$B_s \to \mu^+ \mu^-$ Theory Status

Mikołaj Misiak

University of Warsaw

- 1. Introduction
- 2. Matching calculations in the SM and beyond
- 3. Soft photon emission
- 4. Enhanced QED effects
- 5. Update of the input parameters
- 6. The SM prediction and its error budget
- 7. Summary

B-meson or Kaon decays occur at low energies, at scales $\mu \ll M_W$.

We pass from the full theory of electroweak interactions to an effective theory by removing the highenergy degrees of freedom, i.e. integrating out the W-boson and all the other particles with $m \sim M_W$.

$$\mathcal{L}_{\text{(full EW\times QCD)}} \longrightarrow \mathcal{L}_{\text{eff}} = \mathcal{L}_{\text{QED}\times \text{QCD}} \left(\begin{smallmatrix} \text{quarks} \neq t \\ \text{\& leptons} \end{smallmatrix} \right) + N \sum_{n} C_{n}(\mu) Q_{n}$$

 Q_n – local interaction terms (operators), $\qquad C_n$ – coupling constants (Wilson coefficients).

Operators (dim 6) that matter for $B_s o \mu^+\mu^-$ read:

$$Q_A=\left(ar{b}\gamma^lpha\gamma_5 s
ight)\left(ar{\mu}\gamma_lpha\gamma_5\mu
ight)$$
 — the only relevant one in the SM at the LO in QED

$$Q_{S(P)} = \left(ar{b}\gamma_5 s
ight) \left(ar{\mu}(\gamma_5)\mu
ight) \ = rac{i(ar{b}\gamma^lpha\gamma_5 s)\partial_lpha(ar{\mu}(\gamma_5)\mu)}{m_b + m_s} \ + rac{E}{v_{
m anishing}} + rac{T}{v_{
m constraint}}$$

Necessary non-perturbative input:
$$\langle 0|ar{b}\gamma^{lpha}\gamma_5 s|B_s(p)
angle = ip^{lpha}f_{B_s}$$

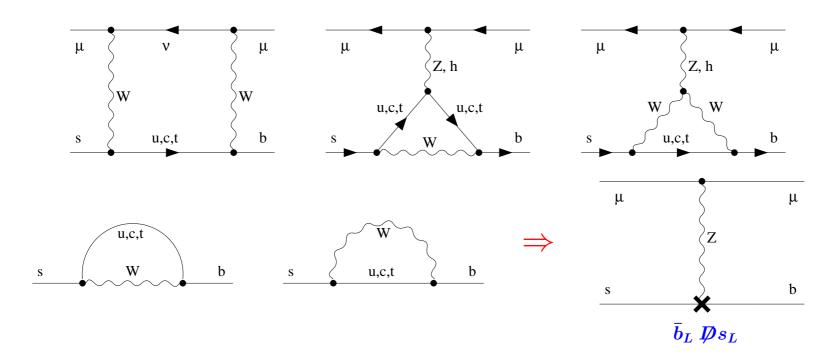
Such a matrix element vanishes for $(ar b\gamma^{lpha}s)$ and (ar bs) because B_s is a pseudoscalar.

It also vanishes for $(ar{b}\sigma^{lphaeta}s)$ because no antisymmetric tensor can be formed from p^{lpha} alone.

$$Q_V = \left(ar{b}\gamma^lpha\gamma_5 s
ight)(ar{\mu}\gamma_lpha\mu)$$
 gives no contribution at the LO in QED because $p^lpha(ar{\mu}\gamma_lpha\mu) = ar{\mu}p\!\!\!/\mu = ar{\mu}(p\!\!\!/_{\mu^+} + p\!\!\!/_{\mu^-})\mu = ar{\mu}(-m_\mu + m_\mu)\mu = 0.$

 Q_S gets generated in the SM via the Higgs exchange, but...—see next page.

