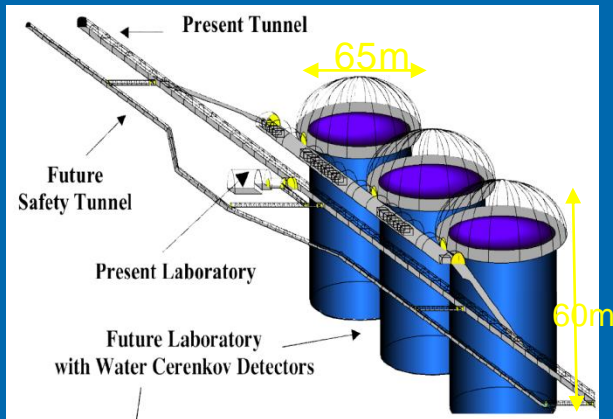
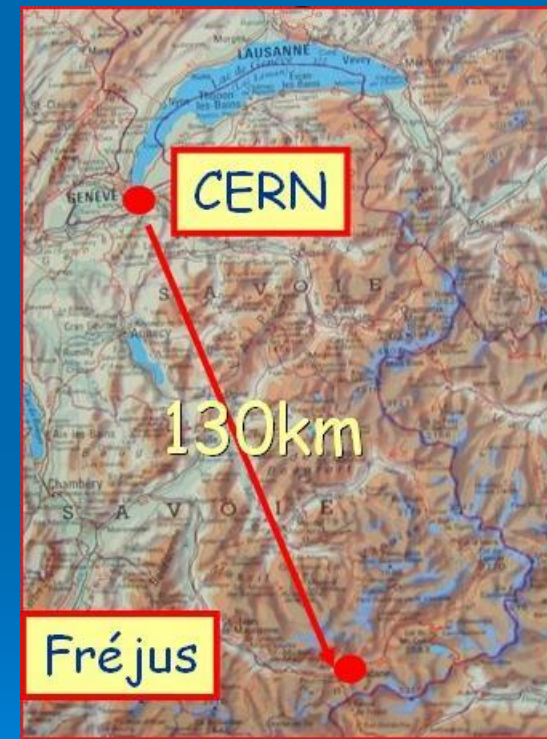
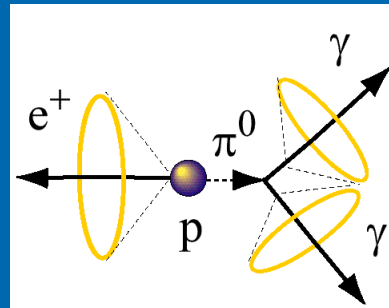


MEMPHYS SIMULATION

N. Vassilopoulos / APC



each shaft $\approx 4 \times \text{SK} = 215 \text{ Kton H}_2\text{O}$

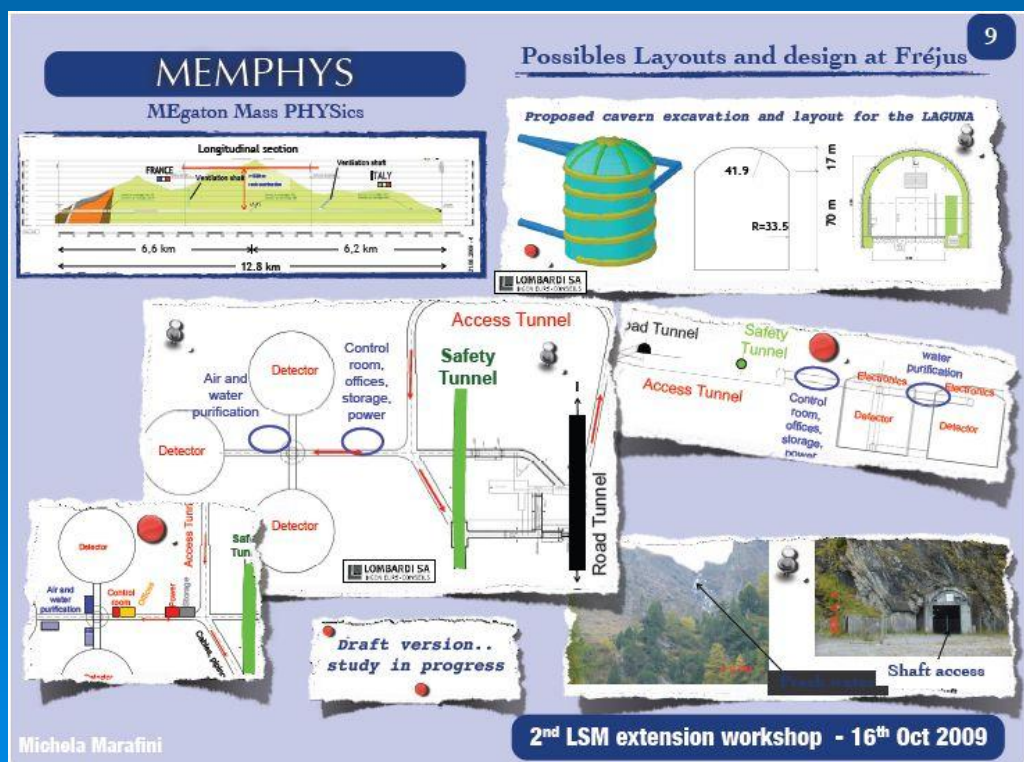


ASTROPARTICULE ET COSMOLOGIE

emphasis: software developments

MEMPHYS: Underground Laboratory and Detector

- underground water Cherenkov at Fréjus (Alps), Modane, France
- total fiducial mass: up to 400 kton: 3 x 60mX65 modules
 - size limited by light attenuation length ($\lambda \sim 80\text{m}$) and pressure on PMTs
 - readout: $\sim 3 \times 81\text{k}$ 12" PMTs, 30% cover (# PEs = 40% cover with 20" PMTs)
- PMT R&D + detailed study on excavation existing & ongoing



Laboratoire
Souterrain de
Modane

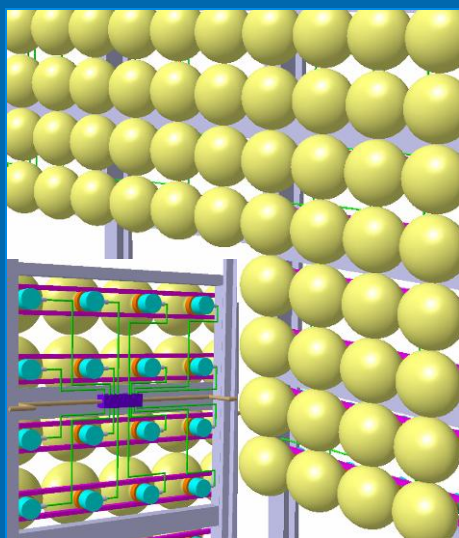
http://2www.apc.univ-paris7.fr/APC_CS/Experiences/MEMPHYS/
arXiv: hep-ex/0607026

Contacts: J.E. Campagne and M. Mezzetto

MEMPHYS physics goals

- Proton decay sensitivity:
 - up to 10^{35} yrs in 10y from the "golden" channel: $p \rightarrow e^+ \pi^0$
 - up to 2×10^{34} yrs in 10y from $p \rightarrow K^+ + \text{anti-}\nu$
- SuperNova core collapse:
 - huge statistics from galactic SN \Rightarrow spectral analysis in E,t, flavour \rightarrow access SN collapse mechanism / neutrino oscillation parameters
 - sensitivity up to ~ 1 Mpc
 - possibility of early SN trigger (from event coincidence) up to ~ 5 Mpc
- SuperNova relic neutrinos:
 - observable in few years with significant statistics, according to most of existing models
 - direct measurement of ν emission parameters possible
- and, of course... NEUTRINO BEAMS ! (*watch next talk from Andrea Longhin*)

R&D towards MEMPHYS : PMm2



“Innovative electronics for array of photodetectors used in High Energy Physics and Astroparticles”.

R&D program funded by French national agency for research (LAL, IPNO, LAPP and Photenis) (2007-2010)

Basic concept: very large photodetection surface → macropixels of PMTs connected to an autonomous front-end electronics.

Replace large PMTs (20") by groups of 16 smaller ones ($1\frac{1}{2}$ ", 8") with central ASIC :

- Independent channels
- charge and time measurement
- water-tight, common High Voltage
- Only one wire out (DATA + VCC)

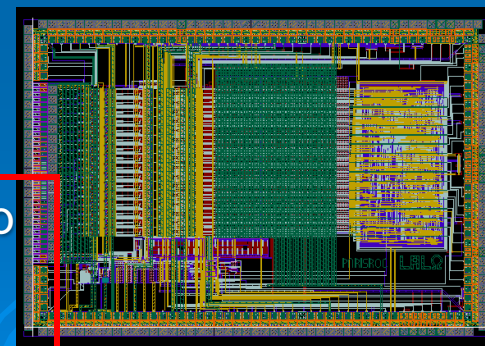
Latest News & detailed description of the R&D: read J.E Campagne's Talk at NNN09

I. studies on 12" PMTs design

- parameter correlation
 - potting
 - pressure resistance
- (collaboration with BNL since NNN07)

II. PARISROC readout chip

- complete front-end chip with 16 channels
- testboard now in layout, soon available



