



The APC Neutrino Group

Conseil Scientifique APC
March 2020

The Neutrino Group

| 11 permanents + 2 volunteers | Position | HDR |
|---------------------------------|----------|-----|
| M. Cribier | Emeritus | |
| J. Dawson | CR | |
| D. Franco | CR | X |
| A. Kouchner | PR | X |
| D. Kryn | benevole | |
| T. Lasserre | CEA | X |
| M. Obolensky | benevole | |
| T. Patzak | PR | X |
| S. Sacerdoti | CR | |
| A. Tonazzo | PR | X |
| V. Van Elewyck | MCF | |
| F. Vannucci | Emeritus | |
| D. Vignaud | Emeritus | |

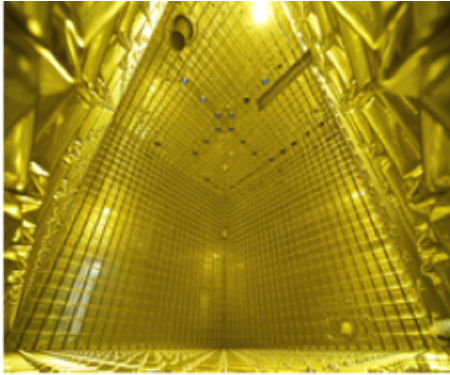
| 2 Postdocs | Experiment | Period | Grant |
|----------------|------------|-----------|-------|
| R. Le Breton | ORCA | 2017/2020 | ANR |
| G. De Gwenhael | ORCA | 2019/2021 | LabEx |

| 9 PhD | Experiment | Period | Cotutelle |
|---------------|------------|-----------|------------|
| R. Bajou | DUNE | 2019/2022 | |
| E. Chardonnet | DUNE | 2018/2021 | |
| Y. Han | JUNO | 2017/2020 | |
| T. Hugues | DarkSide | 2019/2023 | Poland (*) |
| M. Lai | DarkSide | 2017/2020 | Italy |
| L. Maderer | ORCA | 2019/2022 | |
| C. Thien Nhan | ORCA | 2018/2021 | |
| J. Rode | DarkSide | 2019/2022 | |
| F. Versari | ORCA | 2017/2020 | Italy |

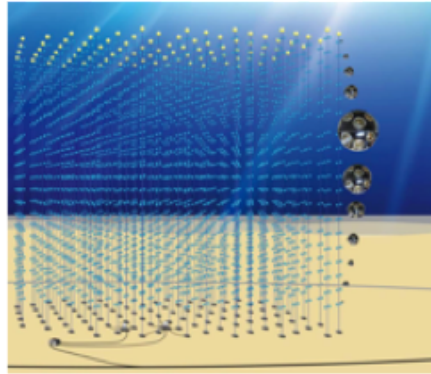
(*) in partnership with Astrocent

Evolution since 2017

DUNE/WA105



ORCA



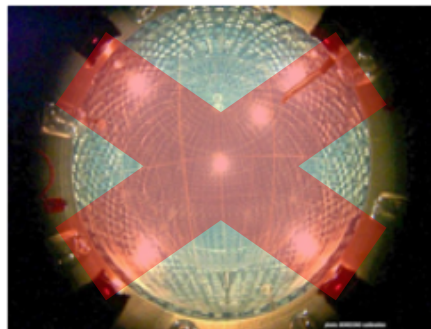
DarkSide



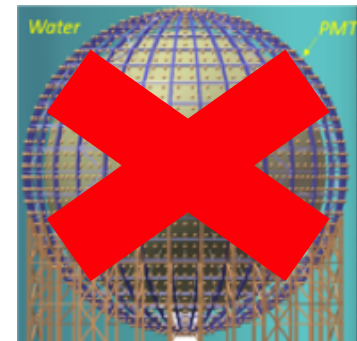
Double Chooz



Borexino/SOX



JUNO



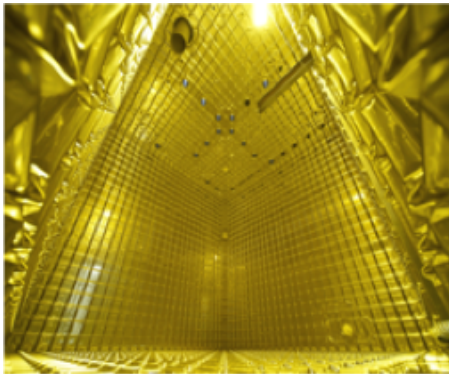
LIQUIDO



ARIS

Neutrino Group Experiments: 2019

DUNE/WA105



Oscillation physics

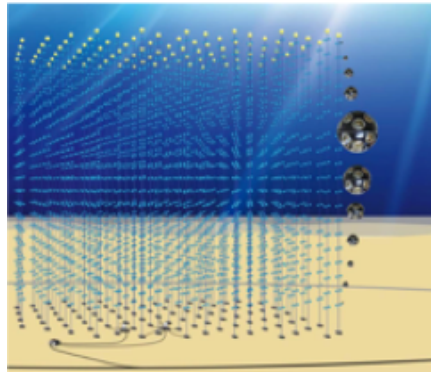
- Search for leptonic CP violation
- Determine the neutrino mass hierarchy
- Precision PMNS measurements

Supernova physics

Solar neutrinos

Baryon number violation

ORCA



Oscillation physics

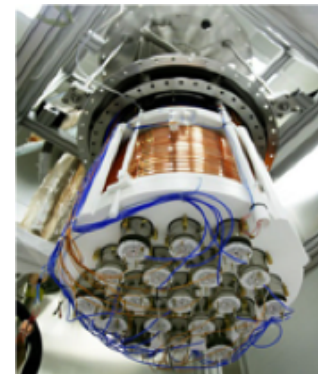
- Determine the neutrino mass hierarchy
- Precision PMNS measurements

Supernova physics

Neutrino astronomy (AHE)

Earth tomography

DarkSide



Direct Dark Matter Search:

- Low/high mass WIMPs
- Boosted dark matter
- Axions

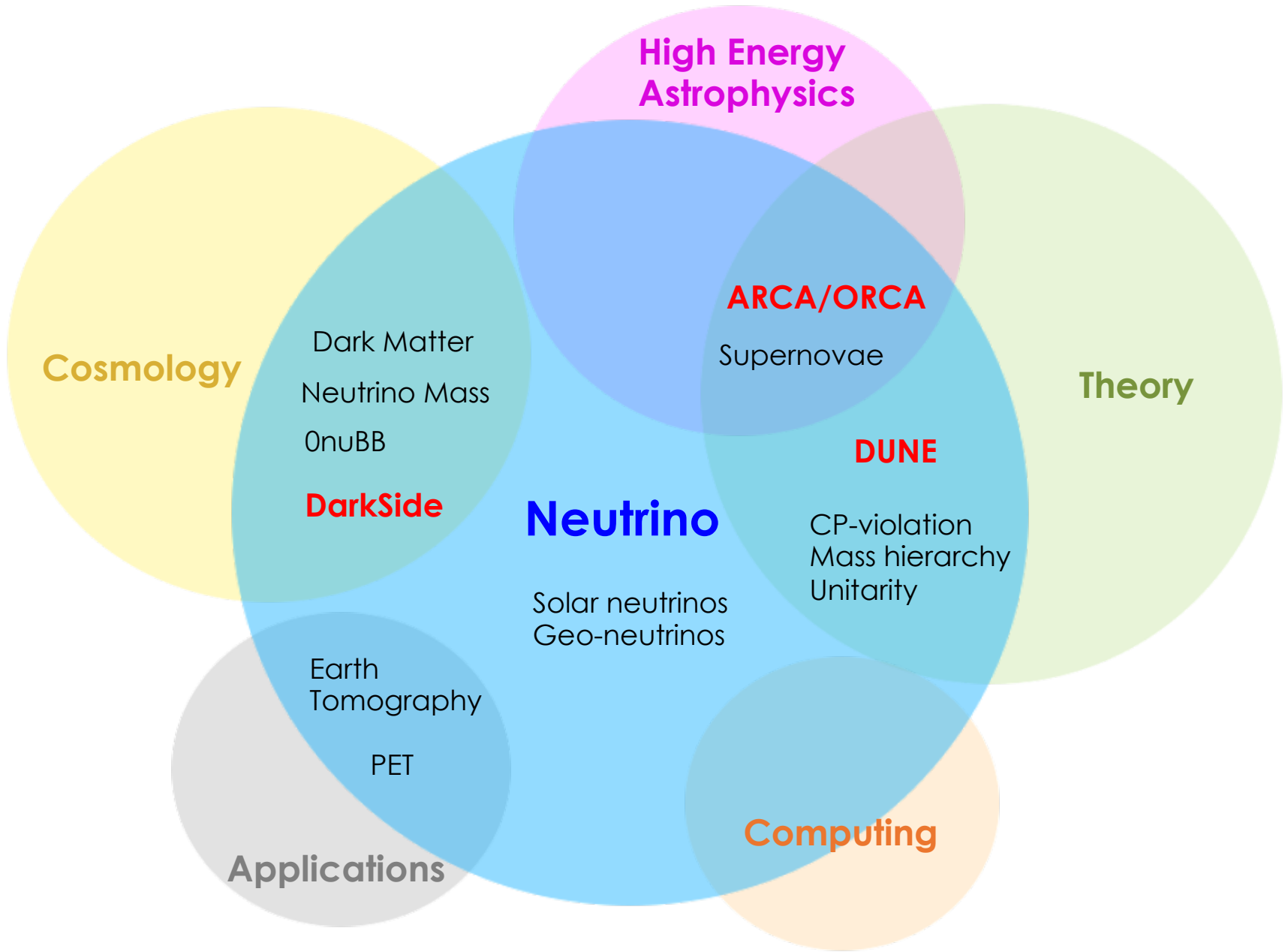
Supernova physics

Solar neutrinos

OnBB

Positron Emission Tomography

Integration in the APC universe



The Timeline

Design

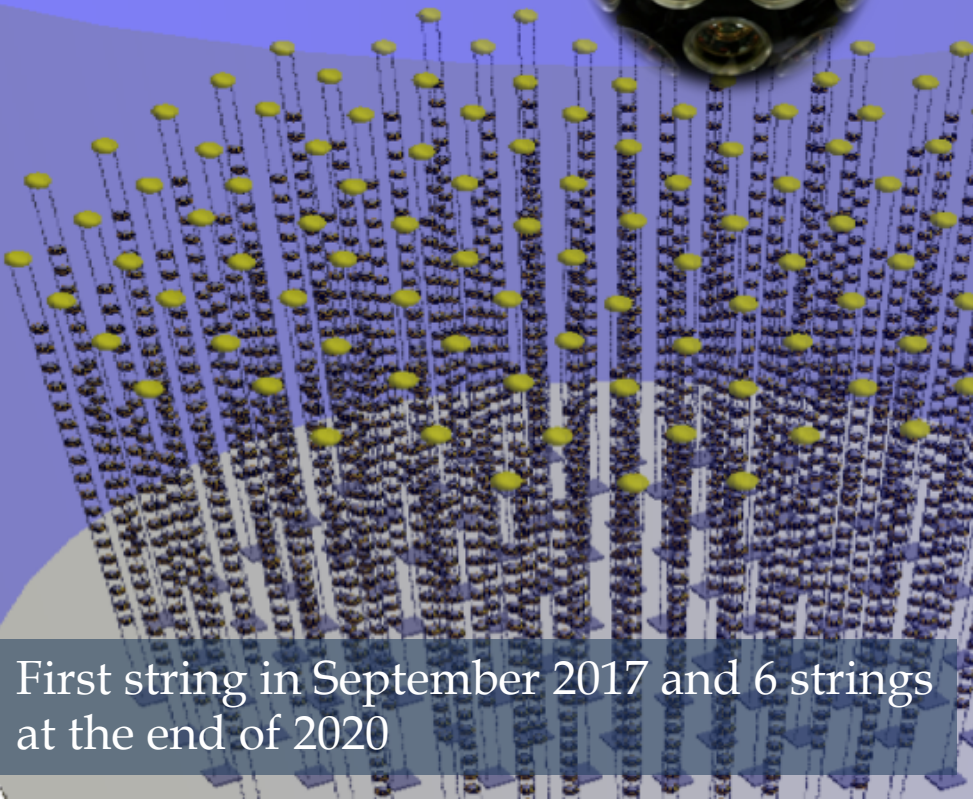
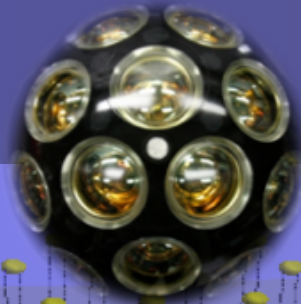
Construction

