



# XeSAT 2022 — Coimbra, 23–26 May 2022

## International Workshop on Applications of Noble Gas Xenon to Science and Technology

### Searching for Majorana neutrinos with nEXO

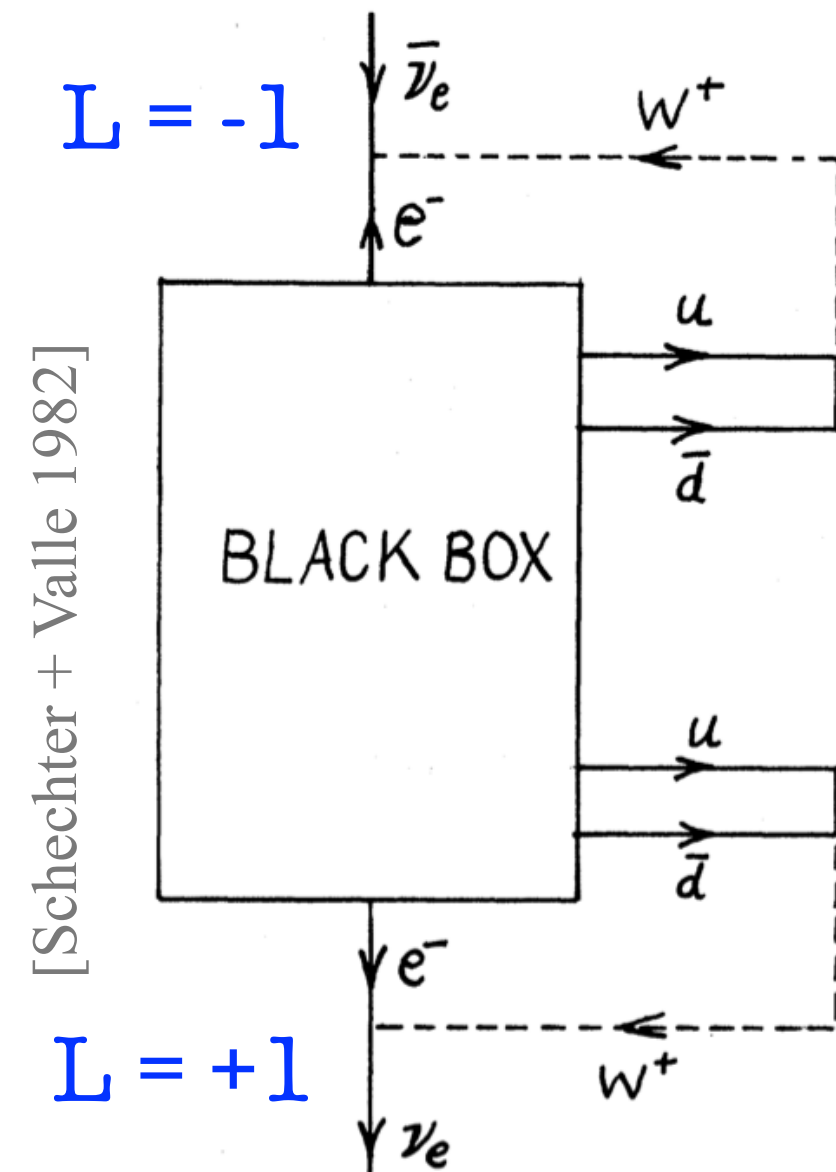
Andrea Pocar @ University of Massachusetts, Amherst

on behalf of the nEXO collaboration





# $0\nu\beta\beta$ decay — motivation

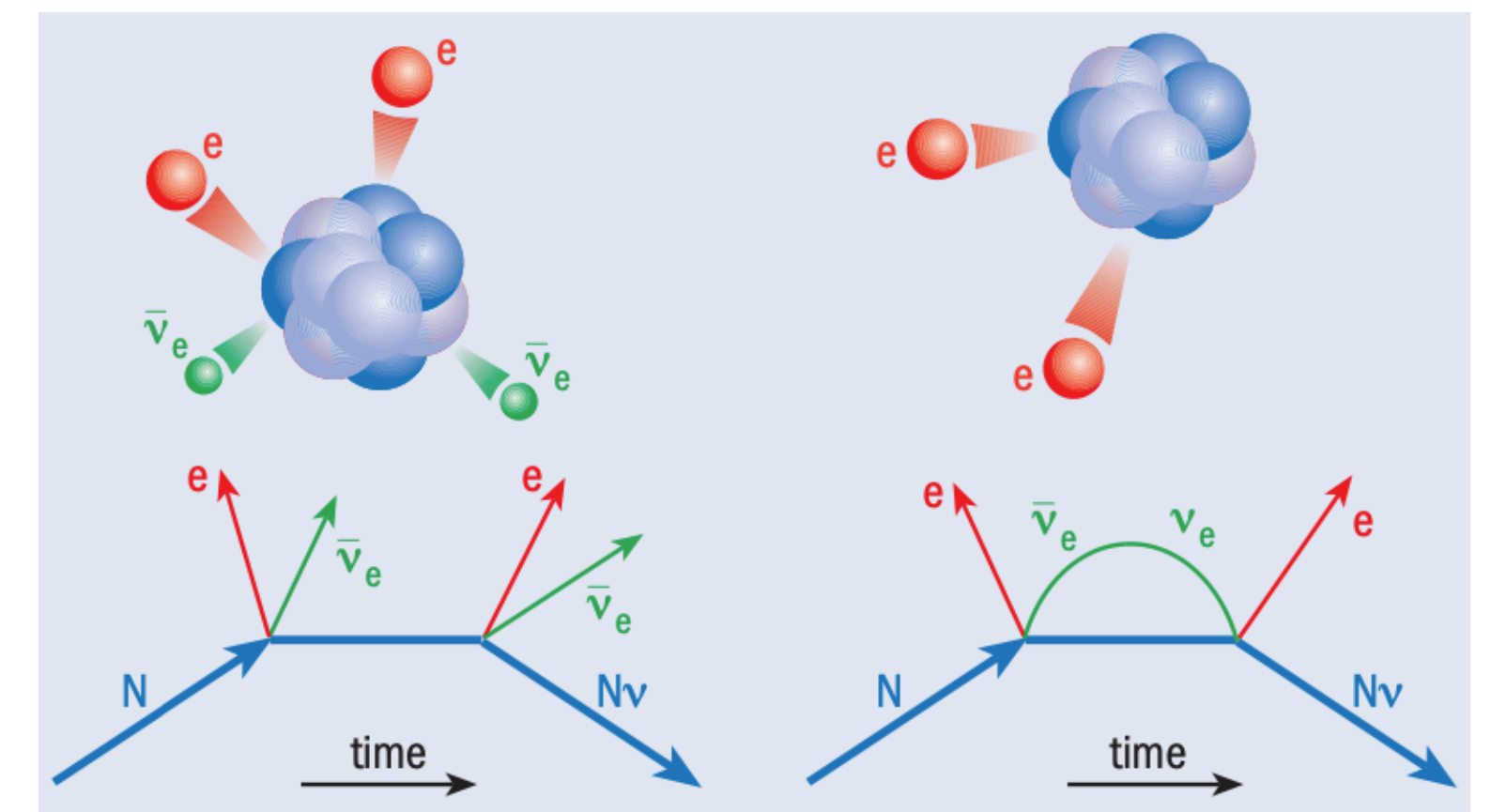
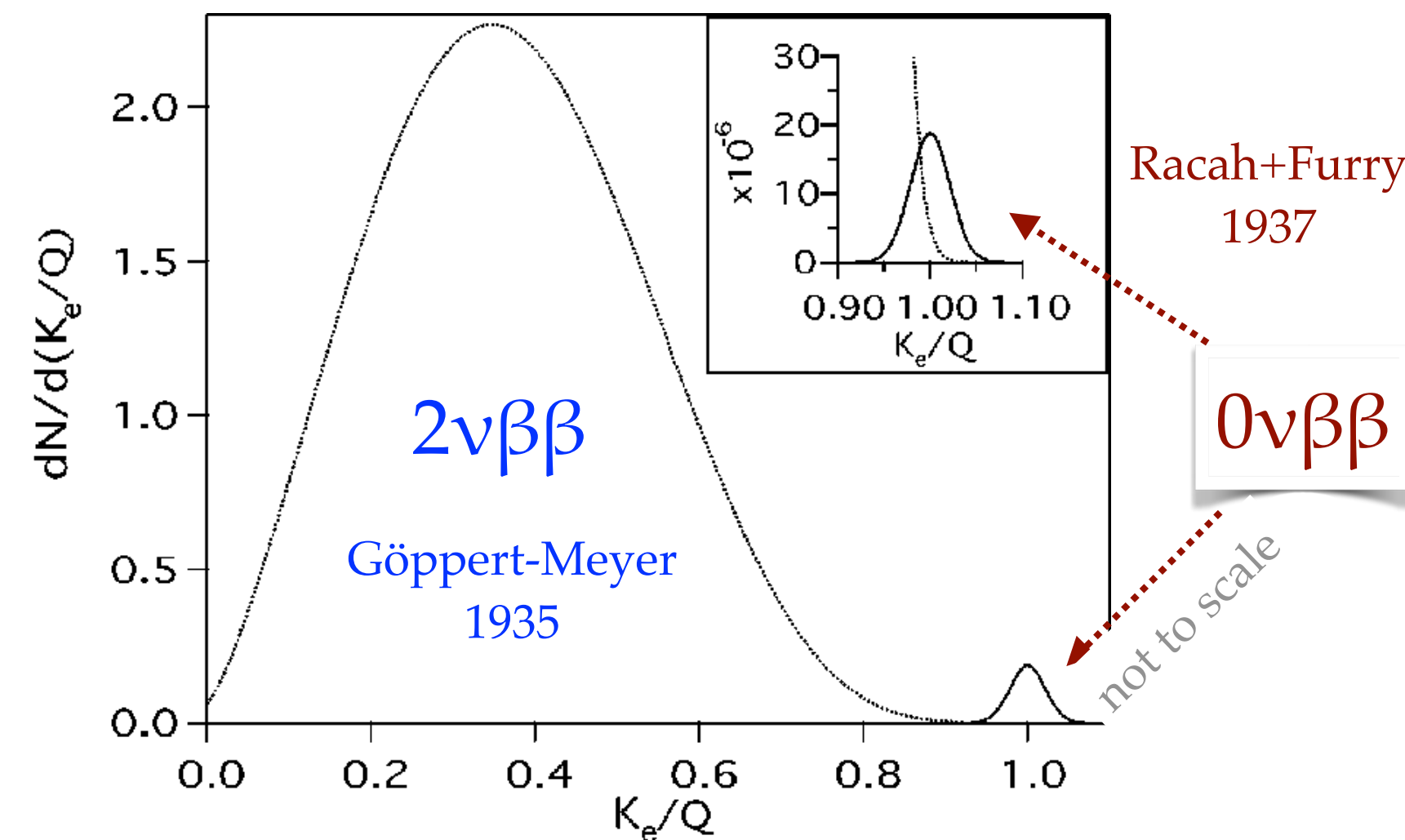
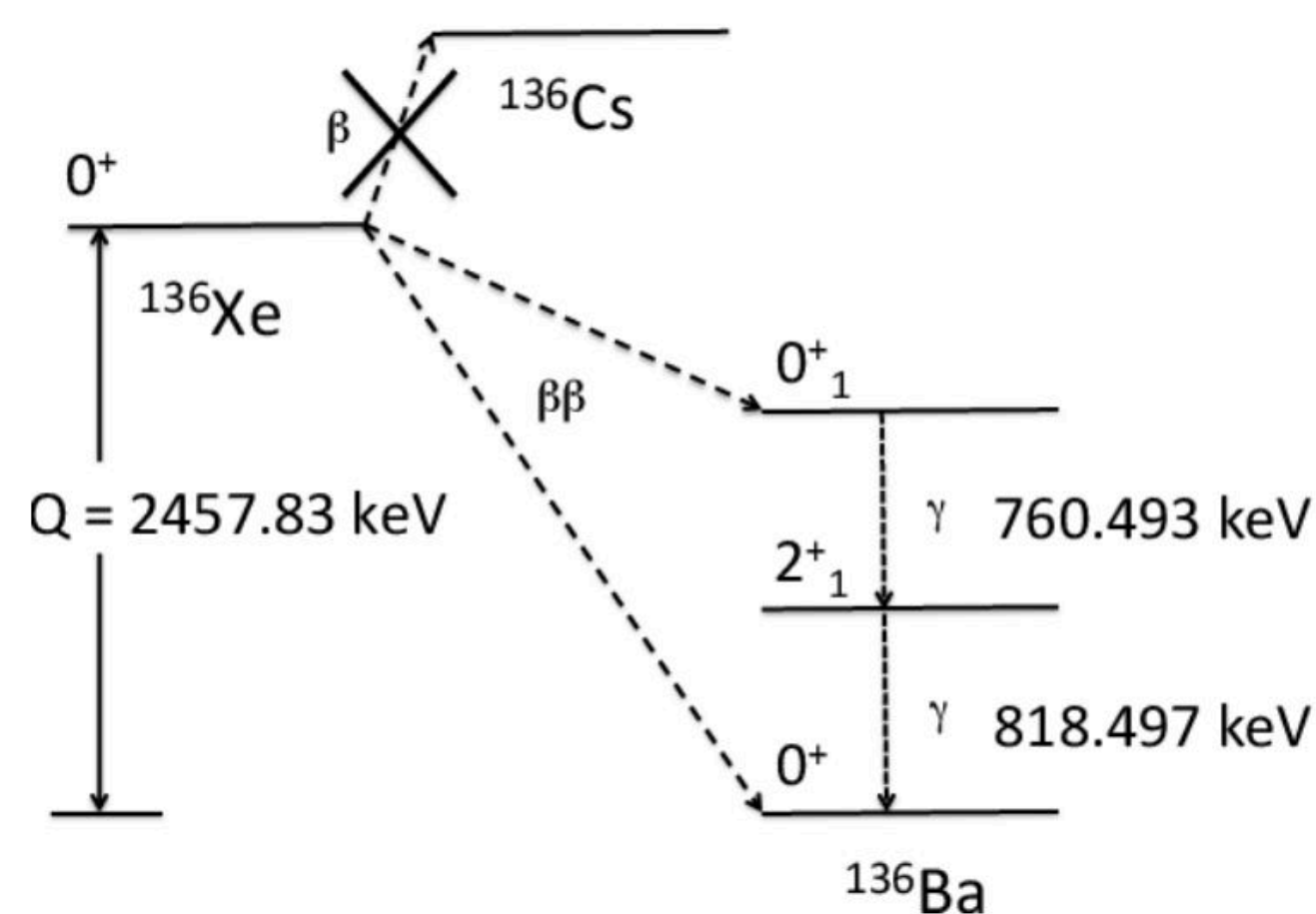


## observation of $0\nu\beta\beta$ decay:

- massive, Majorana neutrinos
- lepton number violation ( $\Delta L=2$ )
- new mass creation mechanism (non Higgs)
- matter dominance in the universe,  $\Delta(B-L)$  leptogenesis
- smallness of neutrino mass  $\rightarrow$  new BSM physics mass scale

## $0\nu\beta\beta$ rate

- absolute neutrino mass (model dependent)





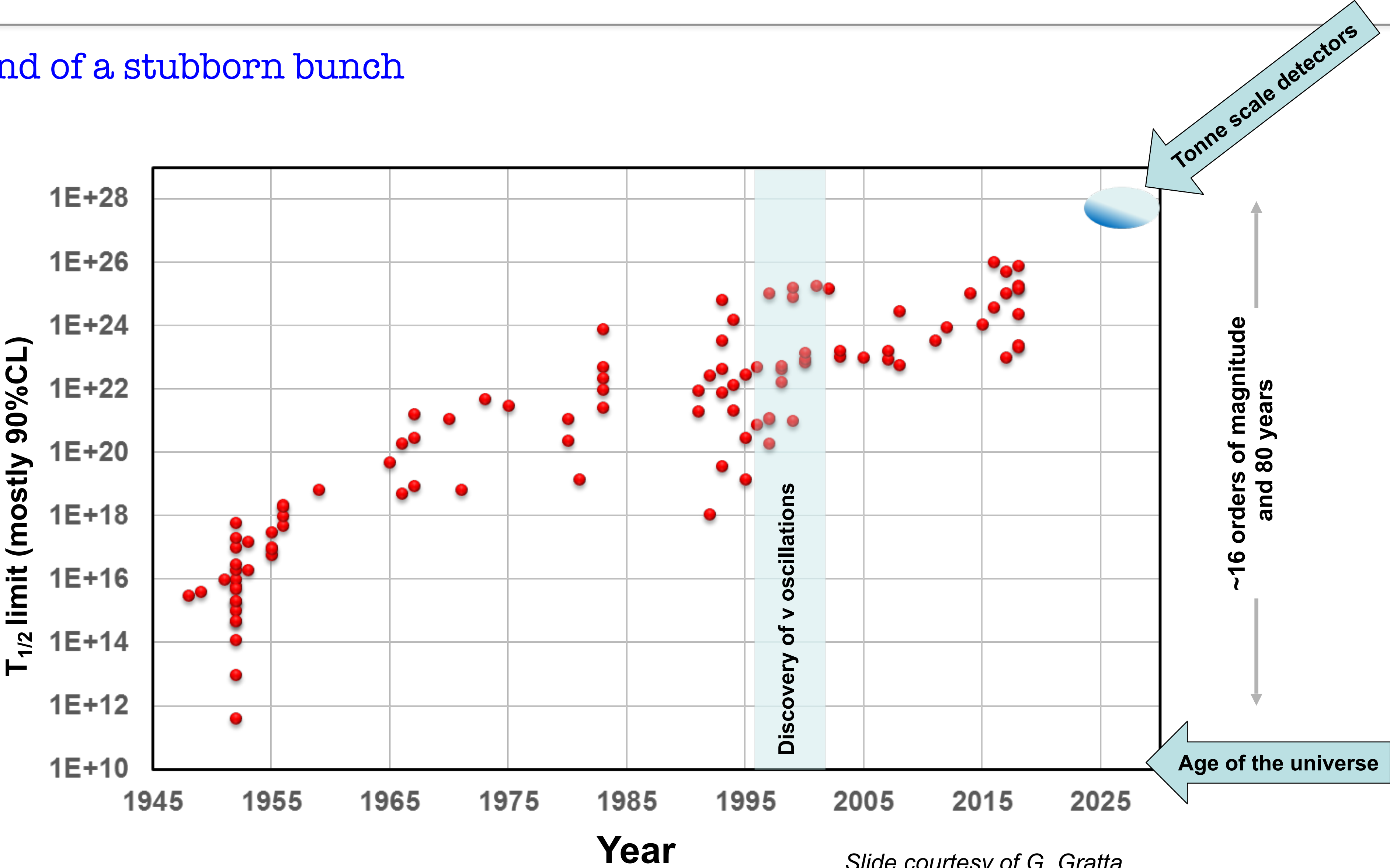
# nEXO : A world wide effort involving 9 countries, 33 institutions, ~200 collaborators





# The history of $0\nu\beta\beta$ decay experiments in one slide

...we are kind of a stubborn bunch



Slide courtesy of G. Gratta  
Data courtesy of S.Elliott and the PDG.  
Not all results are necessarily shown.



- The global  $0\nu\beta\beta$  decay program is gearing up for the so-called “tonne-scale” phase (as communicated by the European and North American funding agencies at the 10/2021 International Forum at LNGS)
- Three major experiments with different isotopes:
  - nEXO ( $^{136}\text{Xe}$ ), Legend-1000 ( $^{76}\text{Ge}$ ), CUPID ( $^{100}\text{Mo}$ )
  - very different experimental techniques, healthy program
- nEXO plans a 5-tonne, single phase, LXe TPC (90% enriched in  $^{136}\text{Xe}$ )  
100-fold increase in sensitivity wrt current experiments
- nEXO builds on the successful EXO-200 program which has demonstrated the key technical features of this technology at scale, such as:
  - Effectiveness of self-shielding and low intrinsic background
  - Energy resolution via collection of ionization and scintillation
  - Event topology ( $\beta$  vs.  $\gamma$ ) and particle ID ( $\beta/\gamma$  vs.  $\alpha$ )
  - Continuous purification of the source



# EXO milestones (2001- ): R&D -> EXO-200 -> nEXO



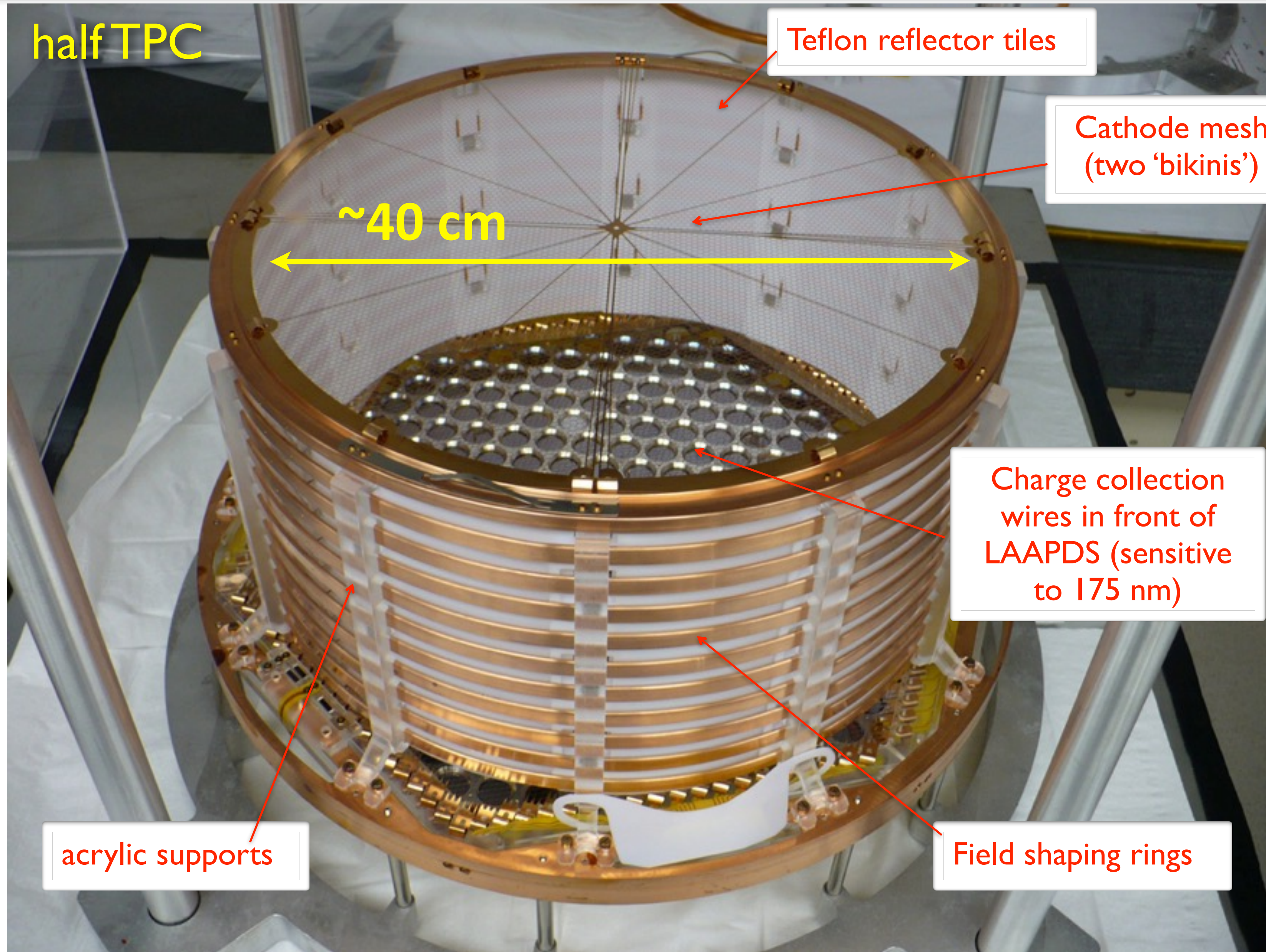
2001	"EXO" started as an R&D towards a $^{136}\text{Xe}$ $\beta\beta$ decay experiment.
2002	Improved energy resolution in LXe using the correlation between scintillation and ionization is discovered.
Ca. 2005	Settled on a LXe TPC design for a "prototype" 200 kg detector.
2007-2010	The EXO-200 detector is designed and built, with major contributions from Canada, Russia and Switzerland.
2012-2016	After EXO-200 started taking data, showing excellent performance, the idea of a 5000 kg was further developed.
2014.	The "nEXO collaboration" was formed.
2014-2016	Five US Nat'l Labs join the collaboration.
May 2018	nEXO pre-CDR posted on the arXiv
Nov 2018	CD-0 for tonne-scale $\beta\beta$ decay
Dec 2018	End of EXO-200 run
2019-now	nEXO project developed; substantial nEXO engineering at SNOLAB
Feb 2020	nEXO MAC review
Feb 2021.	nEXO budget review
Jul 2021	DoE portfolio review
Sept 2021	Europe – North America Summit at LNGS



