

Development and characterization of novel electronics for the search of dark matter for DAMIC-M

Giorgos PAPAPOULOS

DAMIC-M

LPNHE, Sorbonne University, Paris

Supervised by: Antoine Letessier-Selvon

for the ELBERETH Conference 2020



27 February 2020

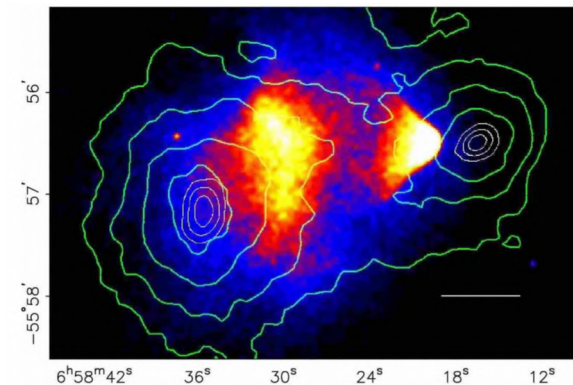
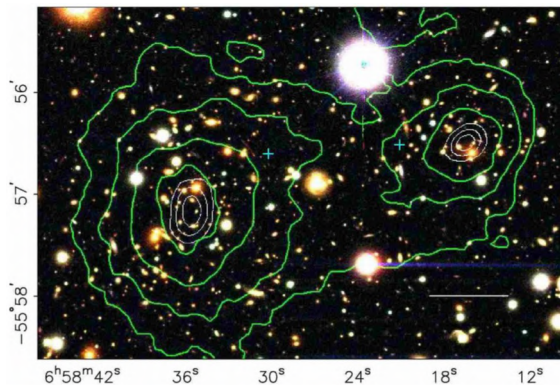
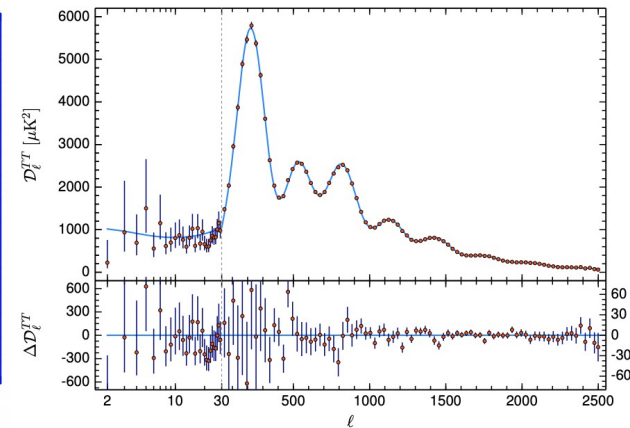
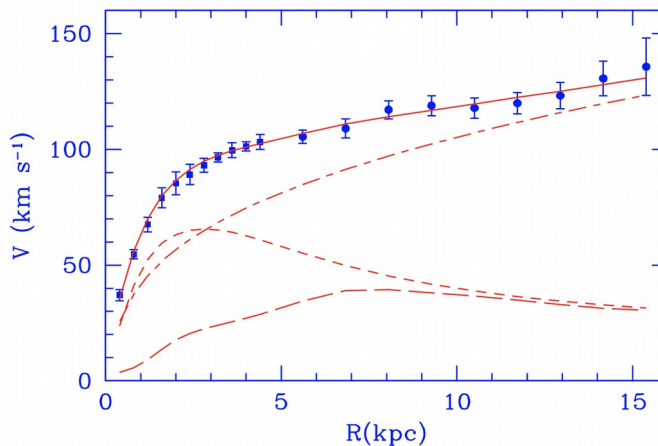
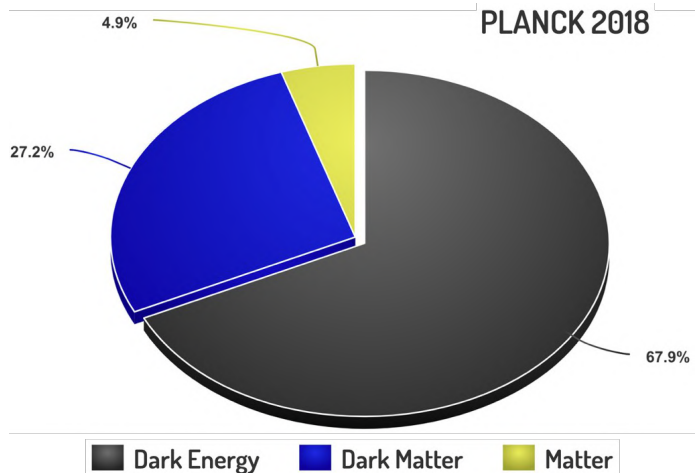
Dark Matter (DM) motivation



Evidence:

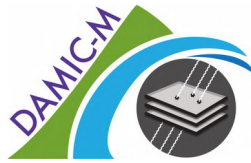
- Galaxy rotation curve [1]
- CMB power spectrum [2]
- Weak field lensing of colliding galaxy clusters [3]
- Mass to luminosity ratio[4]

Up to date status:

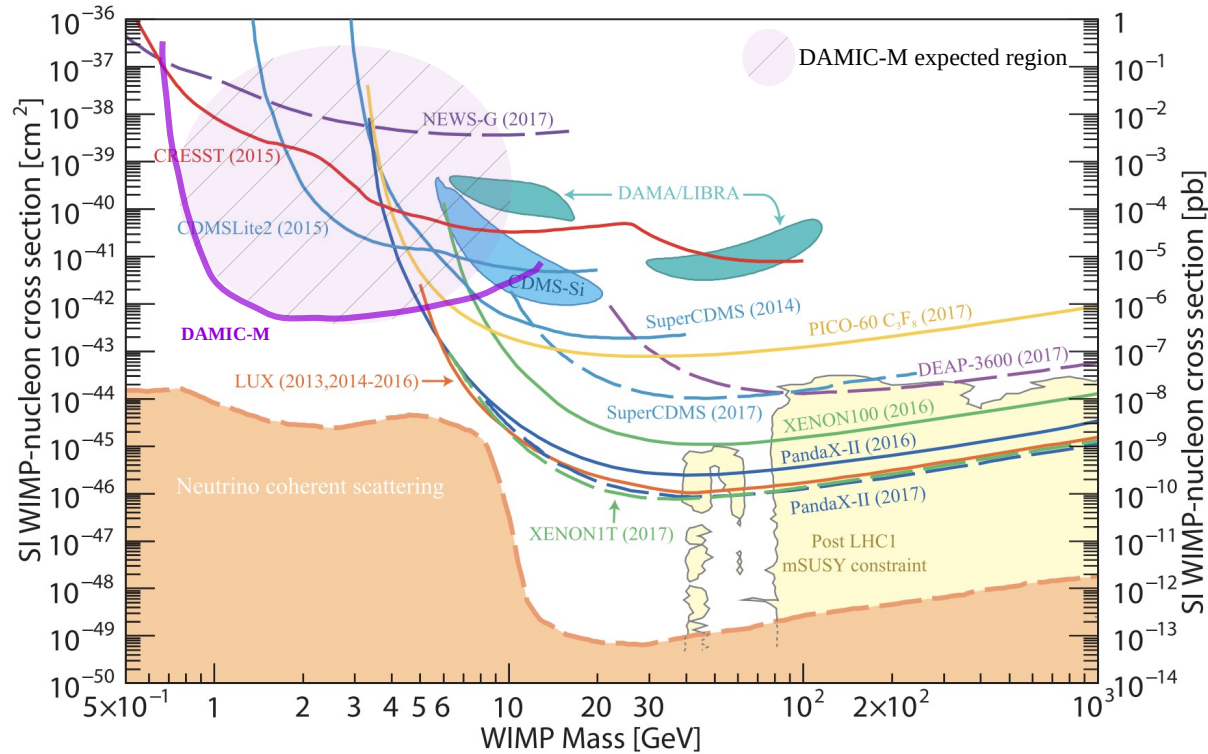


DM current status

DAMIC : *D*ARK *M*atter In *C*CDs



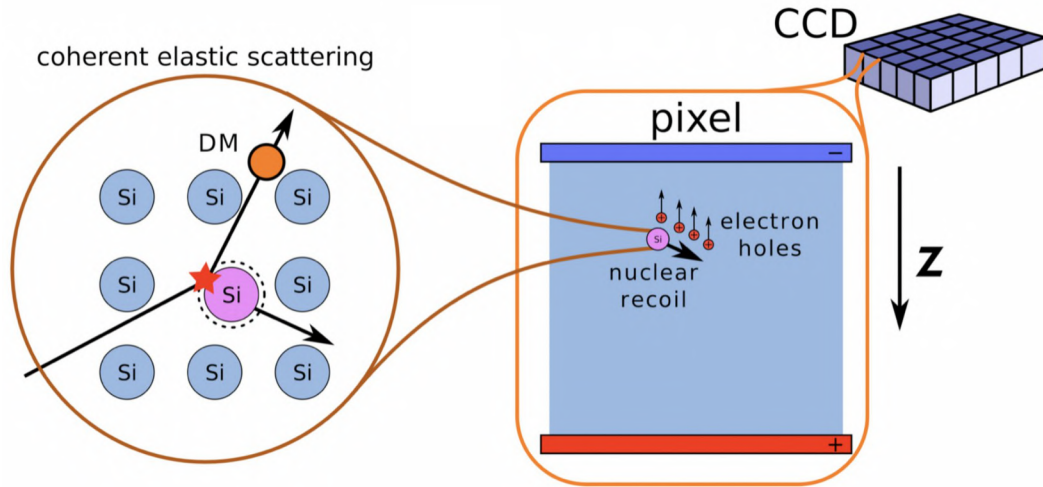
DAMIC-M: search for light WIMPs and hidden sector DM with mass $< 10\text{GeV}$.



