

Comparison Data – MC

Data reconstruction

Acceptance comparison with MC simulation

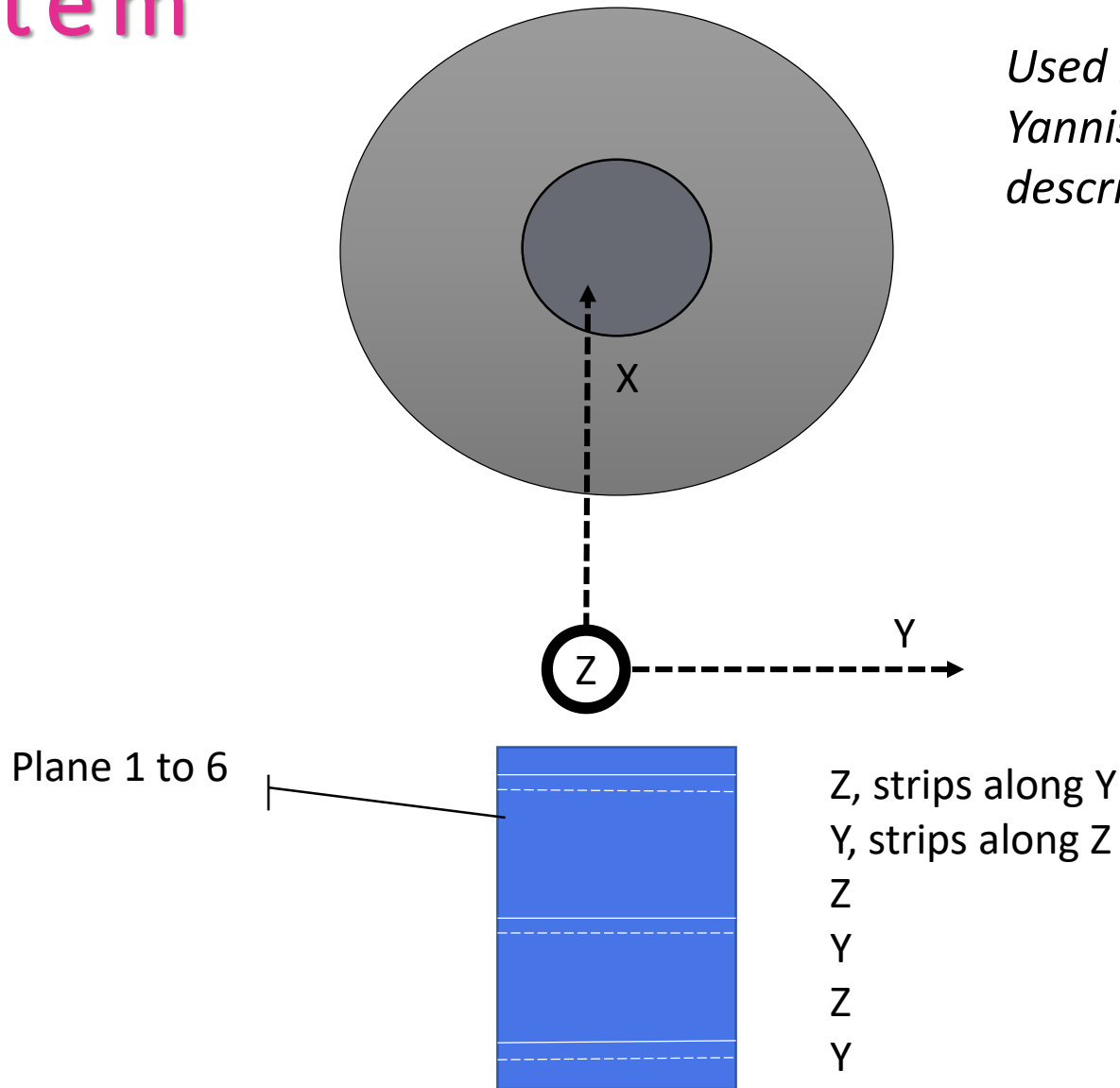
Corinne, Yannis, Theodoros, Sabine

Data

- Preselection by Theodoros → Rootfile
- Data Looking towards the tumulus
- *Evts with at least $\sum ADC > 200$, per plane*
 - No cut on the middle 2 planes

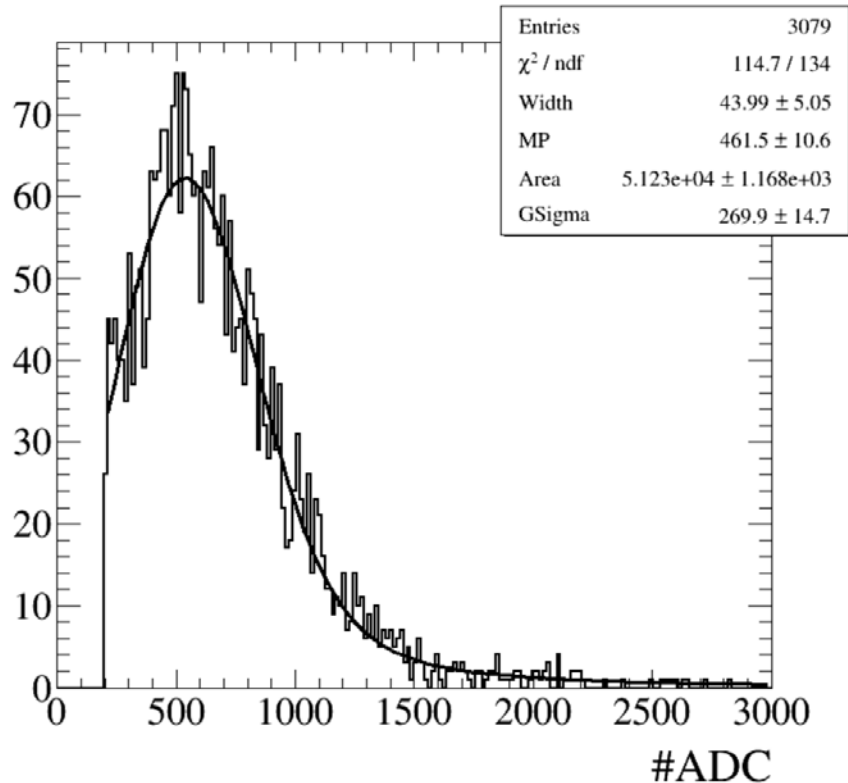
Numbering & Coordinate System

*Used in this presentation
Yannis uses a different
description*

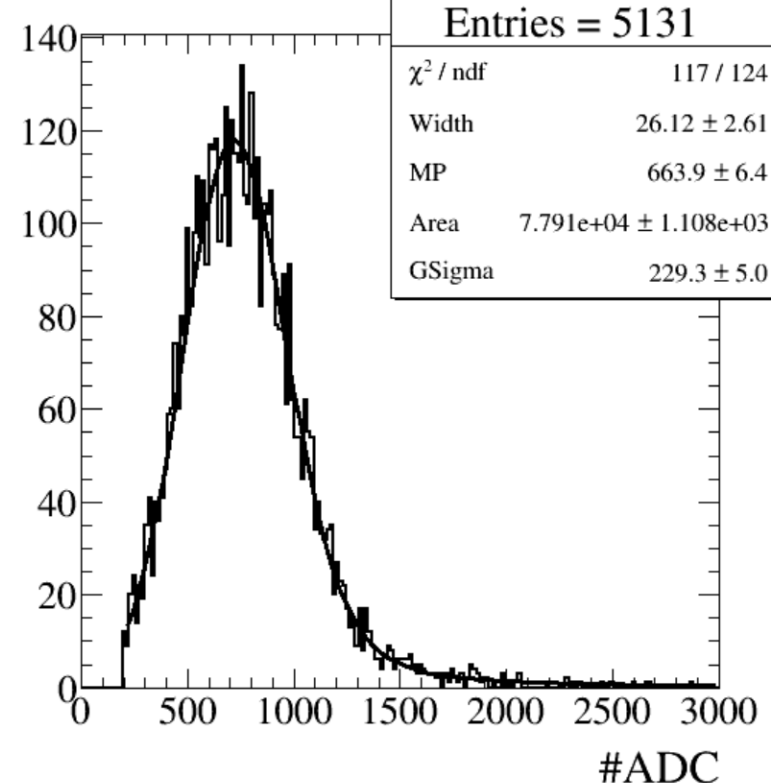


Simulation: E (kev) to ADC

Data (Example)



MC (Example)

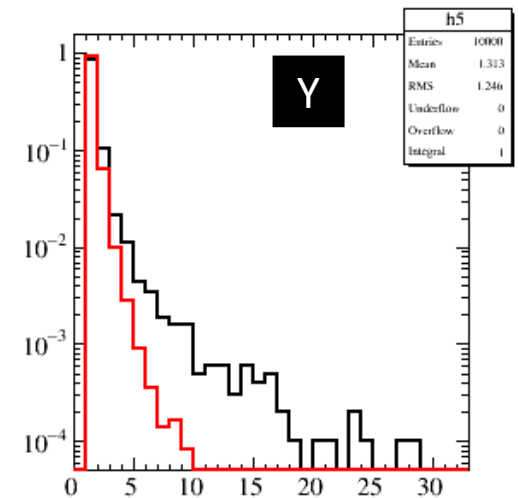
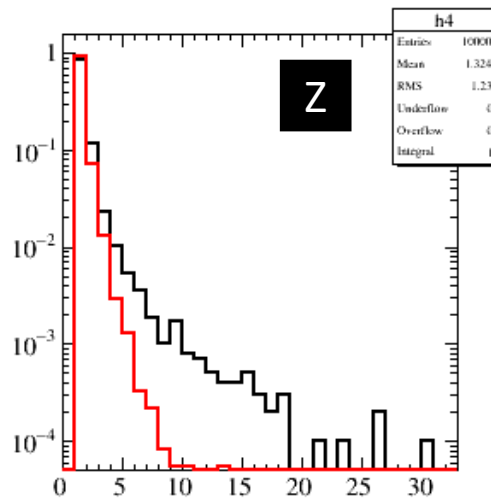
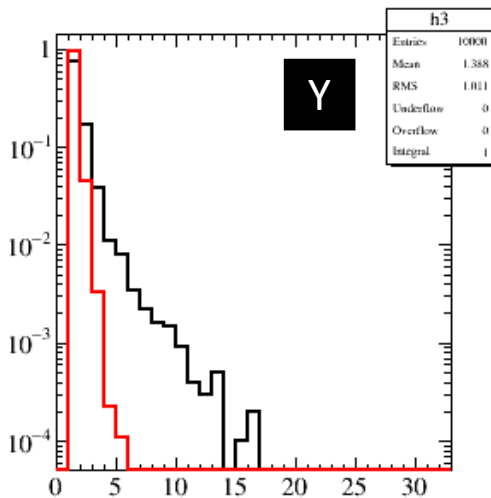
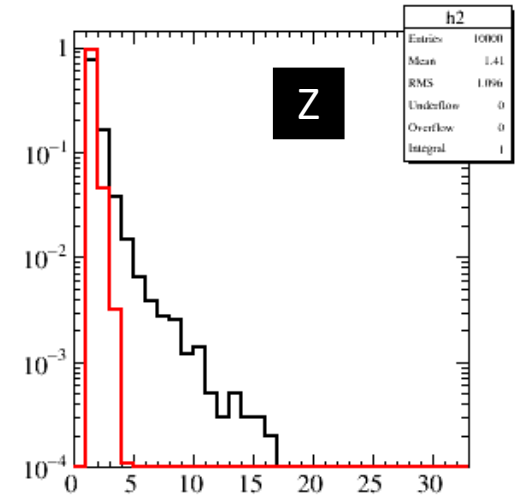
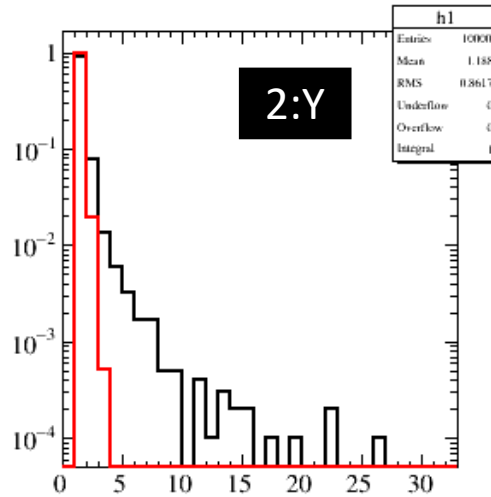
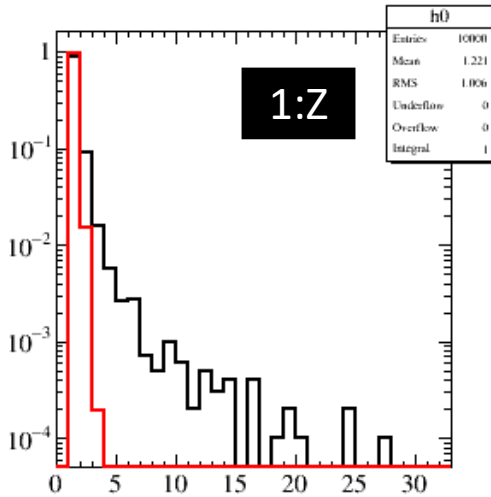


Simulation to be adjusted channel per channel, but qualitatively Ok for now

Raw Multiplicity

Data
Simulation

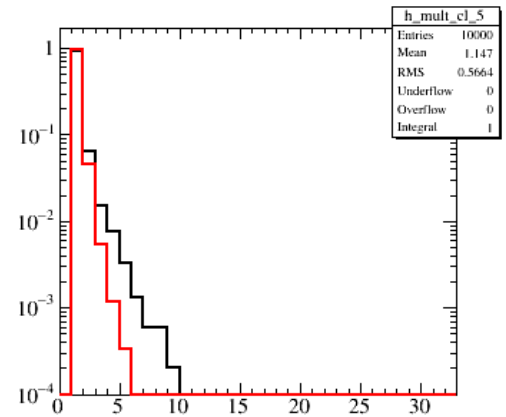
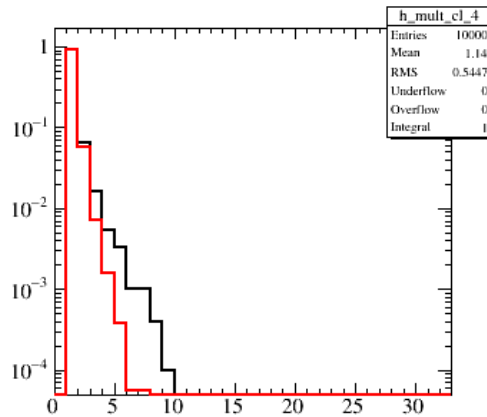
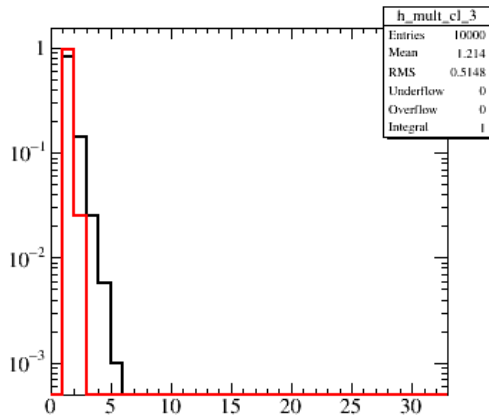
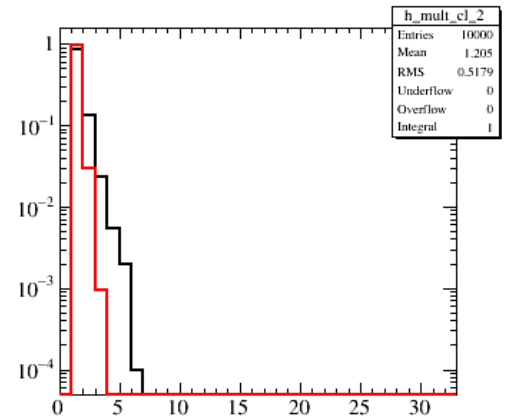
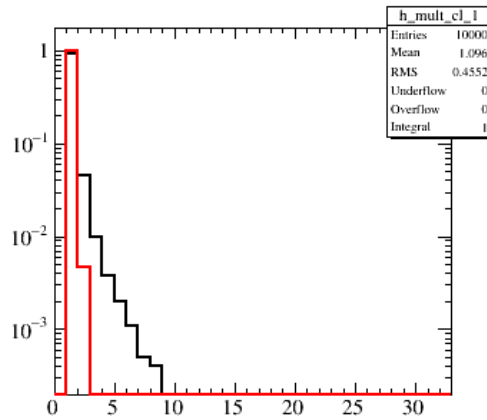
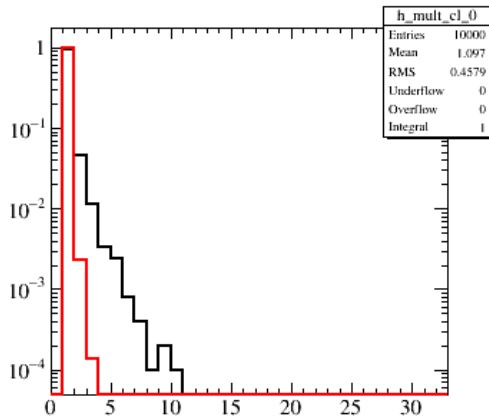
Events



