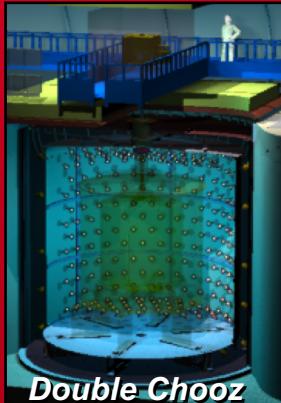


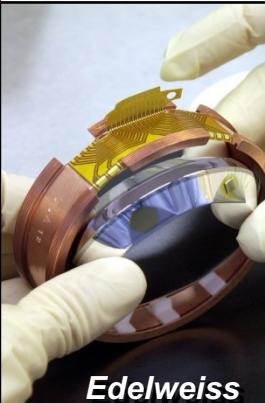
DE LA RECHERCHE À L'INDUSTRIE



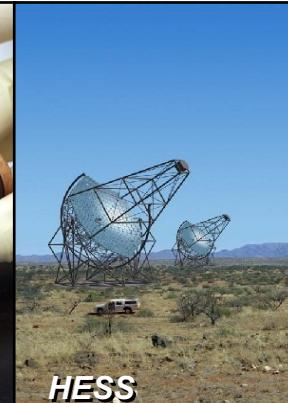
*Double Chooz*



*ALICE*



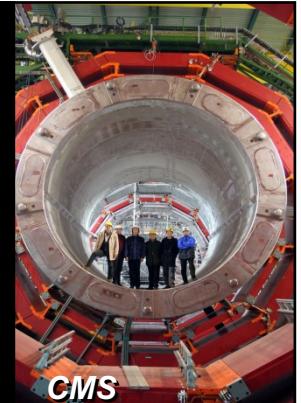
*Edelweiss*



*HESS*



*Herschel*



*CMS*

*Déchiffrer les rayons de l'Univers*



Valérie Gautard

25/09/2020

# Les applications du ML à l'IRFU

# OUTLINE

- Overview of ML activities @ Irfu / IN2P3
  - ML algorithms and tools
  - Computing and hardware resources
  - Covered activities
  - Training and schools
  - Conclusion

## Internal (HEP) tools

- ROOT framework for data storage and processing
- Multivariate Analysis: TMVA for mostly BDT and (deep) NN
- Specific for Neural Networks: NeuroBayes

## External tools

- Data format: text, csv, images, HDF5, ...
- ML libraries: Keras+TensorFlow, Pytorch, scikit-learn (no DL), ...
- All kinds of popular algorithms: CNN, GAN, RNN, LSTM, AE, VAE ...

## Interfaces and middleware

- PyMVA: Interface TMVA and Keras
- Several middleware file format conversion solutions:

<b>PyROOT</b>	Python extension module that allows the user to interact with ROOT data/classes.	[69]
<b>root_numpy</b>	The interface between ROOT and NumPy supported by the Scikit-HEP community.	[65]
<b>root_pandas</b>	The interface between ROOT and Pandas dataframes supported by the DIANA/HEP project.	[70]
<b>uproot</b>	A high throughput I/O interface between ROOT and NumPy.	[71]
<b>c2numpy</b>	Pure C-based code to convert ROOT data into Numpy arrays which can be used in C/C++ frameworks.	[72]
<b>root4j</b>	The hep.io.root package contains a simple Java interface for reading ROOT files. This tool has been developed based on freehep-rootio.	[73]
<b>root2npy</b>	The go-hep package contains a reading ROOT files. This tool has been developed based on freehep-rootio.	[73]
<b>root2hdf5</b>	Converts ROOT files containing TTrees into HDF5 files containing HDF5 tables.	[74]

## ML computing @ IN2P3/Irfu

- Mostly **CPU**, sometimes **GPU**, and some attempts with **FPGA**
- **Local** resources: laptop, lab/university clusters
- **CCIN2P3** resources: lots of CPU, less GPU

## Any other resources ?

- Tensor Processing Units (**TPU**)
- Vision Processing Units (**VPU**)
- Calculation on **cloud** from industry ?Amazon Web Services machines
- Google colab notebook with GPU support

...

- Astrophysics
- Nuclear physics
- Accelerator / Detectors
- High energy physics
- Medical physics
- Ethics

#### Disclaimer

Just a brief outlook of all activities based on input that we received  
And thanks a lot to all who provided material and explanations !

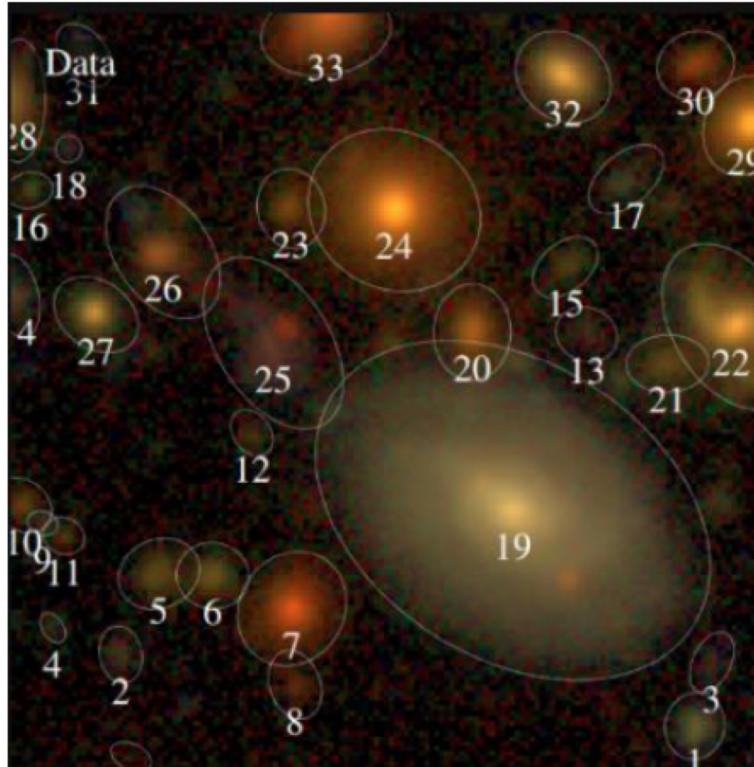
**ASTROPHYSICS**

**DEBLENDING**

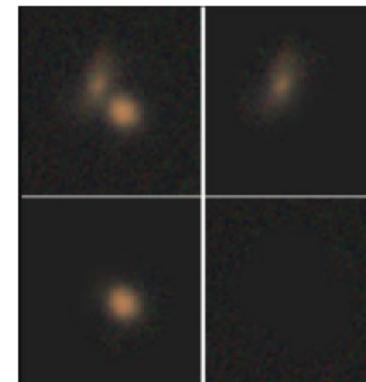
**IDENTIFICATION OF  
RADIONUCLIDES**

- **Cosmostat :**  
<http://www.cosmostat.org/communication>
  - ASTRONOMICAL IMAGE PROCESSING
    - . Weak And Strong Lensing
    - . Dark matter maps
    - . ERedshift
    - . Deblending
  - ACTIVITIES IN PARTNERSHIP WITH NEUROSPIN: THE OBJECTS OF STUDY ARE NOT THE SAME BUT THE METHODS ARE THE SAME

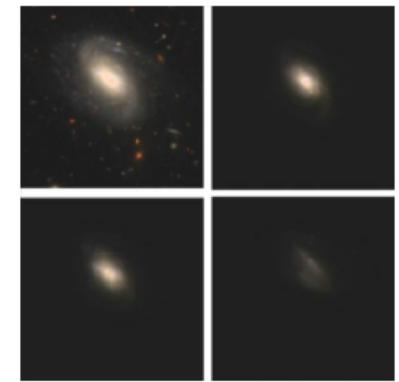
# THE DEBLENDING PROBLEM



- Galaxies overlap on the sky, they are blended, they need to be de-blended in order to measure the properties of each galaxy individually
- Current deblenders struggle on complex images

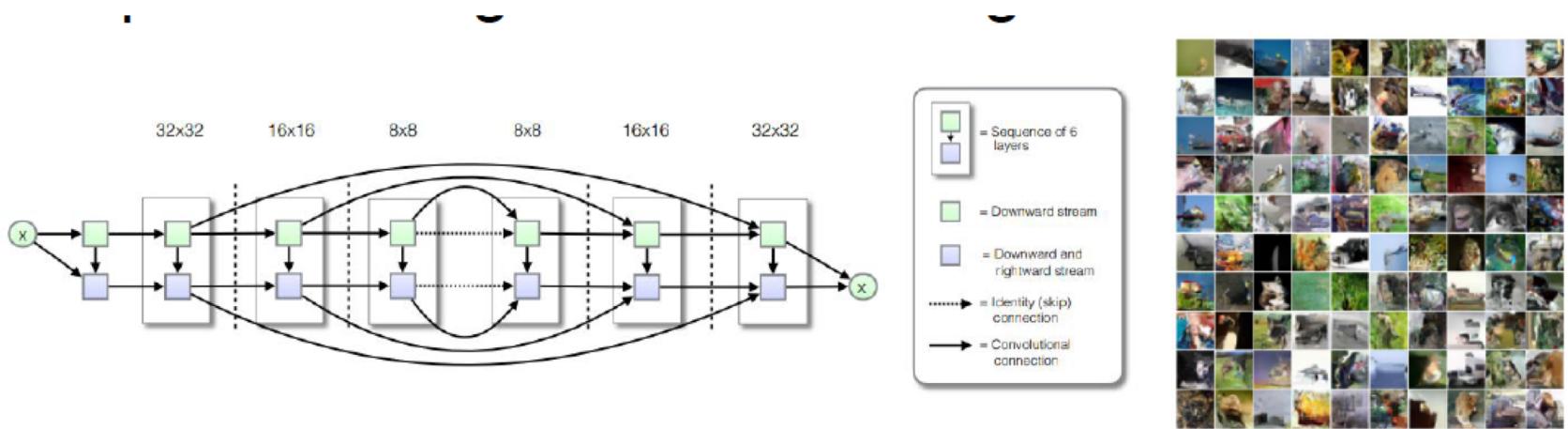


Successful deblending



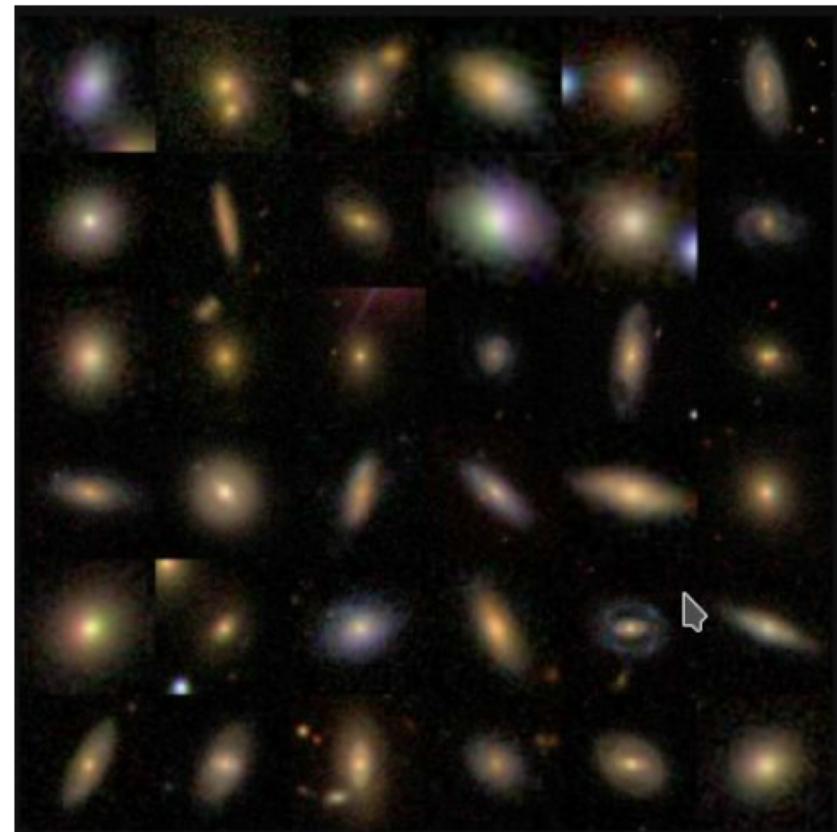
Deblending Failure

# DEEP DEBLENDING WITH LEARNED SIGNAL PRIORS



- The idea: learn what galaxies look like using a deep generative model, here a PixelCNN++ (Salimans, 2017)
- The model learns the probability  $p(x)$  of a given image  $x$  to belong to the training set

Which set of galaxies is real ?