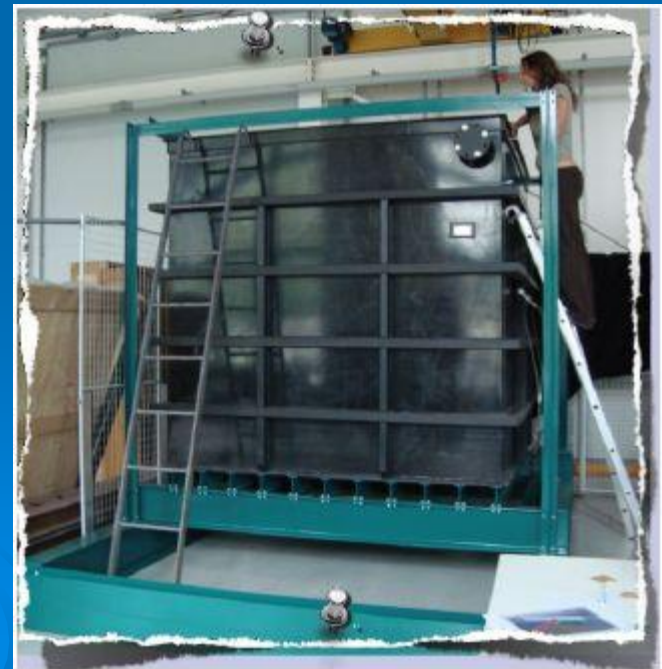
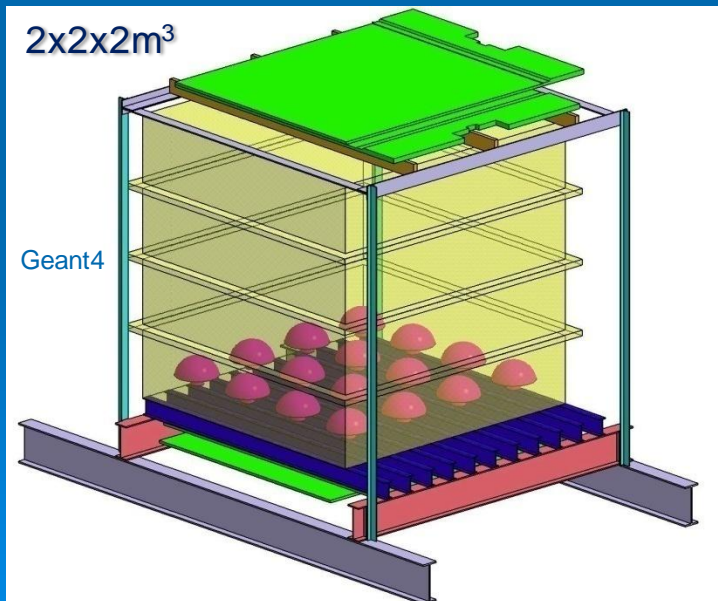
R&D towards MEMPHYS: MEMPHYNO

Goals:

- full test of electronics and acquisition chain with actual physics events
- trigger threshold studies
- self-trigger mode
- track reconstruction performances
- Gd doping: feasibility and performance (if studies still needed...)

Test bench for
photodetection
solutions for large
detectors

-> latest news on MEMPHYNO prototype
read M. Marafini's talk from the 2nd LSM-
EXTENSION WORKSHOP

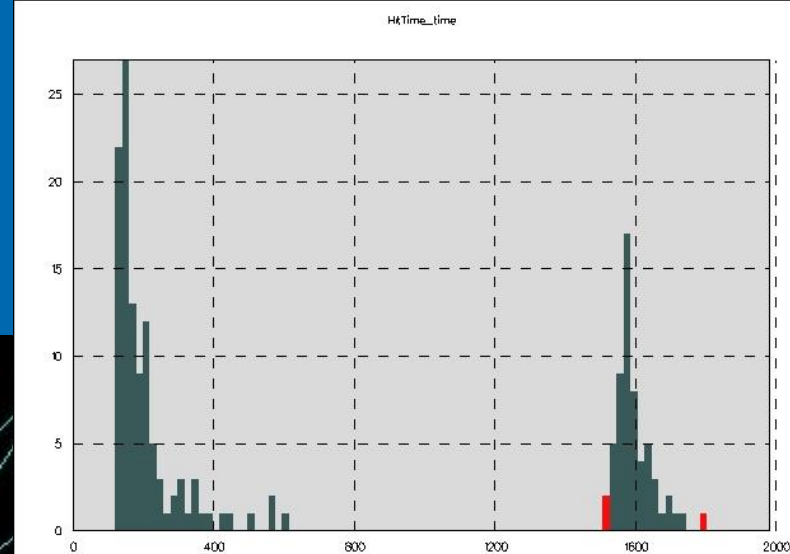
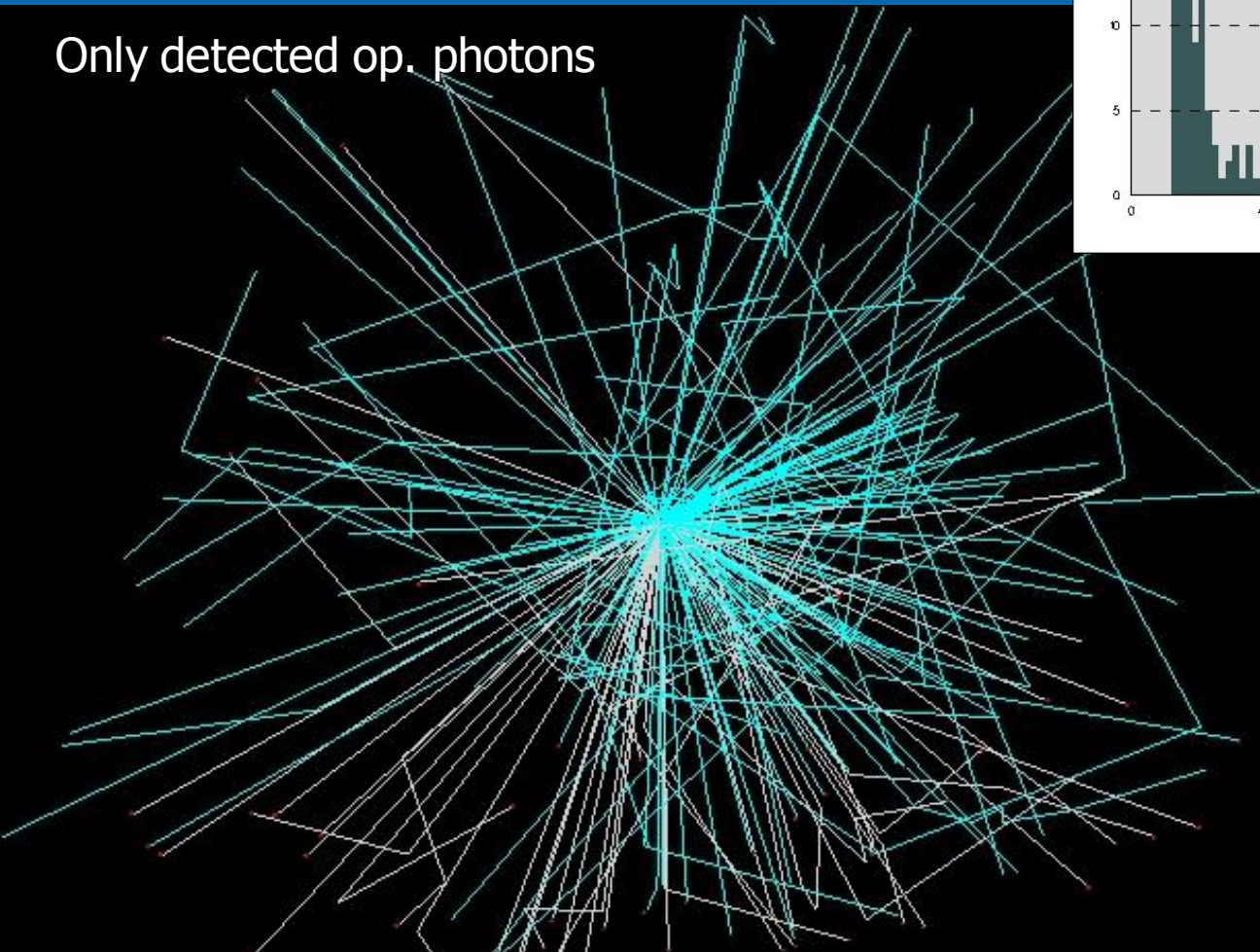


MEMPHYS: MC Present Status

- Event Generator:
 - **NUANCE** for ν beam, ν Atmospheric & Proton Decay
- MEMPHYS Simulation (M. Fechner , J.E. Campagne, N. Vassilopoulos) :
 - Interface with the **OpenScientist v16r0** framework (G. Barrand/LAL) provided using **distribution kits** including **Geant4 & CLHEP & AIDA-IO implementation to Rio** (also HDF5, XML)
 - **3 modes of running in the same framework:**
 - Interactive Viewing, Batch processing, AIDA_ROOT analysis
 - event info from MC, primary + non-Optical photons track infos, track selection, modular detector geometry, etc. ntuples' storage
 - e.g. hits: each PM maintain a list of arrival time of optical photons detected
- Analyses
 - **ROOT- cint**
 - Solo **C++** with **ROOT, AIDA libraries**