# the EXO-200 TPC

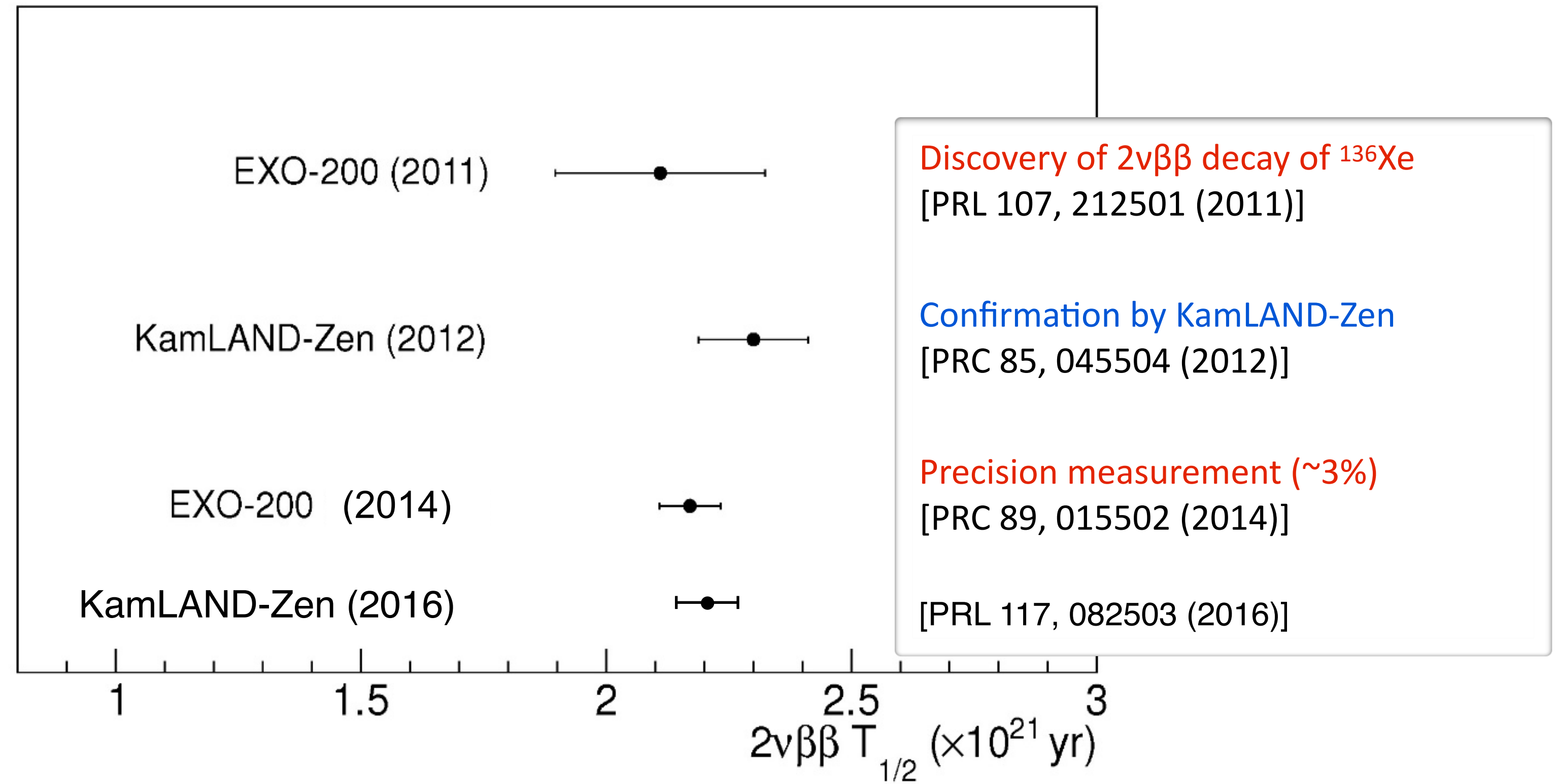
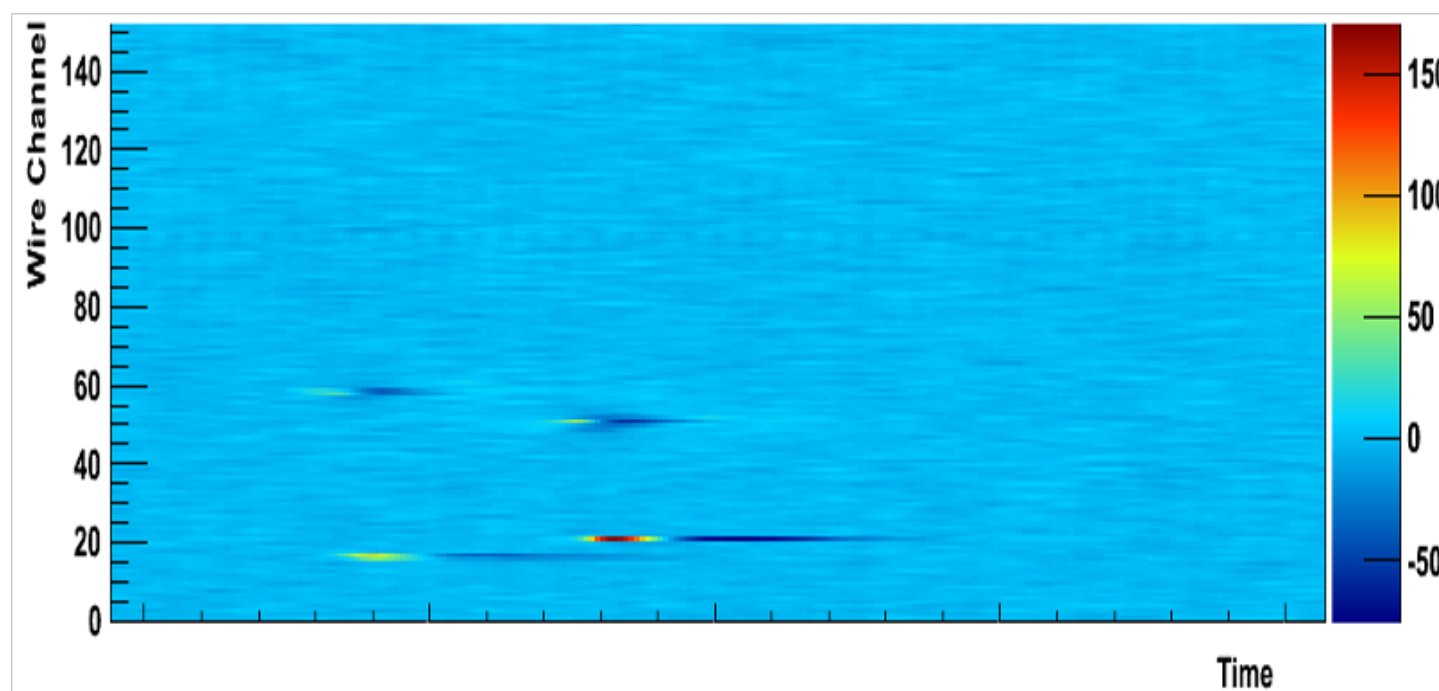
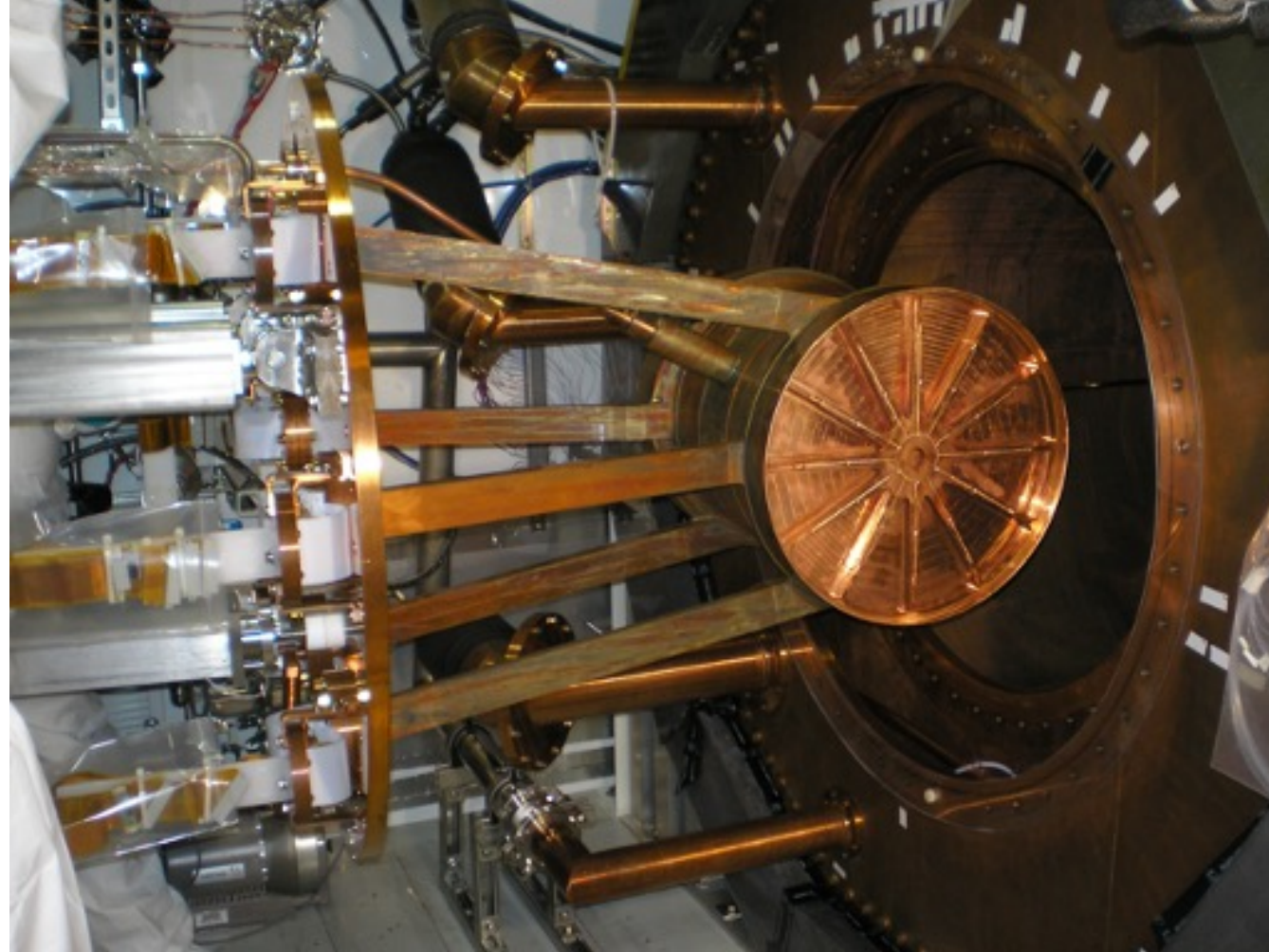


half TPC





# Precision measurement of $2\nu\beta\beta$

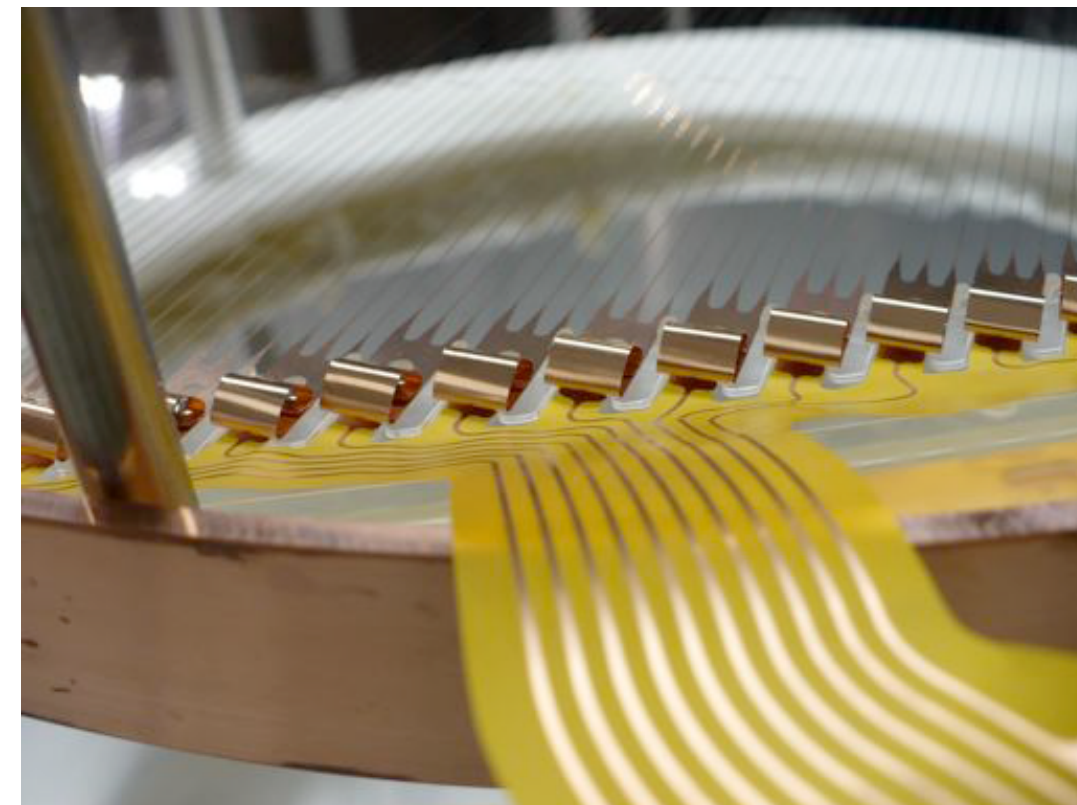
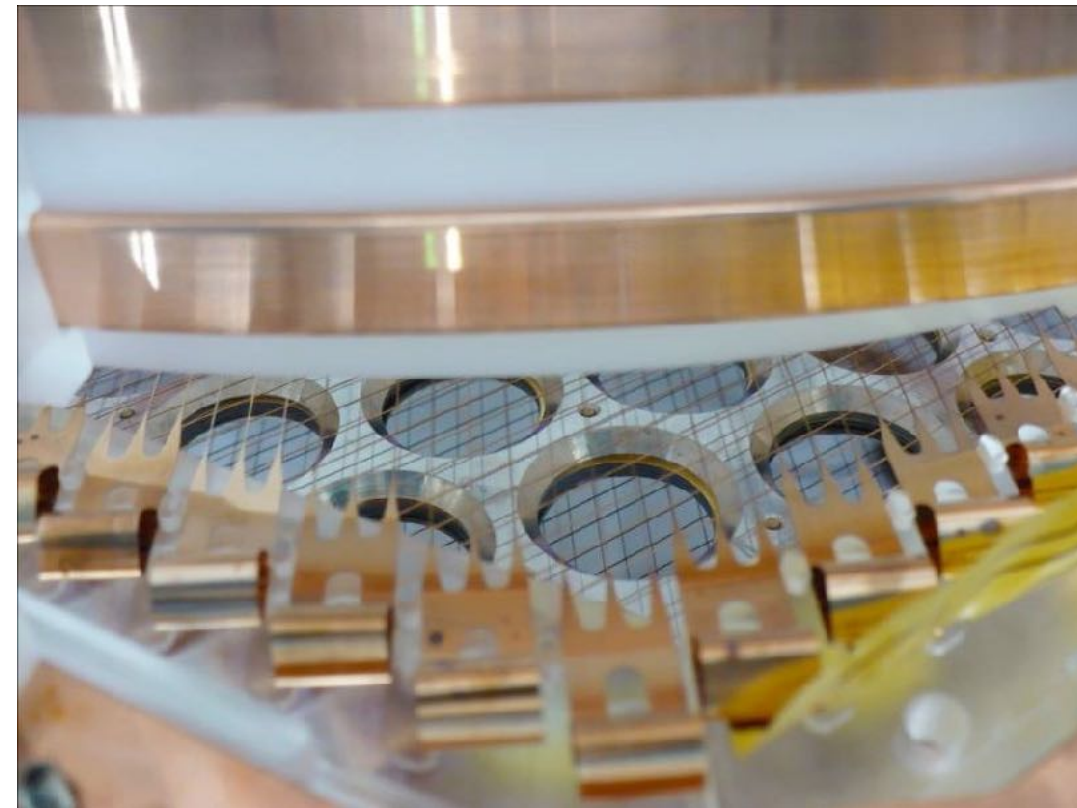


$$T_{1/2}^{2\nu\beta\beta} = (2.165 \pm 0.016(\text{stat}) \pm 0.059(\text{syst})) \times 10^{21} \text{ yr}$$

(longest, first to be precisely\* (directly) measured  $2\nu\beta\beta$  decay of all 'practical' isotopes)  
(\* Ge-76 and Te-130 have similarly precise measurements)



# EXO-200 $0\nu\beta\beta$ decay results



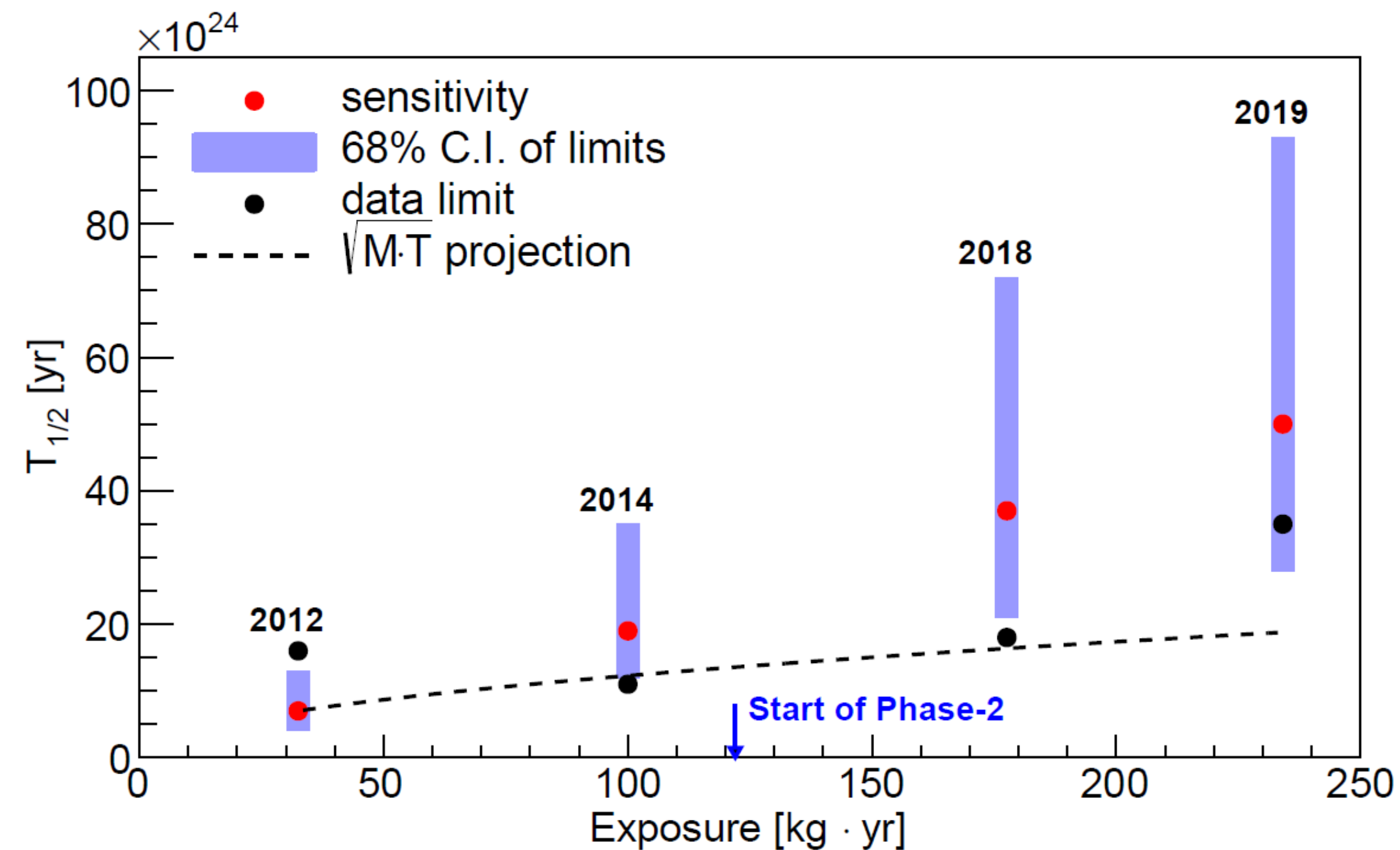
Phase I+II: 234.1 kg·yr  $^{136}\text{Xe}$  exposure

Limit  $T_{1/2}^{0\nu\beta\beta} > 3.5 \times 10^{25}$  yr (90% C.L.)

$\langle m_{\beta\beta} \rangle < (93 - 286)$  meV

Sensitivity  $5.0 \times 10^{25}$  yr

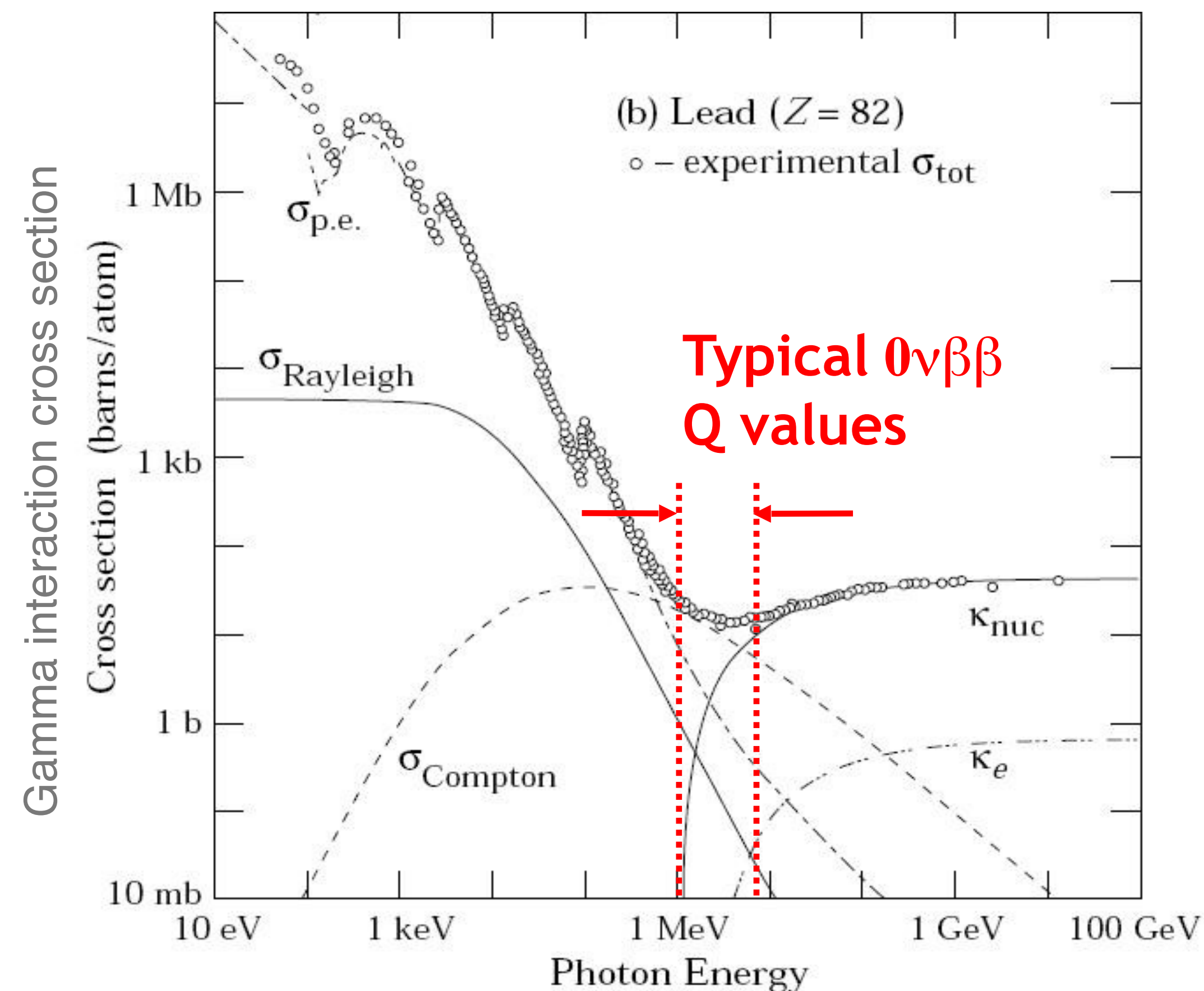
No statistically significant signal observed



2012: Phys.Rev.Lett. 109 (2012) 032505  
 2014: Nature 510 (2014) 229-234  
 2018: Phys. Rev. Lett. 120, 072701 (2018)  
 2019: Phys. Rev. Lett. 123, 161802 (2019)



