

Muon tomography applied to the Apollonia Tumulus (Preliminary results)

Radiographie par particules cosmiques du tumulus d'Apollonia

Antoine CHEVALIER, Jacques MARTEAU, Amelie COHU, Jean-Christophe IANIGRO

IPN Lyon - CNRS - Université de Lyon - Muon Sight

Marina ROSAS-CARBAJAL

IPGP - CNRS - Université Denis Diderot

Grigorios TSOKAS

Aristotle University of Thessaloniki

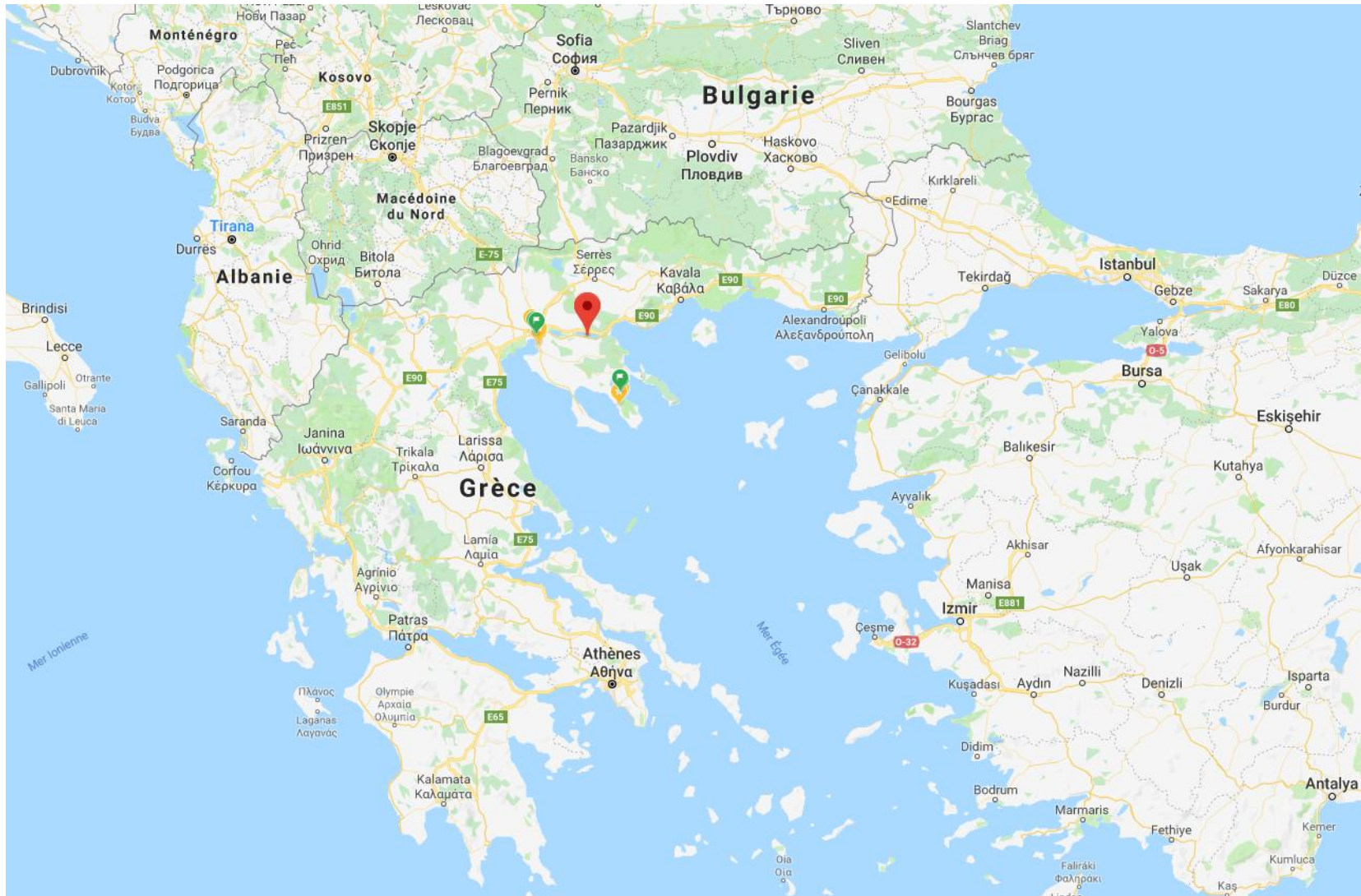
Context of study

Geography

Historical context

Measurements

South Macedonia / North from Chalkidiki peninsula



Between Apollonia & Nea Apollonia

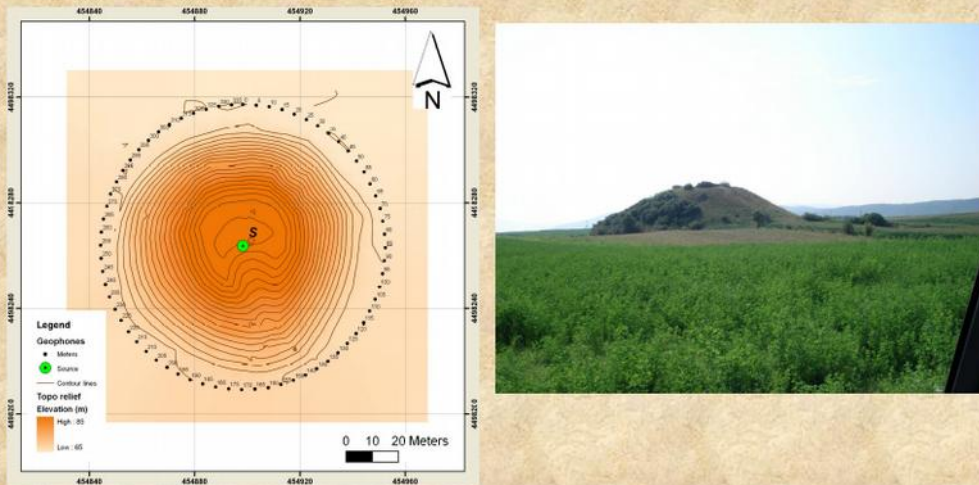


Tumulus & muon-mobile



A massive “Macedonian-type” tumulus

Tumulus in Apollonia North Greece
Diameter 100 m; Height 24 m. The biggest in Greece?



CHALLENGES OF THE GEOPHYSICAL SURVEYS ON TUMULI

The material of the tumulus might be highly inhomogeneous (a fact that renders difficult the data interpretation). In some cases, stratification of the tumulus infill material has been observed.



G. Tsokas et al. (2015)

- **What kind of information can the Muon tomography provide to non-destructive tumulus survey ?**

Muon Tomography applied to tumulus targets

Muons: a secondary product of cosmic particles

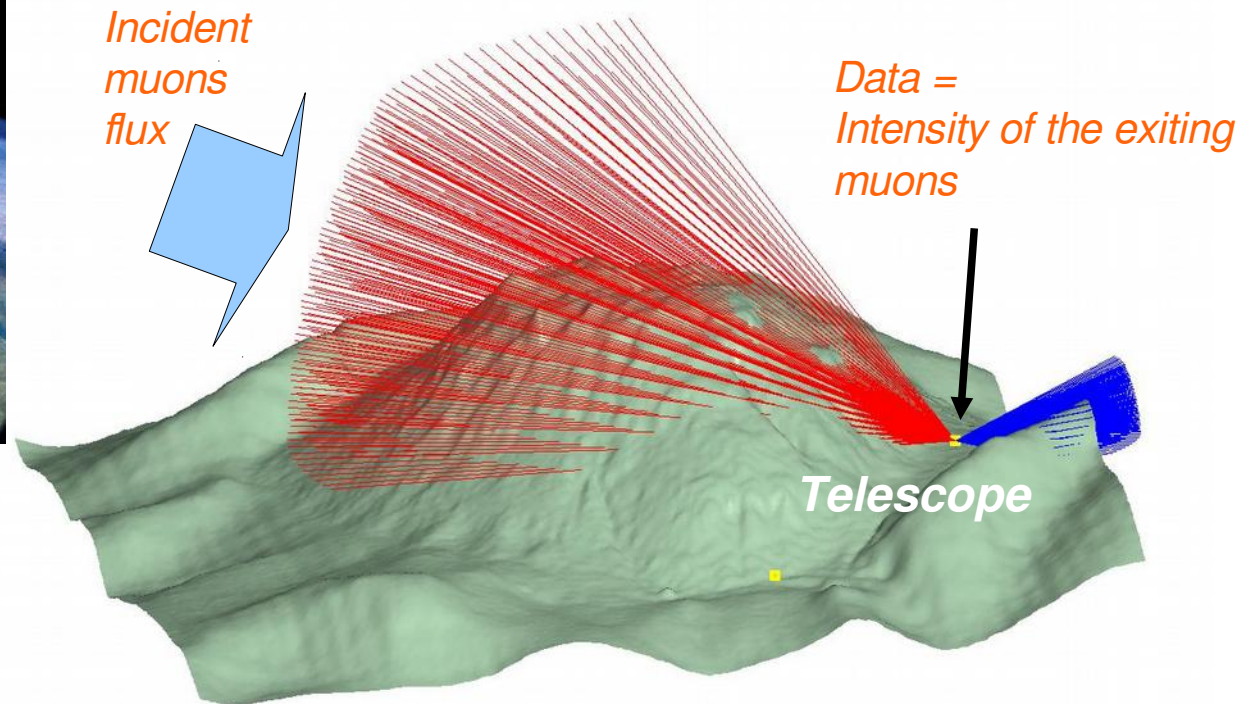
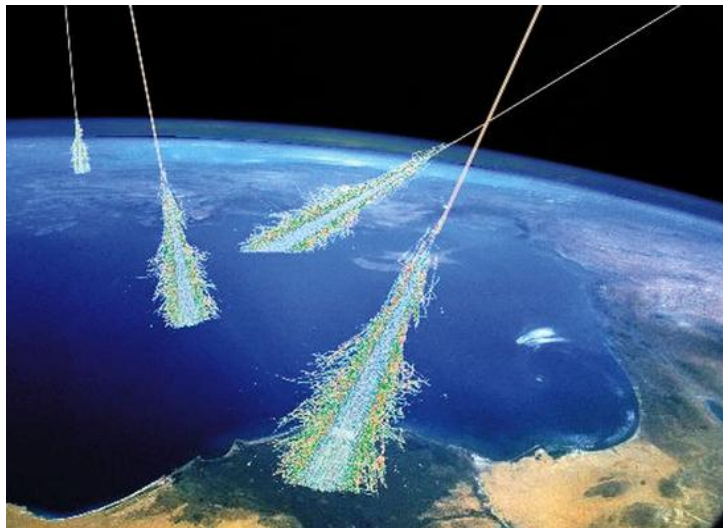
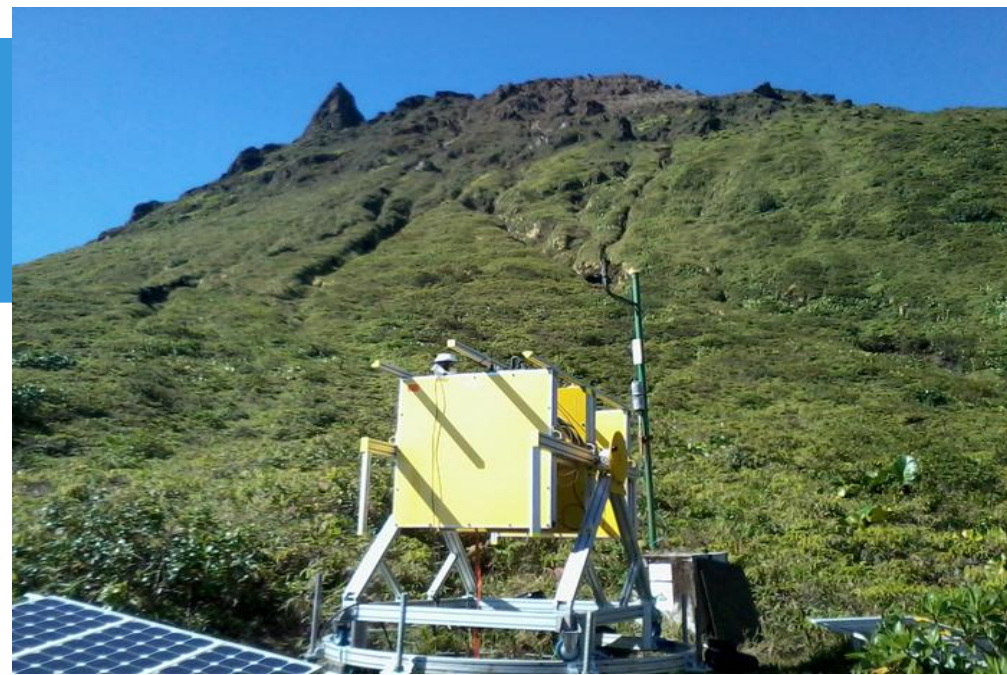
Muon Telescope, physical principles, zenithal dependency