Data taking

| Projets | 2018 | 2019 | 2020 | 2021 | 2022 | 2023 | 2024 |
|-----------------|--------------|----------|----------|--------------|----------|----------|-------|
| DUNE | WA105 | WA105 | WA105 | WA105 | | | |
| | DUNE | DUNE | DUNE | DUNE | DUNE | DUNE | DUNE |
| KM3NeT | ANTARES | | | | | | |
| | ORCA Phase 1 | | | ORCA Phase 2 | | | |
| | | ORCA | ORCA | ORCA | ORCA | ORCA | ORCA |
| DarkSide | DS50 | DS50 | DS50 | | | | |
| | DS-PROTO | DS-PROTO | DS-PROTO | DS-PROTO | DS-PROTO | DS-PROTO | |
| | DS20k | DS20k | DS20k | DS20k | DS20k | DS20k | DS20k |

Some number:

- **5.7 Mt** instrumented volume
- **115** x 200-m long strings (DU)
- **18** DOMs / DU
- **31** PMTs / DOM
- **64k** x 3" PMTs



First string in September 2017 and 6 strings at the end of 2020

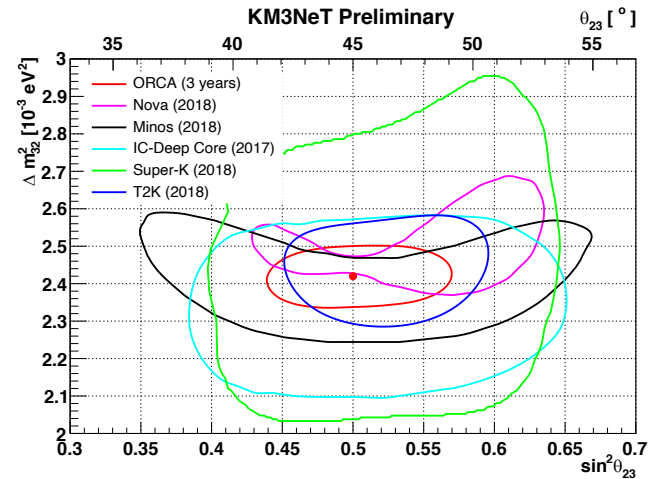
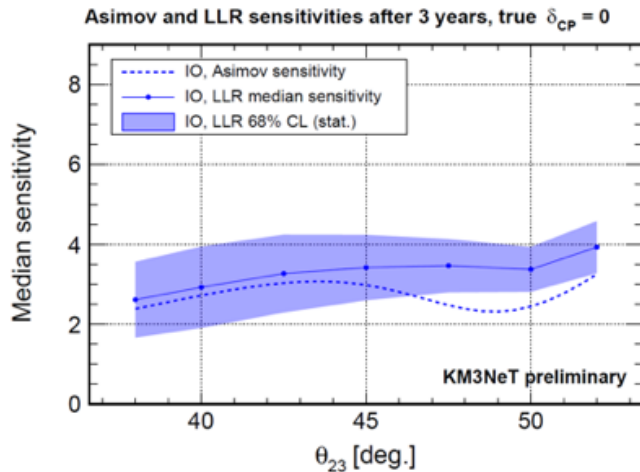
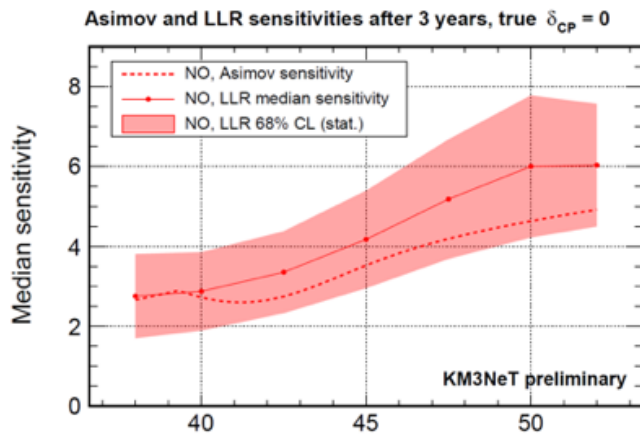
- M. Lindsey Clark **Technical Project Manager** since 2018
- **Neutrino Physics: MH + oscillation parameters** (N. Chau, Christine Nielsen)
- **Earth science: neutrino tomography** (V. Van Elewyck + E. Kaminski and J. Badro from IPGP)
- **Neutrino astronomy: Supernovae neutrinos** (M. Colomer) and **GeV neutrinos** (G. de Wasseige, V. Van Elewyck)
- **Detector monitoring: charge/time calibration** (A. Creusot)
- **Instrumentation: DOM characterization, Calibration Units, Quality Control** (A. Creusot, R. Le Breton, C. Boutonnet, C. Champion, S. Colonges, L. Confucius, A. Ilioni, M. Lindsey Clark)

ORCA: neutrino oscillations

oscillations of 1-100 GeV atmospheric neutrinos crossing the Earth

Neutrino Mass Hierarchy

Theta23



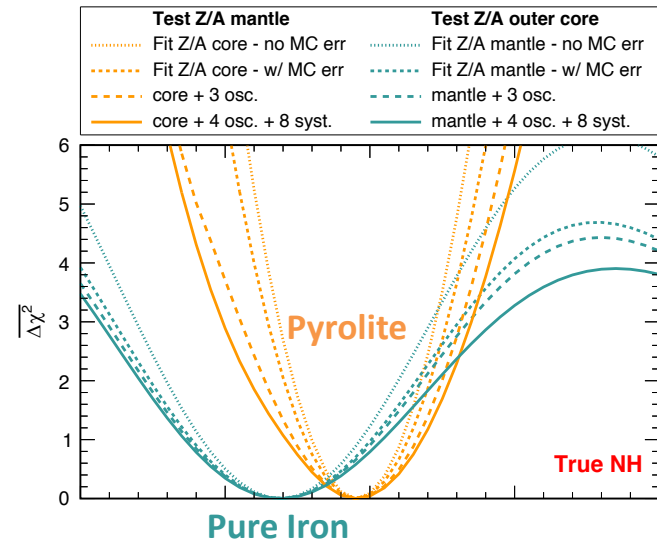
@APC

Development of a complete simulation chain including detector response => official ORCA sensitivity for NMH, oscillation parameters, + systematics studies

Neutrino oscillation tomography of the Earth

sensitive to the electron density => allows for few % accuracy on Earth core and mantle compositions with GeV neutrino oscillations

In collaboration with E. Kaminski, J. Badro (IPGP)

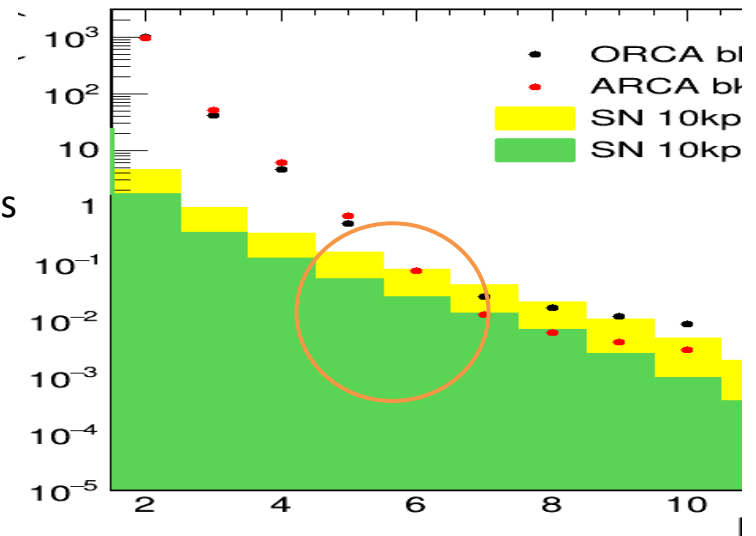


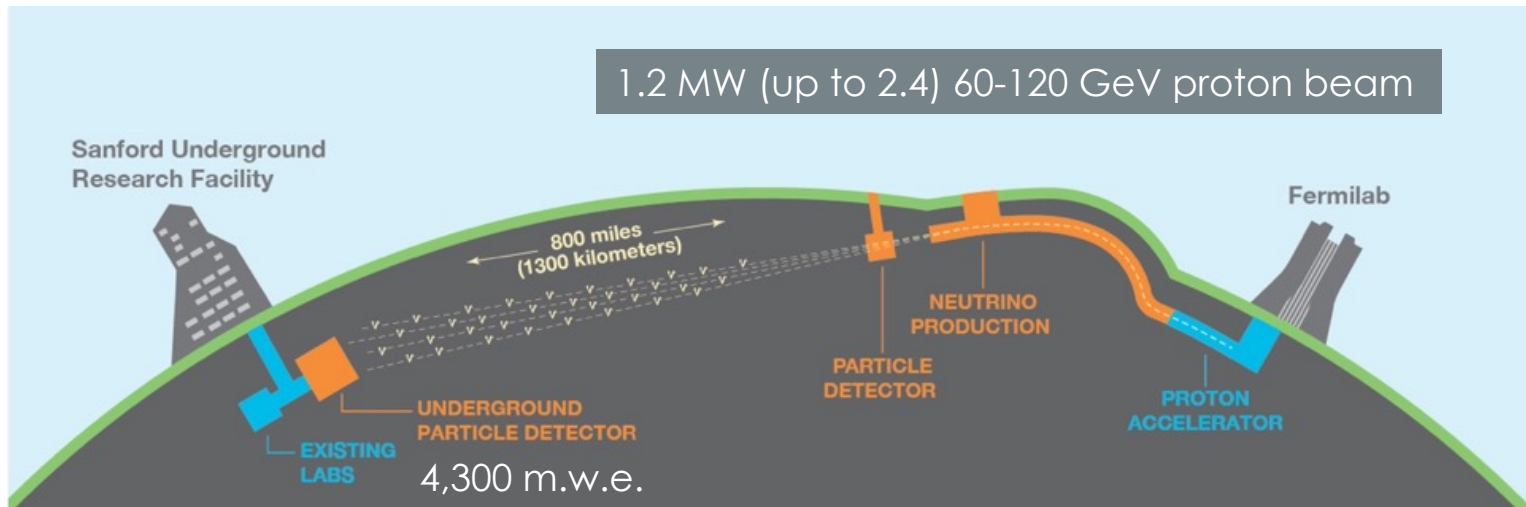
Supernovae neutrino detection

MeV neutrino burst observed with increase in PMT counting rates

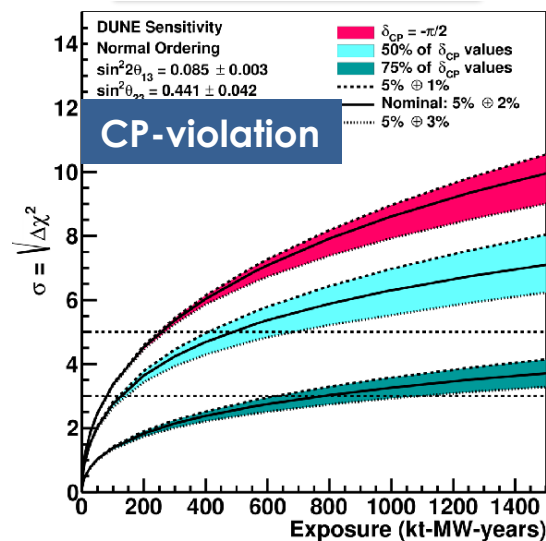
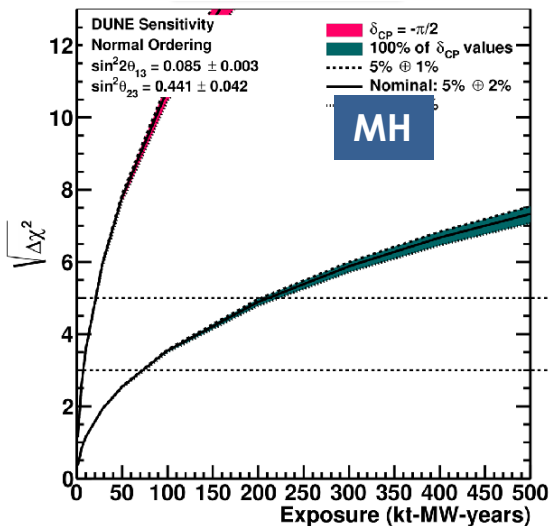
background rejection by requesting O(ns) coincidences between 6-10 PMTs on same DOM

