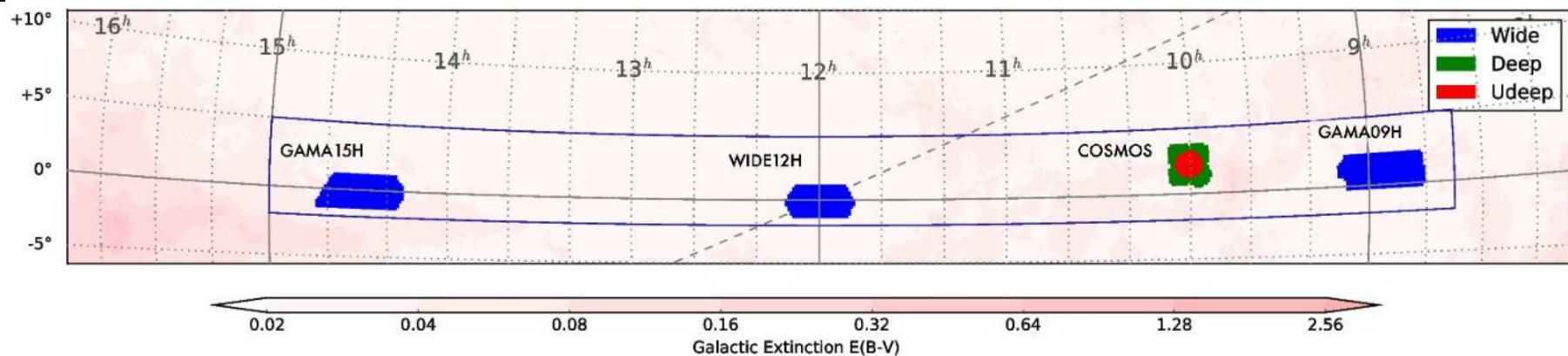
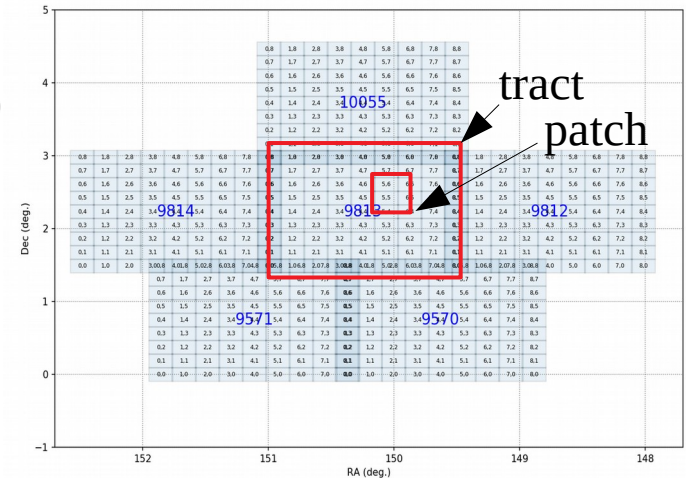
HSC characteristics



Fast, Wide & Deep

300 nights over 5-6 years (started in March 2014)

- grizY+4 narrow bands
- Wide FoV: $\sim 1.5^\circ$ diameter
- Excellent Image quality: $\sim 0.6''$ seeing
- 3 layers:
 - Wide (1400deg^2 , $r\sim 26$)
 - Deep (27deg^2 , $r\sim 27$)
 - Ultradeep ($3\text{-}3.5\text{deg}^2$, $r\sim 27.7$)

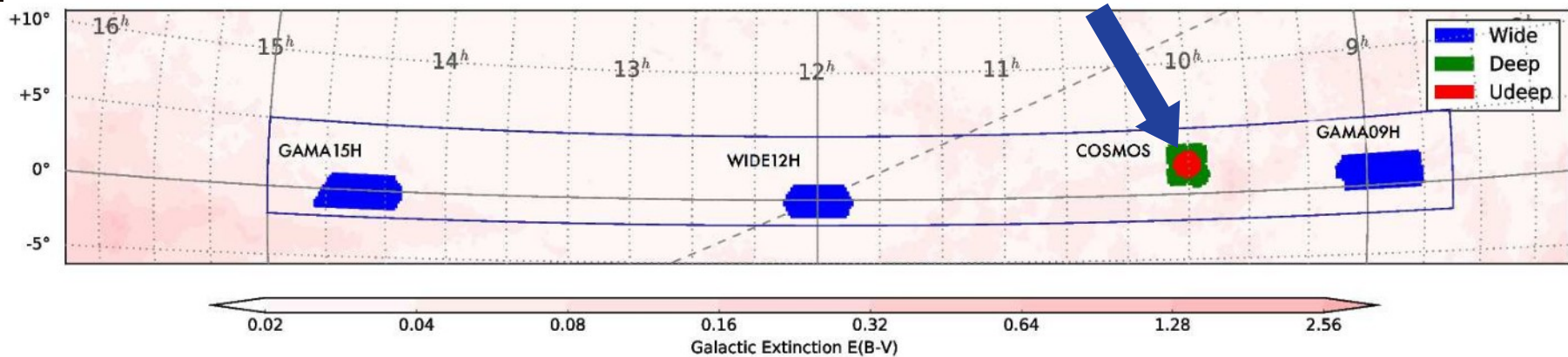
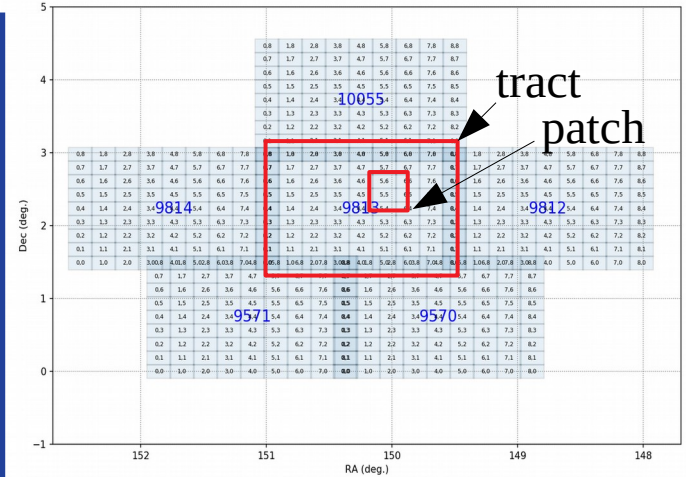