Evaluation of the LO Wilson coefficients in the SM:

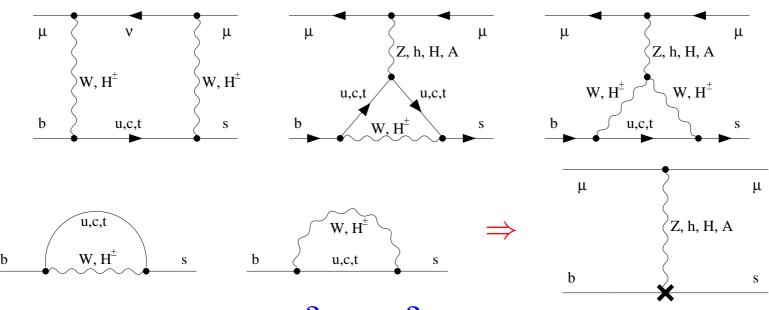


$$egin{align} C_A^{(0)} &= rac{1}{2} Y_0 \left(m_t^2 / M_W^2
ight), \quad Y_0(x) = rac{3 x^2}{8 (x-1)^2} \ln x + rac{x^2 - 4 x}{8 (x-1)}, \ C_S &= \mathcal{O} \left(rac{m_\mu}{M_W}
ight), \quad C_P = 0. \end{aligned}$$

Effects of C_S on the branching ratio are suppressed by $M_{B_s}^2/M_W^2$ \Rightarrow negligible.

Thus, only C_A matters in the SM.

Evaluation of the Wilson coefficients in the Two-Higgs-Doublet Model II



$$aneta = v_2/v_1, ~~ z = M_{H^\pm}^2/m_t^2,$$

$$C_S \simeq C_P \simeq rac{m_\mu m_b an^2 eta}{4 M_W^2} rac{\ln z}{z-1} > 0, {
m H.E.~Logan~and~U.~Nien} \ {
m NPB~586~(2000)~39} \ ({\cal O}(aneta) {
m neglected})$$

H.E. Logan and U. Nierste, $(\mathcal{O}(\tan\beta) \text{ neglected})$

$$\mathcal{B}(B_s o \mu^+ \mu^-) \simeq ext{(const.)} \left[\left| rac{2m_\mu}{M_{B_s}} C_A - C_P
ight|^2 + \left| C_S
ight|^2
ight]$$

$$C_A = C_A^{\text{SM}} + \Delta C_A$$
positive small
$$\Rightarrow \begin{cases} \text{suppression for moderate } C_{S,P} \\ \text{enhancement for huge } \tan \beta \text{ only} \end{cases}$$

Average time-integrated branching ratio:

$$egin{aligned} \overline{\mathcal{B}}(B_s
ightarrow \mu^+ \mu^-) &= rac{|N|^2 M_{B_s}^3 f_{B_s}^2}{8\pi \Gamma_H^s} eta \left(|m{rC_A} - m{uC_P}|^2 m{F_P} + |m{ueta C_S}|^2 m{F_S}
ight) + \mathcal{O}(lpha_{em}), \ & ext{where} \quad m{N} = rac{V_{tb}^* V_{ts} \, G_F^2 M_W^2}{\pi^2}, \quad m{r} = rac{2m_\mu}{M_{B_s}}, \quad m{eta} = \sqrt{1-m{r}^2}, \quad m{u} = rac{M_{B_s}}{m_b+m_s}, \ m{F_P} &= 1 - rac{\Delta \Gamma^s}{\Gamma_L^s} \sin^2\left[rac{1}{2}\phi_s^{ ext{NP}} + rg(m{rC_A} - m{uC_P})
ight] & \stackrel{ ext{SM CP}}{\longrightarrow} & 1 \; , \ m{F_S} &= 1 - rac{\Delta \Gamma^s}{\Gamma_L^s} \cos^2\left[rac{1}{2}\phi_s^{ ext{NP}} + rg C_S
ight] & \stackrel{ ext{SM CP}}{\longrightarrow} & rac{\Gamma_H^s}{\Gamma_L^s} & \stackrel{ ext{derived following [K. de Bruyn \it{et al.}, Phys. Rev. Lett. 109 (2012) 041801]} \ m{F_S} &= 1 - rac{\Delta \Gamma^s}{\Gamma_L^s} \cos^2\left[rac{1}{2}\phi_s^{ ext{NP}} + rg C_S
ight] & \stackrel{ ext{SM CP}}{\longrightarrow} & rac{\Gamma_H^s}{\Gamma_L^s} & rac{ ext{derived following [K. de Bruyn \it{et al.}, Phys. Rev. Lett. 109 (2012) 041801]} \ m{F_S} &= 1 - rac{\Delta \Gamma^s}{\Gamma_L^s} \cos^2\left[rac{1}{2}\phi_s^{ ext{NP}} + rg C_S
ight] & \stackrel{ ext{SM CP}}{\longrightarrow} & rac{\Gamma_H^s}{\Gamma_L^s} & rac{ ext{derived following [K. de Bruyn \it{et al.}, Phys. Rev. Lett. 109 (2012) 041801]} \ m{F_S} &= 1 - rac{\Delta \Gamma^s}{\Gamma_L^s} \cos^2\left[rac{1}{2}\phi_s^{ ext{NP}} + rg C_S
ight] & \stackrel{ ext{NP}}{\longrightarrow} & \frac{1}{\Gamma_L^s} & \frac{1}{2}\phi_s^{ ext{NP}} & \frac{1}{2}\phi_s^{ ext{NP}}$$