Number of strips

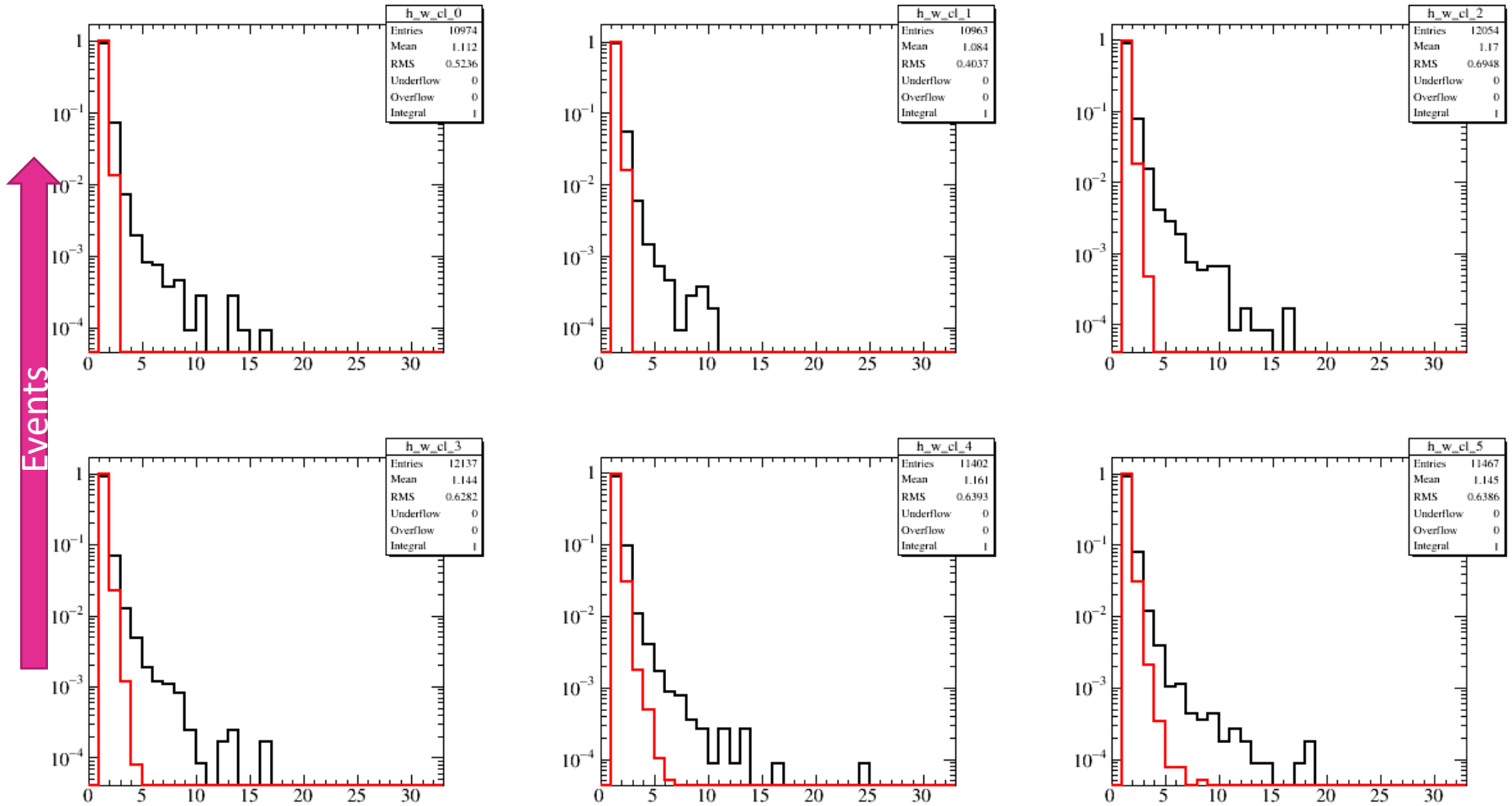
Clustering

- Simple clustering associating neighbors



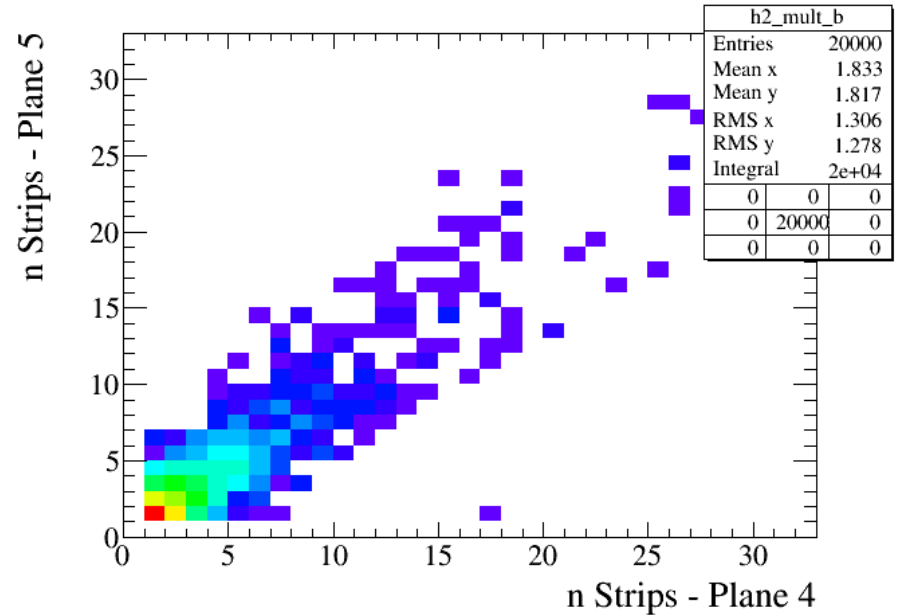
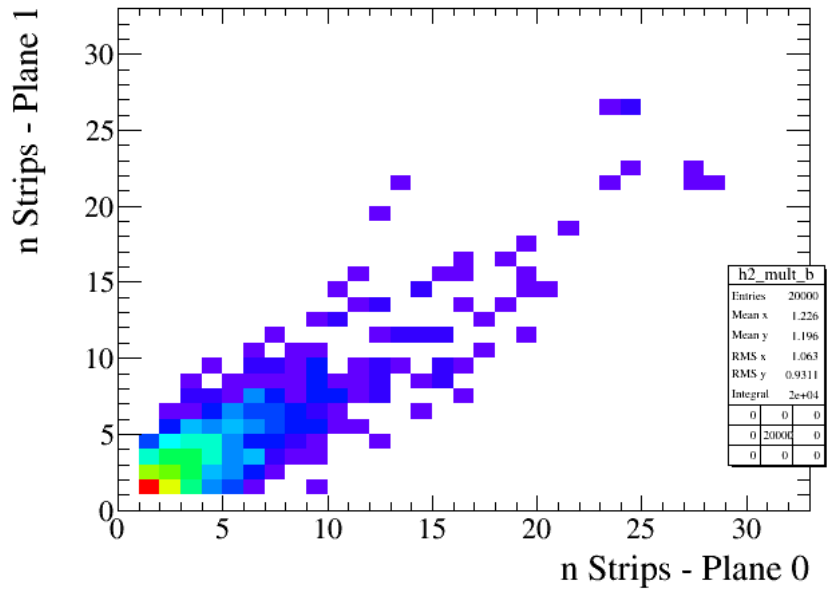
Number of clusters

Cluster's width



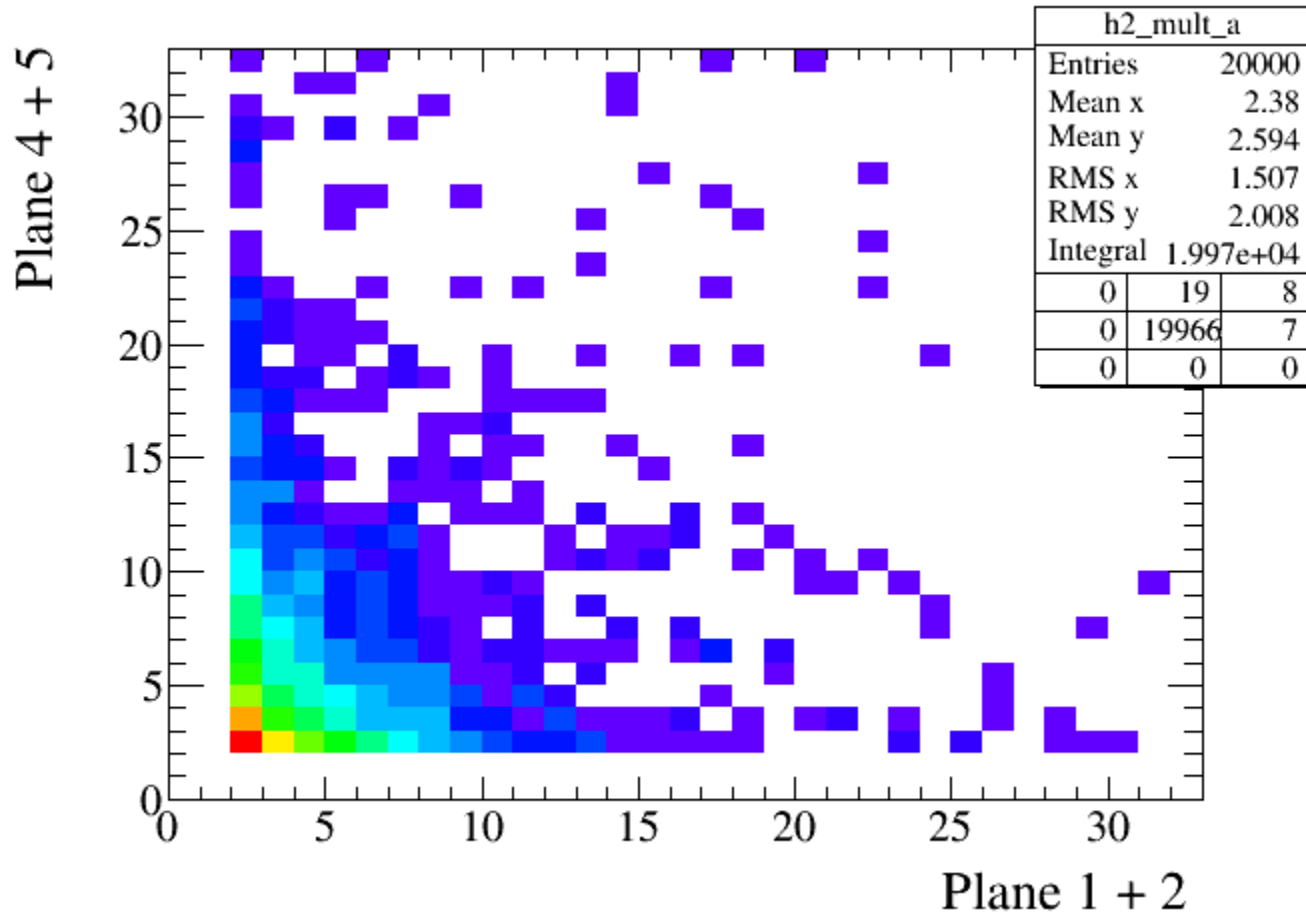
Investigating whether the
larger multiplicity could
be due to electrons

Electrons ?



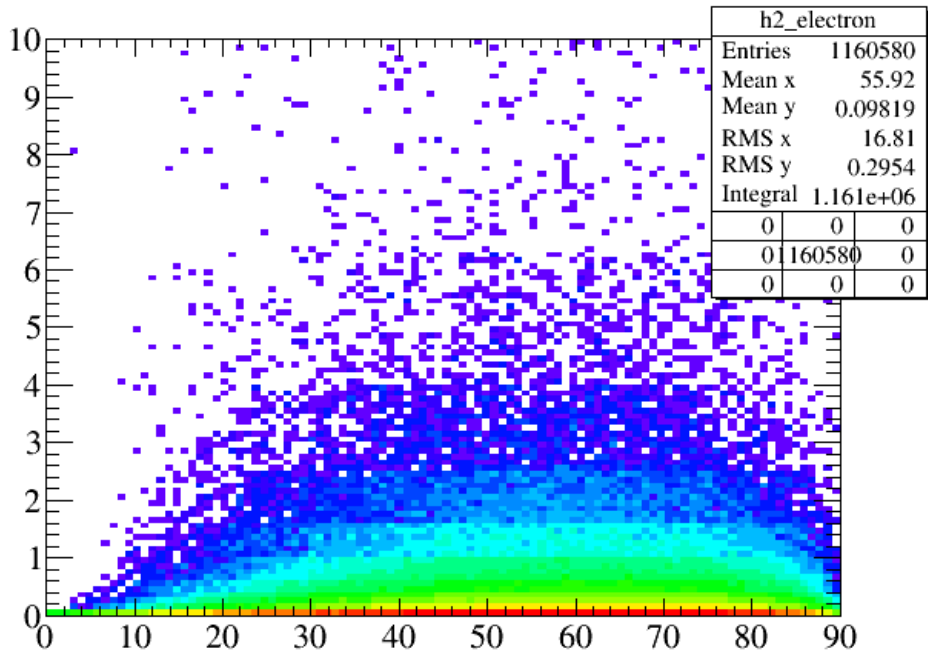
Seems to be of physical origin

Electrons ? ~No correlation Front-Back



Electrons/ Gamma (CRY)

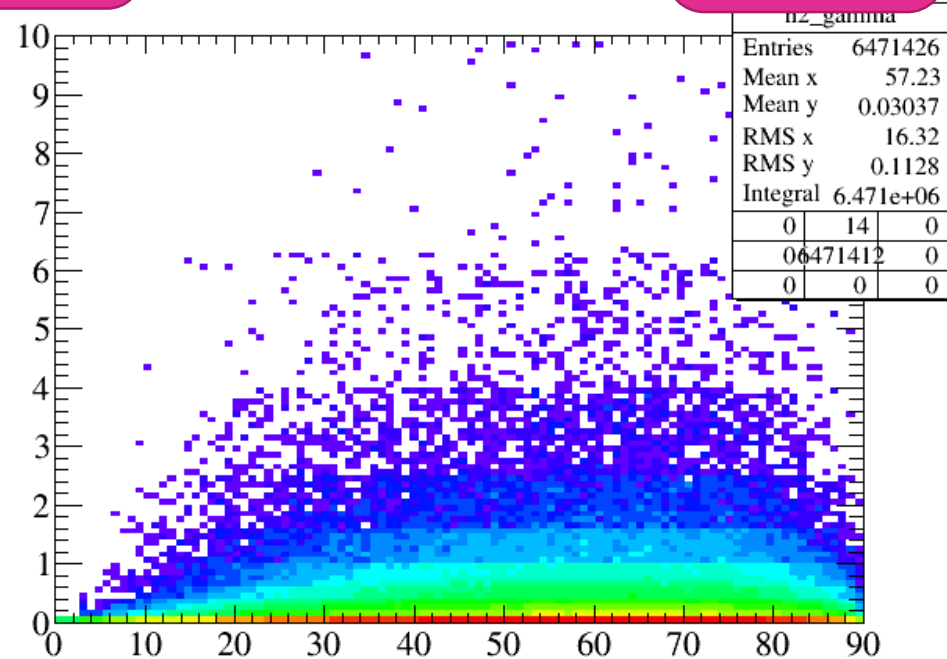
Energy (GeV)



Horizontal

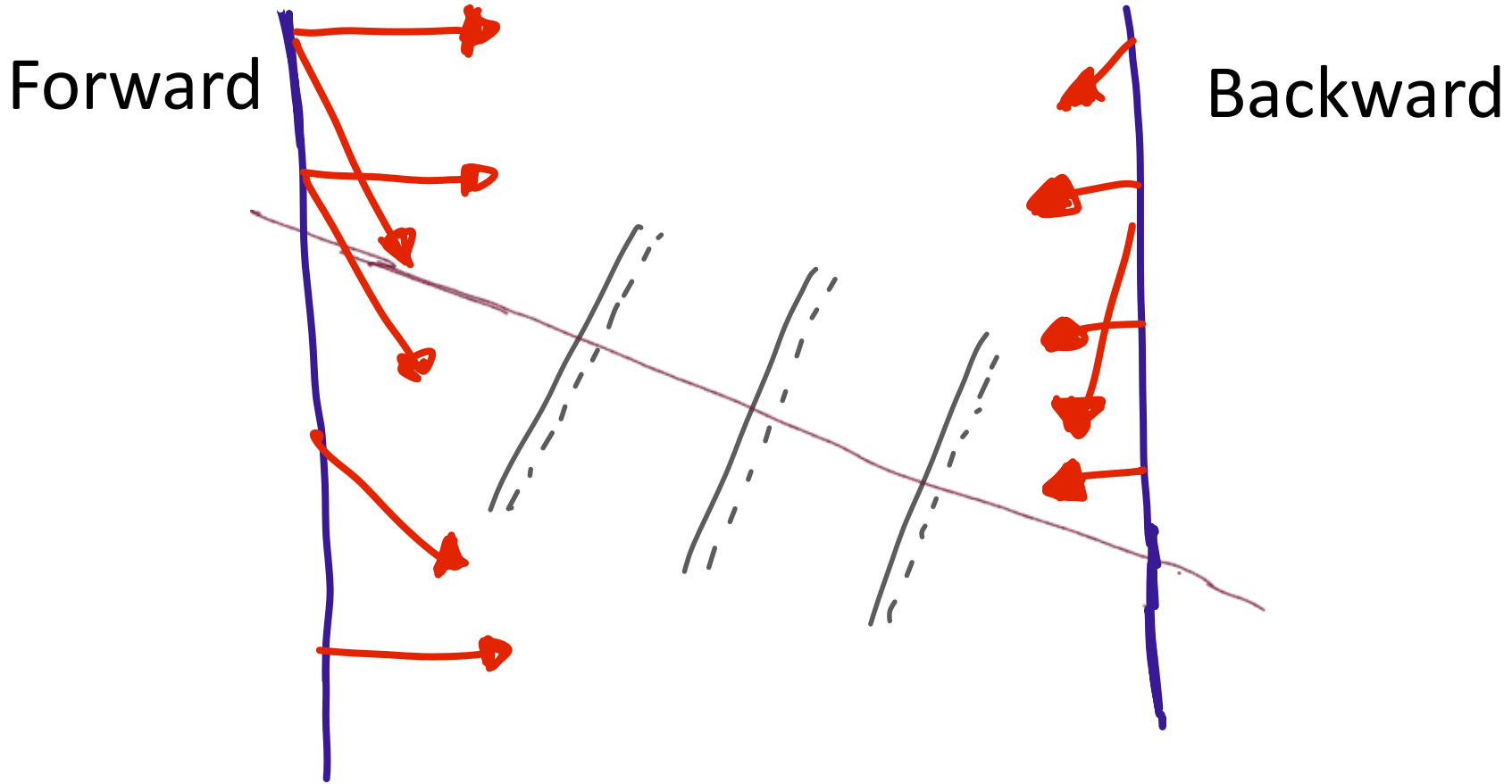
Vertical

Energy (GeV)