# Key requirement: shielding from MeV $\gamma$ -rays



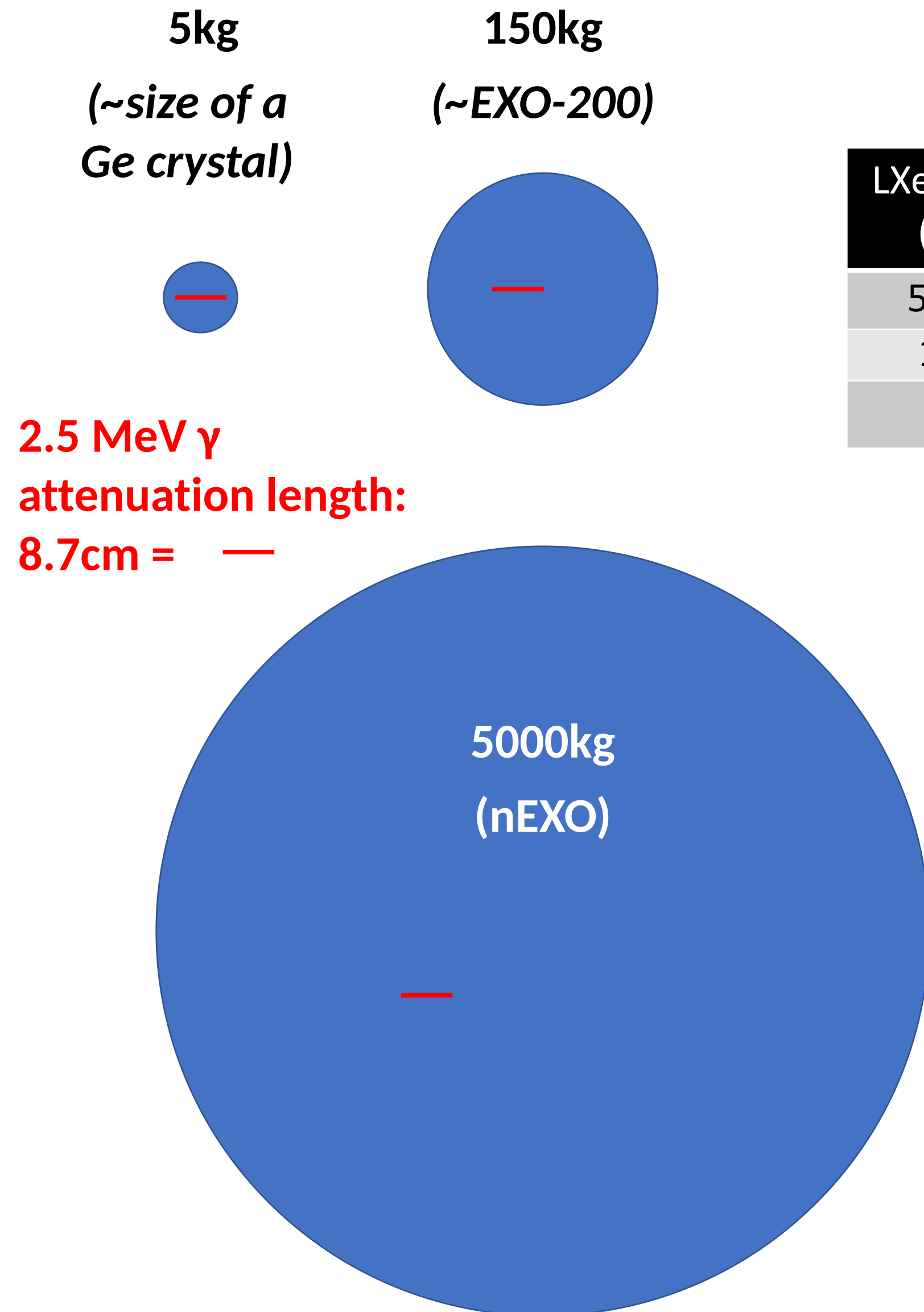
Shielding  $\beta\beta$  decay detectors from external electromagnetic background is harder/different than shielding Dark Matter detectors

We are entering the “golden era” of  $\beta\beta$  decay experiments as detector sizes exceed gamma-ray interaction lengths

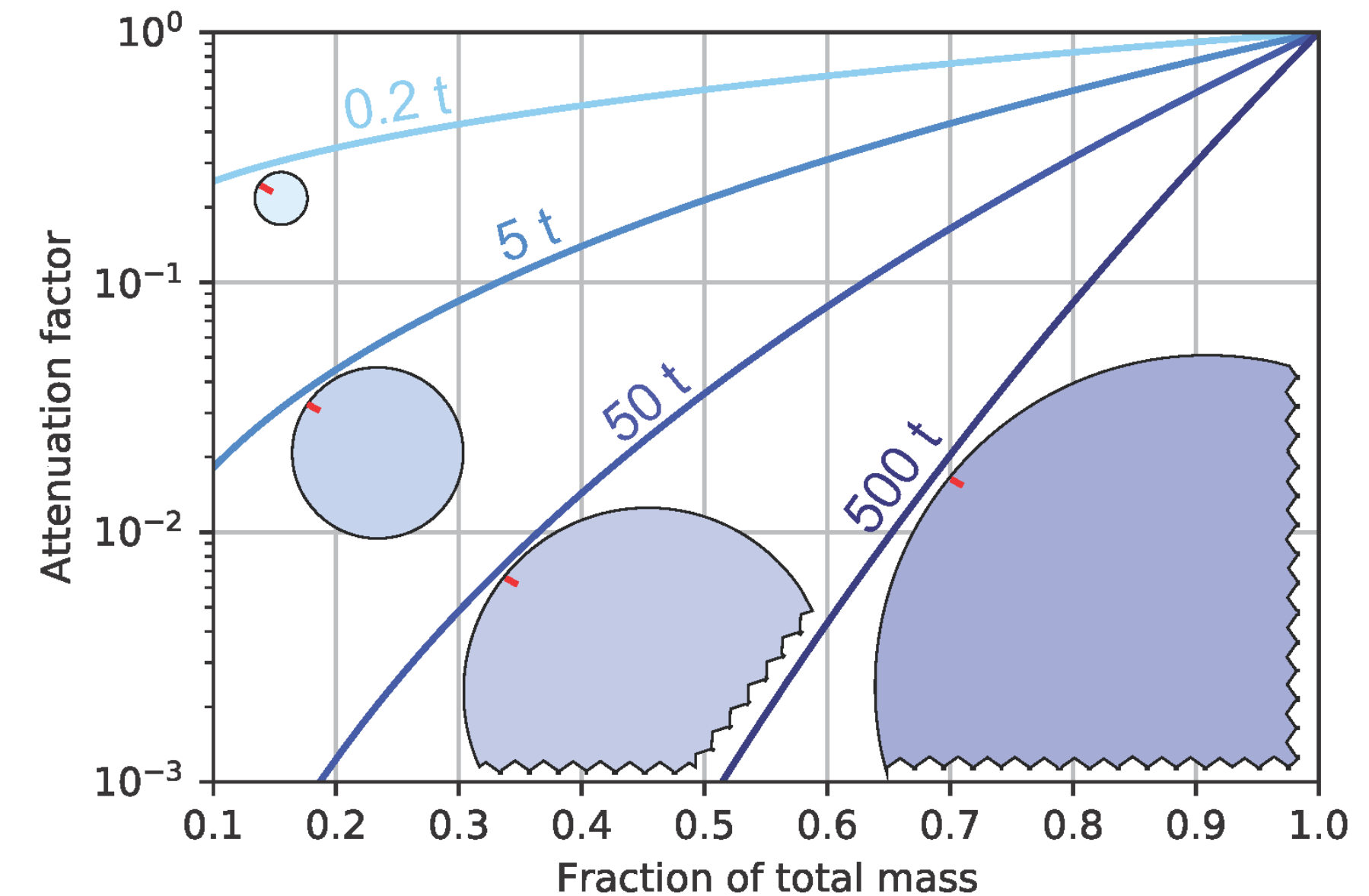
E.g: the  $\gamma$ -ray interaction length in Ge is 4.6 cm, comparable to the size of a germanium detector



# Power of self-shielding, monolithic, homogeneous detector



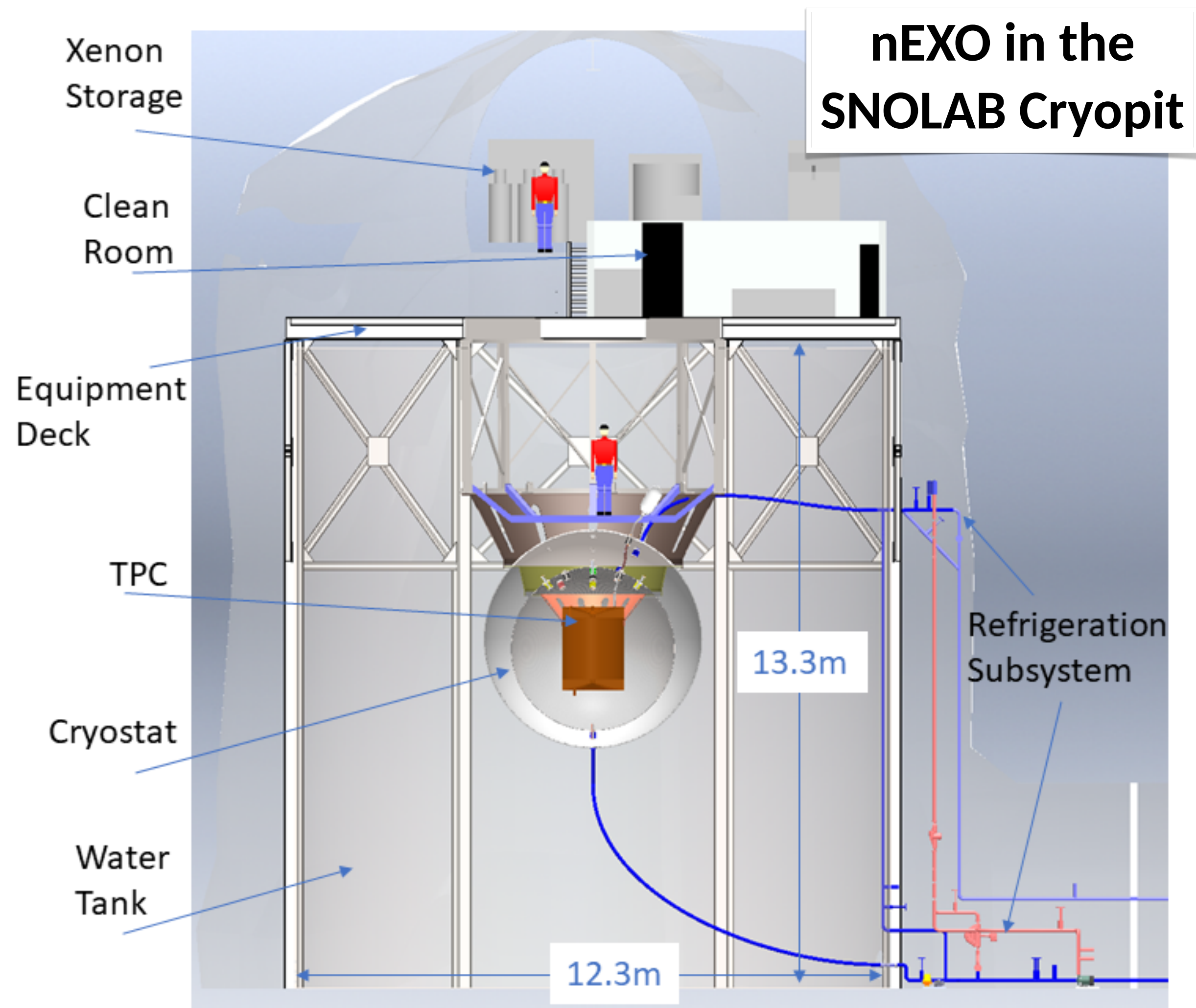
LXe mass (kg)	Linear size (cm)
5000	130
150	40
5	13



## Advantages of LXe technology for $0\nu\beta\beta$ decay:

- Scalable, re-purifiable, transferable between detectors
- Low intrinsic background (fully exploited at the tonne scale)
- Particle ID ( $\beta/\alpha$ ), event topology ( $\beta/\gamma$ )
- Possibility of no-source control experiment



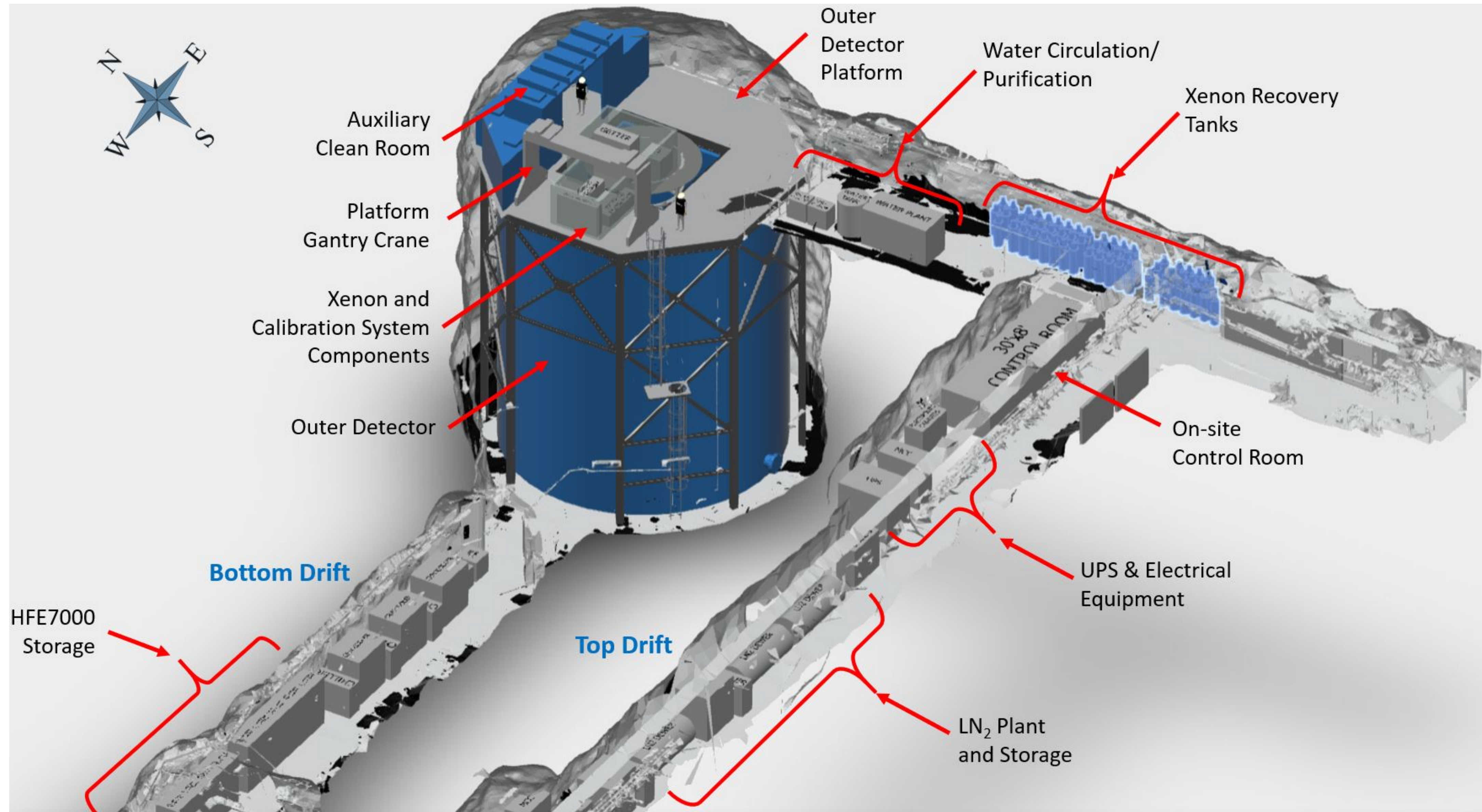


- Double-walled, vacuum-insulated carbon composite cryostat
- Refrigeration via 32 tonnes of Hydro-fluoro ether (HFE) 3M Novec-7000
- Design proven with EXO-200
- HFE-7000 intrinsically ultra radio-pure
- Excellent temperature stability ( $\sim 165\text{K}$ )
- Thermal storage in case of power loss
- Lightweight, ultra-low background Xe vessel (pressure set at the Inner Cryostat)
- Active water-Cherenkov muon veto

*nEXO pre-conceptual Design Report: [arXiv:1805.11142](https://arxiv.org/abs/1805.11142)*

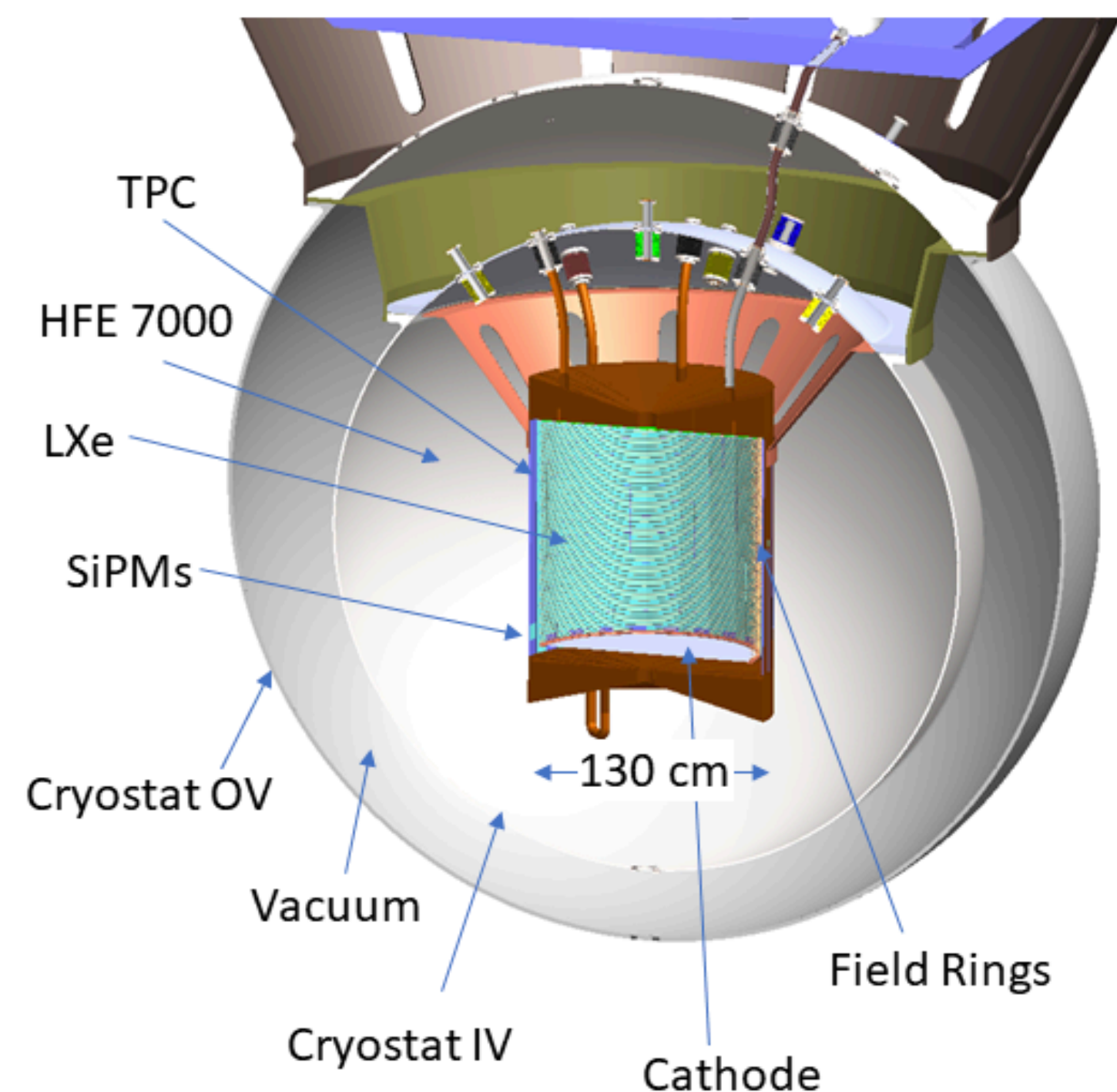


# nEXO conceptual layout at SNOLAB (preferred site)





# The nEXO detector is an evolution from EXO-200

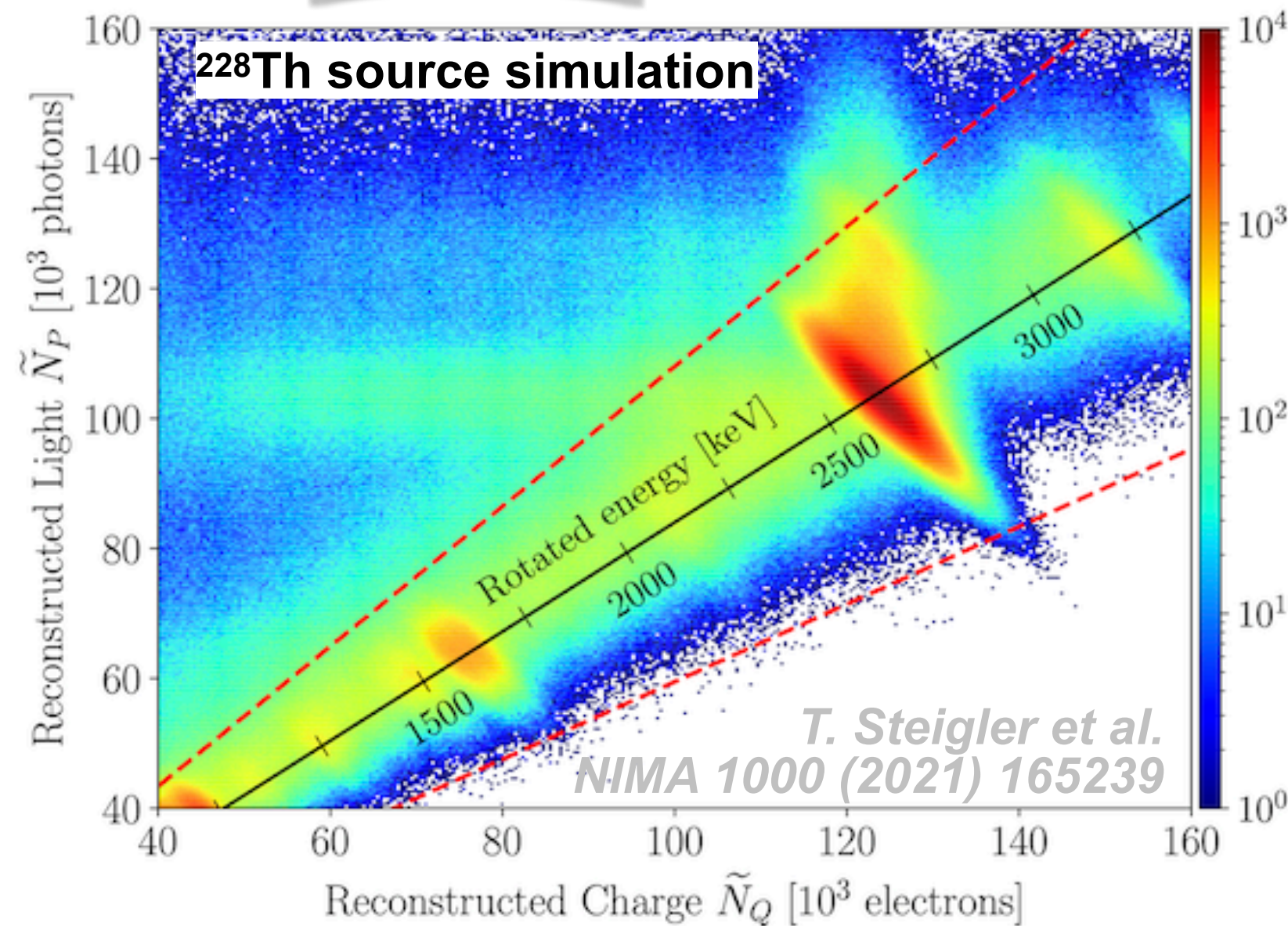


- 5000 kg of 90%-enriched LXe
- Single 120 cm drift volume; 130 cm diameter
- Drift E-field ~400 V/cm
- Ionization electrons collected on charge tile detectors at the anode (no gain), ~6,000 channels
- VUV (178 nm) scintillation light is detected by a large array of SiPMs (~45,000 devices, ~4.5 m<sup>2</sup>)