In collaboration with T. Foglizzo (AIM)





4 x 10 kt fiducial mass liquid argon TPC modules, two technologies: **Single Phase** and **Dual Phase**



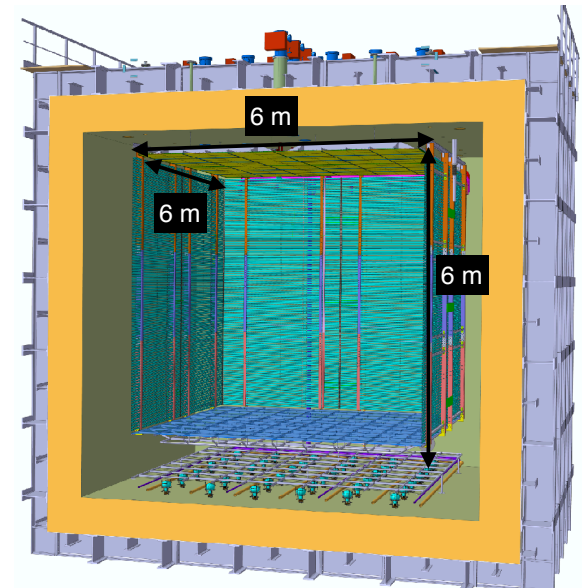
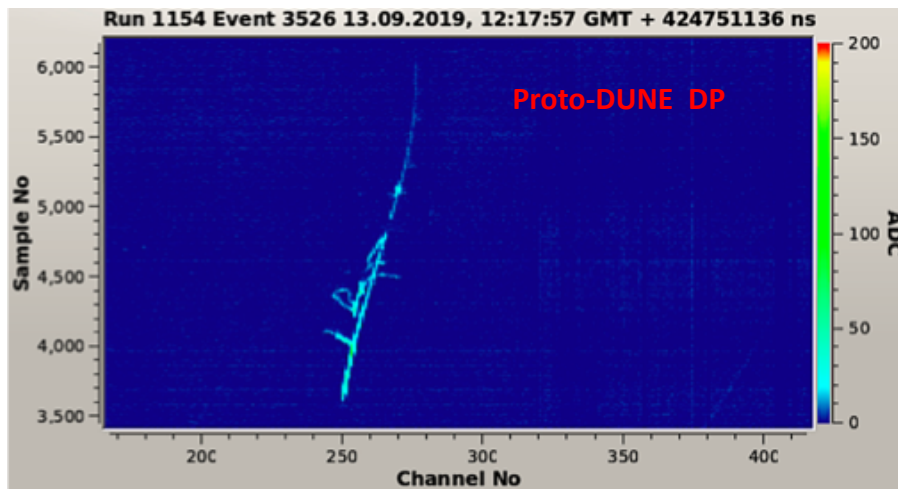
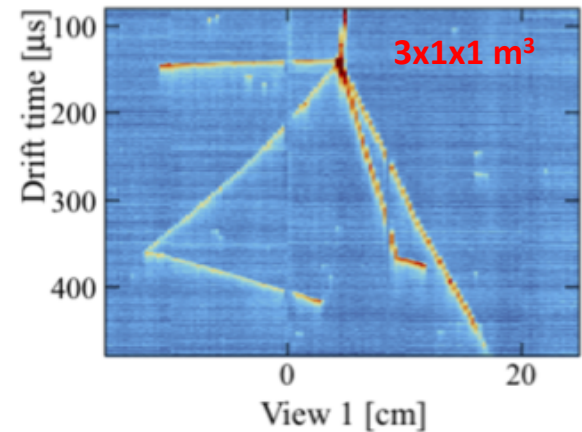
- Unitarity test
- Proton decay
- Supernova neutrinos
- Astroparticle
- Exotic physics

3x1x1 m³ in 2017 => A 4 tonne demonstrator for large-scale dual-phase LAr TPCs, B. Aimard et al.. arXiv:1806.03317, JINST 13 (2018) no.11, P11003.

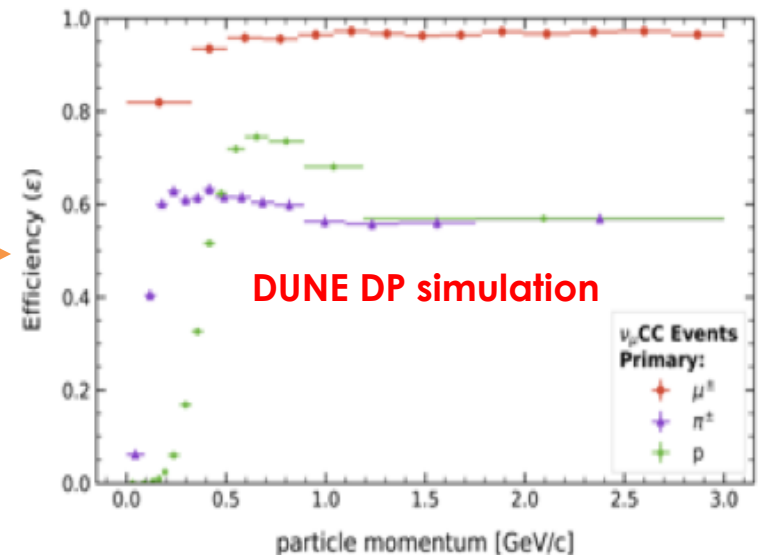
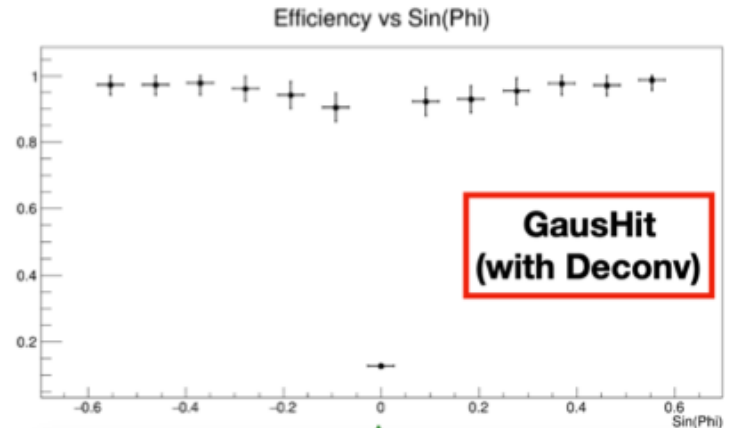
Proto-DUNE DP taking data since September 2019
APC technical contribution with light readout board (J. Dawson) + run coordinators (J. Dawson, S. Sacerdoti, A. Tonazzo)

Preparation for **DUNE**:

- TDR: published (FERMILAB-DESIGN-2018-02 arXiv:1807.1033)
- DP TDR: J. Dawson co-editor of the electronics chapter (to be published)



- **T. Patzak**: member of IB
- **A. Tonazzo**: member of Speaker Committee
- **J. Dawson** co-convener dual-phase proto analysis
- **S. Sacerdoti**: analysis of slow-control parameters / performance
- **R. Bajou** and **E. Chardonnet**:
 - track reconstruction + particle identification with cosmic muons in proto-DUNE DP
 - integration in LArSoft, development of PANDORA algorithms for DP
- **A. Scarpelli** - PhD 2016-2019:
 - Leading role in 3x1x1 analysis
 - First benchmark DP - specific studies for physics performance
 - CNN particle identification and event classification



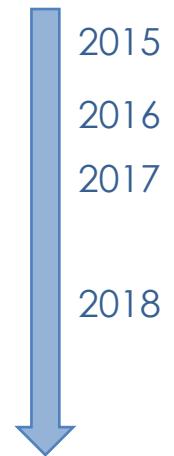


The team:

- ✓ **APC**: D. Franco, A. Tonazzo + 3 PhD (M. Lai, J. Rode, T. Hugues)
- ✓ L1 WBS Manager “Science, Simulation, and Computing”
- ✓ D. Franco: **National Coordinator** in France:
 - ✓ APC + LPNHE + CPPM
 - ✓ CPT – MINES ParisTech (P. Stringari, M. Campestrini) for Ar-Xe mixture
 - ✓ Technical support for beam tests by **IPNO** (M. Lebois, J. Wilson)

French main contributions to (among 18 publications since 2013):

- ✓ Pulse shape discrimination for a **background-free** experiment [PLB 743 \(2015\) 456](#)
- ✓ **Solar neutrino** physics in a large LAr detector [JCAP 1608 \(2016\) 8, 017](#)
- ✓ ^{39}Ar depletion fraction in **underground argon** [PRD 93 \(2017\) 081101](#)
- ✓ **Simulation** and LAr response model [JINST 12 \(2017\) P10015](#)
- ✓ Most accurate constraint of **LAr response** with ARIS, [PRD 97 \(2018\) 11 112005](#)
- ✓ World **best exclusion limit for 1.8-6 GeV WIMPs** [PRL 121 \(2018\) 081307](#)
- ✓ Extension of existing **exclusion limit to WIMP-electrons** [PRL 121 \(2018\) 111303](#)



Detector

- a **50 kg** dual-phase Liquid Argon TPC
- Using **Underground Argon**: depleted in ^{39}Ar
- In a **30 ton** borated liquid scintillator **neutron veto**
- In a **1000 ton Water** Cherenkov Veto
- Underground at Gran Sasso National Lab, Italy

Performance

S1 and S2 Yields:

- S1 Yield ~ 7.9 pe/keV at null field
- S1 Yield ~ 7.0 pe/keV at 200 V/cm
- S2 yield ~ 23 pe / e^-

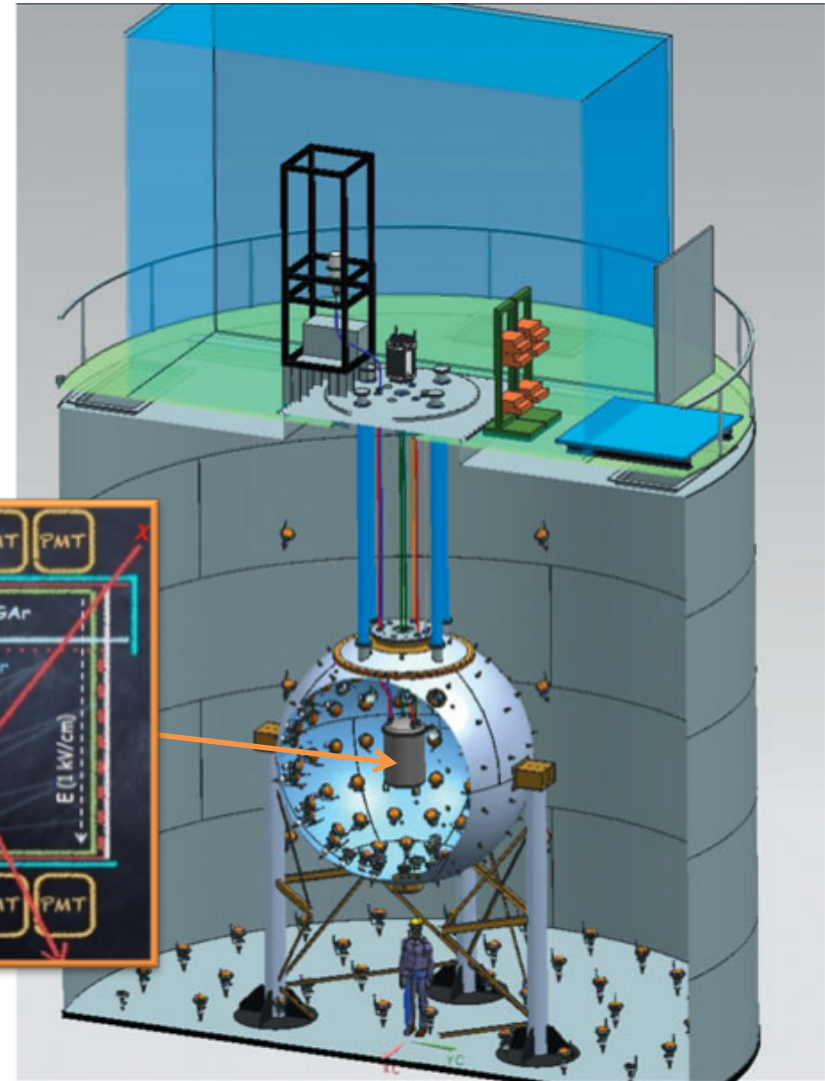
Electron lifetime > 10 ms

Maximum drift time: 376 μs

Position reconstruction:

- Resolution in Z ~ 1 mm
- Resolution in XY < 1 cm

Neutron Veto Rejection Efficiency: 99.6%

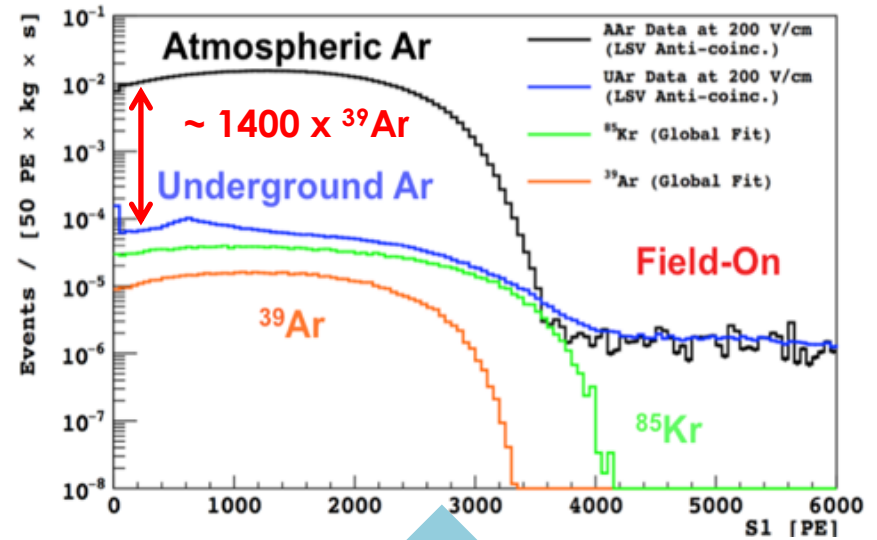
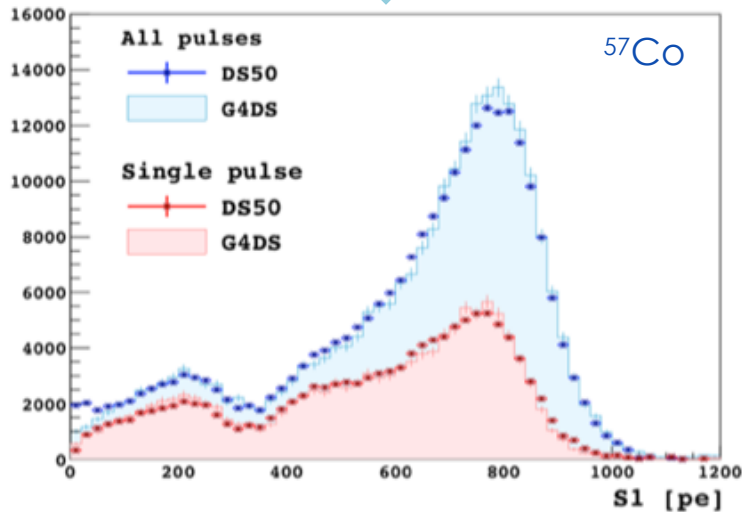


G4DS

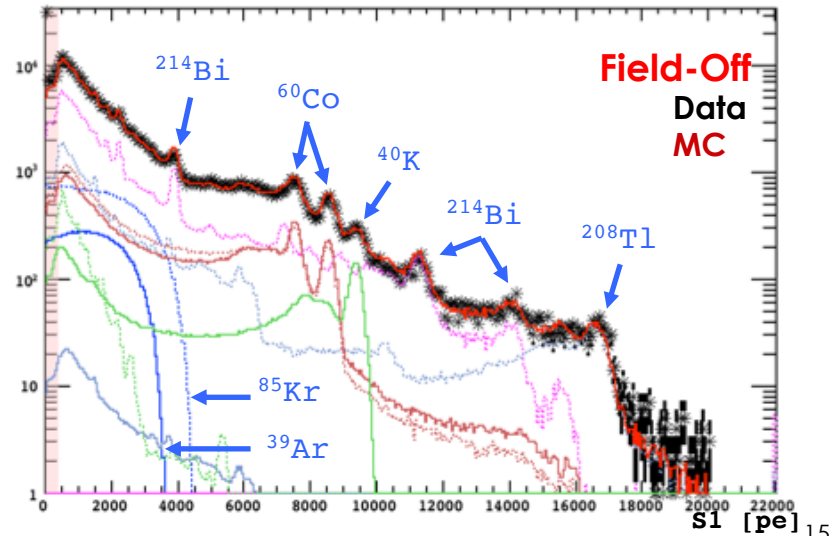
the DarkSide Monte Carlo package

- ✓ **PARIS**: custom made LAr scintillation-ionization response model
- ✓ **Percent level accuracy** in energy scale and resolution

Calibration



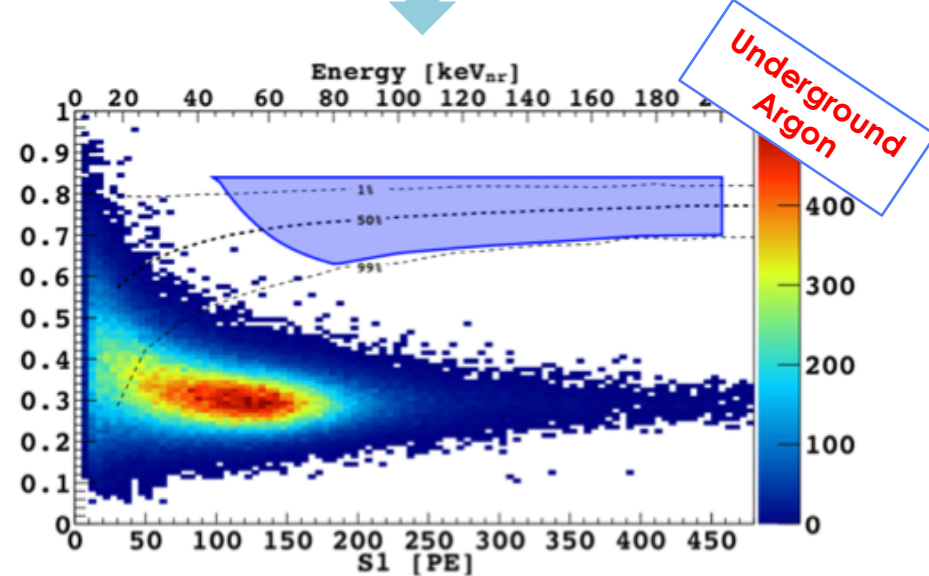
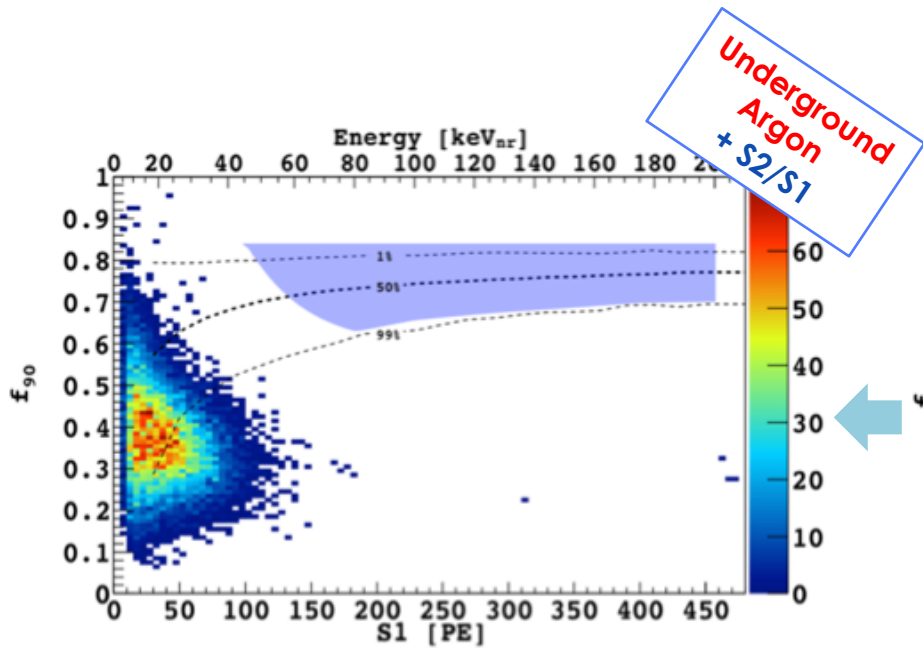
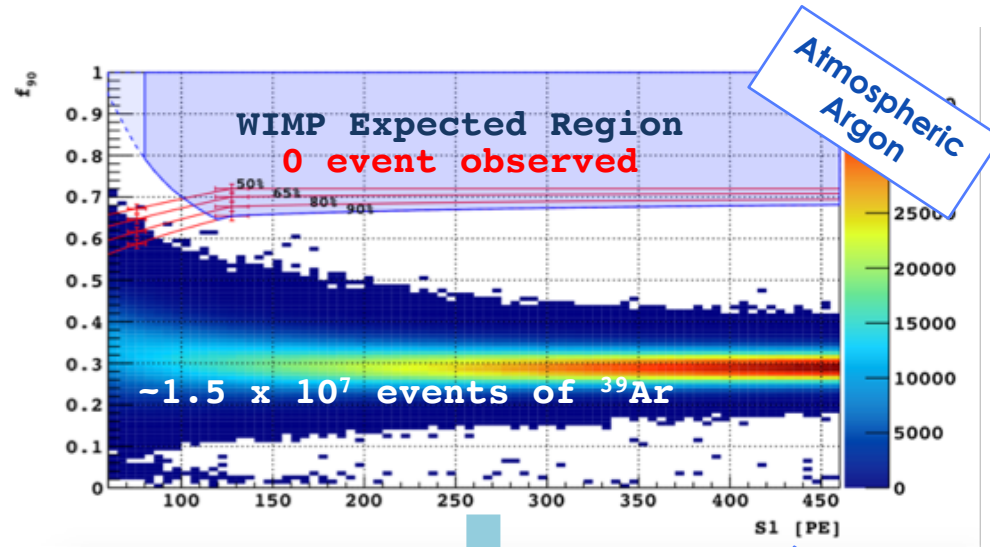
Result