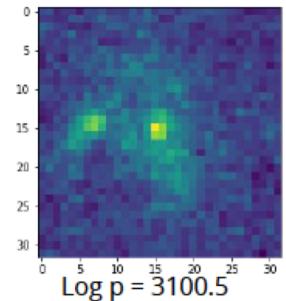
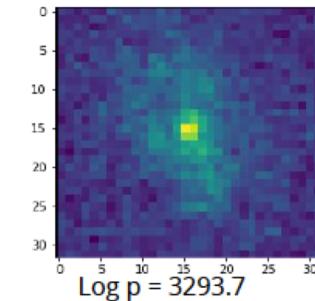
# DEBLENDING AS AN OPTIMIZATION PROBLEM

- We are trying to find deblending solutions that fits the data the best, and at the same time the recovered galaxies should look realistic

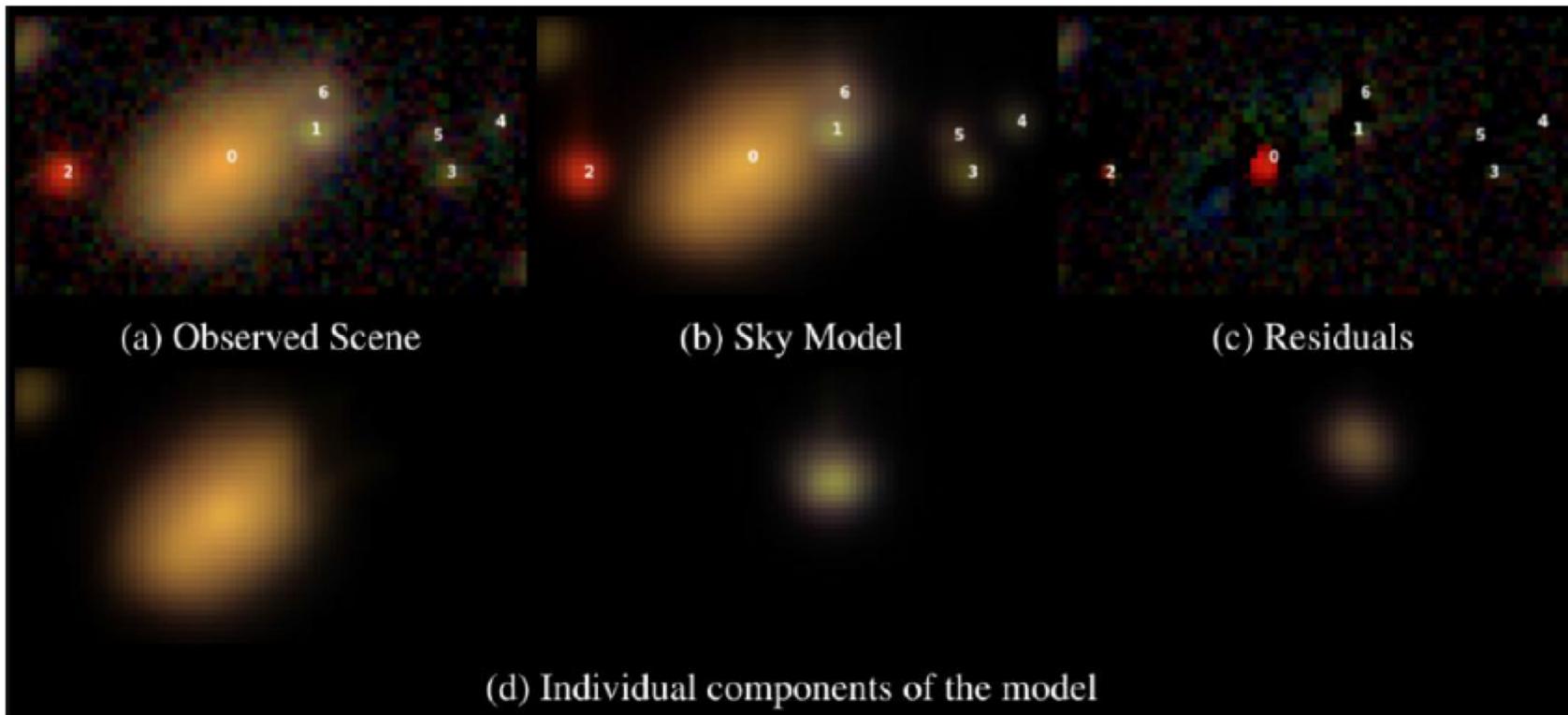
$$\arg \min \frac{1}{2} \| Y - \sum_k S_k \|_{\Sigma}^2 + \log p_{\theta}(S_1) + \log p_{\theta}(S_2)$$

where S1 and S2 are the individual blend components, Y is the observed image, and p is the learned probability of being a galaxy

- Illustration of probabilities for blended and isolated galaxies:  
The probability is higher for the isolated galaxies



# EXAMPLE OF DEBLENDING SOLUTIONS

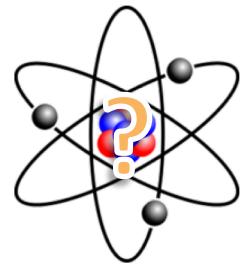


**Nuclear safety monitoring:**

- Decommissioning and dismantling
- Post-accidental scenes

**Nuclear security:** radiation portal monitor

Characterization of radiological scenes



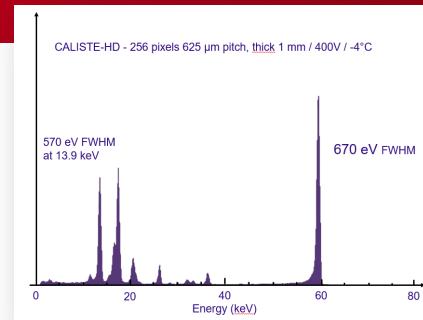
- **Identification of the radionuclides in the scene**
- **Proportion of each identified radionuclide**

Fukushima accident

# CALISTE DETECTOR

**CdTe** semi-conductor crystal

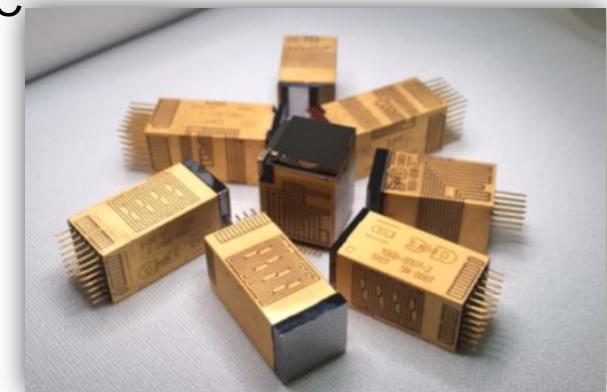
**Miniature pixelated spectro-imager**



Works at **nearly room temperature**: high performance at -15°C

**Low power consumption**: 200 mW

First developments for **astrophysical** application  
 → STIX: Spectrometer Telescope Imaging X-rays  
     Observation of Bremsstrahlung from accelerated  
     electrons near the Sun



Different versions of Caliste: Caliste-SO, Caliste-HD, Caliste-O...

From space applications to **industrial** applications:  
 → Medical application: breast tumor cells detection  
 → **Nuclear safety application**



# PROBLEM : SPECTRAL IDENTIFICATION

**Input : Calibrated event list**

Select	FRAME	MULTIPLICITY	MULT	TIME	PIXEL	X	Y	TYPE	ENERGY
	1I	1B	1B	1D	1B	1I	1I	1I	1D
1	0	1	0	0.000000000000E+000	128	0	8	1	9.794309200000E+001
2	1	1	0	3.062956094742E+000	145	1	9	1	9.333808000000E+001
3	2	1	0	3.107777595520E+000	126	14	7	1	5.661259000000E+001
4	3	1	0	3.311278343201E+000	176	0	11	1	1.204386400000E+002
5	4	1	0	4.261723756790E+000	117	5	7	1	5.908343000000E+001
6	5	1	0	4.685714244843E+000	155	11	9	1	3.126330300000E+001
7	6	1	0	7.426927089691E+000	134	6	8	1	2.661225810000E+002
8	7	2	0	7.764491081238E+000	242	2	15	255	1.749924320000E+002
9	8	2	1	7.764491081238E+000	243	3	15	255	2.266542400000E+001
10	9	1	0	8.099779129028E+000	226	2	14	1	2.473193900000E+001
11	10	2	0	8.122672319412E+000	119	7	7	255	3.066022800000E+001
12	11	2	1	8.122672319412E+000	135	7	8	255	1.780166800000E+001
13	12	1	0	8.208755016327E+000	252	12	15	1	4.164087000000E+001
14	13	1	0	8.331153392792E+000	254	14	15	1	7.991901300000E+001
15	14	1	0	8.772984504700E+000	47	15	2	1	1.588442560000E+002
16	15	1	0	9.839590549469E+000	59	11	3	1	2.487148740000E+002
17	16	1	0	9.900824785233E+000	191	15	11	1	2.493642300000E+001
18	17	1	0	1.208127880096E+001	239	15	14	1	6.909980000000E+001
19	18	1	0	1.288081645966E+001	10	10	0	1	2.669001630000E+002
20	19	1	0	1.323039579391E+001	97	1	6	1	1.040922650000E+003
21	20	1	0	1.323042392731E+001	97	1	6	1	5.779931210000E+002
22	21	1	0	1.323045206070E+001	97	1	6	1	1.885572170000E+002
23	22	1	0	1.470239663124E+001	120	8	7	1	1.665020800000E+001
24	23	1	0	1.4720333093937E+001	77	13	4	1	7.437592800000E+001
25	24	1	0	1.529405093193E+001	52	4	3	1	8.464667200000E+001
26	25	1	0	1.603019905090E+001	217	9	13	1	5.825236000000E+001
27	26	1	0	1.623439192772E+001	169	9	10	1	3.984544000000E+001
28	27	1	0	1.637744331360E+001	202	10	12	1	5.378394600000E+001
29	28	1	0	1.652689838409E+001	48	0	3	1	3.205725600000E+001

**Outputs**



Which **radioelements?** → Classification



In which **proportions?** → Regression



With **uncertainties?**

Some constraints:

- **Real-time** computation
- Identification for **low-statistics** of photon
- Independent on **operational conditions** (temperature, high-voltage... → impact on calibration)
- Not sensitive to environmental conditions (presence of absorbing materials or diffusing materials)

# DATA TRANSFORMATION

**Input**

Calibrated event list

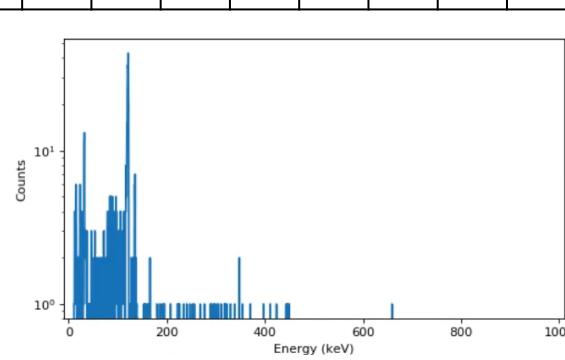
Select	#	NAME	MULT	PNAME	PISEL	Y	TYPE	ENERGY
	Mod1	Mod2	Mod3	Mod4	Mod5	Mod6	Mod7	Mod8
1	1	1	1	1	1	1	1	1
2	2	2	2	2	2	2	2	2
3	3	3	3	3	3	3	3	3
4	4	4	4	4	4	4	4	4
5	5	5	5	5	5	5	5	5
6	6	6	6	6	6	6	6	6
7	7	7	7	7	7	7	7	7
8	8	8	8	8	8	8	8	8
9	9	9	9	9	9	9	9	9
10	10	10	10	10	10	10	10	10
11	11	11	11	11	11	11	11	11
12	12	12	12	12	12	12	12	12
13	13	13	13	13	13	13	13	13
14	14	14	14	14	14	14	14	14
15	15	15	15	15	15	15	15	15
16	16	16	16	16	16	16	16	16
17	17	17	17	17	17	17	17	17
18	18	18	18	18	18	18	18	18
19	19	19	19	19	19	19	19	19
20	20	20	20	20	20	20	20	20
21	21	21	21	21	21	21	21	21
22	22	22	22	22	22	22	22	22
23	23	23	23	23	23	23	23	23
24	24	24	24	24	24	24	24	24
25	25	25	25	25	25	25	25	25
26	26	26	26	26	26	26	26	26
27	27	27	27	27	27	27	27	27
28	28	28	28	28	28	28	28	28
29	29	29	29	29	29	29	29	29
30	30	30	30	30	30	30	30	30
31	31	31	31	31	31	31	31	31
32	32	32	32	32	32	32	32	32
33	33	33	33	33	33	33	33	33
34	34	34	34	34	34	34	34	34
35	35	35	35	35	35	35	35	35
36	36	36	36	36	36	36	36	36
37	37	37	37	37	37	37	37	37
38	38	38	38	38	38	38	38	38
39	39	39	39	39	39	39	39	39
40	40	40	40	40	40	40	40	40
41	41	41	41	41	41	41	41	41
42	42	42	42	42	42	42	42	42
43	43	43	43	43	43	43	43	43
44	44	44	44	44	44	44	44	44
45	45	45	45	45	45	45	45	45
46	46	46	46	46	46	46	46	46
47	47	47	47	47	47	47	47	47
48	48	48	48	48	48	48	48	48
49	49	49	49	49	49	49	49	49
50	50	50	50	50	50	50	50	50
51	51	51	51	51	51	51	51	51
52	52	52	52	52	52	52	52	52
53	53	53	53	53	53	53	53	53
54	54	54	54	54	54	54	54	54
55	55	55	55	55	55	55	55	55
56	56	56	56	56	56	56	56	56
57	57	57	57	57	57	57	57	57
58	58	58	58	58	58	58	58	58
59	59	59	59	59	59	59	59	59
60	60	60	60	60	60	60	60	60
61	61	61	61	61	61	61	61	61
62	62	62	62	62	62	62	62	62
63	63	63	63	63	63	63	63	63
64	64	64	64	64	64	64	64	64
65	65	65	65	65	65	65	65	65
66	66	66	66	66	66	66	66	66
67	67	67	67	67	67	67	67	67
68	68	68	68	68	68	68	68	68
69	69	69	69	69	69	69	69	69
70	70	70	70	70	70	70	70	70
71	71	71	71	71	71	71	71	71
72	72	72	72	72	72	72	72	72
73	73	73	73	73	73	73	73	73
74	74	74	74	74	74	74	74	74
75	75	75	75	75	75	75	75	75
76	76	76	76	76	76	76	76	76
77	77	77	77	77	77	77	77	77
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84	84	84	84	84	84	84	84	84
85	85	85	85	85	85	85	85	85
86	86	86	86	86	86	86	86	86
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91	91	91	91	91	91	91	91	91
92	92	92	92	92	92	92	92	92
93	93	93	93	93	93	93	93	93
94	94	94	94	94	94	94	94	94
95	95	95	95	95	95	95	95	95
96	96	96	96	96	96	96	96	96
97	97	97	97	97	97	97	97	97
98	98	98	98	98	98	98	98	98
99	99	99	99	99	99	99	99	99
100	100	100	100	100	100	100	100	100

**Output**

List of sources:  $^{137}\text{Cs}$ ,  $^{57}\text{Co}$   
Proportions of the detected photons:  $^{137}\text{Cs}$ : 30%,  $^{57}\text{Co}$ : 70%