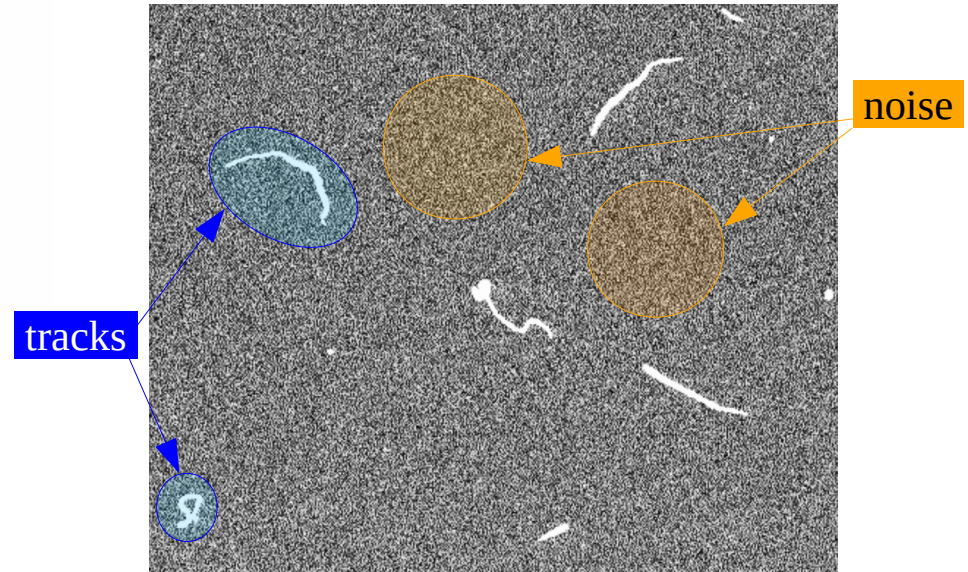
WIMP-nucleon cross sections for Spin-Independent coupling versus mass [5,6]

Charge Coupled Device (CCD)

DAMIC : *D*Ark Matter *In* CCDs

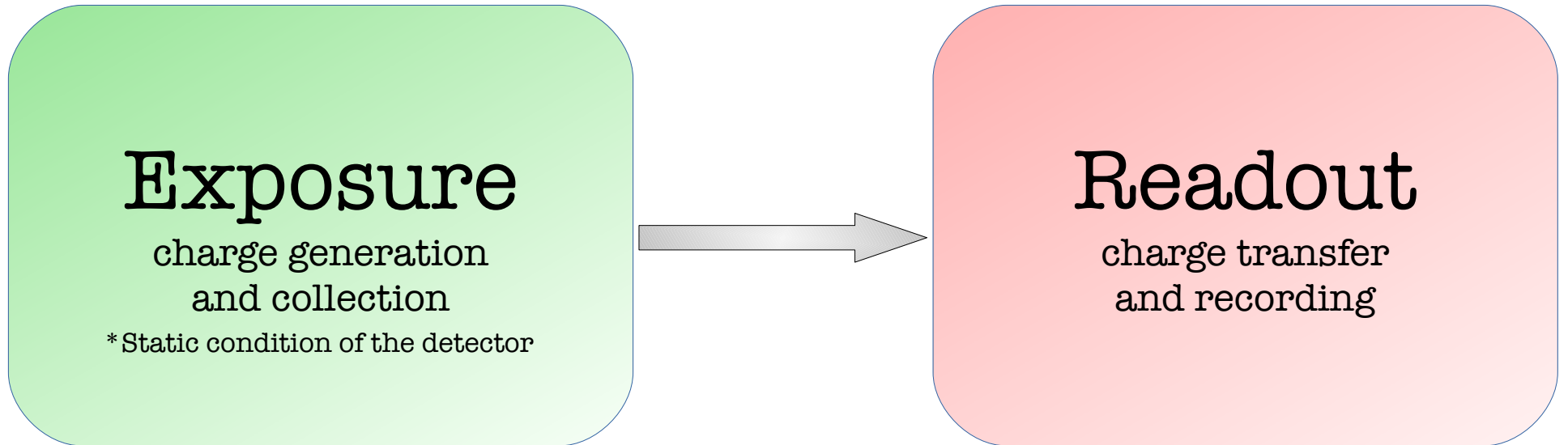


- Direct detection: $\chi + \text{SM} \rightarrow \chi + \text{SM}$
- CCD is made of silicon, separated in pixels
- Incident particles deposit energy in the bulk

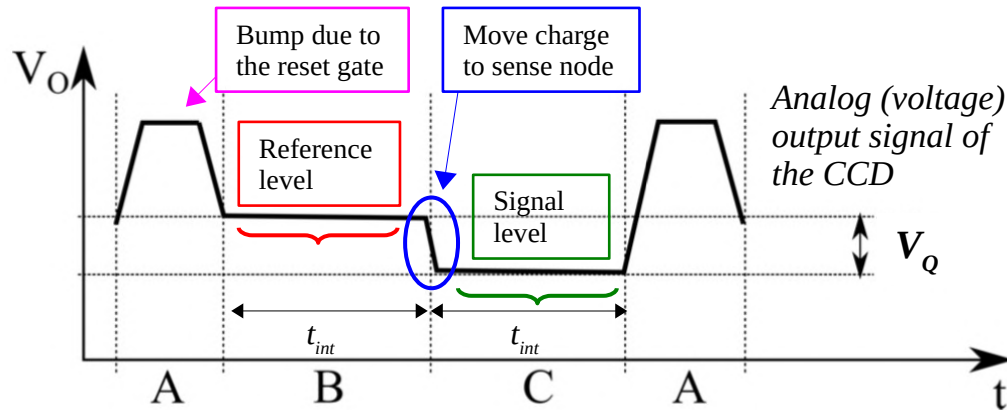
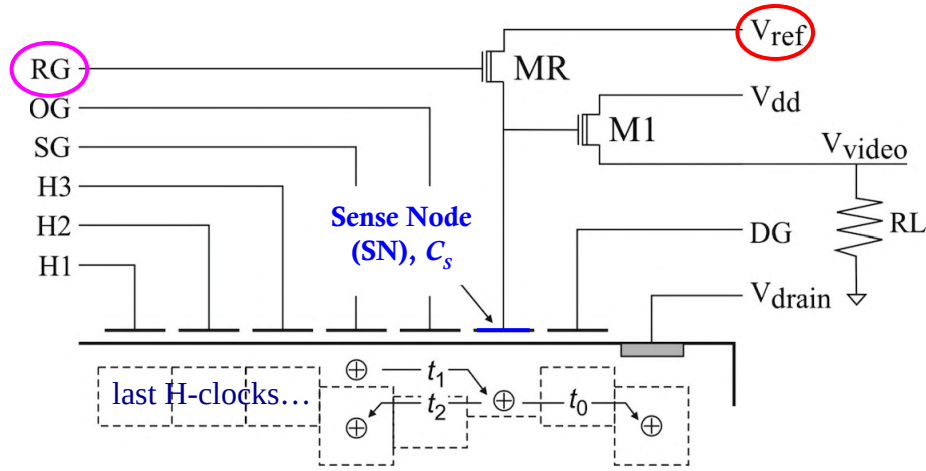