*Limits of Muon tomography applied to archaeological survey
(tumulus example)*

Muography imaging

Basic principles

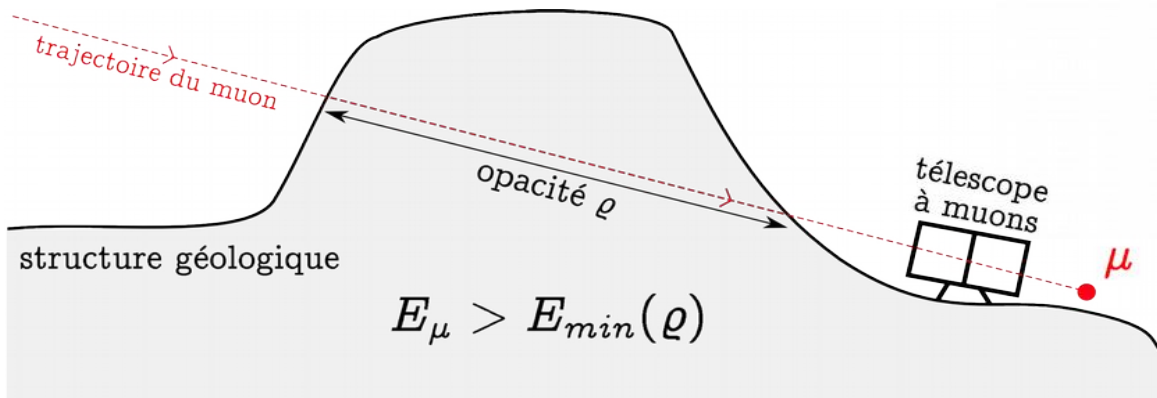
The particles energy loss along their trajectories is controlled by the amount of material encountered
⇒ " clinical radiography " by density contrast



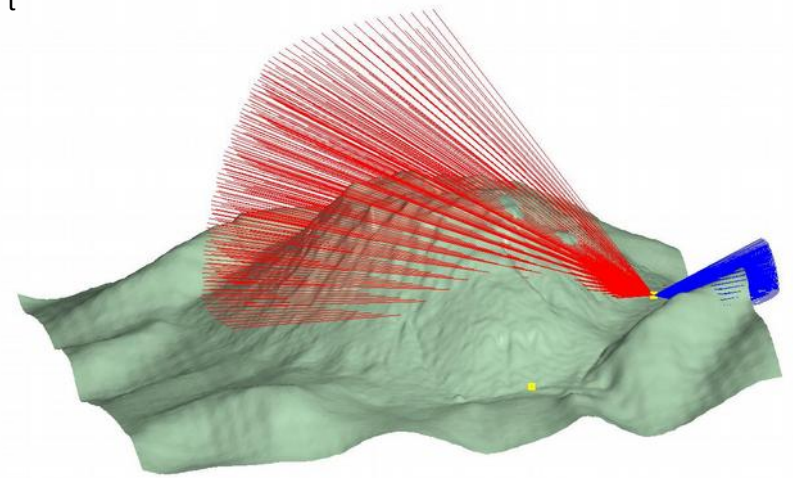
Physical principles

In practice, 2D muography is sensible to the **opacity** ϱ_t

$$\varrho_t = \int_t \rho(\xi) \times d\xi = L \times \bar{\rho}$$

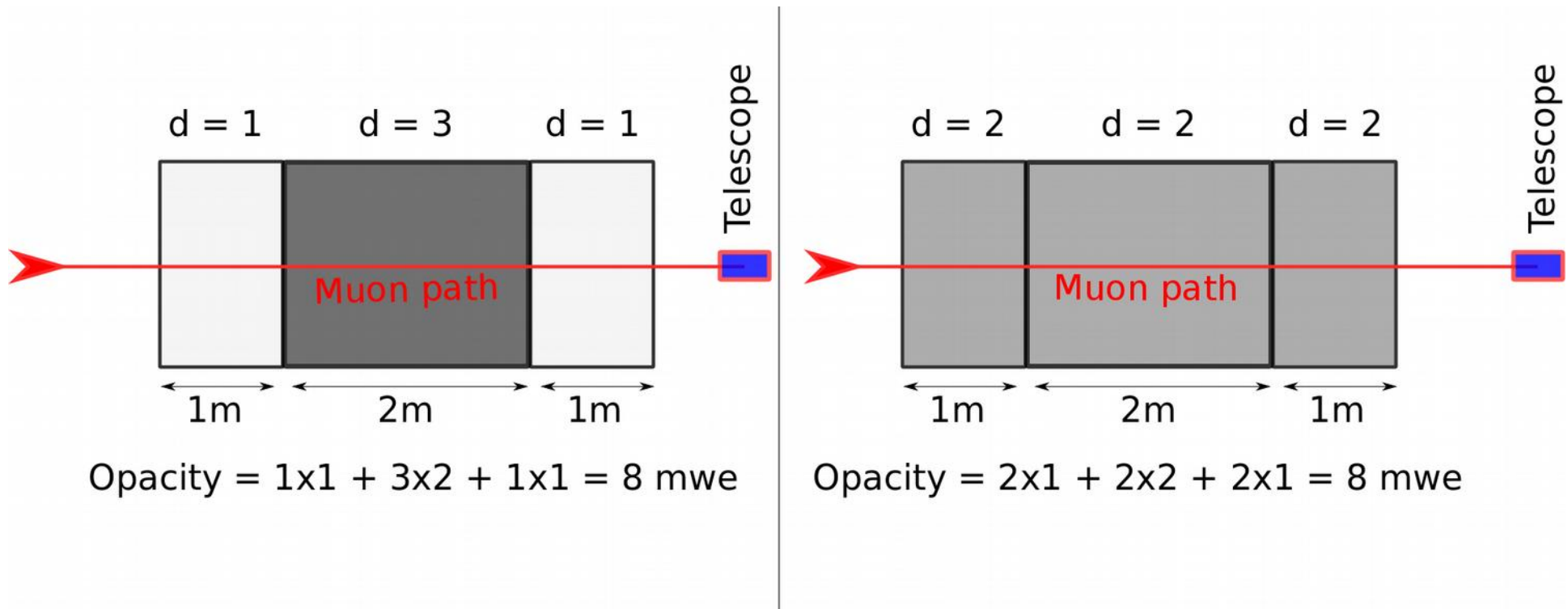


*George (1955), Alvarez et al. (1970),
Nagamine et al. (1995)*



=> strong **equivalences** exist in 2-D muography

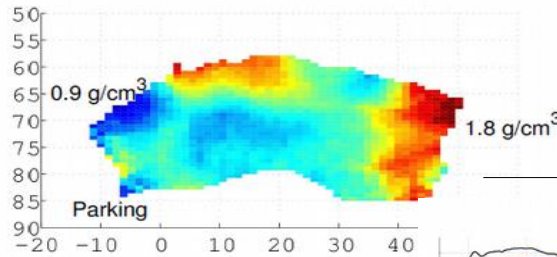
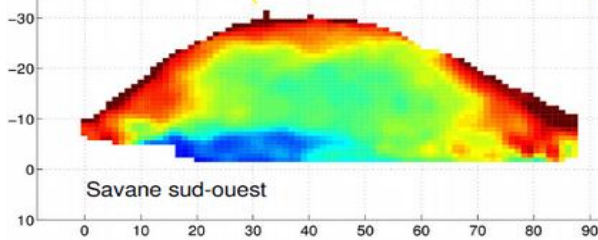
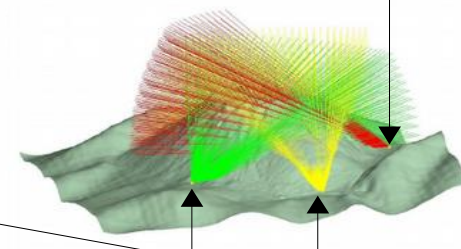
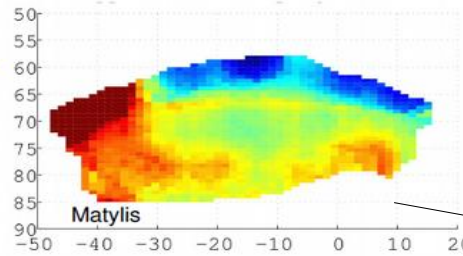
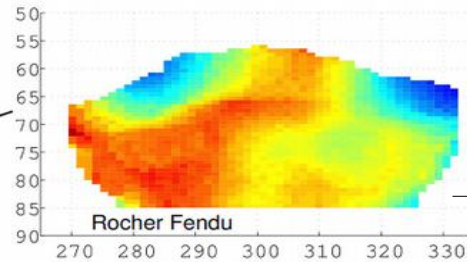
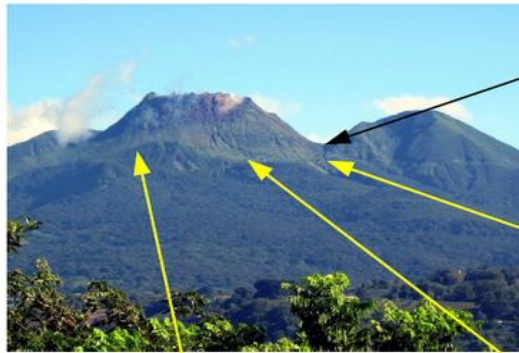
Equivalences example (1st order)



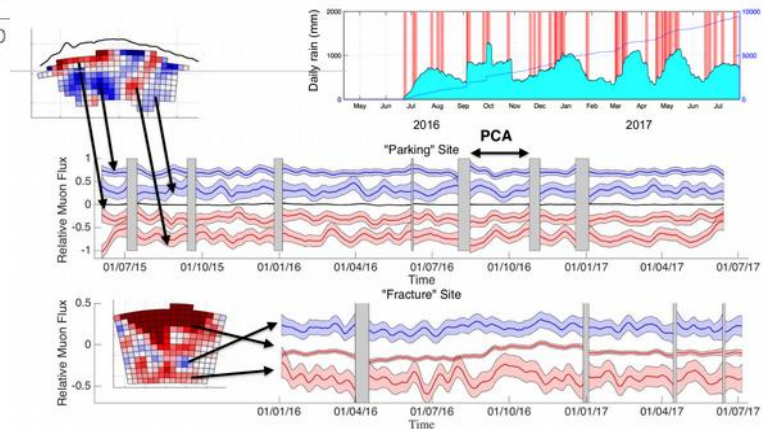
Equivalent !