$$\nu_{\mu} \rightarrow \mu^{-} \rightarrow e^{-}$$

Only detected op. photons



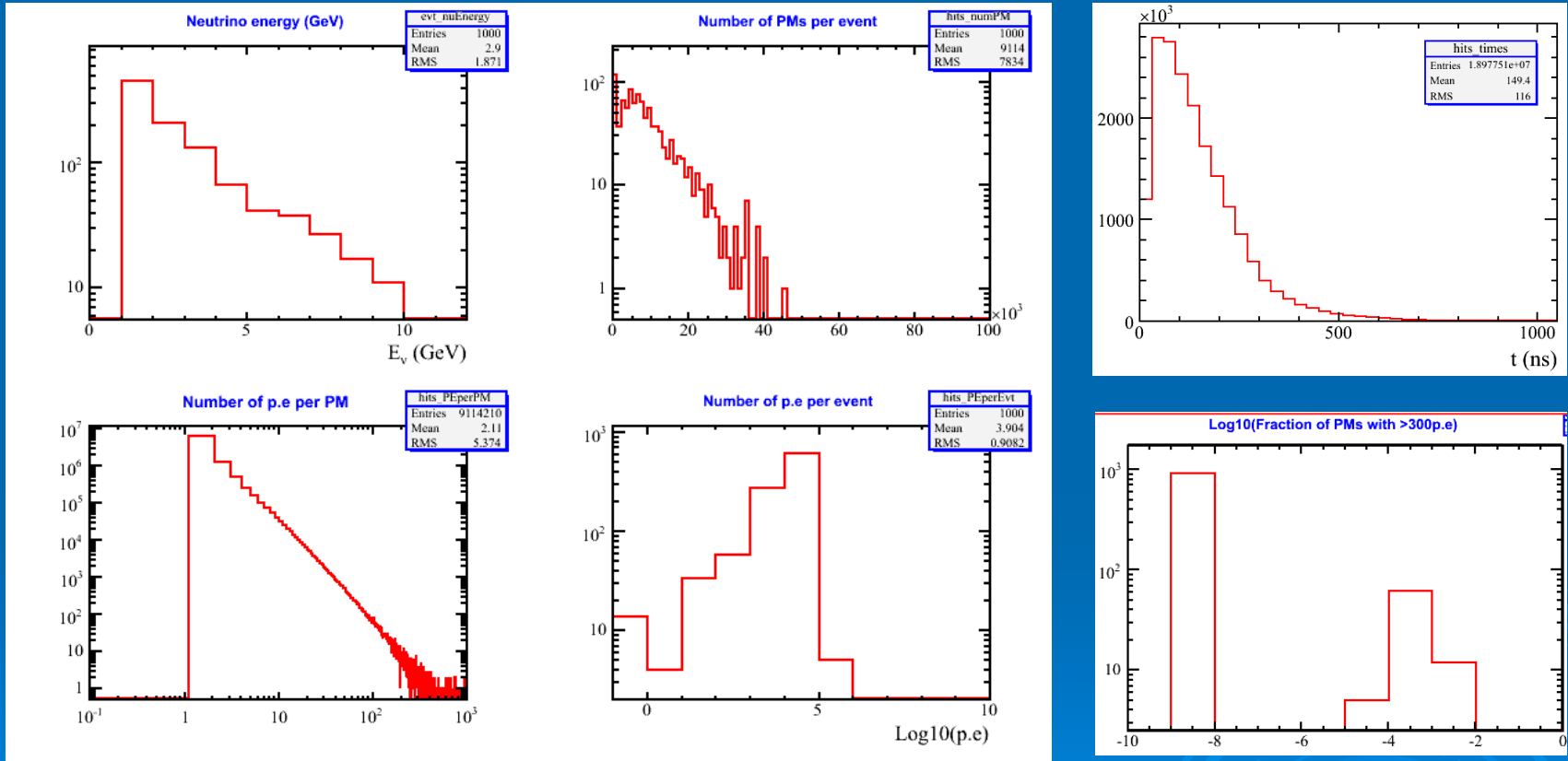
1600ns

Interactive
histogram to identify
the e Michel optical
photons...

transparency by J. E. Campagne

MEMPHYS v7

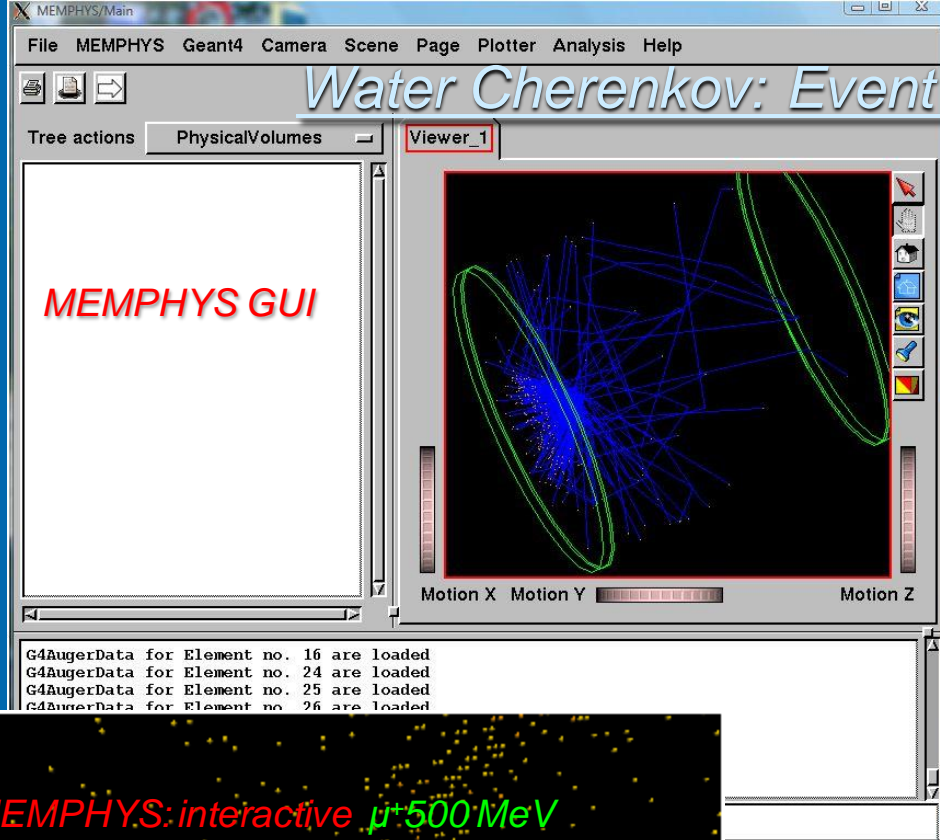
ν atmospheric (1-10GeV)



transparency by J. E. Campagne

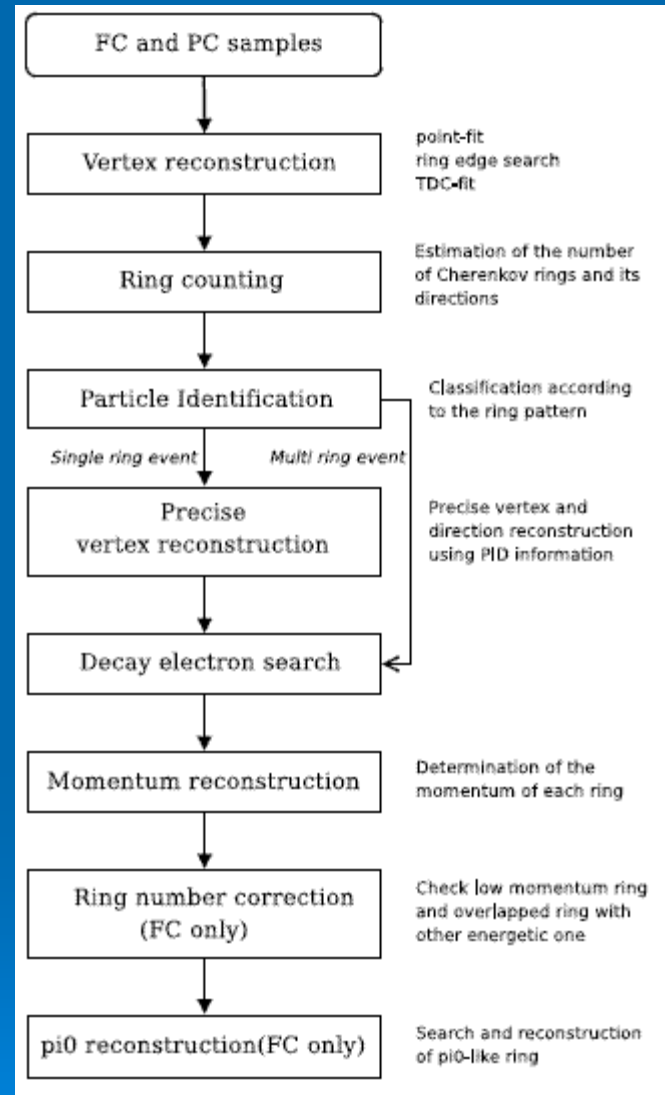
MEMPHYS v7

Water Cherenkov: Event Reconstruction



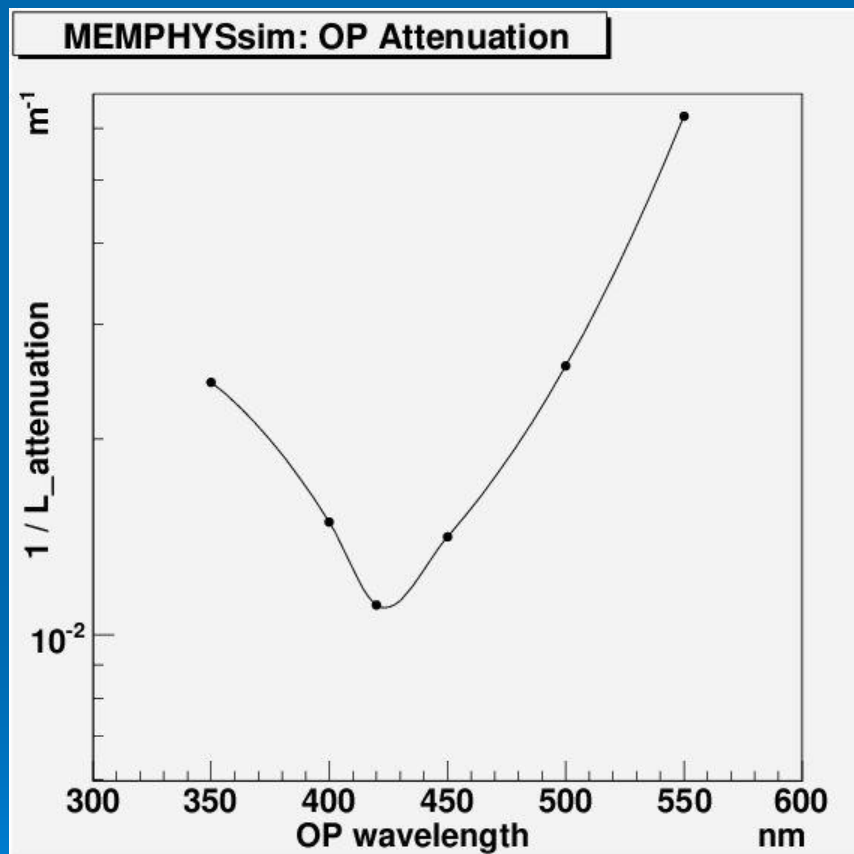
MEMPHYS: interactive μ^+ 500 MeV

GDR09/Strasbourg-N.V.