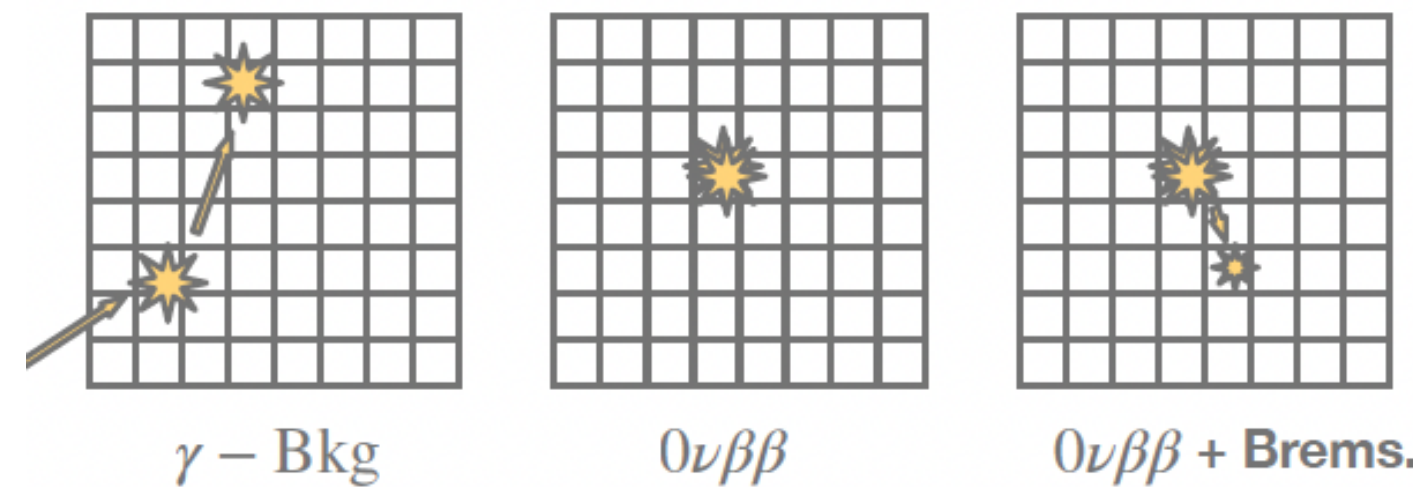
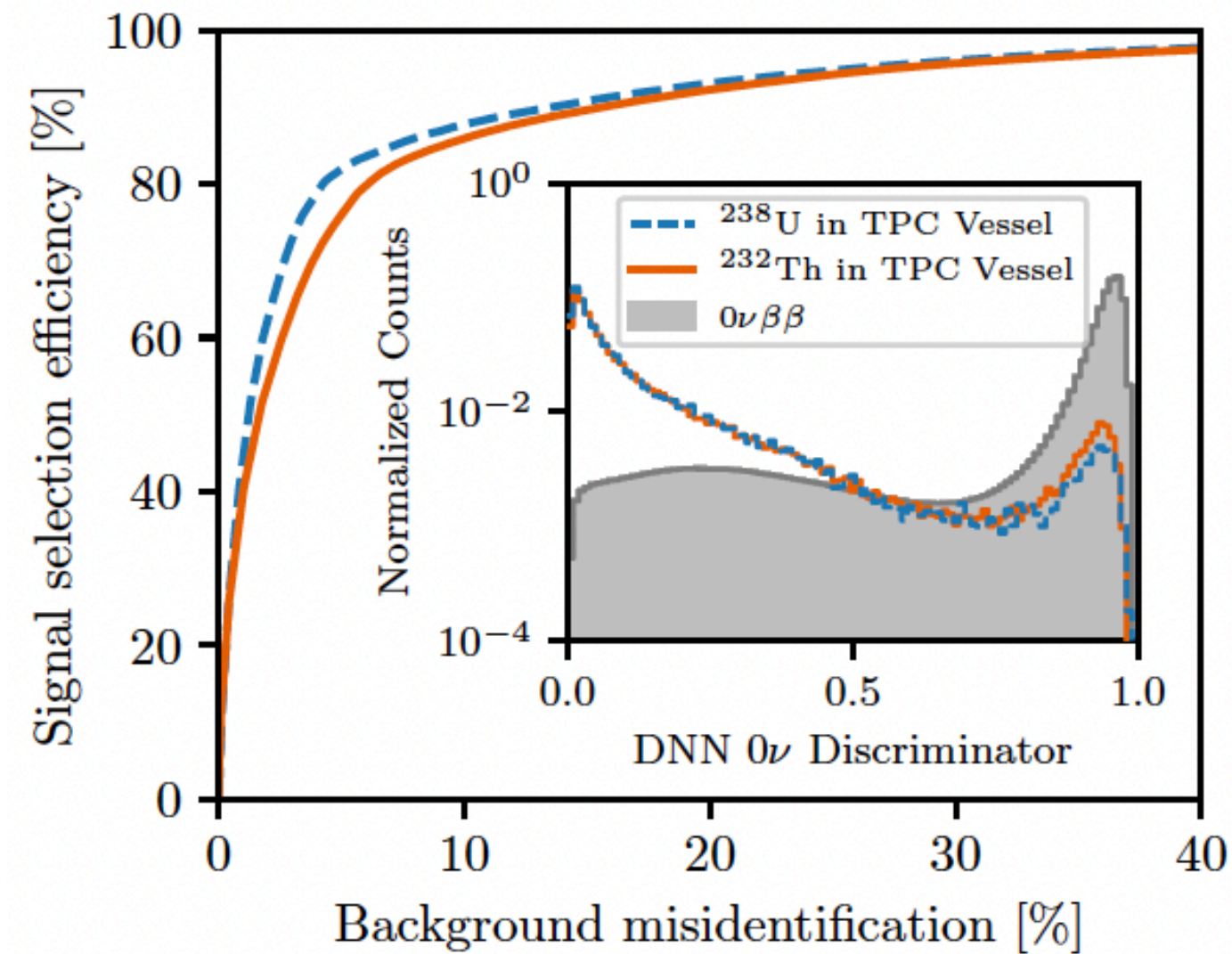
	EXO-200:	nEXO:	Improvements:
Vessel and cryostat	Thin-walled commercia Cu w/HFE	Thin-walled electroformed Cu w/HFE	Lower background
High voltage	Max voltage: 25 kV (end-of-run)	Operating voltage: 50 kV	Full scale parts tested in LXe prior to installation to minimize risk
Cables	Cu clad polyimide (analog)	Cu clad polyimide (digital)	Same cable/feedthrough technology, R&D identified 10x lower bkg substrate and demonstrated digital signal transmission
e <sup>-</sup> lifetime	3-5 ms	5 ms (req.), 10 ms (goal)	Minimal plastics (no PTFE reflector), lower surface to volume ratio, detailed materials screening program
Charge collection	Crossed wires	Gridless modular tiles	R&D performed to demonstrate charge collection with tiles in LXe, detailed simulation developed
Light collection	APDs + PTFE reflector	SiPMs around TPC barrel	SiPMs avoid readout noise, R&D demonstrated prototypes from two vendors
Energy resolution	1.2%	1.2% (req.) 0.8% (goal)	Improved resolution due to SiPMs (negligible readout noise in light channels)
Electronics	Conventional room temp.	In LXe ASIC-based design	Minimize readout noise for light and charge channels, nEXO prototypes demonstrated in R&D and follow from LAr TPC lineage
Background control	Measurement of all materials	Measurement of all materials	RBC program follows successful strategy demonstrated in EXO-200
Larger size	>2 atten. length at center	>7 atten. length at center	Exponential attenuation of external gammas and more fully contained Comptons



## Energy

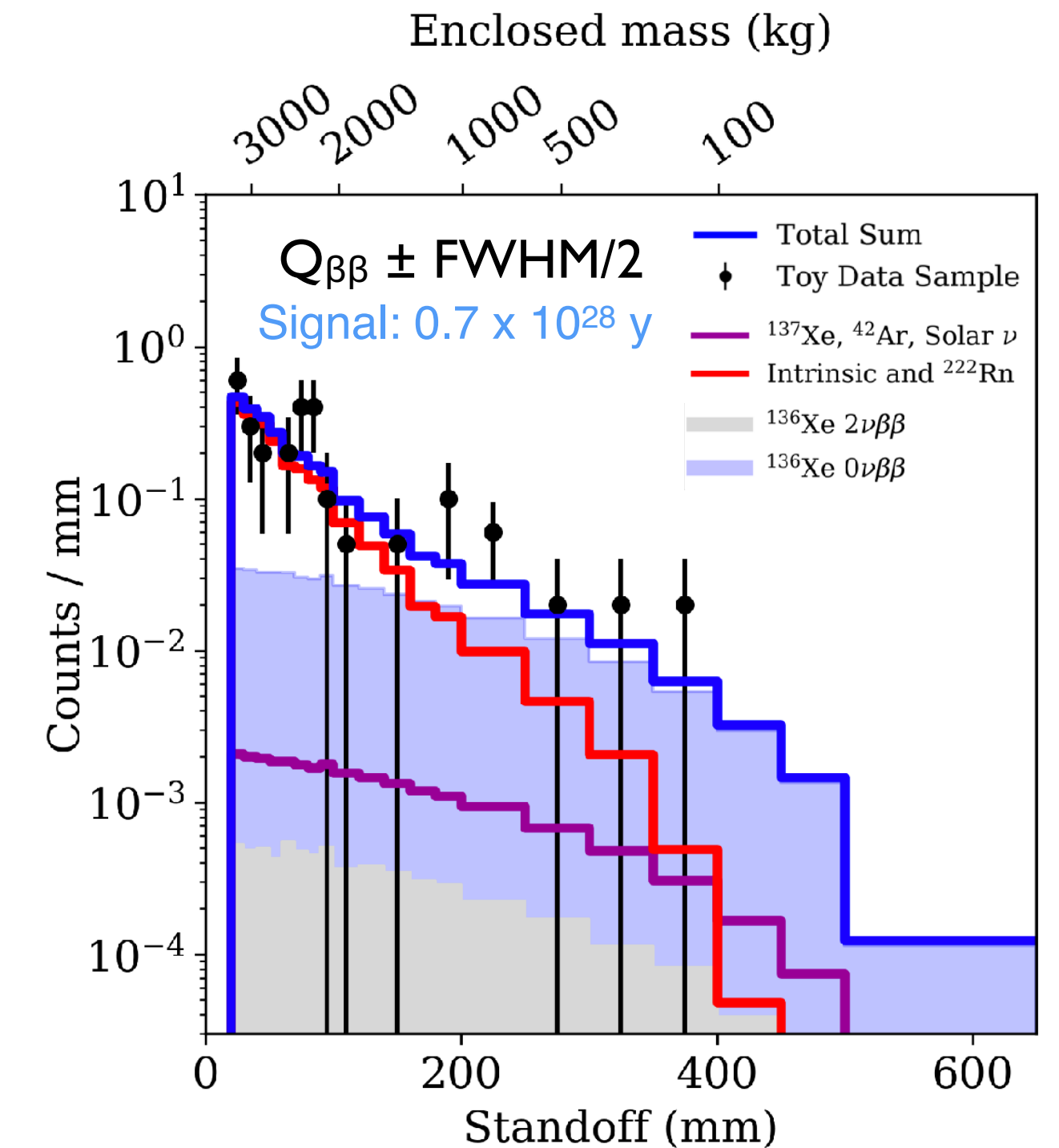


## Event topology



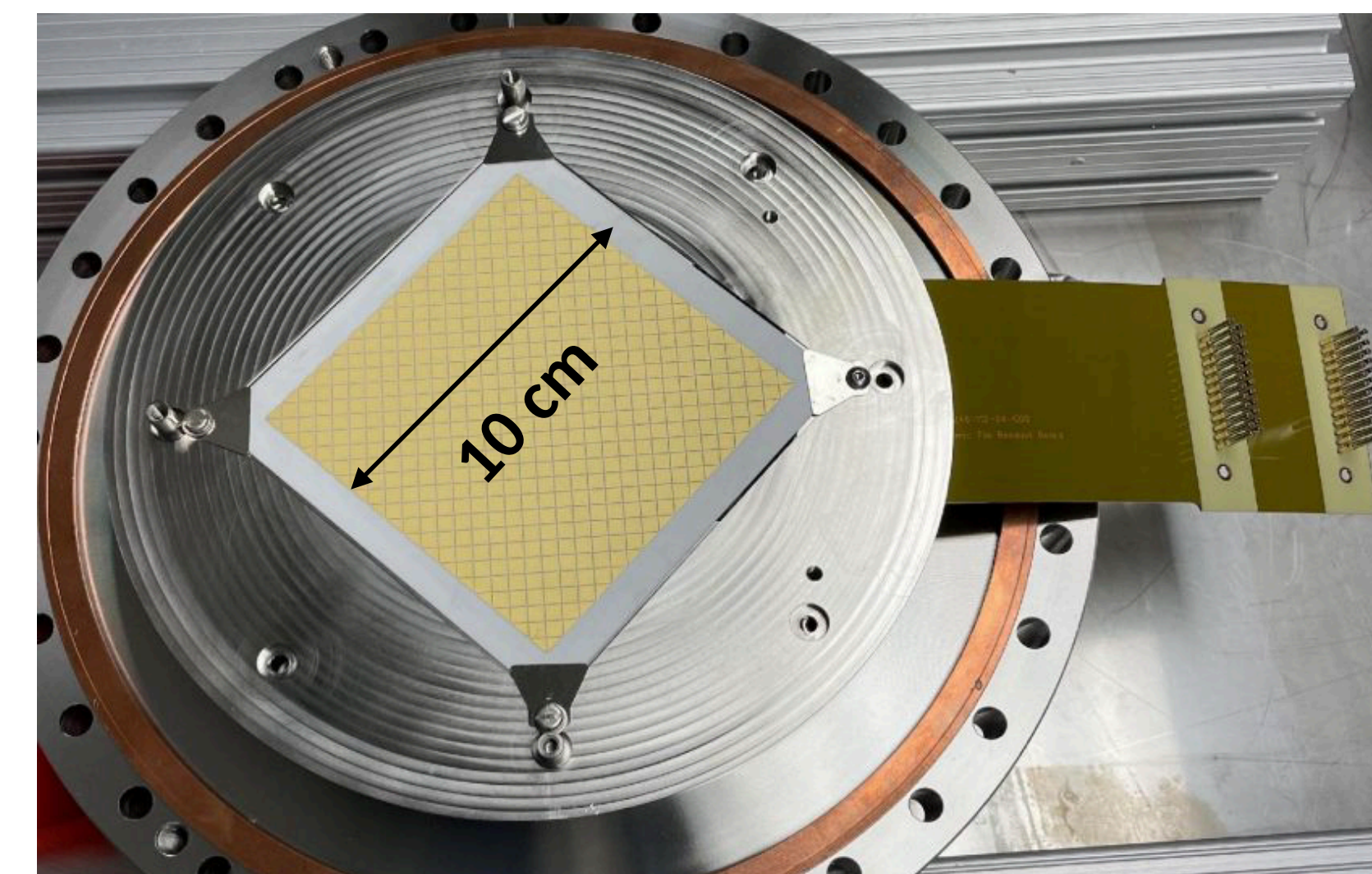
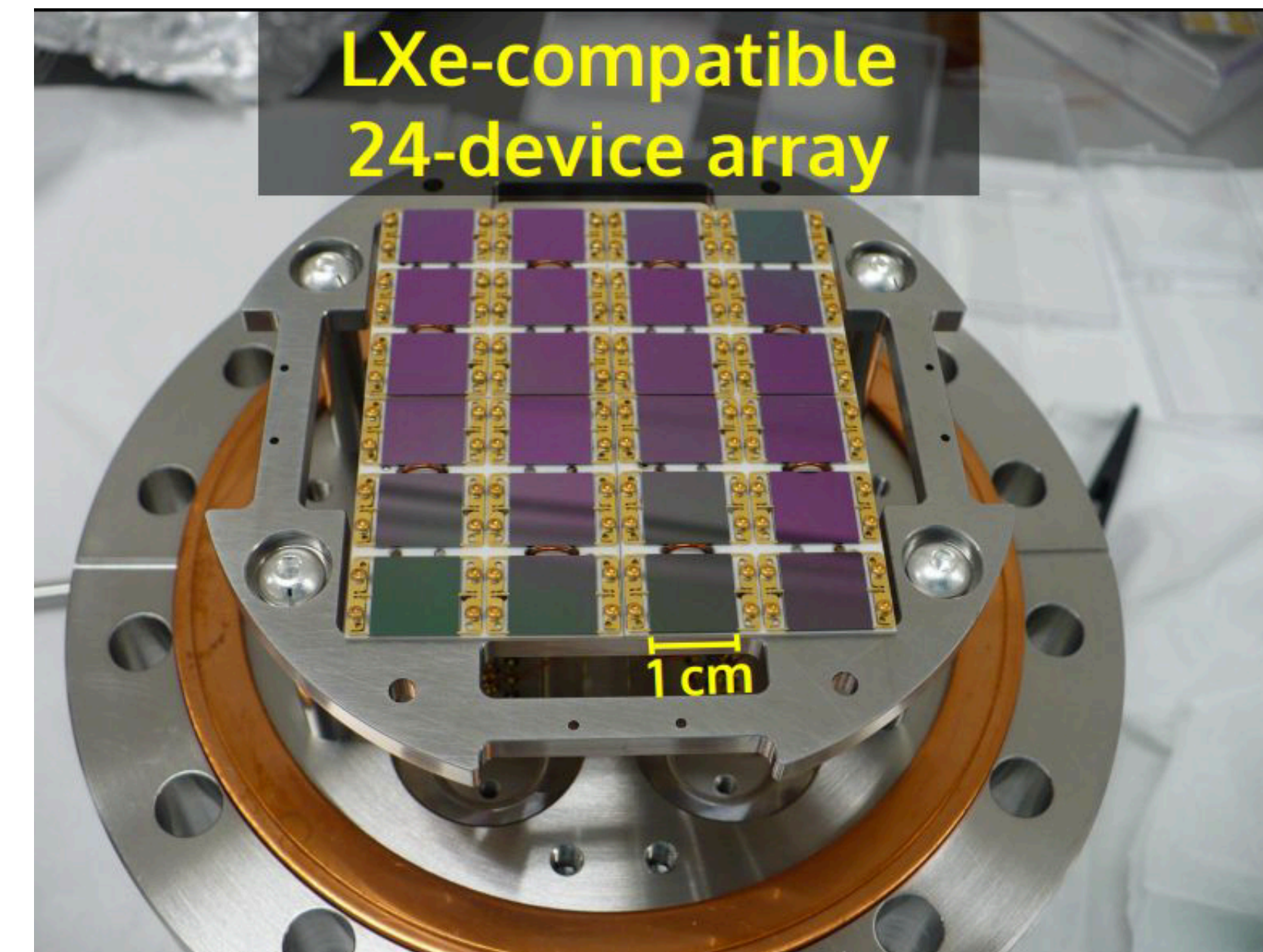
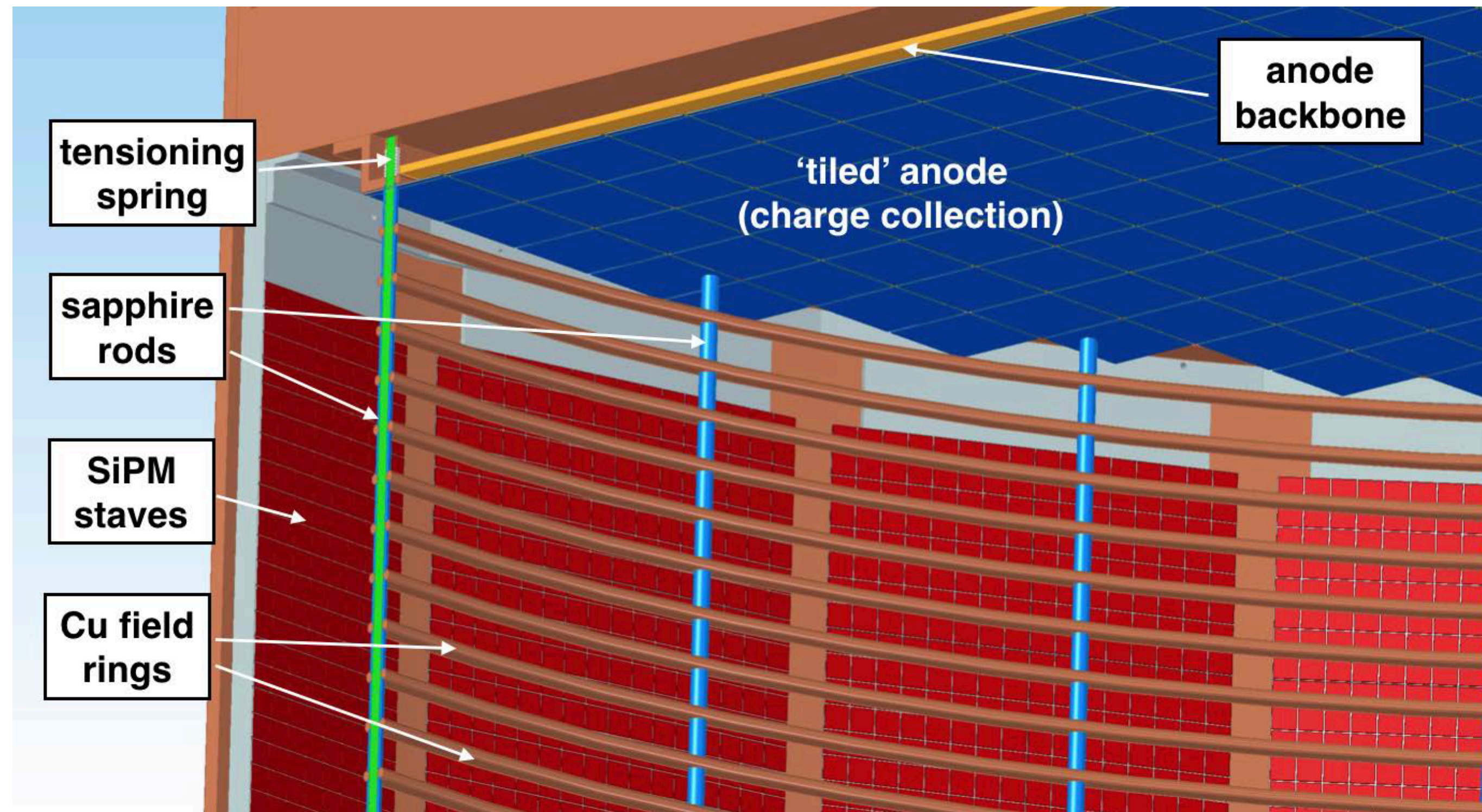
- Single- vs. multi- site energy depositions

## Standoff distance



- $\beta\beta$  events are uniformly distributed in the LXe volume
- Most backgrounds originate from outside of the TPC