Average time-integrated branching ratio:

$$egin{aligned} \overline{\mathcal{B}}(B_s
ightarrow \mu^+ \mu^-) &= rac{|N|^2 M_{B_s}^3 f_{B_s}^2}{8\pi \Gamma_H^s} oldsymbol{eta} \left(|m{rC_A} - m{uC_P}|^2 m{F_P} + |m{ueta C_S}|^2 m{F_S}
ight) + \mathcal{O}(lpha_{em}), \ & ext{where} \quad N &= rac{V_{tb}^* V_{ts} G_F^2 M_W^2}{\pi^2}, \quad m{r} = rac{2m_\mu}{M_{B_s}}, \quad m{eta} = \sqrt{1-m{r}^2}, \quad m{u} = rac{M_{B_s}}{m_b + m_s}, \ m{F_P} &= 1 - rac{\Delta \Gamma^s}{\Gamma_L^s} \sin^2 \left[rac{1}{2} m{\phi}_s^{ ext{NP}} + rg(m{rC_A} - m{uC_P})
ight] \stackrel{ ext{SM CP}}{\longrightarrow} \mathbf{1}, \ m{F_S} &= 1 - rac{\Delta \Gamma^s}{\Gamma_L^s} \cos^2 \left[rac{1}{2} m{\phi}_s^{ ext{NP}} + rg m{C}_S
ight] \stackrel{ ext{SM CP}}{\longrightarrow} rac{\Gamma_H^s}{\Gamma_L^s} & \stackrel{ ext{derived following [K. de Bruyn \it{et al.}, Phys. Rev. Lett. 109 (2012) 041801]} \ m{F_S} &= 1 - rac{\Delta \Gamma^s}{\Gamma_L^s} \cos^2 \left[rac{1}{2} m{\phi}_s^{ ext{NP}} + rg m{C}_S
ight] & \stackrel{ ext{SM CP}}{\longrightarrow} rac{\Gamma_H^s}{\Gamma_L^s} & \stackrel{ ext{derived following [K. de Bruyn \it{et al.}, Phys. Rev. Lett. 109 (2012) 041801]} \ m{F_S} &= 1 - rac{\Delta \Gamma^s}{\Gamma_L^s} \cos^2 \left[rac{1}{2} m{\phi}_s^{ ext{NP}} + rg m{C}_S
ight] & \stackrel{ ext{SM CP}}{\longrightarrow} rac{\Gamma_H^s}{\Gamma_L^s} & \stackrel{ ext{derived following [K. de Bruyn \it{et al.}, Phys. Rev. Lett. 109 (2012) 041801]} \ m{E_S} &= 1 - rac{\Delta \Gamma^s}{\Gamma_L^s} \cos^2 \left[rac{1}{2} m{\phi}_s^{ ext{NP}} + rg m{C}_S
ight] & \stackrel{ ext{SM CP}}{\longrightarrow} rac{\Gamma_H^s}{\Gamma_L^s} & \stackrel{ ext{derived following [K. de Bruyn \it{et al.}, Phys. Rev. Lett. 109 (2012) 041801]} \ \m{E_S} &= 1 - rac{\Delta \Gamma^s}{\Gamma_S} \cos^2 \left[rac{1}{2} m{\phi}_s^{ ext{NP}} + rg m{C}_S
ight] & \stackrel{ ext{SM CP}}{\longrightarrow} rac{\Gamma_S}{\Gamma_S} & \frac{\Gamma_S}{\Gamma_S} & \frac{\Gamma_S}{\Gamma_S}$$

In the limit of no CP-violation, mass eigenstates are CP eigenstates:

Heavier, CP-odd:
$$B_s^H = \frac{1}{\sqrt{2}}(B_s + \bar{B}_s)$$
, annihilated by $\bar{b}\gamma_5 s + \bar{s}\gamma_5 b$, $(\tau_H = 1.616(10) \text{ ps})$

