

Population and decay of states of ^{12}C

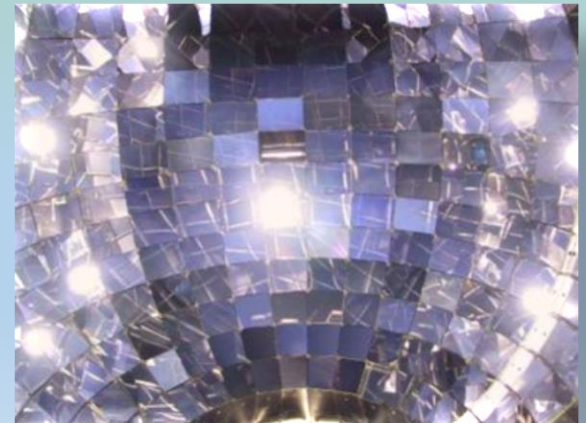
IWM-EC 2021

G. Cardella

INFN Sezione di Catania

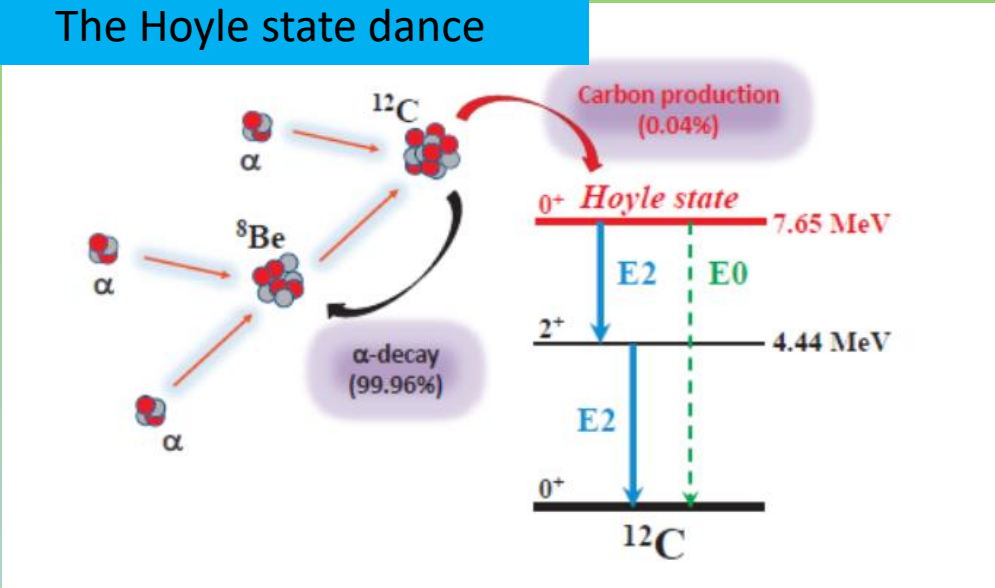
For Chirone collaboration

IWM-EC 2021 November 24th CAEN

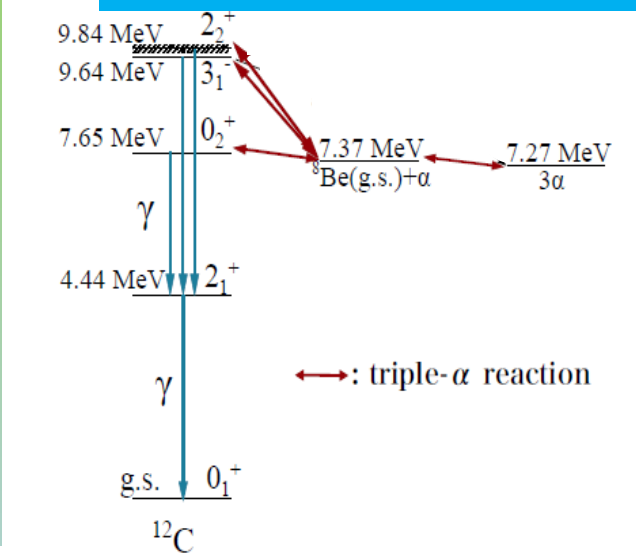


Production of Carbon in stars: The role of Hoyle and 9.64 MeV states γ -decay

T8 - 100 million of kelvin deg
The Hoyle state dance



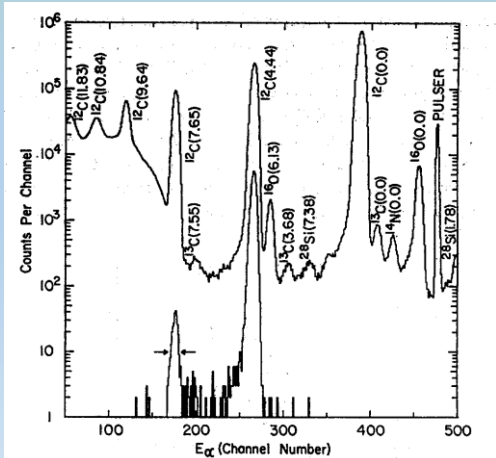
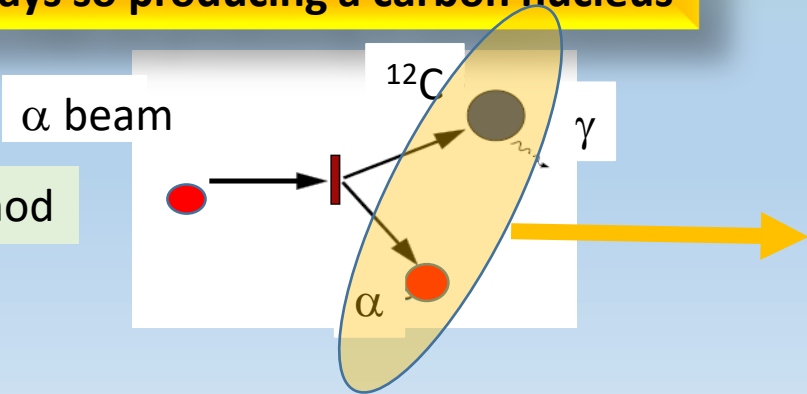
T9 - 1 Billion of kelvin deg
Other levels come into play