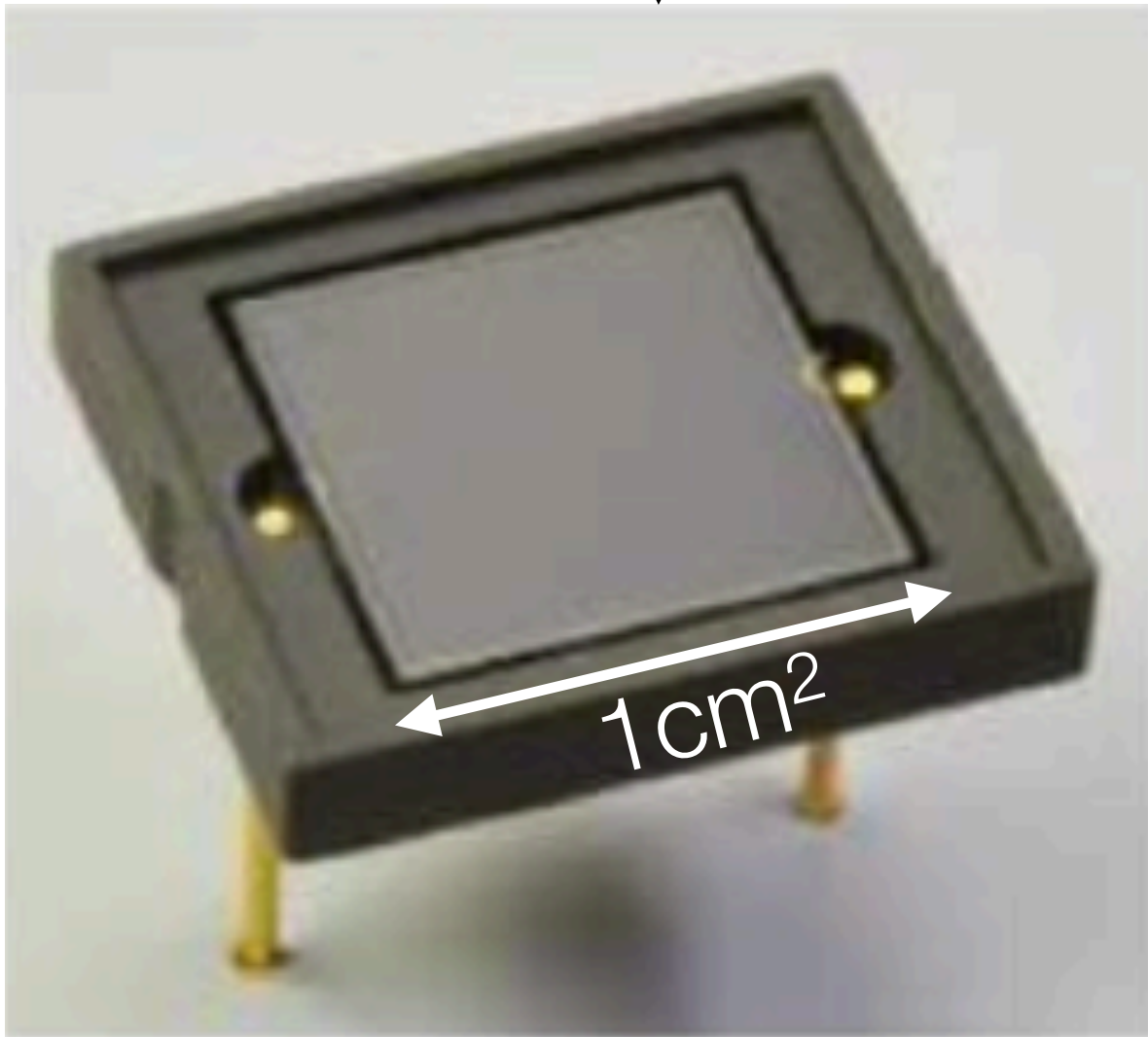




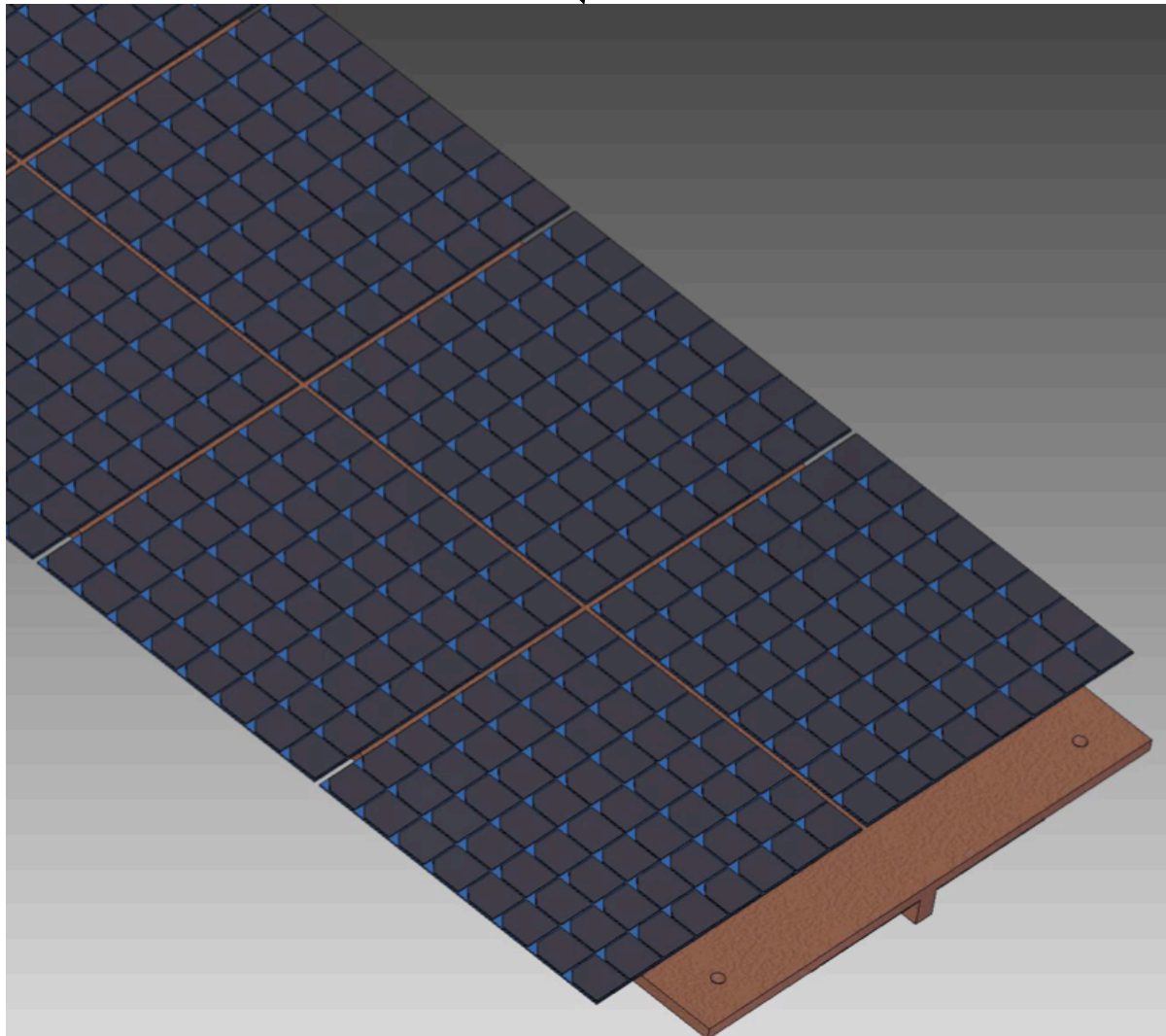
- SiPMs instead of PMTs
- no reflector panels
- Field rings and cathode coated with reflective aluminum deposition (capped by fluoride)



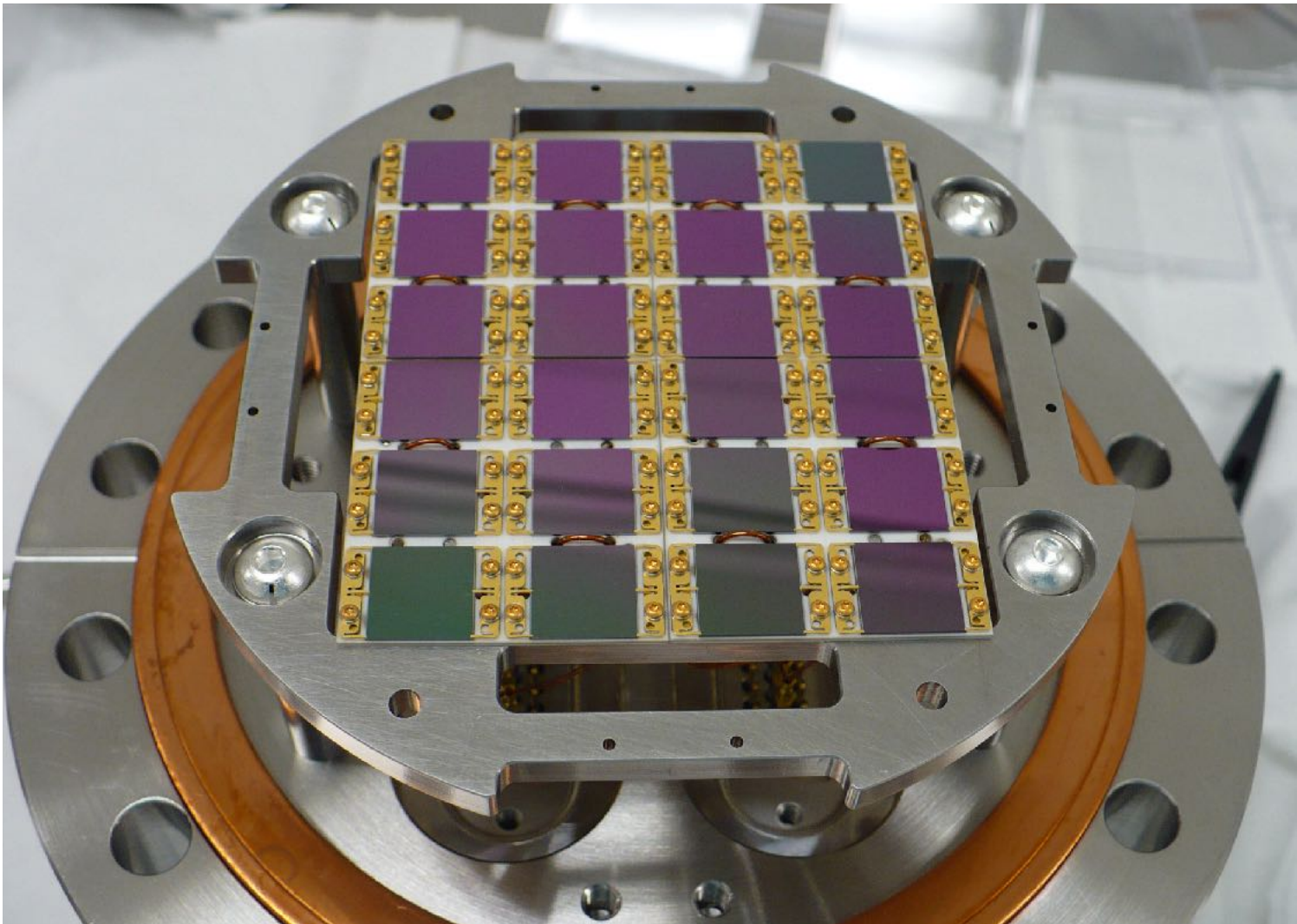
Single Sipm



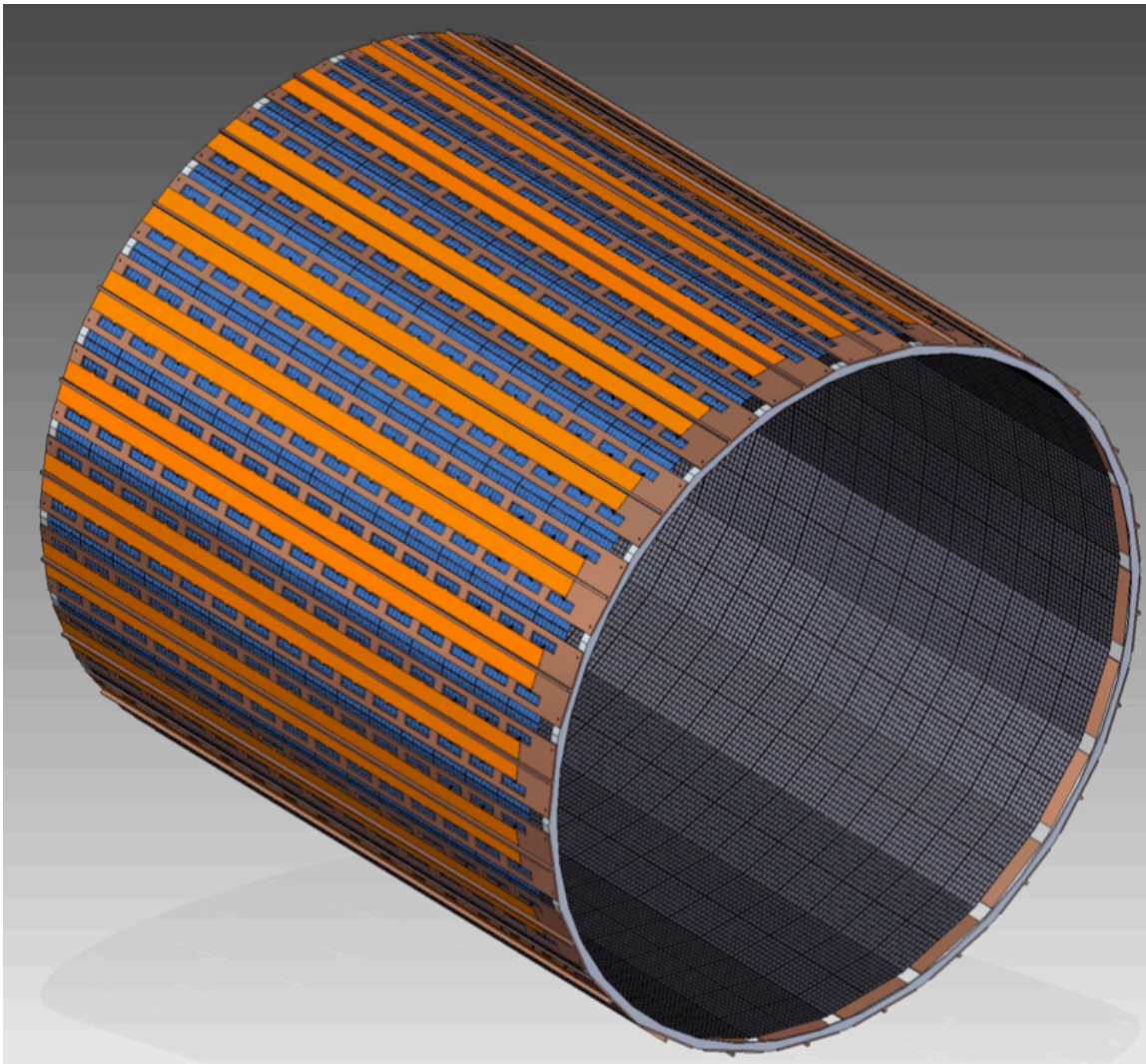
SiPM Stave



SiPM Array



Full light detection system





- Crucial to optimize Photon Transport Efficiency for increased light collection

$$\epsilon = \text{PDE} \cdot \text{PTE}$$

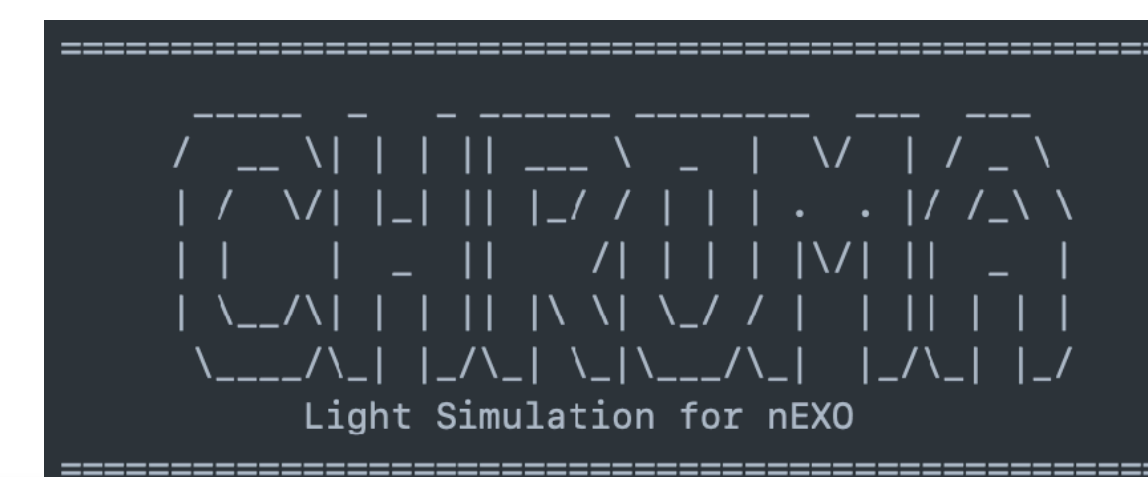
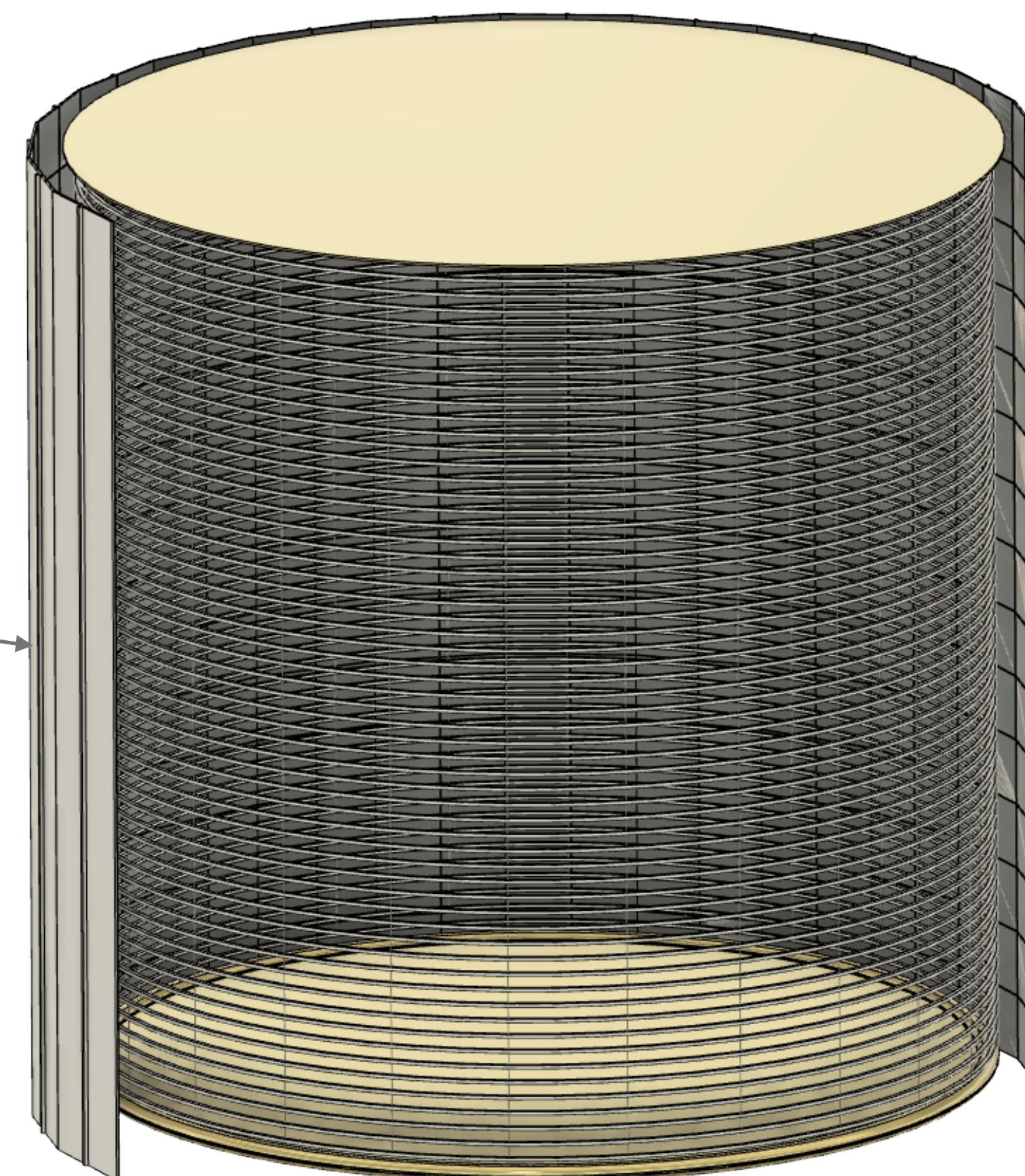
- Highly dependent on reflectivity of TPC components
- Developed new light simulation of nEXO with GPU-based Chroma software

<https://github.com/nEXO-collaboration/chroma>

- > 300x faster

- More detailed geometry

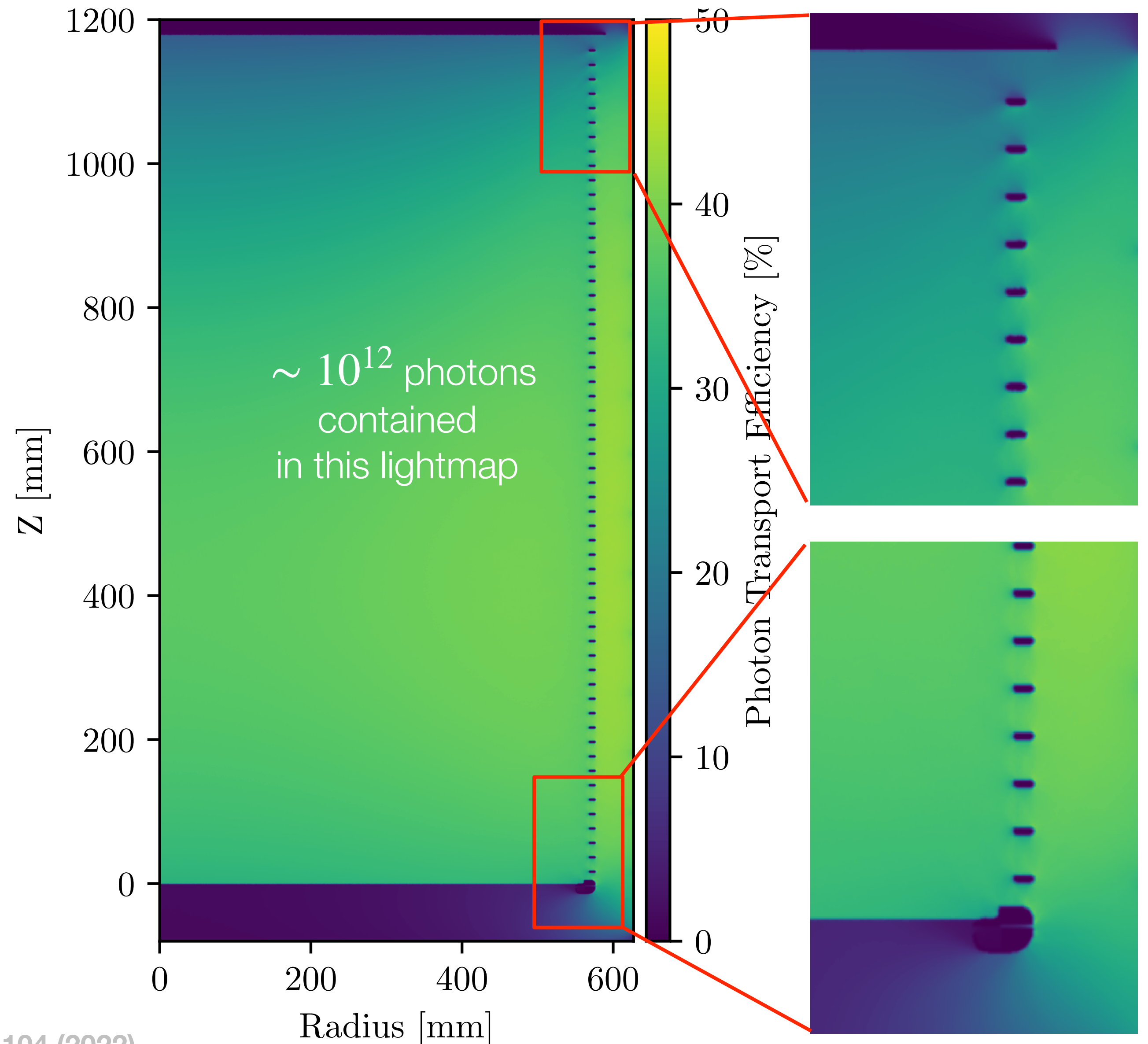
CAD Model of  
nEXO





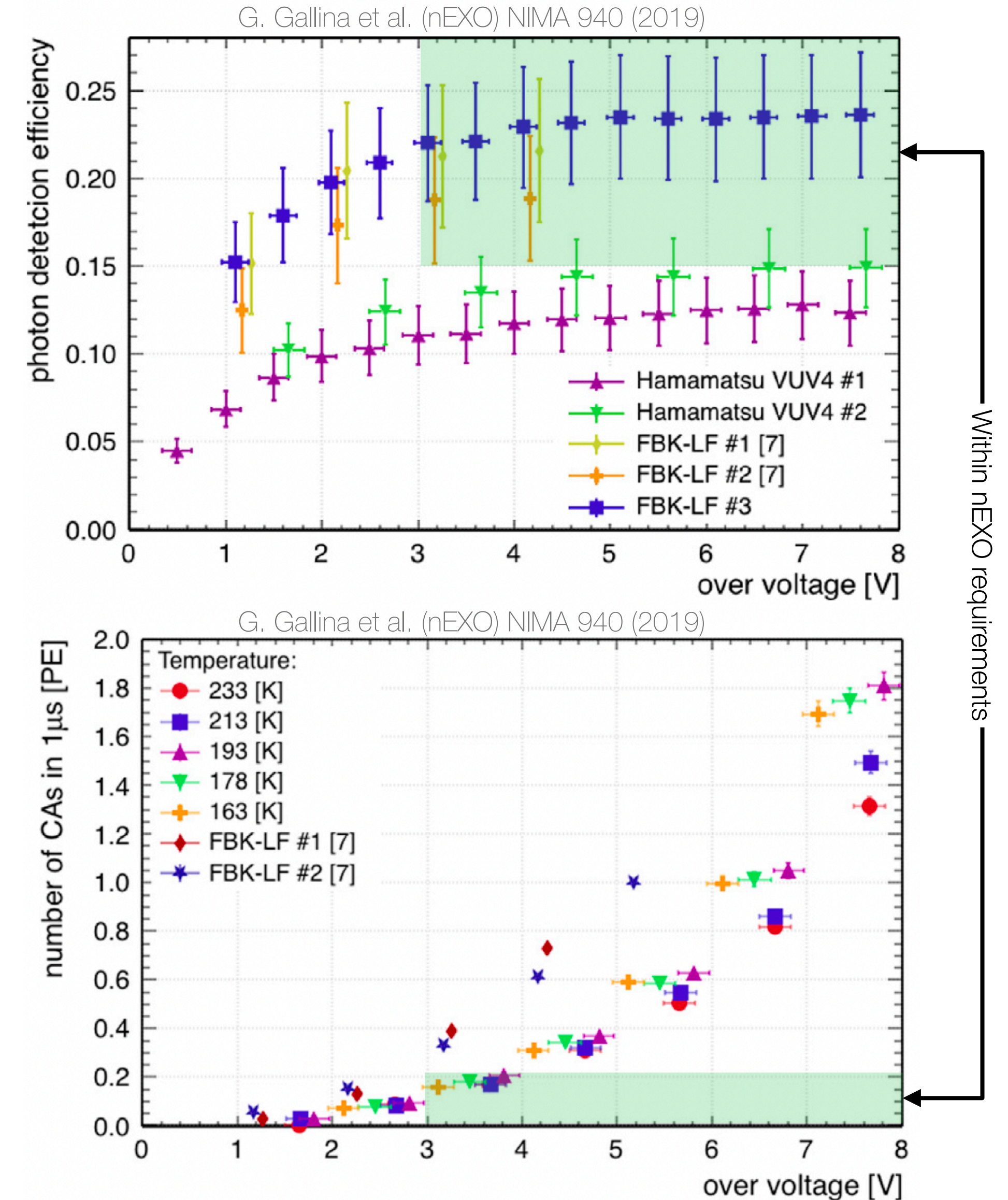
- Varied optical parameters and evaluated systematic error
- Estimated PTE combined with measured PDE results in