Real CCD image on surface level

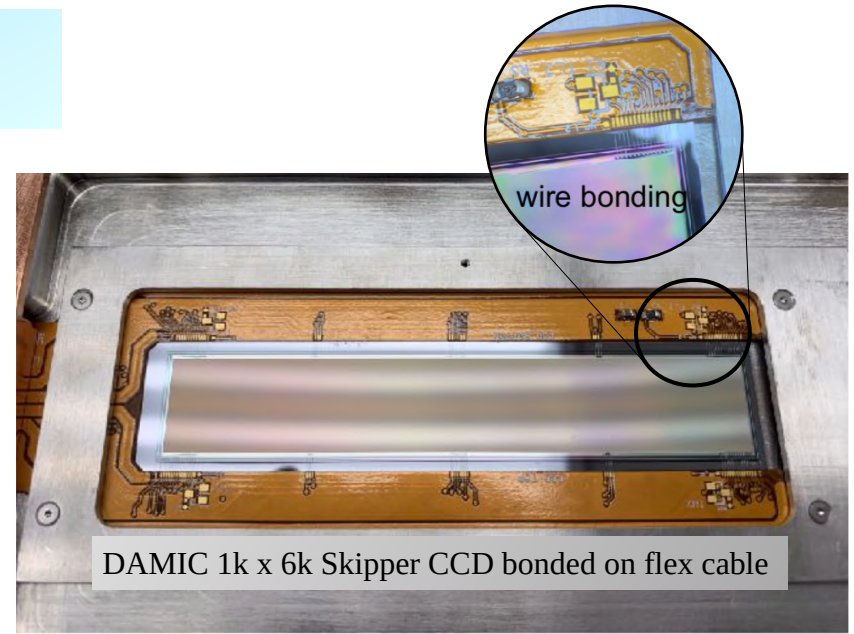
CCD Operation



CCD readout circuit



Giorgos PAPAPOPOULOS



DAMIC 1k x 6k Skipper CCD bonded on flex cable

Reaching the end of the CCD, there is a circuit to convert the pixel charge into voltage.

- The Reset Gate sets the Sense Node at a voltage reference value. This will cause a bump in the output signal which will decay quickly, resulting in an outcome reference level around V_{ref} .
- The charge Q is injected to the SN changing the voltage by $V_Q = Q / C_s$, where C_s is the capacity of the sense node which is known.
- Measure and subtract the reference and signal levels to find the V_Q .

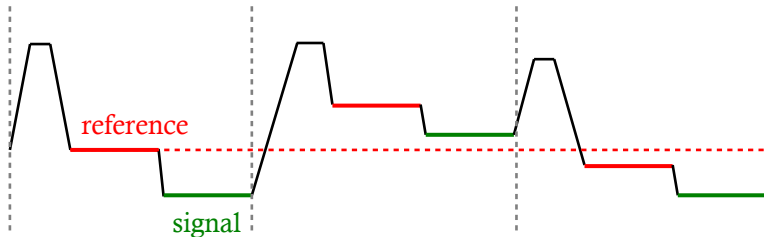
CCD noise sources

Dark current

- Thermally generated electrons in the bulk of the CCD
- Linear dependence on time → limits the exposure duration. The longer the exposure, the worse the Signal to Noise Ratio.
- Lower the temperature (~100-140K) to decrease the dark current

Reset or kTC noise

- After reset pulse thermal noise is generated by the resistance of the reset FET. The small capacitance of the Sense Node \sim fF leads to a significant uncertainty of the reference level.
- *Correlated Double Sampling* (CDS): measure both the reference and signal levels and subtract them to eliminate the reset noise.