Only about 4 times each 10000 events (R.G.Markham et al Nucl.Phys.A270(1976)489) the Hoyle state decays trough the emission of two γ -rays so producing a carbon nucleus

$$\Gamma_{\text{rad}}/\Gamma < 8.2 \times 10^{-7} \text{ (95\%C.L.)}$$

Chamberlin method



STATUS OF THE ART: recent attempts to improve results

T. Kib'edi et al PHYSICAL REVIEW LETTERS 125, 182701 (2020)

$p+^{12}\text{C}$ reaction at 10.7 MeV
 $p\text{-}\gamma\text{-}\gamma$ coincidence at OSLO cyclotron
 p measured with SIRI strip array – γ with CACTUS array of 26 NaI detectors

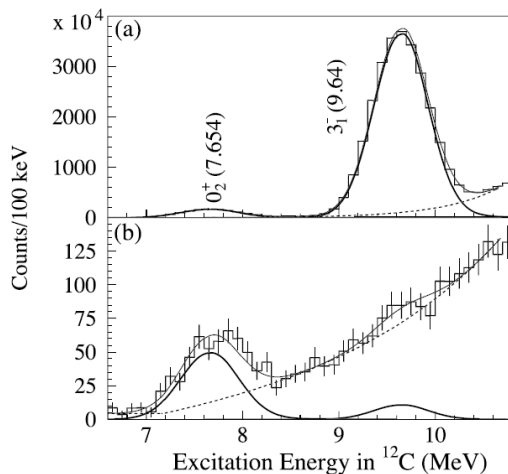
Hoyle status new surprising result

$$\Gamma_{\text{rad}}/\Gamma = 6.2(6) \times 10^{-4}.$$

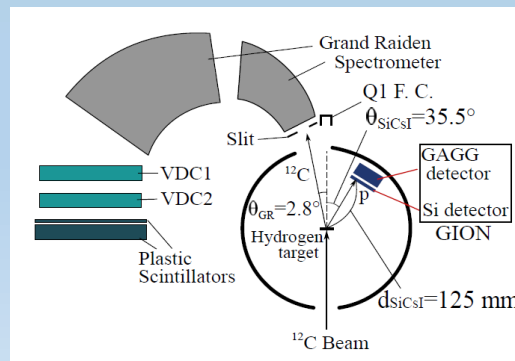
Tsumura et al Phys.Lett. B817(2021)136283

9.64 Yield larger than expectation also measured

$$\Gamma_{\text{rad}}/\Gamma_{\text{tot}} = 1.3_{-1.1}^{+1.2} \times 10^{-6}$$

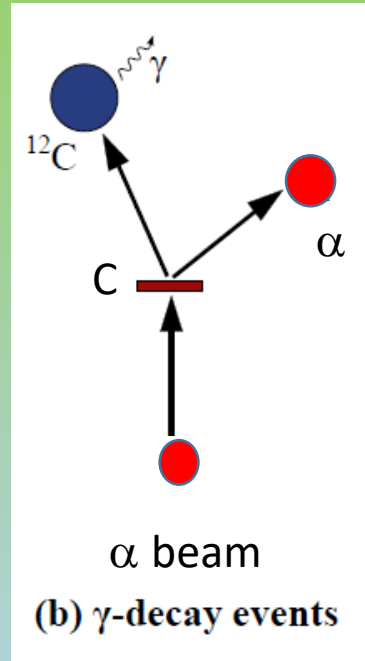
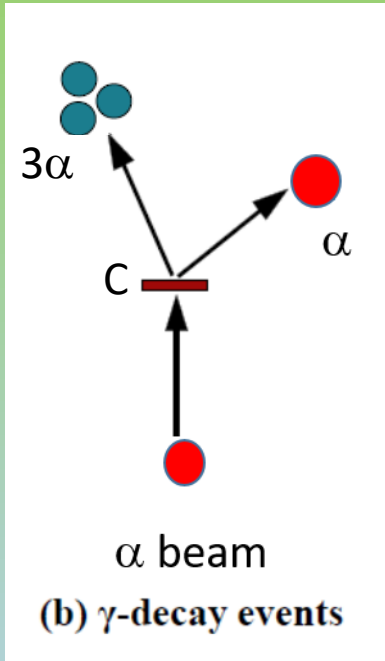


Very high background
 notwithstanding the big technological
 efforts



Solid hydrogen target

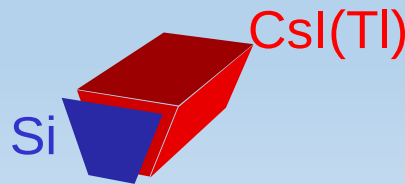
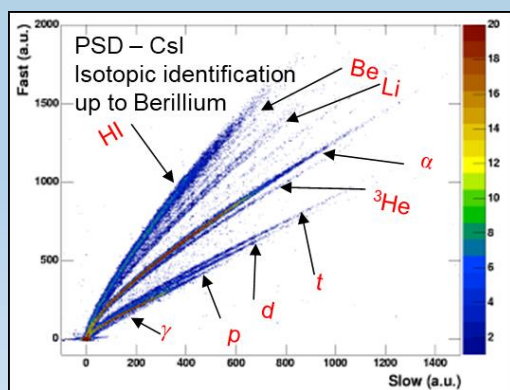
BACKGROUND REDUCTION : THE CRM method



We can reduce the background using more constraints: measure all particles and γ – use energy and momentum conservation