Vector of counts → Spectrum

Histogram of measured energies  
Classical way for these studies



Mathematical representation

$^{241}\text{Am}$	$^{133}\text{Ba}$	$^{57}\text{Co}$	$^{137}\text{Cs}$	$^{152}\text{Eu}$	$^{22}\text{Na}$
0	0	1	1	0	0
0	0	0,7	0,3	0	0

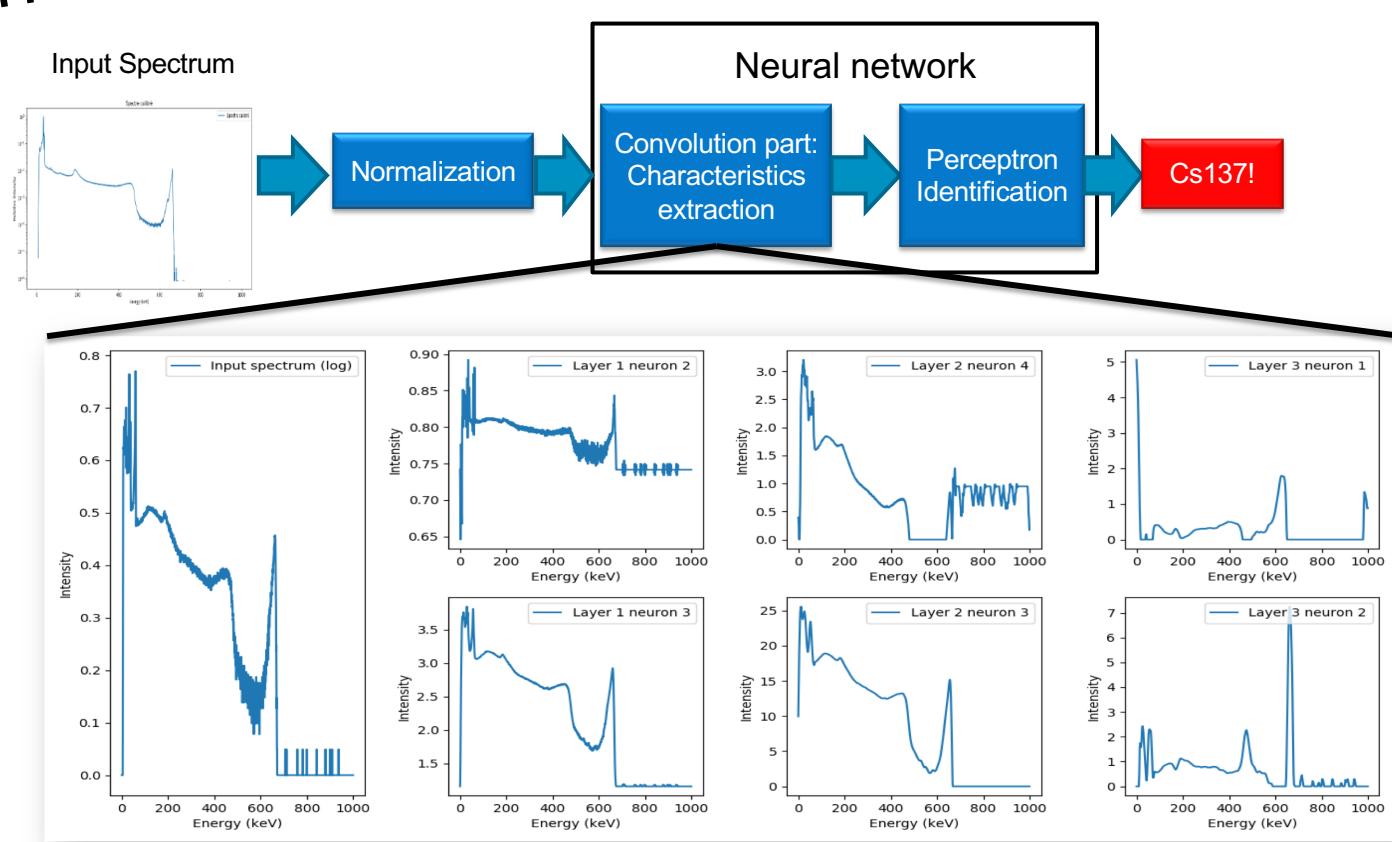
Normalisation  
Log-normalization, max = 1

**Output of ML algorithm**

**Convolutional Neural Network**

# CONVOLUTIONNAL NEURAL NETWORK

**One identification network  
for each radio-element:**



# PERFORMANCE

**Precision:** False positive influence

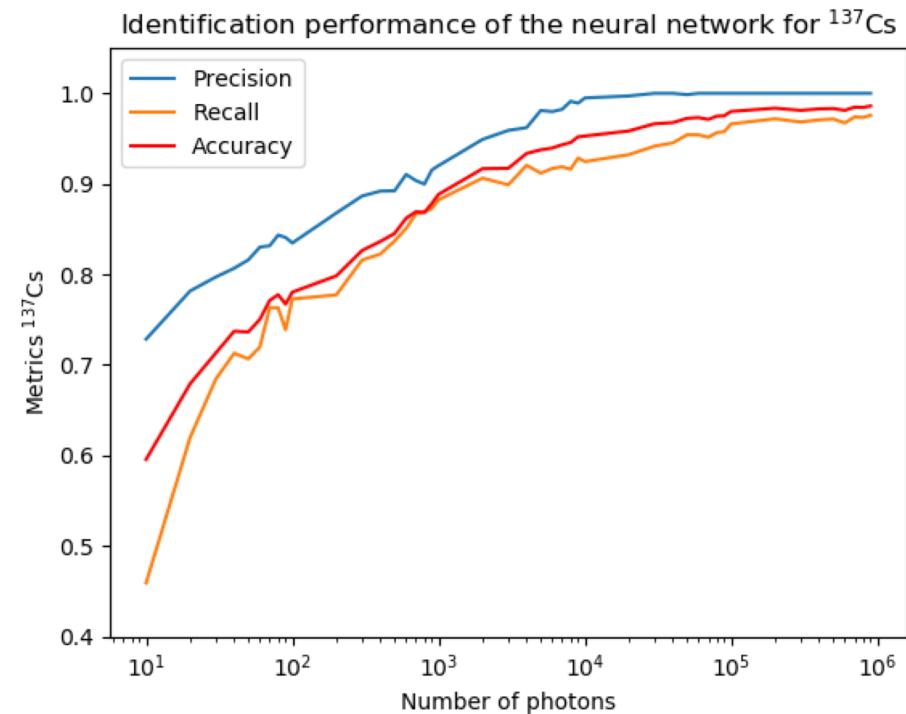
**Recall:** False negative influence

**Accuracy:** Right identification rate

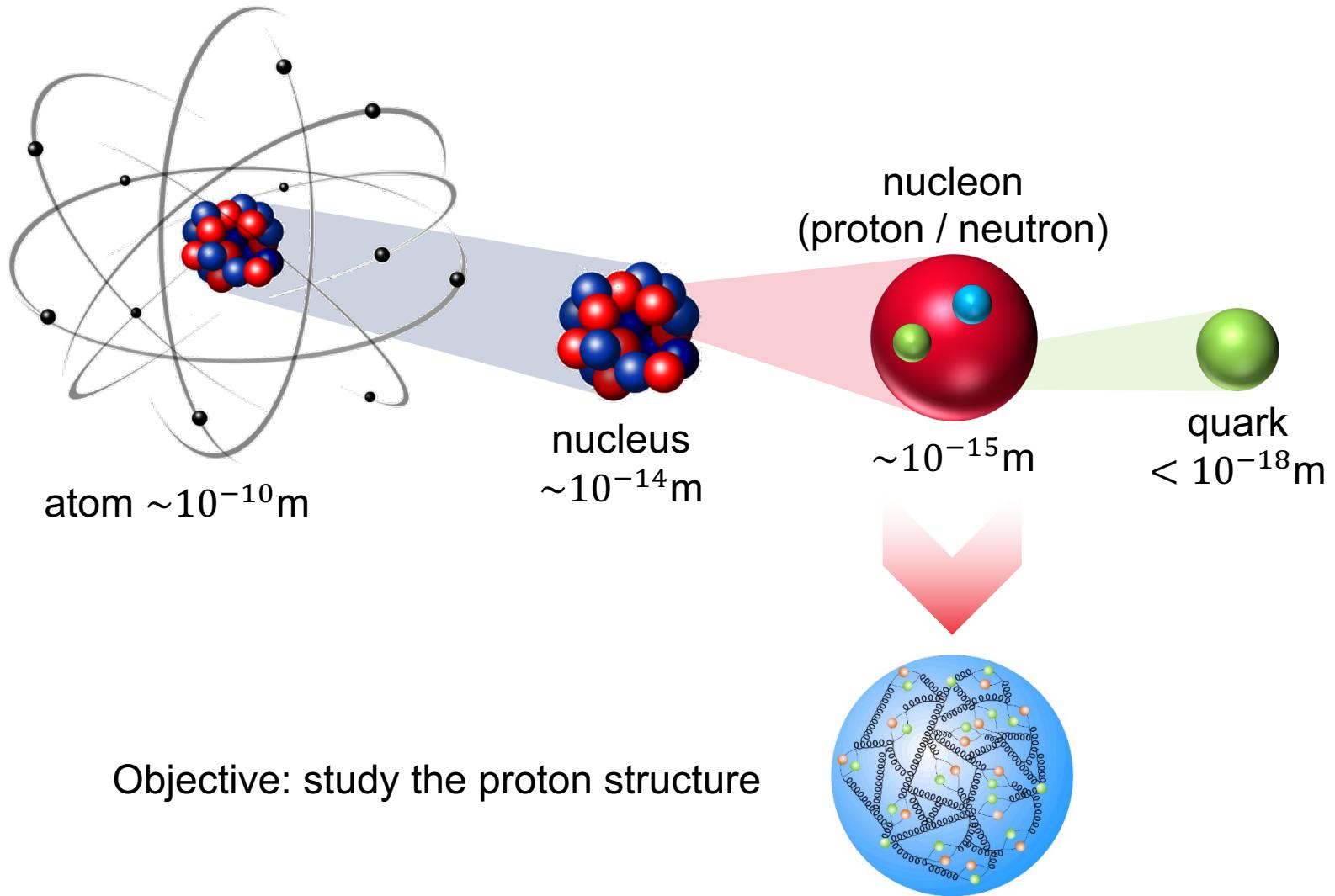
## Accuracy:

- > 80 % with more than 200 photons
- > 90 % with at least 1000 photons
- > 95 % with at least some thousands of photons
  
- Similar performance for other radionuclides:  $^{241}\text{Am}$ ,  $^{133}\text{Ba}$ ,  $^{57}\text{Co}$ ,  $^{152}\text{Eu}$ ,  $^{22}\text{Na}$