HSC characteristics



COSMOS: Cosmic Evolution Survey

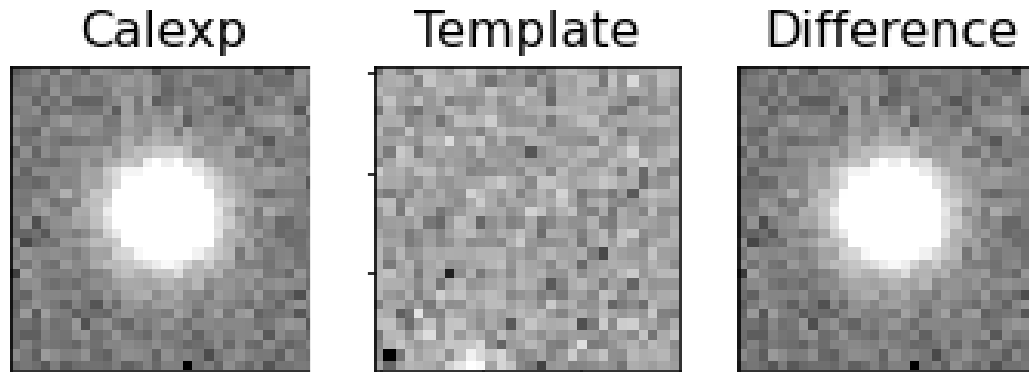
- Astronomical survey designed to probe the formation and evolution of galaxies as a function of redshift
- It covers a 2 deg² equatorial field
- ~ over 2 million galaxies are detected



Ongoing work

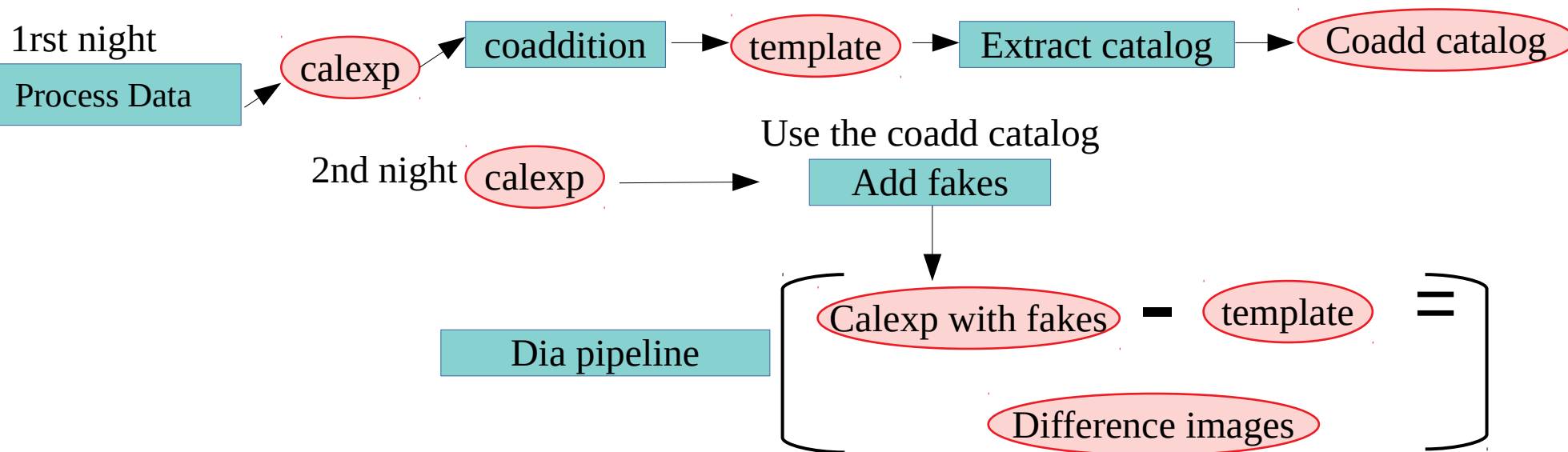
Goal: evaluate the performance of the difference imaging pipeline