Complete Redundant Measurement - CRM

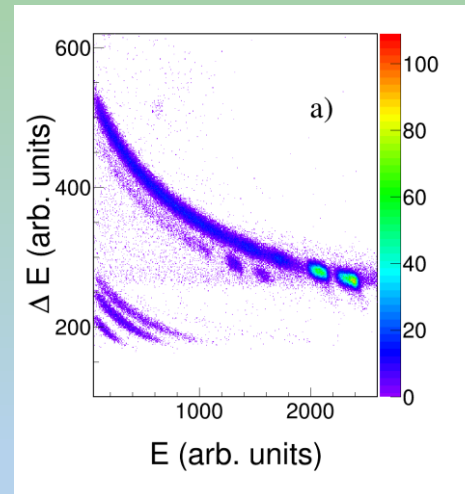
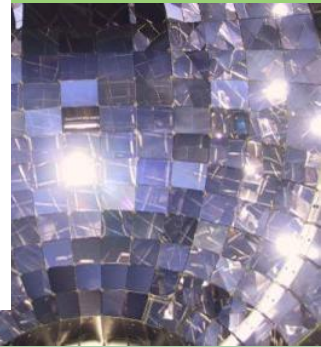
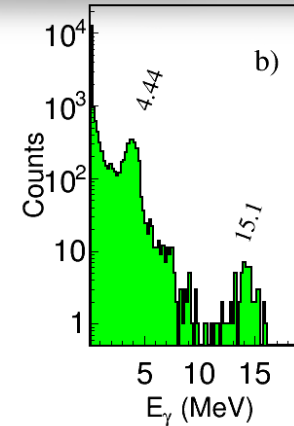
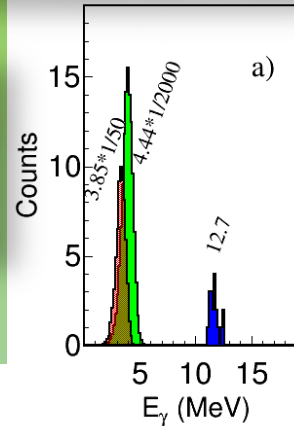
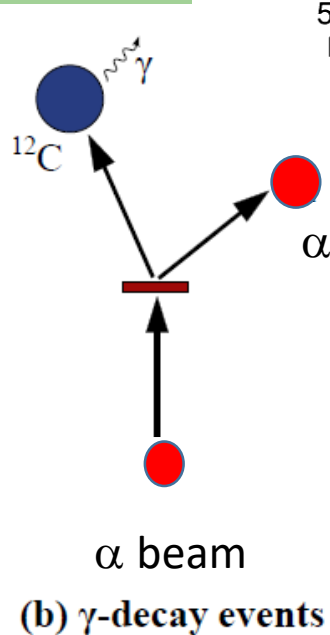
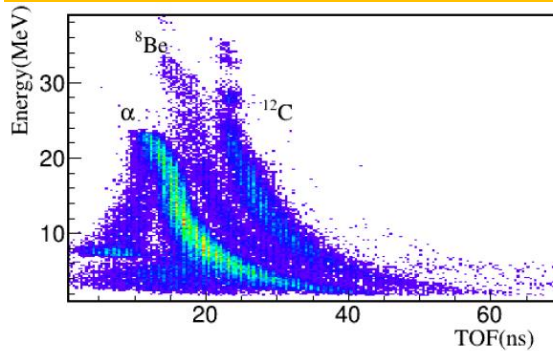
**Need a complete detector :
CHIMERA**



THE EXPERIMENT : detection method

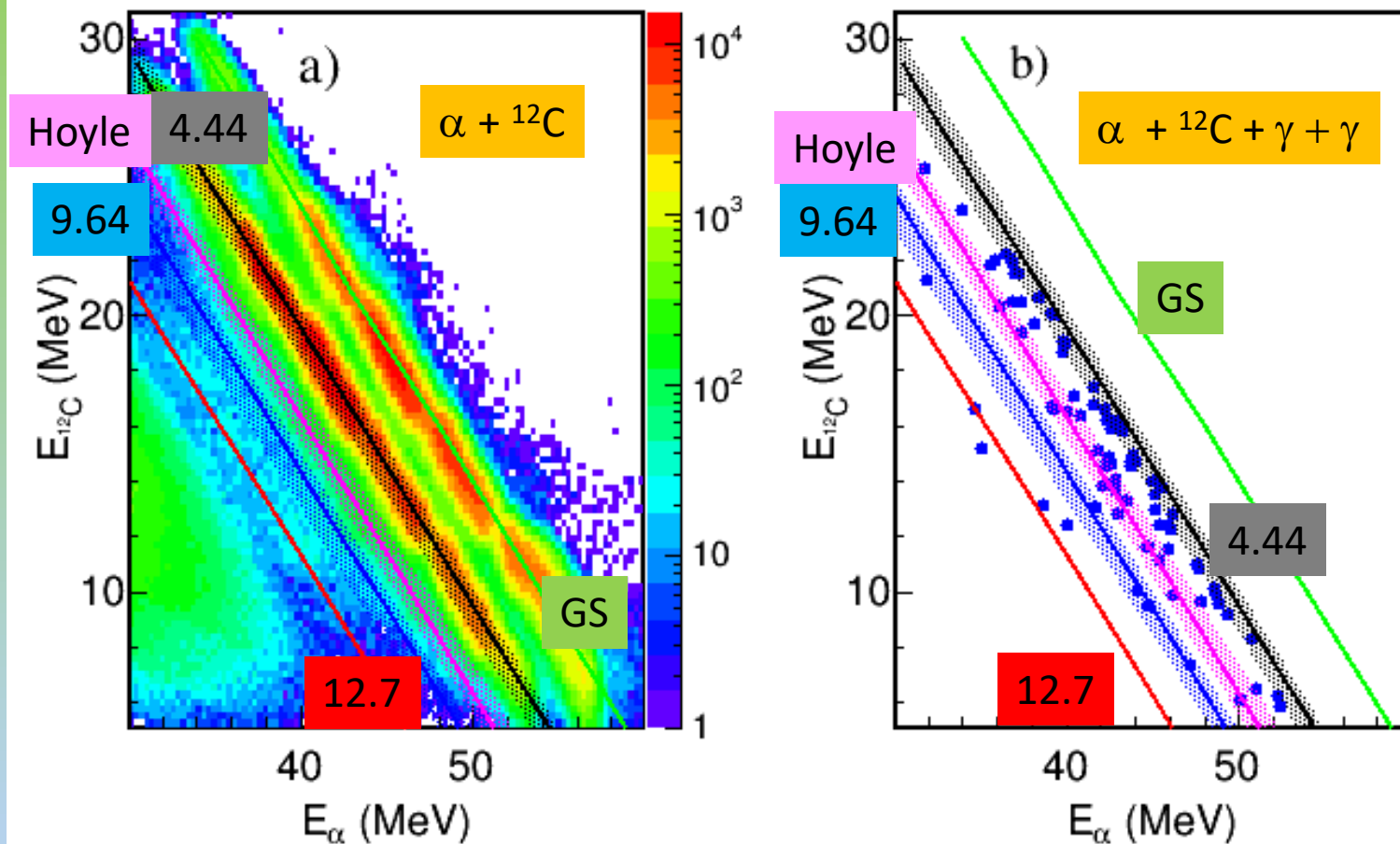
We performed the reaction
 $\alpha + {}^{12}\text{C}$ at 64 MeV at LNS

G.Cardella et al PRC in print



We detect and identify in energy both scattered α -particles(by ΔE -E) and carbon (by time of flight) – detection efficiency of kinematical coincidences is 100% in a 4π detector – γ rays in CsI(Tl)