Test on real data of mixtures with **random decalibration**

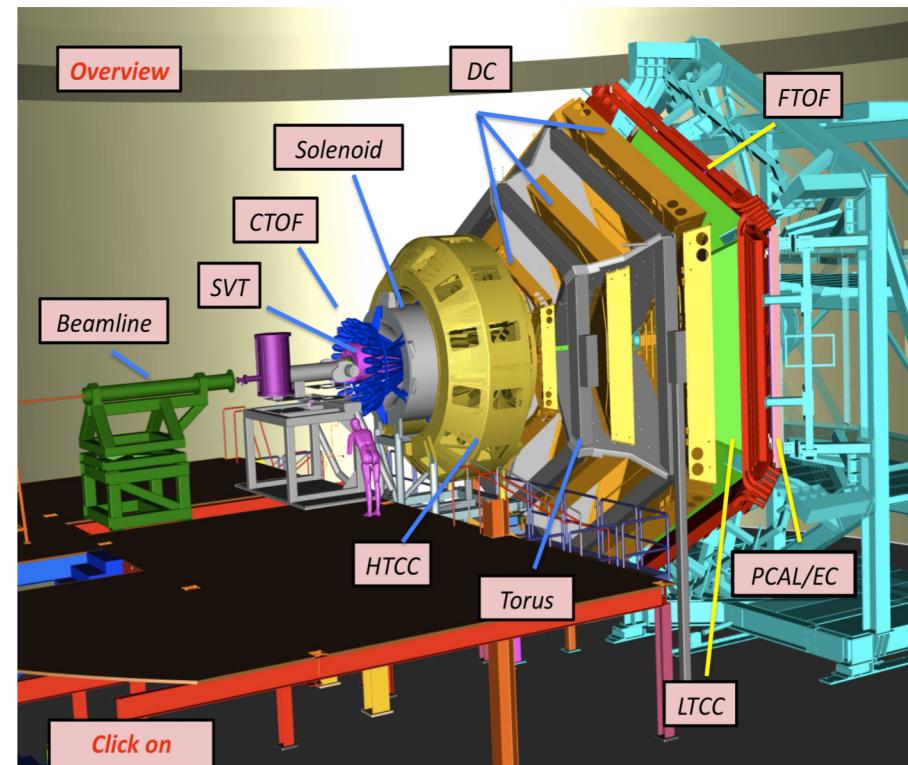


# **PHYSIQUE NUCLÉAIRE**

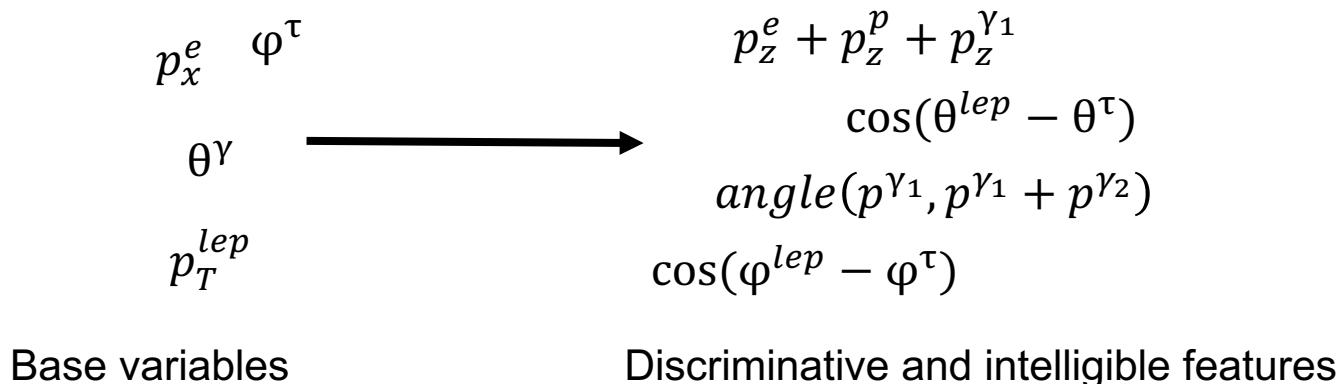


**JEFFERSON LAB: 10.6 GEV ELECTRON BEAM  
CLAS12 DATA TAKING SINCE 2018: HYDROGEN TARGET**

**EVENT CLASSIFICATION TASK: ISOLATE DVCS EVENTS ( $ep \rightarrow ep\Gamma$ )**

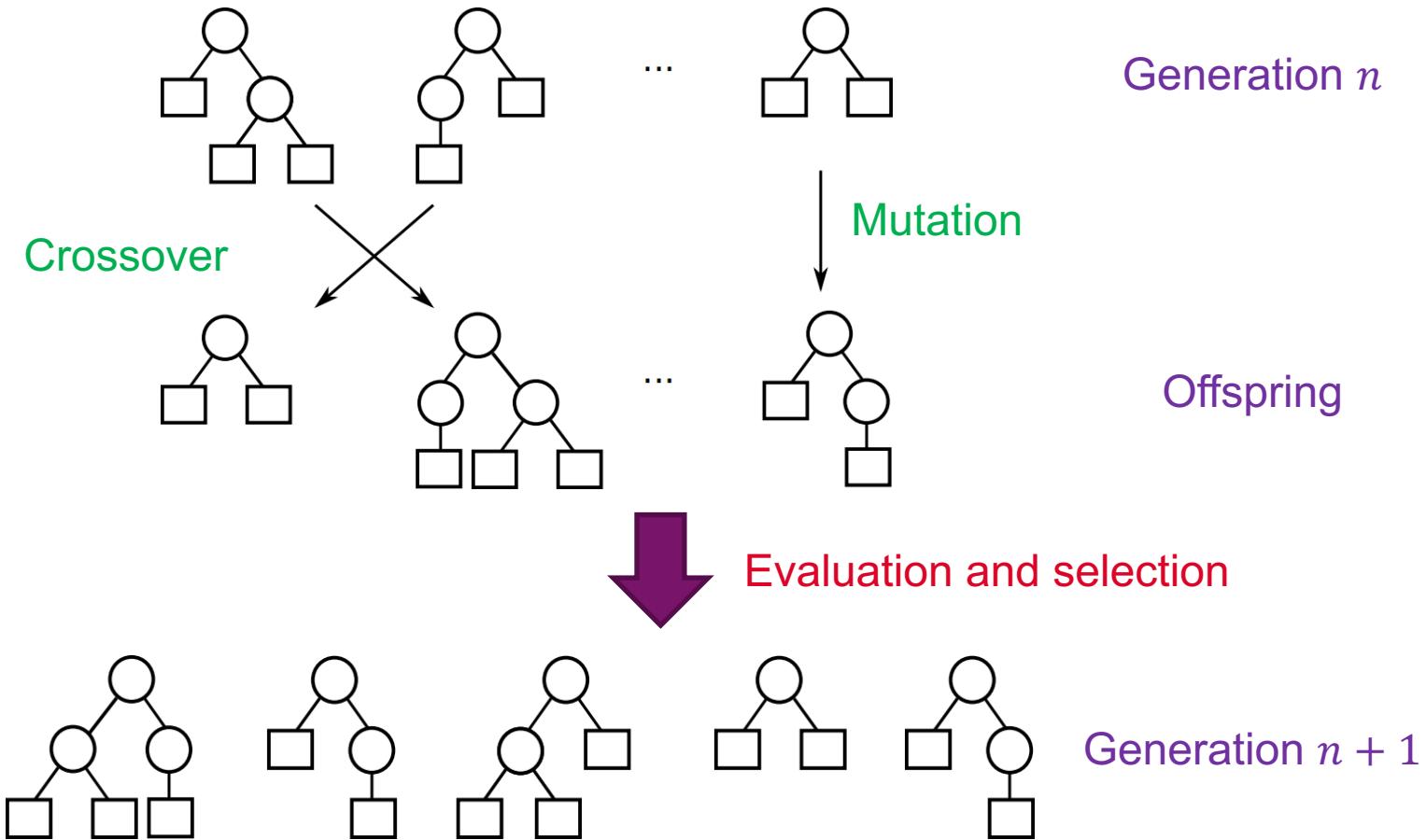


MOTIVATION: THESE MODELS DO NOT BUILD A SUFFICIENTLY COMPLEX INTERNAL REPRESENTATION OF THE DATA



IN MACHINE LEARNING: FEATURE ENGINEERING, FEATURE CONSTRUCTION

# FEATURE CONSTRUCTION



Cherrier, N., Defurne, M., Poli, J. P., & Sabatié, F. (2019). Embedded Constrained Feature Construction for High-Energy Physics Data Classification. In *Workshop on Machine Learning for the Physical Sciences, NeurIPS 2019*.

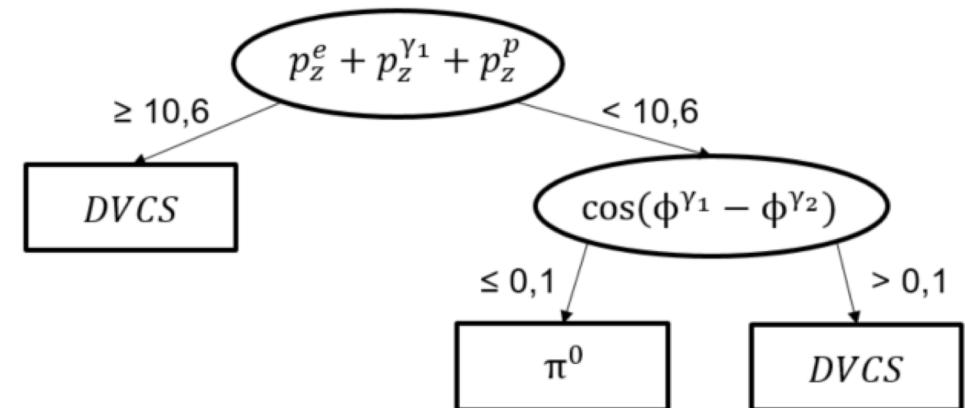
	$p_x^e$	$\varphi^\tau$	$\theta^\gamma$	Constrained Genetic Programming	Discriminative and intelligible features
					$p_z^e + p_z^p + p_z^{\gamma_1}$
					$\cos(\theta^{lep} - \theta^\tau)$
					$\text{angle}(p^{\gamma_1}, p^{\gamma_1} + p^{\gamma_2})$
					$\cos(\varphi^{lep} - \varphi^\tau)$

Features in the dataset

**embedded into tree node induction**

Fitness: tree splitting criterion (information gain, Gini index, ...)

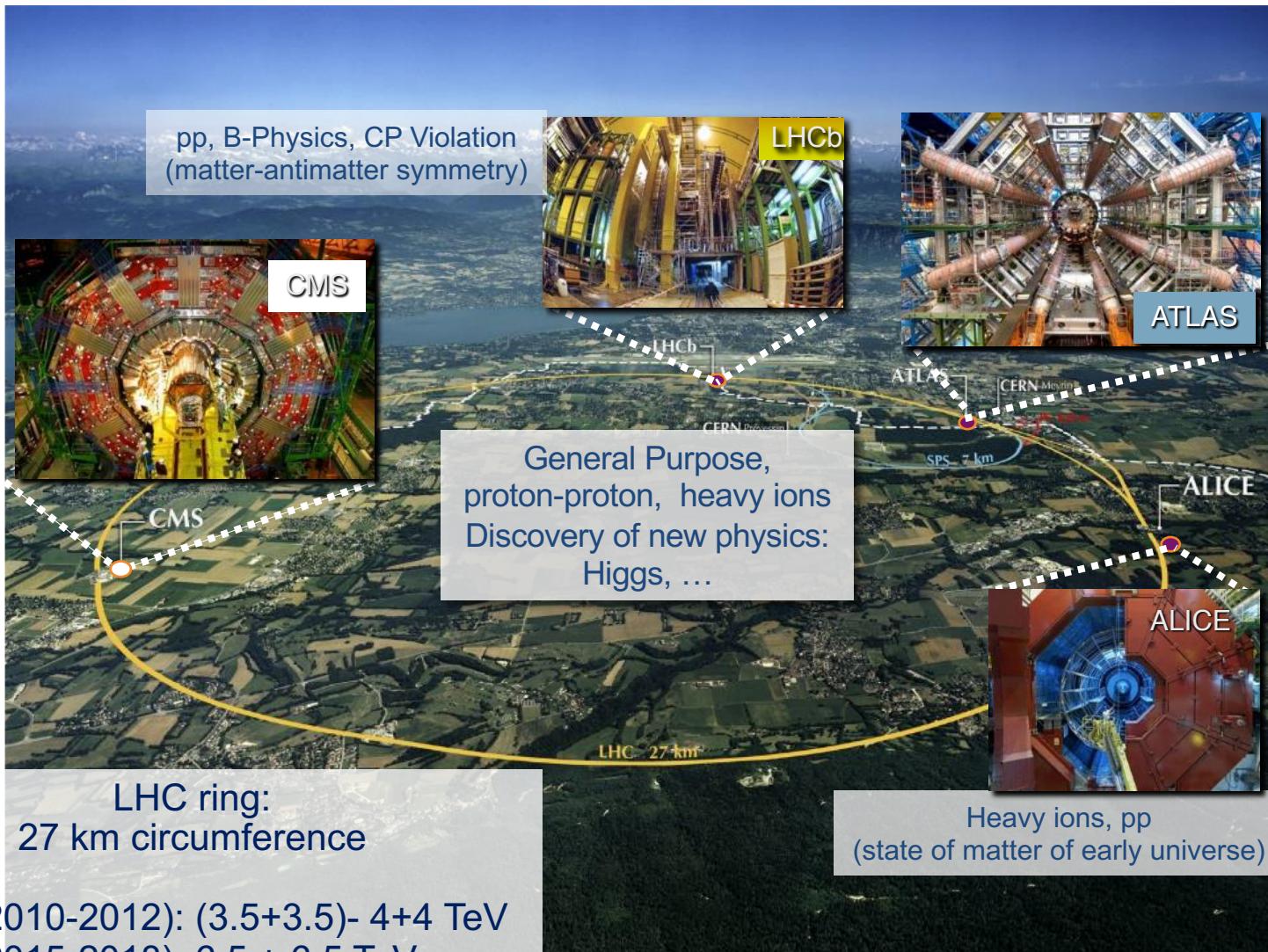
- ✓ Faster construction (“filter” methods are faster to evaluate than training a whole ML model)



**ACCELERATOR**

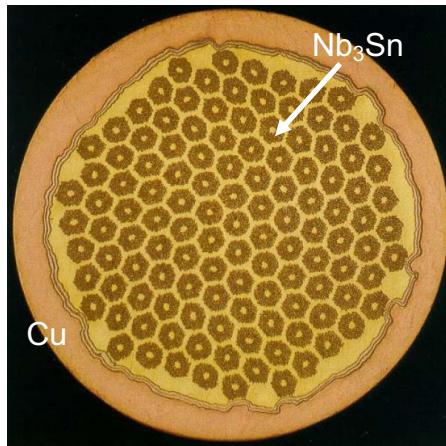
**DESIGNING  
SUPERCONDUCTING MAGNET**

**MODELLING DYNAMIC  
APERTURE**



# CHALLENGES IN DESIGNING SUPERCONDUCTING MAGNETS

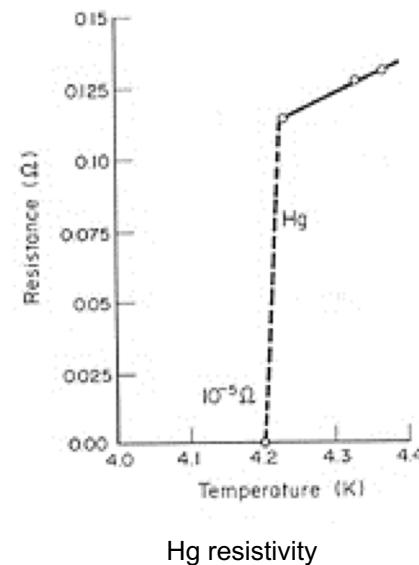
*Why superconductivity?*



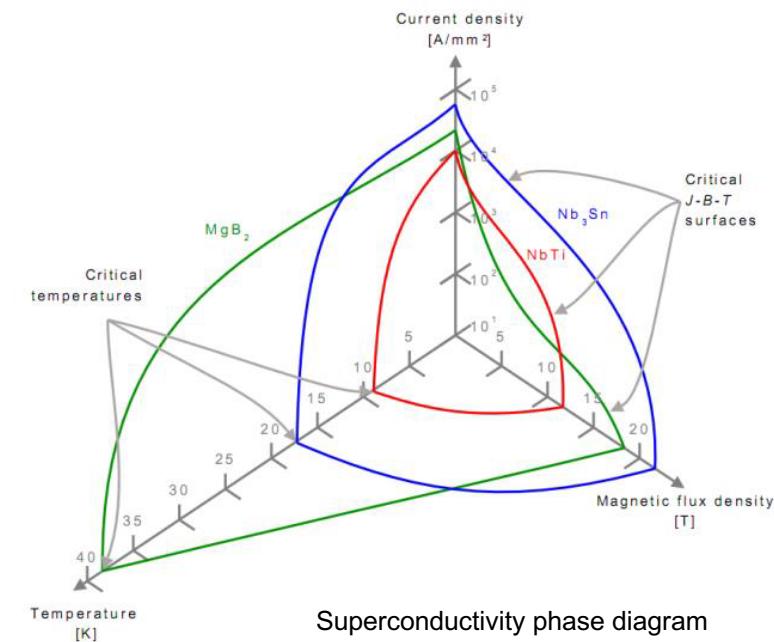
Nb<sub>3</sub>Sn bronze-process wire [1]