$$\epsilon = \text{PDE} \cdot \text{PTE} \approx 6.7 \%$$



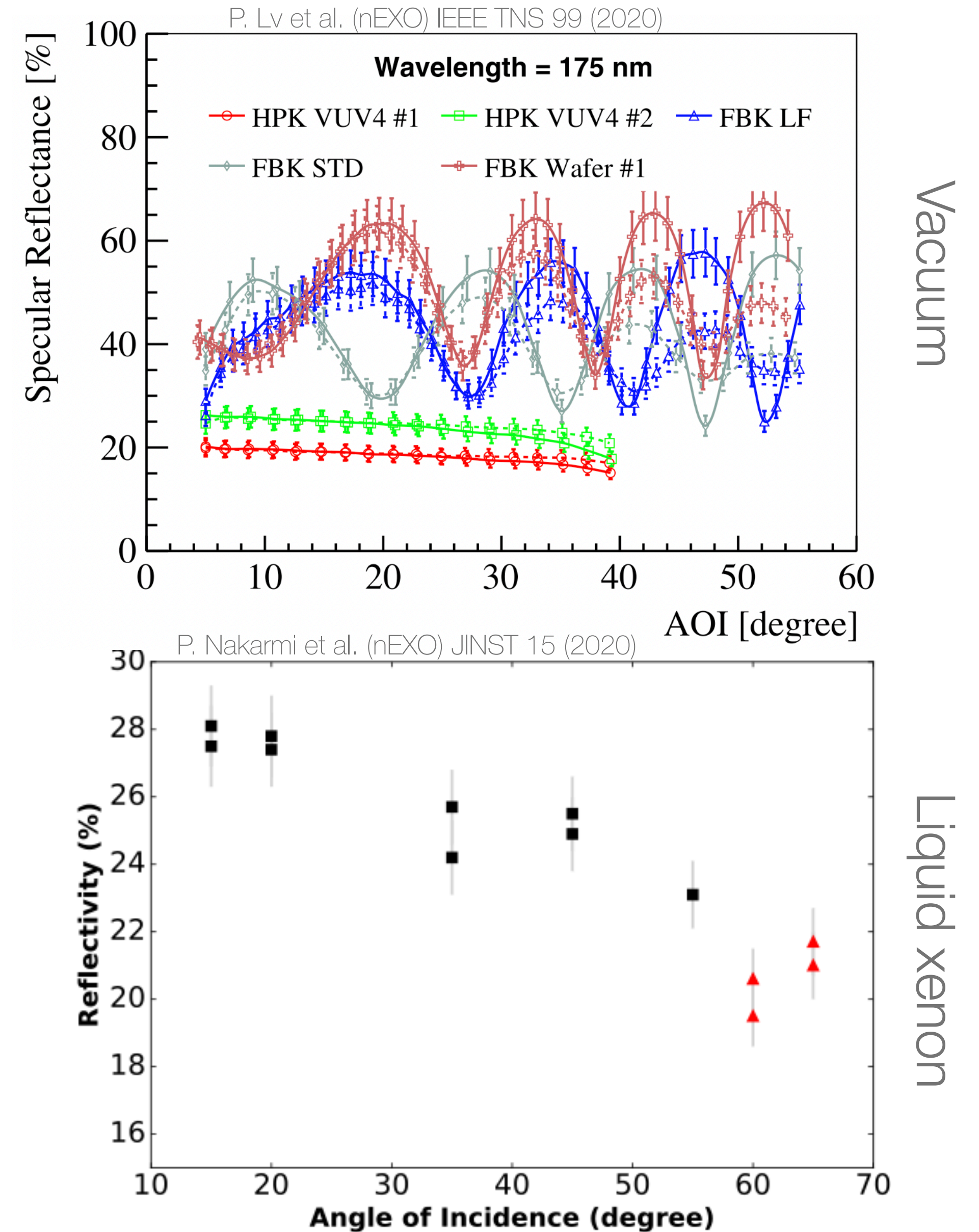


- nEXO is running an extensive characterization campaign with several setups measuring
  - Absolute PDE in vacuum
    - Ostrovskiy et al. (nEXO) IEEE TNS 62 (2015)
    - A. Jamil et al. (nEXO) IEEE TNS 65 (2018)
    - G. Gallina et al. (nEXO) NIMA 940 (2019)
  - Have identified devices from two manufacturers (FBK and HPK) that meet our requirements
  - Working together with vendors to increase operational range
  - More results soon





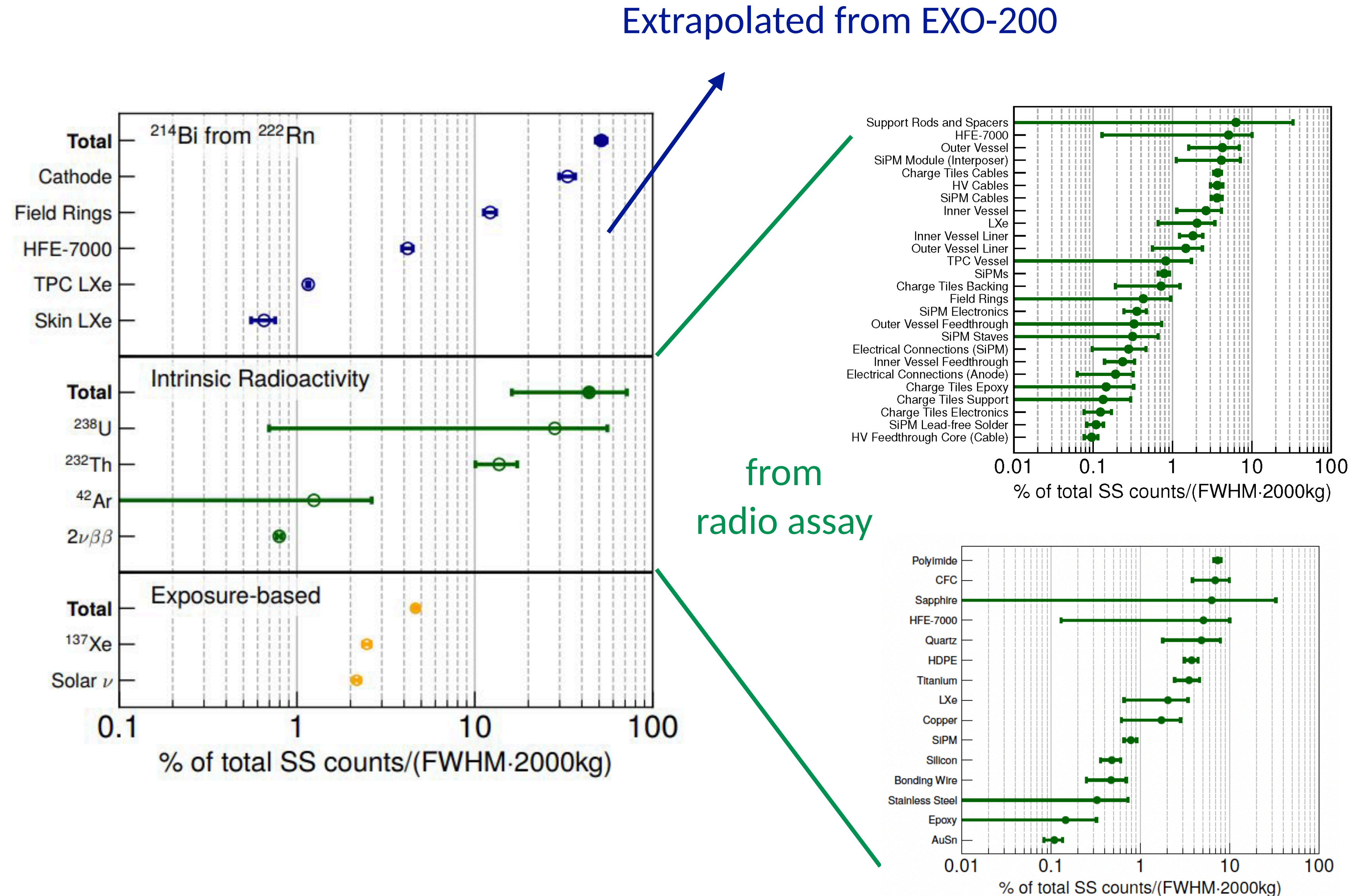
- nEXO is running an extensive characterization campaign with several setups at various institutions
- Reflectivity in vacuum and LXe
  - P. Nakarmi et al. (nEXO) JINST 15 (2020)
  - P. Lv et al. (nEXO) IEEE TNS 99 (2020)
  - M. Wagenpfeil et al. (nEXO) In prep. (2021)
- Photons reflected from SiPM surface can be detected by other SiPMs
- Reflectivity of passive TPC components (e.g. field shaping rings, cathode, ...) crucial for good light collection efficiency





# Robust and validated background model

- Materials screening
- Bottom-up background model
- Optimized: no detector component dominates the bg
- Electro-formed copper for TPC components
- R&D to further reduce radon
- $^{137}\text{Xe}$  can be vetoed with  $\epsilon > 70\%$  at ~no exposure loss





© Cartoonbank.com



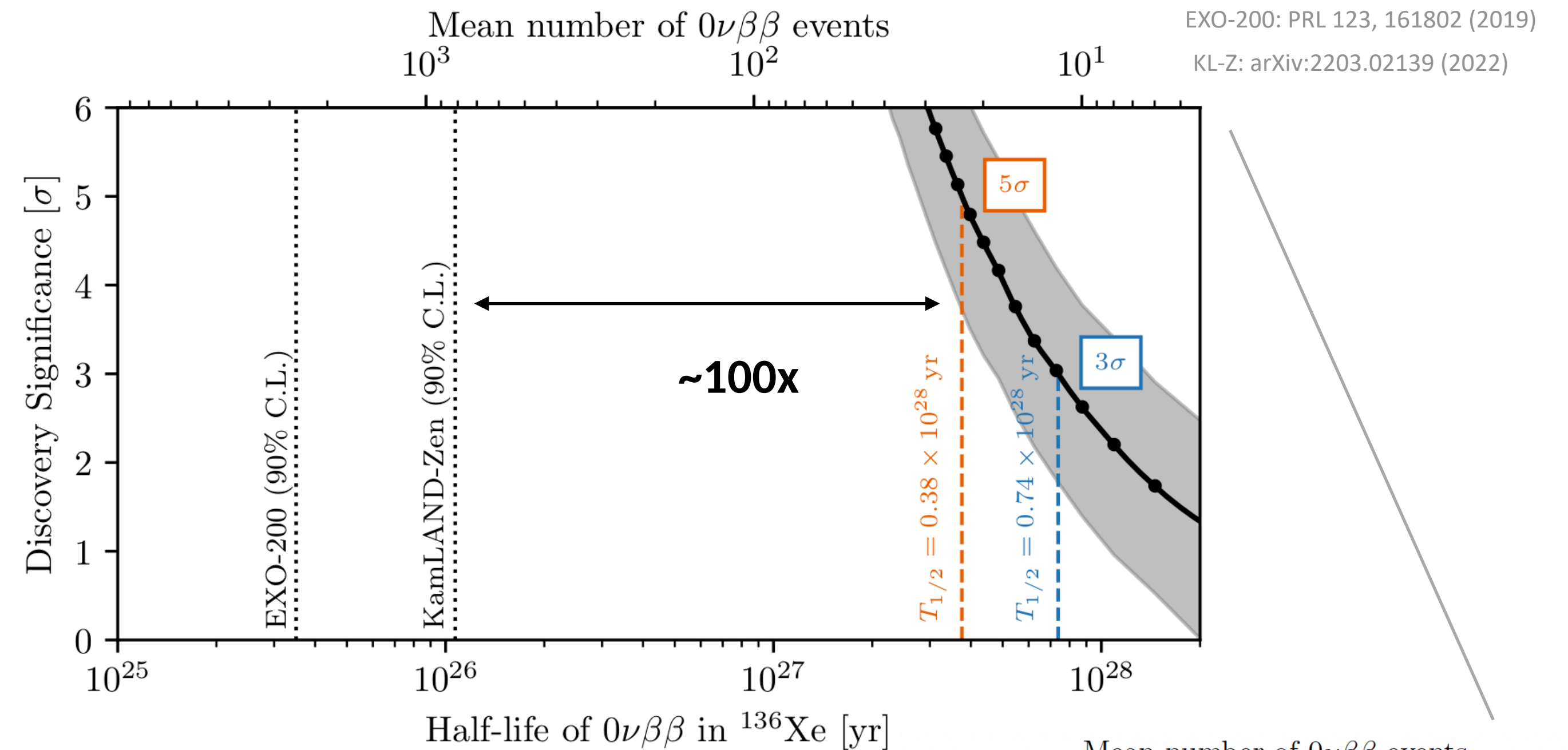
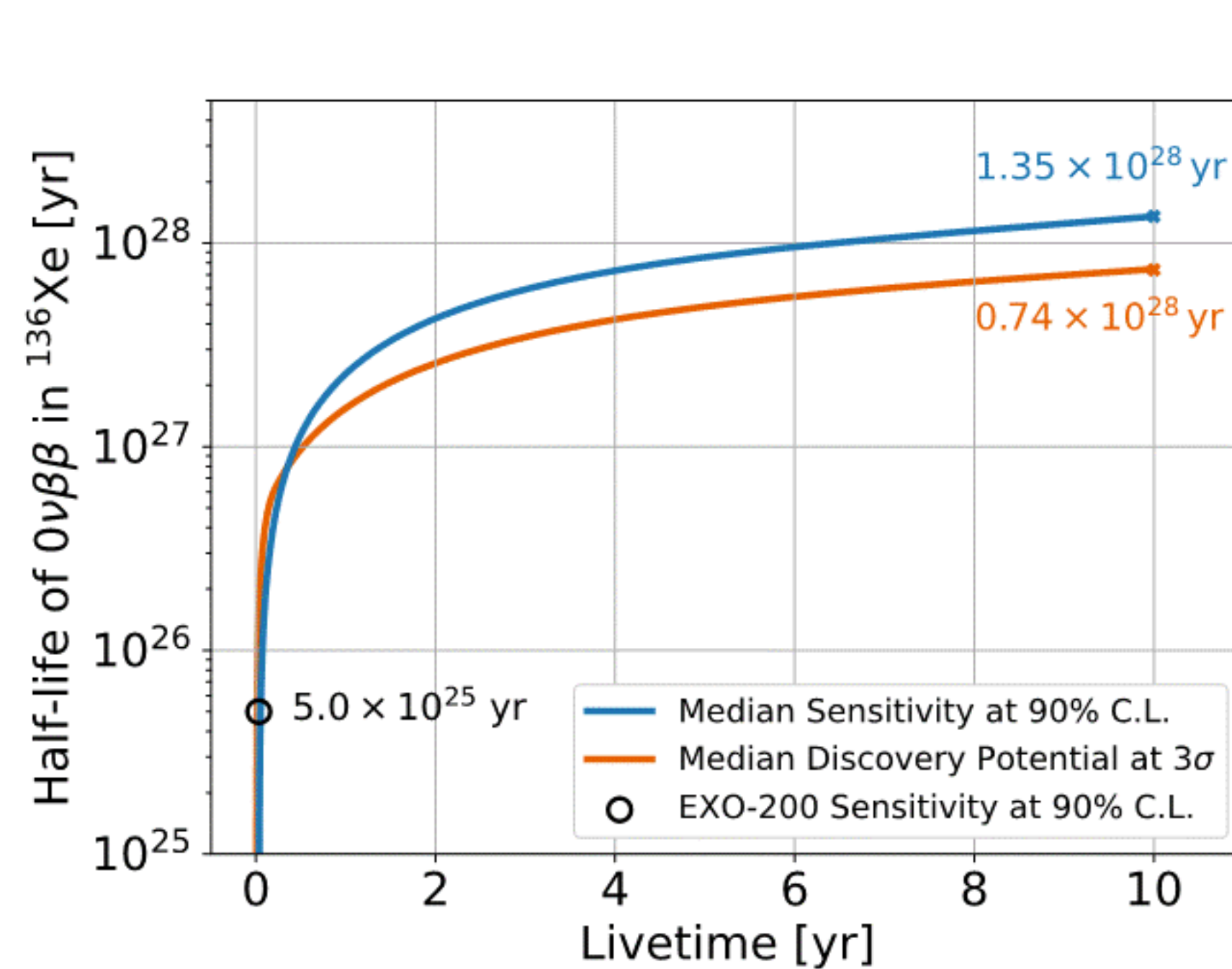
*"And then I thought, What better  
hedge than a ~~uranium~~ centrifuge?"*

*xenon*

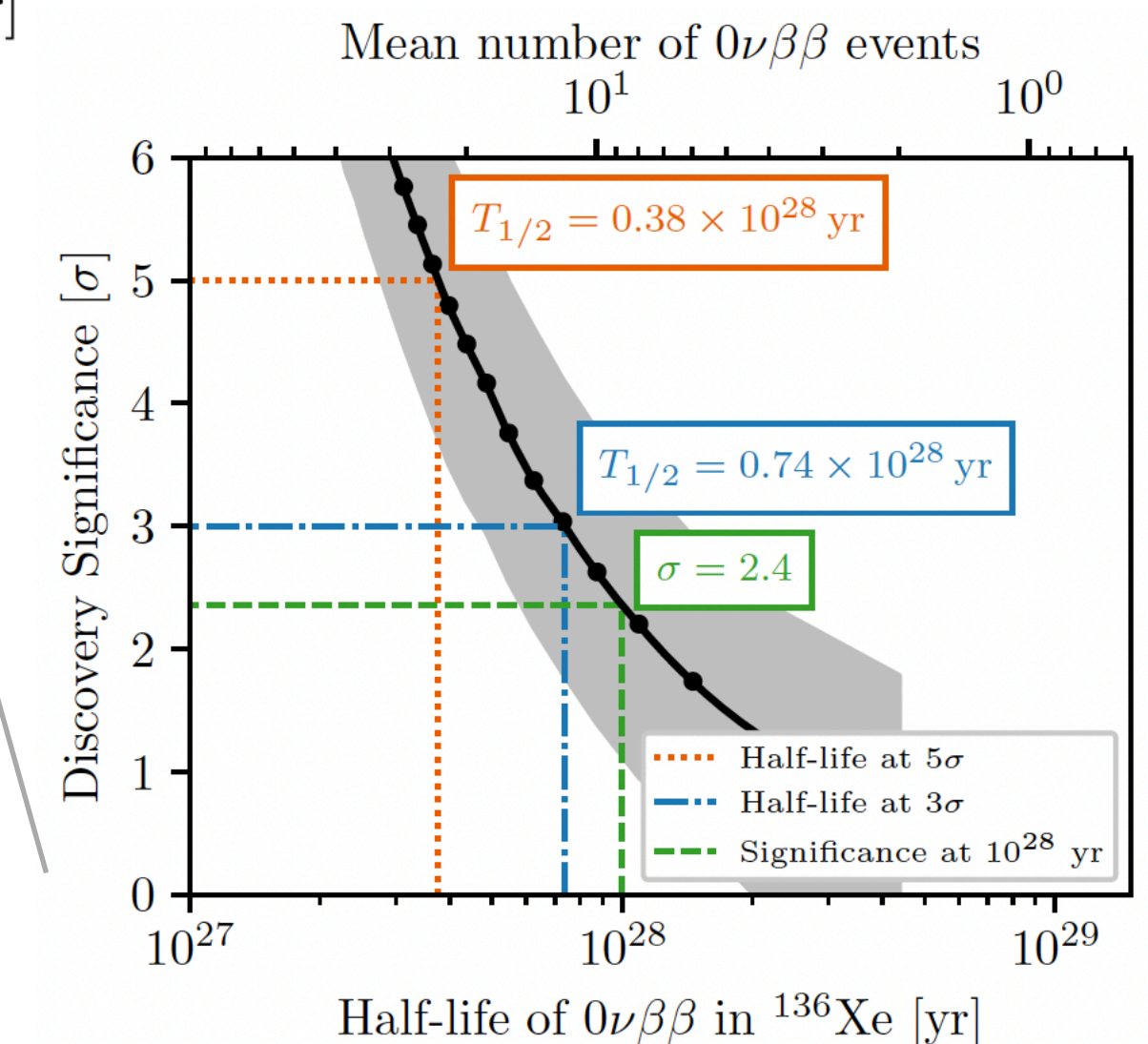
**Xenon enrichment is well understood and cost effective:**

- ▶ EXO-200 used 200 kg of Xe enriched to 80% in 136, at the time a pioneering production.
- ▶ KamLAND-ZEN more recently purchased ~800 kg of xenon enriched to 90% in mass 136 isotope.
- ▶ The nEXO need is only 5x of what already available.
- ▶ nEXO has identified at least two western suppliers each with enough enrichment capacity for the entire production at competitive price.
- ▶ We also have two backup options (neither in eastern Europe)



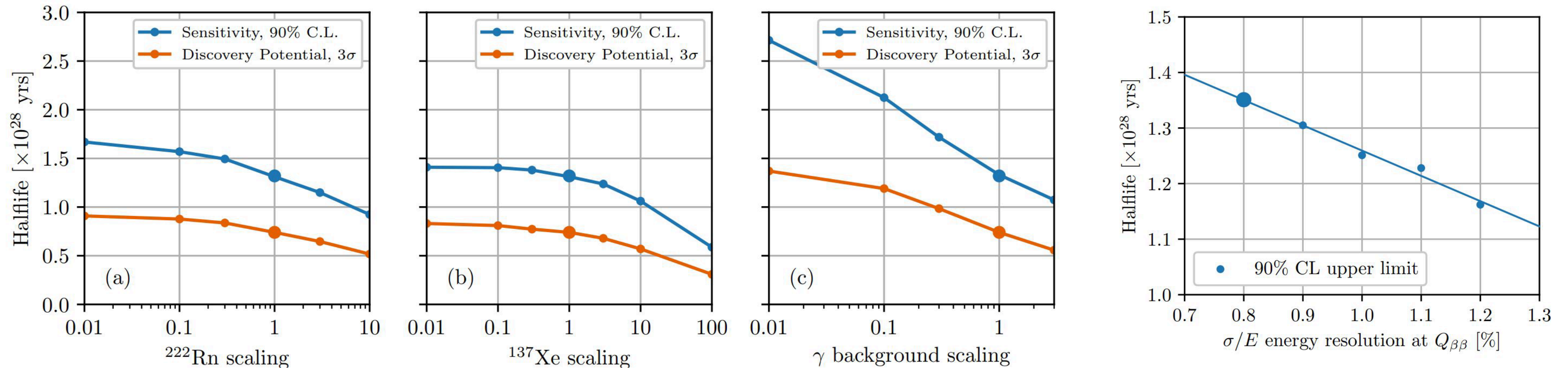


- $>10^{28}$  year sensitivity reach in 10 years
- Can provide compelling evidence of  $0\nu\beta\beta$  decay discovery
- Probes  $m_{\beta\beta} \sim 15$  meV (model and NME dependent)