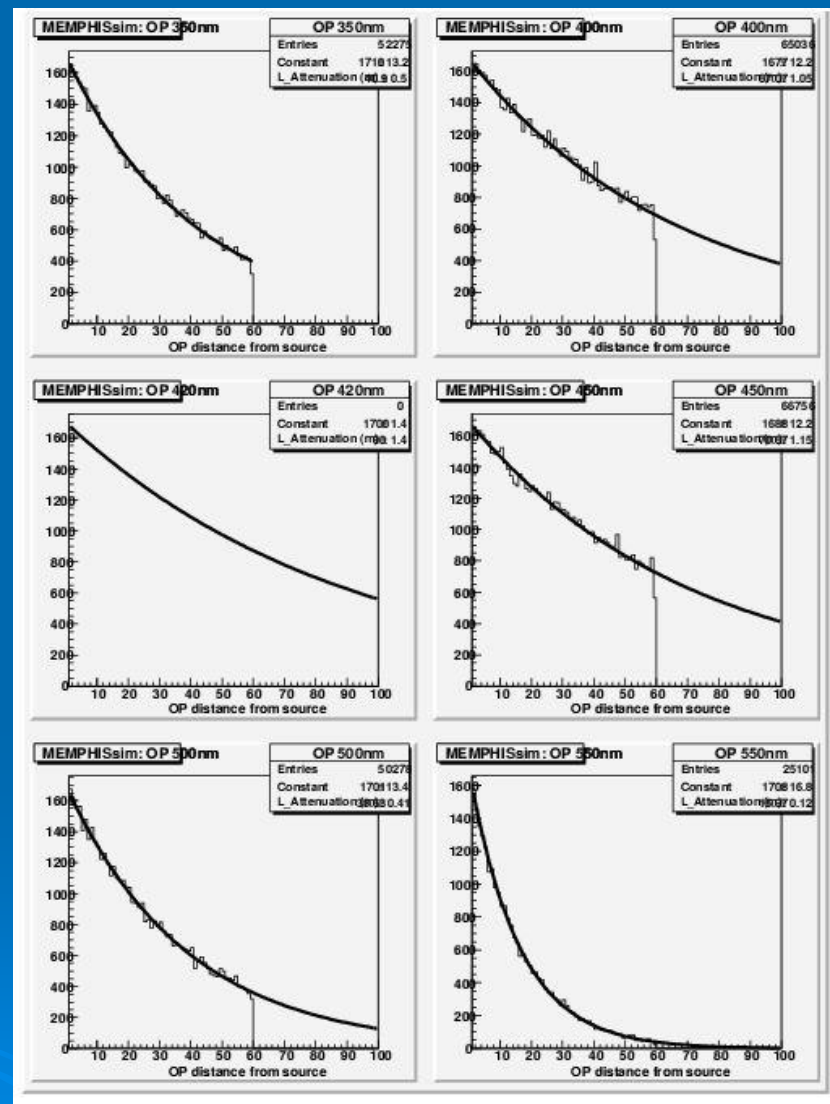


Attenuation length studies

- correct modelling of light propagation



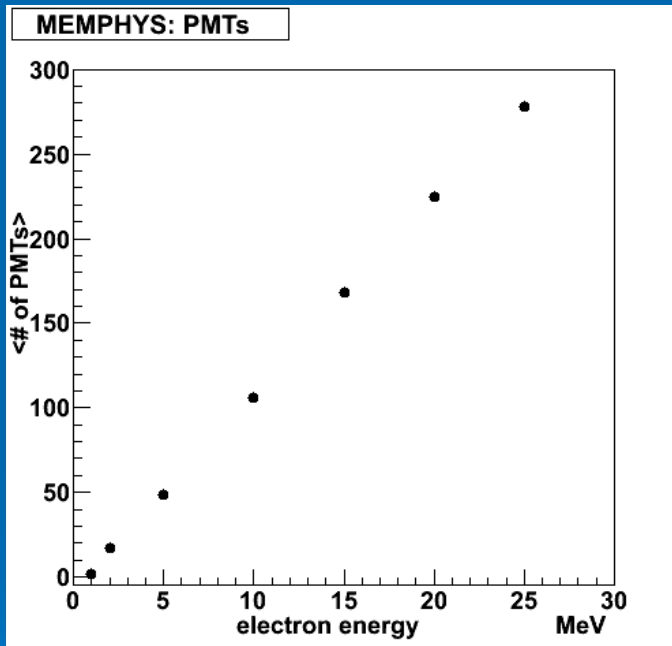
attenuation length in water as a function of the wavelength in the Gean4 simulation



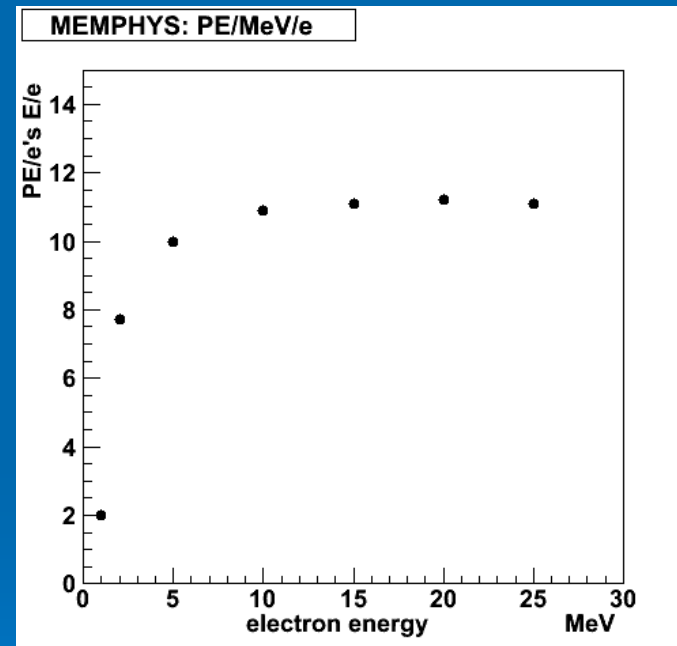
✓ looks comparable to SK data

MEMPHYS Single ring studies, electrons

- single e- events from 1 to 25 MeV: PMTs and PE infos



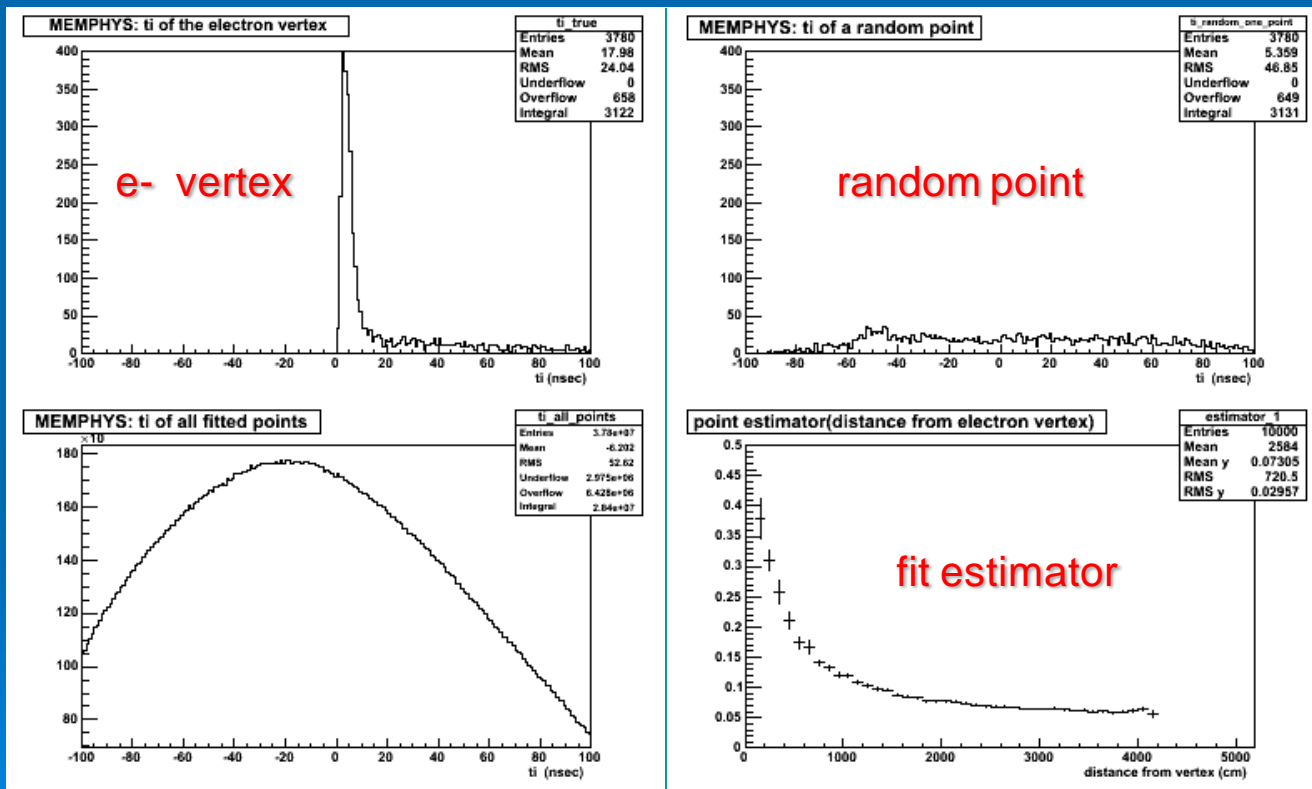
Number of PMTs with at least one photoelectron as a function of electron energy



Number of detected photoelectrons per MeV as a function of electron energy

Single rings: electrons primary vertex fit

- pick up a 400 MeV electron, assume point like track length
- primary vertex fit based only on each PMT's timing info: $t_{i \text{ PMT}} = t_i + \text{TOF}_i \Rightarrow t_i = t_{i \text{ PMT}} - \text{TOF}_i$, where $\text{TOF}_i = (n / c) \times D$, D = distance between each PMT and grid's coordinates
- maximize estimator E a la SK to find the true vertex of electron :