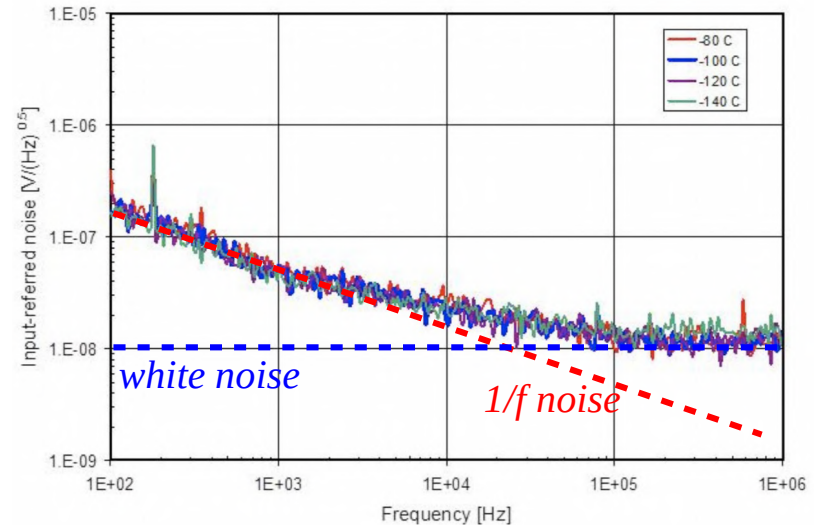


White noise

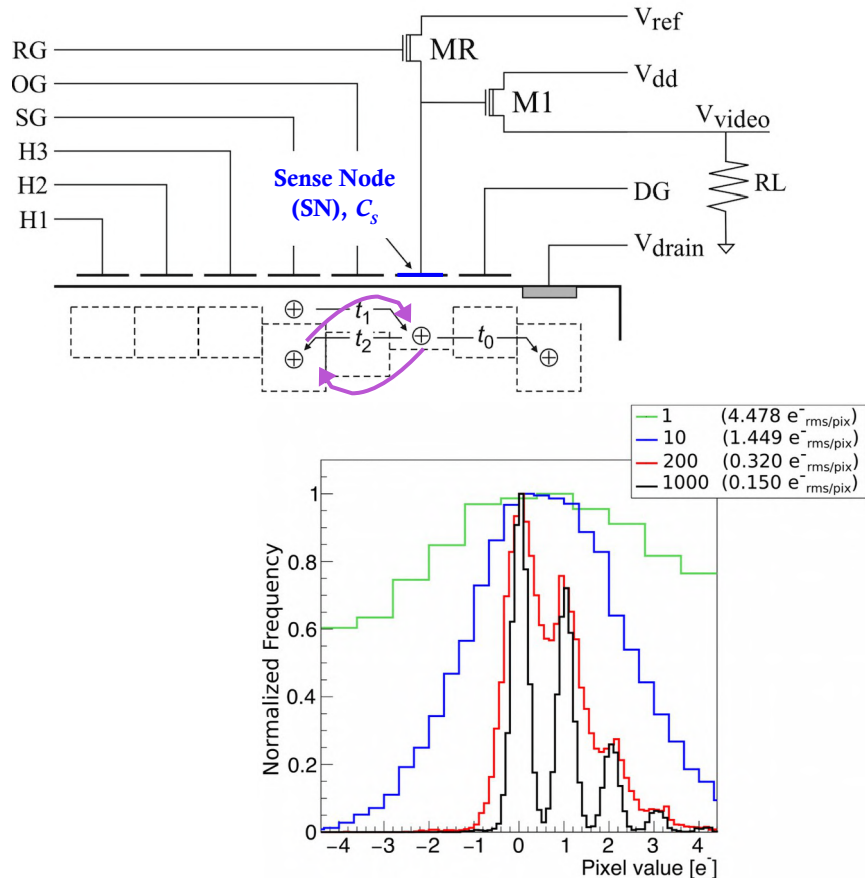
Thermal noise generated by the output amplifier MOSFET.

Flicker or $1/f$ noise

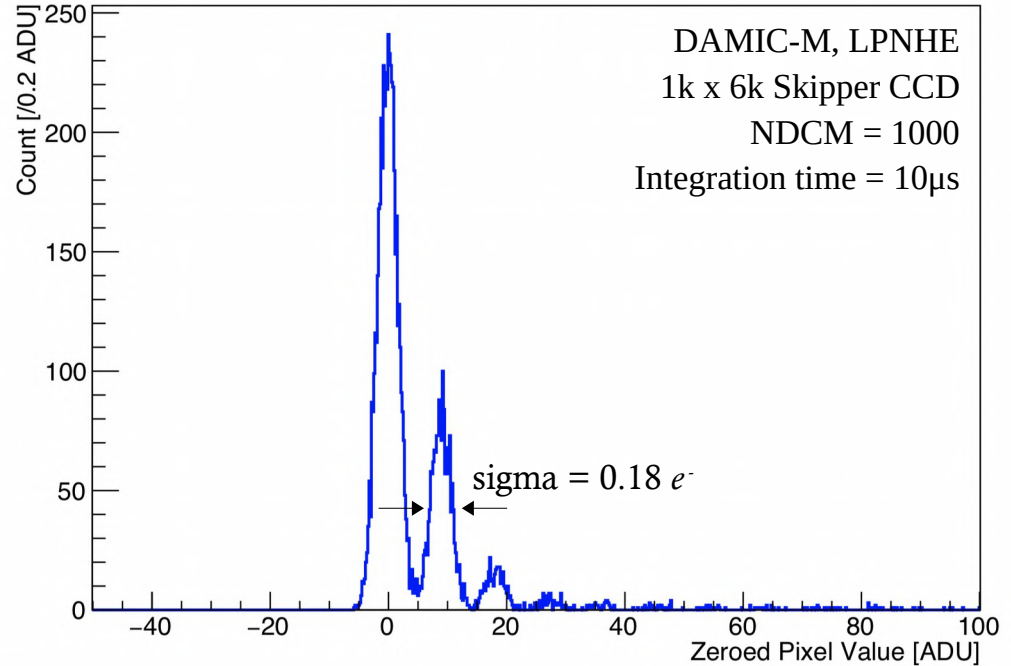
- The *flicker* noise depends approximately on the inverse of the sampling frequency.
- Dominant up to \sim 0.1MHz readout speed