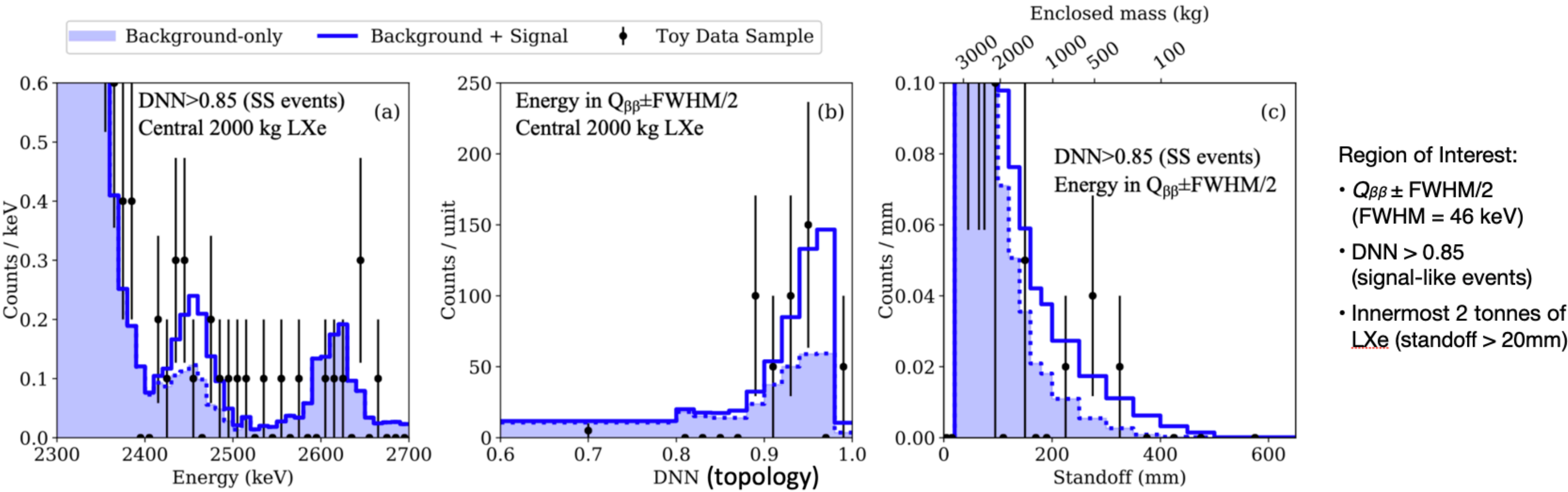
# Robust against background and energy resolution scaling



- nEXO sensitivity is robust against background mis-estimations
- Good data/MC agreement was demonstrated with EXO-200
- Point design (=1):
  - 600 steady state radon atoms (EXO-200 had  $\sim 200$ )
  - $^{137}\text{Xe}$ :  $0.85 \times 10^{-3}$  atoms/kg/year



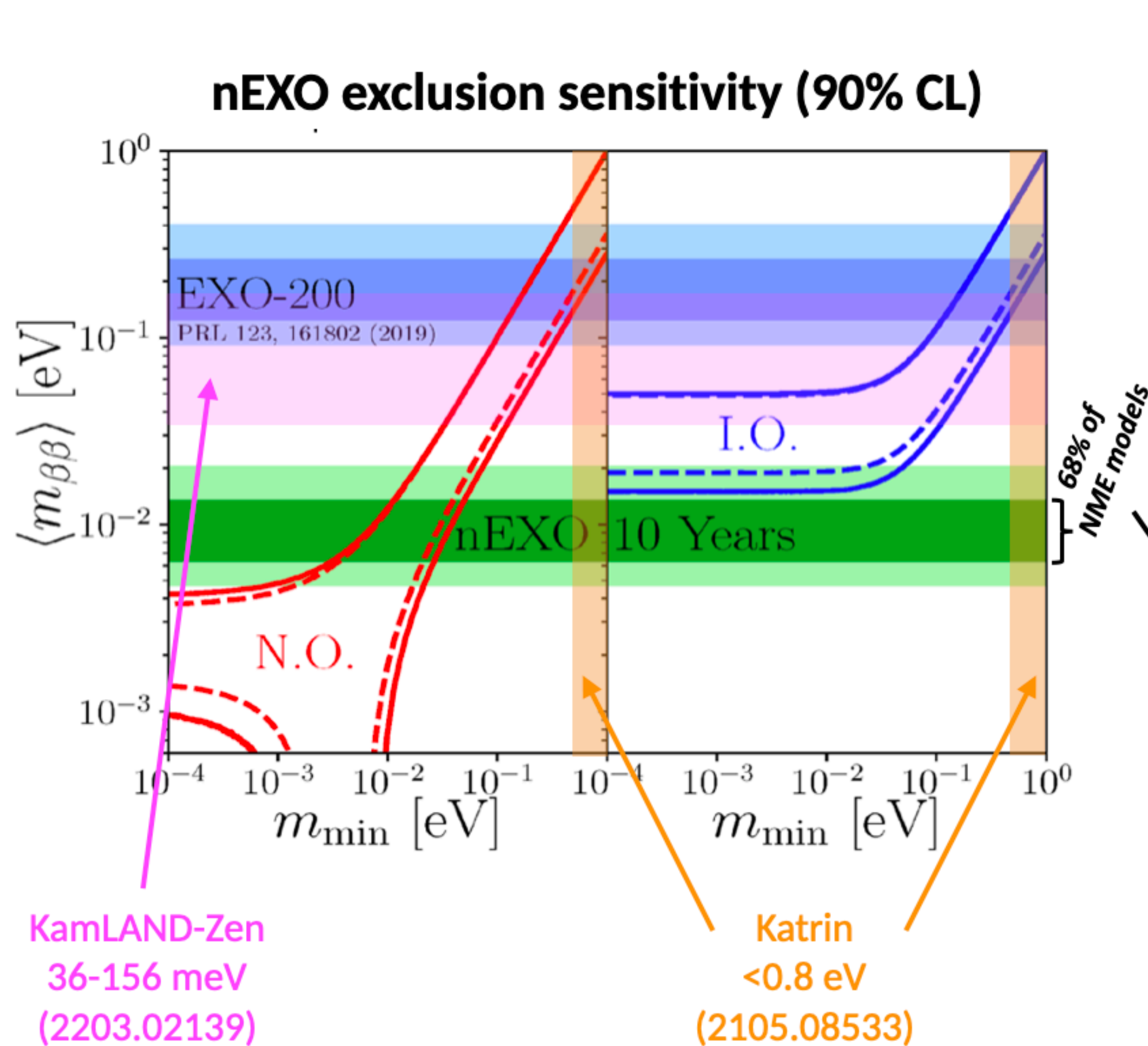
# A robust discovery signal



$$T_{1/2} = 0.74 \times 10^{28} \text{ years}$$

- A positive signal unlikely to be mistaken for an unforeseen background
- A control run with depleted xenon could be run to confirm a discovery signal



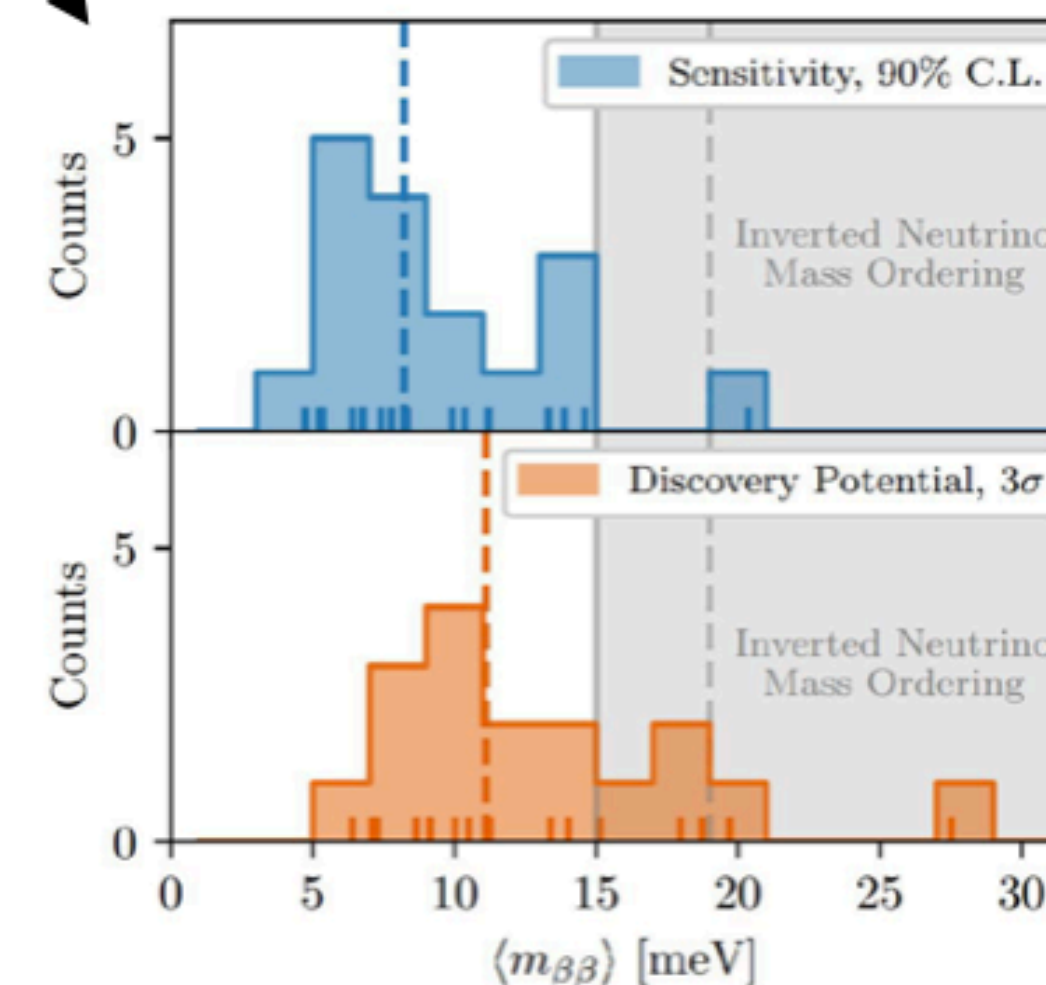


Phase-space factor  
*J. Kotila and F. Iachello,  
Phys Rev C 85, 034316 (2012)*

Axial coupling  
 $g_A = 1.27$

$$\frac{1}{T_{1/2}^0} = \frac{\langle m_{\beta\beta} \rangle^2}{m_e^2} G^{0\nu} g_A^4 |\mathcal{M}^{0\nu}|^2$$

Nuclear Matrix Element



- Agnostic approach: show all published NMEs
- nEXO  $3\sigma$  discovery sensitivity for the median NME is 11.1 meV, below/beyond I.O.



## If nEXO discovers $0\nu\beta\beta$ decay:

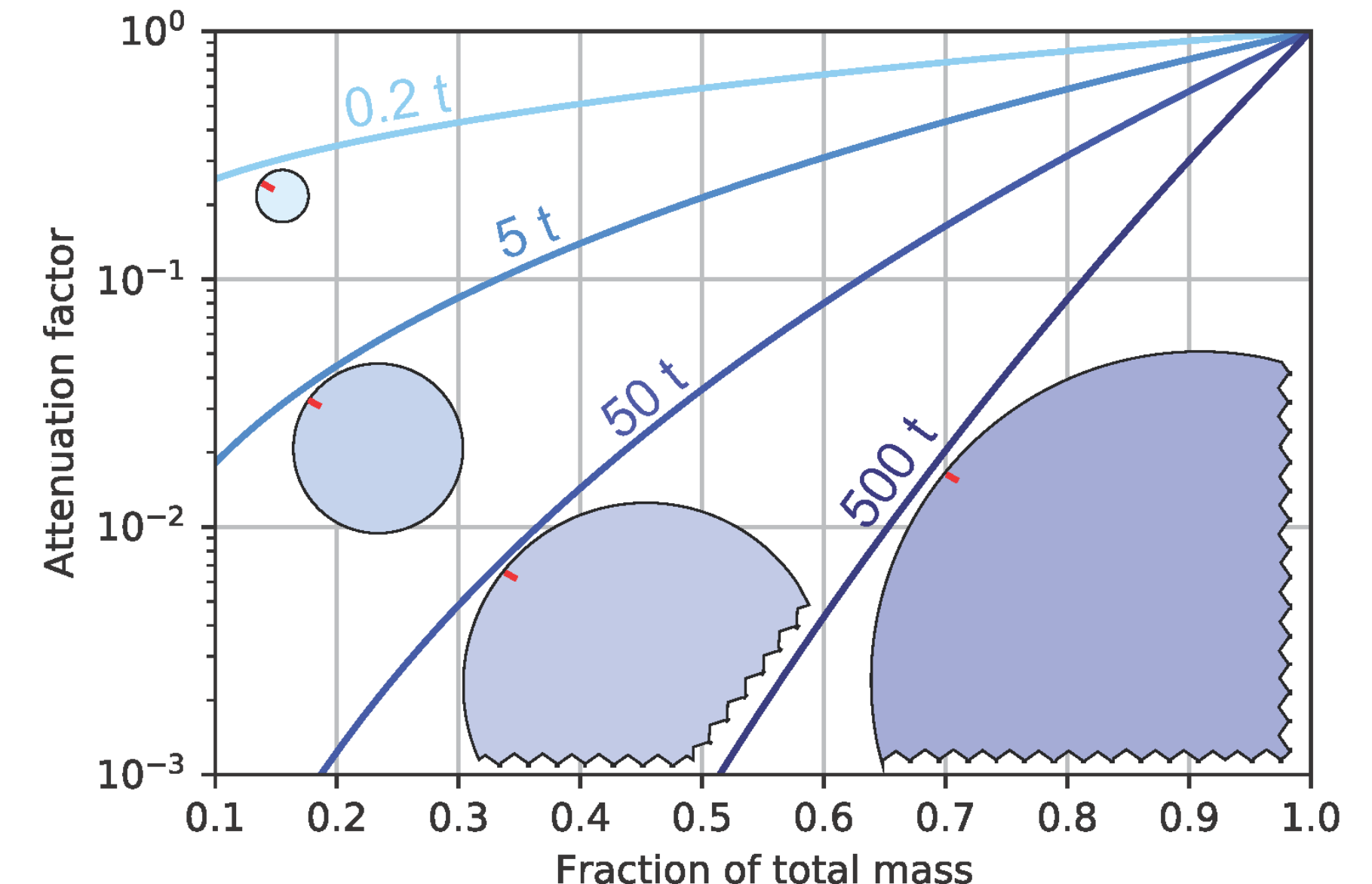
The enriched xenon is NOT “frozen” in a particular detector. Should  $0\nu\beta\beta$  decay be discovered by nEXO, the xenon could be re-used in a different experimental configuration to investigate the underlying physics.

*This is particularly important at the tonne scale, given the cost of the material.*

## If nEXO does not discover $0\nu\beta\beta$ decay:

The advantages of the homogeneous detector keep improving with size. Should  $0\nu\beta\beta$  decay not be discovered by nEXO, larger detectors using the same technology are plausible. There is enrichment capacity for this, although the feed stock will need to be directly extracted from air. Again, this is plausible, with proper planning.

*A clear avenue for the future.*





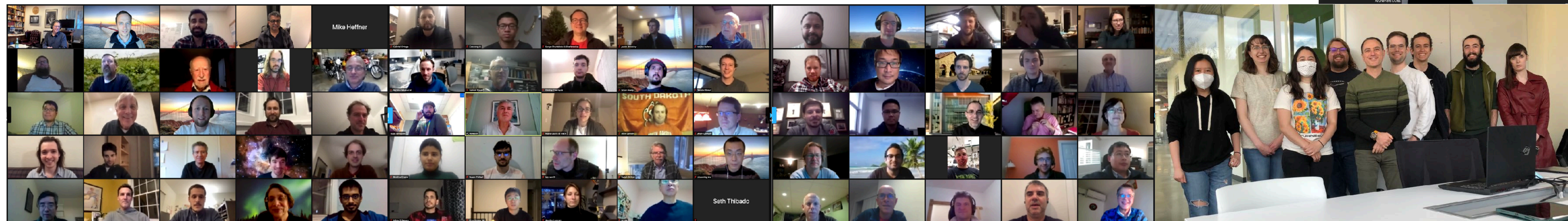
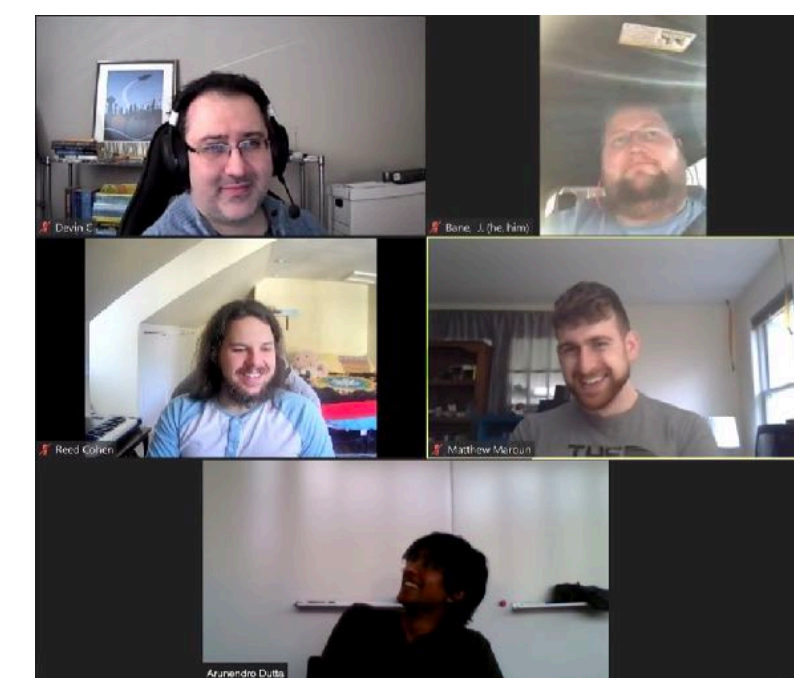


# Summary and Outlook

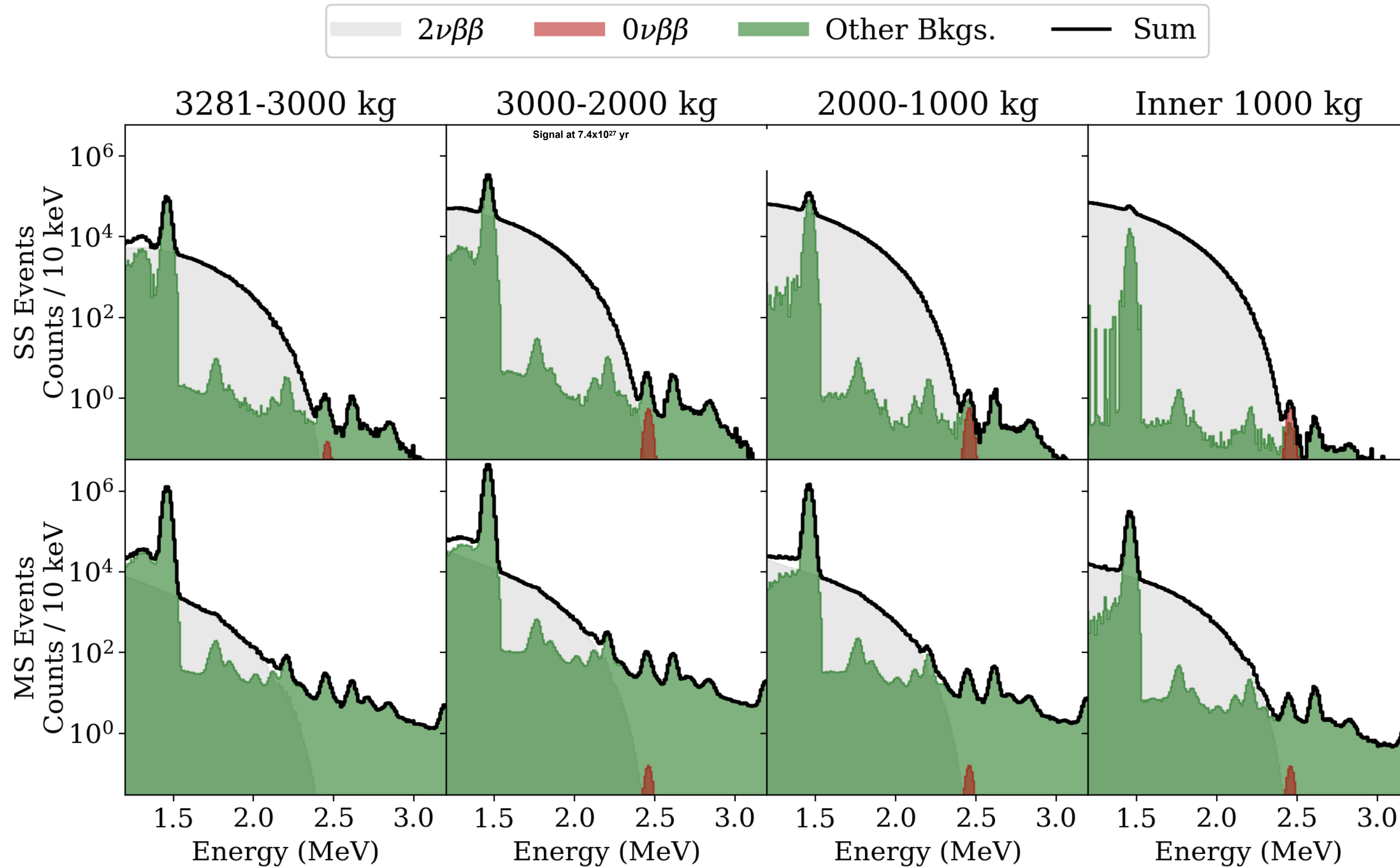


- nEXO will explore beyond the current  $0\nu\beta\beta$  decay sensitivity 100-fold
  - Are neutrinos their own antiparticles?
  - Origin of neutrino mass
  - Matter/antimatter asymmetry
- 5 tonnes, enriched LXe TPC
  - Low background design rooted in EXO-200 experience
  - Robust bg model, powerful bg discrimination
  - Benefits from and contributes to rapid evolution of LXe technology

- URL: [nexo.llnl.gov](https://nexo.llnl.gov)
- Twitter: @nEXOexperiment
- Code of Conduct: <https://nexo.llnl.gov/diversity-equity-and-inclusion>