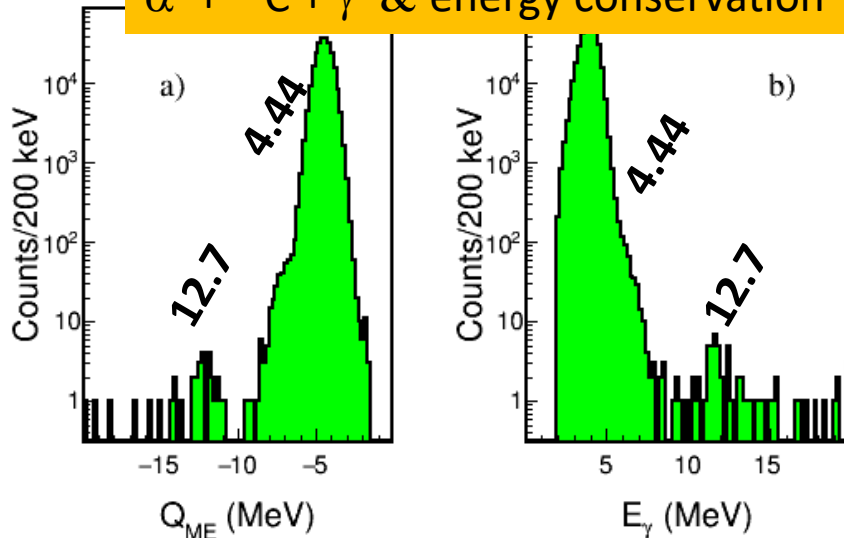
THE EXPERIMENT : CRM cleaning effect



This plot shows how one can further clean Chamberlin method results imposing also the coincidence with 2- γ and all conservation rules even with a relatively scarce angular and energy resolution

THE EXPERIMENT : particle-gamma coincidences

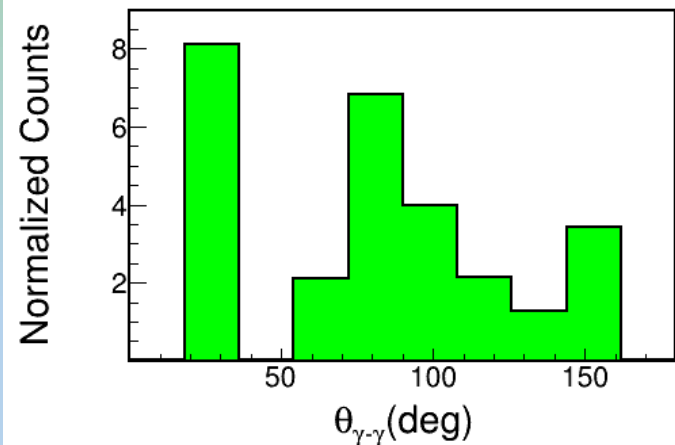
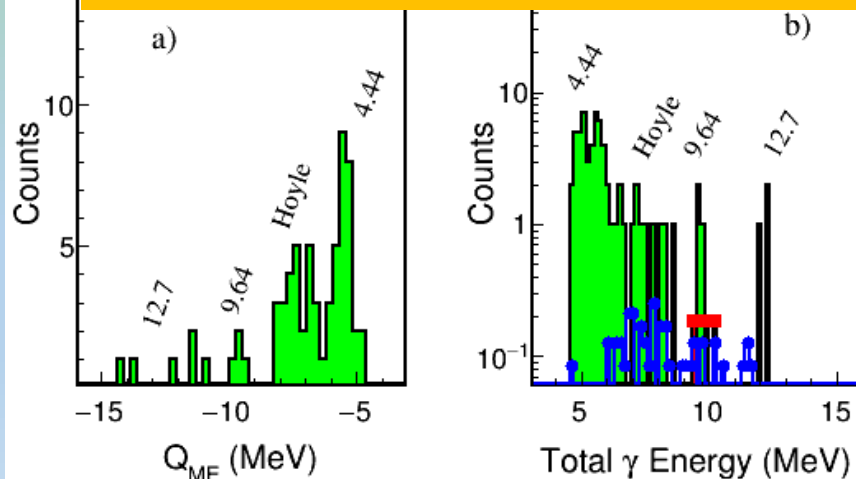
$\alpha + {}^{12}\text{C} + \gamma$ & energy conservation



In coincidence with 1 γ we see only the 4.44 and the 12.7 MeV level both in Q-val and in γ -energy

In coincidence with 2 γ we see a little residual background of the 4.44 the Hoyle, the 9.64 MeV and 12.7 MeV level both in Q-val and in γ -energy

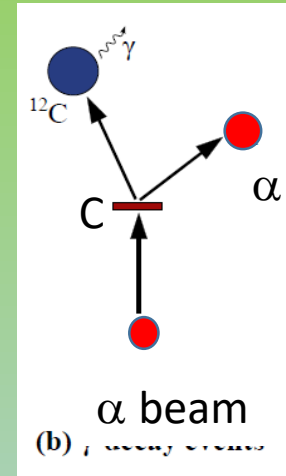
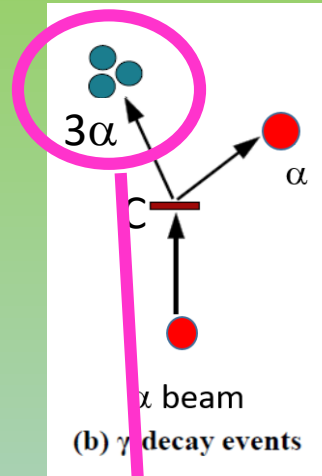
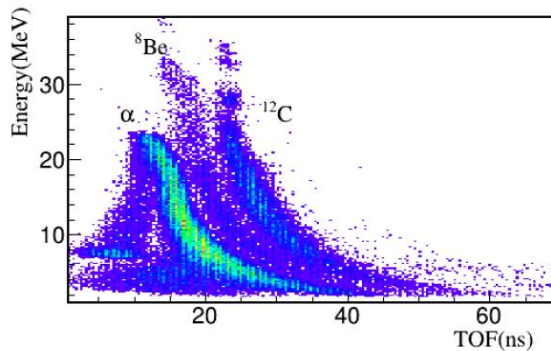
$\alpha + {}^{12}\text{C} + \gamma + \gamma$ & energy conservation