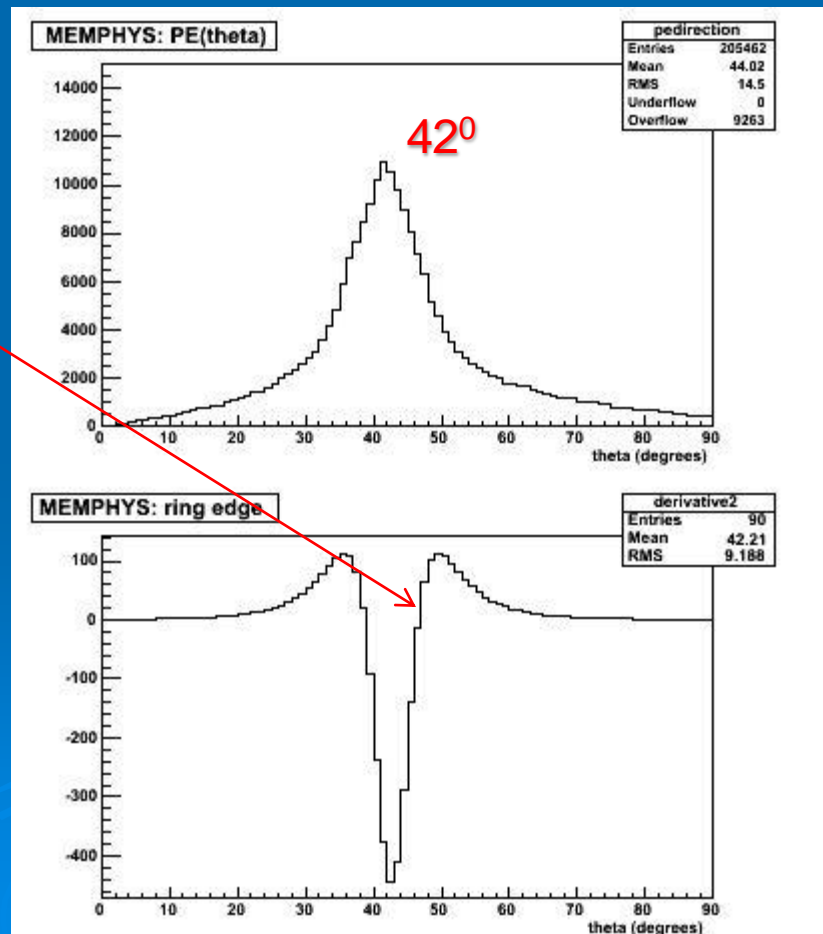


single rings: particle direction, outer ring edge

- keep the 400 MeV e-
- calculate roughly the direction:
- θ_{edge}

$$\vec{d}_0 = \sum_i q_i \times \frac{\vec{P}_i - \vec{O}_0}{|\vec{P}_i - \vec{O}_0|}$$

- $\theta_{\text{edge}} > \theta_{\text{peak}}$
- $d^2\text{PE}(\theta) / d^2(\theta) = 0$

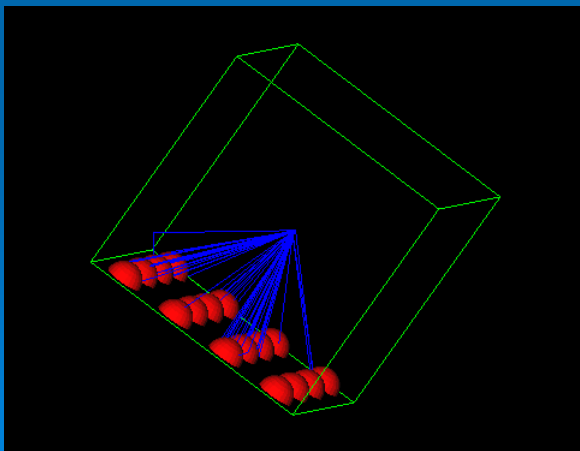


MEMPHYS: Simulation Studies for the small scale Prototype

MEMPHYNO

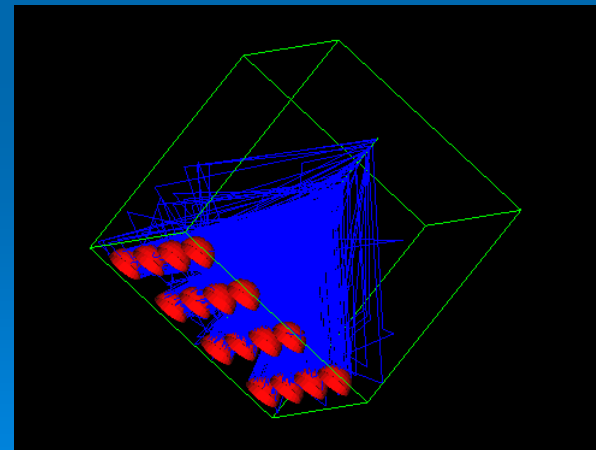
Alessandra Tonazzo, Nikos Vassilopoulos / APC-PARIS

- tests with radioactive sources (monoenergetic, point-like) and cosmic muons (direction selected with hodoscope) on surface
- measure background level @ underground site
- MEMPHYS simulation & visualization code
- 4x4 12in PMTs = ~35% coverage (for one side)



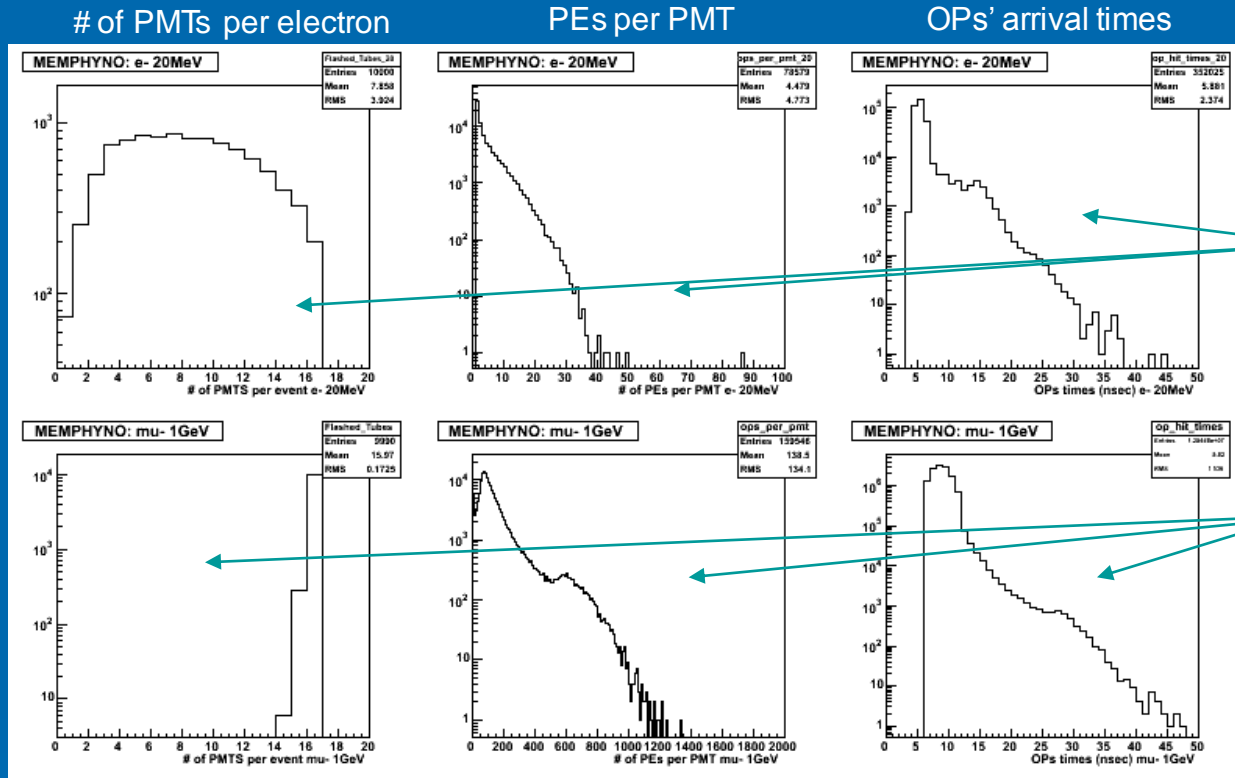
electron 10 MeV, $p_z / p = -1$, vtx : centre

MEMPHYS GUI for
MEMPHYNO and
detected optical
photons



muon 1 GeV, $p_z / p = -1$, vtx : top centre

MEMPHYS: MEMPHYNO e^- , μ^- studies



| | | | | | | | |
|----------------|-----|-----|-----|------|------|------|------|
| e- E (MeV) | 1 | 2.5 | 5 | 10 | 15 | 20 | 25 |
| PEs / MeV / el | 0.2 | 1.1 | 1.5 | 1.7 | 1.8 | 1.8 | 1.8 |
| X 6 (sides) | 1.2 | 6.6 | 9 | 10.2 | 10.8 | 10.8 | 10.8 |
| MEMPHYS | 2 | 7.7 | 10 | 10.9 | 11.1 | 11.2 | 11.1 |