A background-free experiment

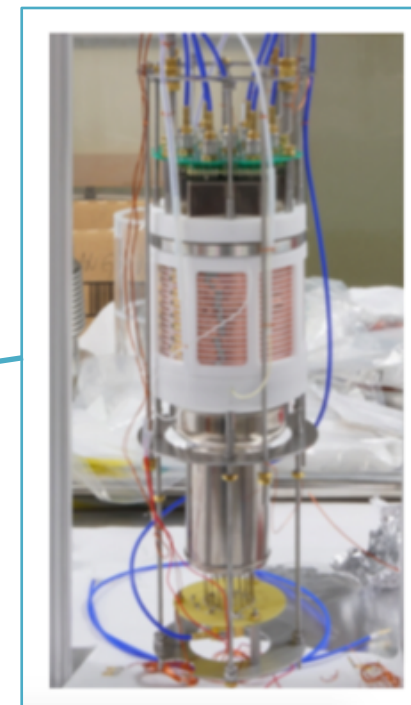
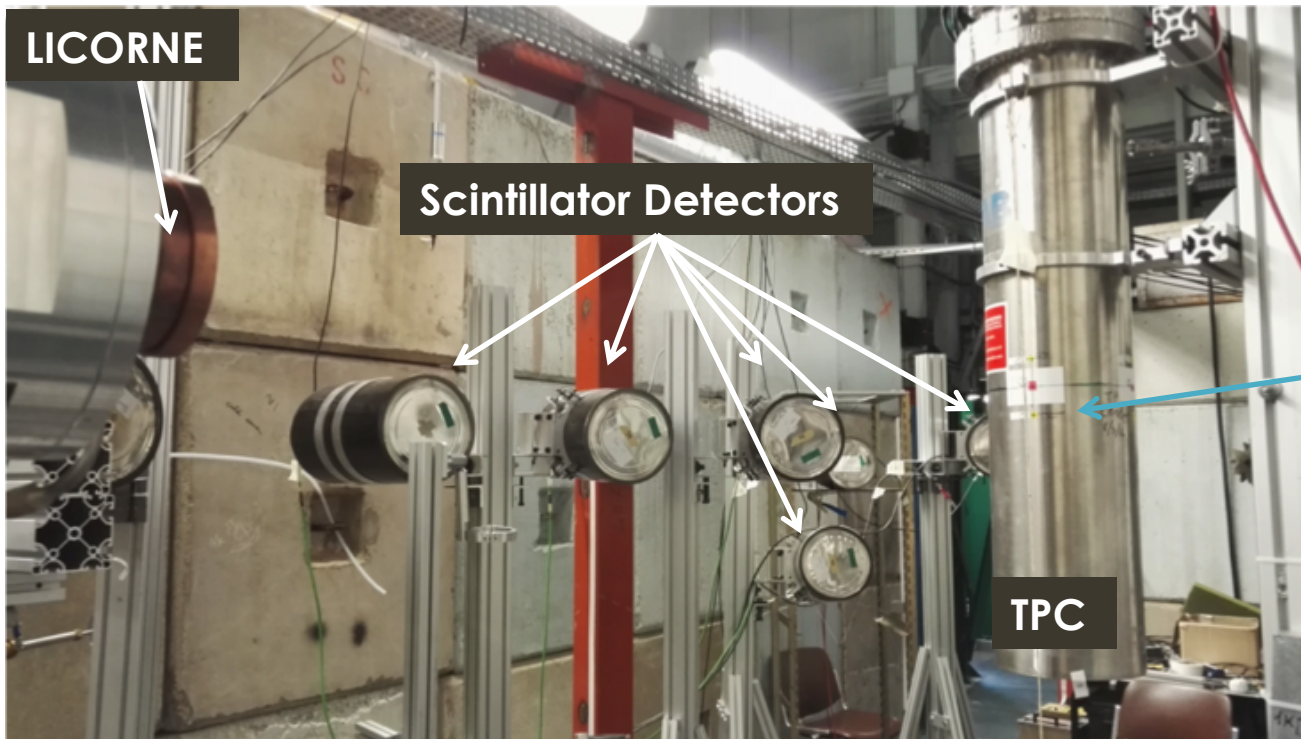
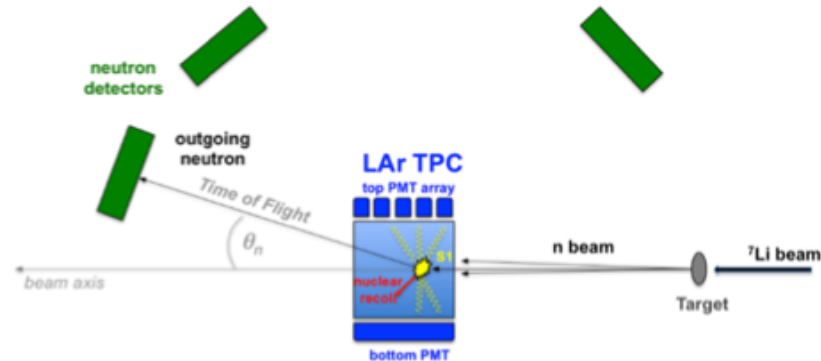
- ✓ **Pulse shape discrimination (f90)** power demonstrated with atmospheric argon
- ✓ **Blind analysis** in the ~532 days analysis
- ✓ **f90 analytical model by APC**

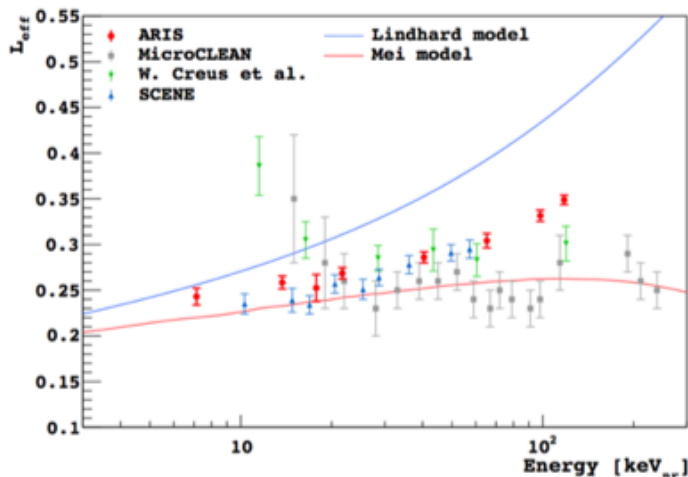


Background-free over more than 530 days!

LICORNE source: **inverted ${}^7\text{Li}(p,n){}^7\text{Be}$** reaction

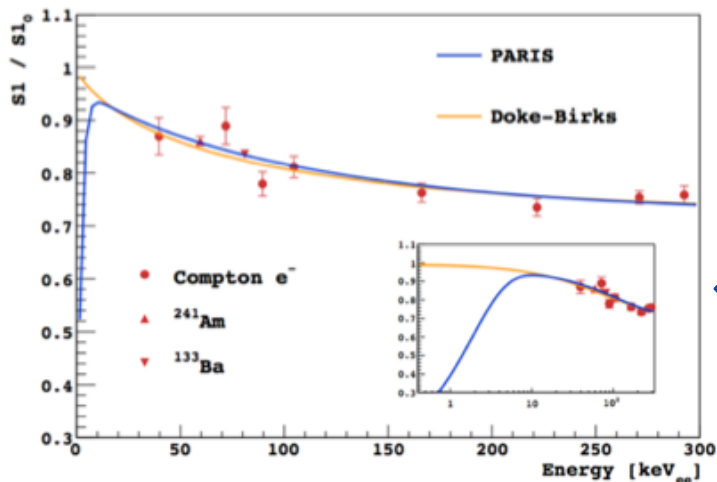
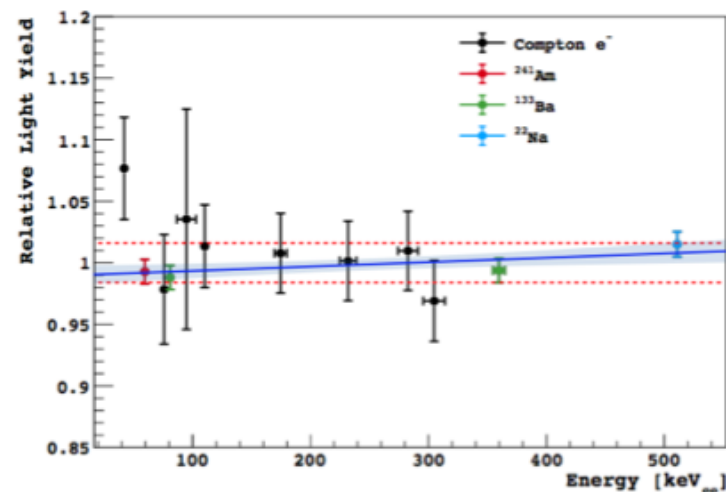
- ☑ Pulsed (1.5 ns width)
- ☑ Monochromatic: $<6\%$ ($\mu \sim 1450$ keV $\sigma \sim 85$ keV)
- ☑ Collimated: < 2 degrees
- ☑ Correlated 478 keV gammas: ER calibration





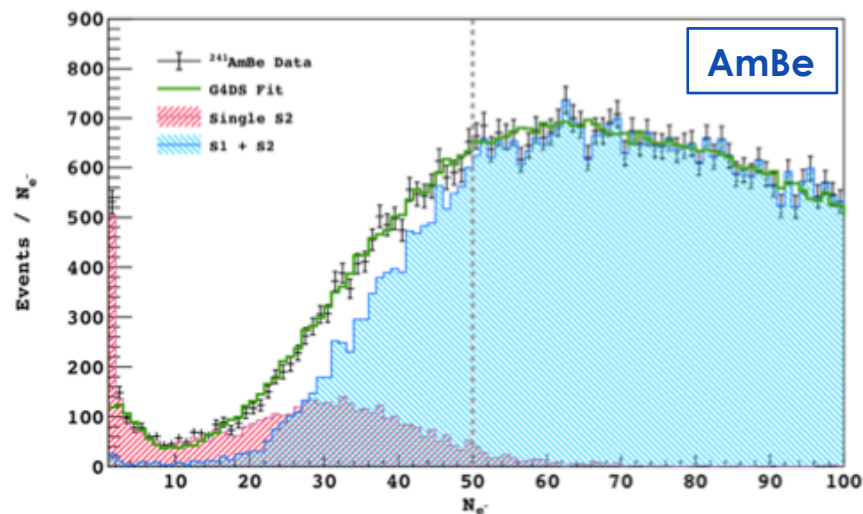
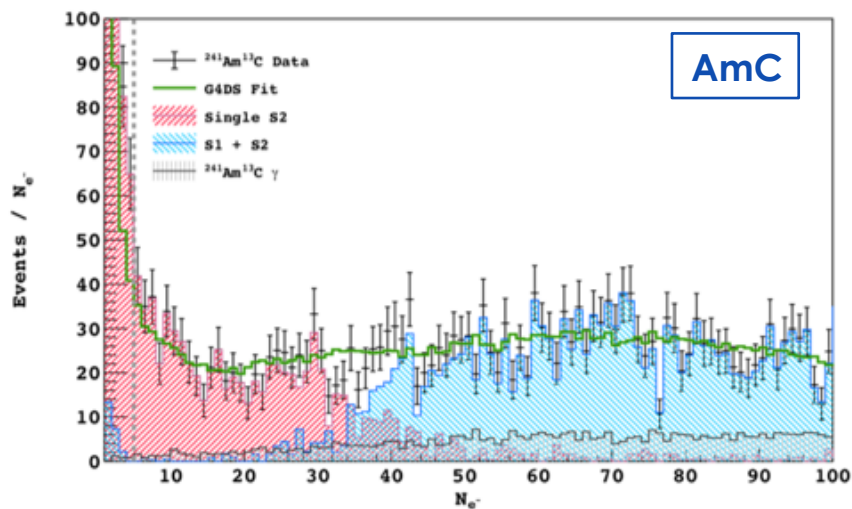
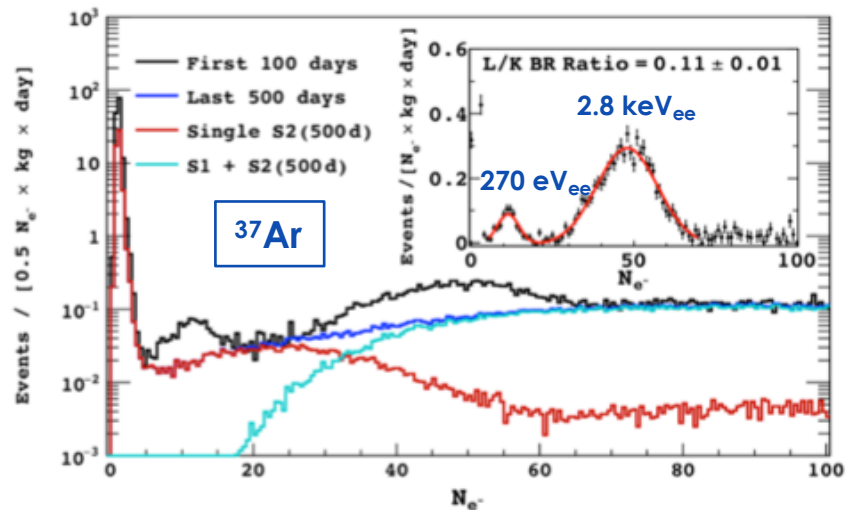
Most accurate measurement of the **quenching** effect: **crucial for the low-mass analysis**