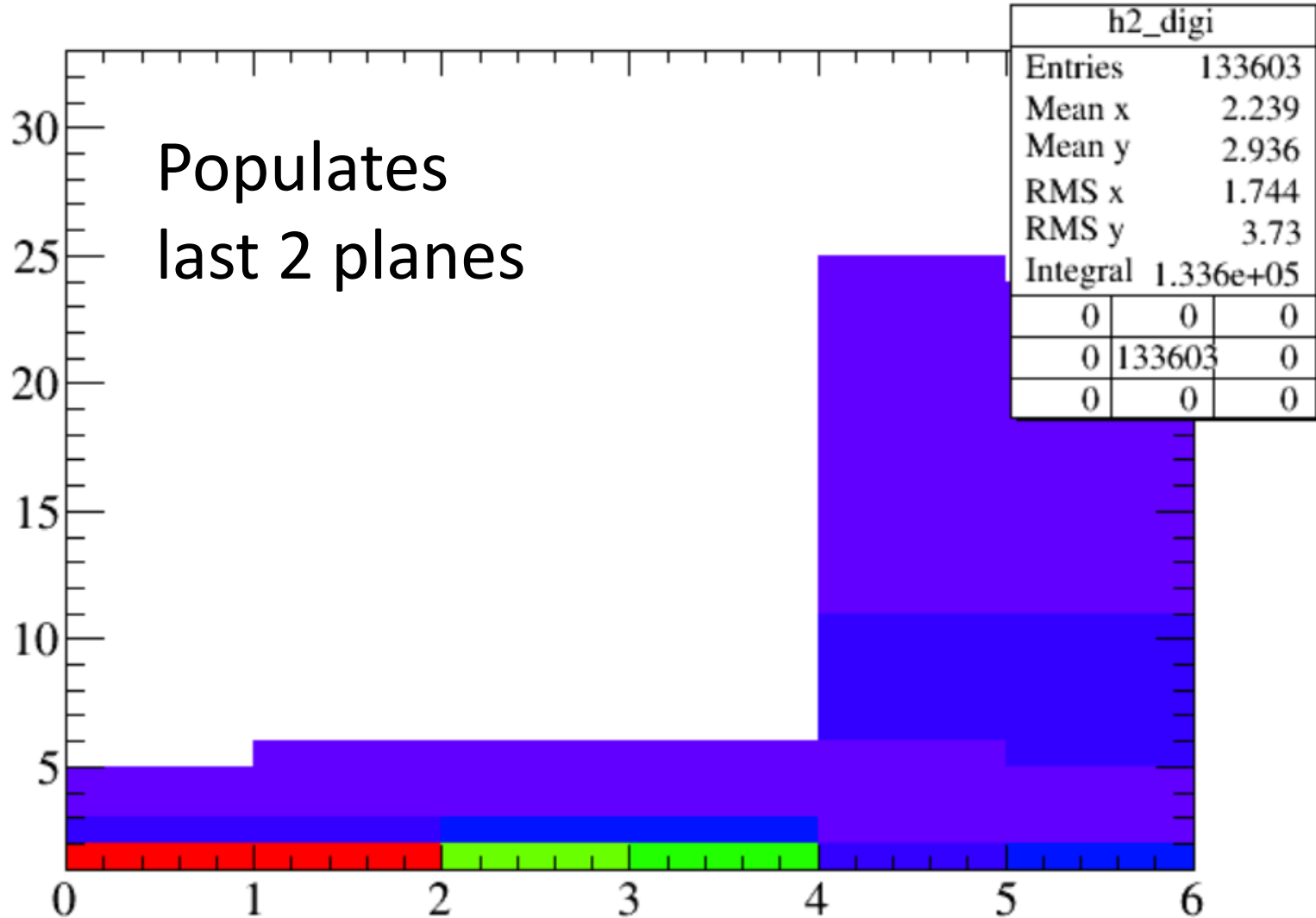


Fast simulation

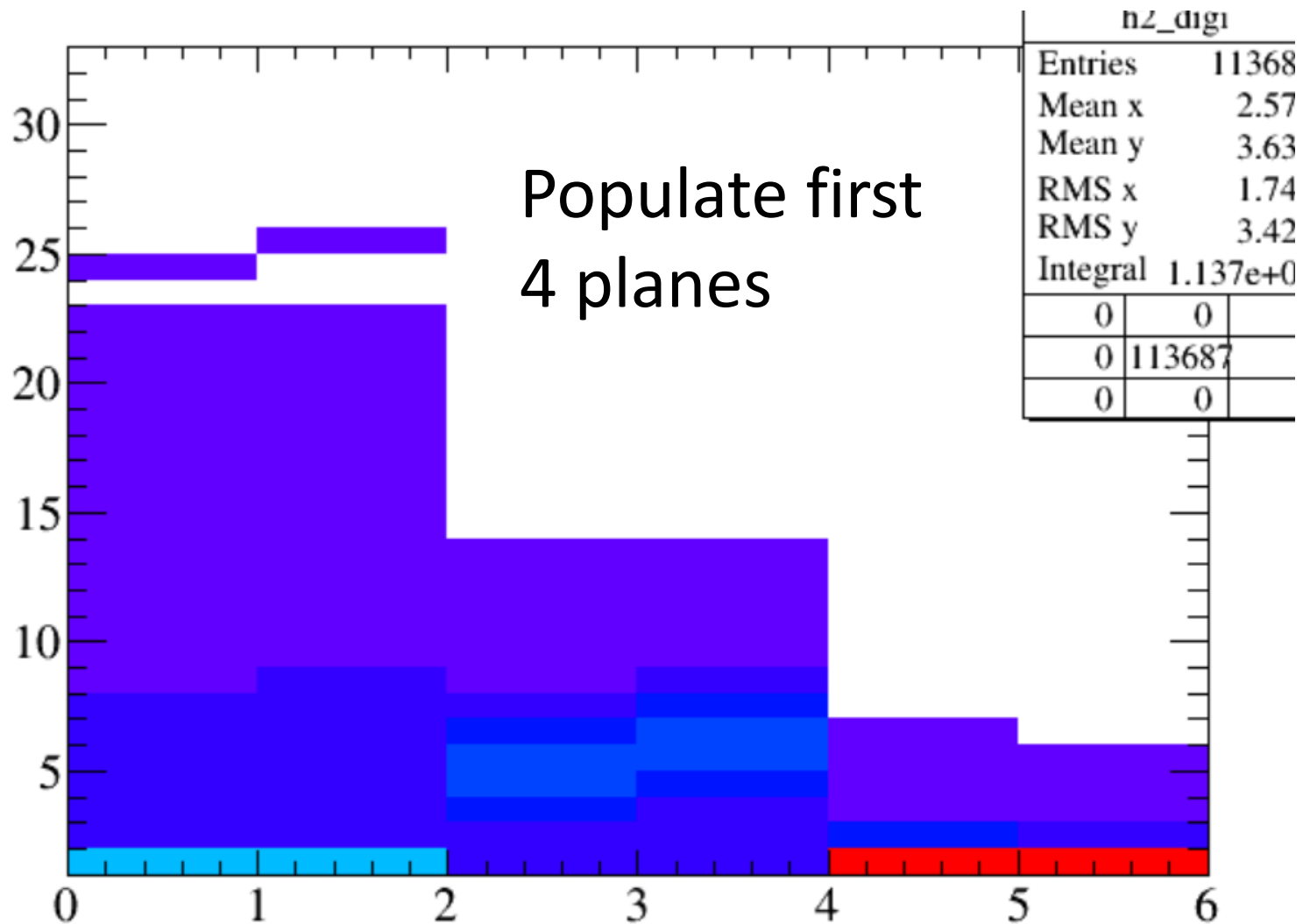
$$\left\{ \begin{array}{ll} \cos^2 \theta & \mu^- \quad 10 \text{ GeV} \\ \text{isotropic} & e^- \quad 5 \text{ GeV} \end{array} \right.$$



5 GeV forward electrons



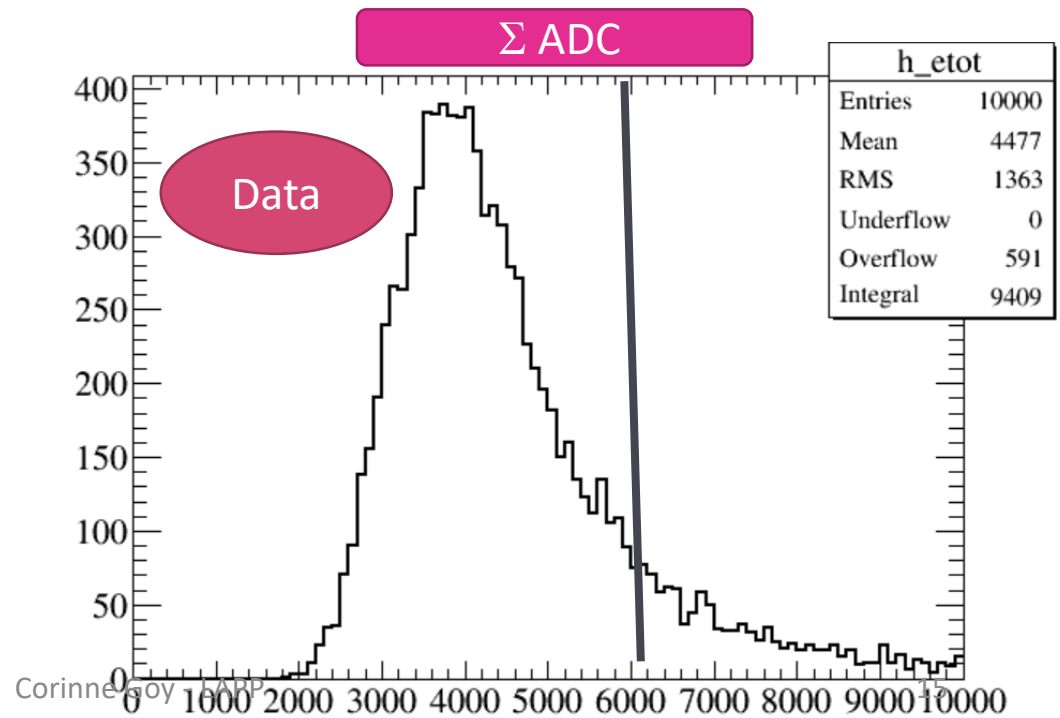
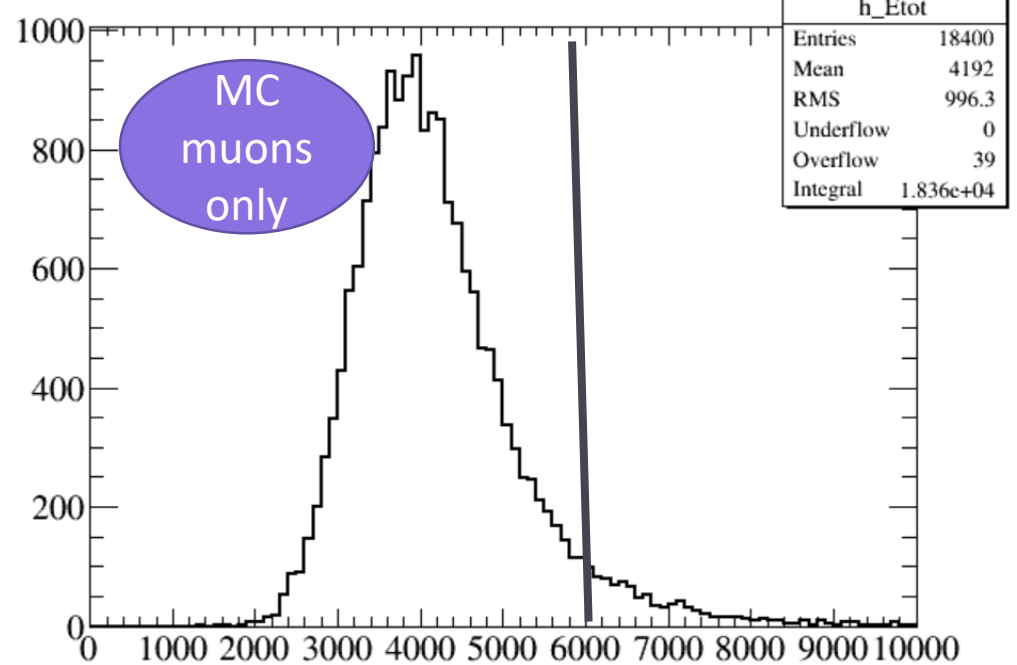
5 GeV backward electrons



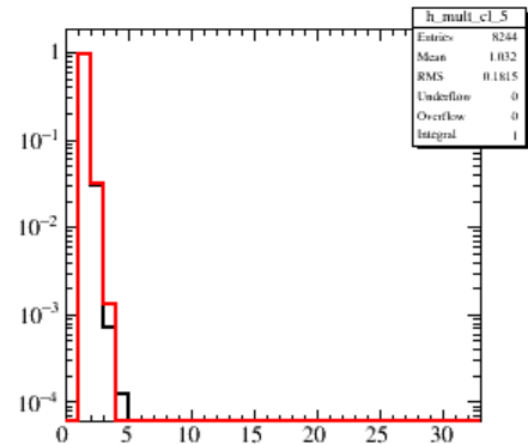
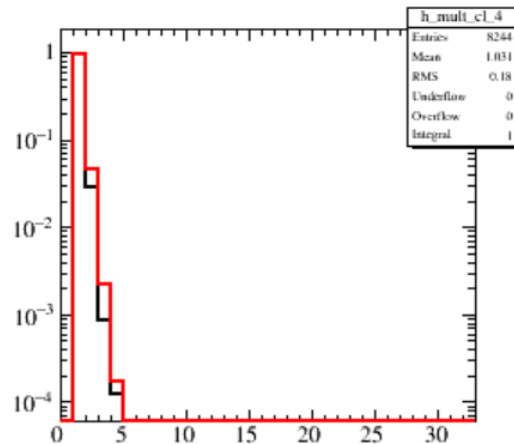
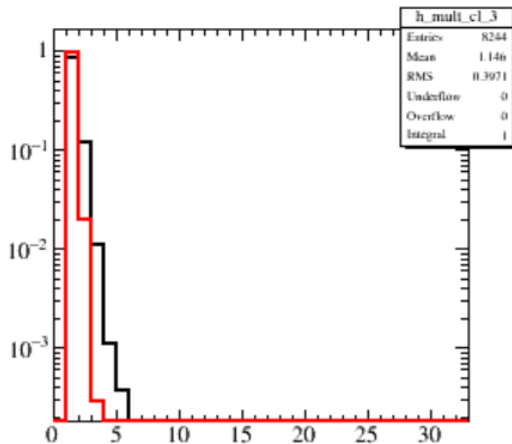
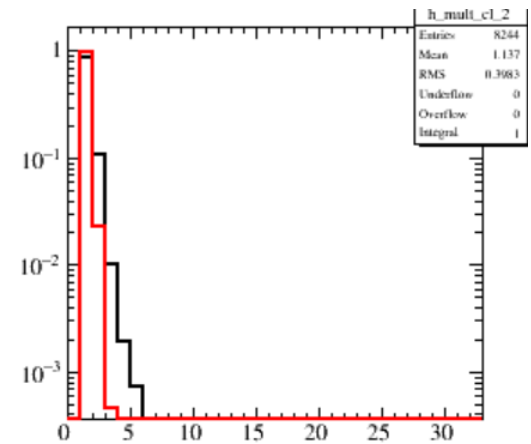
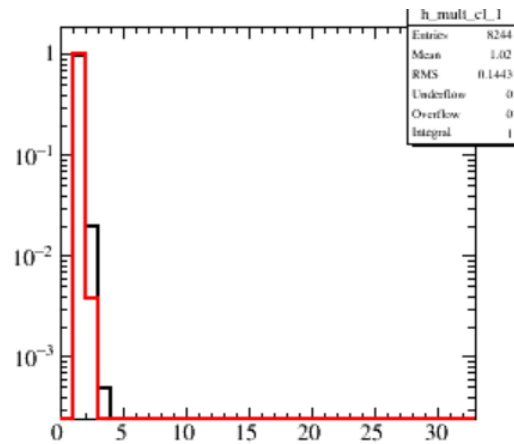
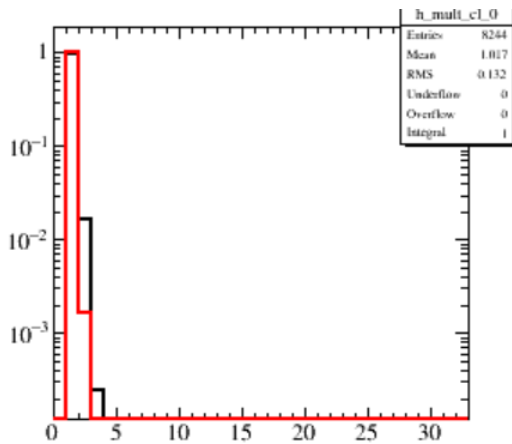
Classical cut!

Σ ADC < 6000 counts

To mitigate the background from electrons

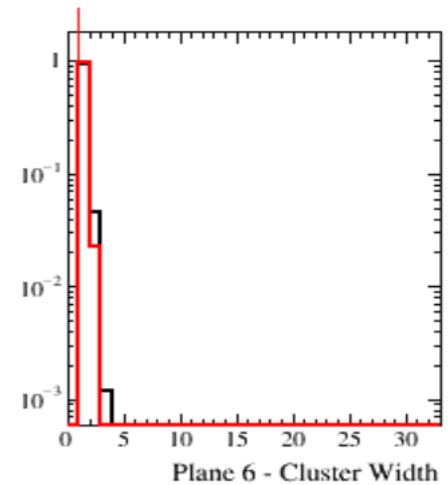
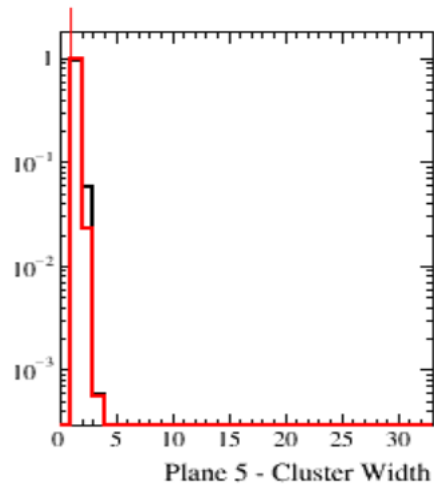
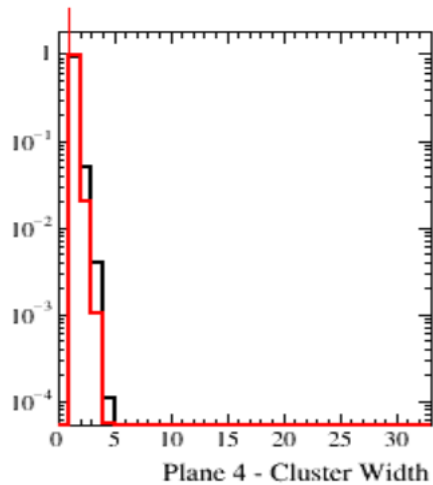
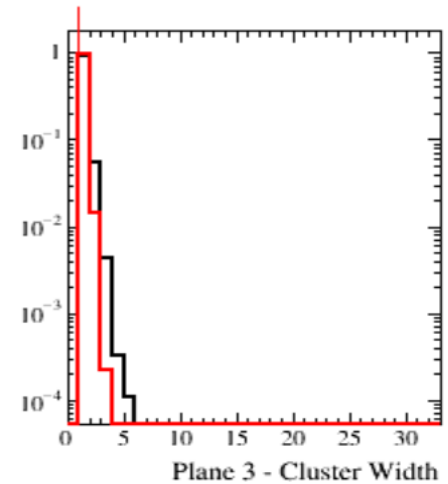
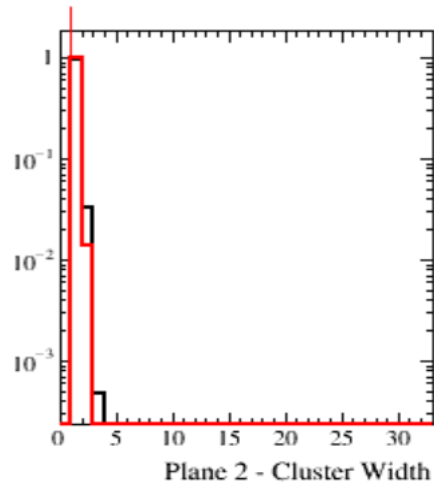
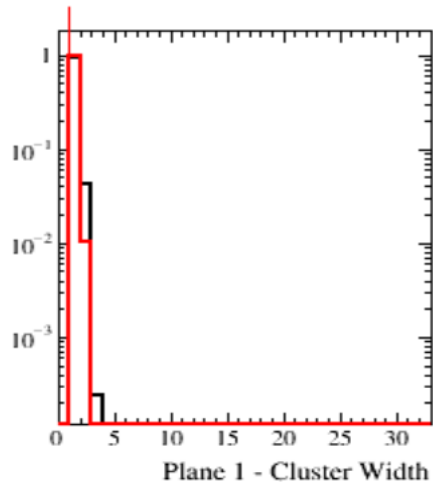