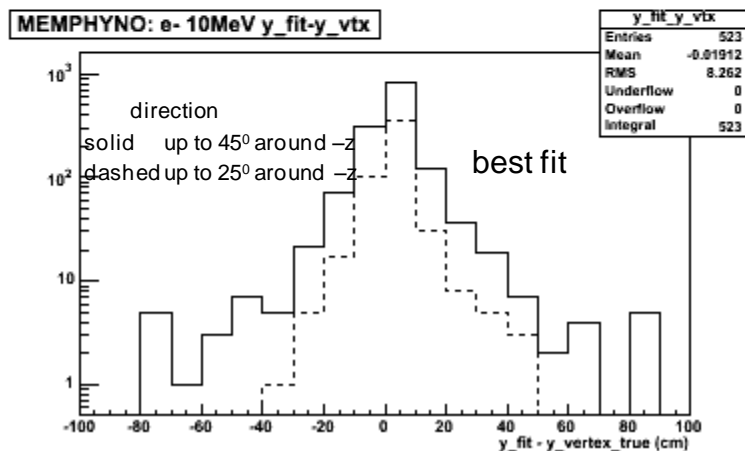
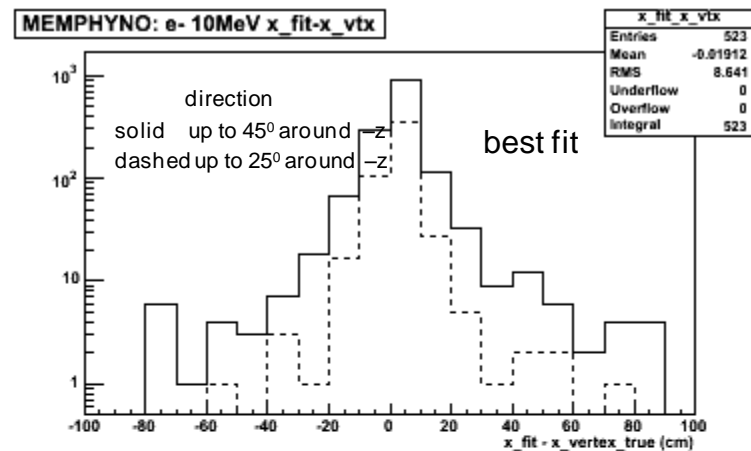
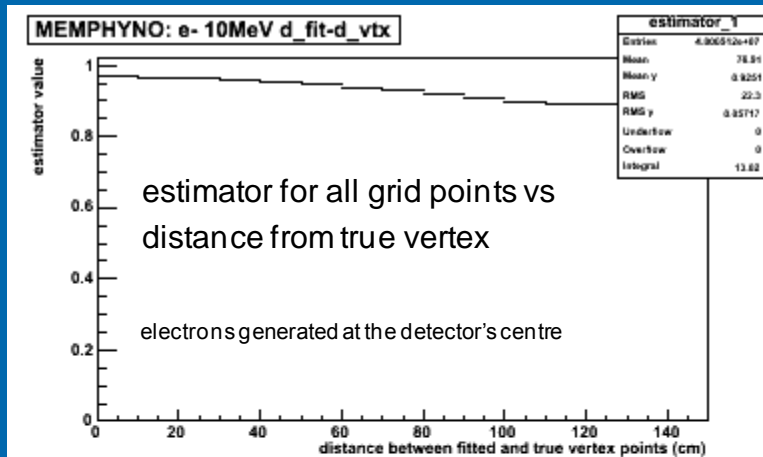
Table: MEMPHYSNO's PEs per MeV per electron

10k per energy electrons generated at the detector's centre with random direction

MEMPHYNO

electrons 10 MeV : vertex finding

- primary vertex fit based only on each PMT's timing info: $t_{i \text{ PMT}} = t_i + \text{TOF}_i \Rightarrow t_i = t_{i \text{ PMT}} - \text{TOF}_i$, where $\text{TOF}_i = (n / c) \times D$, D = distance between each PMT and grid's coordinates
- maximize estimator E a la SK to find the true vertex of electron :



for the primary fit :

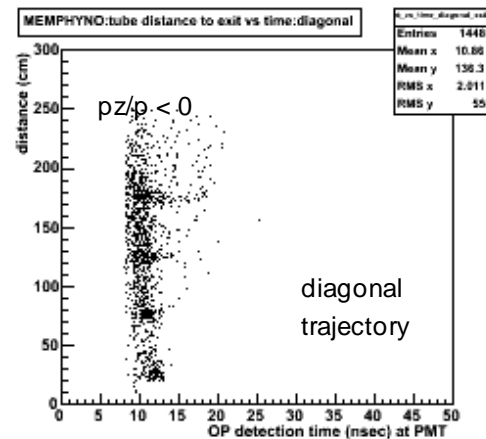
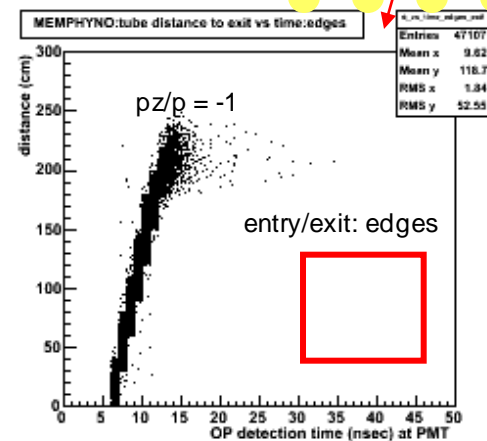
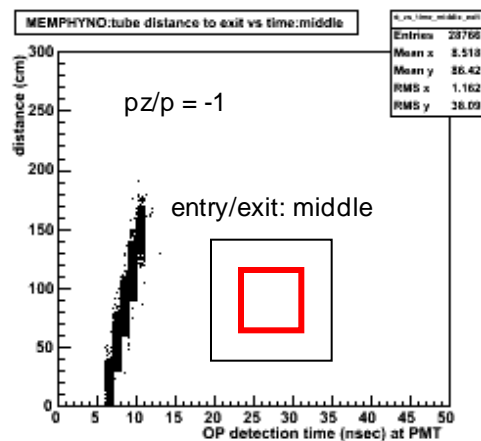
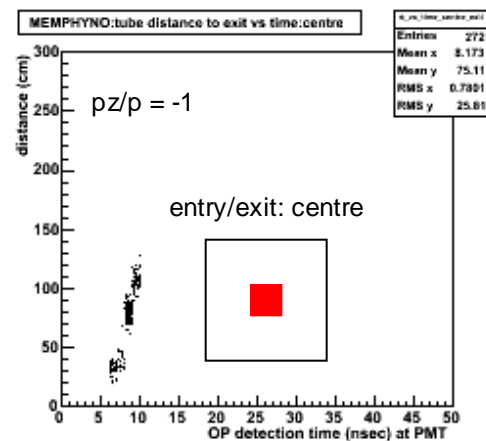
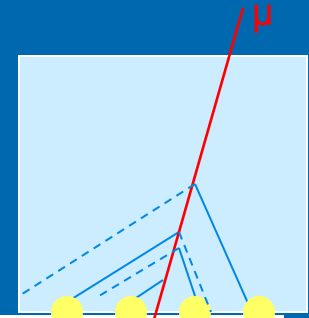
1. grid analysis (5cm spacing) in MEMPHYNO
2. good resolution for downwards electrons in x-y plane where is the PMTs' module (shown)
3. resolution becomes worse as pz/p increases due to one PMTs' module : best for perpendicular electrons

MEMPHYNO

muons 1 GeV

light propagation effect of OPs :

- check correlation of PMT time with distance between muon's exit point and detection PMT's coordinates



- pz/p = -1 : later produced OPs are detected first
- pz/p < 0 : relation not clean

MEMPHYS Simulation

- Always on going

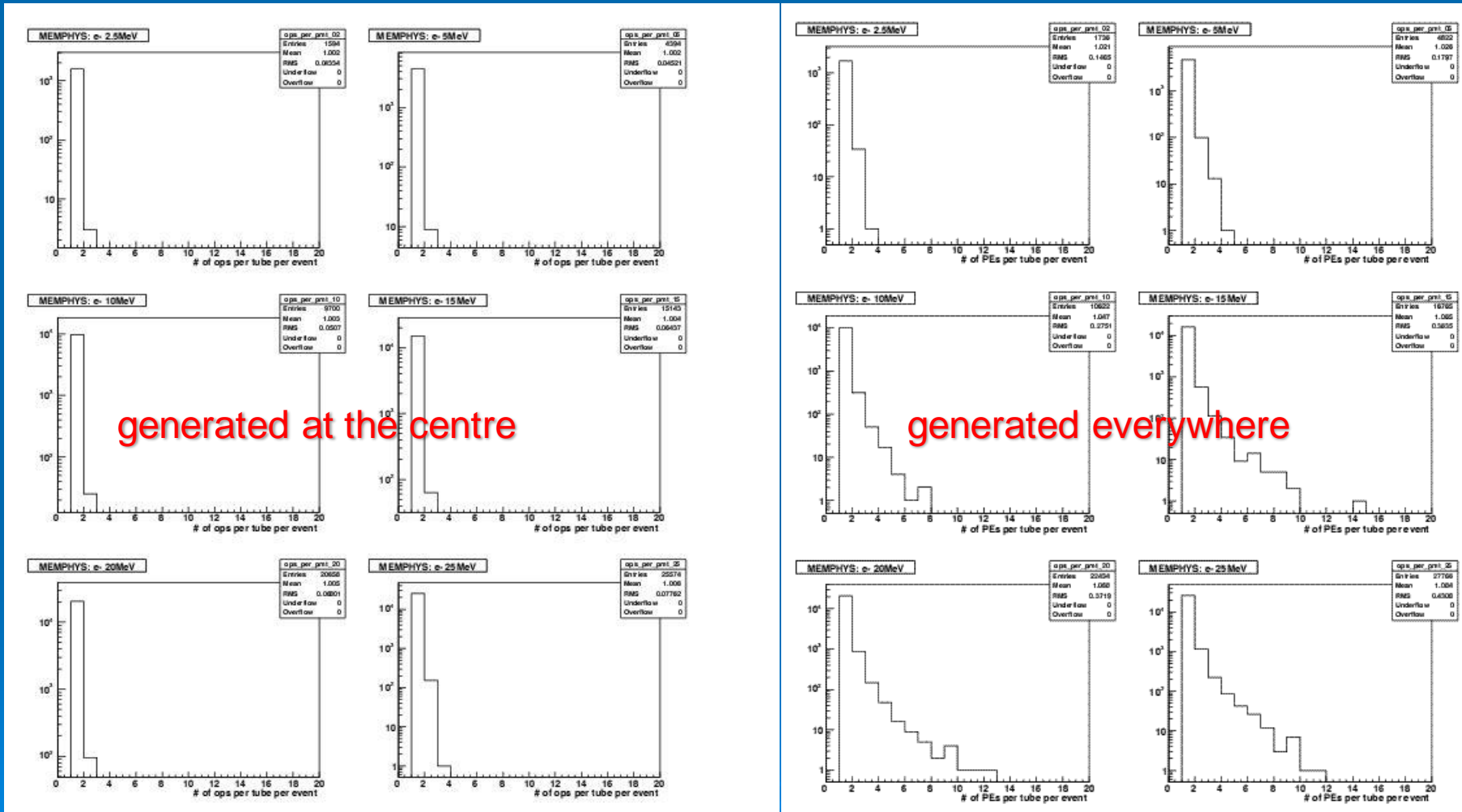
next steps:

- vertex fit considering the track's length
- ring separation
- particle identification

THANKS

MEMPHYS Single ring studies, electrons

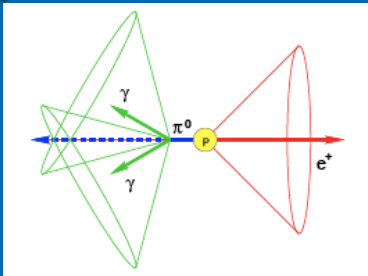
- single e- events from 1 to 25 MeV: PEs per PMT



Distribution of the number of PEs in each PMT for electrons

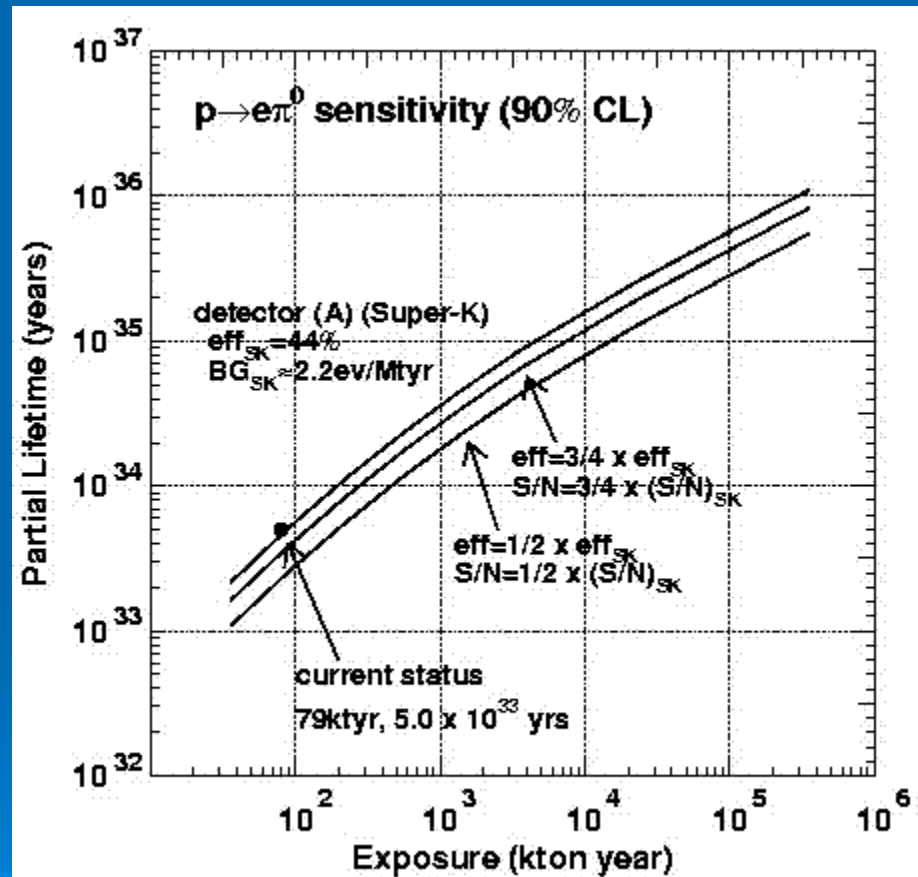
MEMPHYS: Proton Decay, $p \rightarrow e^+ \pi^0$

- search for $p \rightarrow e^+ \pi^0$ (3 showering event), efficiency $\sim 43\%$



- predictions: $\tau_p \sim 10^{34} - 10^{36}$ yrs
- bdg: $\text{atm-} \bar{\nu}_e + N \rightarrow e^+ + N' + \pi^0$
- $\tau_p > 1.6 \times 10^{33}$ yrs by SK
- $\tau_p > 1.6 \times 10^{35}$ yrs at 90% CL for 5Mt.yr by MEMPHYS

MEMPHYS coverage 30% with 12" PMTs is equivalent to coverage 40% with 20" PMTs in terms of #PE.
MEMPHYS relies on the study by UNO, adapting the results

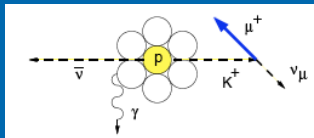


H₂O is best for this channel

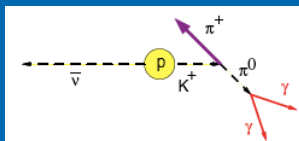
MEMPHYS: Proton Decay, $p \rightarrow K \bar{\nu}$

- search for $p \rightarrow K \bar{\nu}$ (K^+ below cherenkov threshold so the channel is detected via the decay products of K^+)

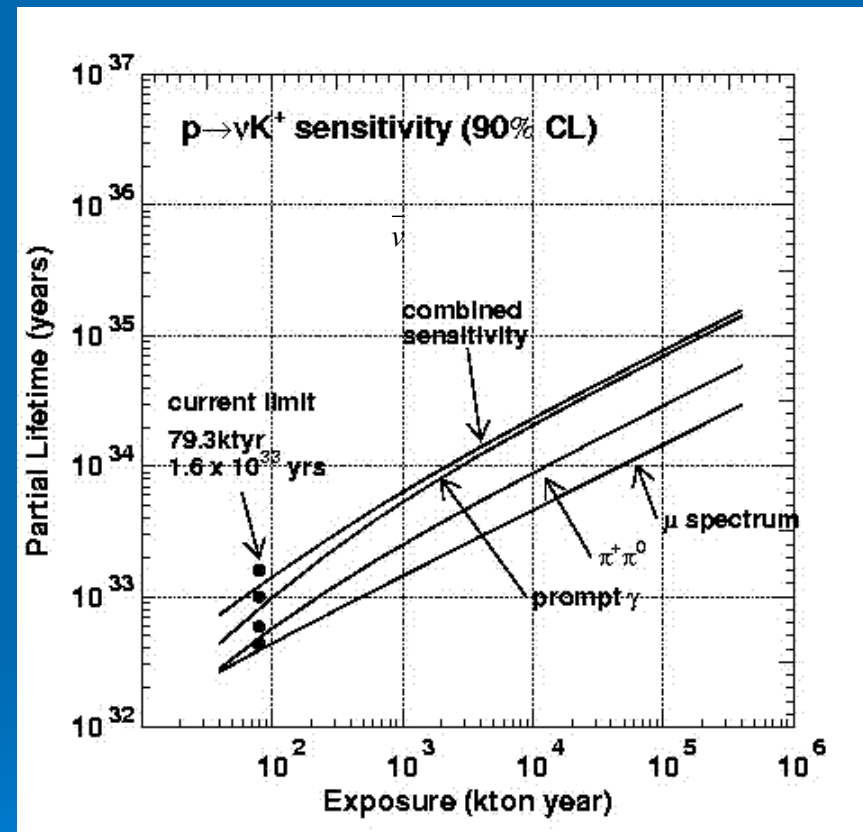
- $K \rightarrow \mu \nu$, monoenergetic μ + 6.3 MeV prompt- γ from capture



- $K \rightarrow \pi^+ \pi^0$ with $\pi^0 \rightarrow \gamma\gamma$



- predictions: $\tau_p \sim 3 \times 10^{33} - 3 \times 10^{34}$ yrs
- $\tau_p > 2.3 \times 10^{33}$ yrs by SK
- $\tau_p \sim 2 \times 10^{34}$ yrs for 5Mt.yr by MEMPHYS

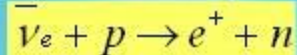


H₂O not as good as LAr, Lsint

MEMPHYS: SN ν in water Čerenkov detectors

Detection of SN neutrinos

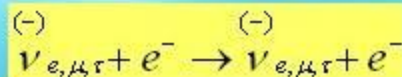
⊙ Inverse-beta (89%)



- Large statistics in detectors with lots of free p
- Good determination of ν time and energy
- Option: add Gd to tag neutron from delayed- γ [see talk by Vagins]



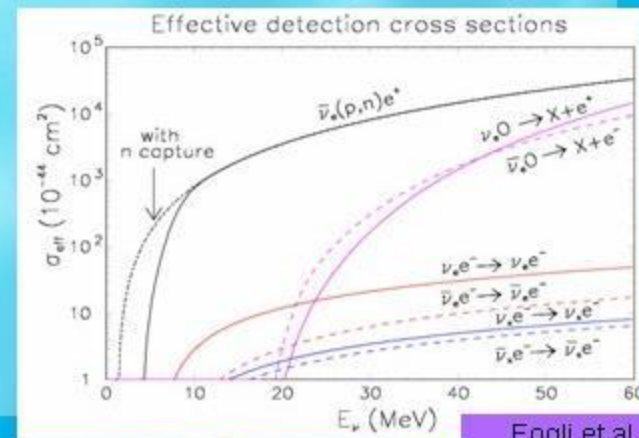
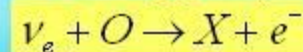
⊙ Elastic scattering (~3%)



- Pointing $\sigma \sim 25^\circ / \sqrt{N}$



⊙ NC on Oxygen (8%)



NOV06 15/09/06

A.Tonazzo - MEMPHYS: non-oscillation physics

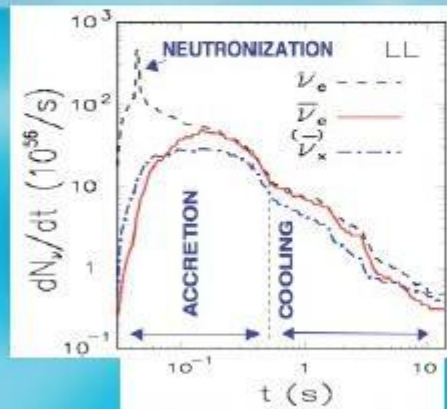
Fogli et al.,
hep-ph/0412046

MEMPHYS: SN and θ_{13} information

SN neutrinos

[see talk by Cardall]

- ⊙ Neutronization burst $\Rightarrow \nu_e$
 - $E \sim 10^{51}$ erg $t \sim 25$ ms
- ⊙ Accretion + K-H cooling $\Rightarrow \nu_{e,\mu,\tau}, \bar{\nu}_{e,\mu,\tau}$
 - $E \sim 10^{53}$ erg $t \sim 10$ s
 - 99% of total explosion energy