Results: structural & temporal imagery

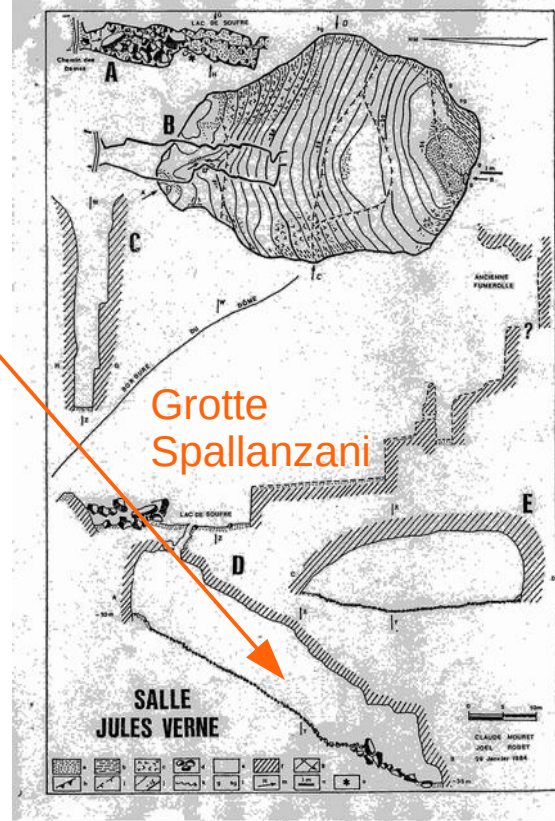
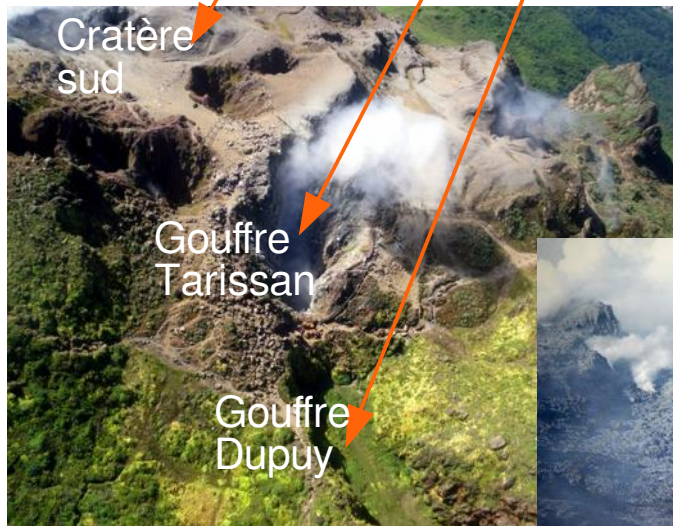
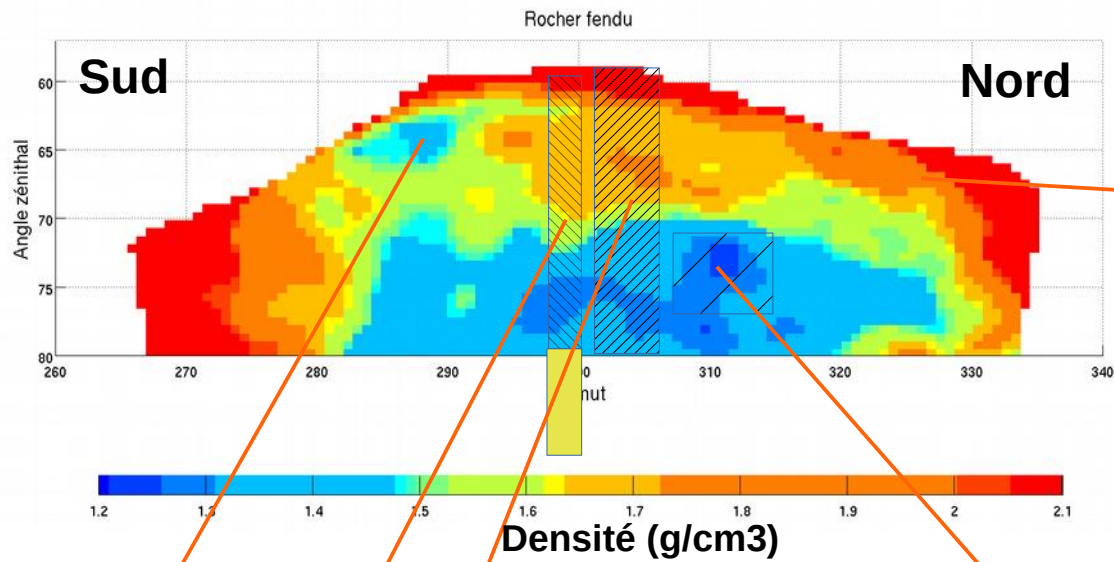
Radiographies de la Soufrière



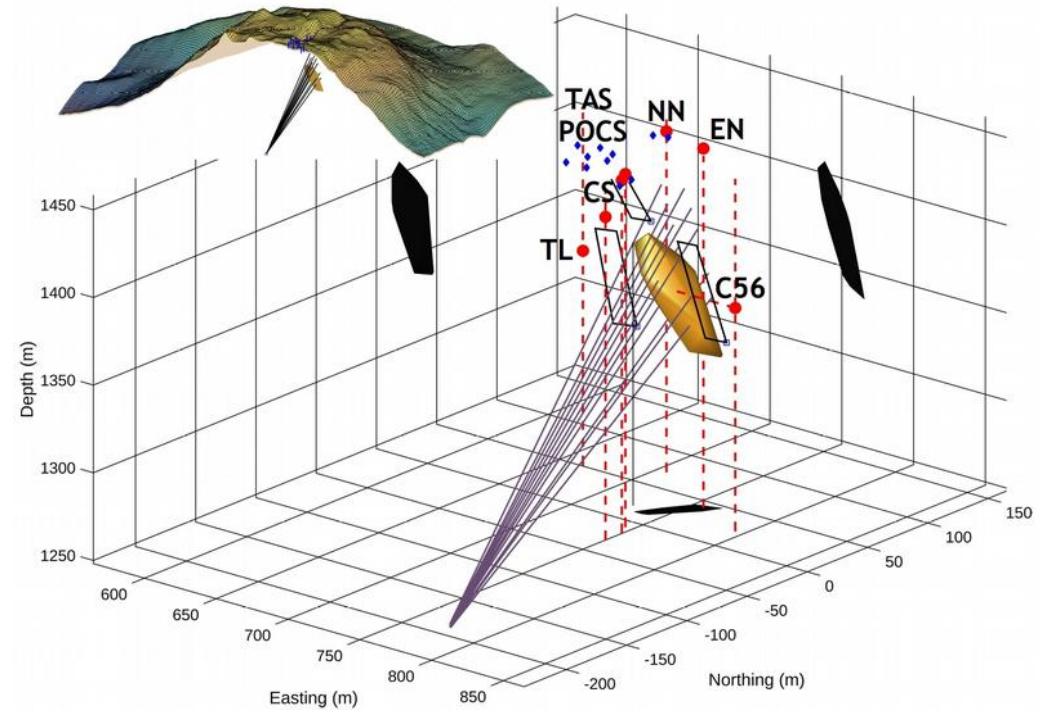
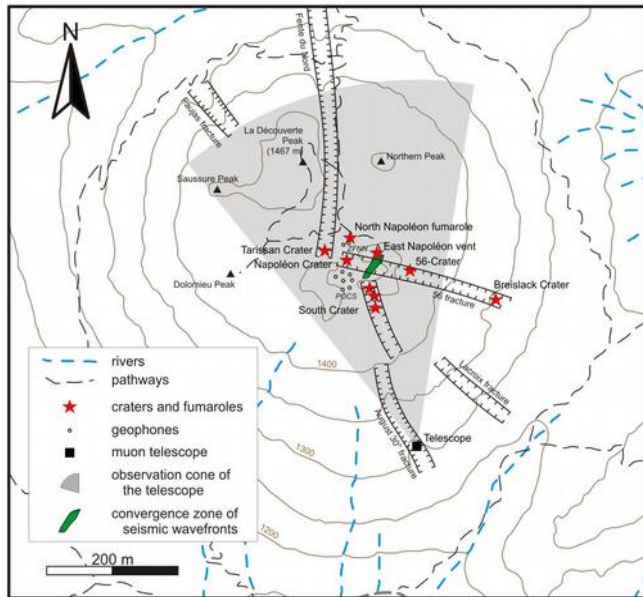
Lesparre, N. et al. 2012
Jourde, K. et al. 2013