Cluster multiplicity

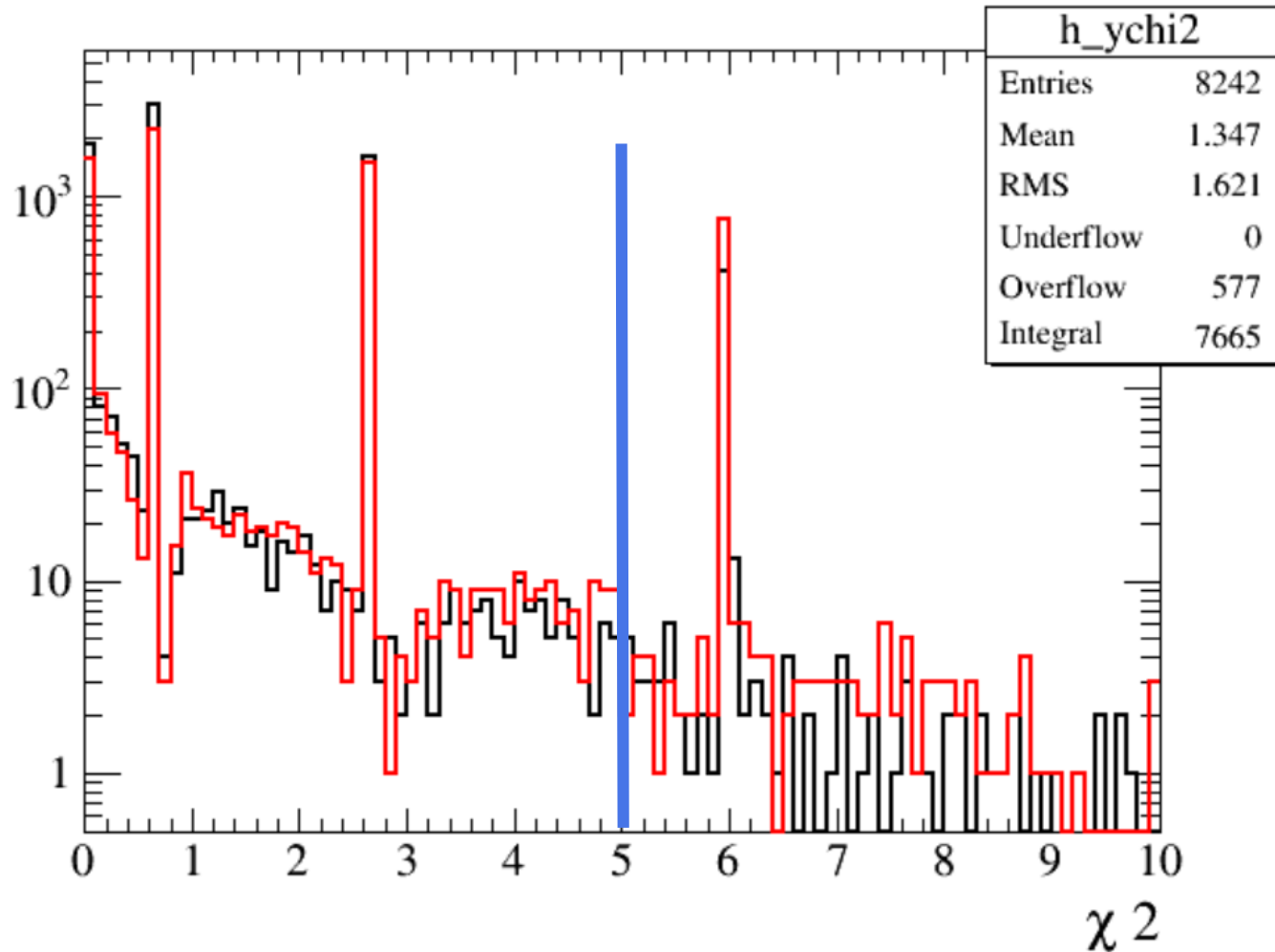


Nb of clusters

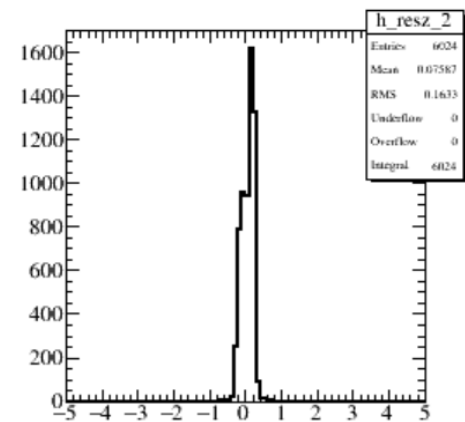
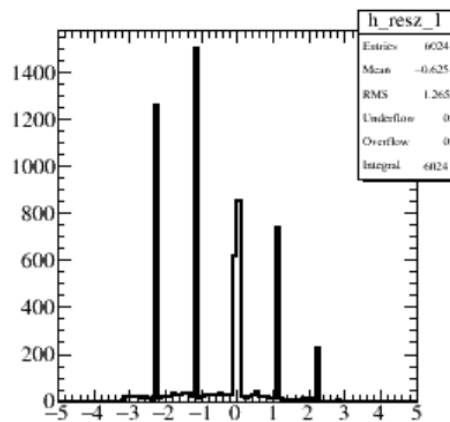
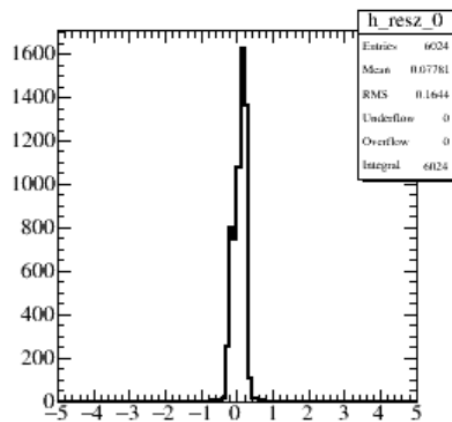
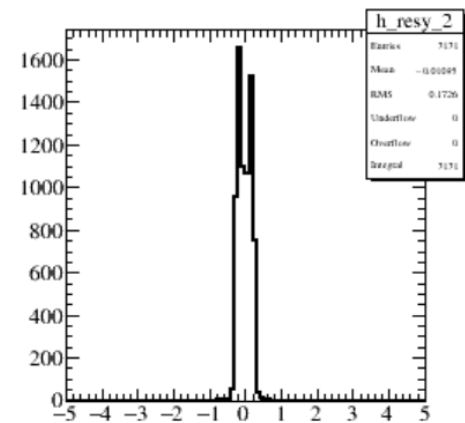
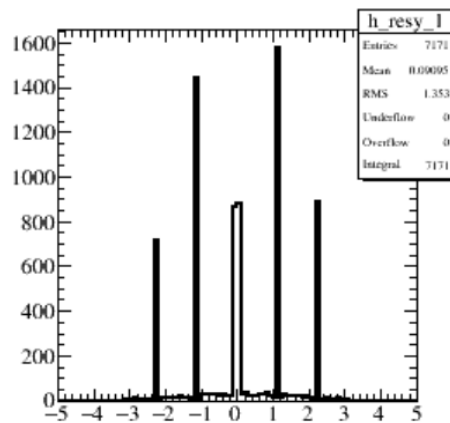
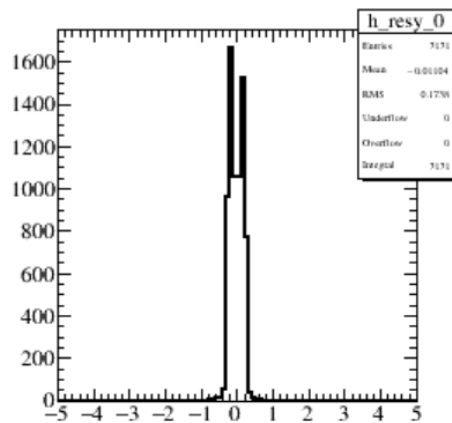
Cluster Width $\rightarrow W \leq 3$



χ^2 of the fitted track in y & z direction

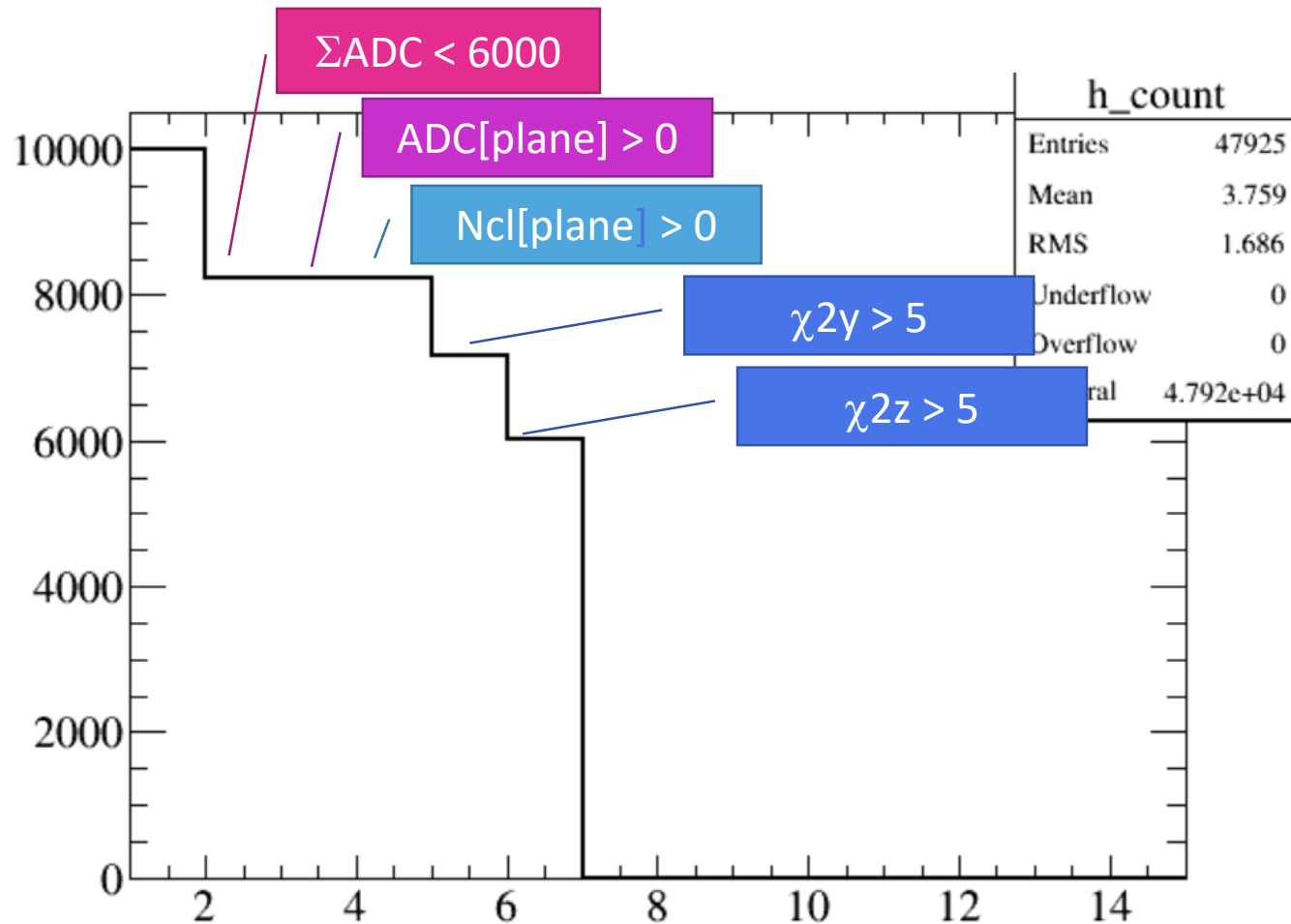


Residuals after χ^2 cut < 5



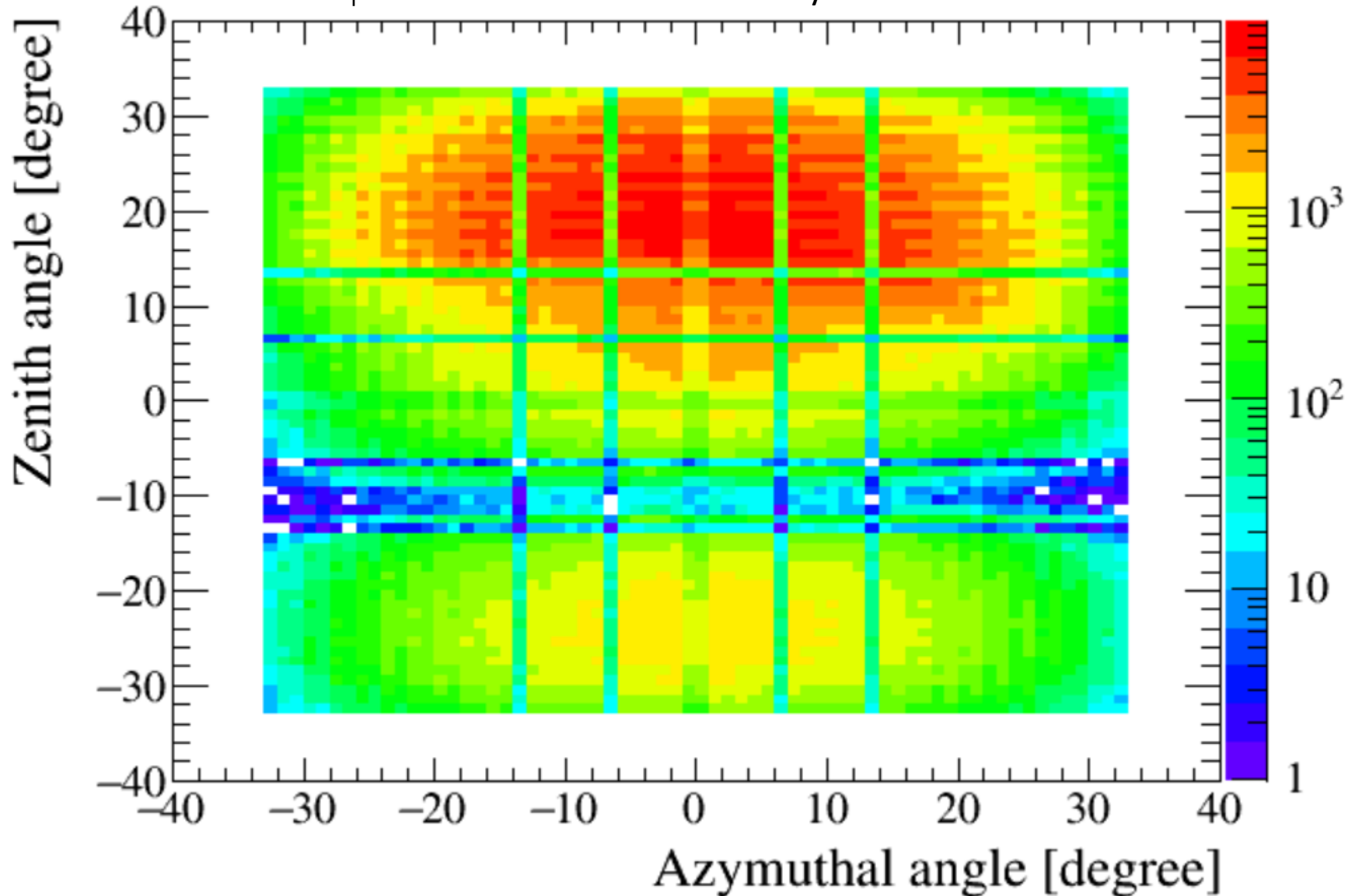
cm

Cut flow (data)

