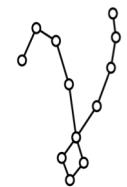


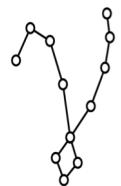
GammaLearn: Deep Learning applied to the CTA data analysis

Mikaël Jacquemont and Thomas Vuillaume
for the GammaLearn team

IN2P3/IRFU Machine Learning workshop, 22/01/2020



- Gamma astronomy and the CTA
- Why DL for CTA data analysis ?
- Deep Multitask Learning
- γ -PhysNet performances
- Going further



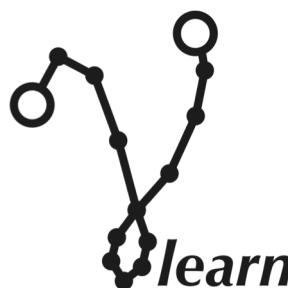
Mikaël Jacquemont
thesis



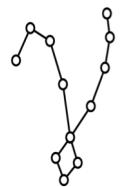
- Thomas Vuillaume
- Gilles Maurin
- Giovanni Lamanna



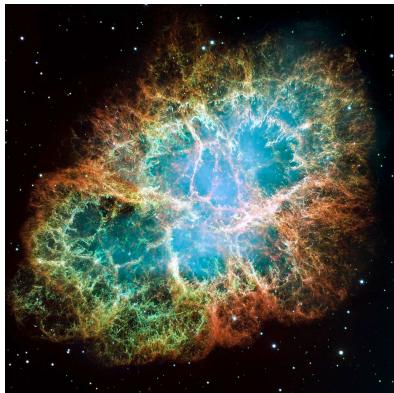
- Patrick Lambert
- Alexandre Benoit



*ASTERICS: european H2020 project



Supernova



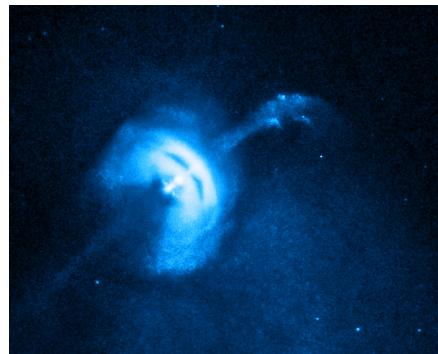
wikipedia

AGN



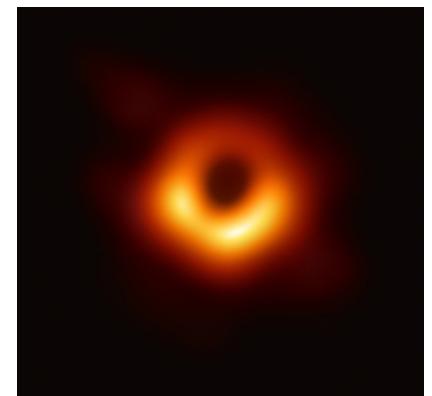
ESA/NASA

Pulsar



NASA

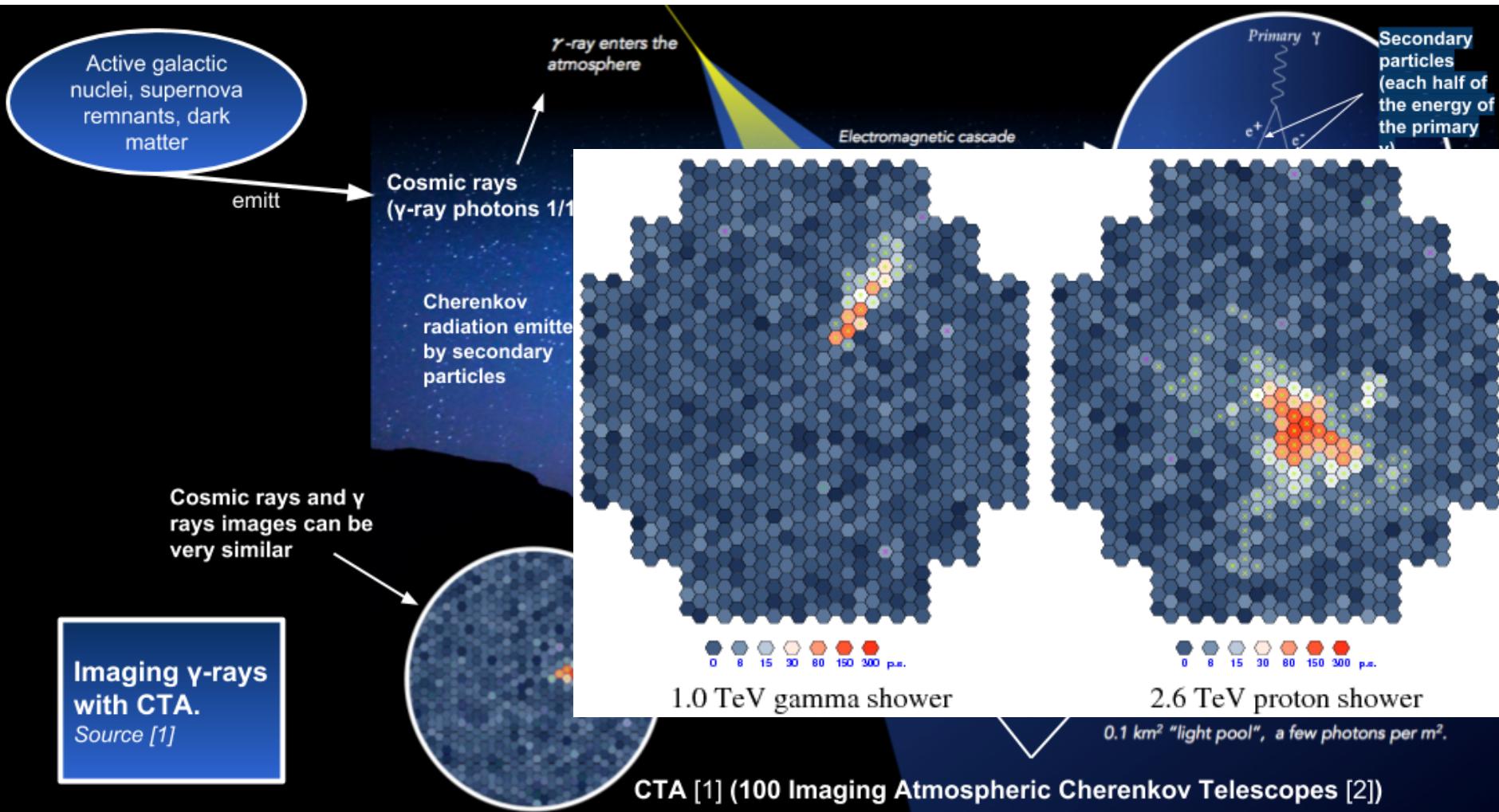
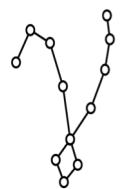
Black hole

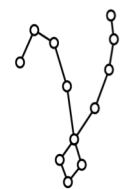


ESO

...