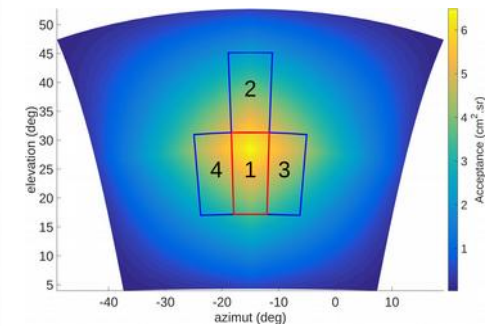
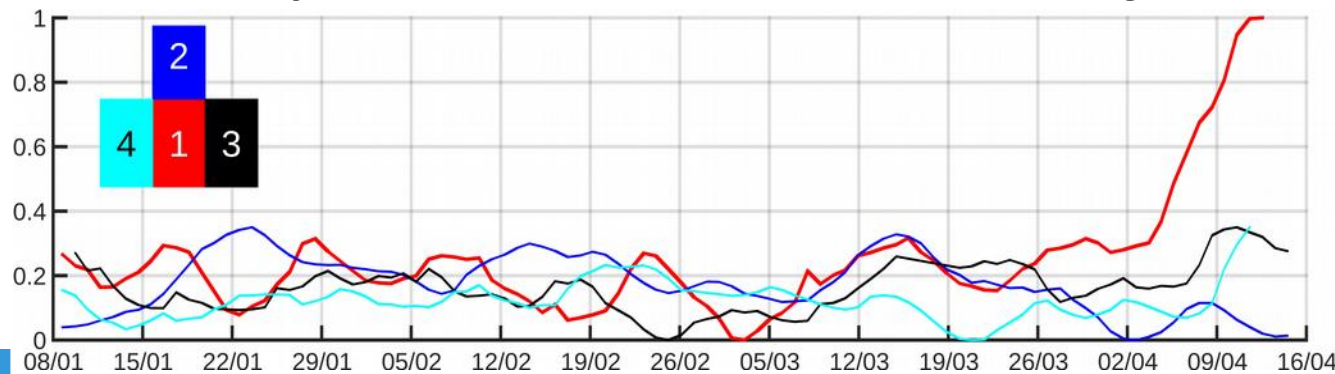
Results: validation with observations



Joint monitoring: seismic / muons

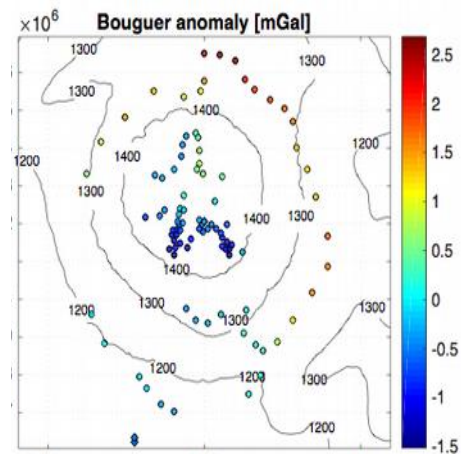


Global analysis of muon and seismic monitoring

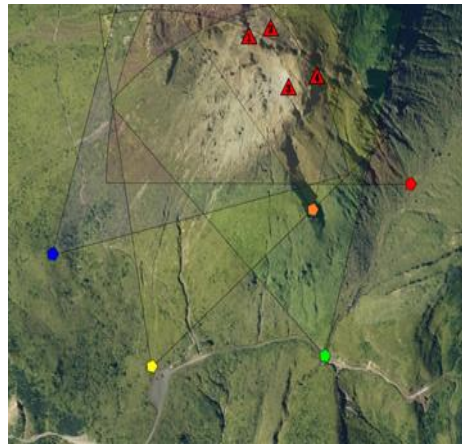


Joint monitoring: gravimetry / muons

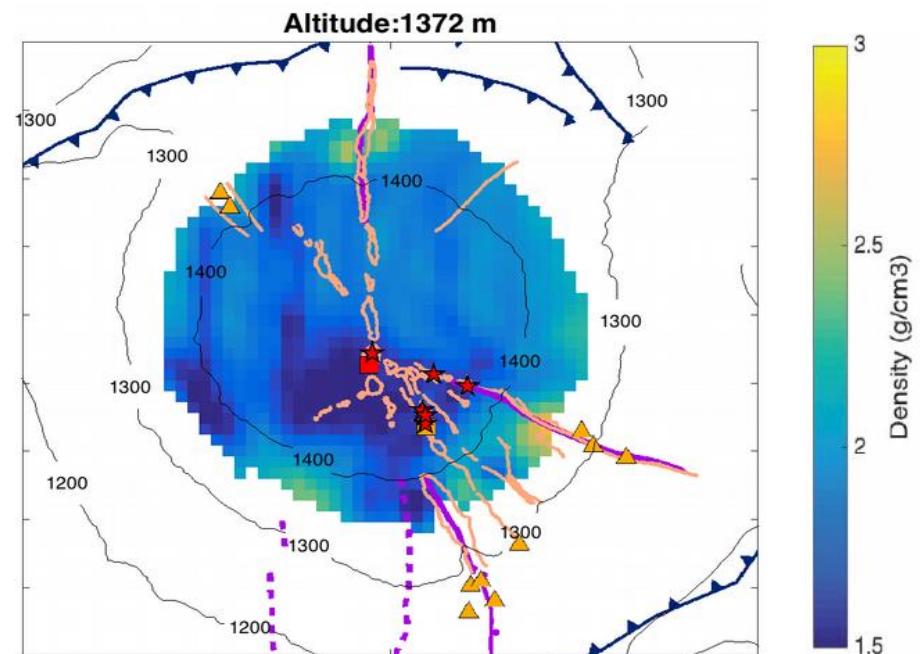
Gravimetry



Muons telescopes



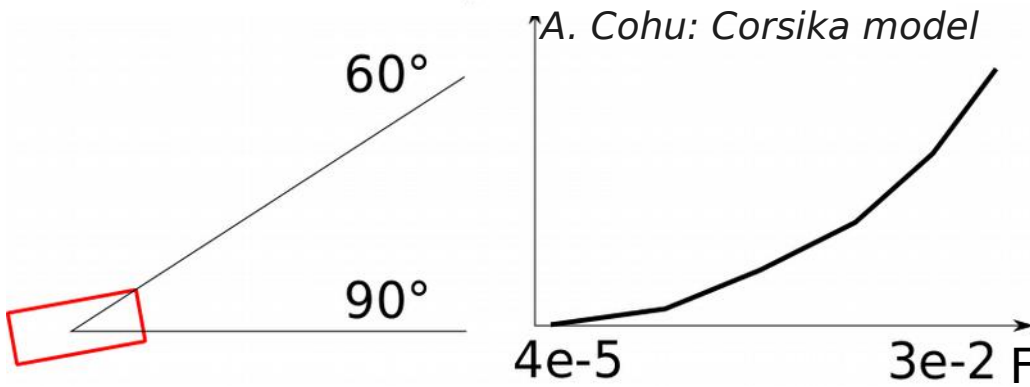
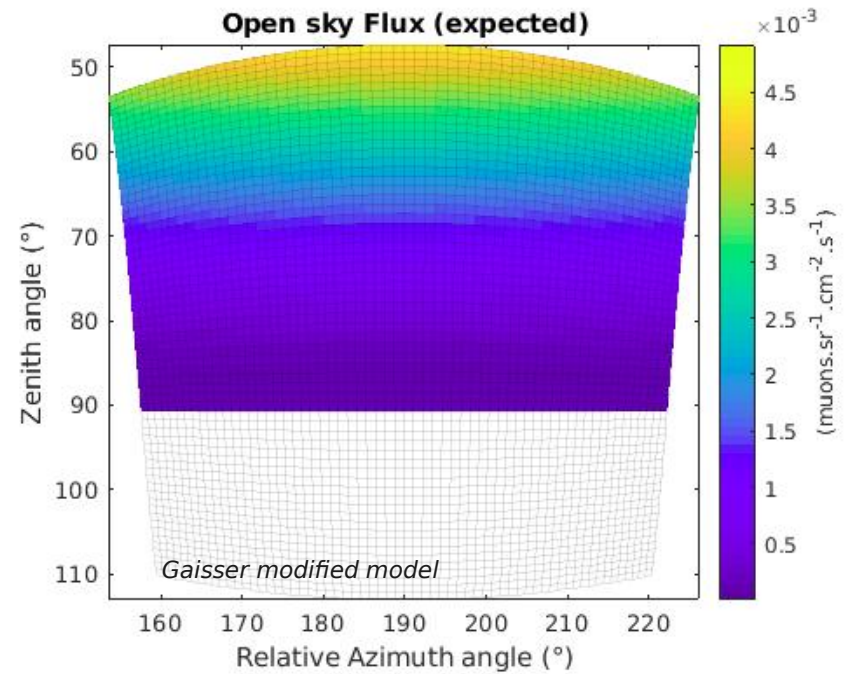
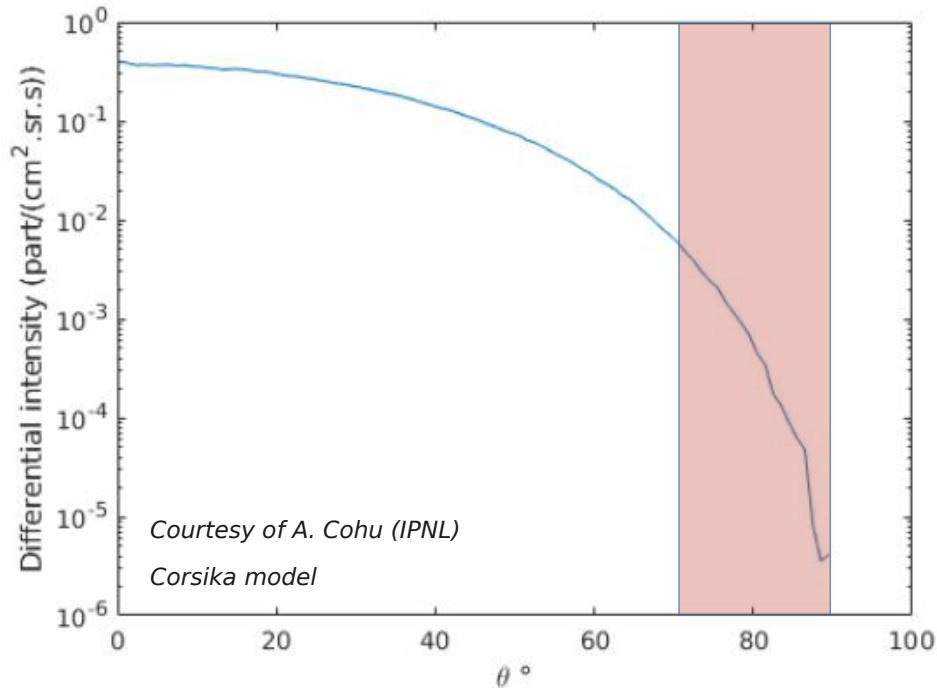
Joint inversion



$$G \begin{bmatrix} \rho_{\mu} \\ \Delta\rho \end{bmatrix} = \begin{bmatrix} G_g \\ G_{\mu} \end{bmatrix} \begin{bmatrix} \rho_{\mu} \\ \Delta\rho \end{bmatrix} = \begin{bmatrix} d_g \\ d_{\mu} \end{bmatrix} = d$$