- Use the fake injection technique:
 - Inject fakes onto science images (in calexps)
 - Subtract images from a templates of stacked exposures (coadd)
 - Check whether these fakes are detected from the difference image: detection efficiency



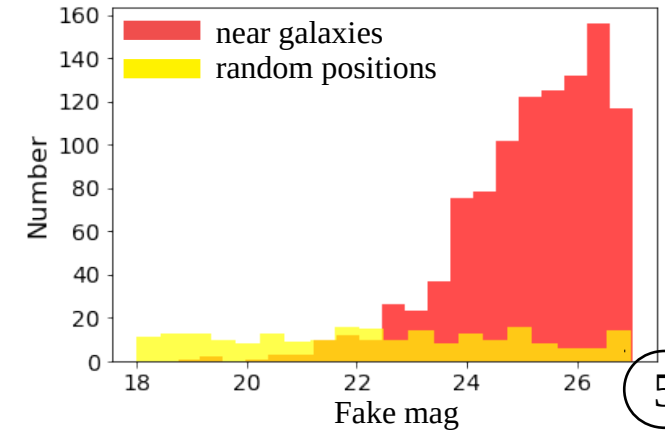
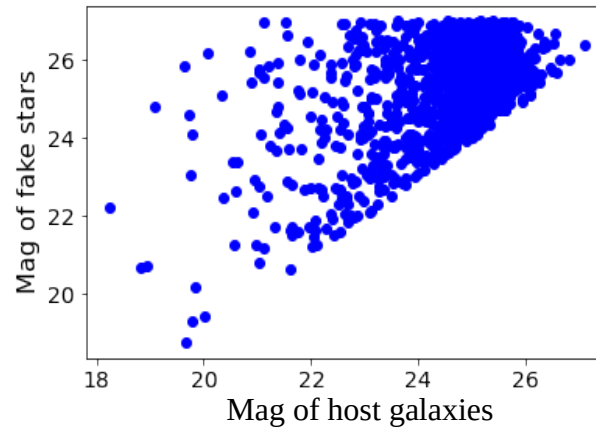
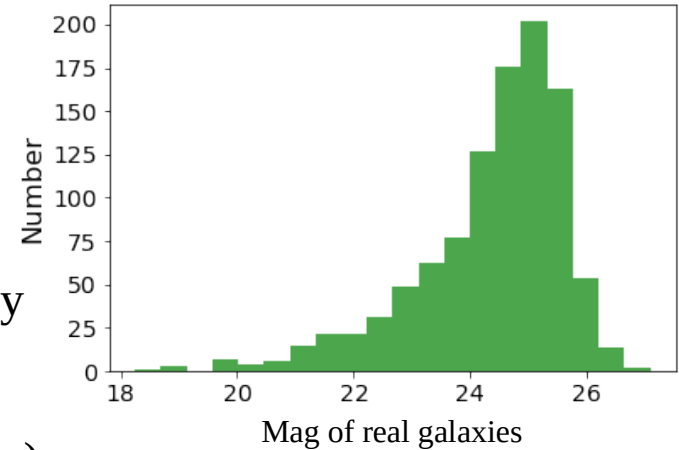
Used data & pipeline

- from 2 nights, PDR1 UDEEP COSMOS field
 - Filter: Z filter
 - Process exposures for 2 nights (2014-03-28, 2015-01-16)
 - Coaddition over 1 night (2015-01-16) → produce template
 - Add fakes on 3 visits (night 2014-03-28, visits: 1166, 1168, 1170)



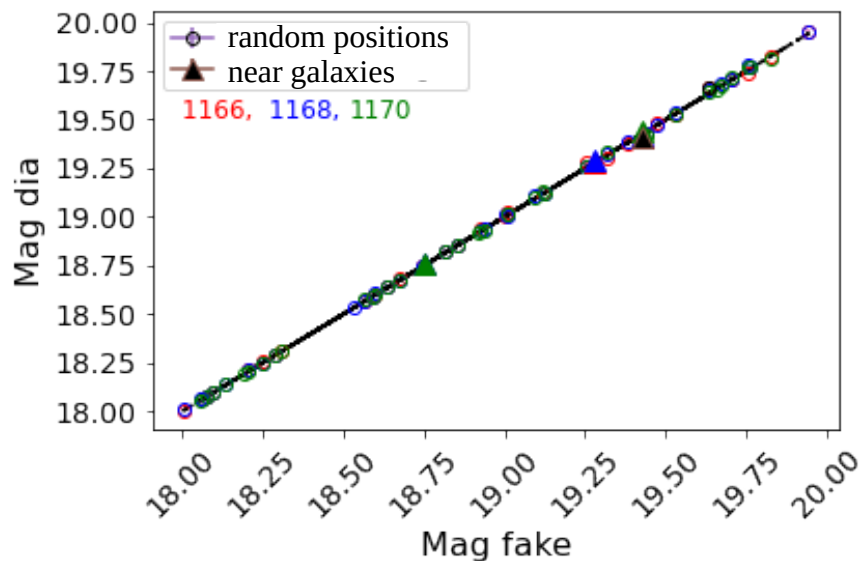
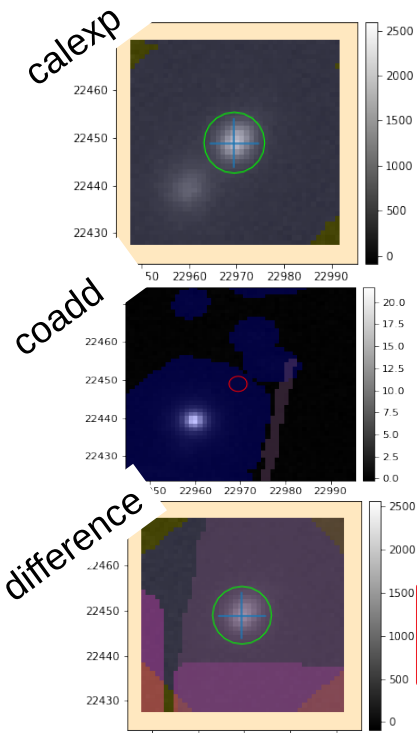
Fake injection