Lighter, CP-even:
$$B_s^L = \frac{1}{\sqrt{2}}(B_s - \bar{B}_s)$$
, annihilated by $\bar{b}\gamma_5 s - \bar{s}\gamma_5 b$, $(\tau_L = 1.519(4) \text{ ps})$

Our interactions in this <u>limit</u> are all CP-even:

$$egin{align*} Q_A + Q_A^\dagger &= \left[\left(ar{b} \gamma^lpha \gamma_5 s
ight) + \left(ar{s} \gamma^lpha \gamma_5 b
ight)
ight] \left(ar{\mu} \gamma_lpha \gamma_5 \mu
ight) \ Q_P + Q_P^\dagger &= \left[\left(ar{b} \gamma_5 s
ight) + \left(ar{s} \gamma_5 b
ight)
ight] \left(ar{\mu} \gamma_5 \mu
ight) \ Q_S + Q_S^\dagger &= \left[\left(ar{b} \gamma_5 s
ight) - \left(ar{s} \gamma_5 b
ight)
ight] \left(ar{\mu} \mu
ight) \end{array}
ight.
ight.$$
annihilates $m{B}_s^L$, produces CP-even dimuons

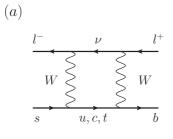
With SM-like CP-violation – still $Q_{A,P}$ annihilate B_s^H and Q_S annihilates B_s^L .

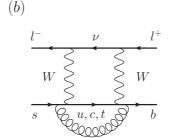
Beyond SM – interesting time-dependent observables, see arXiv:1303.3820, 1407.2771.

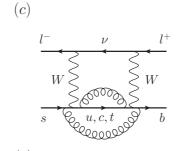
Evaluation of the NNLO QCD matching corrections in the SM

[T. Hermann, MM, M. Steinhauser, JHEP 1312 (2013) 097]

W-boxes: (1LPI)

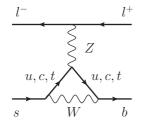


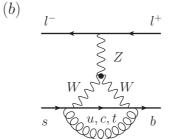


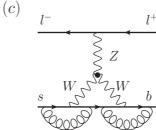


Z-penguins: (1LPI)

(a)



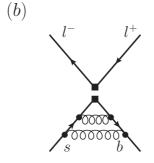


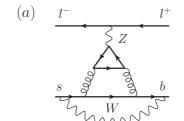


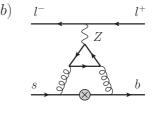
Subtleties: (i) counterterms with finite parts $\sim \bar{b}_L \not\!\! D s_L$

(ii) evanescent operators: $E_B = (\bar{b}\gamma_{\nu}\gamma_{\rho}\gamma_{\sigma}\gamma_{5}s)(\bar{\mu}\gamma^{\sigma}\gamma^{\rho}\gamma^{\nu}\gamma_{5}\mu) - 4(\bar{b}\gamma_{\alpha}\gamma_{5}s)(\bar{\mu}\gamma^{\alpha}\gamma_{5}\mu)$

 $E_T = {
m Tr}\, (\gamma^
u \gamma^
ho \gamma^\sigma \gamma^lpha \gamma_5) (ar b \gamma_
u \gamma_
ho \gamma_\sigma s) (ar \mu \gamma_lpha \gamma_5 \mu) + 24 (ar b \gamma_lpha \gamma_5 s) (ar \mu \gamma^lpha \gamma_5 \mu)$







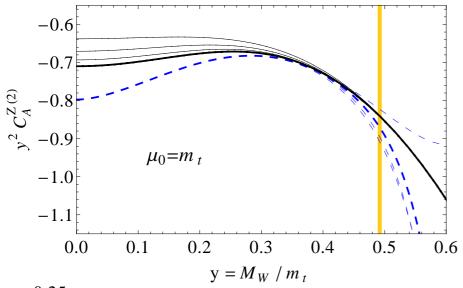
Renormalization of E_B

Diagrams generating E_T

Perturbative series for the Wilson coefficient at $\mu = \mu_0 \sim m_t, M_W$:

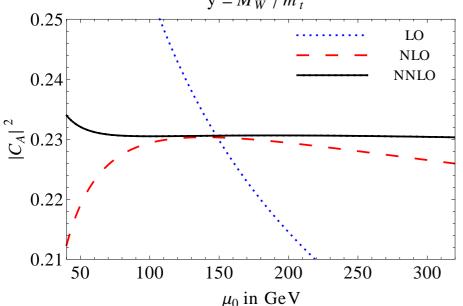
$$C_A(\mu_0) = C_A^{(0)}(\mu_0) + rac{lpha_s}{4\pi} C_A^{(1)}(\mu_0) + \left(rac{lpha_s}{4\pi}
ight)^2 C_A^{(2)}(\mu_0) + rac{lpha_{em}}{4\pi} \Delta_{
m EW} C_A(\mu_0) + \ldots$$

The top quark mass is $\overline{\text{MS}}$ -renormalized at μ_0 with respect to QCD, and on shell with respect to the EW interactions. Both α_s and α_{em} are $\overline{\text{MS}}$ -renormalized at μ_0 in the effective theory.



$$C_A^{(n)} = C_A^{W,(n)} + C_A^{Z,(n)}$$

To deal with single-scale tadpole integrals, we expand around y=1 (solid lines) and around y=0 (dashed lines), where $y=M_W/m_t$. The expansions reach $(1-y^2)^{16}$ and y^{12} , respectively. The blue band indicates the physical region.



Matching scale dependence of $|C_A|^2$ gets significantly reduced. The plot corresponds to $\Delta_{\rm EW}C_A(\mu_0)=0$. However, with our conventions for m_t and the global normalization, μ_0 -dependence is due to QCD only.

NNLO fit (with $\Delta_{\text{EW}}C_A(\mu_0) = 0$):

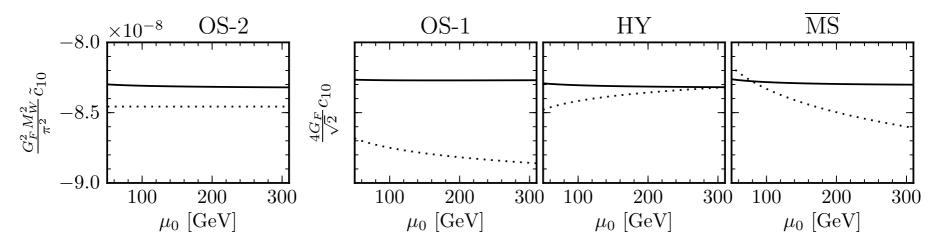
$$C_A = 0.4802 \left(rac{M_t}{173.1}
ight)^{1.52} \left(rac{lpha_s(M_Z)}{0.1184}
ight)^{-0.09} + \mathcal{O}(lpha_{em})$$

Evaluation of the NLO EW matching corrections in the SM

[C. Bobeth, M. Gorbahn, E. Stamou, Phys. Rev. D 89 (2014) 034023]

Method: similar to the NNLO QCD case. Two-loop integrals with three mass scales are present.

Dependence of the final result on μ_0 in various renormalization schemes (dotted – LO, solid – NLO):



In all the four plots: no QCD corrections to C_A included, $m_t(m_t)$ w.r.t. QCD used.

OS-2 scheme: Global normalization factor in \mathcal{L}_{eff} set to $N = V_{tb}^* V_{ts} \, G_F^2 M_W^2 / \pi^2$

Masses at the LO renormalized on-shell w.r.t. EW interactions (including M_W in N)

Plotted quantity: $-2C_A G_F^2 M_W^2/\pi^2$ in GeV⁻²

NLO EW matching correction to the BR: -3.7%

other schemes: Global normalization factor in \mathcal{L}_{eff} set to $4V_{tb}^*V_{ts}\,G_F/\sqrt{2}$

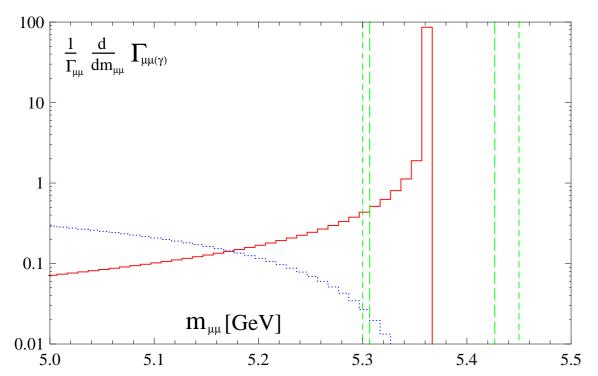
At the LO, $\alpha_{em}(\mu_0)$ used

 $\overline{\mathrm{MS}}$: Masses and $\sin^2 \theta_W$ renormalized at μ_0

OS-1: Masses as in OS-2, $\sin^2 \theta_W$ on-shell

HY (hybrid): Masses as in OS-2, $\sin^2 \theta_W$ as in $\overline{\text{MS}}$.