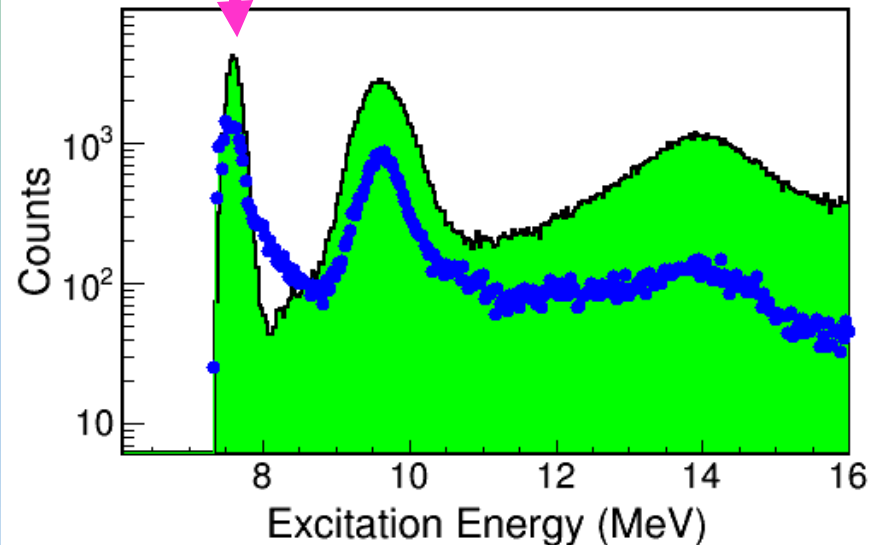
γ rays angular correlation for the Hoyle state similar to M.Munch et al PRC 93, 065803(2016)

THE EXPERIMENT : YIELD EVALUATION

How to extract the yield of Hoyle and 9.64 Mev Level?



We detect and identify α particles from the decay of ^{12}C so we can evaluate the excitation energy in the 3- α CM and produce the excitation energy spectrum giving the yields



THE EXPERIMENT : evaluated yields

level	singles (counts/ 1000)	two γ counts	efficiency 2 γ %	yield 2 γ
12.7	7 $\pm$ 1.4	1 $\pm$ 1	5.1	0.0028 $\pm$ 0.0028
9.64	778.2	2 $\pm$ 1.6	4.0	6.4 $\pm$ 5.1 $\times 10^{-5}$
7.65	199.8	12 $\pm$ 3.8	3.3	1.8 $\pm$ 0.6 $\times 10^{-3}$

9.64 Tsumura

$$\Gamma_{\text{rad}}/\Gamma_{\text{tot}} = 1.3_{-1.1}^{+1.2} \times 10^{-6}$$

Very large error bars
the two results are
compatible

Hoyle - kibedi

$$\Gamma_{\text{rad}}/\Gamma = 6.2(6) \times 10^{-4}.$$

**Hoyle decay
too large
WHY?**

THE EXTRA-YIELD : efimov state?

Look to some old paper

Volume 33B, number 8

PHYSICS LETTERS

21 December 1970

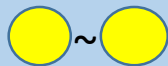
ENERGY LEVELS ARISING FROM RESONANT TWO-BODY FORCES IN A THREE-BODY SYSTEM

V. EFIMOV

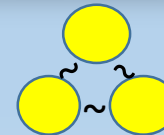
A.F.Ioffe Physico-Technical Institute, Leningrad, USSR

Received 20 October 1970

Again in the 70s of last century Efimov suggested that when two bodies have a resonance, in the system formed by 3 bodies, one will observe a bound status and in principle also an infinite number of excited levels over it. Its goal was to demonstrate that the Hoyle level could be explained by this quantum effect but he failed

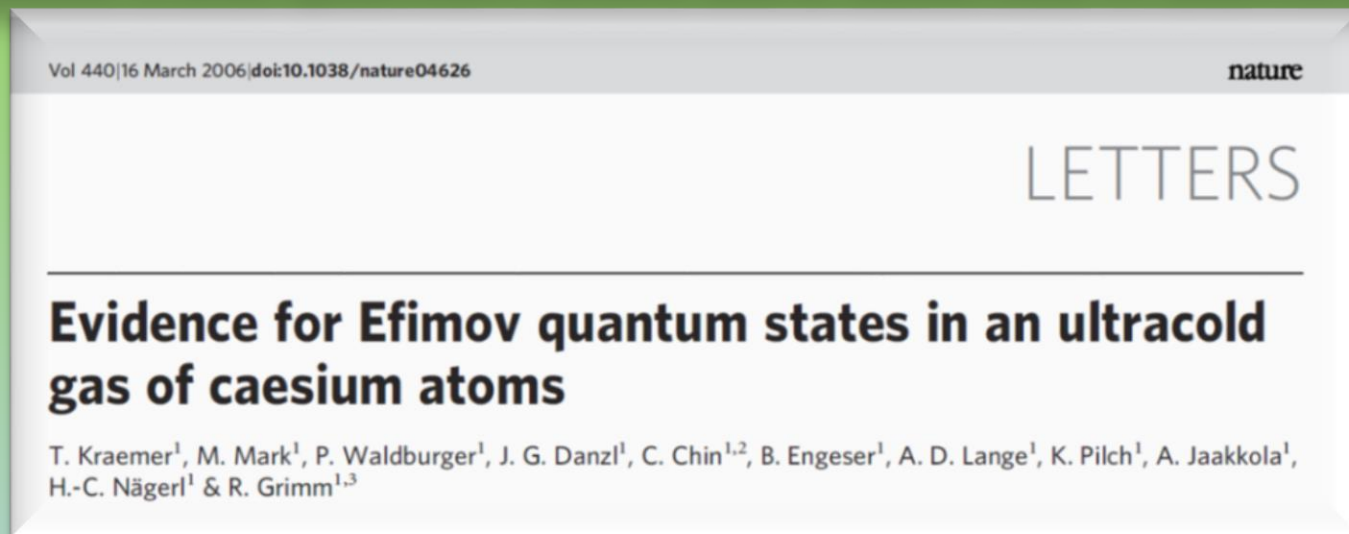


2 α resonance
92 keV (^8Be)