- Among the coadd catalog of patch 5,5, tract 9813
 - ➔ Select galaxies (without deblending)
 - ➔ Excluded flagged sources ('base_PixelFlags_flag', 'base_PsfFlux_flag')
- Generate fakes associated to a galaxy (on 10% of randomly chosen galaxies)
- Gaussian distribution to choose random position according to galaxy radius
- Flat distribution to choose mag ([galaxy mag -1, 27])
- Generate 2% fakes with random positions (without galaxy association)
- Check the detection within magnitude intervals

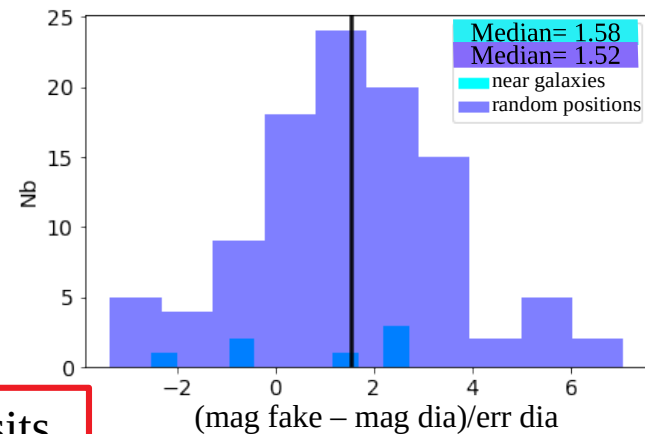
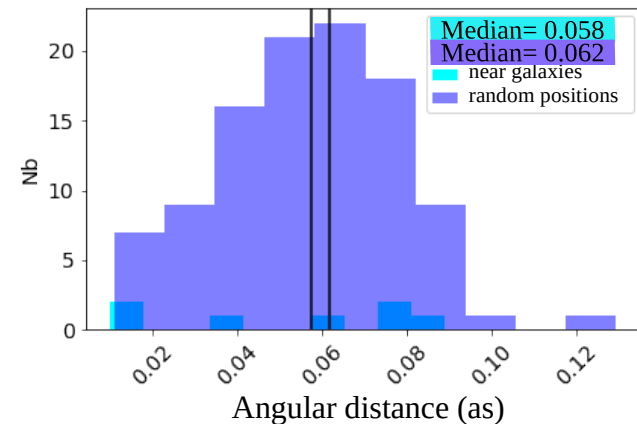


Bright fakes [18 - 20]

- Input 52 fakes with $18 < \text{mag} < 20$
- 46 among 52 are retrieved with angular distance < 0.12 as
- Only 3 fakes are associated with galaxies (and brighter than hosts)

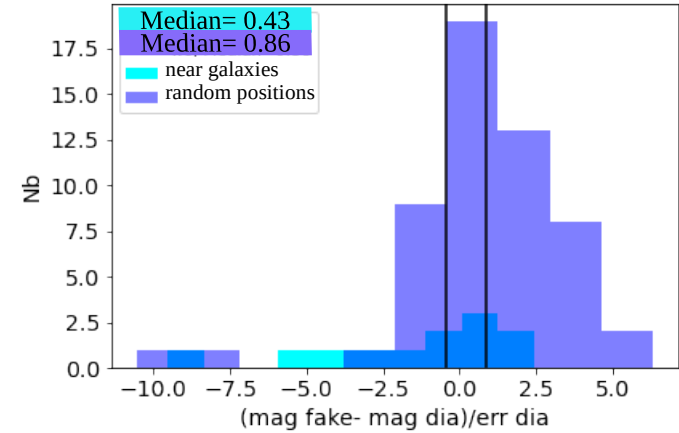
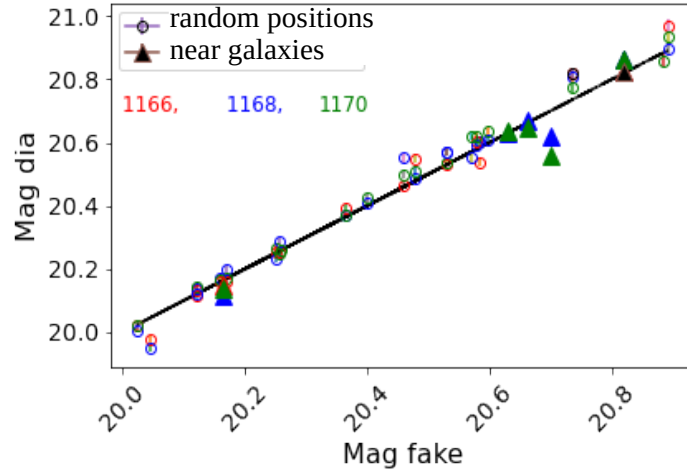