Radiative tail in the dimuon invariant mass spectrum



Ggreen vertical lines – experimental "blinded" windows [CMS and LHCb, Nature 522 (2015) 68]

Red line – no real photon and/or radiation only from the muons. It vanishes when $m_{\mu}
ightarrow 0$.

[A.J. Buras, J. Girrbach, D. Guadagnoli, G. Isidori, Eur.Phys.J. C72 (2012) 2172]

[S. Jadach, B.F.L. Ward, Z. Was, Phys.Rev. D63 (2001) 113009], Eq. (204) as in PHOTOS

Blue line – remainder due to radiation from the quarks. IR-safe because $m{B}_s$ is neutral.

Phase-space suppressed but survives in the $m_{\mu}
ightarrow 0$ limit.

[Y.G. Aditya, K.J. Healey, A.A. Petrov, Phys.Rev. D87 (2013) 074028]

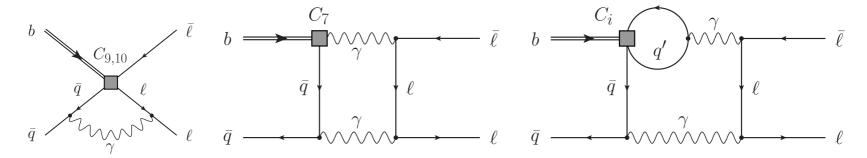
[D. Melikhov, N. Nikitin, Phys.Rev. D70 (2004) 114028]

Interference between the two contributions is negligible – suppressed both by phase-space and $m_{\mu}^2/M_{B_s}^2$.

The leading contribution to the decay rate is suppressed by $\frac{m_\ell^2}{M_{B_g}^2}$.

The leading contribution to the decay rate is suppressed by $\frac{m_{\ell}^2}{M_{B_q}^2}$.

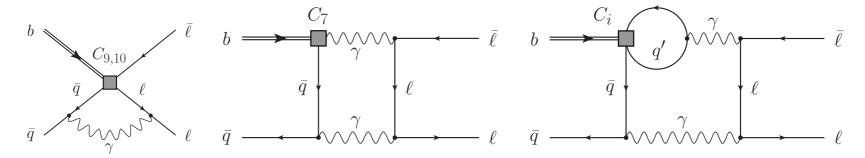
As observed by M. Beneke, C. Bobeth and R. Szafron in arXiv:1708.09152, some of the QED corrections receive suppression by $\frac{m_\ell^2}{\Lambda M_{B_g}}$ only:



See also the lecture by RS at the Paris-2019 workshop: https://indico.in2p3.fr/event/18845/sessions/12137/attachments/54326/71064/Szafron.pdf

The leading contribution to the decay rate is suppressed by $\frac{m_{\ell}^2}{M_{B_q}^2}$.

As observed by M. Beneke, C. Bobeth and R. Szafron in arXiv:1708.09152, some of the QED corrections receive suppression by $\frac{m_\ell^2}{\Lambda M_{Bq}}$ only:

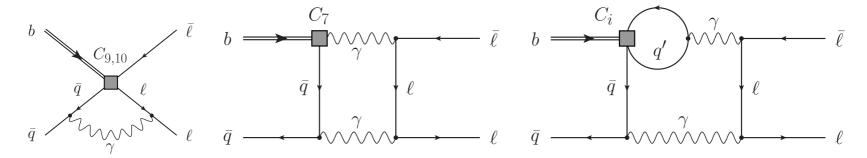


See also the lecture by RS at the Paris-2019 workshop: https://indico.in2p3.fr/event/18845/sessions/12137/attachments/54326/71064/Szafron.pdf

Consequently, the relative QED correction scales like $\frac{\alpha_{em}}{\pi} \frac{M_{Bq}}{\Lambda}$

The leading contribution to the decay rate is suppressed by $\frac{m_{\ell}^2}{M_{B_0}^2}$.