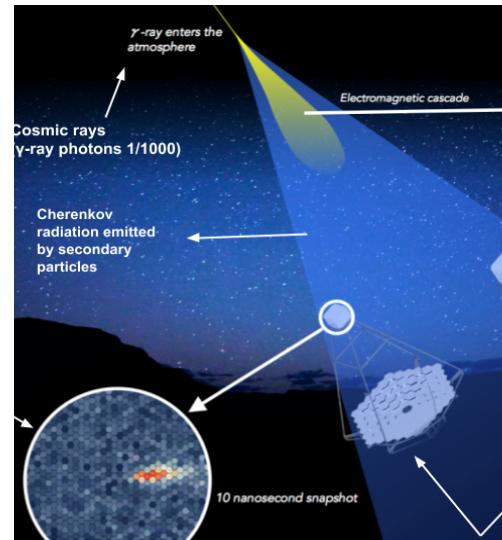
Gamma rays

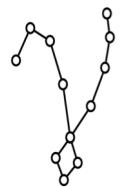




Data specificities

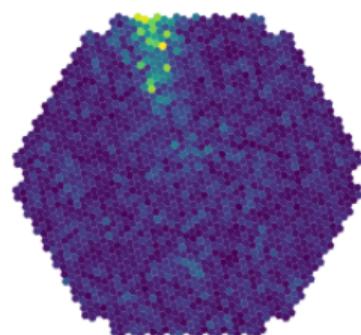
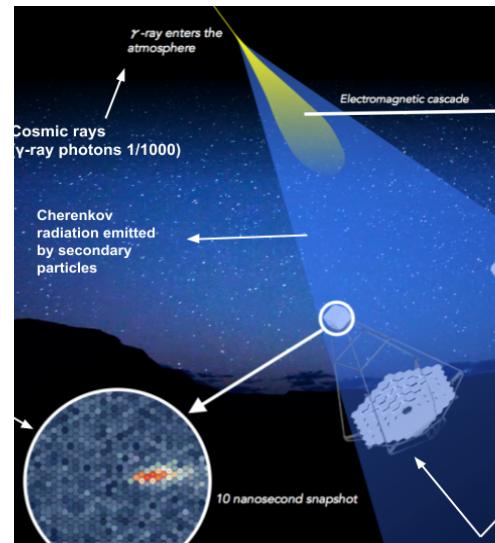
- Hexagonal pixels
- High SNR ($\sim 1/1000$)
- Sequence of images (some ns)
 - Calibration
 - Integration



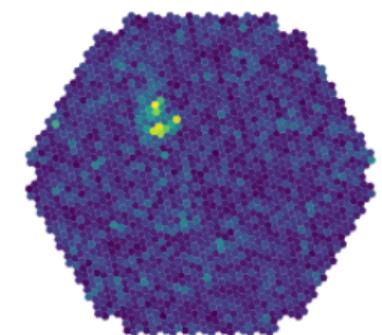


Strong interdependence

- Image shape and intensity
- Energy
- Arrival direction
- Impact point
- Particle type



Gamma 7,3 TeV



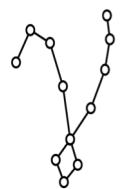
Proton 27 GeV



- Next generation gamma astronomy observatory
- 2 sites (La Palma and Paranal)
- 108 telescopes
- First prototype on site: LST1
- 210 PB per year !



Credit: Gabriel Pérez Diaz, IAC, SMM



CTA data analysis

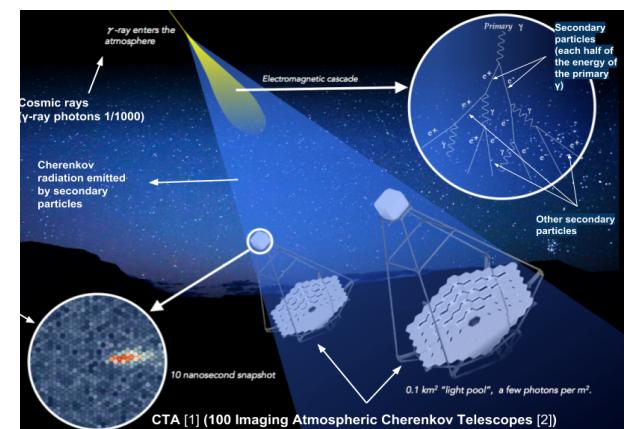
- Gamma / proton separation
- Energy and Direction reconstruction

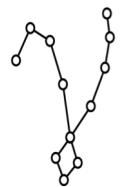
Standard methods

- Hillas + RF
- Impact / Model++

Why Deep Learning ?

- Potentially more performant
- Minimal images treatment before analysis
- Fast execution, thanks to GPUs