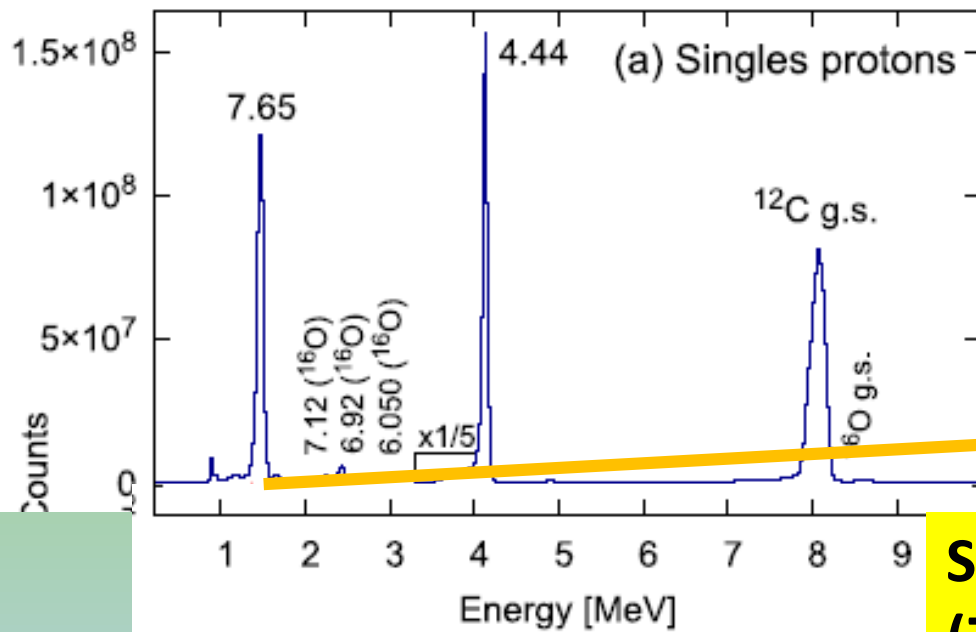
3 α bound
level

On 2006 these quantum levels were observed in atomic physics



We discuss a hypothetical 'Efimov state' at $E^*(^{12}\text{C}) = 7.458 \text{ MeV}$,

THE EXTRA-YIELD : Efimov state?



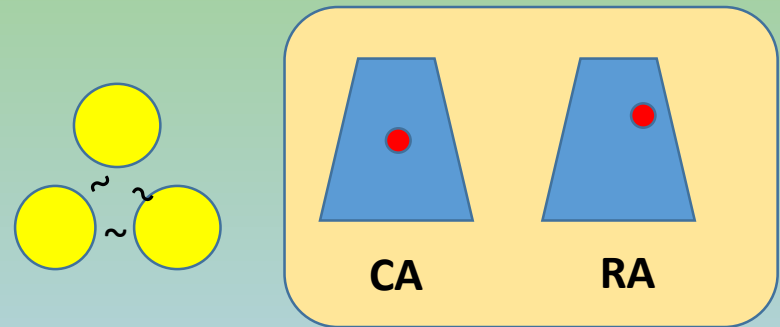
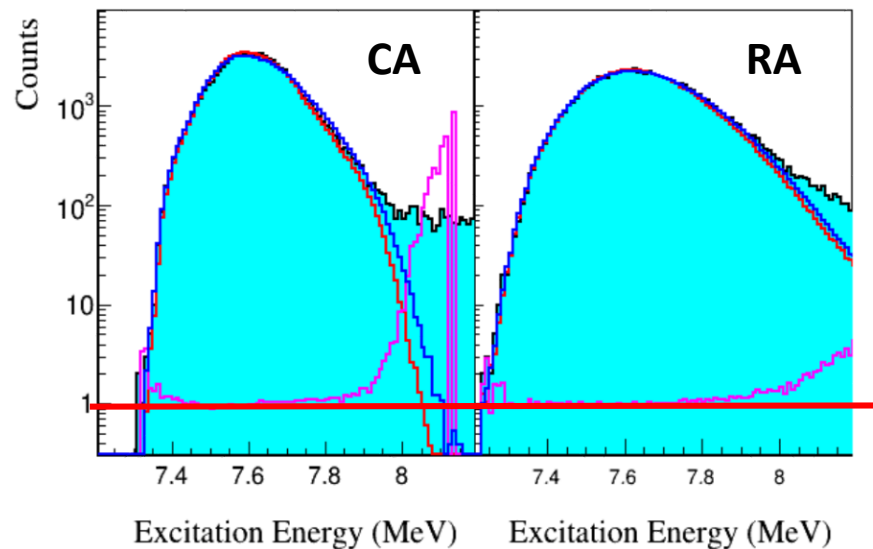
Suppose the Efimov level (7.458 MeV) is populated near the Hoyle state (7.64 MeV)

Efimov state should have near 100% γ decay width against the 4×10^{-4} of the Hoyle state due to large barrier for α decay

It is enough a very low population probability to explain our yield enhancement

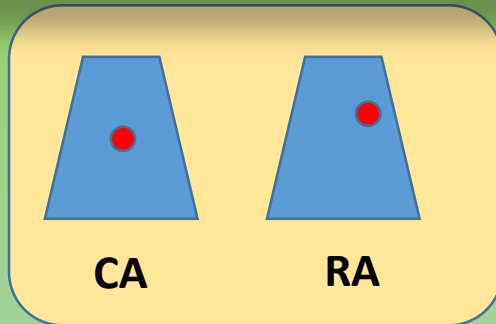
THE EXTRA-YIELD : looking for Efimov state in 3- α data?

In order to search for events different from the Hoyle state decay we simulate the response of our detection system

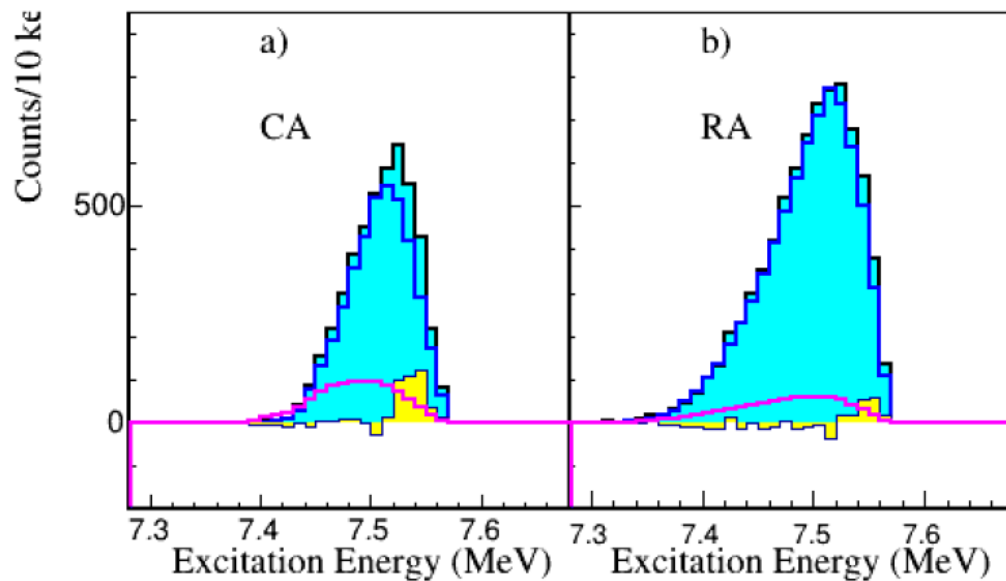
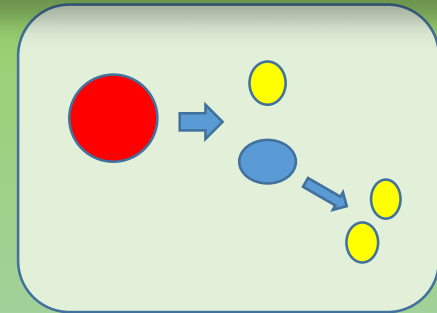