Propagation to Earth:

- ⊙ Matter effects $P_{ee}(\theta_{12})$
- ⊙ Level-crossing probability $P_H(E, V(x,t), \Delta m^2, \theta_{13})$
- ⊙ Survival prob. $p = P_{ee} * P_H$

| Scenario | Hierarchy | $\sin^2 \vartheta_{13}$ | ν_e p | $\bar{\nu}_e$ \bar{p} |
|----------|-----------|-------------------------|--------------------------|----------------------------|
| A | Normal | $\gtrsim 10^{-3}$ | 0 | $\cos^2 \vartheta_\odot$ |
| B | Inverted | $\gtrsim 10^{-3}$ | $\sin^2 \vartheta_\odot$ | 0 |
| C | Any | $\lesssim 10^{-5}$ | $\sin^2 \vartheta_\odot$ | $\cos^2 \vartheta_\odot$ |

“Sensitivity to θ_{13} one order of magnitude better than planned terrestrial experiments”
[see for ex. Lunardini-Smirnov hep-ph/0302033]

Hierarchy of interaction strength \Rightarrow

$$\begin{aligned}
 E_0(\nu_e) &\approx 10\text{--}12 \text{ MeV} \\
 E_0(\bar{\nu}_e) &\approx 13\text{--}16 \text{ MeV} \\
 E_0(\nu_x) &\approx 15\text{--}25 \text{ MeV}
 \end{aligned}$$

Fogli et al., hep-ph/0412046
Raffelt et al., astro-ph/0303226

MEMPHYS: SN neutronization burst

SN spectral analyses (4)

⊙ Neutronization burst

Signal: $\nu_e + e^- \rightarrow \nu_e + e^-$

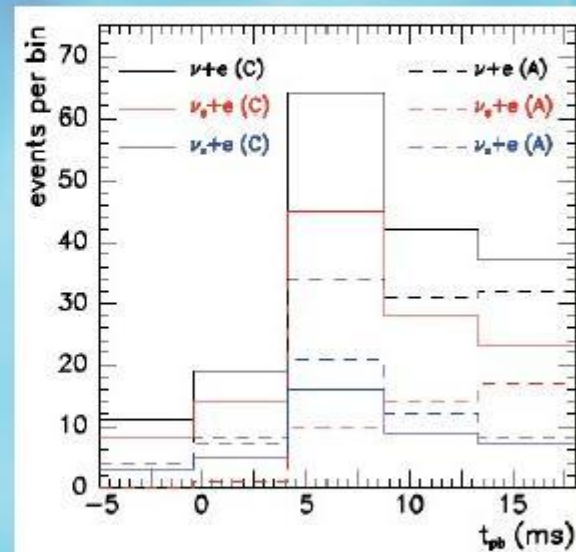
Bkg: —

- mainly $\bar{\nu}_e + p \rightarrow e^+ + n$
- rejected by angle and E cuts + Gd n-tag
- ES of other ν flavours

Observation of time peak
depends on oscillation scenario

| Scenario | Hierarchy | $\sin^2 \vartheta_{13}$ | p | \bar{p} |
|----------|-----------|-------------------------|-------------------------|-------------------------|
| A | Normal | $\gtrsim 10^{-3}$ | 0 | $\cos^2 \vartheta_{13}$ |
| B | Inverted | $\gtrsim 10^{-3}$ | $\sin^2 \vartheta_{13}$ | 0 |
| C | Any | $\lesssim 10^{-5}$ | $\sin^2 \vartheta_{13}$ | $\cos^2 \vartheta_{13}$ |

Kachelrieß et al.,
astro-ph/0412082



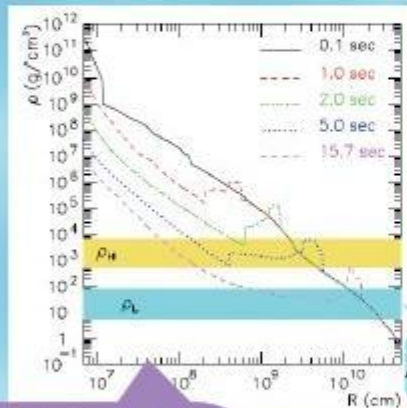
- Burst / no-burst \Rightarrow break degeneracy A/C if θ_{13} unknown
- Measurement of SN distance $D \sim 1/N^{1/2}$ @10kpc within 5%

MEMPHYS: SN shockwave information

SN spectral analyses (2)

⊙ Learning about the shock wave

Crossing of resonances can induce time-dependent matter effects in neutrino oscillations



$\Delta m^2_{\text{atm}}, \theta_{13}$

$\nu_e \leftrightarrow NH$

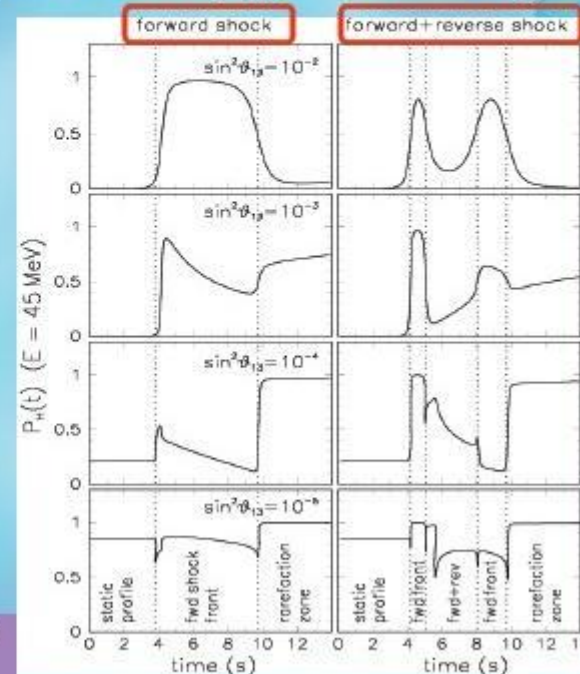
$\nu_e \leftrightarrow IH$

$\Delta m^2_{\text{sol}}, \theta_{\text{sol}}$

ν self-interactions?
Duan et al., 0606616
Raffelt et al., 0608050

Schirato and Fuller, astro-ph/0205390
Fogli et al., hep-ph/0304056

Shock-wave effects on survival probabilities (P_H) depend on θ_{13} .

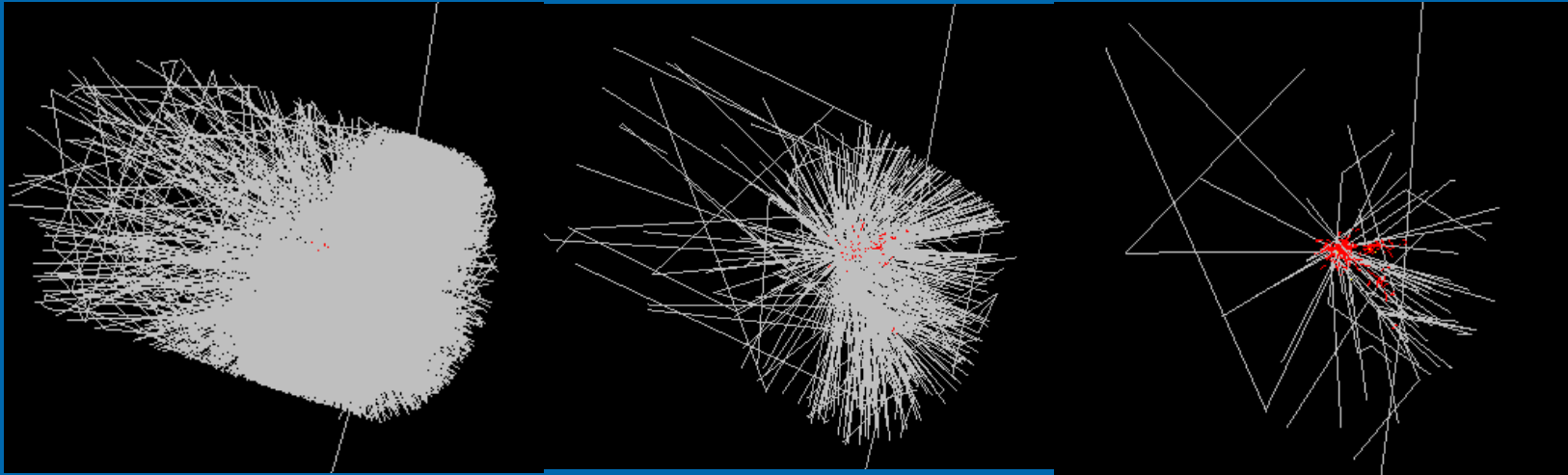


NOW06 15/09/06

A.Tonazzo - MEMPHYS: non-oscillation physics

2

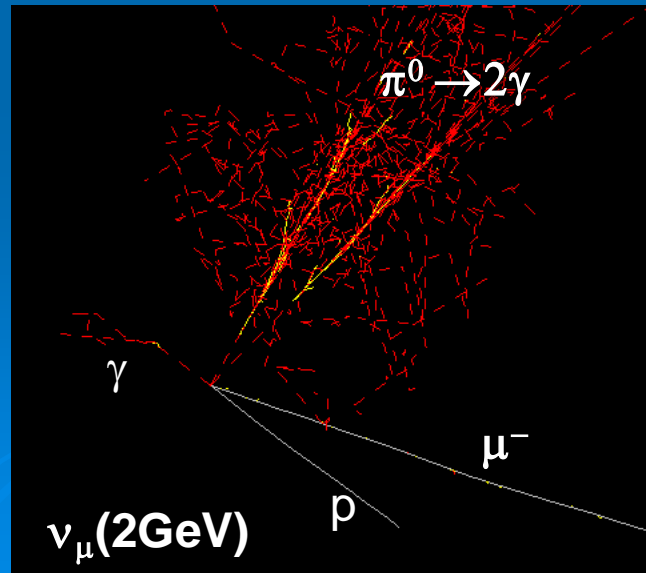
2km WC Geometry



10% de γ Č

1% de γ Č

0.1% de γ Č



transparency by J. E. Campagne