Well retrieved magnitudes for all fakes in the 3 visits

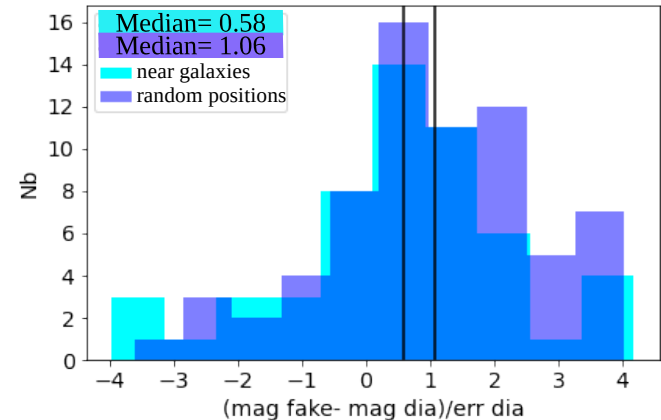
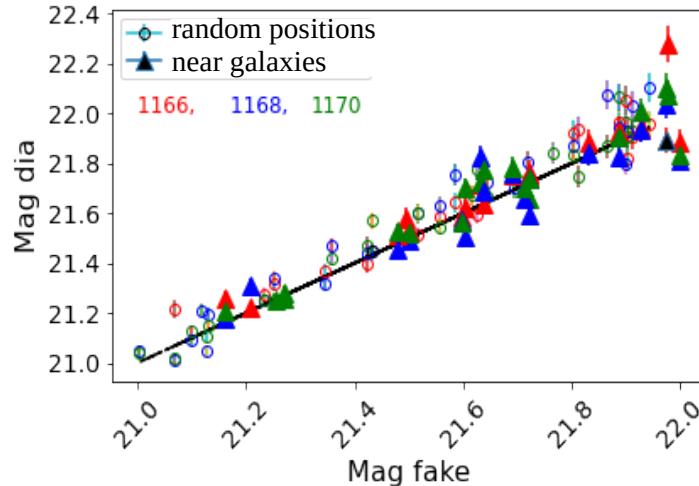


Bright fakes [20 - 22]

- Mag [20 - 21]
- 26/27 retrieved with angular distance < 0.12 as

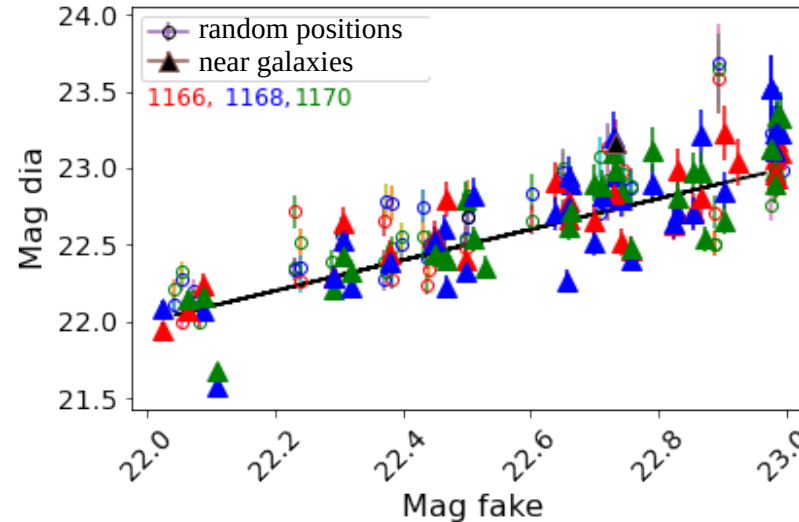
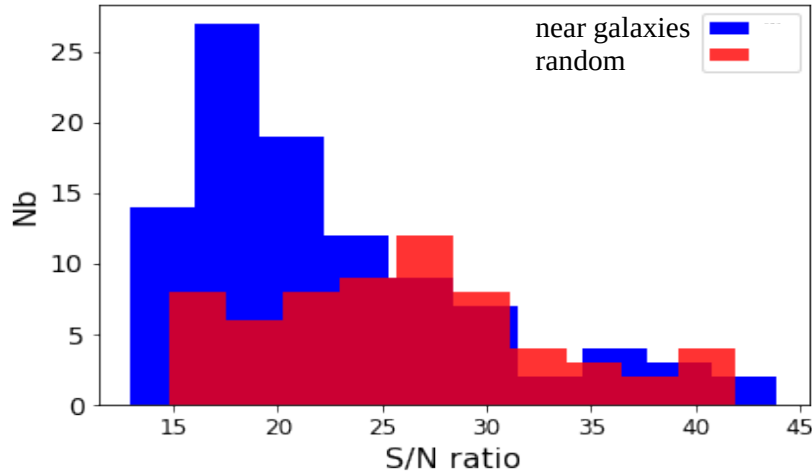
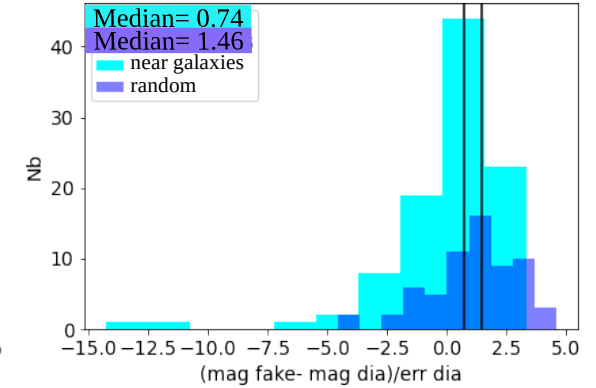
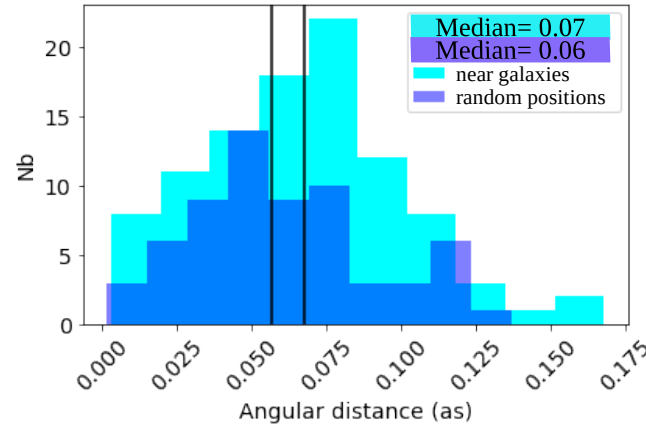