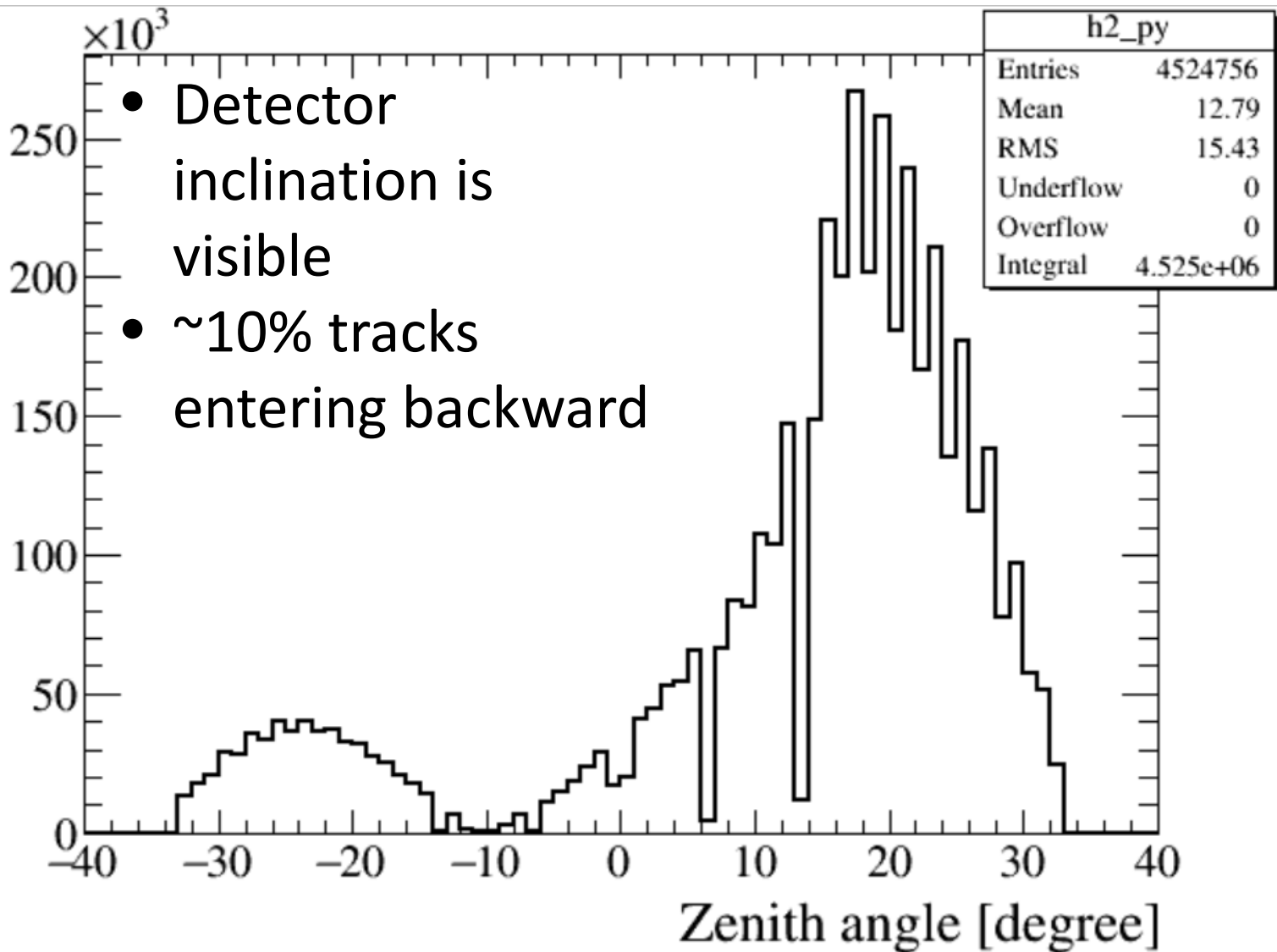


All events : ~ 5 millions tracks

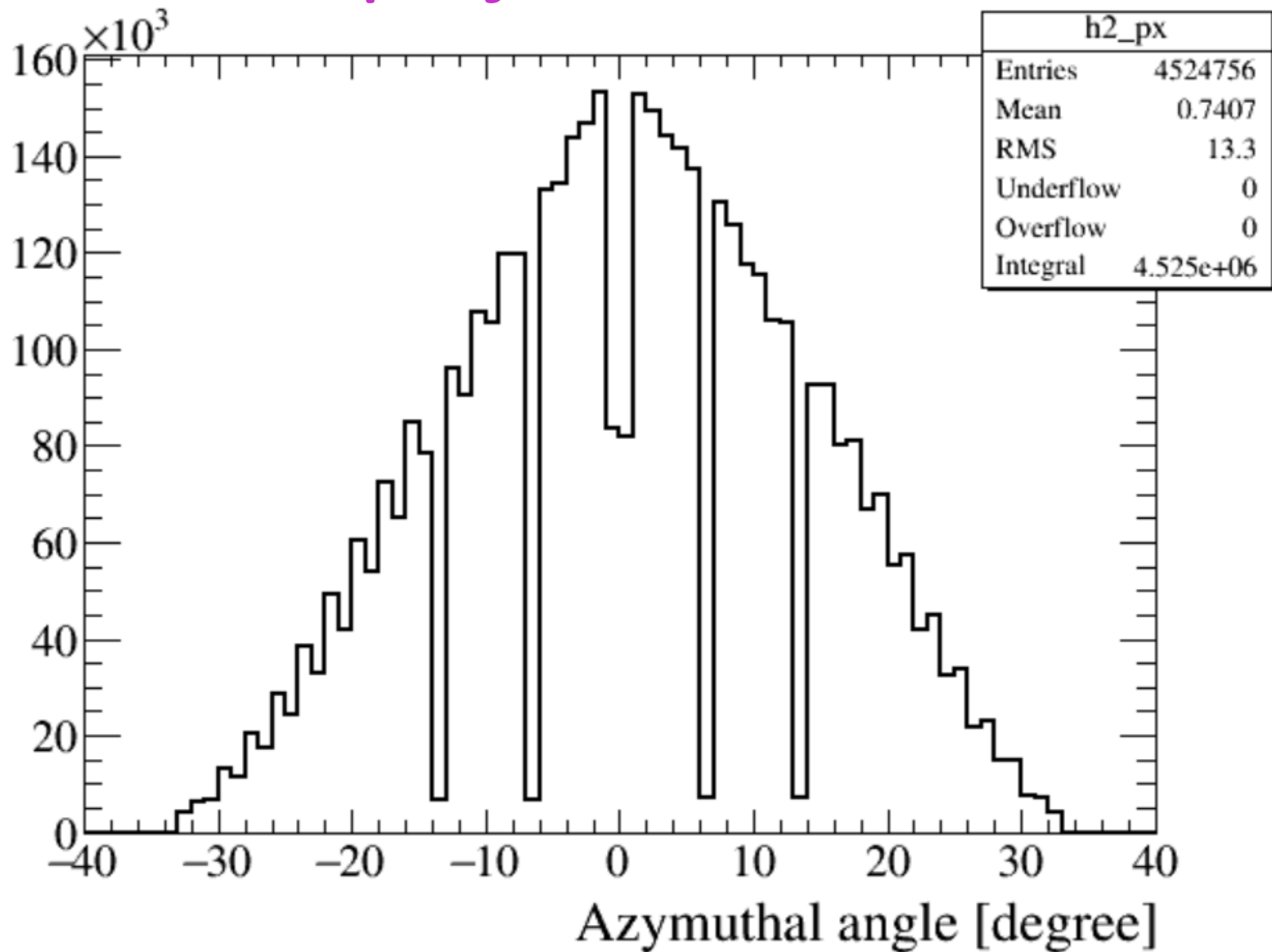
φ and θ are in the coordinate system of the detector



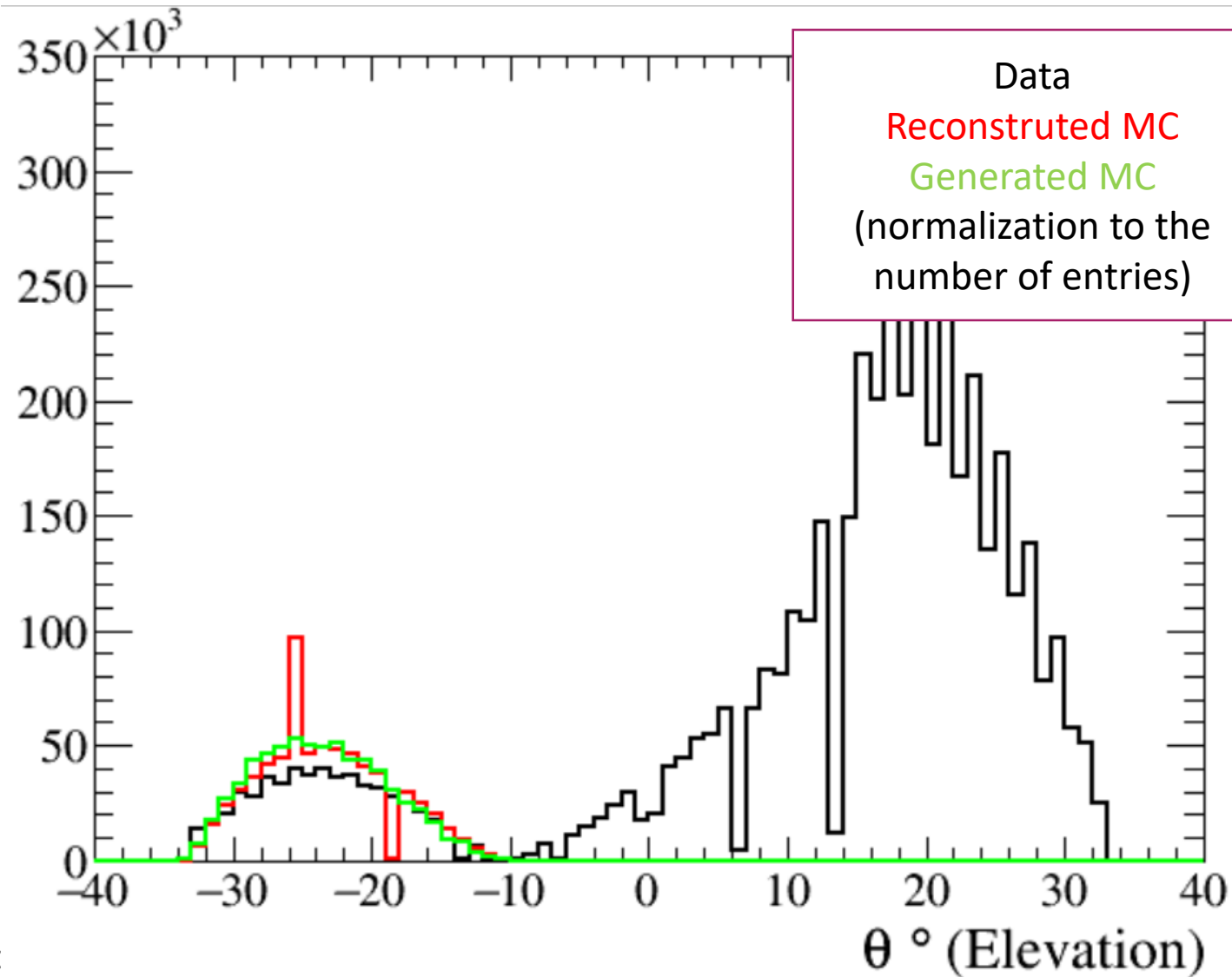
Data :Theta projection



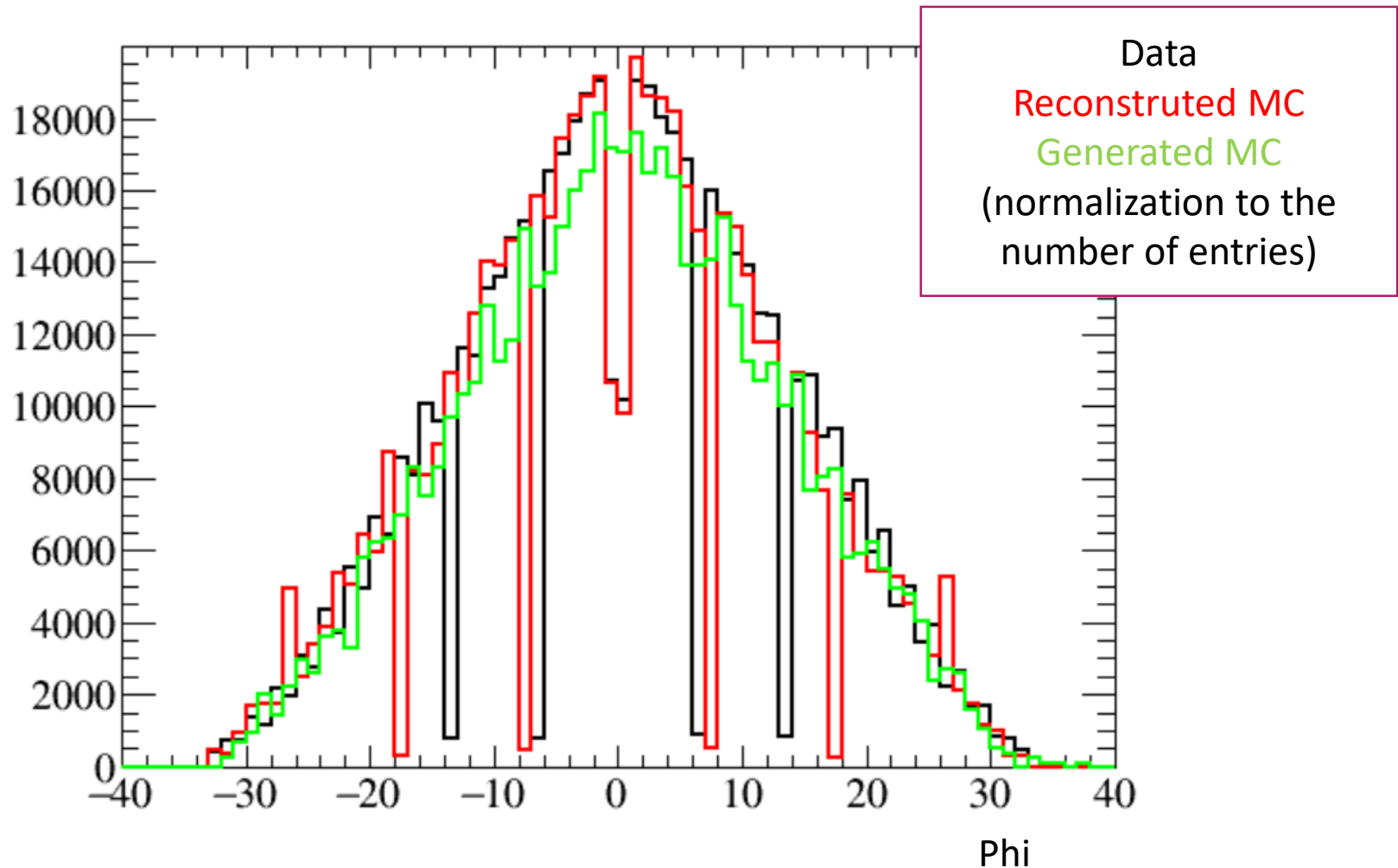
Data : Phi projection



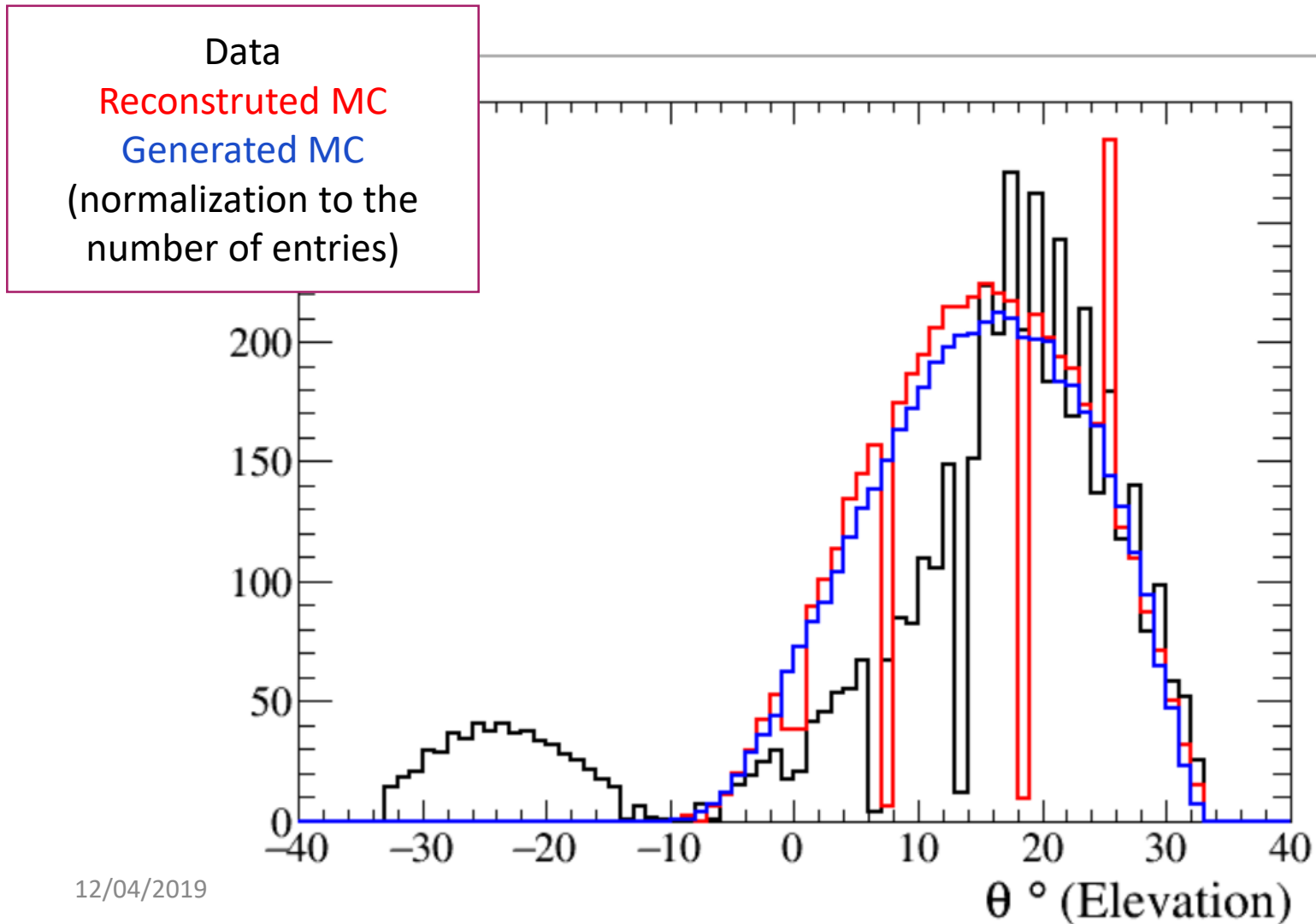
Bwd gen. tracks : theta



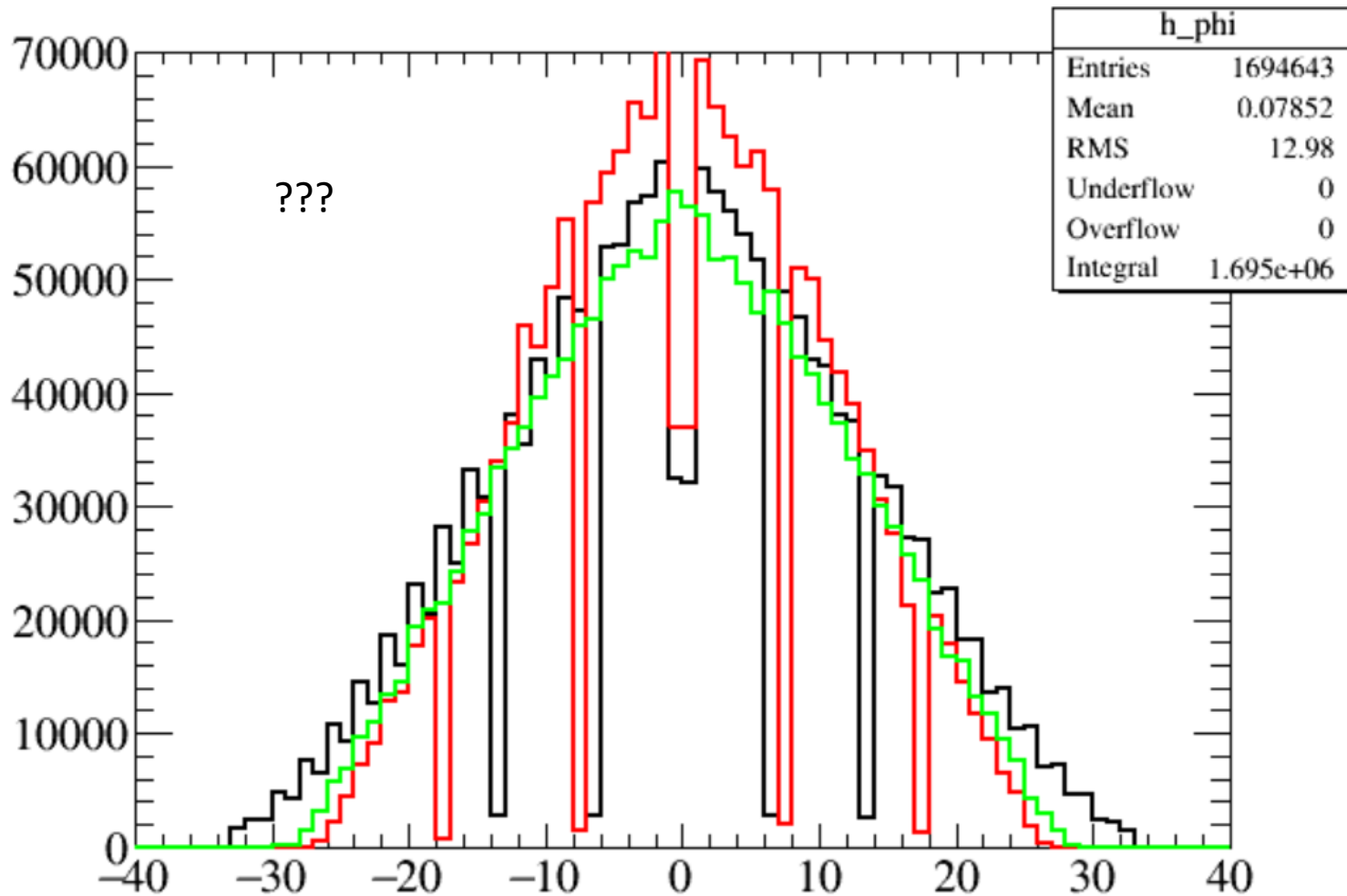
Bwd tracks and comparison with fast simulation: φ