Convolution (and pooling): a key element of DL

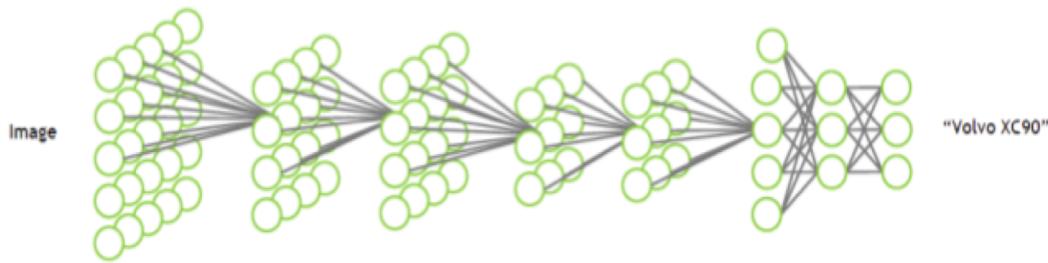
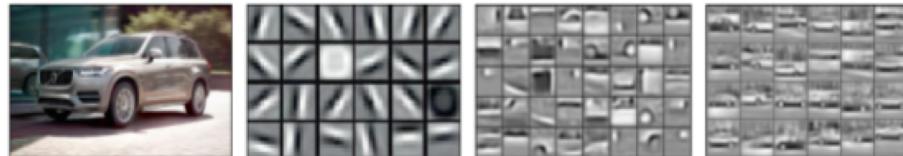
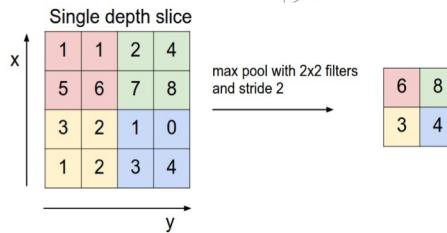
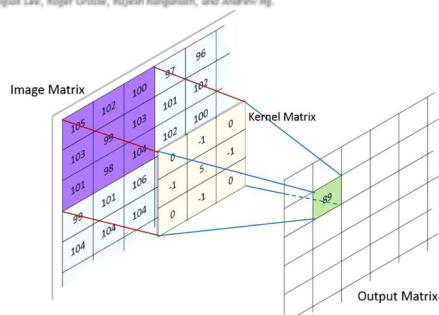


Image source: "Unsupervised Learning of Hierarchical Representations with Convolutional Deep Belief Networks" ICML 2009 & Comm. ACM 2011.
Hengki Lee, Roger Grosse, Rajesh Ranganath, and Andrew Ng.

12 NVIDIA



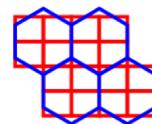
- Reducing the # of parameters
- (Locally) translation invariance
- Larger field of view
- Consideration of the context

Hexagonal pixels ?



Hexagonal pixels

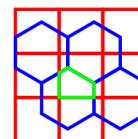
- Oversampling



(c) oversampling

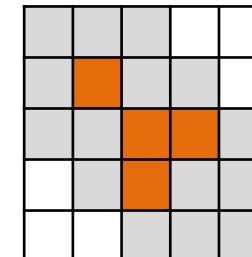
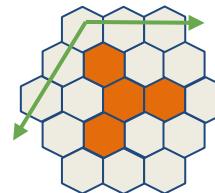
Credits: T. Holch et al.