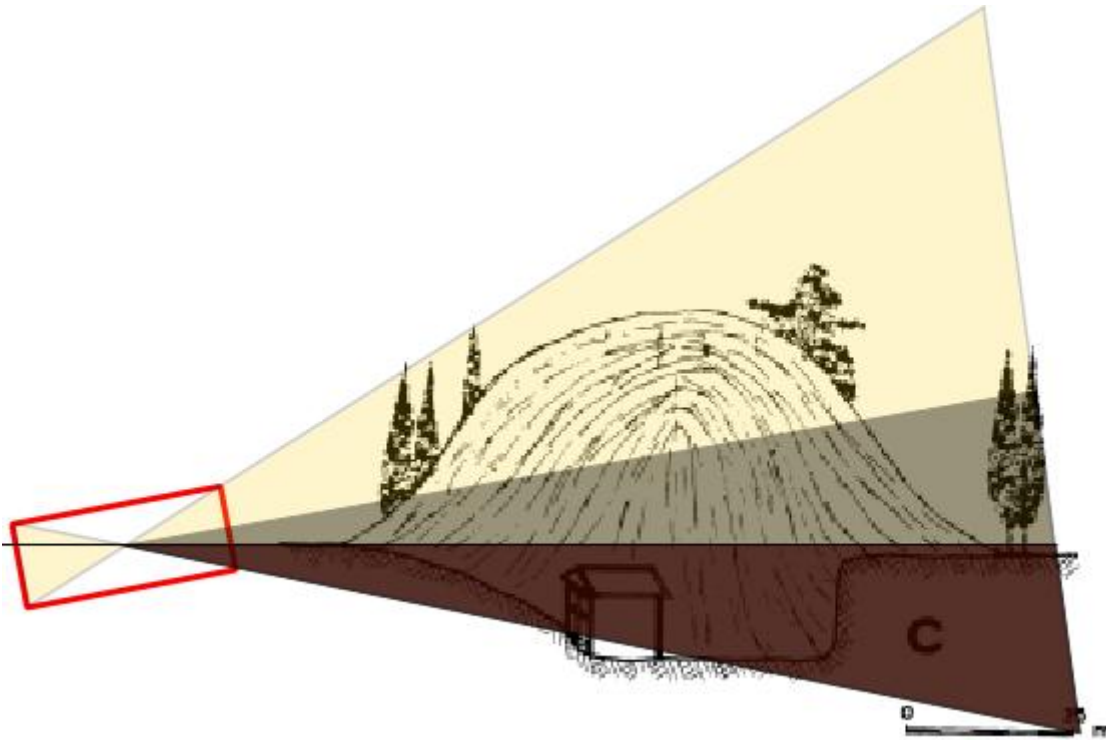
Sensible to the same parameters (density/mass)

Zenithal dependency (limitation)



The closest to 90° the target
The more challenging it is.
Lower Flux (less muons)
Lower Flux estimate precision

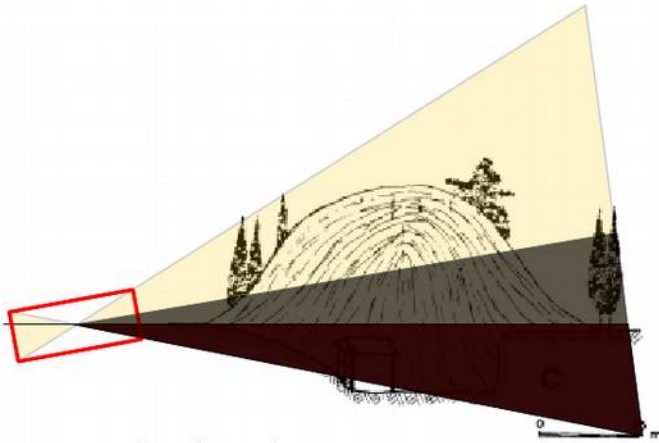
Limits of Muon tomography applied to archaeological survey (e.g. tumulus)



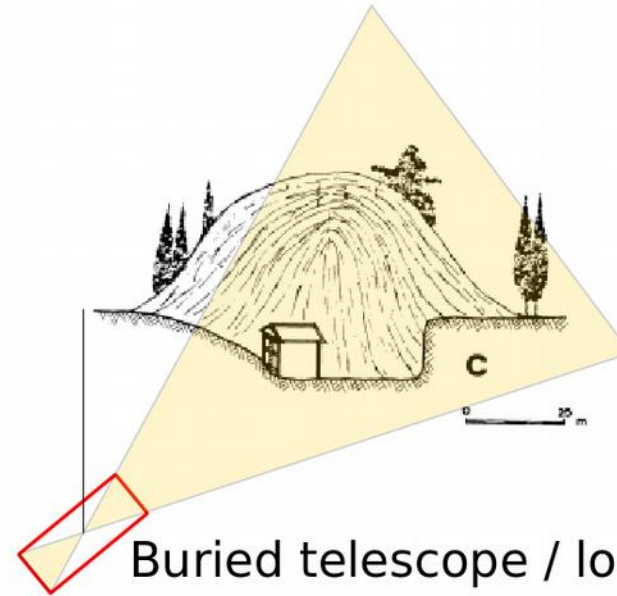
Modified after G. Tsokas et al. (Geophysics, 1995.)

- **Efficiency zone (0-70°)**
- **Challenging zone (70-90°)**
 - => **Weak muon flux (time)**
 - => **Theoretical flux estimate ~**
 - => **Other targets (Mountains...)**
 - => **Scattering (low energy)**
- **Dead angle (Zenith = 90° & below)**

Limits of Muon tomography applied to archeological survey (tumulus example)



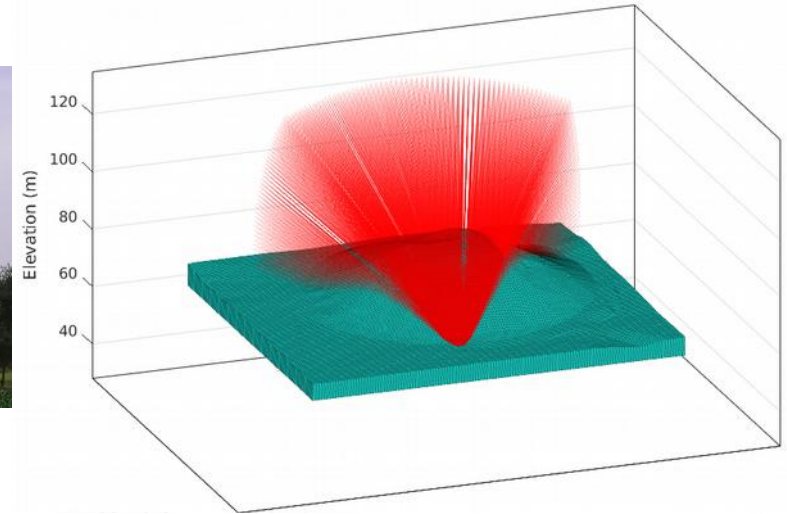
Grounded telescope
=> challenging angles



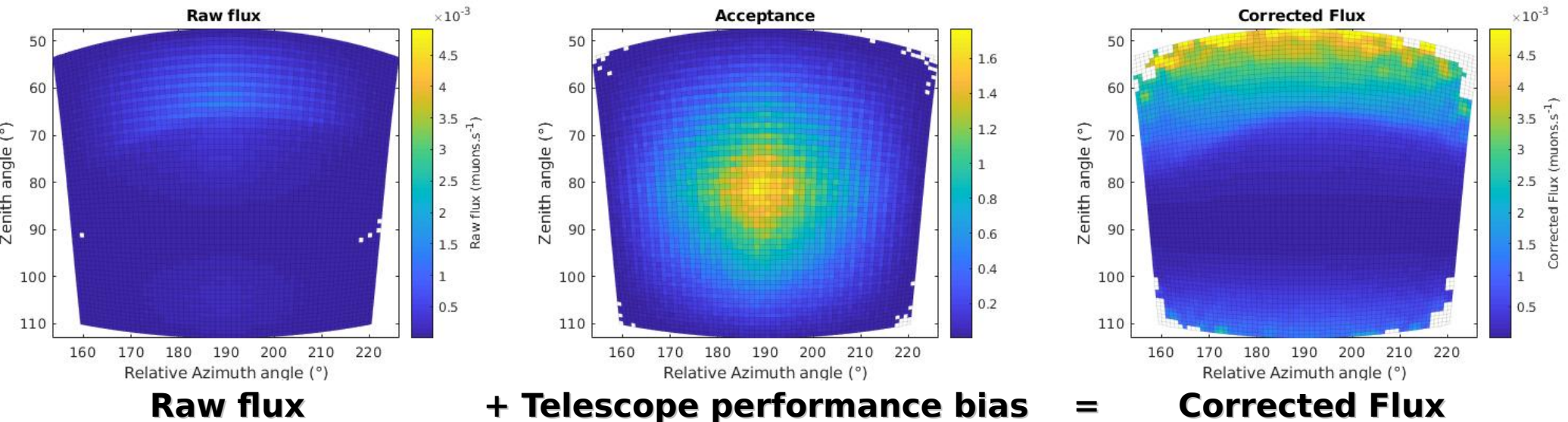
Buried telescope / local depression
=> no more challenging angles
=> telescope security