- Mag [21 - 22]
- 53/53 retrieved with angular distance < 0.12 as



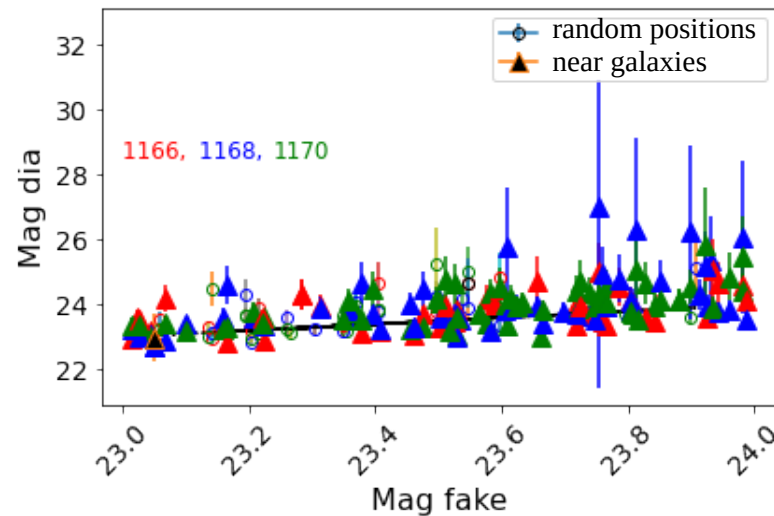
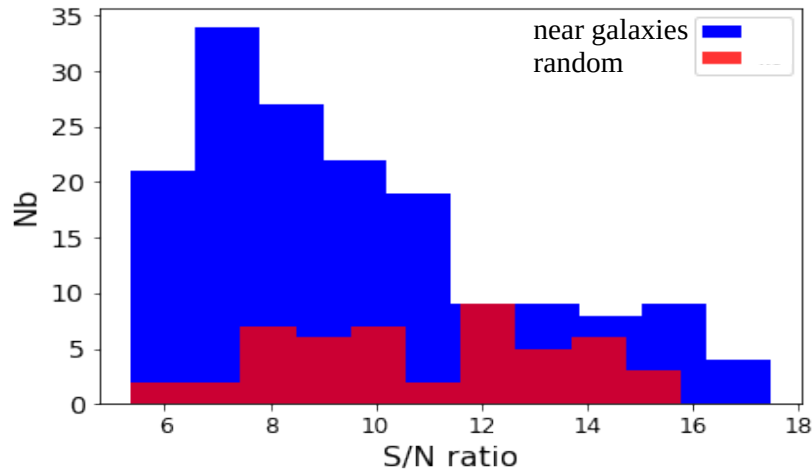
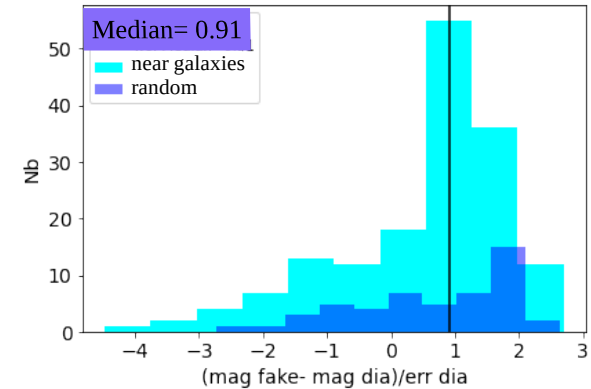
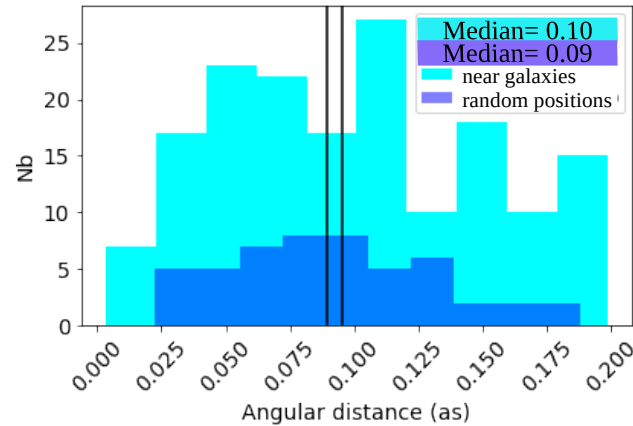
Fakes [22 - 23]