Skipper CCD – sub- e^- resolution

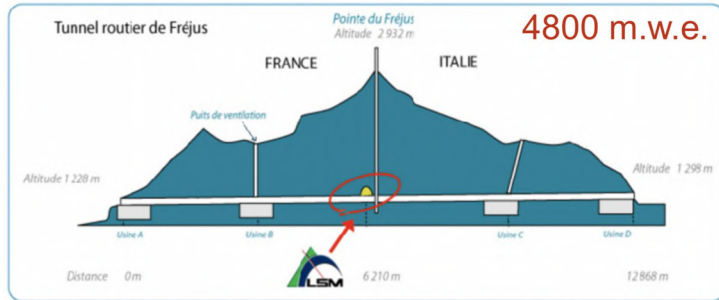
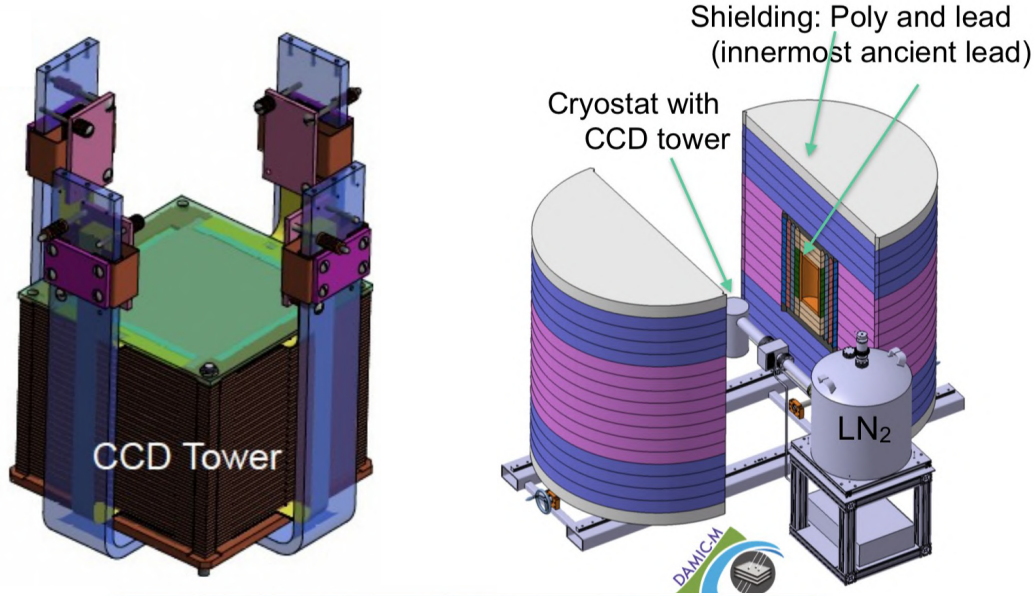


Single e^- resolution spectrum



The Skipper technique allows for a Non-Destructive Multiple pixel Charge Measurement (NDCM) which reduce the electronic noise which decreases as $1/\sqrt{\text{NDCM}}$.

DAMIC-M Overview



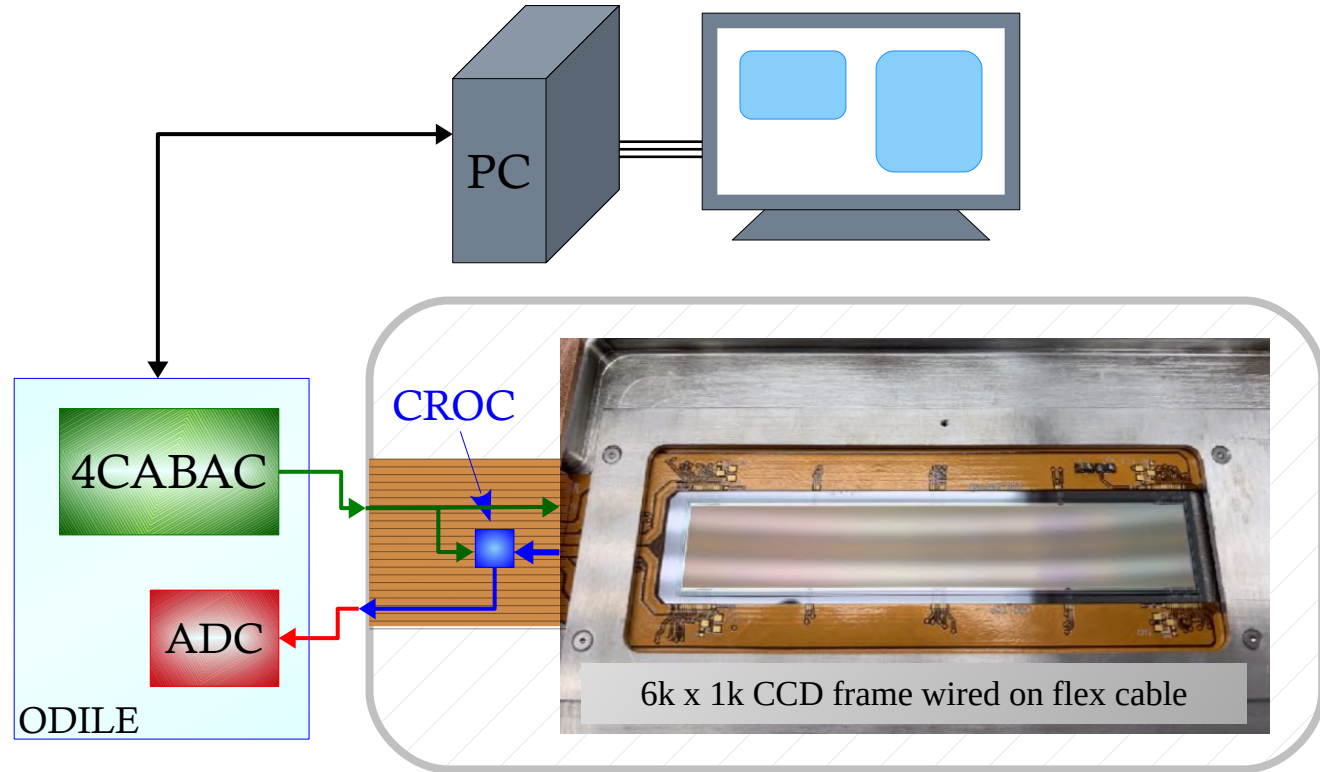
Dark Matter In CCDs at Modane (DAMIC-M):