Best constraint on **linearity** of LAr response to ER at field-off

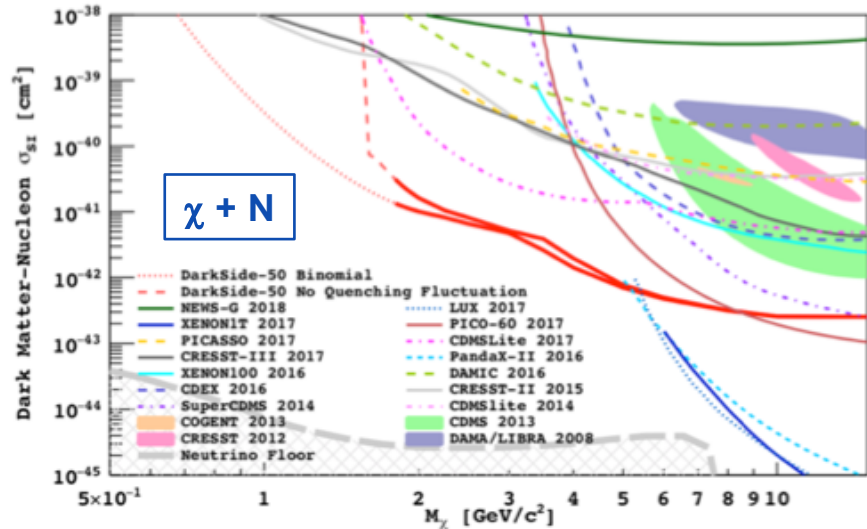
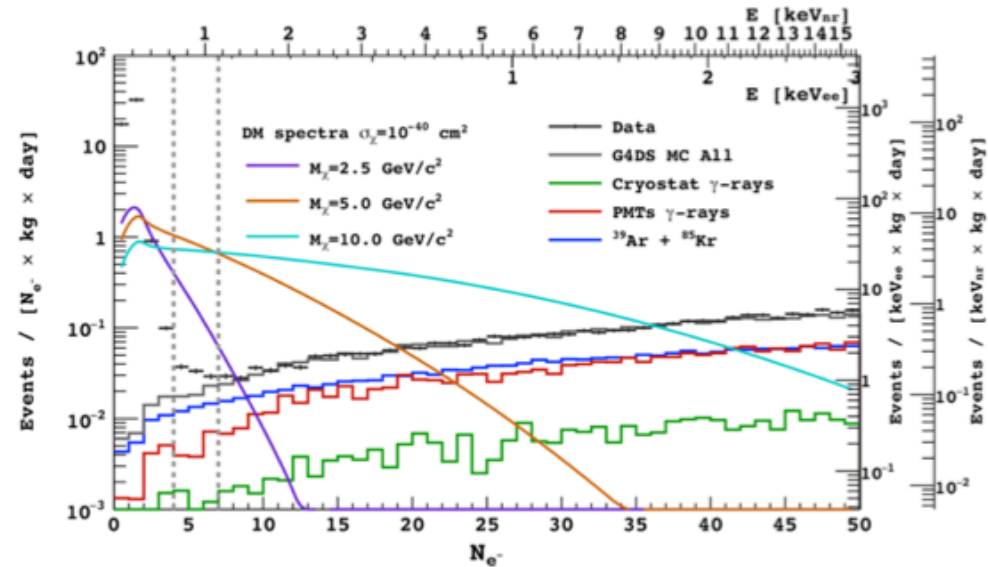
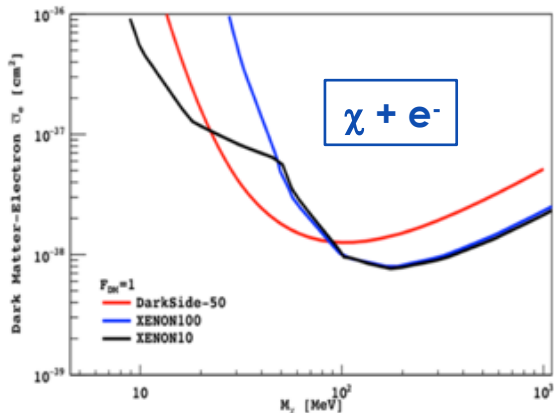


Excellent agreement with the **PARIS** model

- ☑ **100%** trigger efficiency at $\sim 1.5 e^-$
- ☑ Analysis threshold at $4 e^-$: **$\sim 100 eV$** or $\sim 600 eV_{nr}$
- ☑ ER calibration with ^{37}Ar : $270 eV$ and $2.8 keV$
- ☑ NR calibration with *in situ* AmC and AmBe neutron sources
- ☑ Beam experiment results (**ARIS** + SCENE)



- ✓ **Background model** simulated with G4DS, extrapolated from the high energy range
- ✓ **Excellent agreement** down to 7 electrons and excess between 4 and 7 electrons conservatively attributed to DM
- ✓ **Best limits to WIMPs-nucleus $< 6 \text{ GeV}/c^2$** limited at $1.8 \text{ GeV}/c^2$ because of the unknown quenching fluctuation model
- ✓ Improved limits in the **WIMP-electron scattering** assuming heavy mediator



Limits with:

- HistFactory
- Bayesian
- Yellin (in progress)

Models:

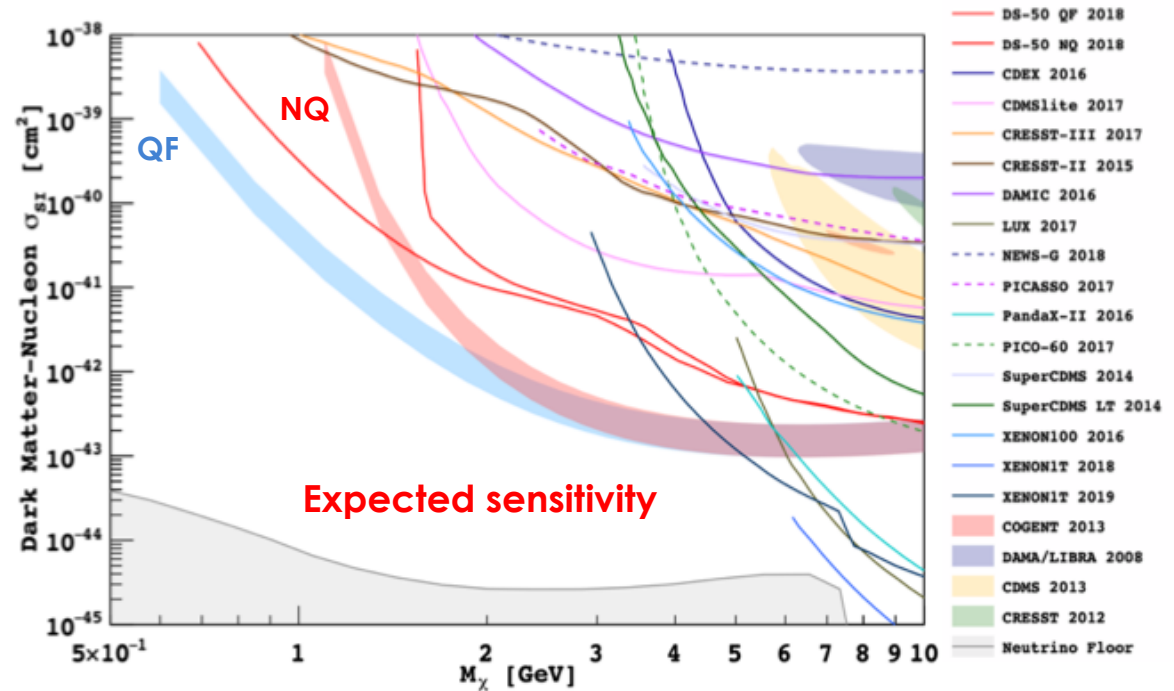
- WIMP – nucleus interactions
- WIMP – electron interactions
- Migdal effect
- Modulations
- Solar and galactic axions via axio-electric effect

DS50:

- New ER calibration
- New NR calibration
- New background model from screening materials
- New detector response
- New data selection
- 50% more statistics

Next generation (DS-Proto/DS-20k/ARGO):

- Ready to evaluate sensitivities for each model

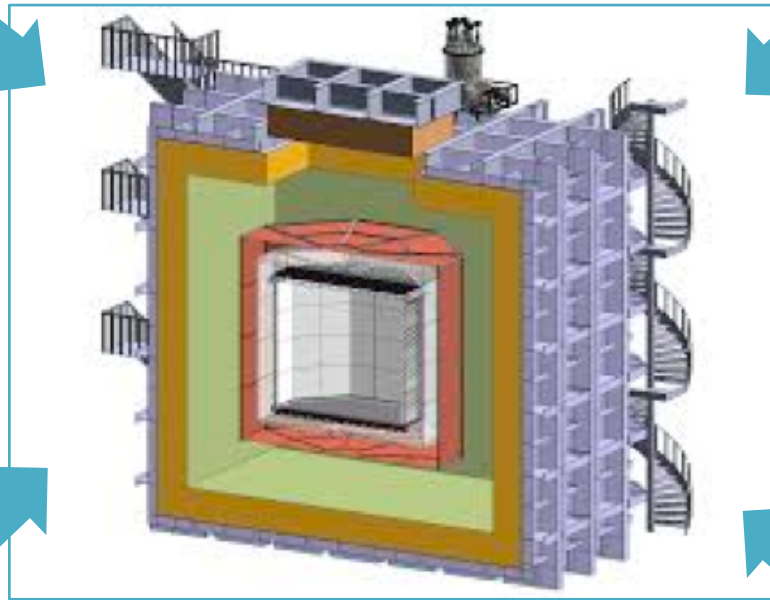
**3 papers expected in 2020:**

- Low-energy calibration (down to ~ 75 eVee)
- New low-mass limits
- New “exotic” limits

The Global Argon Dark Matter Collaboration

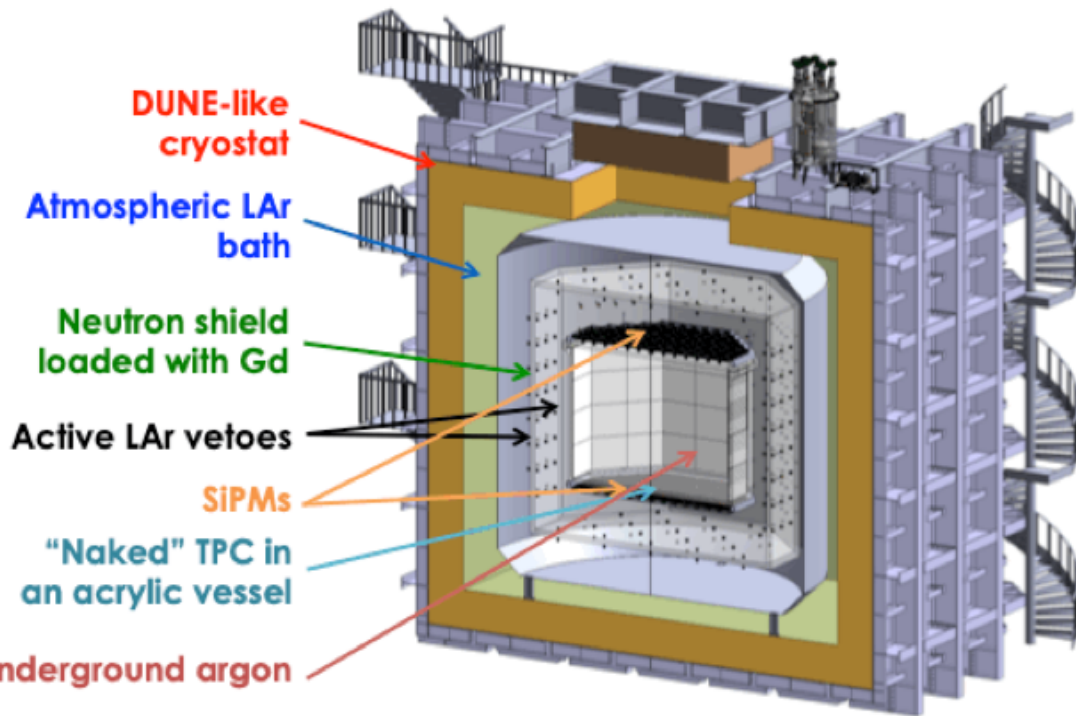


DarkSide-20k



~450 scientists
~70 institutes
+ Neutrino Platform

Design for a large mass **bg-free** LAr TPC



TPC: cryostat + teflon vessel => **only acrylic vessel**

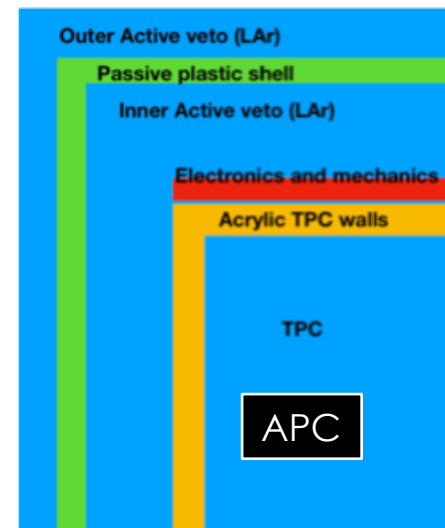
- ☑ Almost doubled the target mass (no UAr buffer): ~35 tonnes
- ☑ Removed cryostat, among the main sources of radiogenic neutrons

Photo-sensors: ~15 m² of SiPMs

- ☑ **Radiopure** and limited amount of material
- ☑ High photodetection **efficiency**