Superconductive materials show zero electric resistance  
→ they can transport a lot of current without heating



Hg resistivity



Superconductivity phase diagram

The superconducting state is limited by

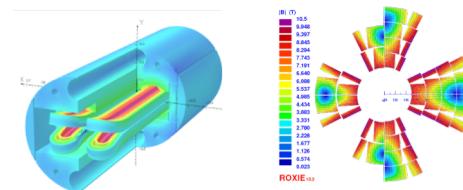
- $\vec{B}$  magnetic field
- $\vec{J}$  current density inside the material
- $T$  temperature

# CHALLENGES IN DESIGNING SUPERCONDUCTING MAGNETS

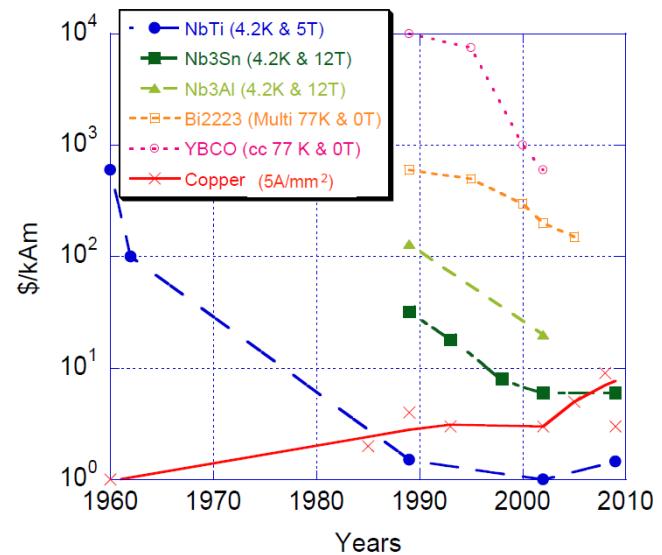
*Which challenges?*

- **Field quality**  
 $\Delta B/B \sim 10^{-4}$  for particle accelerators  
 $\Delta B/B \sim 10^{-8}$  for MRI
- **Mechanics**  
 Stress can reach 300 MPa on the conductor
- **Cryogenics**  
 Must cool down tons of materials
- **Protection**  
 Transition to the normal state means 1000s A in  $\sim \text{mm}^2$  wires (normal Cu wire carries some A)
- **Cost**  
 Superconductors can cost up to 1000\$/kAm

FCC Dipole & Quadrupole



US\$ 150/kg  $\leftrightarrow$  ~1 \$/kAm



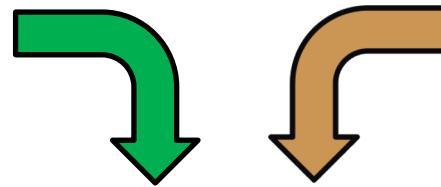
P. Fazilleau – Ecole des accelerateurs 2016

# WHY USING AI METHODS?

## Field Quality

Depends on:

- Position of the coils → linear contribution to  $B$
- Position of the ferromagnetic materials → non linear contribution to  $B$

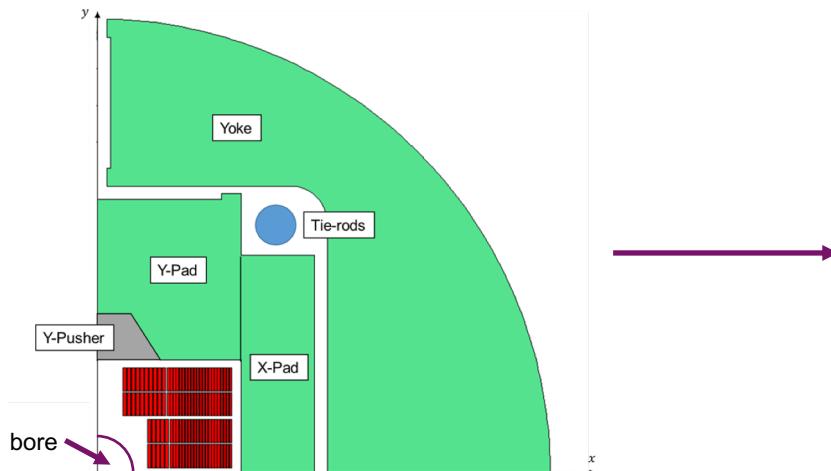


## Cost

Depends on:

- The volume of conductor needed
- The quantity of superconducting material needed in the cable to stand the magnetic field peak

**Not possible to write down an analytical function**



**Problem is :**  
**Given a certain  $B$  according to specs,**

### Field Quality

Minimization of the ratio  $\Delta B / B$  in the bore region

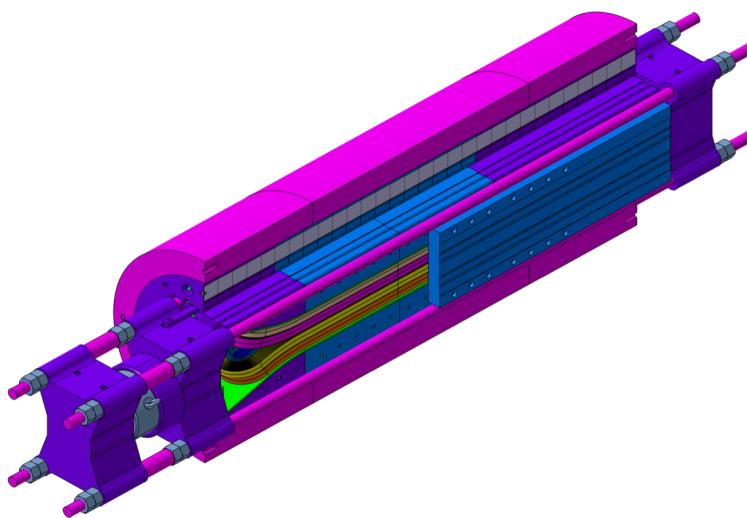
### Cost

Minimization of the ratio  $\Delta B / B$  on the conductors

## Field Quality

### F2D2

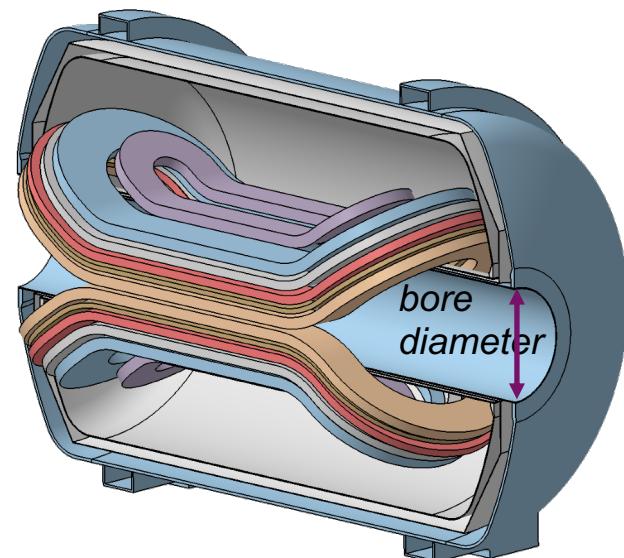
The 16 T graded Nb<sub>3</sub>Sn flared ends dipole to prove the feasibility of the Future Circular Collider at CERN.



## Cost

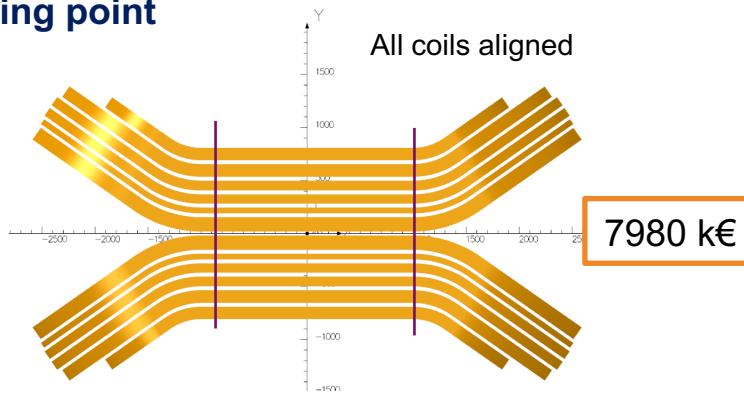
### MADMAX

The 9 T NbTi, 1.3 m of useful diameter dipole (the biggest in the world) to find axions-like particles.



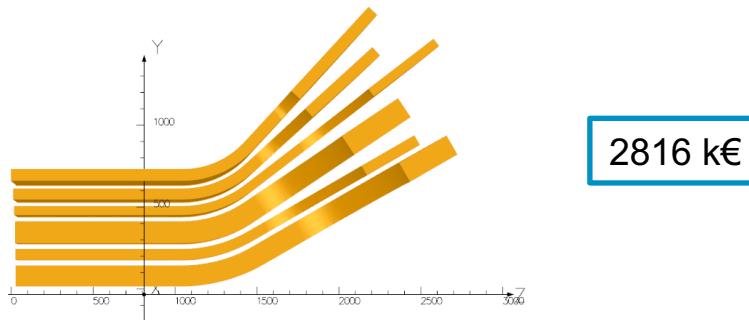
# RESULTS ON MADMAX VERSION 9.0

## Starting point



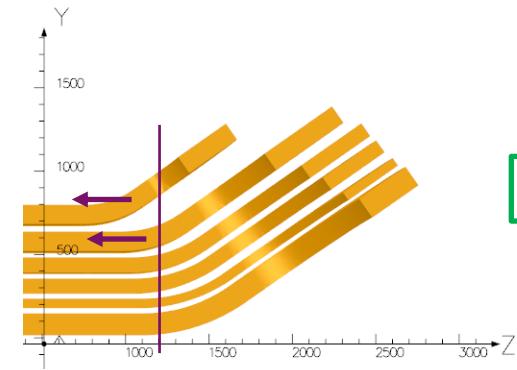
## Intermediate solution

GA study varying many parameters

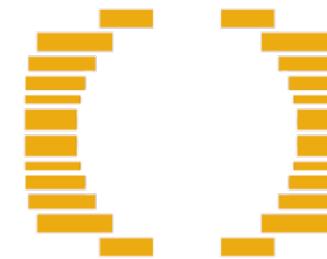


but not feasible for assembly

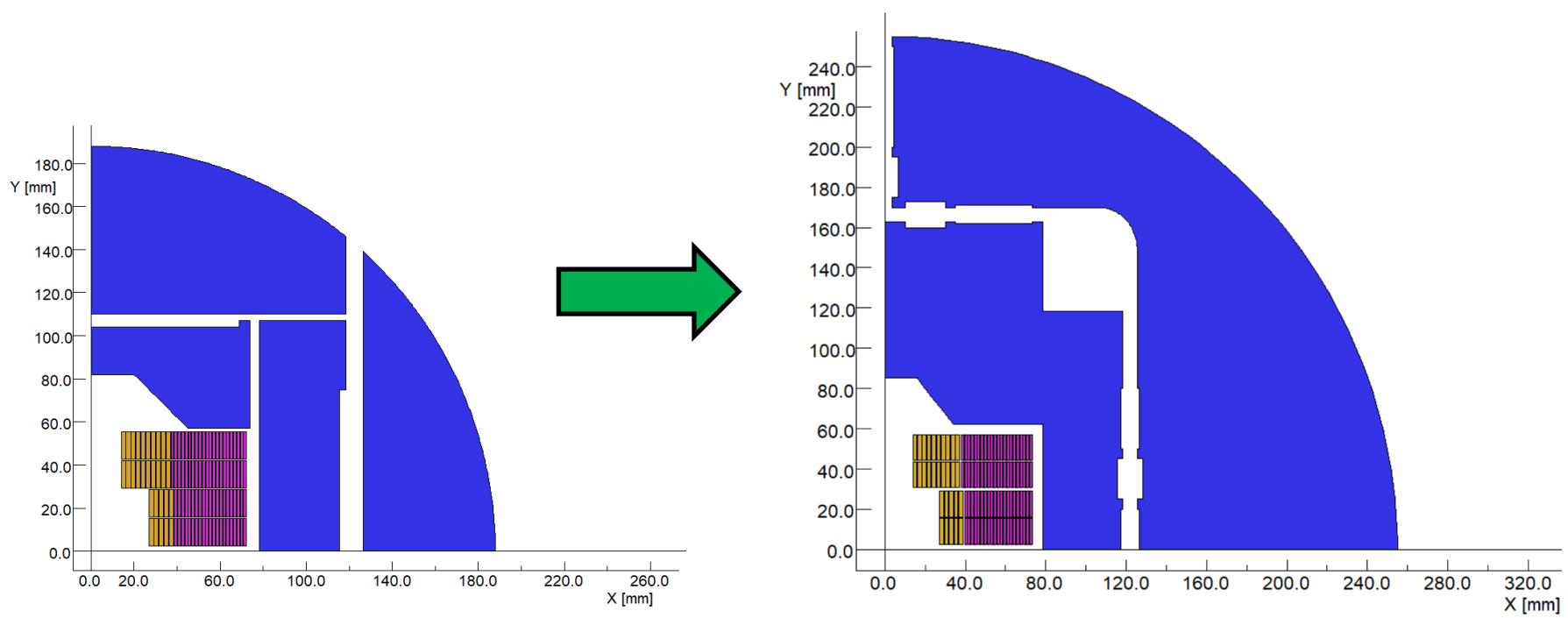
## Final solution



GA + parametric study  
to understand the sensitivity of parameters



# F2D2 – FROM BEGINNING TO FINAL DESIGN

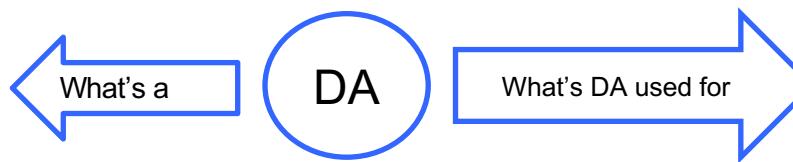


# MODELLING DYNAMIC APERTURE (DA) OF FUTURE CIRCULAR PARTICLE ACCELERATORS



- Focus on the long-term DA

region of stable motion of a particle in an accelerator



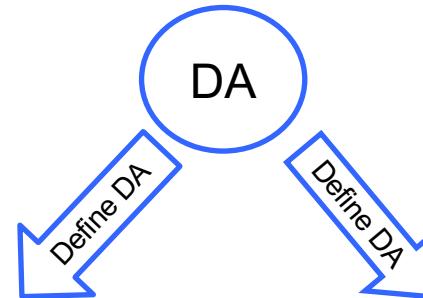
used to define tolerances on field quality

used to define non-linear corrections schemes in the design phase of the accelerator