As observed by M. Beneke, C. Bobeth and R. Szafron in arXiv:1708.09152, some of the QED corrections receive suppression by $\frac{m_\ell^2}{\Lambda M_{Bq}}$ only:



See also the lecture by RS at the Paris-2019 workshop: https://indico.in2p3.fr/event/18845/sessions/12137/attachments/54326/71064/Szafron.pdf

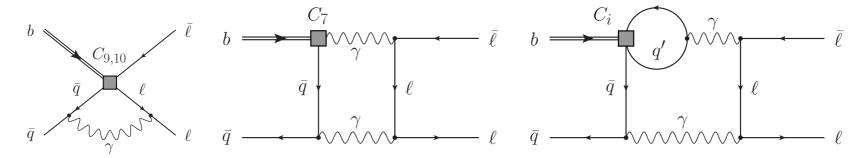
Consequently, the relative QED correction scales like $\frac{\alpha_{em}}{\pi} \frac{M_{Bq}}{\Lambda}$.

Their explicit calculation implies that the previous results for all the $B_q \to \ell^+ \ell^-$ branching ratios need to be multiplied by

$$\eta_{
m QED} = 0.993 \pm 0.004.$$

The leading contribution to the decay rate is suppressed by $\frac{m_{\ell}^2}{M_{B_q}^2}$.

As observed by M. Beneke, C. Bobeth and R. Szafron in arXiv:1708.09152, some of the QED corrections receive suppression by $\frac{m_\ell^2}{\Lambda M_{Bq}}$ only:



See also the lecture by RS at the Paris-2019 workshop: https://indico.in2p3.fr/event/18845/sessions/12137/attachments/54326/71064/Szafron.pdf

Consequently, the relative QED correction scales like $\frac{\alpha_{em}}{\pi} \frac{M_{Bq}}{\Lambda}$.

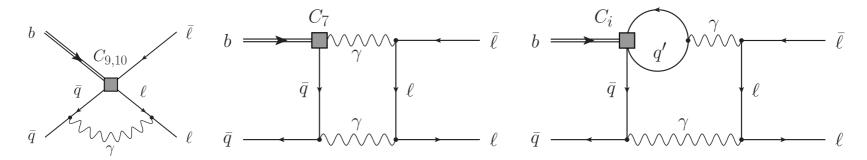
Their explicit calculation implies that the previous results for all the $B_q \to \ell^+ \ell^-$ branching ratios need to be multiplied by

$$\eta_{
m QED} = 0.993 \pm 0.004.$$

Thus, despite the $\frac{M_{Bq}}{\Lambda}$ -enhancement, the effect is well within the previously estimated $\pm 1.5\%$ non-parametric uncertainty.

The leading contribution to the decay rate is suppressed by $\frac{m_{\ell}^2}{M_{B_q}^2}$.

As observed by M. Beneke, C. Bobeth and R. Szafron in arXiv:1708.09152, some of the QED corrections receive suppression by $\frac{m_\ell^2}{\Lambda M_{Ba}}$ only:



See also the lecture by RS at the Paris-2019 workshop: https://indico.in2p3.fr/event/18845/sessions/12137/attachments/54326/71064/Szafron.pdf

Consequently, the relative QED correction scales like $\frac{\alpha_{em}}{\pi} \frac{M_{Bq}}{\Lambda}$.

Their explicit calculation implies that the previous results for all the $B_q \to \ell^+ \ell^-$ branching ratios need to be multiplied by

$$\eta_{
m QED} = 0.993 \pm 0.004.$$

Thus, despite the $\frac{M_{Bq}}{\Lambda}$ -enhancement, the effect is well within the previously estimated $\pm 1.5\%$ non-parametric uncertainty.

However, it is larger than $\pm 0.3\%$ due to scale-variation of the Wilson coefficient $C_A(\mu_b)$.

SM predictions for all the branching ratios $\overline{\mathcal{B}}_{q\ell} \equiv \overline{\mathcal{B}}(B_q^0 \to \ell^+\ell^-)$ including 2-loop electroweak and 3-loop QCD matching at $\mu_0 \sim m_t$