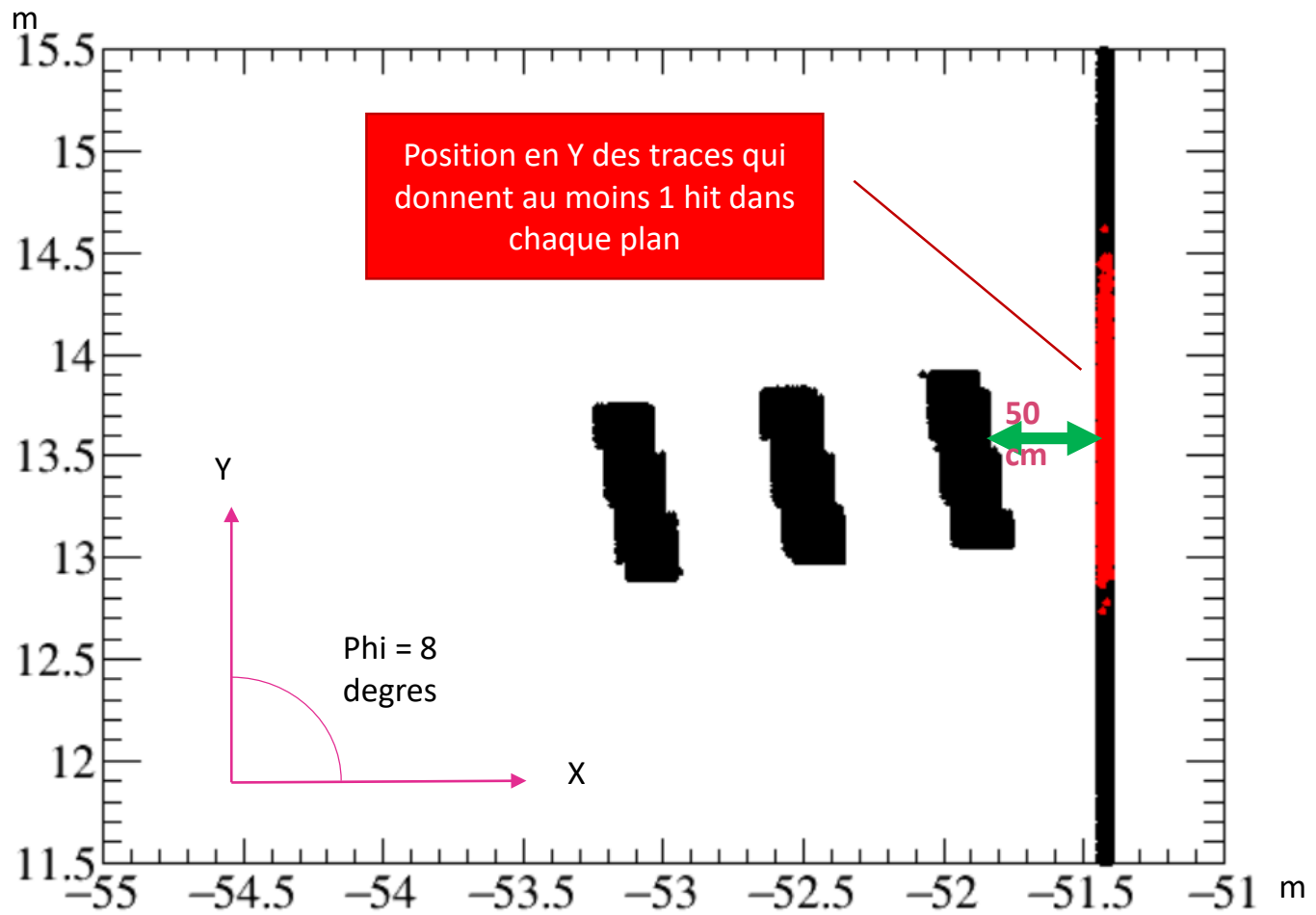


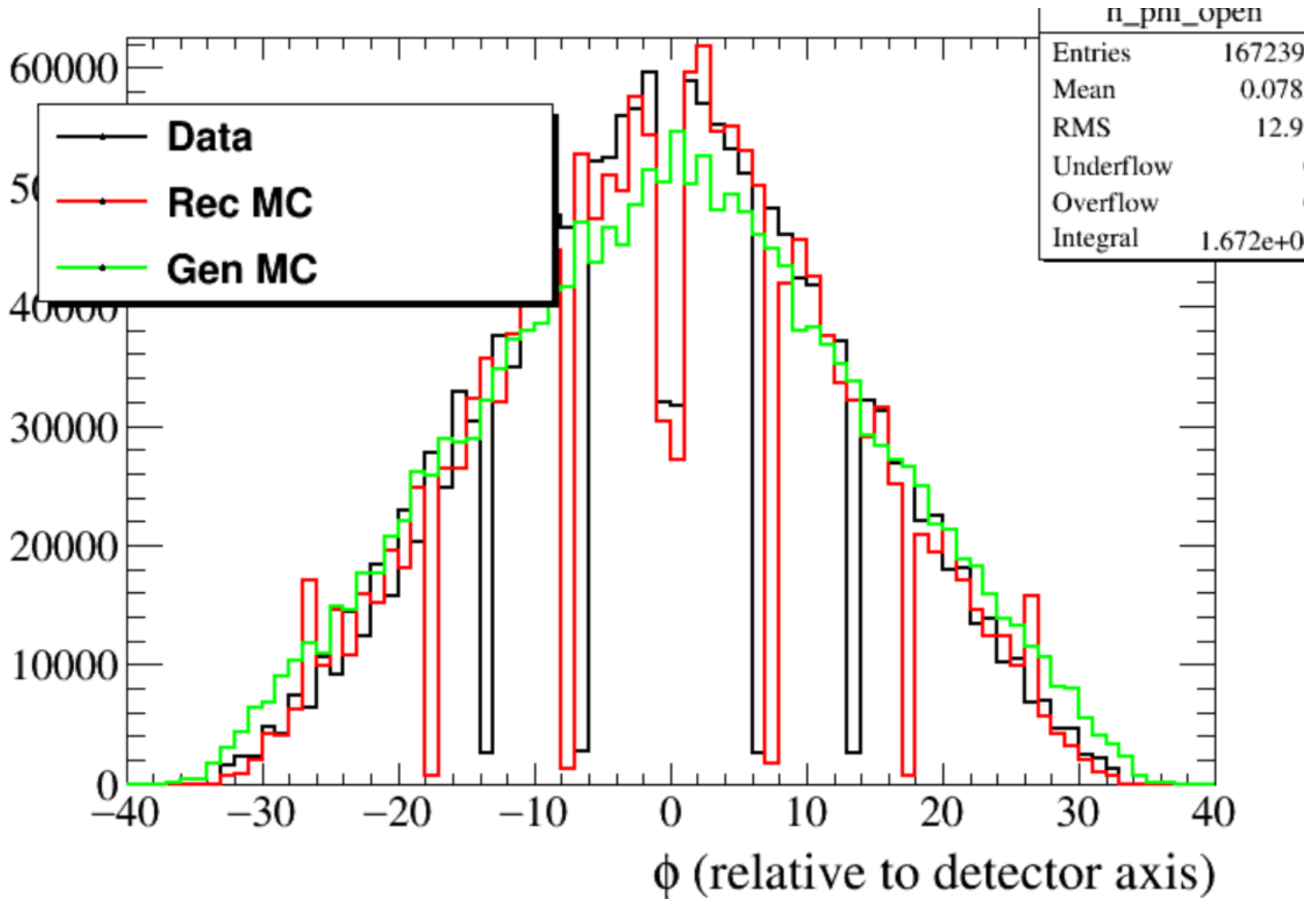
Forward track: theta comparison



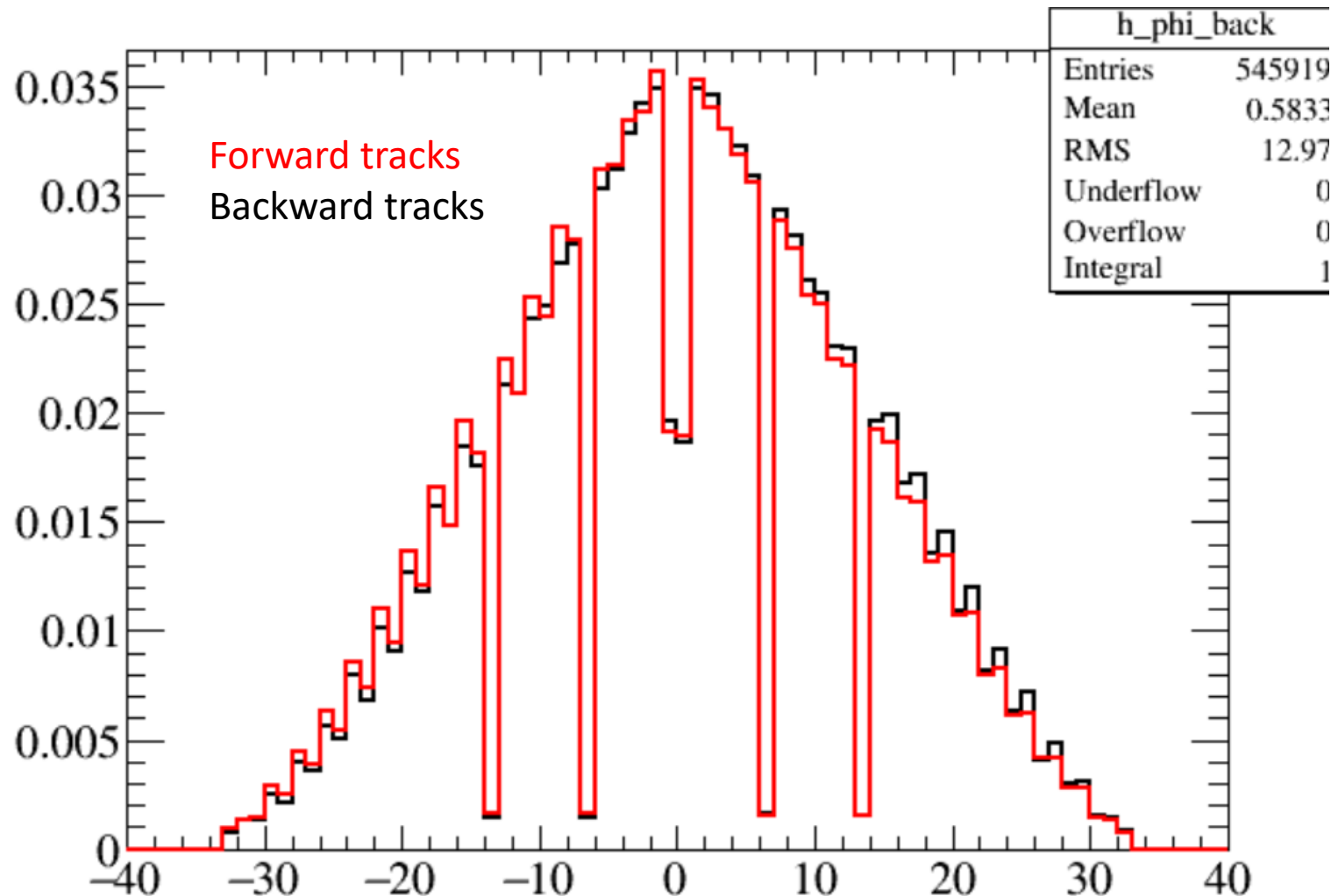
Phi – Open sky (Theta > 20 deg)







Comparison data phi fwd open sky/phi bwd



Outlook

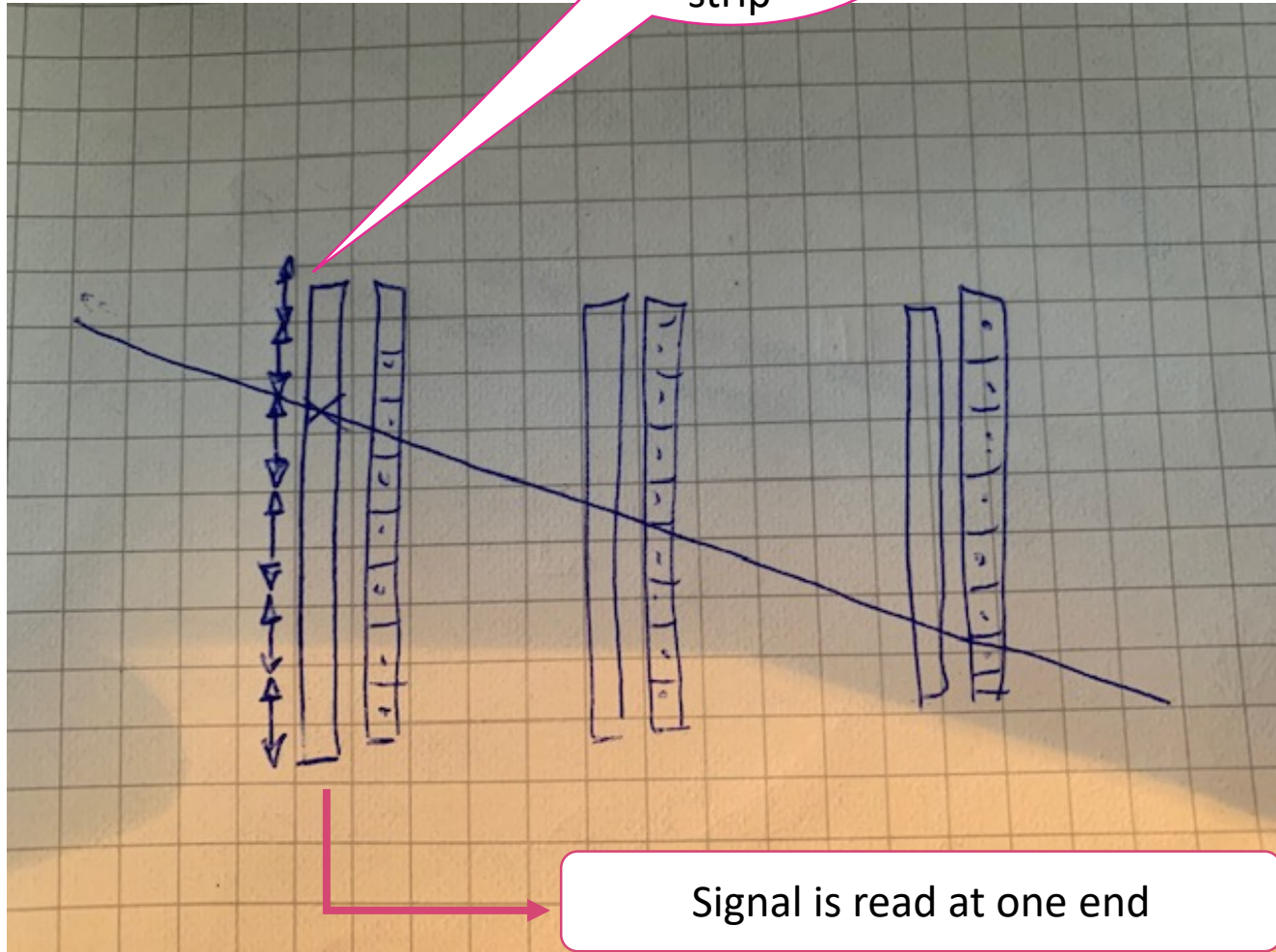
- Fast simulation useful to understand the basic principles
- Detailed simulation (see Yannis)