- Different possibilities to define DA:

Particle amplitude

DA is the amplitude corresponding to particle losses after the  $10^n$  revolutions in the accelerator



Trajectory

DA is the divergence of nearby trajectories in the phase space to define the limit between chaotic and non-chaotic motion

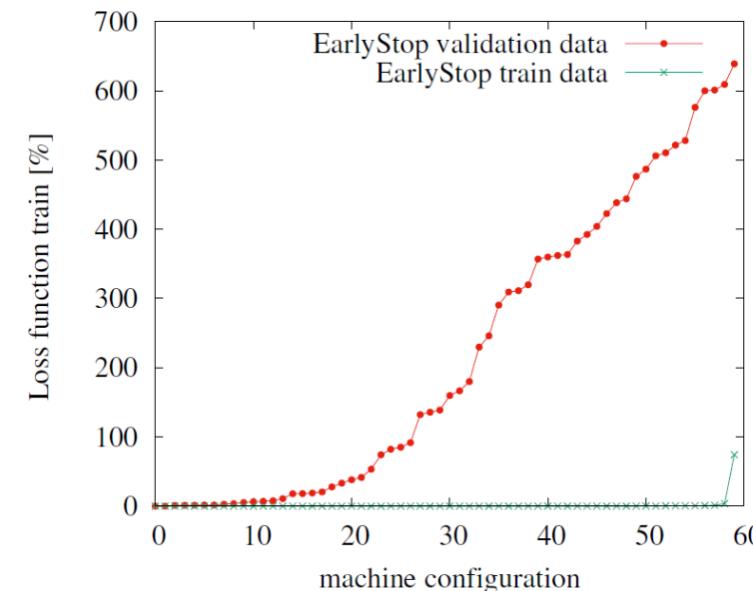
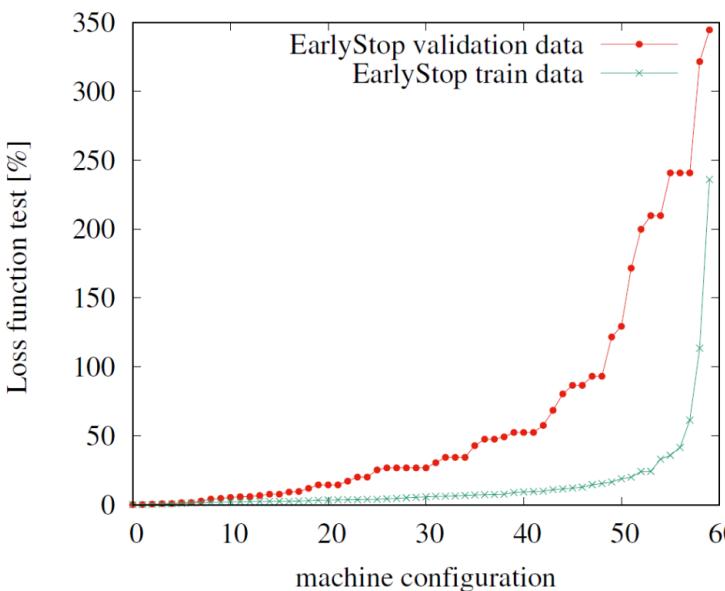
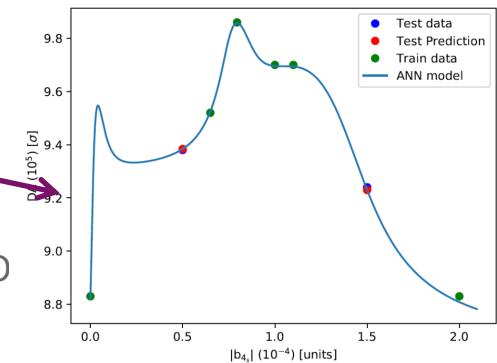
Simulate DA: From 1 day to one week or more...

# TUNING ANN TO INTERPOLATE DA VALUES

Minimum DA over 60 machine configurations as a function of octupole errors in the ring

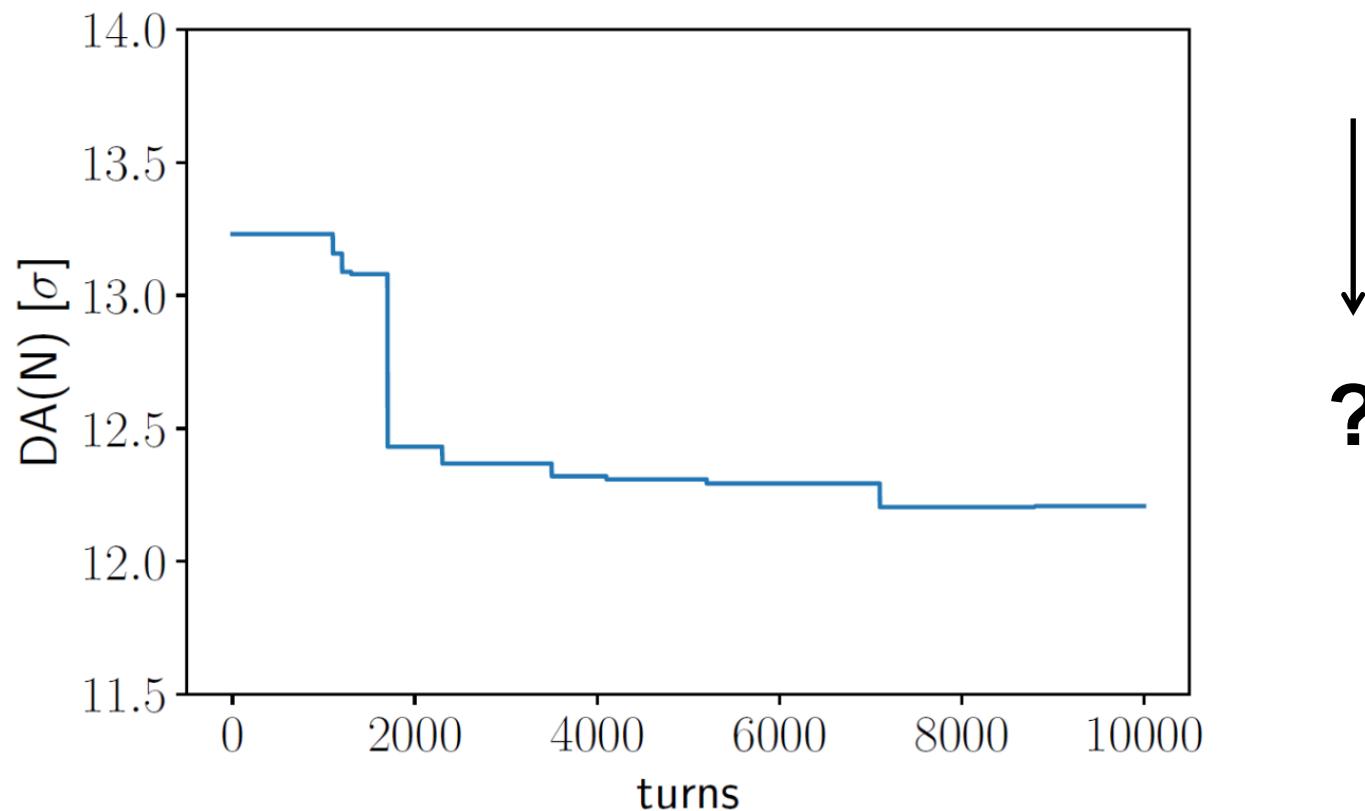
Same ANN on each machine configuration

- PERFORMANCE ARE NOT THE SAME FOR EACH MACHINE CONFIGURATION
- EARLY STOP ON TEST DATA NOT GOOD AS EARLYSTOP ON TRAIN
- STILL SOME OVERFITTING REMAINS
- LOSS FUNCTION VALUES DEPENDS ALSO ON SEED OF KERAS-TENSORFLOW INITIALIZATIONS



# TUNING ESN/RNN TO PREDICT DYNAMIC APERTURE VALUES

- Predict DA values at  $10^5$  turns knowing its values up to  $10^4$  turns, i.e. number (N) of revolutions in the machine



**DETECTORS**

**ATLAS NSW**

**CLEARMIND**

# RAINFROG: PROCESS AUTOMATION ON THE ATLAS/NSW FACILITY

## Context

Fabrication & test of detector elements in the CICLAD clean room facility at Saclay

NSW Micromegas detectors are scanned for planarity on 2 instrumented gantry granite tables

Planarity scans occur at various stages of construction and for final QC acceptance

A large variety of scans requiring different settings, a total of 31 distinct configurations

## Problem

Can we automatically find the active configuration, just by looking at the current part with a camera?

Algorithmic image processing would be very complicated.

**Can we handle this by machine learning?**

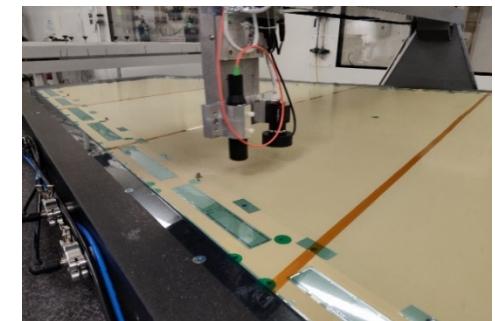


*A completed module being scanned for planarity with a contactless optical sensor*

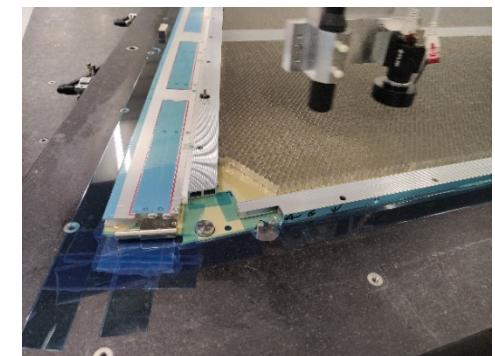
# LOOK AND GUESS

What needs  
to be  
discovered

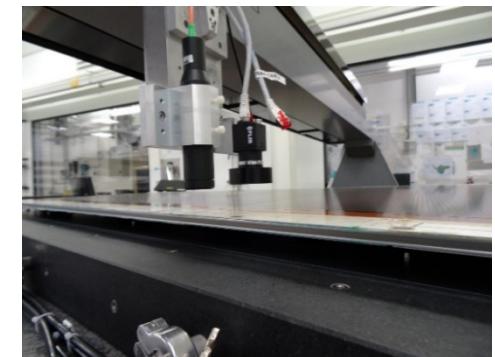
The type and subtype of the object: drift panel  
*front/central/back*, readout panel *eta/stereo*, module



The nature of the surface *pcb/honeycomb/resist/copper*



Which side of the part is exposed for measurement  
*A/B*



A non  
standard  
image  
classification  
problem

A total of 20 categories in 5 families

Will pick the highest score in each family

Categories are non exclusive, showing inclusion  
and intersection across families

*A variety of configurations of  
parts, sides, surfaces, heights*

# IMAGE COLLECTION STRATEGY

Take photos at 6 strategic locations

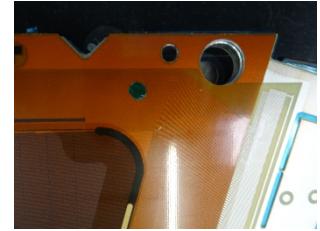
- A Pcb frontier & edge
- B, C, E, F Corners
- D Interconnect