[C. Bobeth, M. Gorbahn, T. Hermann, MM, E. Stamou, M. Steinhauser, PRL 112 (2014) 101801]

$$egin{aligned} \overline{\mathcal{B}}_{se} imes 10^{14} &= m{\eta}_{ ext{QED}}(8.54 \pm 0.13) \, R_{tlpha} \, R_s, \ \overline{\mathcal{B}}_{s\mu} imes 10^9 &= m{\eta}_{ ext{QED}}(3.65 \pm 0.06) \, R_{tlpha} \, R_s, \ \overline{\mathcal{B}}_{s au} imes 10^7 &= m{\eta}_{ ext{QED}}(7.73 \pm 0.12) \, R_{tlpha} \, R_s, \ \overline{\mathcal{B}}_{de} imes 10^{15} &= m{\eta}_{ ext{QED}}(2.48 \pm 0.04) \, R_{tlpha} \, R_d, \ \overline{\mathcal{B}}_{d\mu} imes 10^{10} &= m{\eta}_{ ext{QED}}(1.06 \pm 0.02) \, R_{tlpha} \, R_d, \ \overline{\mathcal{B}}_{d au} imes 10^8 &= m{\eta}_{ ext{QED}}(2.22 \pm 0.04) \, R_{tlpha} \, R_d, \end{aligned}$$

where

$$egin{aligned} R_{tlpha} &= \left(rac{M_t}{173.1~{
m GeV}}
ight)^{3.06} \left(rac{lpha_s(M_Z)}{0.1184}
ight)^{-0.18}, \ R_s &= \left(rac{f_{B_s}[{
m MeV}]}{227.7}
ight)^2 \left(rac{|V_{cb}|}{0.0424}
ight)^2 \left(rac{|V_{tb}^{\star}V_{ts}/V_{cb}|}{0.980}
ight)^2 rac{ au_H^s~[{
m ps}]}{1.615}, \ R_d &= \left(rac{f_{B_d}[{
m MeV}]}{190.5}
ight)^2 \left(rac{|V_{tb}^{\star}V_{td}|}{0.0088}
ight)^2 rac{ au_d^{
m av}~[{
m ps}]}{1.519}. \end{aligned}$$

Update of the input parameters

	2014 paper	this talk	source
$M_t [{ m GeV}]$	173.1(9)	172.69(30)	PDG 2022, http://pdglive.lbl.gov
$lpha_s(M_Z)$	0.1184(7)	0.1179(9)	arXiv:2203.08271, Eq.(9.2)
$f_{B_s} [{ m GeV}]$	0.2277(45)	0.2303(13)	FLAG, arXiv:2111.09849
$f_{B_d} [{ m GeV}]$	0.1905(42)	0.1900(13)	FLAG, arXiv:2111.09849
$ V_{cb} imes 10^3$	42.40(90)	42.16(50)	inclusive, arXiv:2107.00604
$ V_{tb}^*V_{ts} / V_{cb} $	0.9800(10)	0.9819(5)	derived from CKMfitter 2019, http://ckmfitter.in2p3.fr
$ V_{tb}^*V_{td} imes 10^4$	88(3)	$87.1^{+0.86}_{-2.46}$	CKMfitter 2019, http://ckmfitter.in2p3.fr
$ au_H^s [ext{ps}]$	1.615(21)	1.616(10)	HFLAV 2022, https://hflav.web.cern.ch
$ au_H^d [ext{ps}]$	1.519(7)	1.519(4)	HFLAV 2022, https://hflav.web.cern.ch
$\overline{\mathcal{B}}_{s\mu} imes 10^9$	3.65(23)	3.66(12)	
$\overline{\mathcal{B}}_{d\mu} imes 10^{10}$	1.06(9)	$1.02^{+0.03}_{-0.06}$	

Sources of uncertainties	f_{B_q}	CKM	τ_H^q	M_t	$lpha_s$	other parametric	non- parametric	Σ
$\overline{\mathcal{B}}_{s\ell}$	1.1%	2.4%	0.6%	0.5%	0.2%	< 0.1%	1.5%	3.2%
$\overline{\mathcal{B}}_{d\ell}$	1.4%	$\binom{+2.0}{-5.6}\%$	0.3%	0.5%	0.2%	< 0.1%	1.5%	$\binom{+3.1}{-6.2}\%$

Summary

- Uncertainties in the SM predictions for $\overline{\mathcal{B}}_{q\ell}$ are dominated by the parametric ones, mainly due to the decay constants and CKM factors.
- In the $\overline{\mathcal{B}}_{s\ell}$ case, resolving the inclusive-exclusive tension in $|V_{cb}|$ would help a lot.
- The central values of the SM predictions for $\overline{\mathcal{B}}_{s\mu}$ and $\overline{\mathcal{B}}_{d\mu}$ are in good agreement with the data from LHCb, CMS and ATLAS.
- Some of the QED corrections involve non-perturbative physics beyond what is contained in the decay constants. Despite the powerlike enhancement factors in such corrections, the non-parametric uncertainty can be retained at the $\pm 1.5\%$ level.