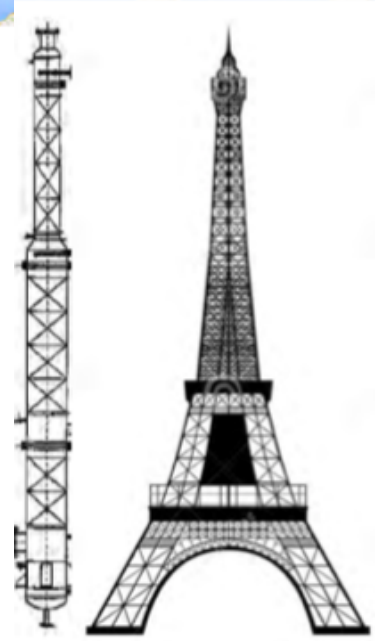
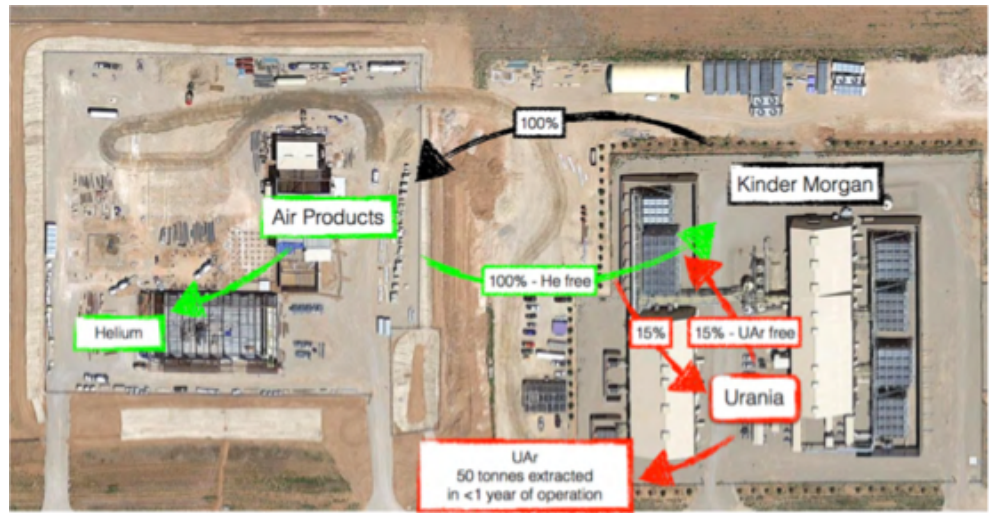
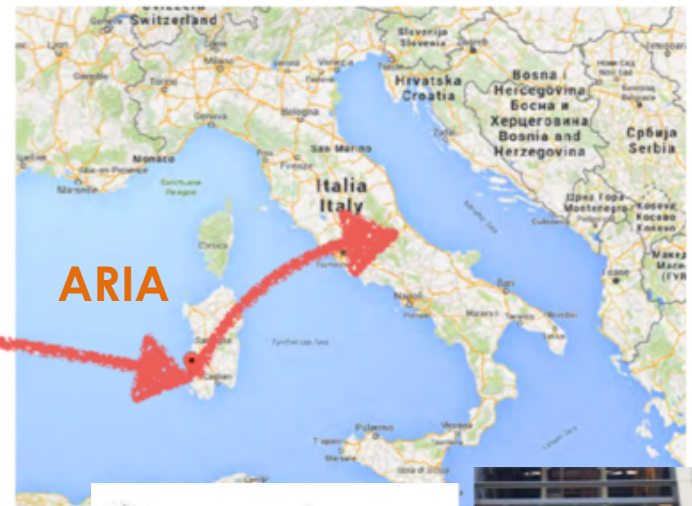
Vetoes: LAr bath + moderator in acrylic loaded with Gd

- ☑ **Neutron veto conceptual design by APC&LPNHE**
- ☑ **DUNE-like cryostat** (GTT patent)
- ☑ CERN Neutrino Platform technical support
- ☑ No organic scintillators





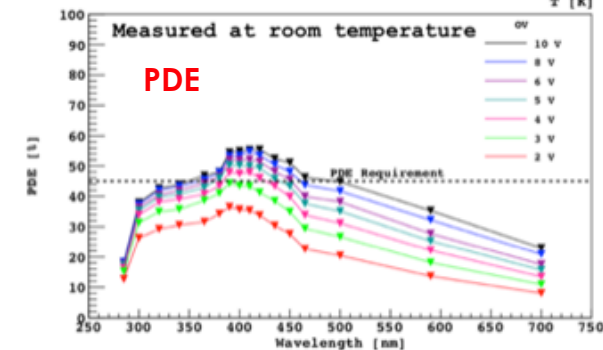
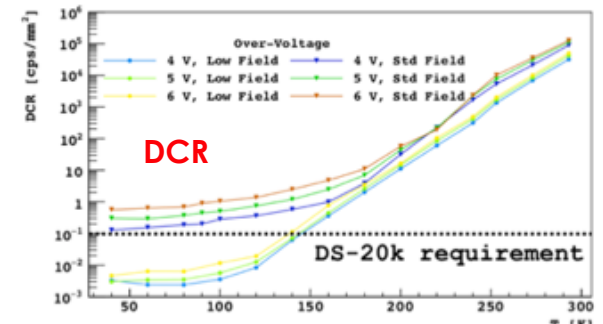
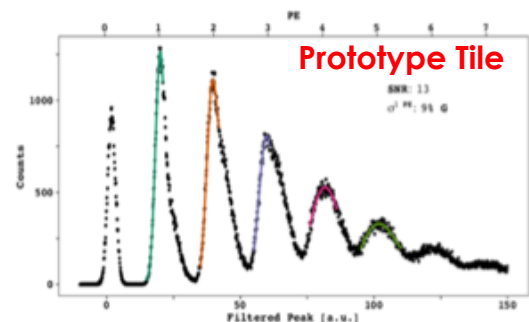
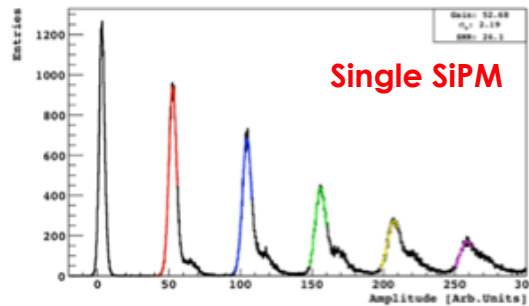
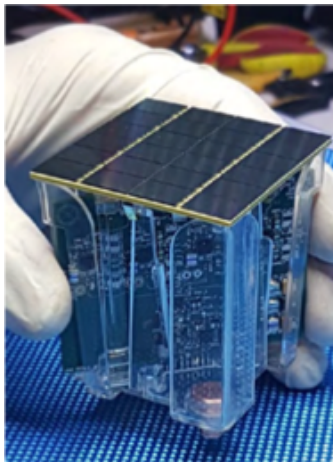
Underground/Depleted Argon

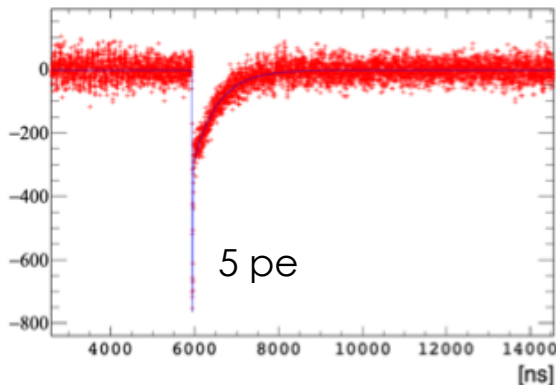


FBK NUV-HD SiPM

- ✓ Strict collaboration with Fondazione Bruno Kessler (**FBK**): development of specific SiPM for LAr (50 PDM under way)
- ✓ The FBK technology on transfer to **LFoundry** for mass production (starting April 2019)
- ✓ Packaging of 240,000 SiPMs at **NOA**, a facility funded at LNGS

| | DS-20k requirement | SiPM tile (PDM) | |
|-------------------|-------------------------|--|---|
| Surface | 5x5cm ² | 24cm ² prototype 25cm ² final PDM | ✓ |
| Power dissipation | <250mW | ~170mW | ✓ |
| PDE | >40% | 50% · ε _{geom} = 45% | ✓ |
| Noise Rate | <0.1cps/mm ² | 0.004cps/mm ² | ✓ |
| Time Resolution | 0(10ns) | 16ns | ✓ |
| Dynamic Range | >50 | ~100 | ✓ |

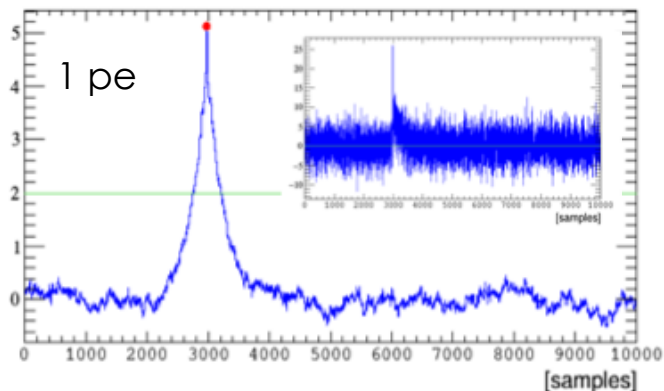




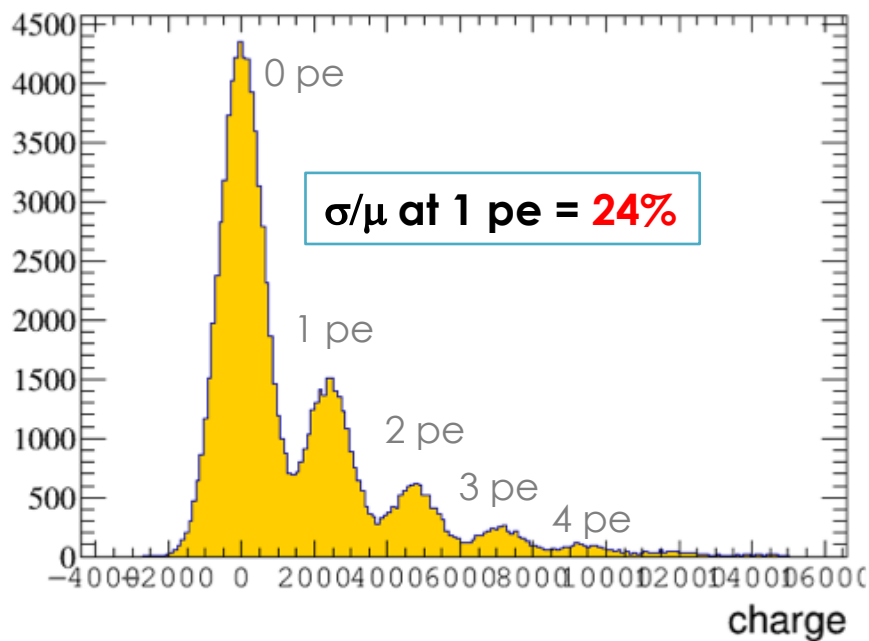
Matched filter



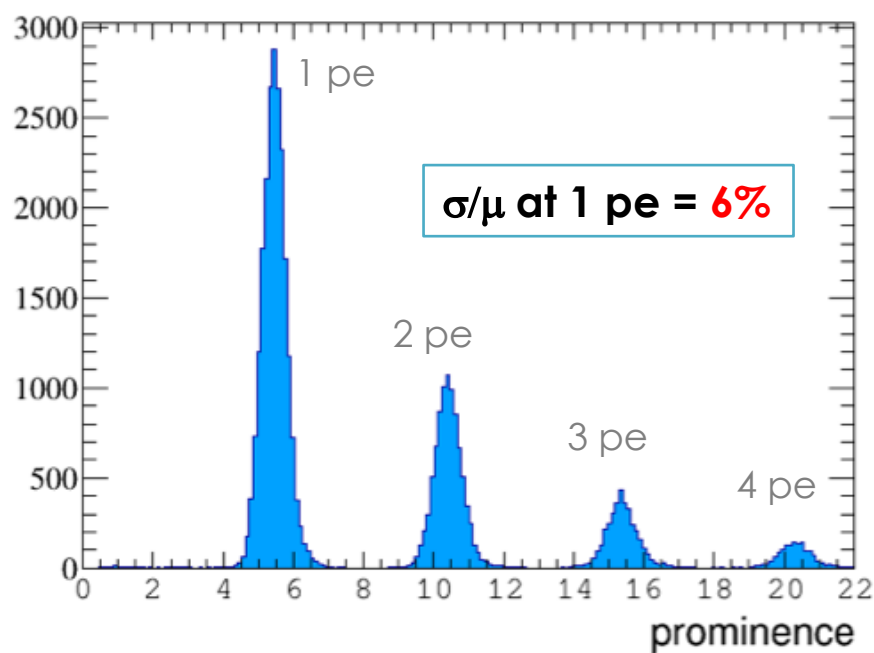
Sigma = 7.5 ns
 Tau Slow = 575 ns
 Slow component = 94%