- 50 scientific grade Charge Coupled Devices (CCDs) made of Si with Skipper readout
 - 36Mpix large and ~20g each, a total kg-size target mass
 - Sub-electron resolution
 - 0.1 event/(keV·kg·day) background
- R&D of the electronics to control the CCD
 - Simulations of the detector design and shielding
 - Underground Laboratory in Modane (LSM). ~2km of rock to stop cosmic background.
 - Low Background Chamber (2020) test detector to evaluate new CCDs, test electronics and measure background
 - Final experiment data taking will start in 2022.
 - DAMIC-M will pioneer in the low mass WIMPs and hidden-sector DM research.
 - ERC grant 2017

Full Setup & New Electronics

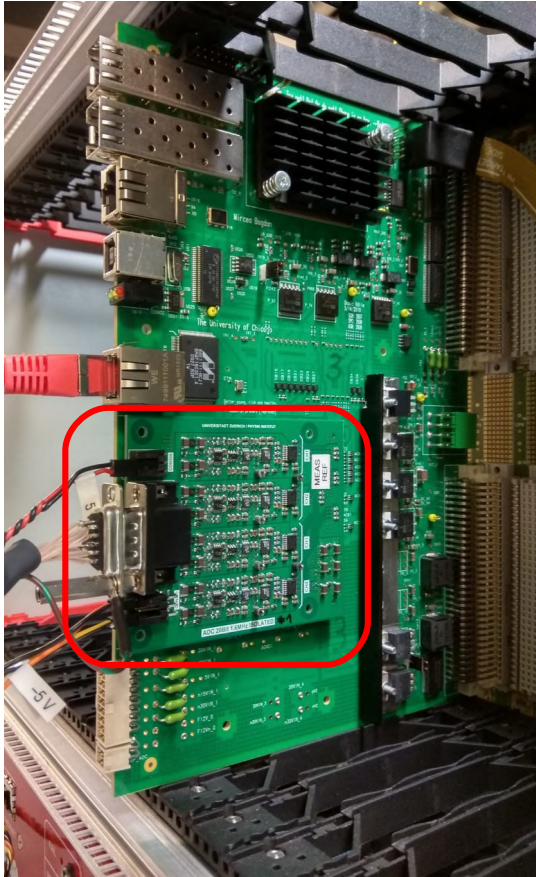
The CCD is placed in a cryostat and is controlled and readout by external electronics.

- All voltages and clocks during the expose and readout phases will be provided by the **Clocks And Biases ASIC for CCD (CABAC)** board.
- The **CCD ReadOut Chip (CROC)** will preprocess and amplify the signal to improve the Signal-to-Noise Ratio.
- An **Analog to Digital Converter (ADC)** will perform the transition from the analog domain to the digital one. The ADC can only apply the conversion in certain specified moments.
- Everything is going to be controlled by the FPGA, the brain of the, so called, **ODILE - Online Digital Interface for Low-noise Electronics** - motherboard