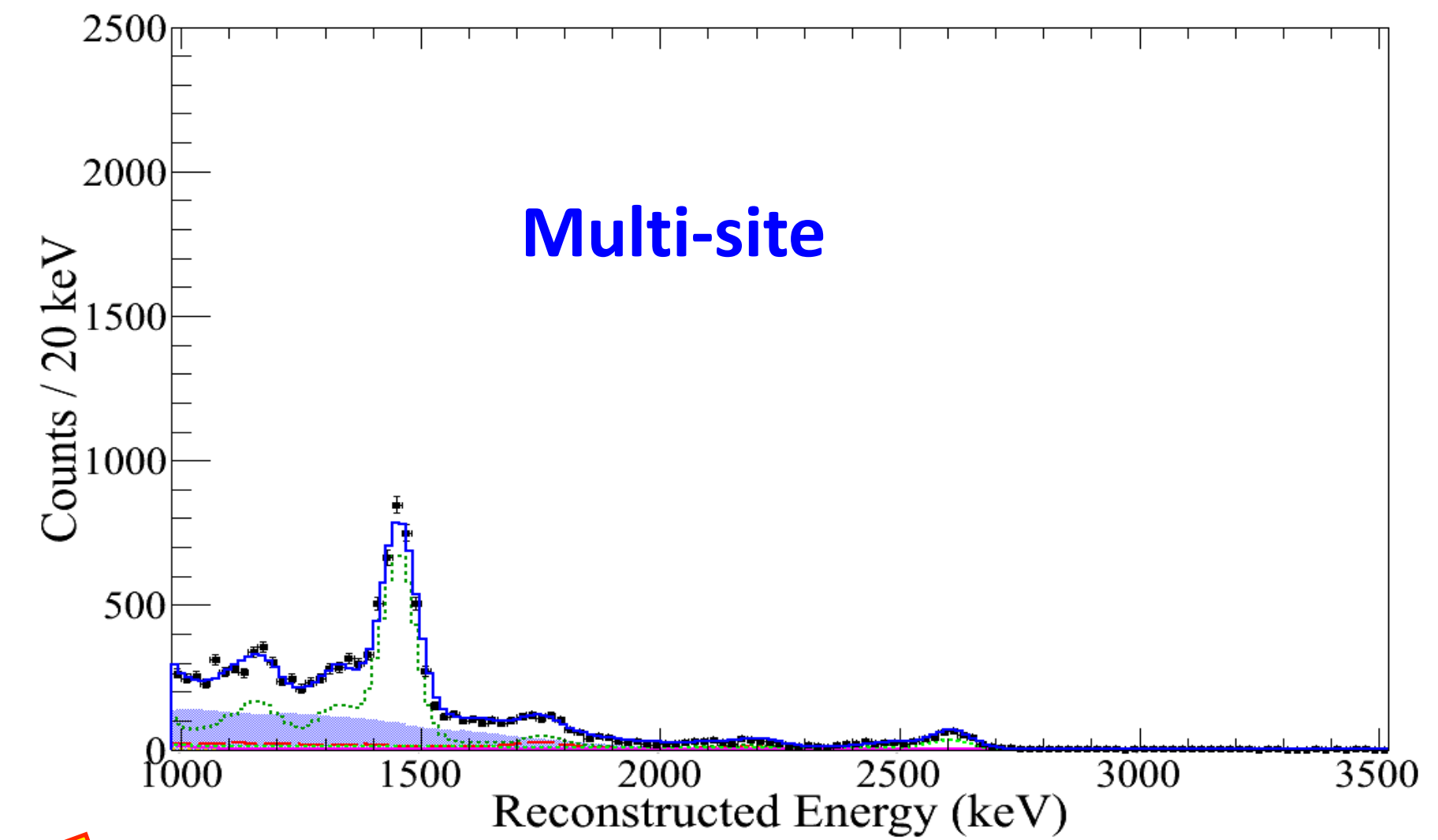
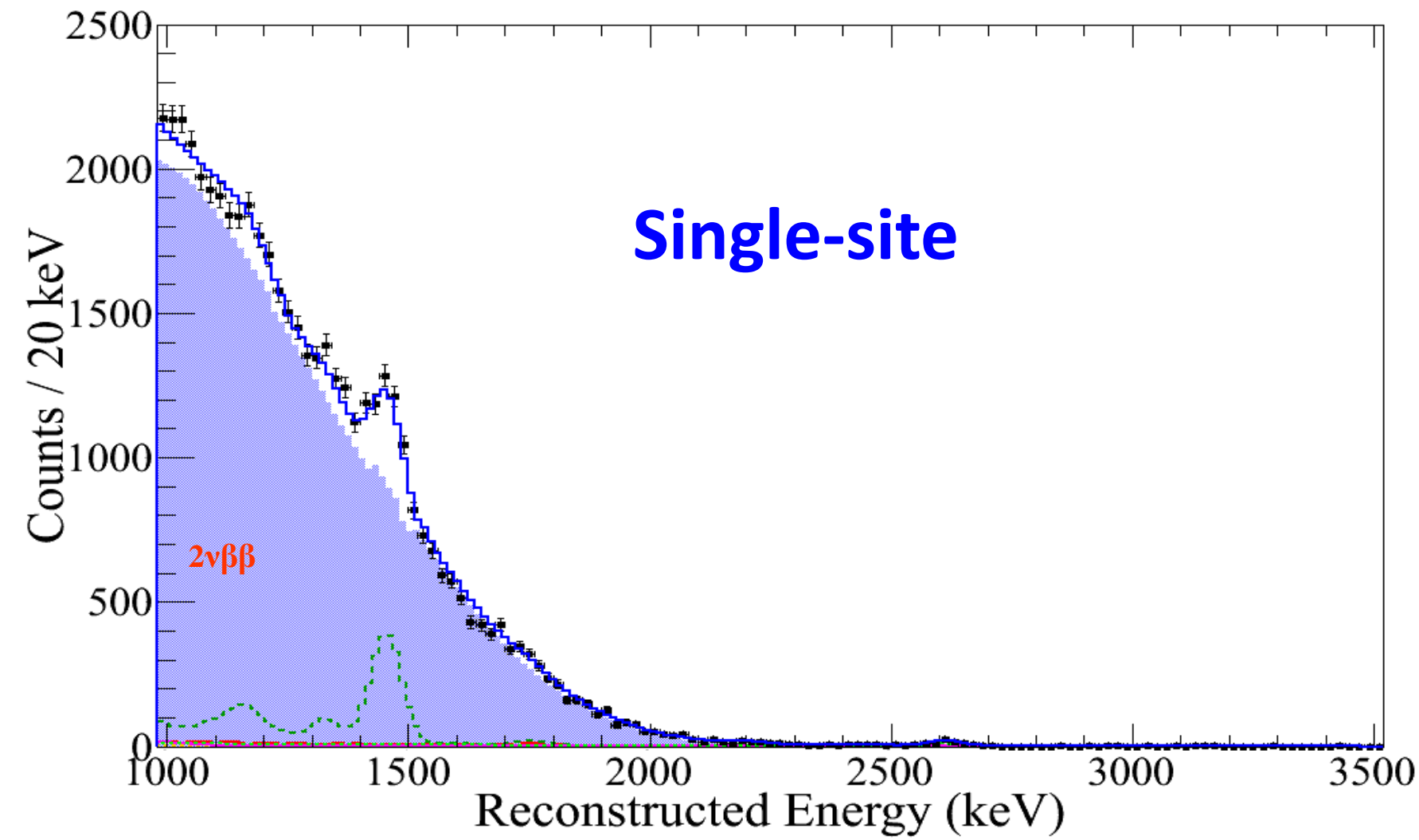




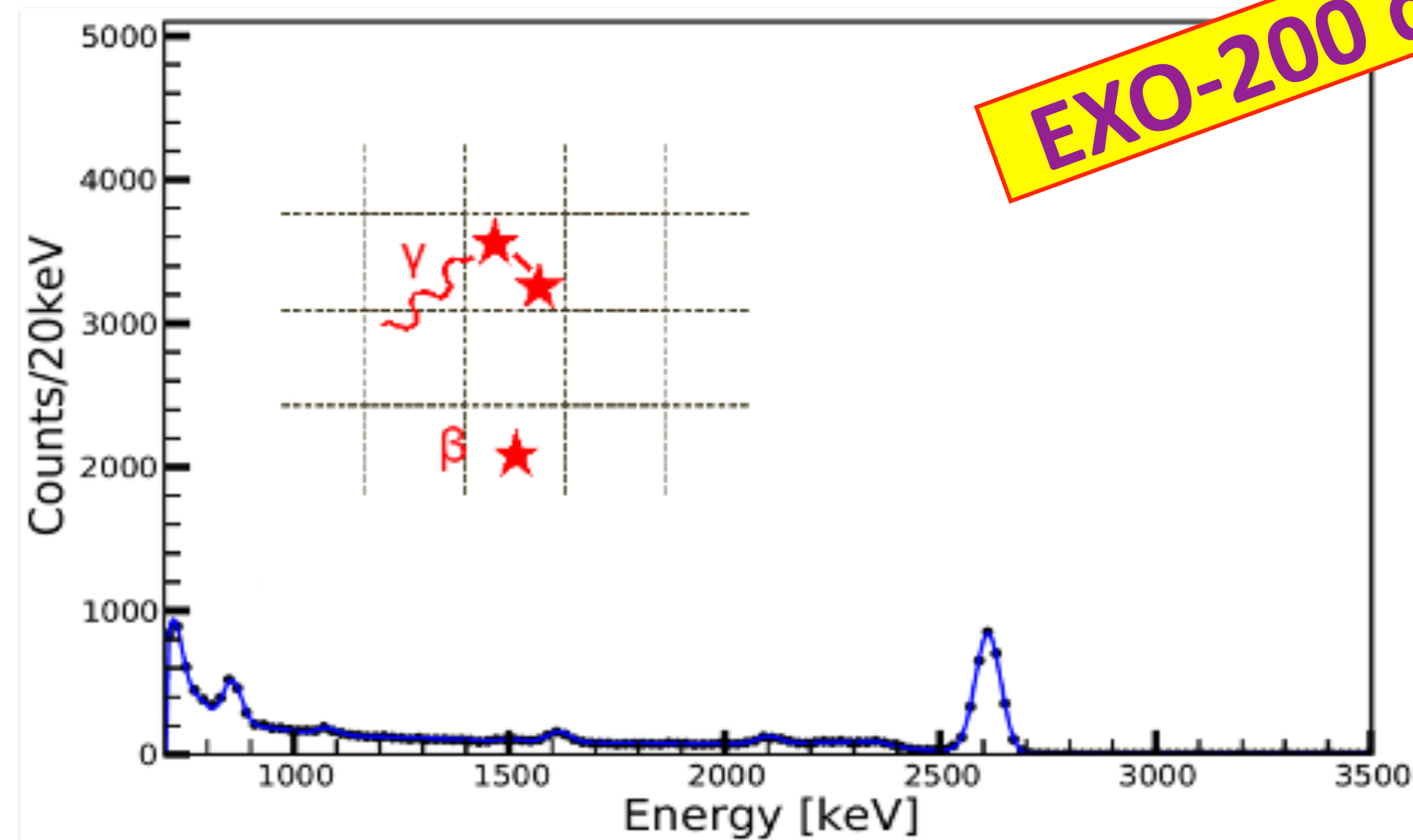
# Event topology and background discrimination in EXO-200



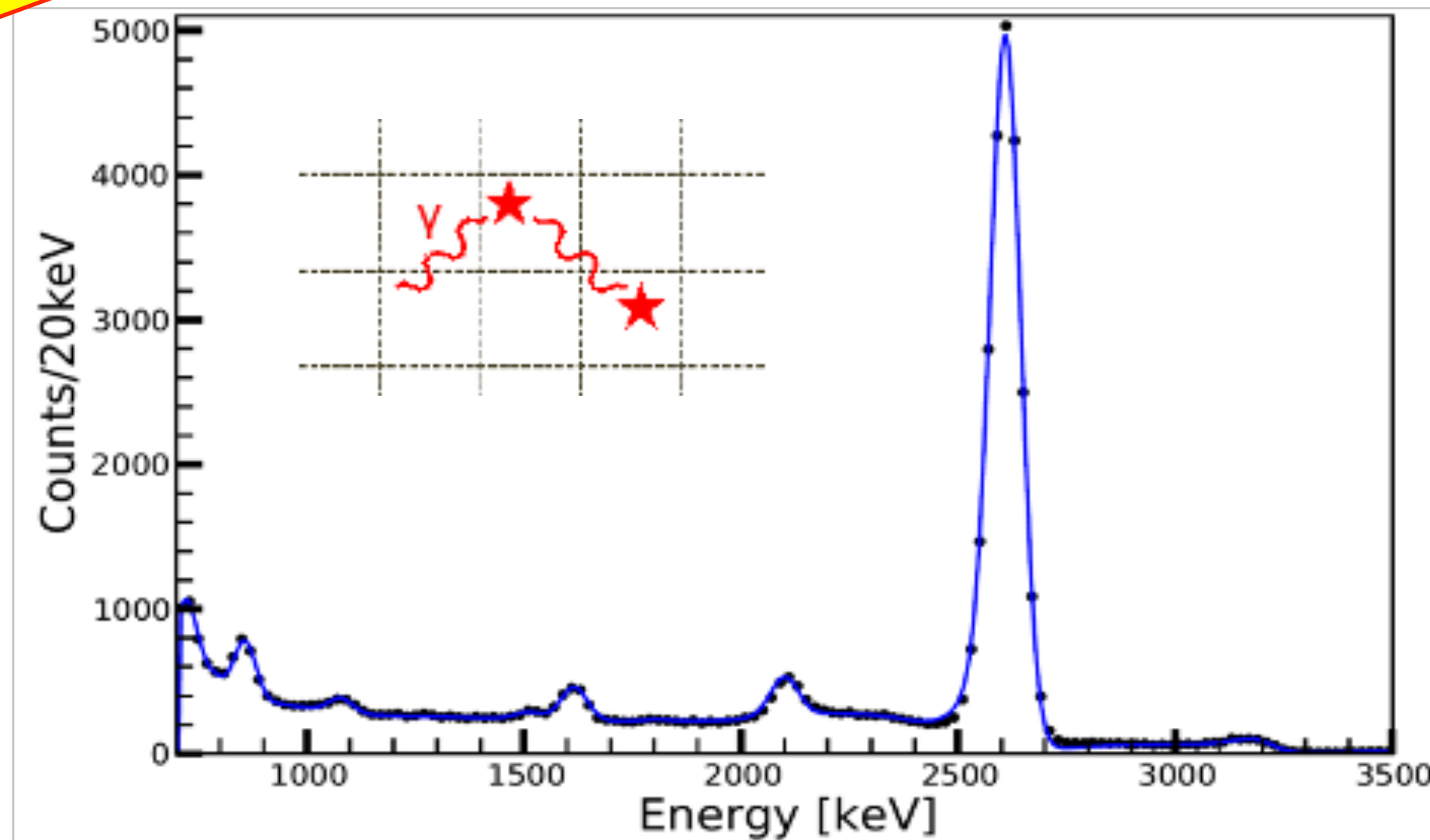
Low background  
data



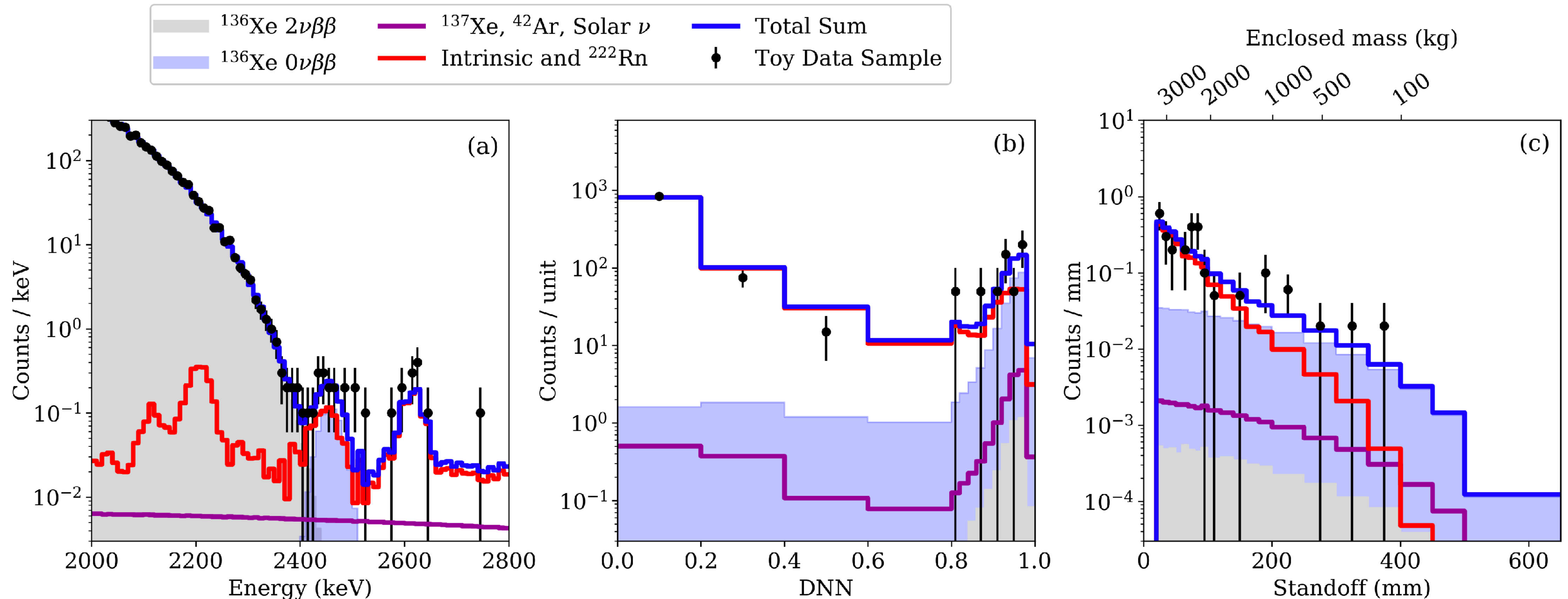
$^{228}\text{Th}$  calibration  
source



EXO-200 data



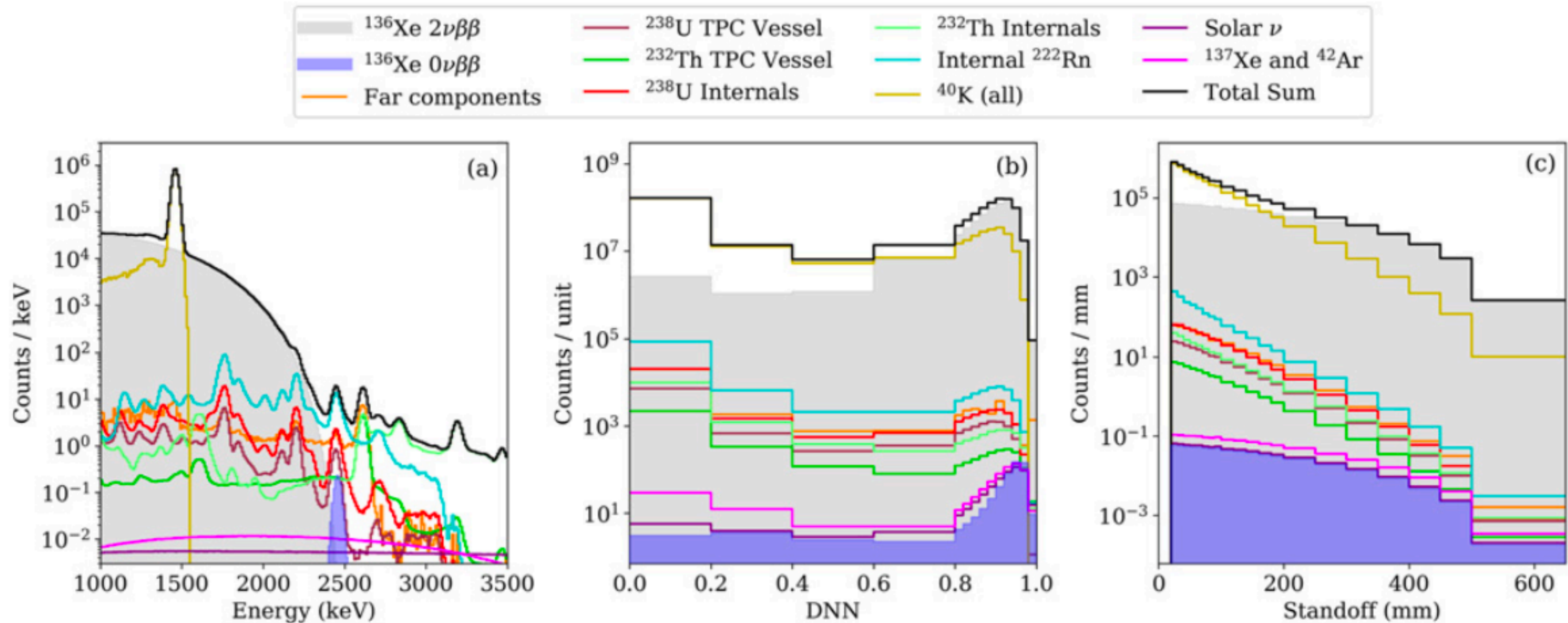




**Figure 11.** Event distributions for an example of toy dataset (black points) and combined groups of the fitted PDFs projected onto the three axes used in the sensitivity analysis. In (a) the event energy distribution is shown for SS-like events ( $\text{DNN} > 0.85$ ) in the central 2000 kg LXe and in the 2000–2800 keV region; (b) the DNN  $0\nu\beta\beta$  discriminator distribution is shown for events with energy within  $Q_{\beta\beta} \pm \text{FWHM}/2$  and in the same central volume; and (c) the standoff distance distribution is shown for SS-like events within  $Q_{\beta\beta} \pm \text{FWHM}/2$ . The  $0\nu\beta\beta$  decay signal corresponds to a half-life of  $0.74 \times 10^{28}$  yr.

Signal:  $0.7 \times 10^{28}$  y





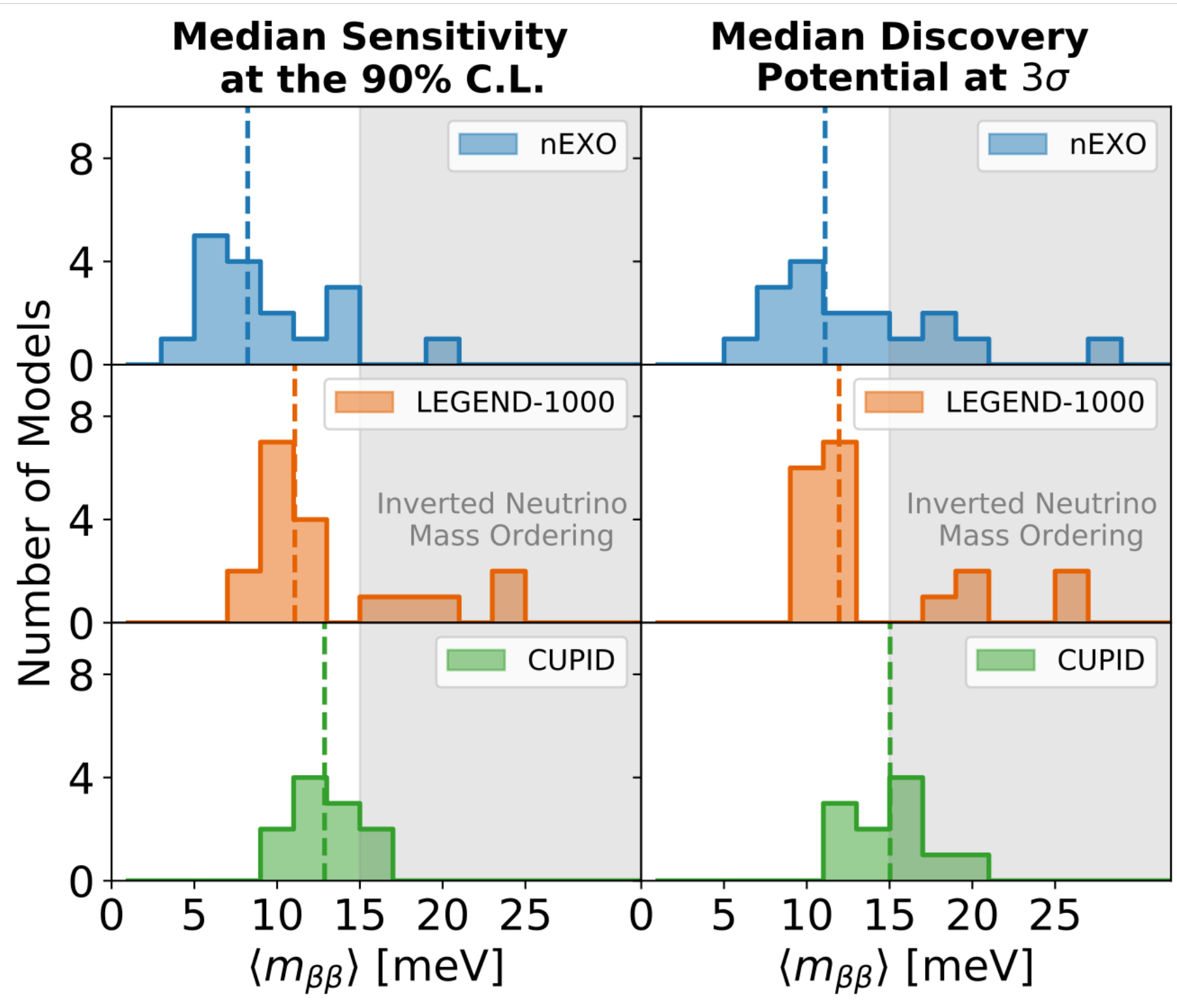
**Figure 10.** Nominal model of event distributions in nEXO, projected onto each of the three axes used in the sensitivity analysis: (a) event energy, (b) DNN  $0\nu\beta\beta$  discriminator, and (c) standoff distance. The  $0\nu\beta\beta$  decay signal corresponds to a half-life of  $0.74 \times 10^{28}$  yr.

Signal:  $0.7 \times 10^{28}$  y

*nEXO Sensitivity and Discovery Potential: J. Phys. G: Nucl. Part. Phys. 49, 015104 (2022)*



Method	Year	Citation
IBM	2015	<a href="#">PRC 91, 034304 (2015)</a>
NSM	2008	<a href="#">PRL 100, 052503 (2008)</a>
IBM	2020	<a href="#">PRD 102, 095016 (2020)</a>
QRPA	2014	<a href="#">PRC 89, 064308 (2014)</a>
NSM	2016	<a href="#">PRC 93, 024308 (2016)</a>
QRPA	2015	<a href="#">PRC 91, 024613 (2015)</a>
QRPA	2018	<a href="#">PRC 98, 024608 (2018)</a>
NSM	2018	<a href="#">JPS Conf. Proc. 23, 012036 (2018)</a>
QRPA	2013	<a href="#">J. High Energ. Phys. 2013, 25 (2013)</a>
QRPA	2013	<a href="#">PRC 87, 064302 (2013)</a>
QRPA	2013	<a href="#">PRC 87, 045501 (2013)</a>
QRPA	2018	<a href="#">PRC 97, 034315 (2018)</a>
QRPA	2010	<a href="#">Nucl.Phys.A 847 (2010) 207</a>
EDF	2013	<a href="#">PRL 111, 142501 (2013)</a>
EDF	2015	<a href="#">PRC 91, 024316 (2015)</a>
QRPA	2018	<a href="#">PRC 97, 045503 (2018)</a>
EDF	2017	<a href="#">PRC 96, 054310 (2017)</a>
QRPA	2015	<a href="#">PRC 91, 024613 (2015)</a>
EDF	2010	<a href="#">Prog.Part.Nucl.Phys. 66 (2011) 436</a>



	[meV] ( <i>median NME</i> )	
	90% excl. sens.	discov. potential
nEXO	8.2	11.1
LEGEND	11.1	12.0
CUPID	12.9	15.0

**\* $T_{1/2}$  values used [ $\times 10^{28}$  yr]:**  
nEXO: 1.35 (90% sens.), 0.74 ( $3\sigma$  discov.) [1]  
LEGEND: 1.4 (90% sens.), 1.2 ( $3\sigma$  discov.) [2]  
CUPID: 0.15 (90% sens.), 0.11 ( $3\sigma$  discov.) [3]