Results of muon tomography in Apollonia

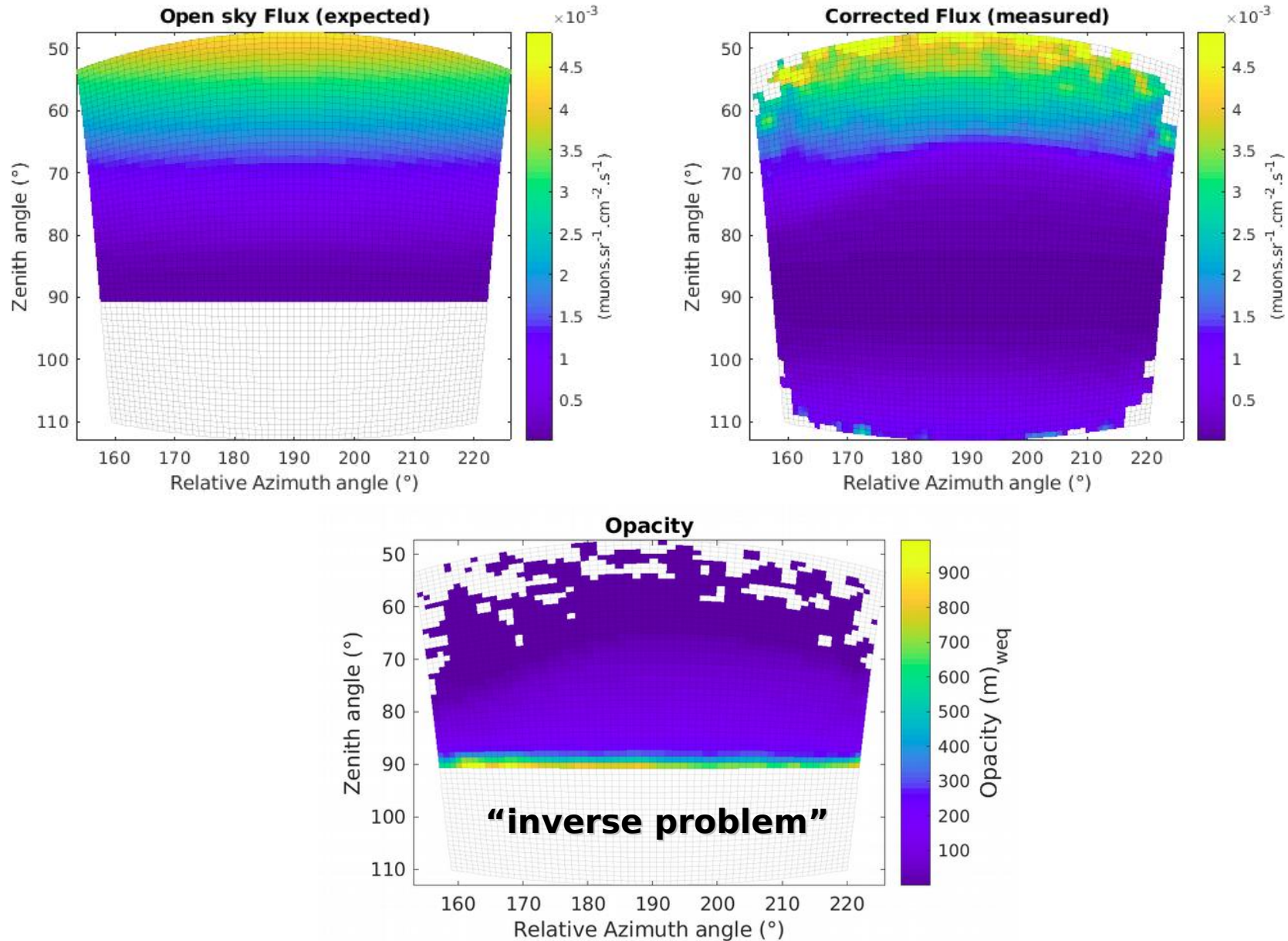
Data processing 1/3: Raw flux to flux



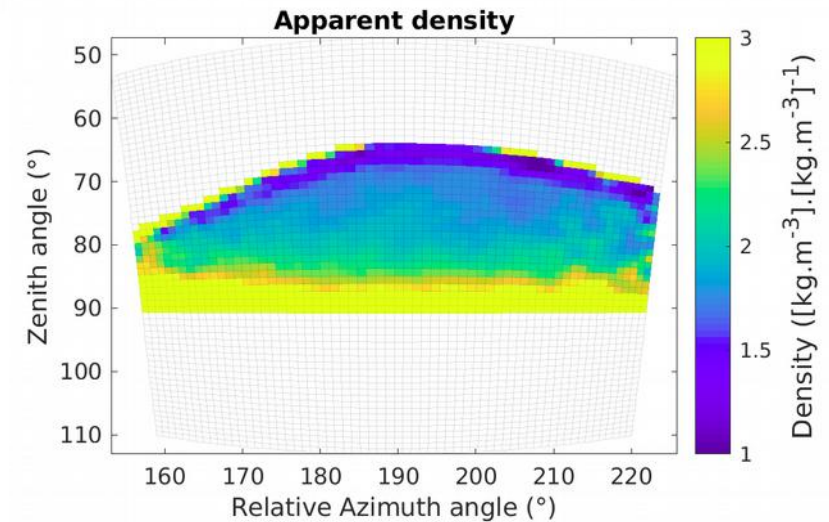
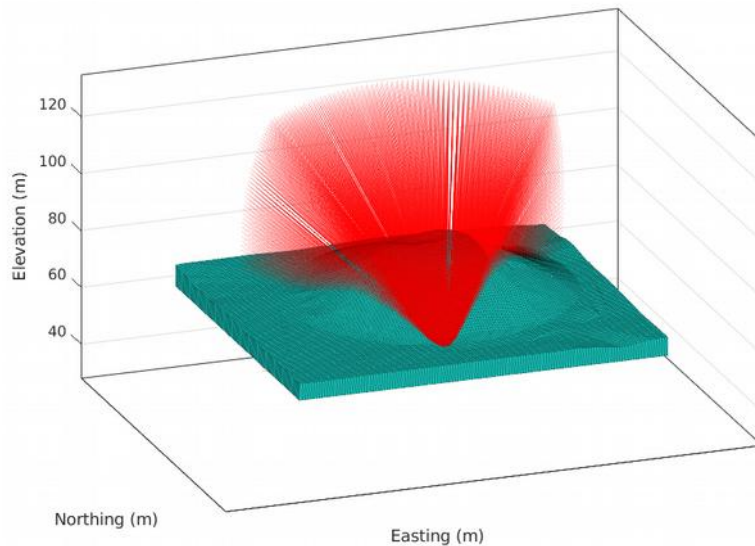
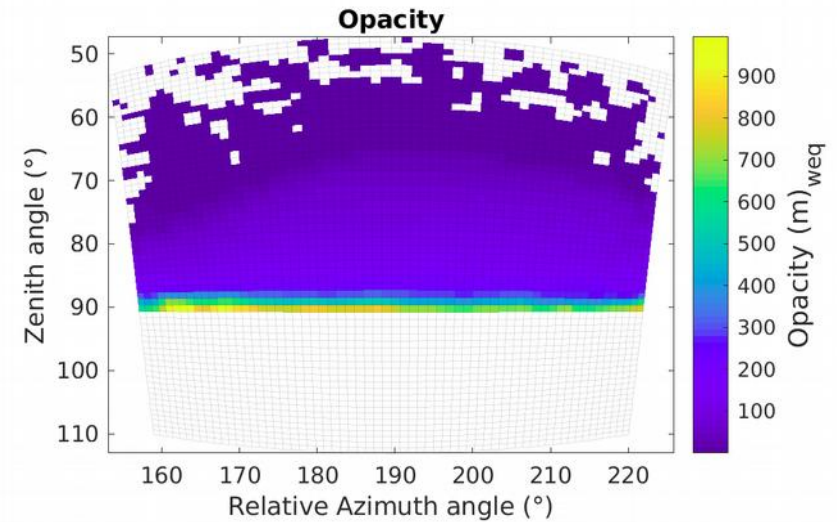
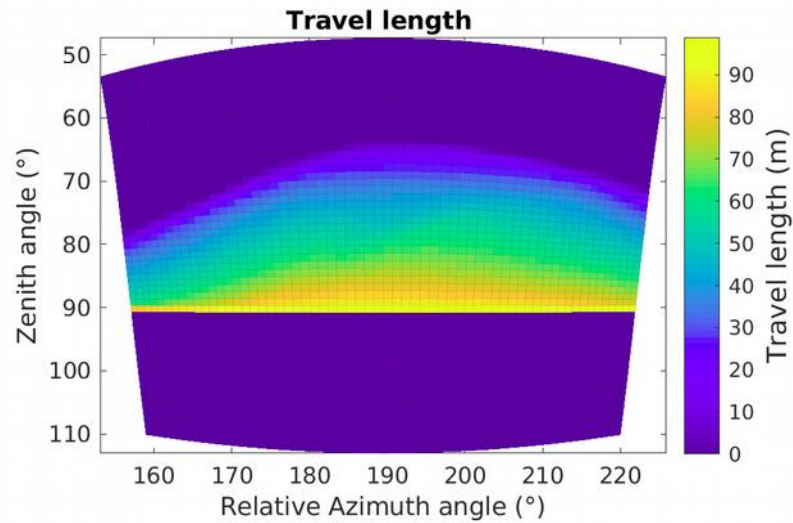
- 3 months data



Data processing 2/3: flux to opacity

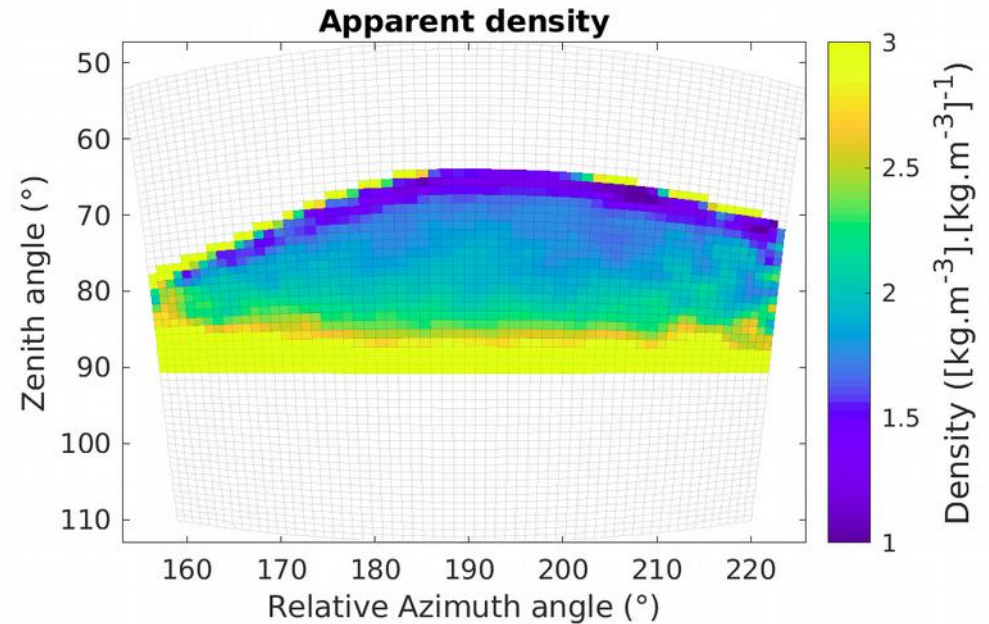


Data processing 3/3: Opacity to density



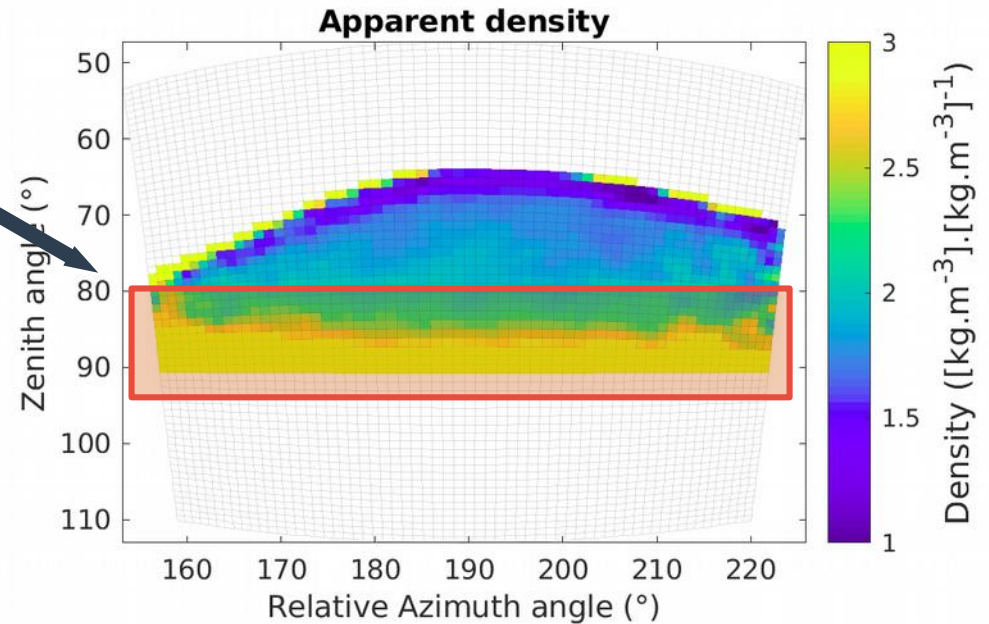
Understanding the Apparent density

- The density is “apparent”
- The structure is 3D, yet the measurement is 2D
- The “apparent” conveys the information that it is a density “averaged” over the path of a muon in a structure