To get the 3- α CM energy we must assign an angle to each event the center of the detector CA or a random angle RA

THE EXTRA-YIELD : looking for Efimov state in 3- α data – sequential decay



SD –Sequential Decay



Experimental data selected with cuts (cyan filled)

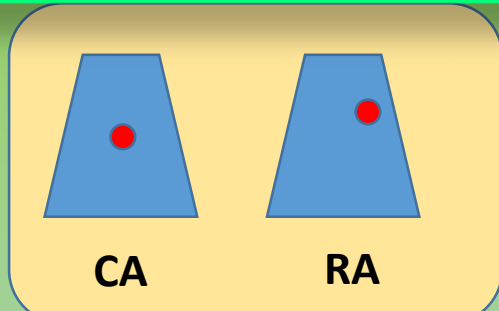
simulations of hoyle decay wrongly measured.

experimental data minus wrong simulated events and random coincidences

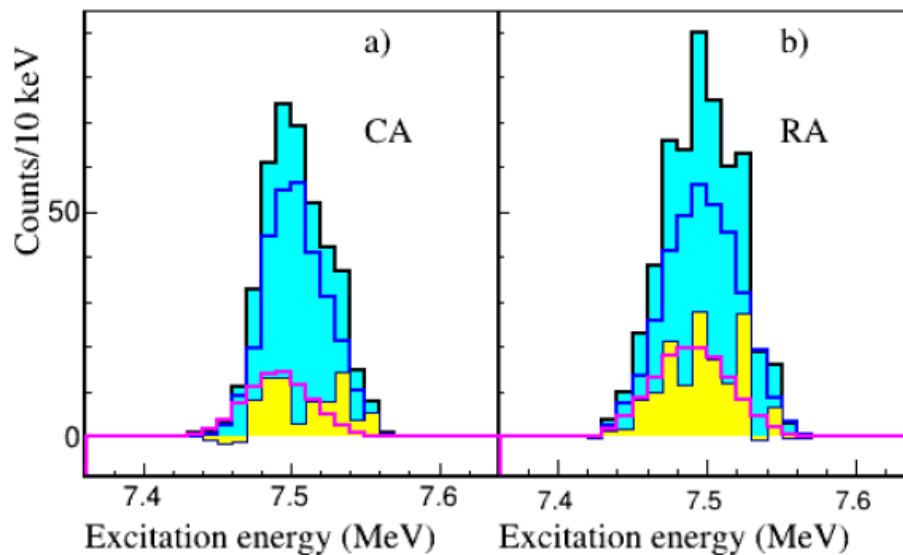
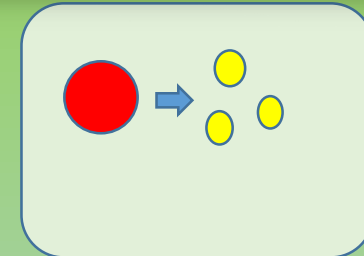
Violet is expected EFIMOV spectrum

No evidence for EFIMOV sequential decay

THE EXTRA-YIELD : looking for efimov state in 3- α data – direct decay



DD - Direct Decay



Experimental data selected with cuts (cyan filled)

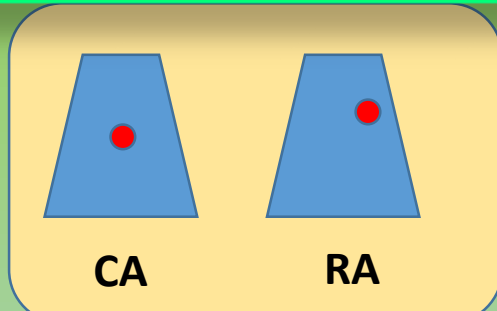
simulations of hoyle decay wrongly measured.

experimental data minus wrong simulated events and random coincidences

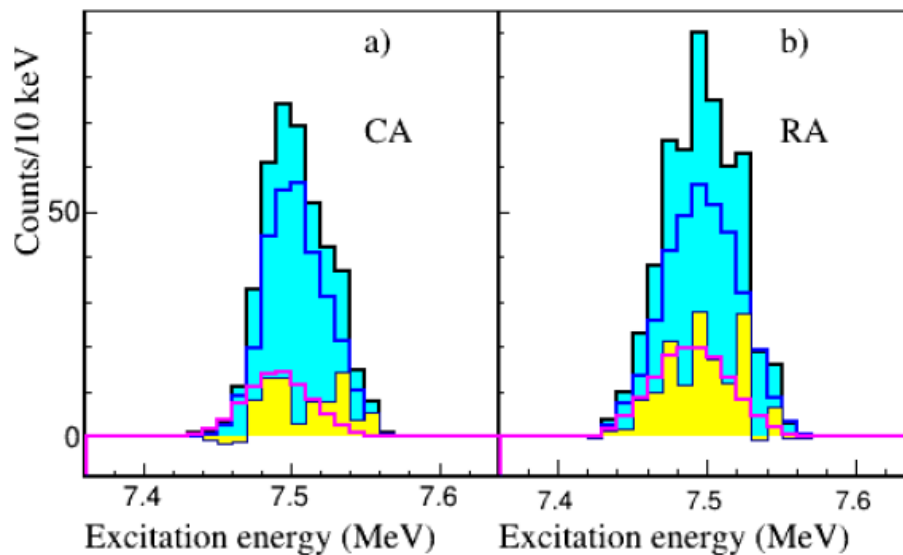
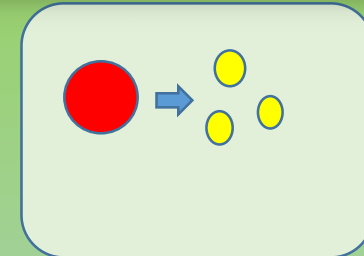
Violet is expected EFIMOV spectrum