- S/N: PSF flux/PSF flux error
- 67/74 were retrieved
angular distance < 0.18 as



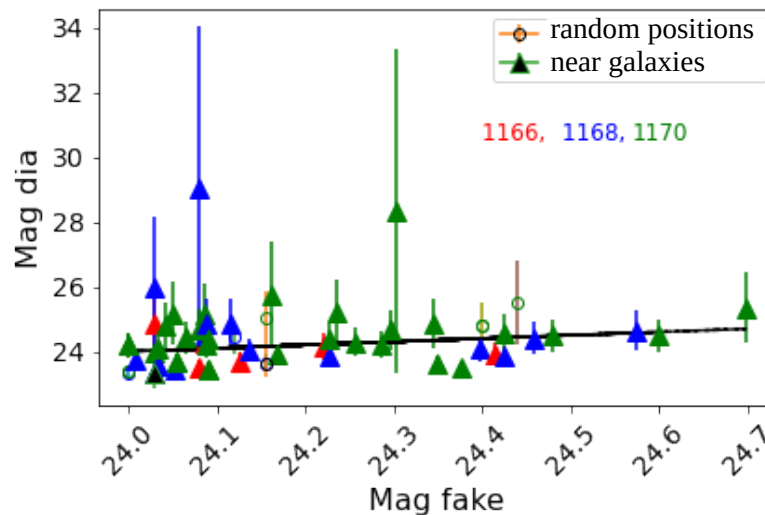
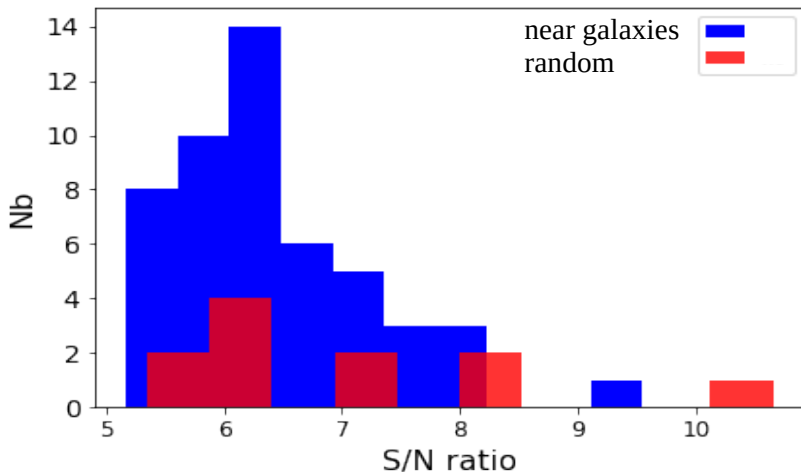
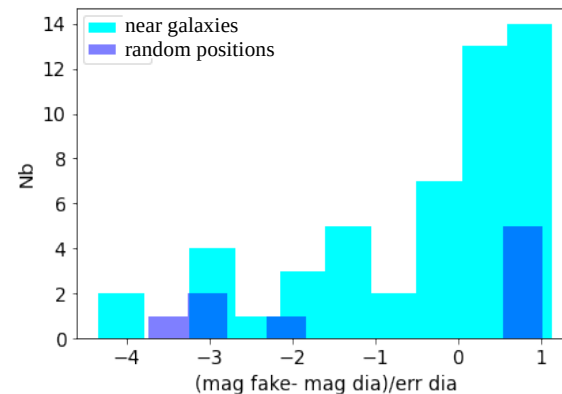
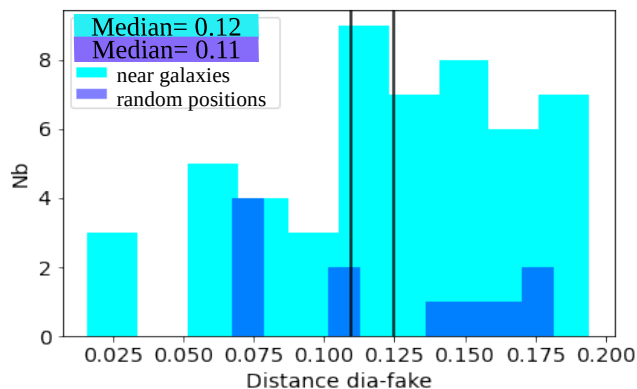
Fakes [23 - 24]