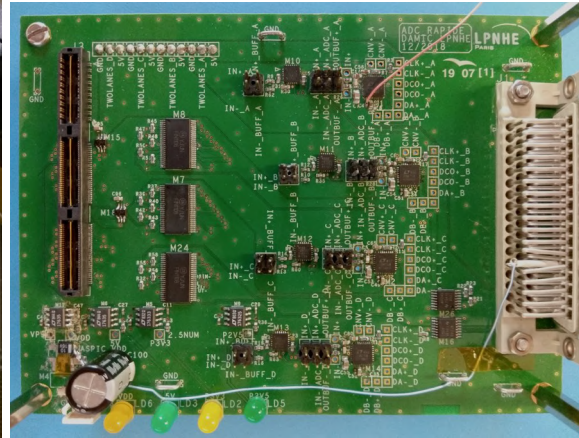


Setup of a single DAMIC-M CCD

New Electronics

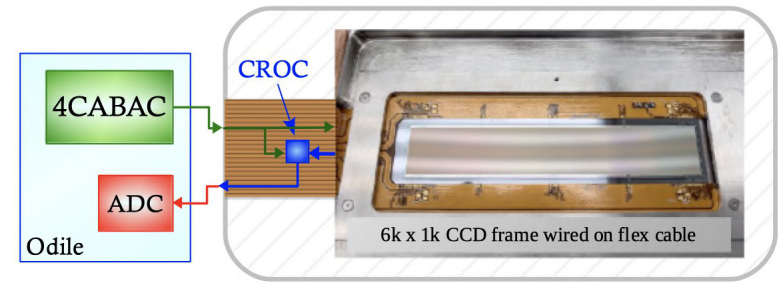
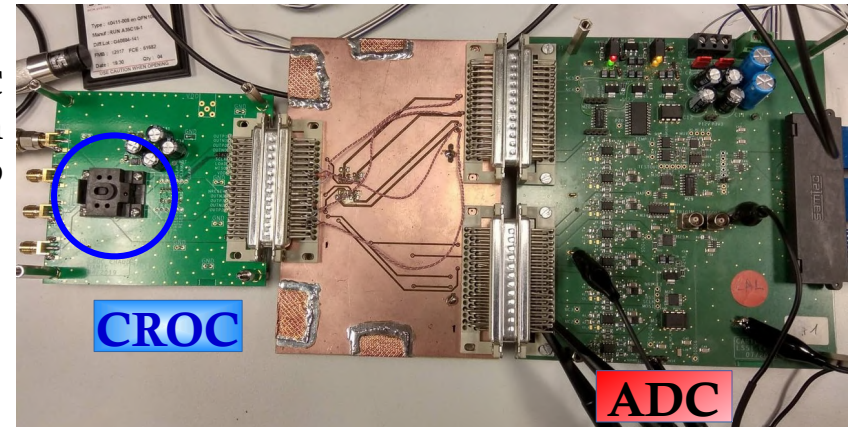


Giorgos PAPAPOPOULOS



ODILE board with
MAX11905 20-bit
1.6 MHz ADC
candidate

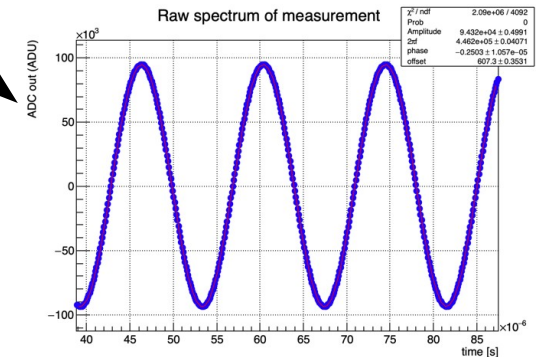
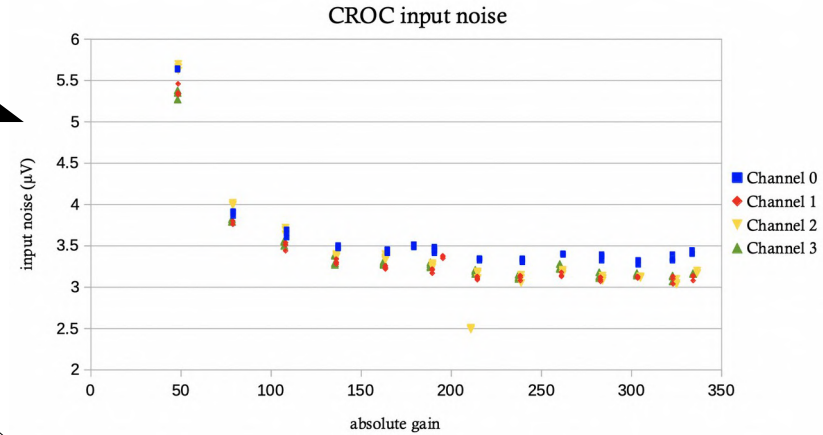
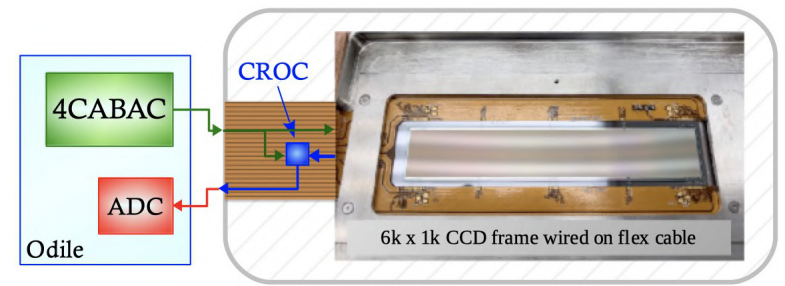
CROC
evaluation
setup



LTC2387-18 18-bit
15 MHz ADC
candidate prototype
board

Near Future

- The CROC chip is operational with low input noise at room temperature
- Test in low temperature → to be done
- Integrate in a setup for real CCD input signals → to be done
- A first version of the fast ADC was evaluated verifying the expected behaviour
- New design compatible with ODILE → in process
- 4CABAC board will be evaluated soon
- An ODILE board with a 20-bit ADC is functional and soon will be tested with a CCD.





References :

- [1] arXiv:astro-ph/9909252
- [2] arXiv:1507.02704
- [3] arXiv:astro-ph/0608407
- [4] *On the masses of nebulae and of clusters of nebulae*, F. Zwicky, 1937
- [5] *The Review of Particle Physics*, M. Tanabashi et al. (Particle Data Group), Phys. Rev. D 98, 030001 (2018)
- [6] *DAMIC-M Experiment: Thick, Silicon CCDs to search for Light Dark Matter*, N. Castello-Mor for the DAMIC-M Collaboration, 2020