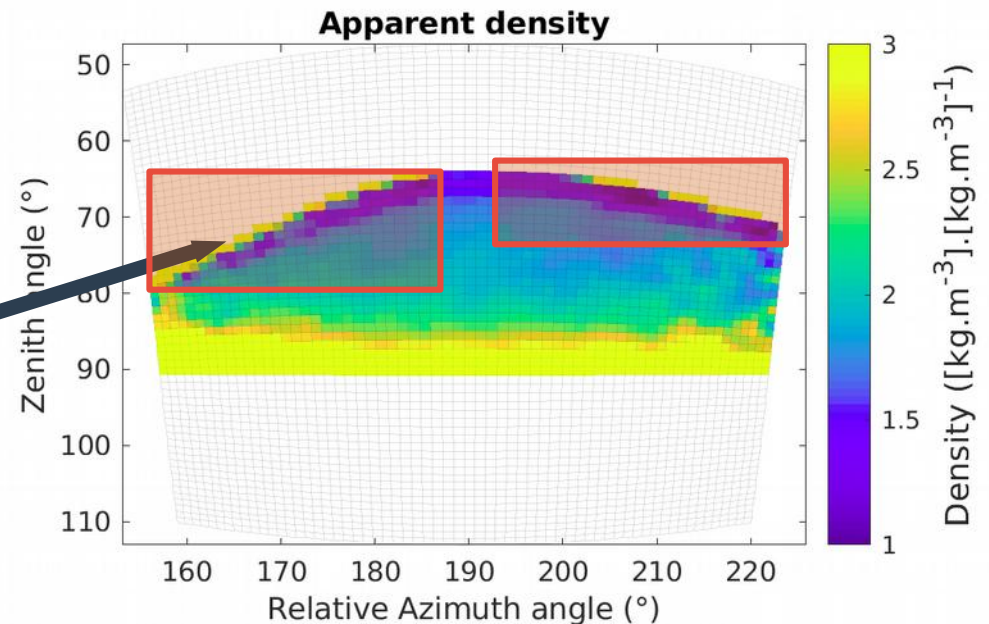
Issues & pre-analysis

- **Bottom ? (over-estimate)**
=> **Open-sky flux estimate**
- Radial edges (under + over-estimate)
=> Open-sky flux estimate
=> TL slightly off (unlikely)
- Inner-lateral contrasts
=> unless TL slightly off (unlikely)
=> not related to open-sky flux
=> confident



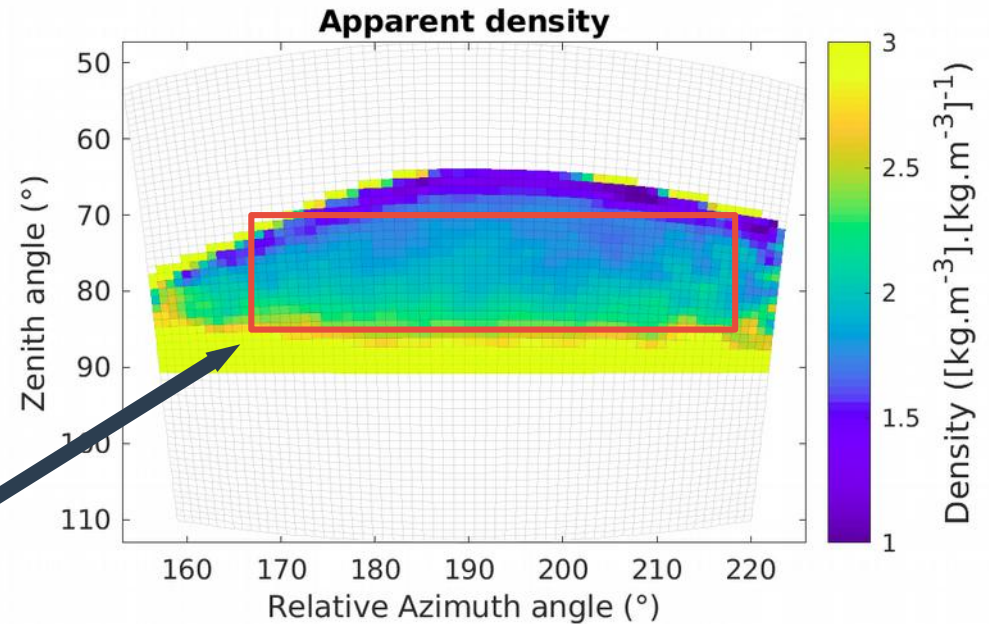
Issues & pre-analysis

- Bottom (over-estimate)
=> Open-sky flux estimate
- Radial edges (under + over-estimate) ?
=> Open-sky flux estimate
=> TL slightly off (unlikely)
- Inner-lateral contrasts
=> unless TL slightly off (unlikely)
=> not related to open-sky flux
=> confident

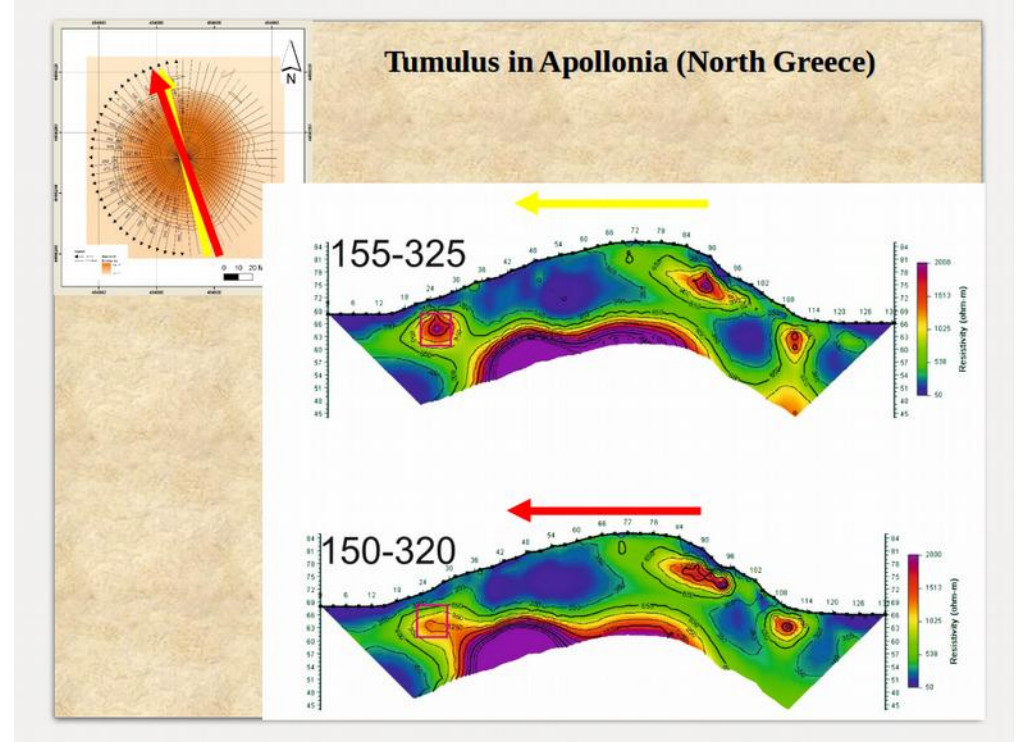
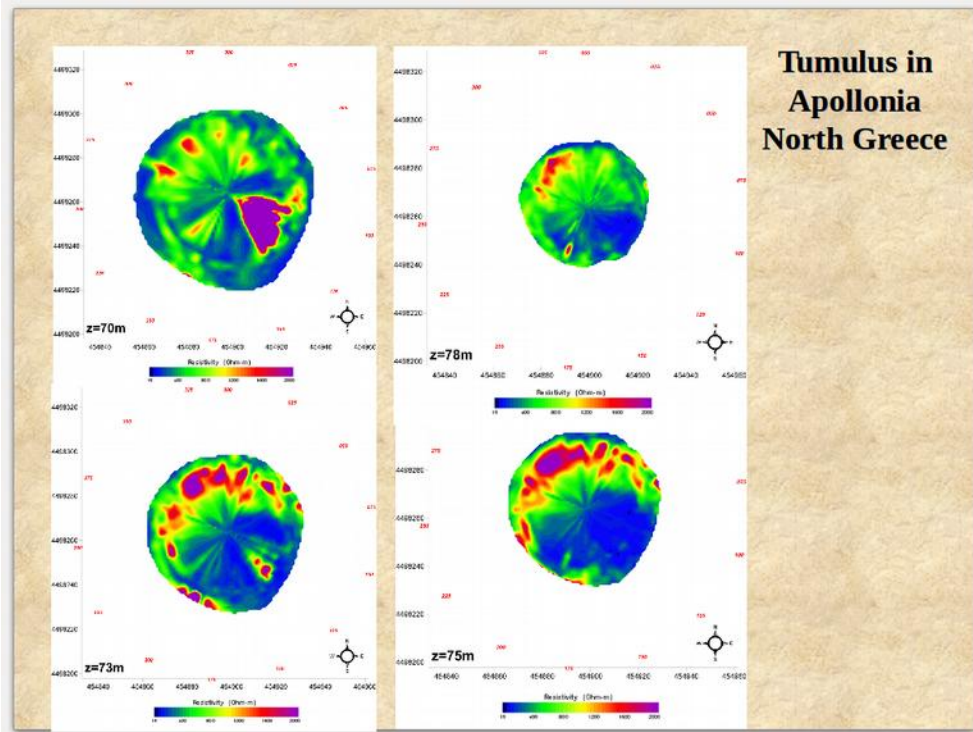


Issues & pre-analysis

- Bottom (over-estimate)
=> Open-sky flux estimate
- Radial edges (under + over-estimate)
=> Open-sky flux estimate
=> TL slightly off (unlikely)
- Inner-lateral contrasts ?
=> unless TL slightly off (unlikely)
=> not related to open-sky flux
=> confident