- 102/121 were retrieved angular distance < 0.2 as
- Large magnitude errors for fakes with $\text{mag} > 23.5$

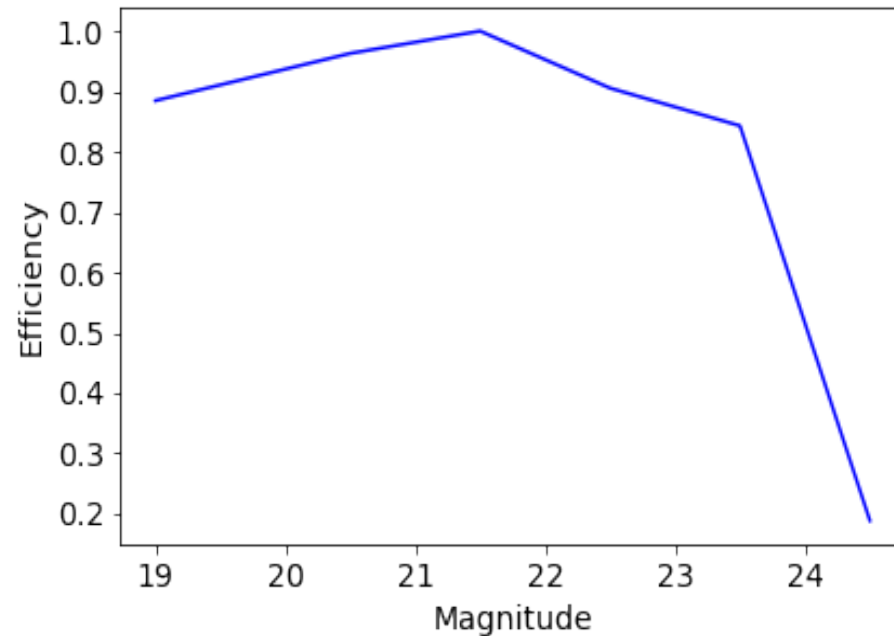
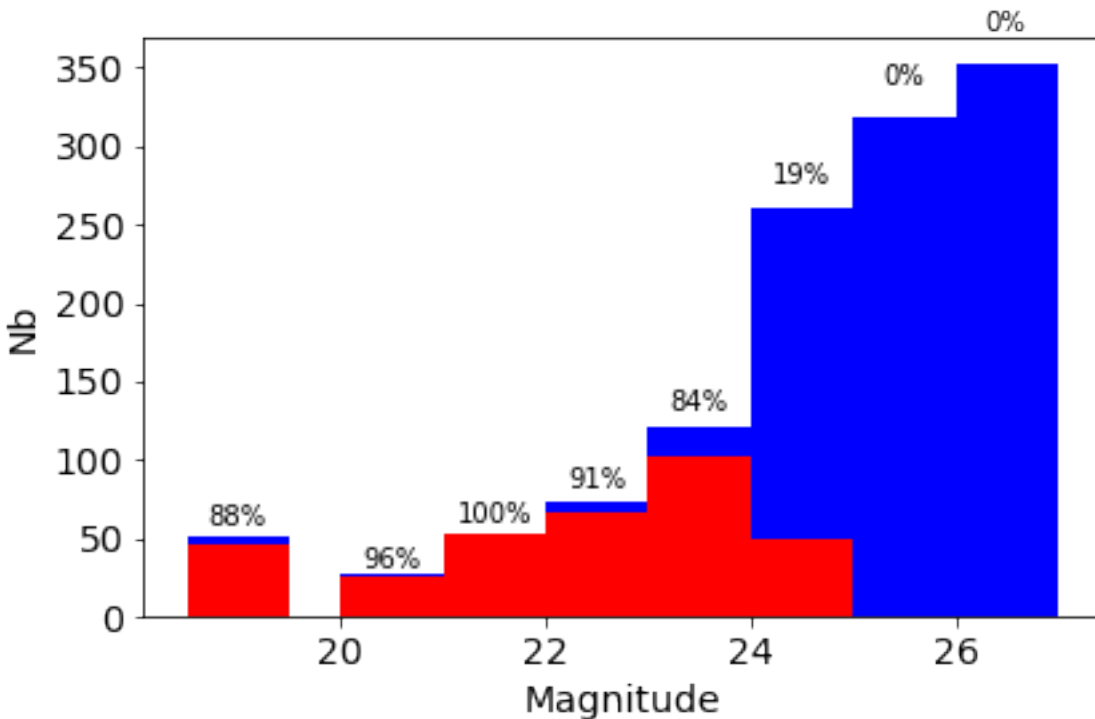


Faint fakes [24 - 25]

- Retrieved only 49/260 fakes with angular distance < 0.2 as
- Large magnitude errors in the difference images
- S/N close to the detection threshold in the difference image (5.5)



Detection efficiency



- Check the positions of the missing bright fakes (masks?)
- Increase the detection threshold to detect fakes with mag > 24

Perspectives

- Optimize the detection threshold
- Consider flags, masked regions in the images
- Generate fakes on the whole focal plane
- Use data from the full PDR2 UDEEP COSMOS broad bands (642 visits)
- Study of false transient detection rates (artefacts)

Thank you