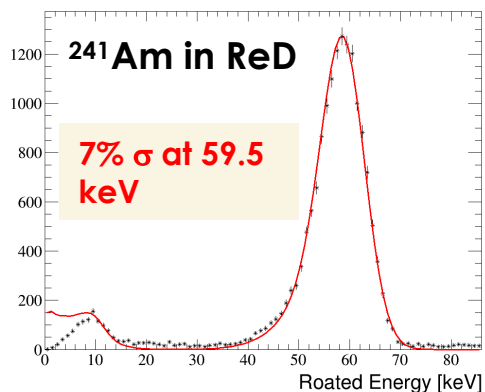
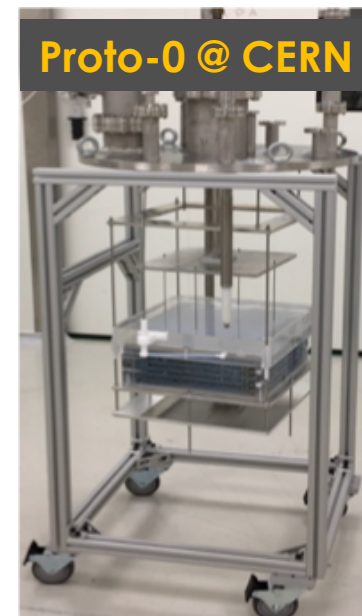
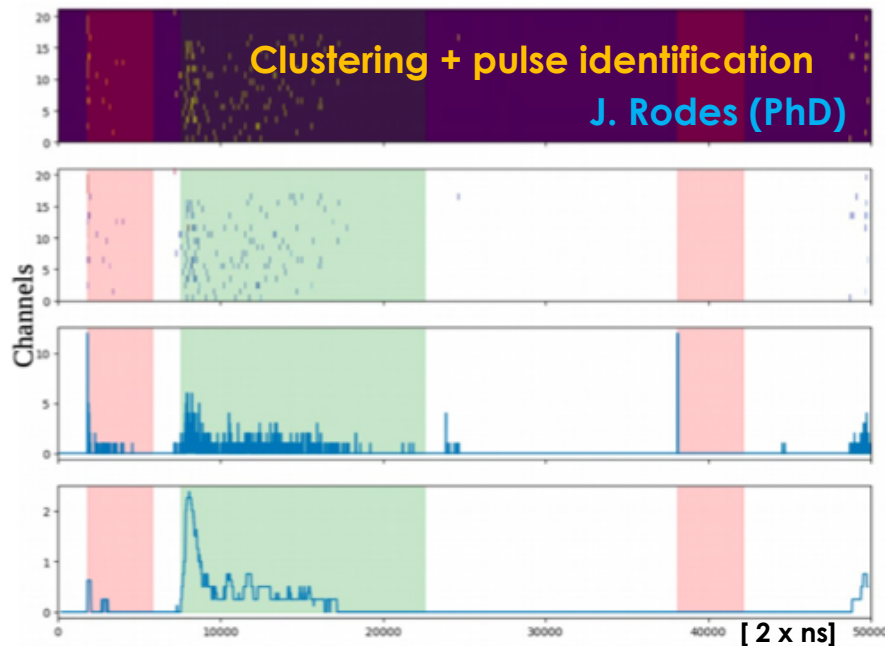
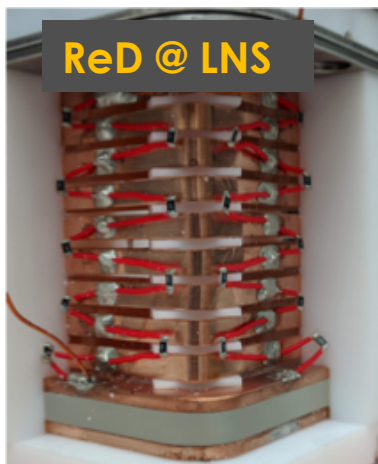
Charge Integration



Filtering



Offline software reconstruction for **ReD** => **PROTO-0** => **PROTO-1** => **DS-20k**

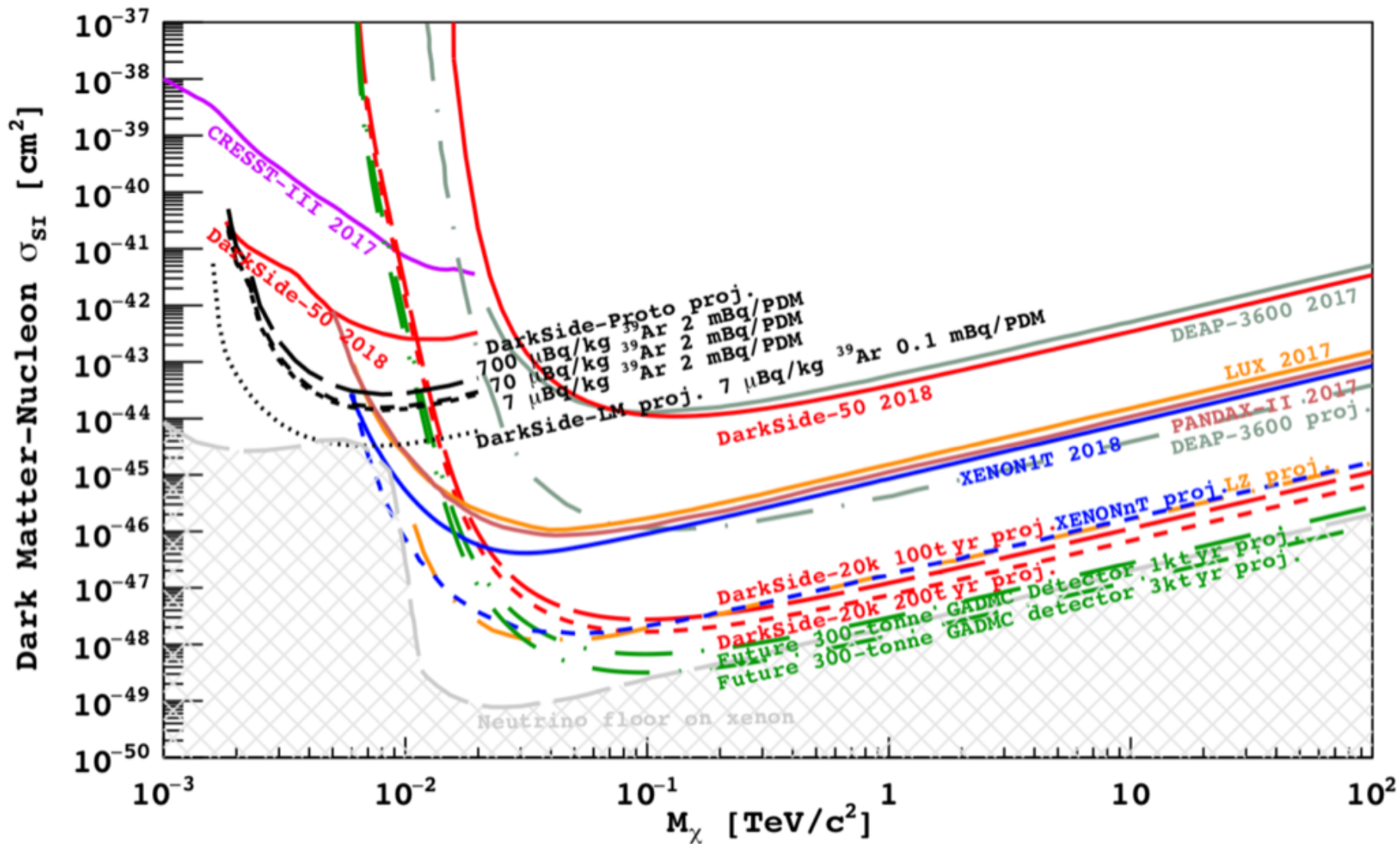


Successful reconstruction but need optimization in CPU time

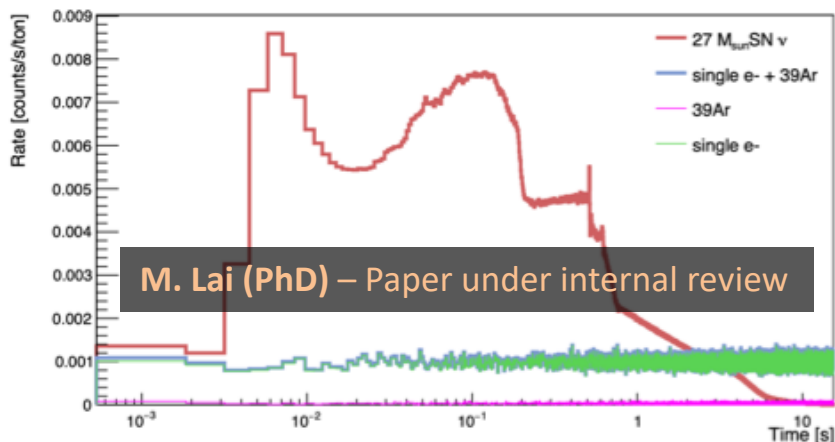
New request for adapting this approach to the “online”

Big effort in computing for handling 8,000 channels in DS-20k

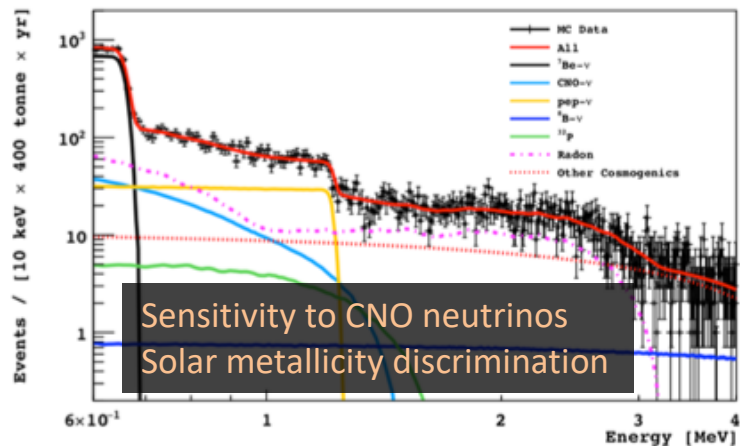
Toward the Neutrino Floor



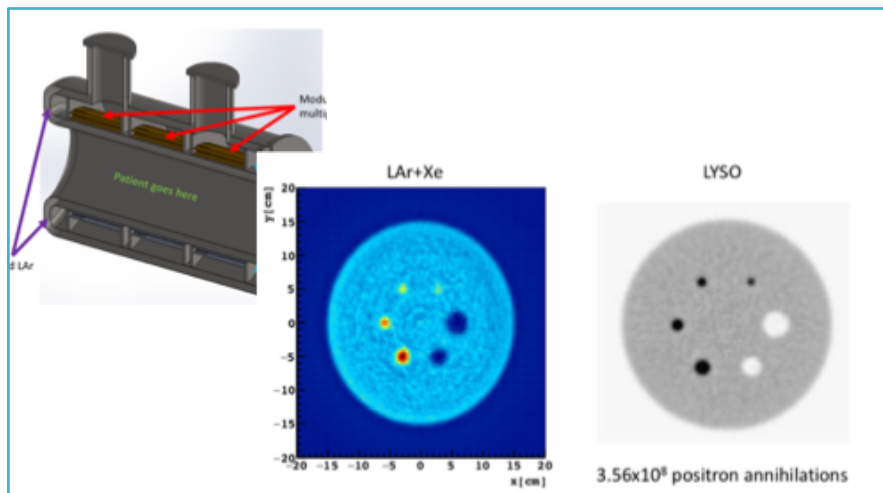
Supernova neutrinos via coherent scattering



Solar neutrino with ARGO



PET based on LAr + O(ppm) Xe

LAr + 5-10% ¹³⁶Xe: $0\nu\beta\beta$

- Radon suppression
- Low temperature => SiPM
- Cryogenic veto => material minimization
- Xe acts as wavelength shifter => no need of TPB
- how to dissolve Xe at 20% molar fraction?
=> MINES ParisTech / CPT



Conclusions and Prospects

DarkSide

- ☑ **Background-free** experiment for high mass WIMP thanks to pulse shape discrimination in S1 (unique to LAr)
- ☑ **Best limit** in the low mass 1.8-6 GeV window
- ☑ Strong WIMP **discovery potential** in the next decade (from 1 GeV to 10 TeV)
- ☑ New opportunities also in looking for **low-energy astroparticle and medical applications**
- ☑ Potential **synergies with DUNE**: astroparticle physics, cryostat and cryogenics, optical simulations, photo-collection efficiency, calibrations