Improvement of simulation

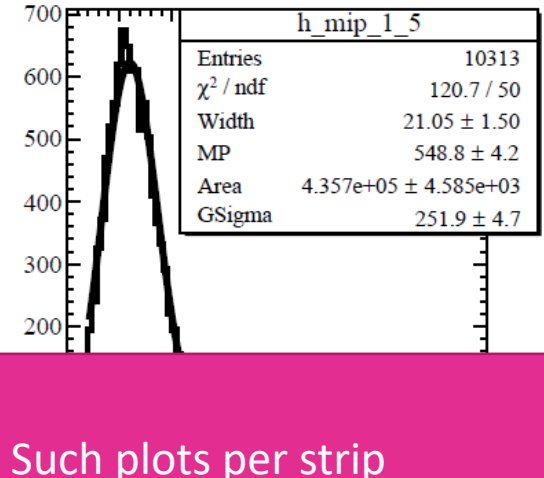
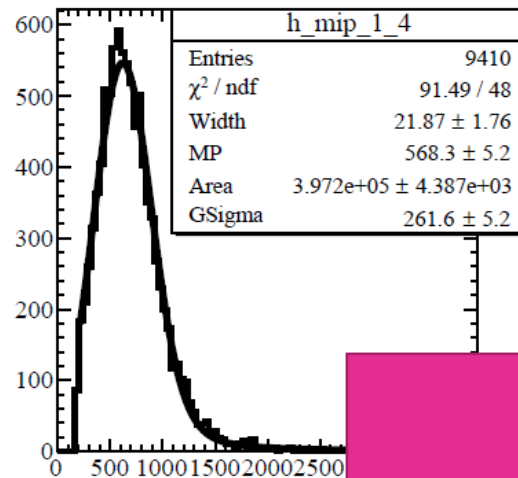
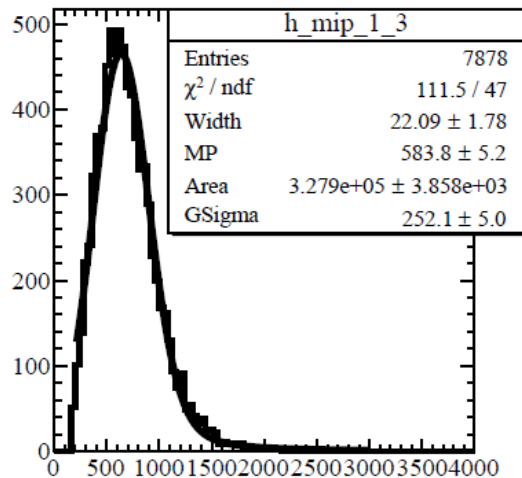
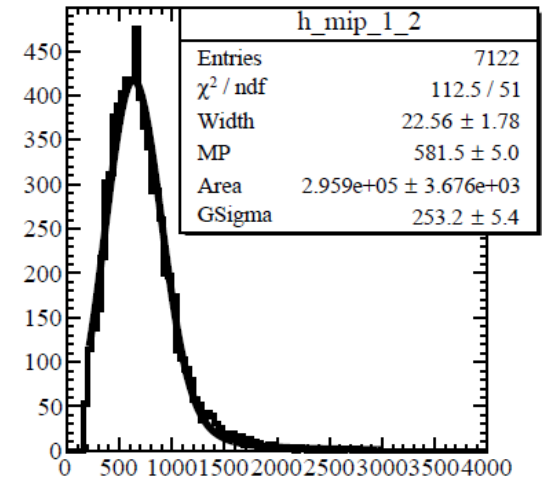
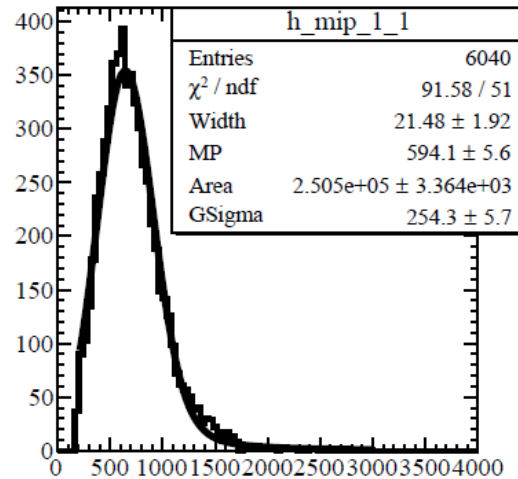
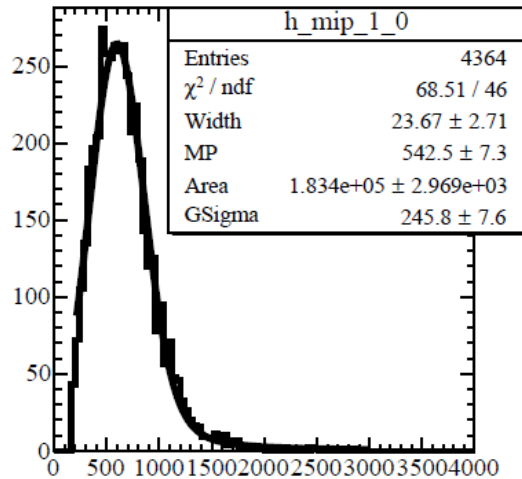
Status of Mip/attenuation studies

Principle

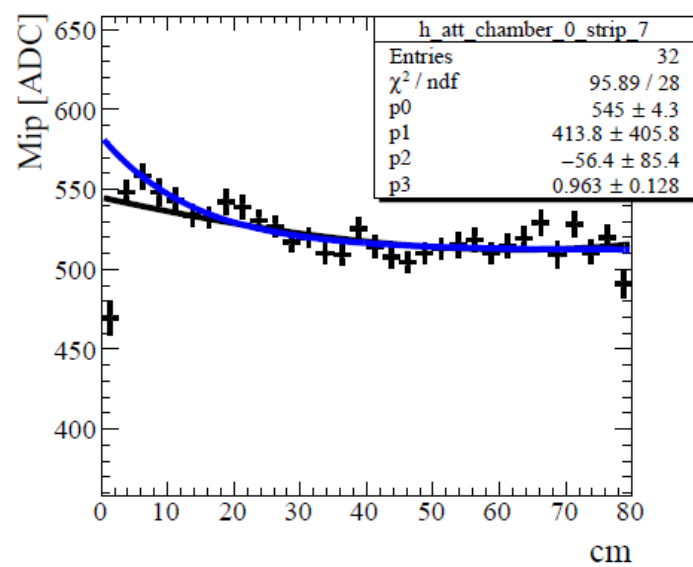
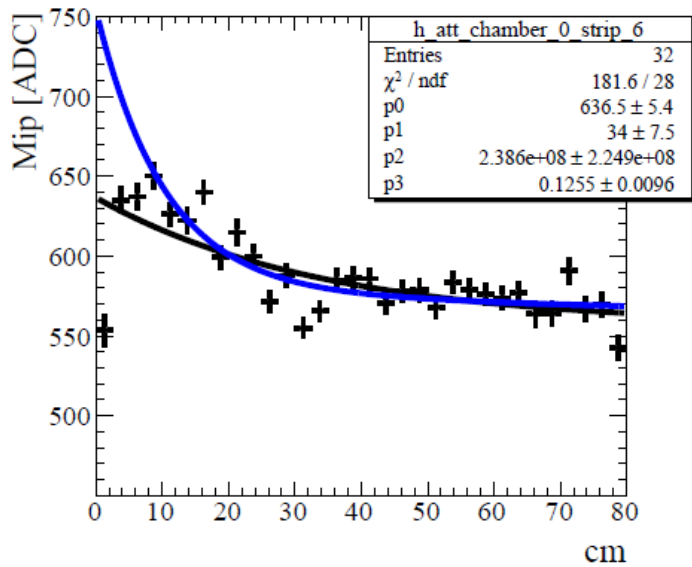
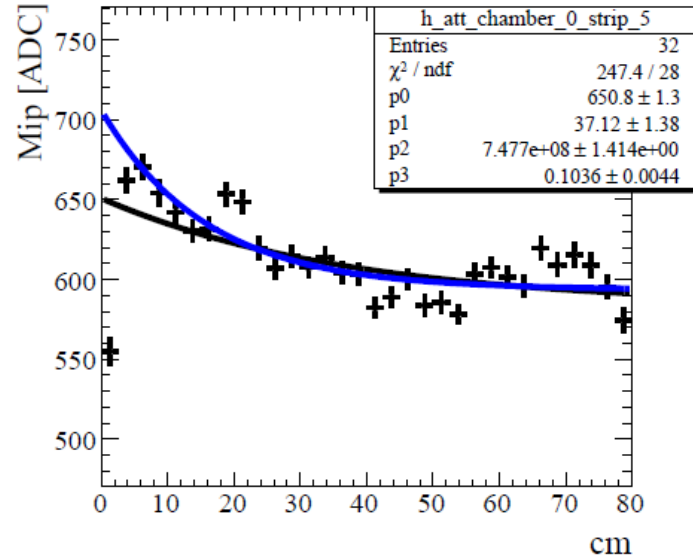
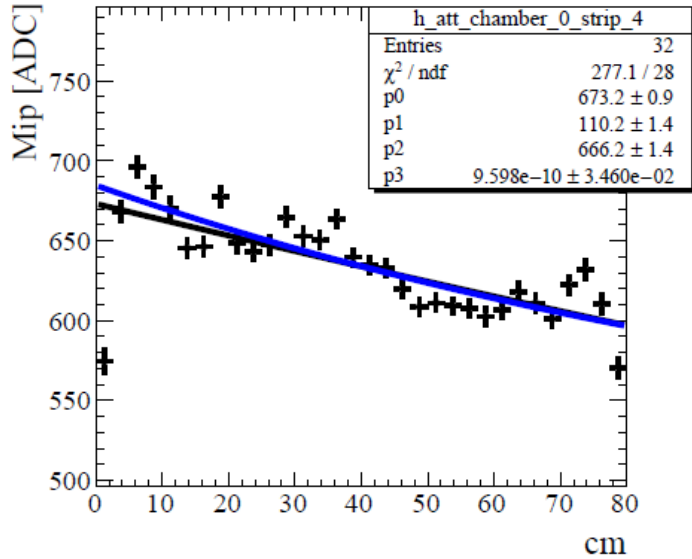
Events are grouped by 5 cm in each strip



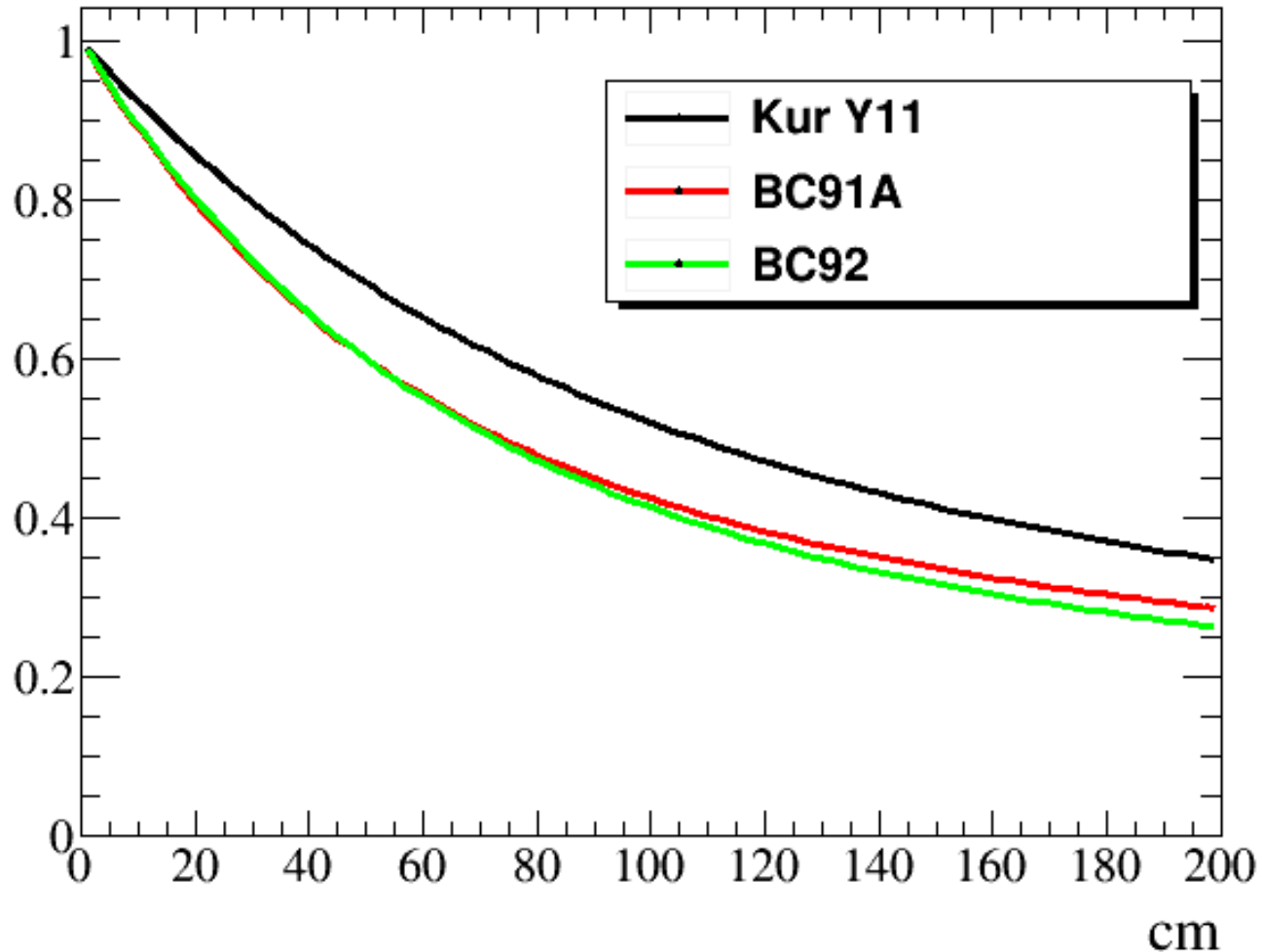
MIP distribution in 1 strip every 5 cm



16 Such plots per strip



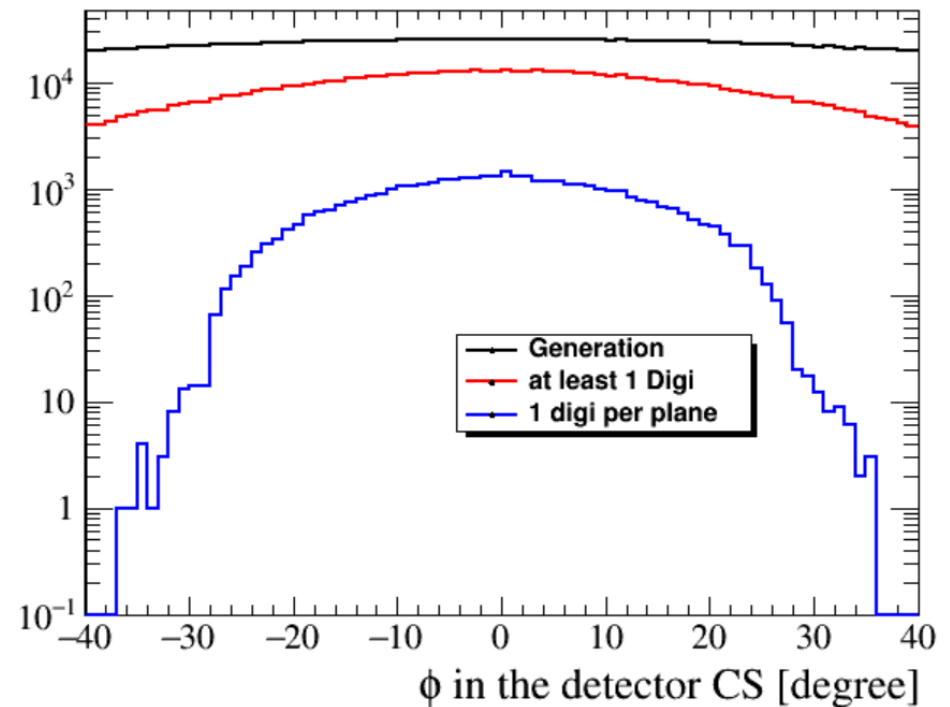
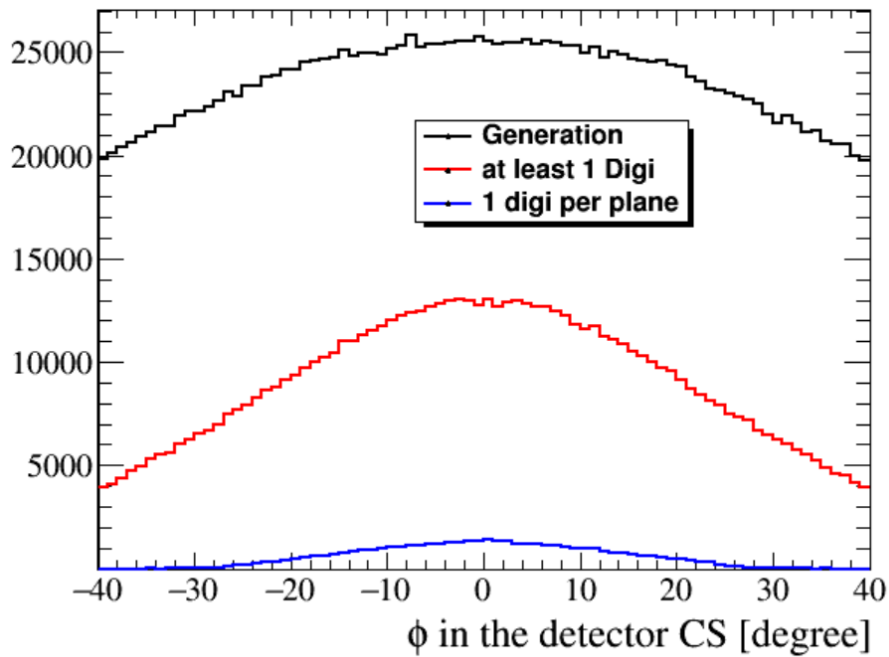
Attenuation from the spec.



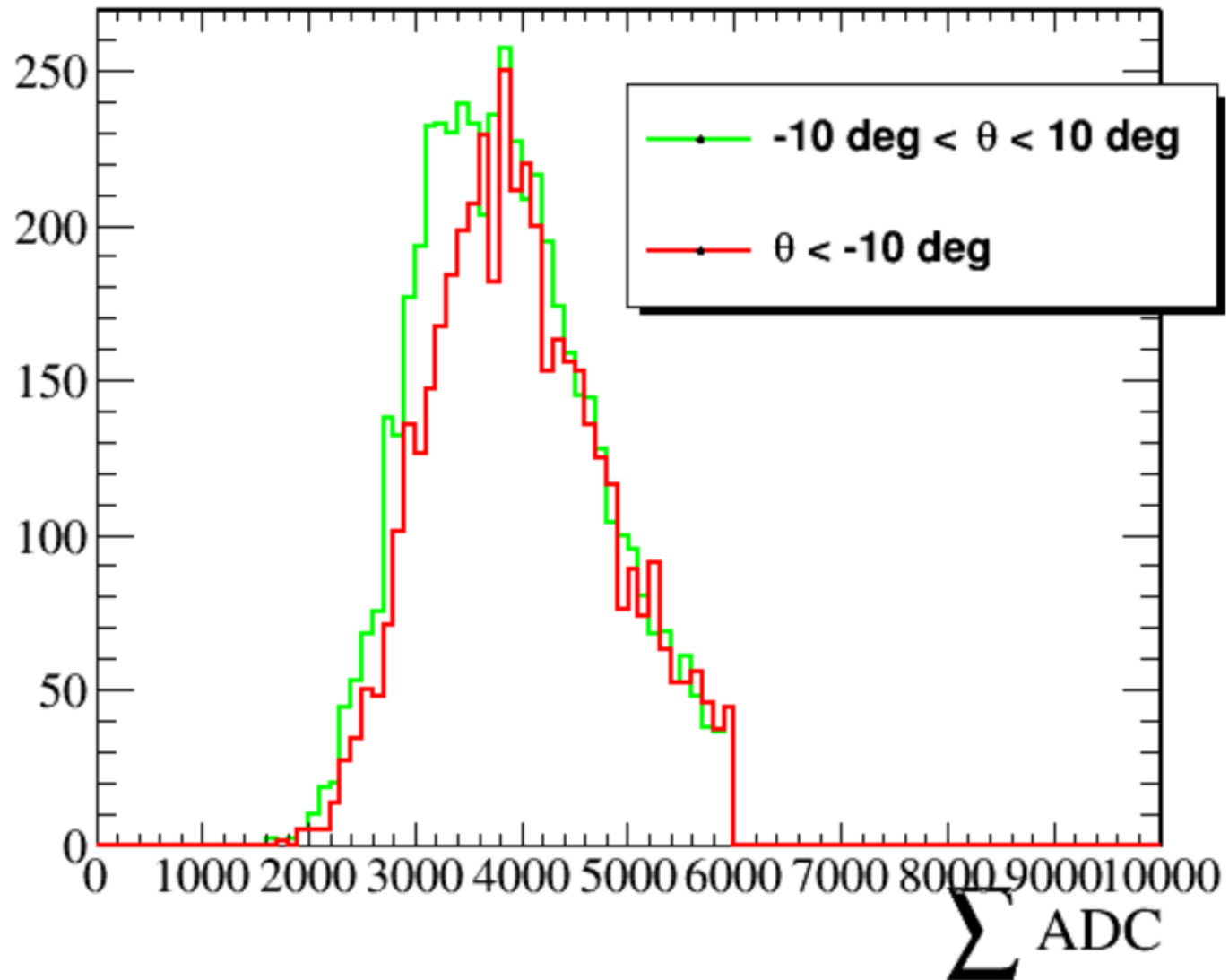
On going work

Generated phi :