Independantly infer for each photo

Fuse 6 scores for final decision

*This example session*

- S23\_BonPads Stereo readout panel
- Side B
- Sitting on pads

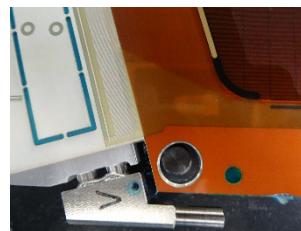


**B**

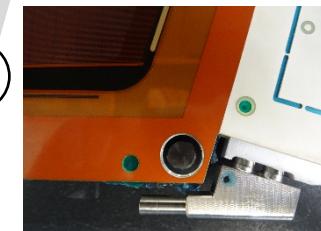
**A**

**D**

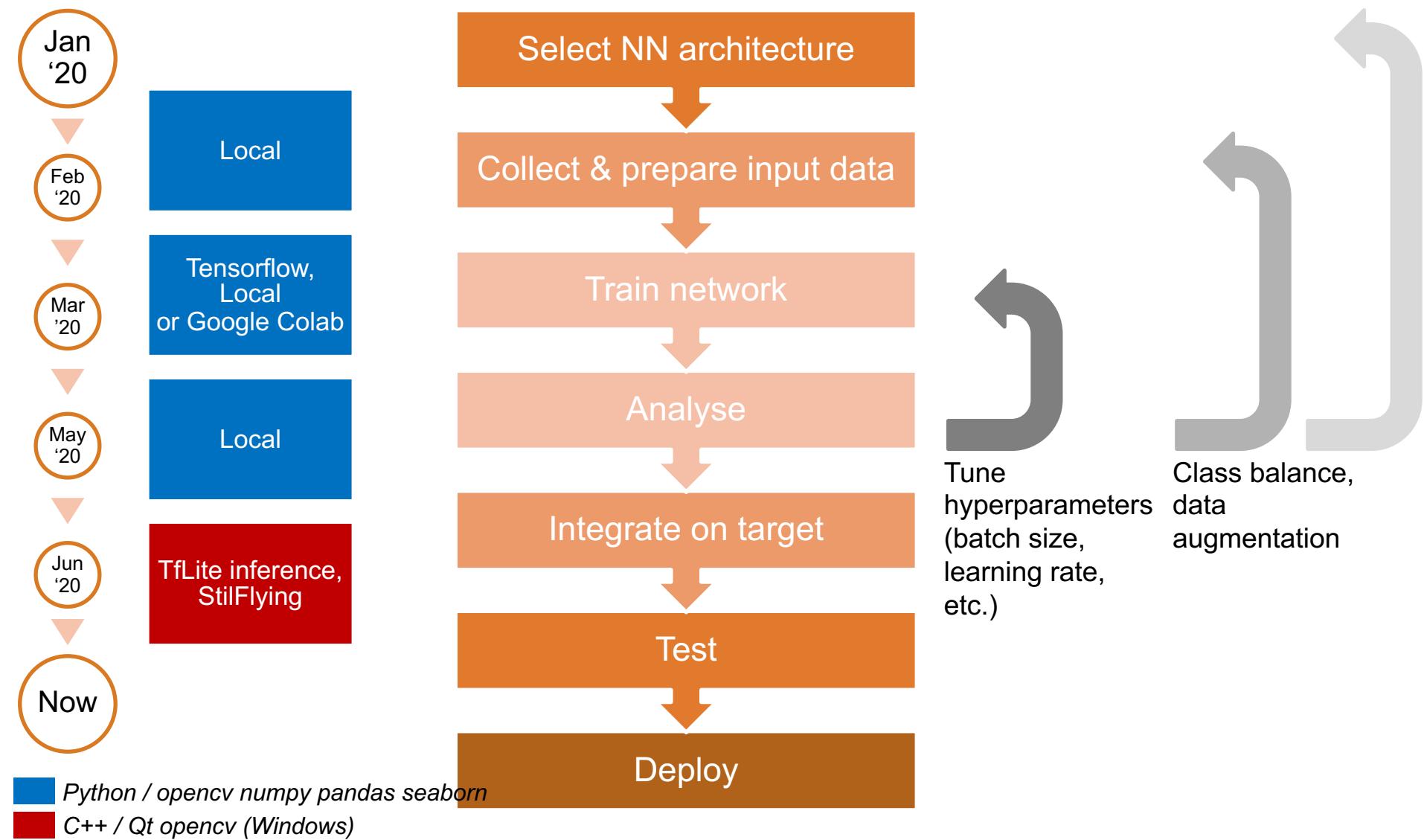
**C**



**F**



# WORKFLOW



# NN SELECTION, TRANSFER LEARNING & TRAINING

## NN

(headless)

MobileNetV2

feature extractor

A Google development

Aimed at « small » inference  
machines (mobile phones...)

Pretrained on  
ImageNet public  
dataset

Publicly available repository

**2.3 million trained parameters**

Fully connected  
classification head  
added (2 layers)

Will produce our own category label  
scores

**1.3 million trainable  
parameters**

## Training

750 source images  
in 60 distinct  
sessions *Too few!*

Hand taken Jan to Mar '20

Dataset augmented  
to 7000+ images

Improve category balance

Shift, rotate, adj. brightness,  
saturate, smooth, crop and scale

Dedicated loss  
function needed

Maximise soft-F1,  
harmonic mean of precision and recall

... when handling false/true  
positive/negative category matches

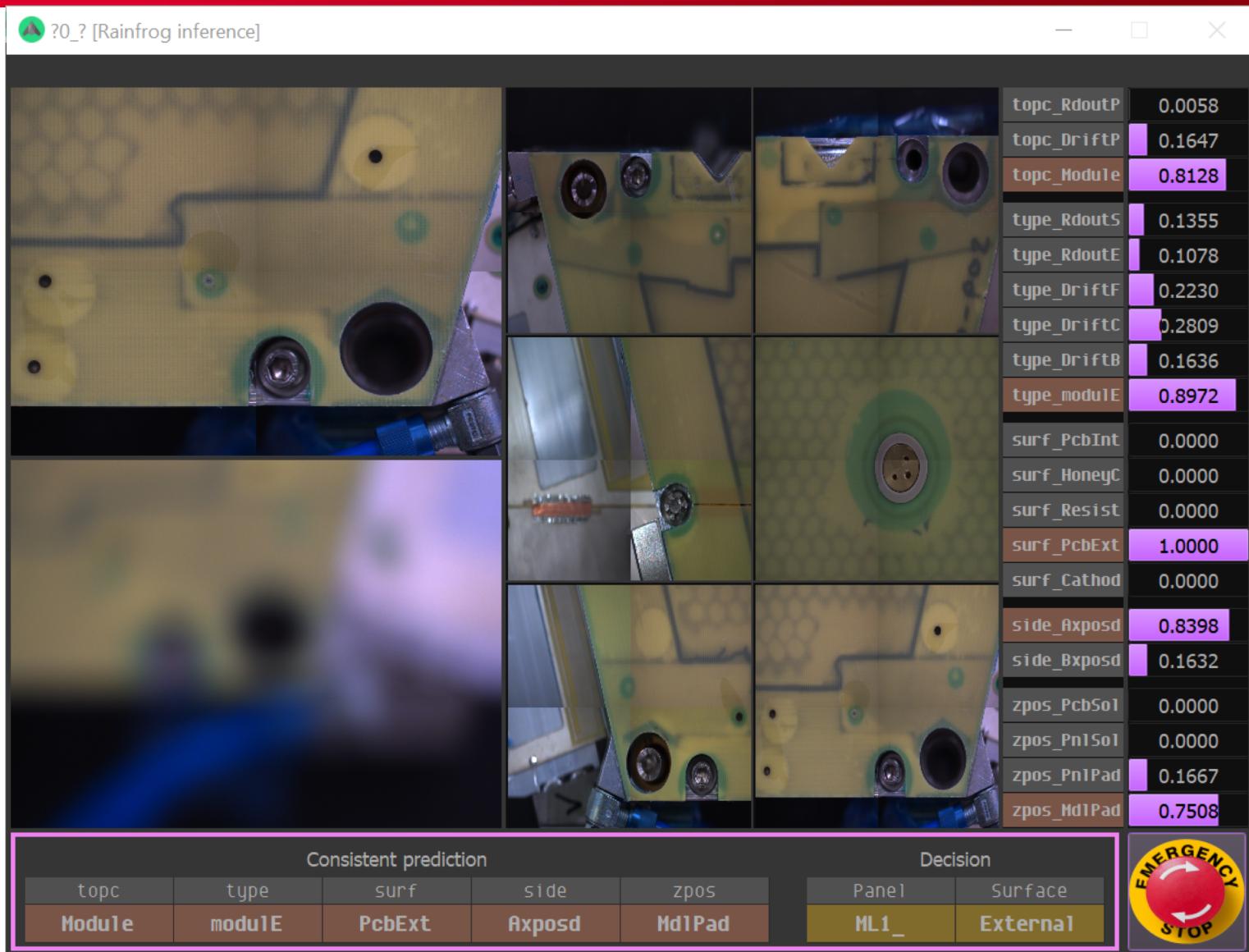
Input	Operator	<i>t</i>	<i>c</i>	<i>n</i>	<i>s</i>
$224^2 \times 3$	conv2d	-	32	1	2
$112^2 \times 32$	bottleneck	1	16	1	1
$112^2 \times 16$	bottleneck	6	24	2	2
$56^2 \times 24$	bottleneck	6	32	3	2
$28^2 \times 32$	bottleneck	6	64	4	2
$14^2 \times 64$	bottleneck	6	96	3	1
$14^2 \times 96$	bottleneck	6	160	3	2
$7^2 \times 160$	bottleneck	6	320	1	1
$7^2 \times 320$	conv2d 1x1	-	1280	1	1
$7^2 \times 1280$	avgpool 7x7	-	-	1	-
$1 \times 1 \times 1280$	conv2d 1x1	-	k	-	-

*MobileNetV2 feature extractor  
architecture (21 layers)*



*Training and validation  
loss*

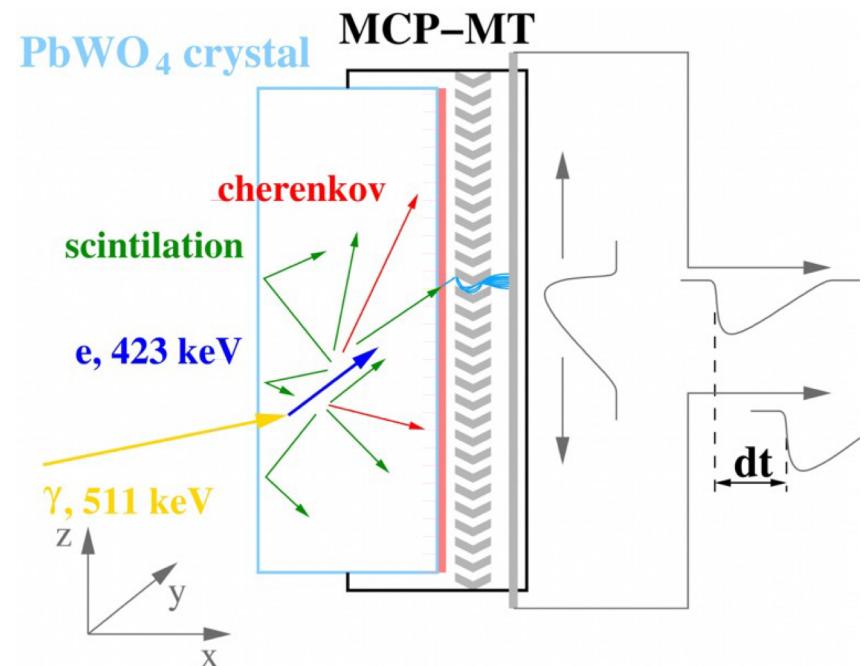
# FIELD TEST: A CORRECT INFERENCE, M22 MODULE



## a $\gamma$ -detector for PET

Spec. list :

- Accurate interaction positioning, 3D : Down to few mm<sup>3</sup>
- Accurate Timing: Few tens of ps (1s)
- Efficiency: Highest can be - 9 mm att. Length @ 511 keV
- Plug and Play: Once developed



# SIMULATED SHALLOW EVENT

## Geant 4 Simulation

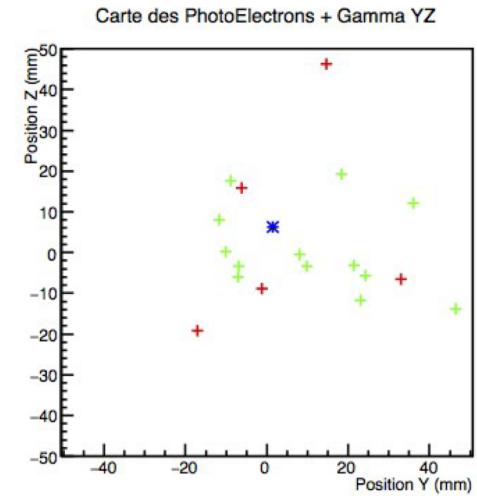
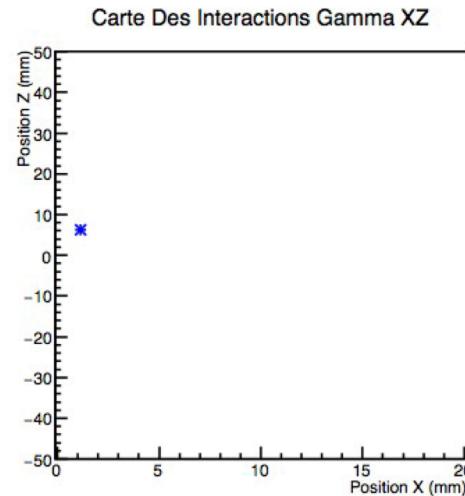
Using Multivariate analysis techniques from MC simulation :

- Positioning accuracy of  $\gamma$  interaction : 4 mm $^3$  ==> ? 1 mm $^3$  ?
- Timing accuracy : few 10 ps (1s)

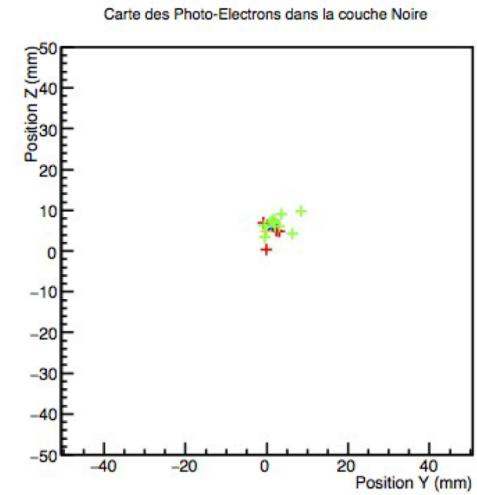
After decorrelation from Depth of Interaction in crystal