The shape of the spectra is similar to expectation
The yield is too large in the case of RA
While for the detection efficiency we expect the contrary

THE EXTRA-YIELD : looking for efimov state in 3- α data – direct decay



DD - Direct Decay



The shape of the spectra is similar to expectation
The yield is too large in the case of RA
While for the detection efficiency we expect the contrary

108 events RA efficiency 8%
means 1350 events against
85000 Hoyle events

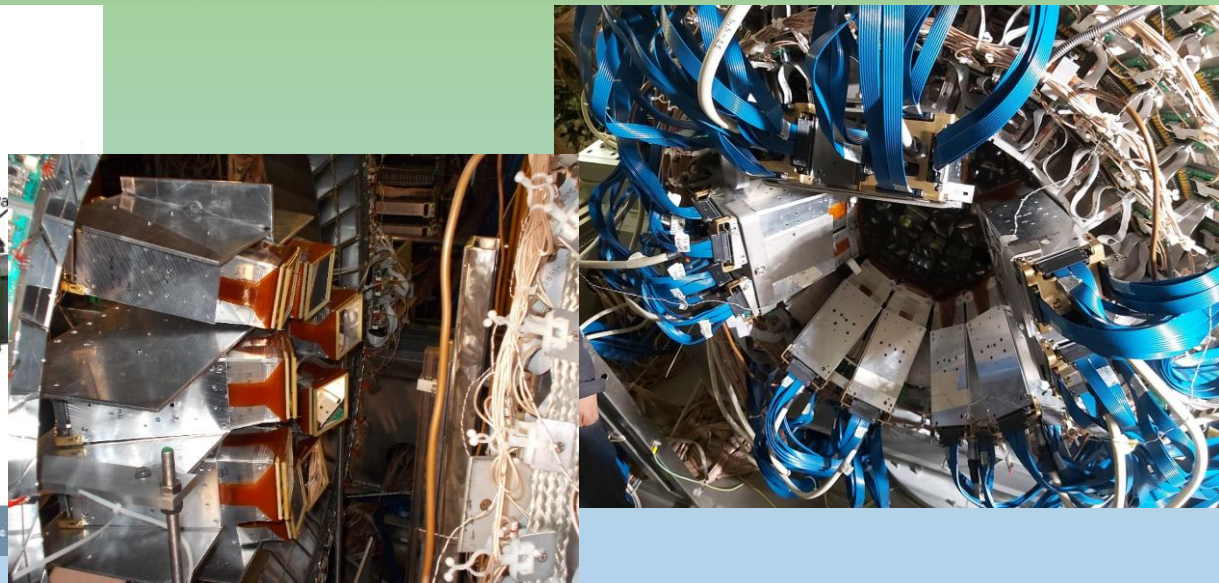
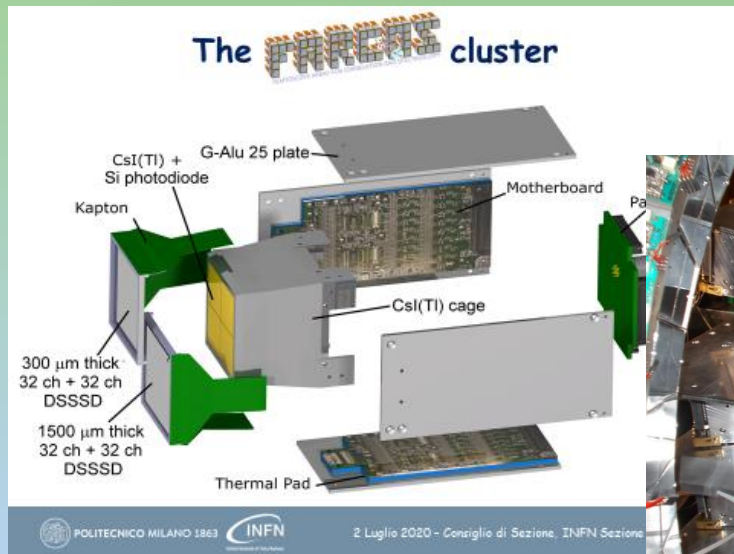
47 events CA efficiency 24% ->
195 events over 85000

For the gamma experiment we
required also the coincidence
with high energy alpha
scattered having only 28500
events this means around 65
events compatible with Efimov
what expected if we have 80% γ
decay probability of the Efimov
state.

THE EXTRA-YIELD : new measurement?

I have not conclusive convincing data we must perform new measurements able to evidence the presence of 100 Efimov events near to 100000 Hoyle events we need much more precise measurements

We will improve the experiment using the new FARCOS telescopes to measure better q-values



Hope to do the new experiment beginning 2023

γ decay of
Hoyle and
9.64 MeV
levels

**FIRST DIRECT OBSERVATION OF γ -DECAY of 9.64
MeV LEVEL**

**TOO LARGE γ -DECAY WIDTH OF THE HOYLE
STATE**

**EFIMOV
STATE?**

**NO EVIDENCE OF SEQUENTIAL DECAY OF EFIMOV
STATE**

**POSSIBLE PRESENCE OF DIRECT DECAY OF SUCH A
STATE**

**HOYLE EXTRA-YIELD COMPATIBLE WITH
CONTAMINATION FROM EFIMOV DECAY ASSUMING
 $\Gamma_\gamma/\Gamma \sim 80\%$**

**Future
opportunities**

**NEW MEASUREMENTS WITH BETTER
RESOLUTION ON Q-VALUE
MEASUREMENT USING FARCOS**

The collaboration

G. Cardella(1), A. Bonasera(2)(3), N.S. Martorana(2)(4), L. Acosta(1)(5), E. De Filippo(1), E. Geraci(1)(4), B. Gnoffo(1)(4), C. Guazzoni(6), L.Lo Monaco(4), C. Maiolino(2), A. Pagano(1), E.V. Pagano(2), M. Papa(1), S. Pirrone(1), G. Politi(1)(4), F. Rizzo(2)(4), P. Russotto(2), D. V. Sicari(6), and M. Trimarchi(1)(7)

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