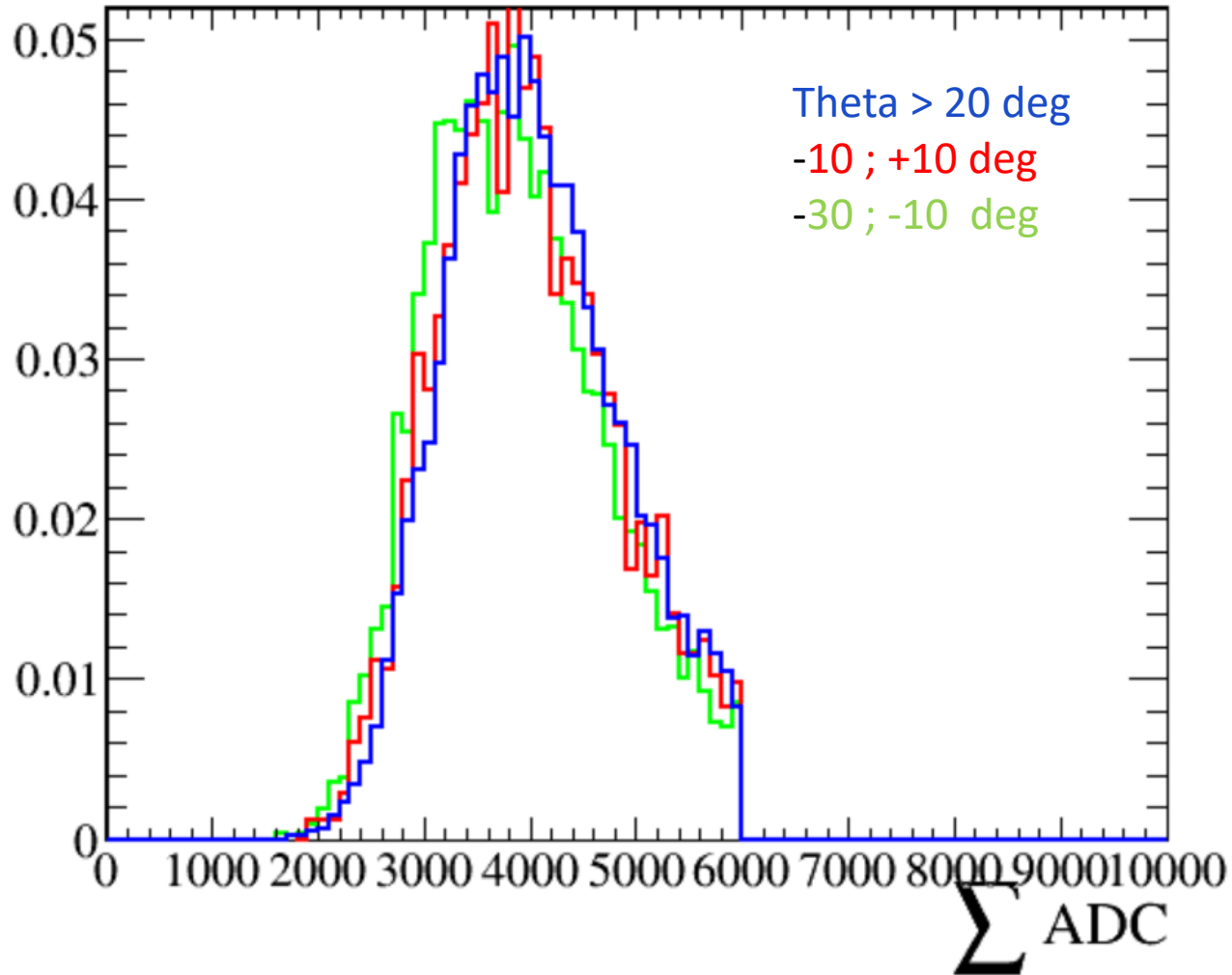
1 strip/plane seems to create the cut-off



Etot distribution

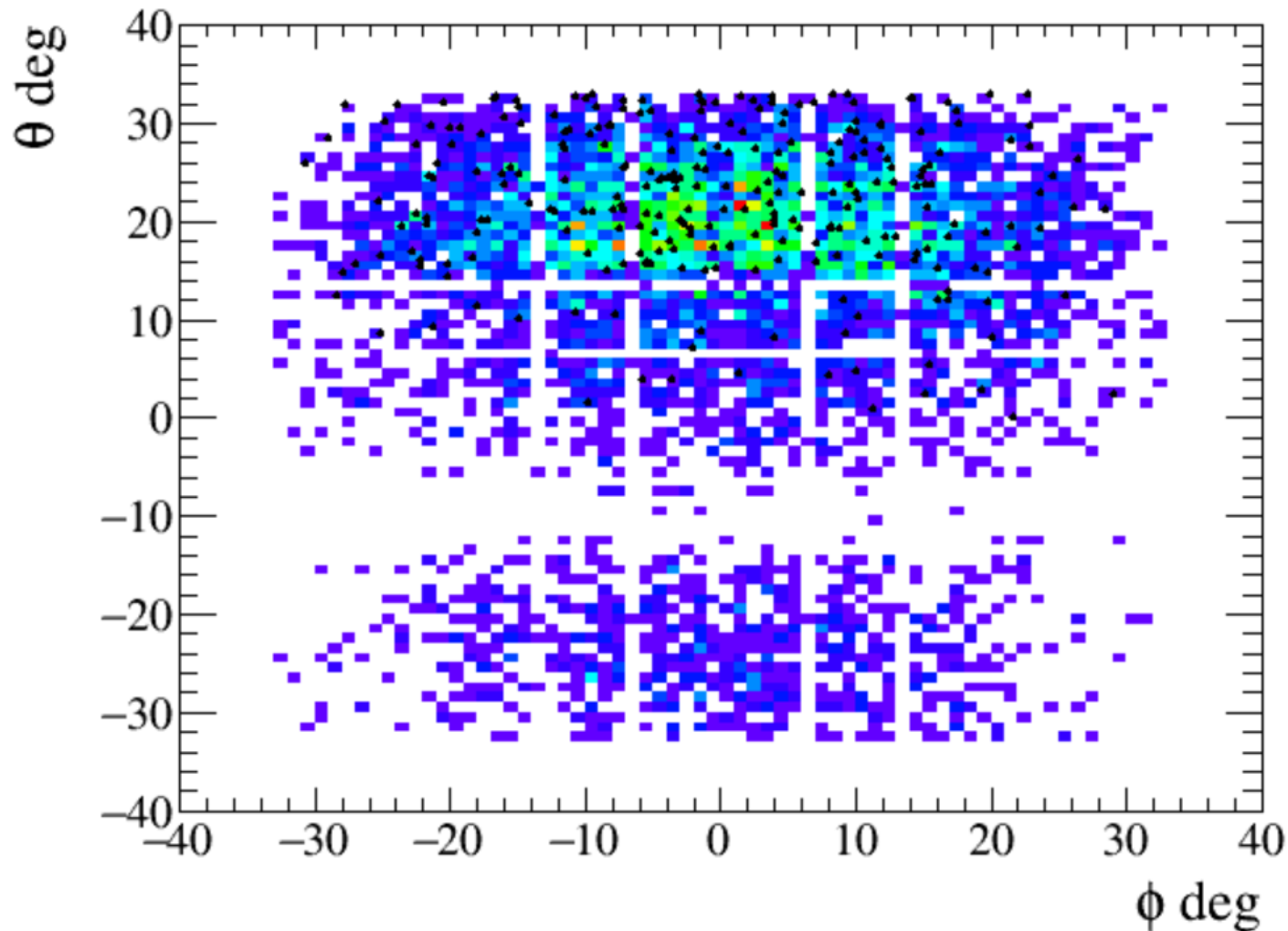


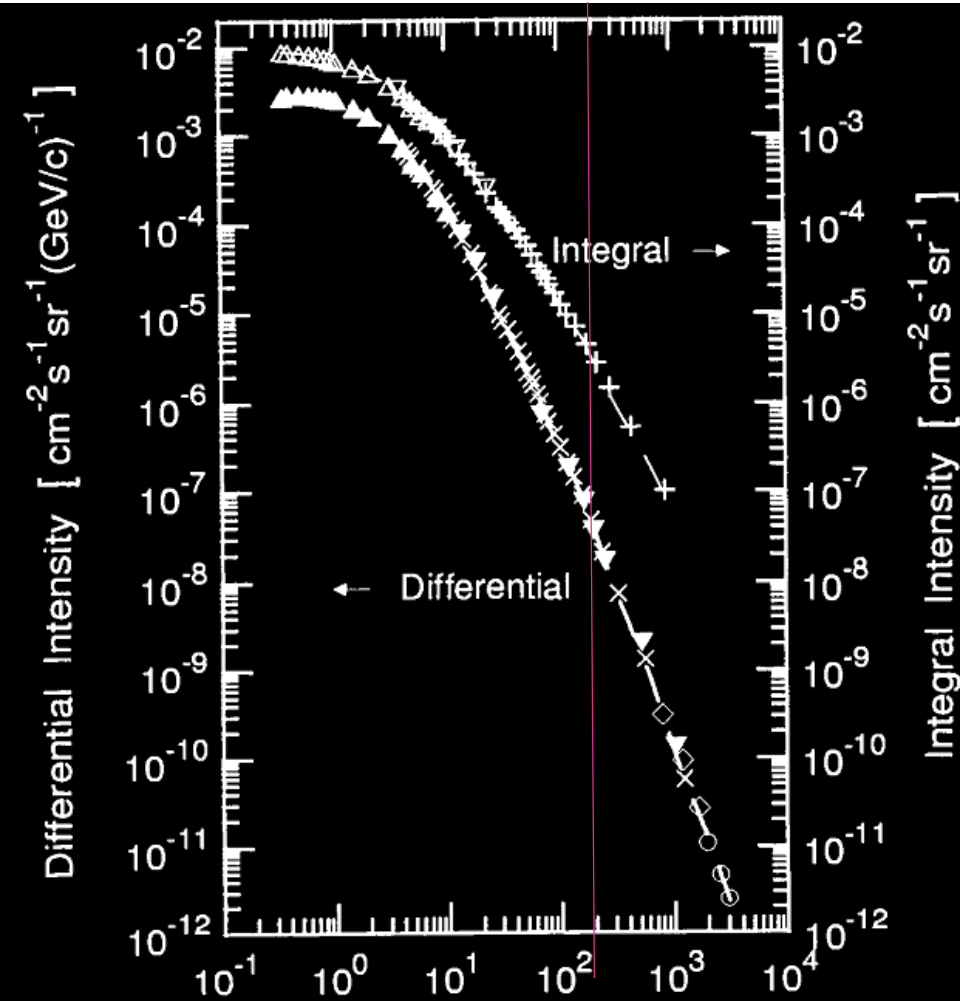
Normalized



Backup

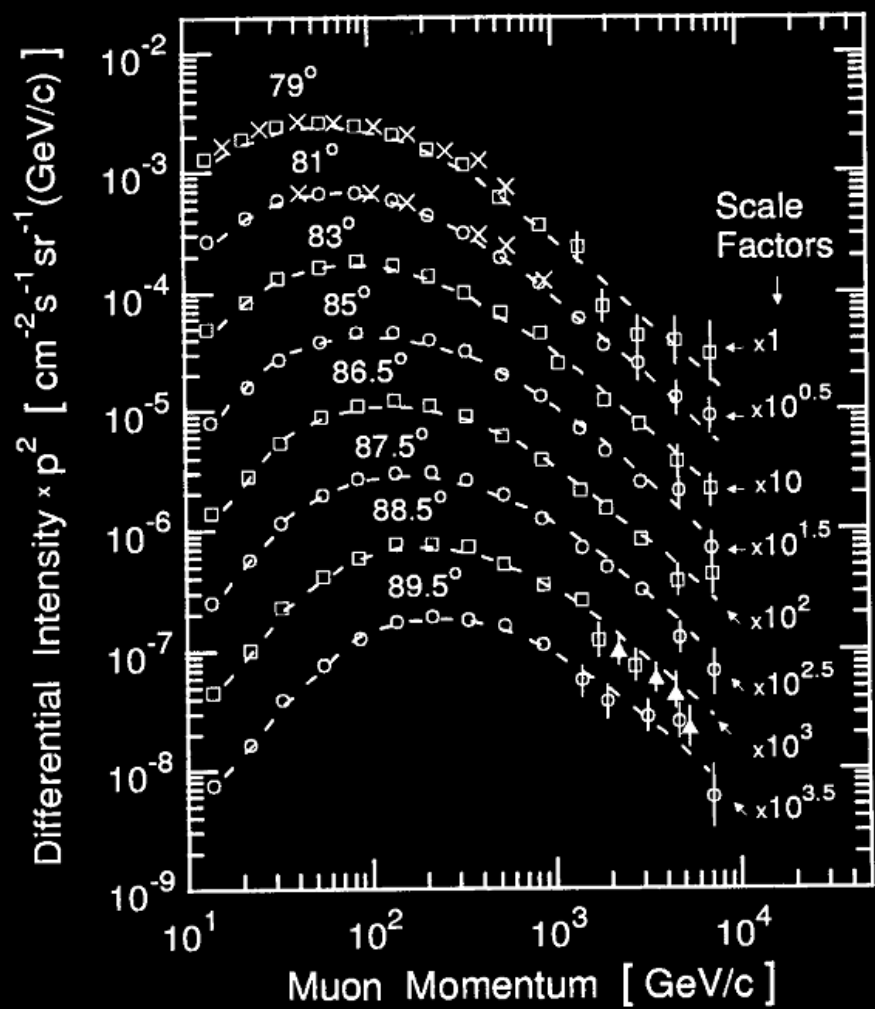
Phase space corresponding to the first strip in the first layer



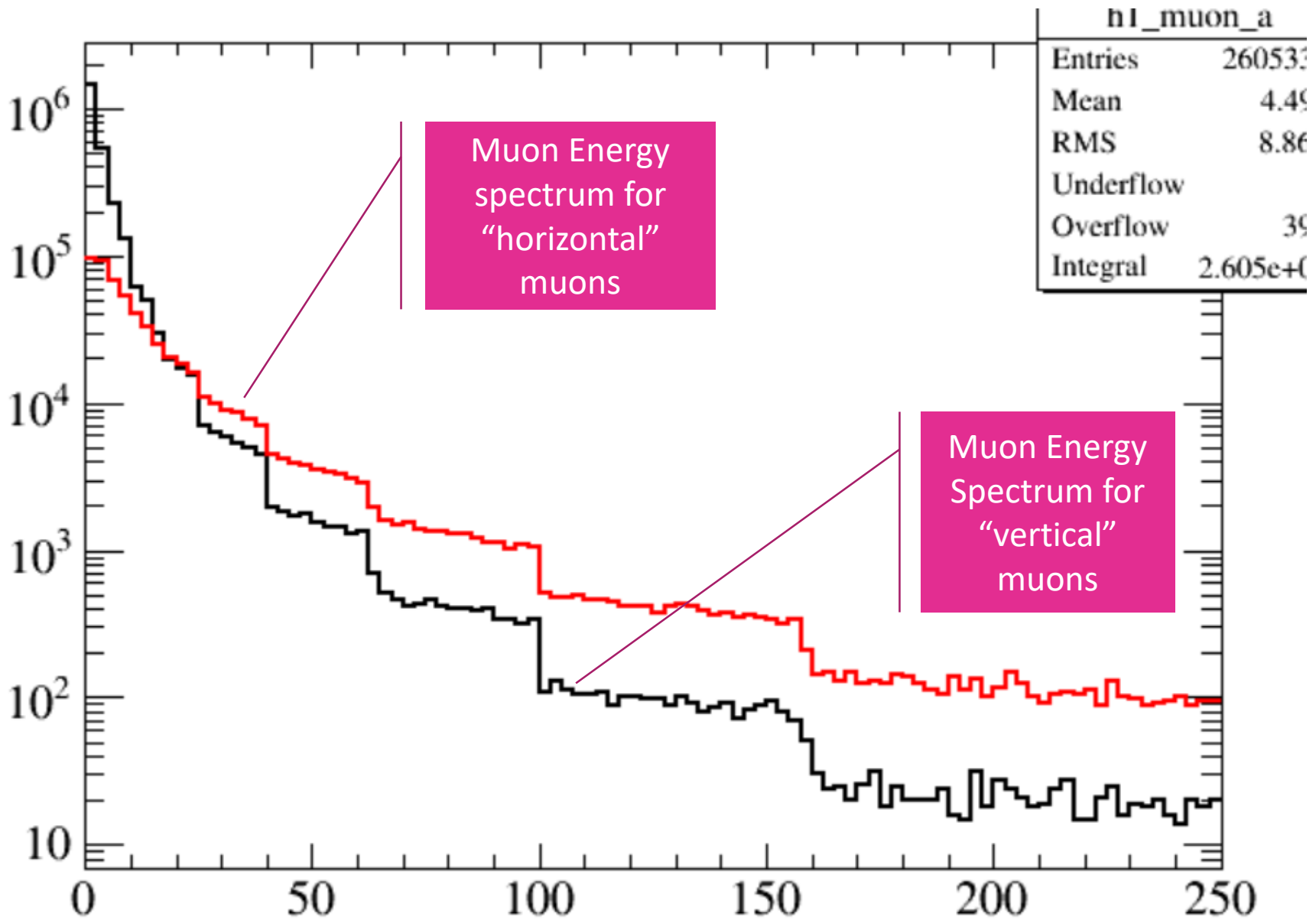


Muon Momentum [GeV/c]

- | | | | |
|------|--------------------------|------|---|
| ▼, ▽ | Appleton et al. (1971) | ▲, △ | Barbouti and Rastin (1983) |
| ×, + | Rastin (1984a) | ○ | Komori (1977) and
Komori and Mitsui (1979) |
| ◇ | Thompson et al. (1977b) | | |
| - - | best fit, Rastin (1984a) | | |



- | | |
|-------|---|
| □, ○ | DEIS 78° - 90°, Allkofer et al. (1979b, 1981) |
| × | Kiel-DESY, Allkofer et al. (1977b) |
| ▲ | MUTRON, Muraki et al. (1979) |
| - - - | model calculation, Maeda (1970, 1973) |



Generator : CRY