Will be limited by MCP-PMT, single g TTS

- Energy resolution: ~ 15 -20 % (FWHM) @ 511 keV

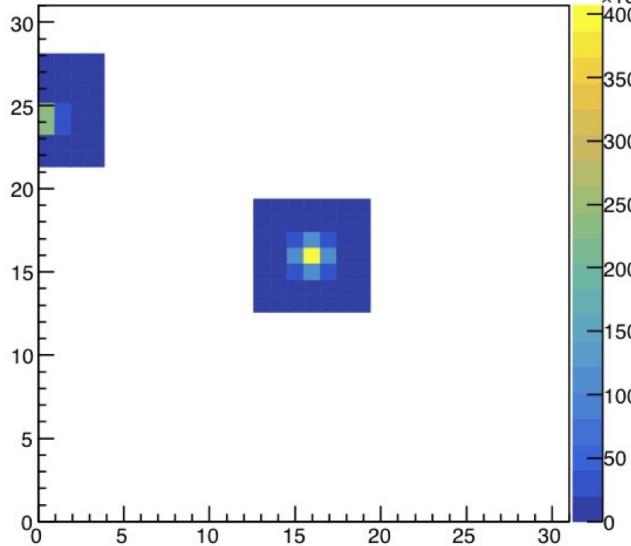


**Energie deposee: 511 keV**  
**Nbre Interactions Gamma: 2**  
**Nbre Photons Optiques: 152**  
**Nbre Photo-Electron: 18**  
**Nbre Photo-Electron Back: 19**

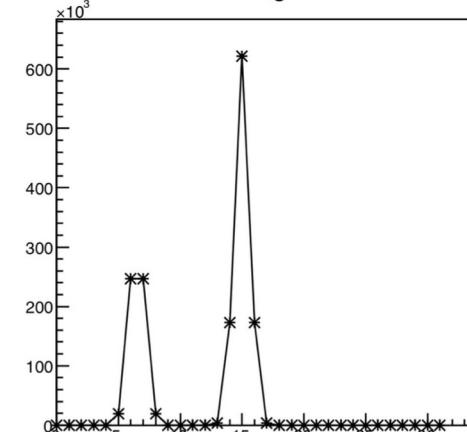


# PHOTO-DETECTOR SIMULATION

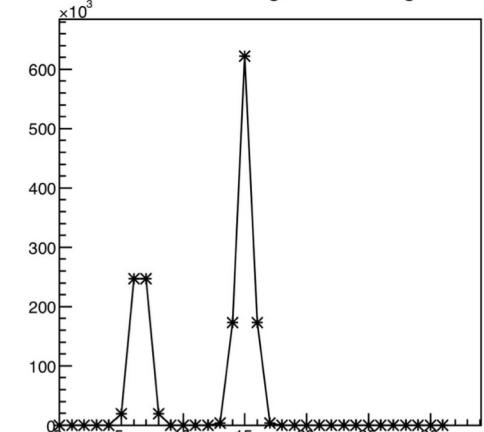
Map of Charges collected on anodes



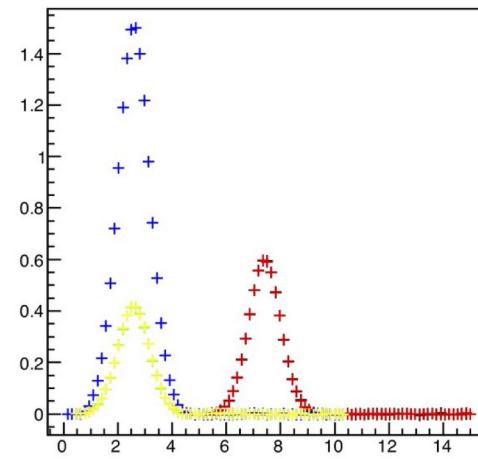
Collected Charge Profile Left



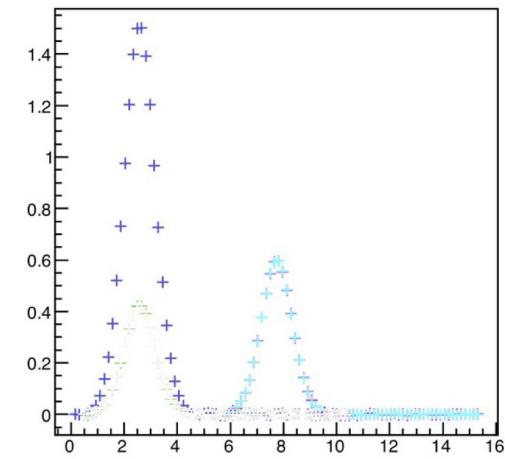
Collected Charge Profile Right



Pulses Profile Left

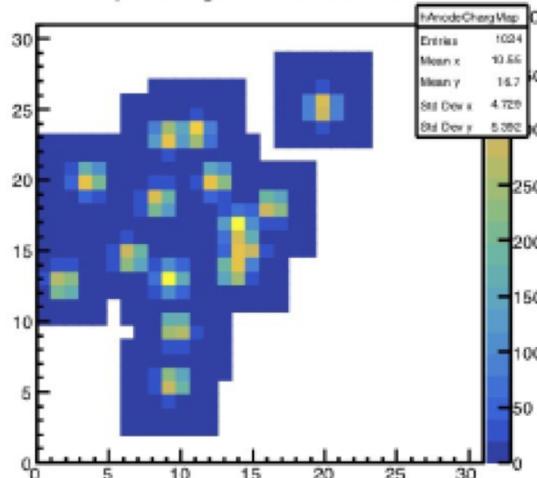


Pulses Profile Right

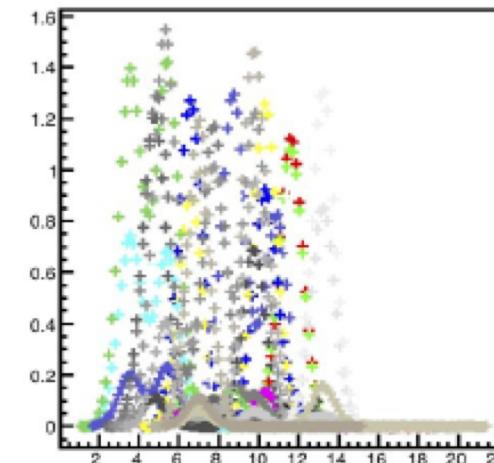


# EVENT VISUALISATION

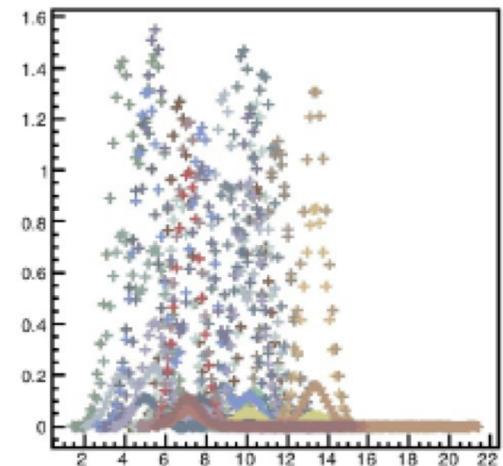
Map of Charges collected on anodes



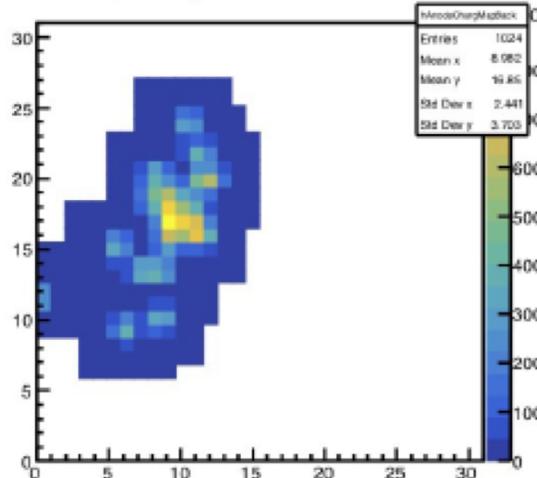
SAMPIC Pulse Left



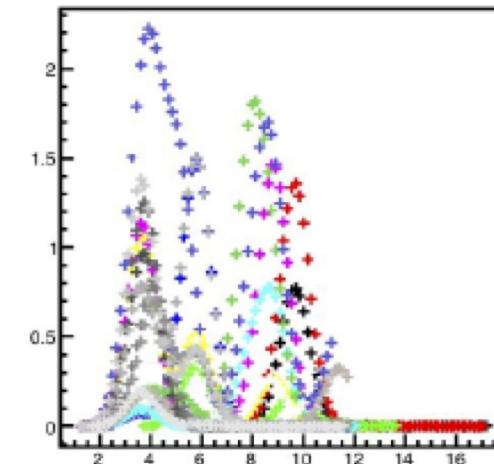
SAMPIC Pulse Right



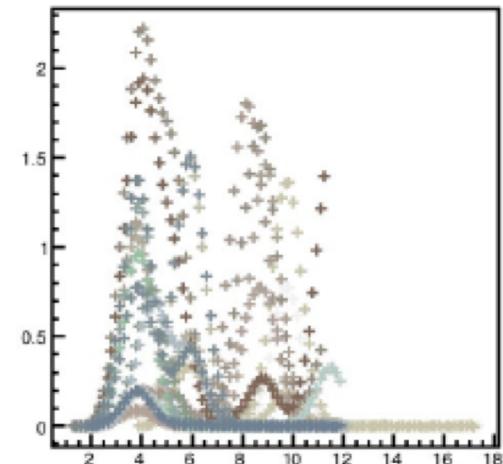
Map of Charges collected on Back anodes



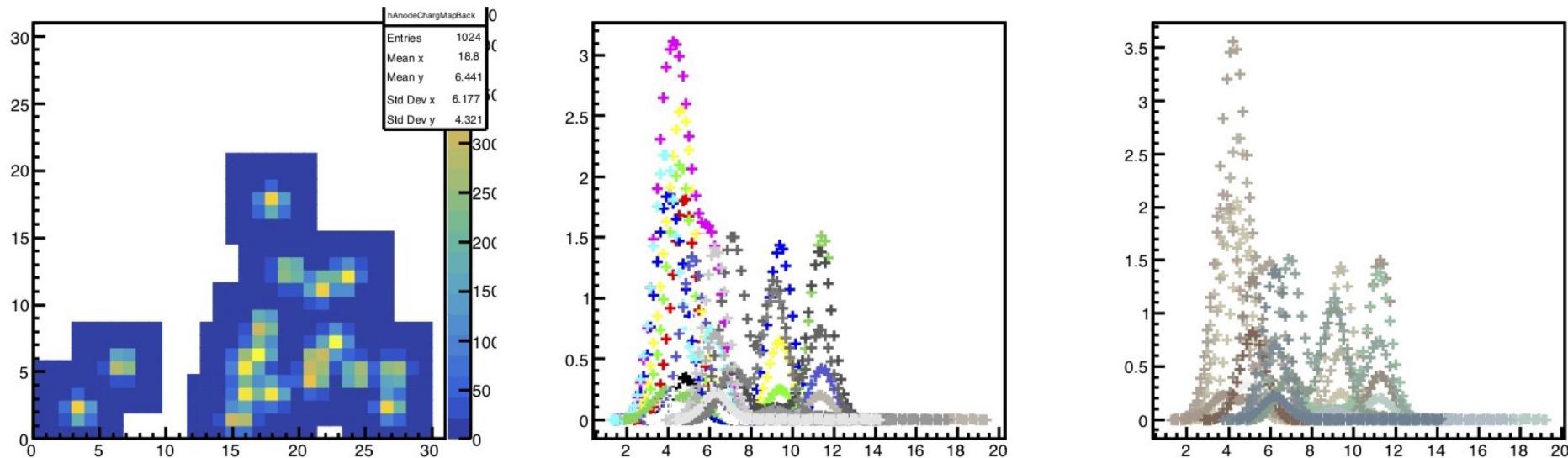
SAMPIC Back Pulse Left



SAMPIC Back Pulse Right



# EVENTS RECONSTRUCTION



- From an registered waveform list
  - Information is complex and intricated
- To Gamma interaction properties
  - 3D positionning in Crystal.
  - Time of interactions.
  - Gamma interaction multiplicity
    - weighted by a probability

## Using Geant4/Gate simulated data

Calibrated on lab measurements

## Specific methods and algorithms

- Machine learning methods
- Convolutional neural networks adapted to the detector spatio-temporal data.
- May require online pre-processing
  - for signals dimension reduction

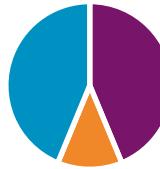
## Validation of Algorithms on detector data

# ÉTHIQUE

- Collège éthique
  - Enjeux moraux dans les processus de décisions automatiques
  - C. Gamrat, C. Castille, M. Duranton, V. Gautard, J.M. Martinez, F. Schindele, V. van Wassenhove
- Larsim
  - Ethique de l'IA
  - Participation a la commission CERNA (commission de réflexion sur l'Ethique de la Recherche en sciences et technologies du Numérique d'Allistene (Alliance des sciences et technologies du numérique))
  - Formation sur l'éthique de l'IA (INRIA, PSL et ENSTA)
  - Parution de livres : les robots et le mal, Alexei Grinbaum

To subscribe : mail at  
valerie.gautard@cea.fr

## About 100 persons from

- ▶ LSCE, CNRGH, IRFU
- ▶ DES, DRT
- ▶ Paris-Sud university, etc.
  
- ▶ Roadmap:
  - Pedagogic presentation
  - Seminars
  


■ Biology
■ Physics
■ Chemistry
■ Mathematic, Machine learning

- Research project
- Training

- **Training**

- ▶ For internship persons from LSCE and IRFU
- ▶ Machine Learning technics for cahotics system

- **PhD thesis**

- ▶ Artificial Intelligence for a gamma-detector for high resolution PET imaging  
*(dir. V. Sharyy, IRFU)*

- **Workshop**

- **Web site:**

<https://indico.in2p3.fr/event/17858/page/1967-intheart>

# CONCLUSION

- Wide variety of applications at IRFU
- Difficulty in choosing a priori methods
- Quite targeted applications that give precise results
- Need for expertise before beginning on a project
- Interest in comparing experiences and thus having an active community (IntheArt forum)
- It would be interesting to also work on the methods themselves for example in the validation of results
- Users must be made aware of ethics from the start of a project

Thanks a lot to :

V. Calvelli - N. Cherrier - B. Dalena - G. Daniel - A. Grinbaum -  
F. Lanusse - M. Mur - D. Yvon

# MERCI !



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Institut de recherche  
sur les lois fondamentales de l'Univers  
Service