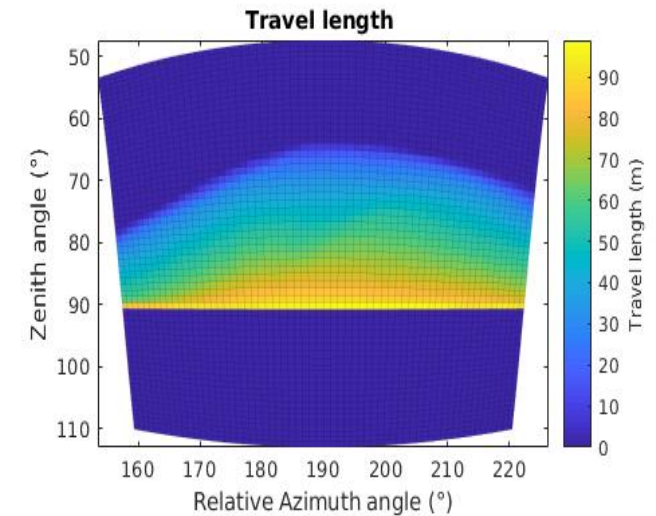
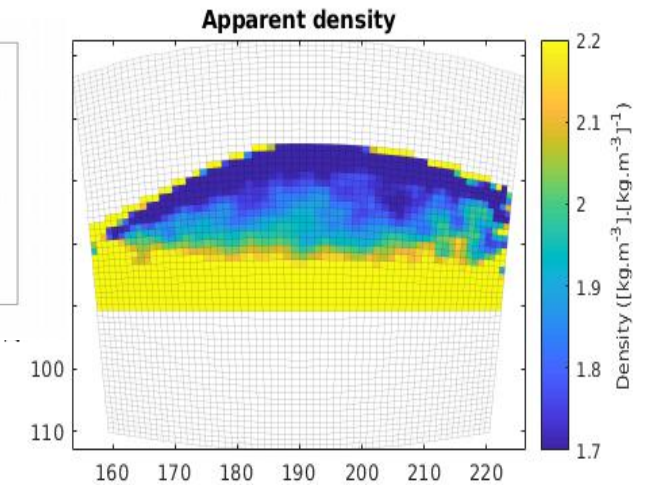
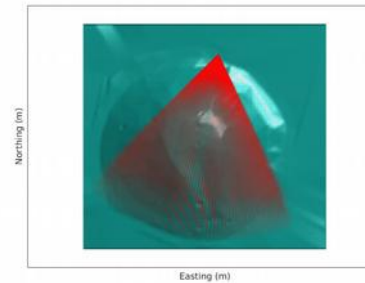
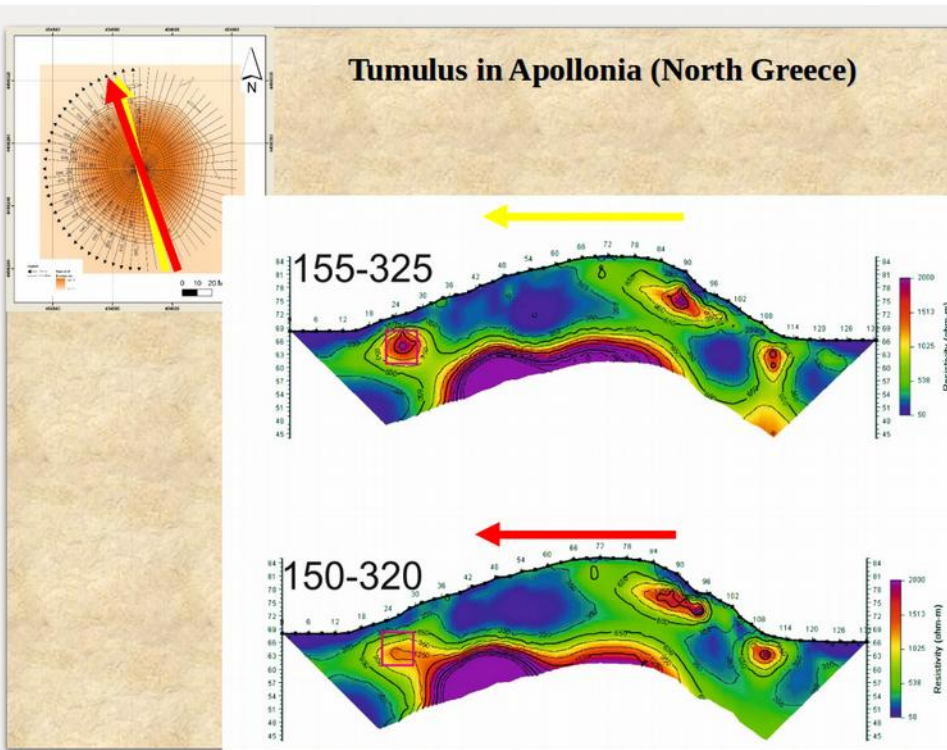


Electric tomography



G. Tsokas et al. (2015)

Electrical & Apparent density



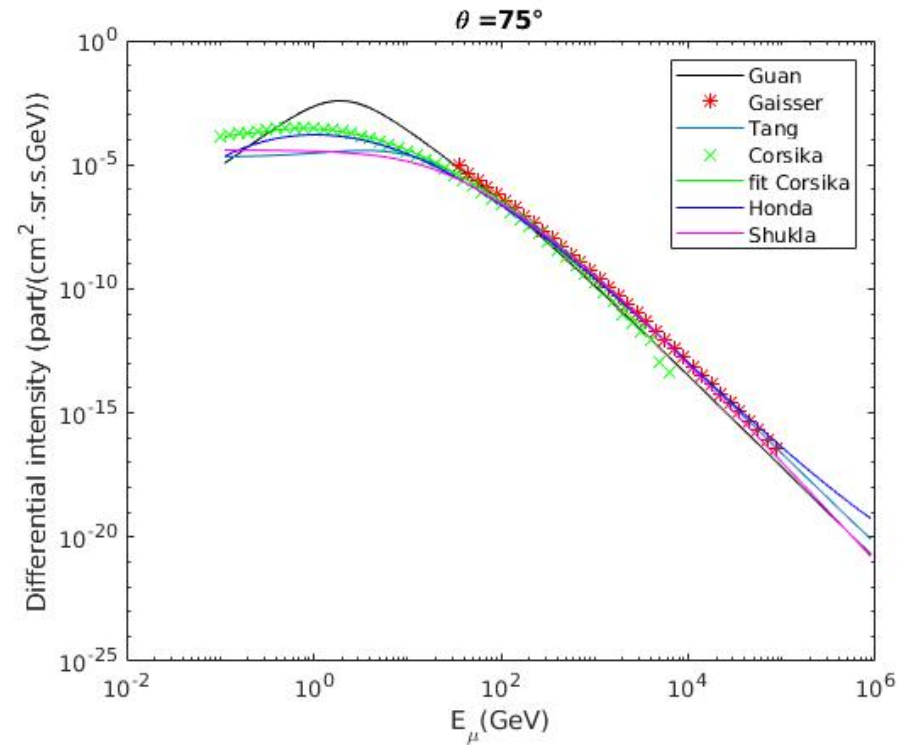
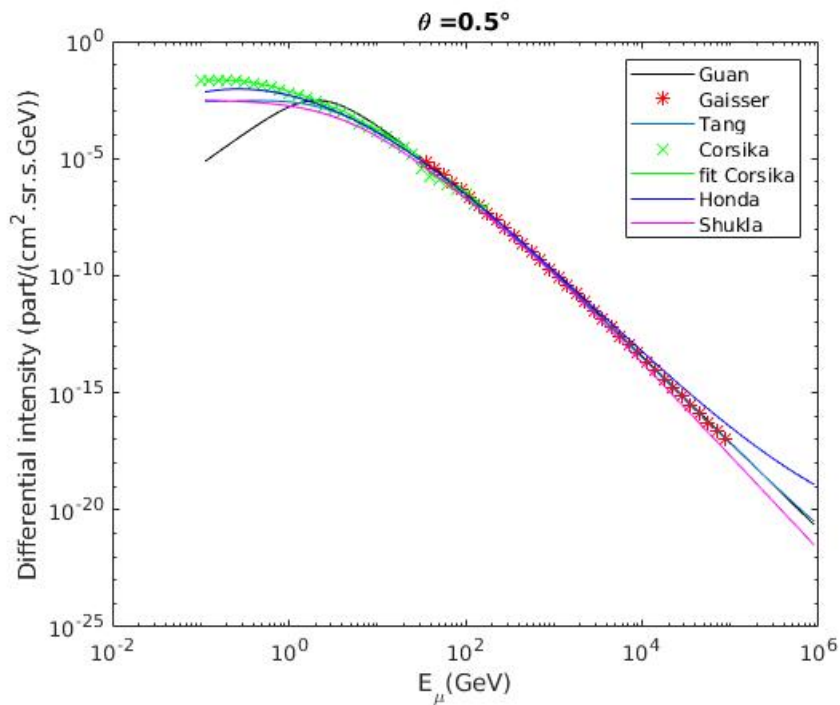
- **Topography variations seem correlated to density variation => Suspicious (for now)**

Extending the actual limitations

Dealing with the low-angle flux issue

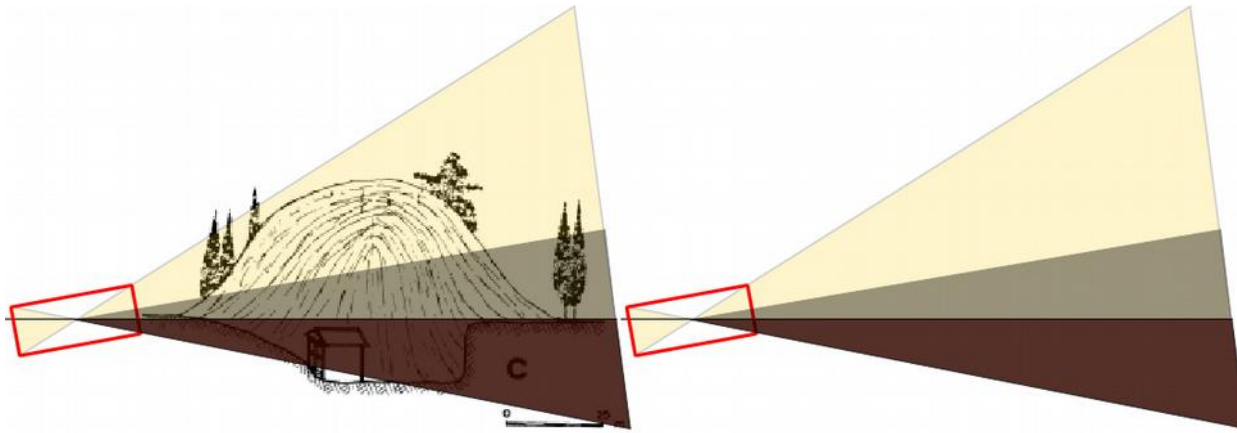
2D limits

For this project: improving the flux estimate



- **Less muons at low angle than actually expected**

In the future: dealing with the flux issue



Survey telescope

Telescope dedicated to raw flux measurements

- => Deals with :
- Meteorological variations
- Far away structures
- Theoretical flux

- Require a common open-sky calibration
- => joined 3 panels telescope



Conclusions

On going project ...

Thank you for your attention...