- Rebinning



- Image shifting + masked convolution

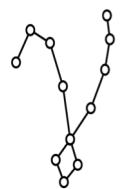
Axial
addressing
system



1	1	0
1	1	1
0	1	1

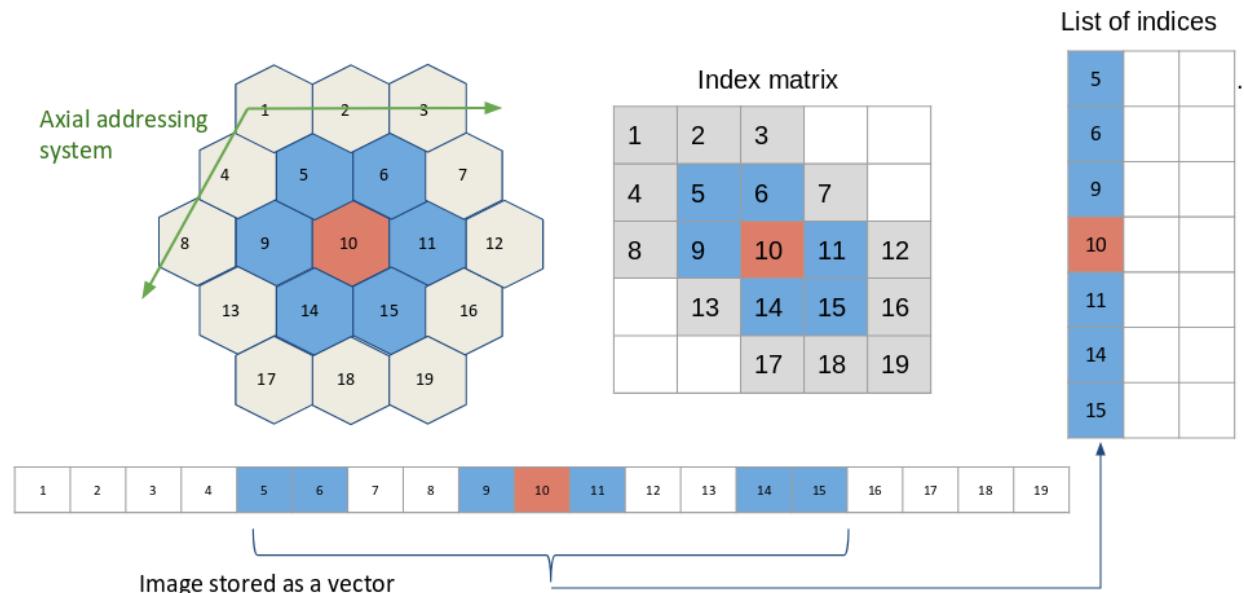
Convolution
kernel mask

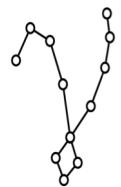
- Preprocessing
- Image distortion



IndexedConv

- GEMM for arbitrary lattices
 - im2col operation based on the list of pixel neighbors
- Hexagonal case





IndexedConv

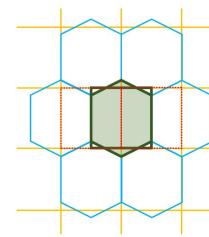
CIFAR-10

Hexagonal kernels (i.c.)	Square kernels
88.51 ± 0.21	88.39 ± 0.48



AID

Hexagonal kernels (i.c.)	Square kernels
79.81 ± 0.73	79.85 ± 0.50



"Indexed operations for non-rectangular lattices applied to convolutional neural networks"

VISAPP 2019



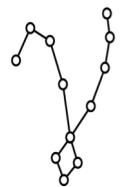
Objectives

- Gamma / proton separation
- Gamma parameters reconstruction: energy and arrival direction

Challenges

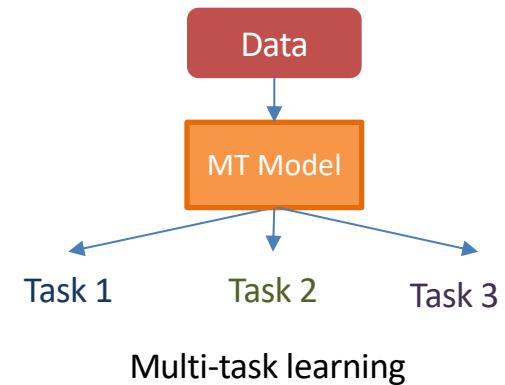
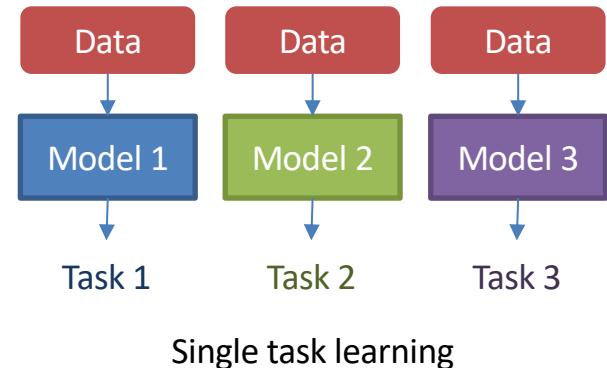
- Data volume
- Camera and pixels shapes
- Parameters interdependence

Multi-task problem



Deep Multi-task Learning

- Interests
 - Implicit data augmentation
 - Eavesdropping
 - Attention focusing
 - Representation bias
- Architecture
 - Hard parameter sharing
 - Soft parameter sharing
- Task balancing
 - Handcrafted
 - GradNorm
 - **Uncertainty estimation**
 - Multi-objective (Pareto)
 - Dynamic task prioritization



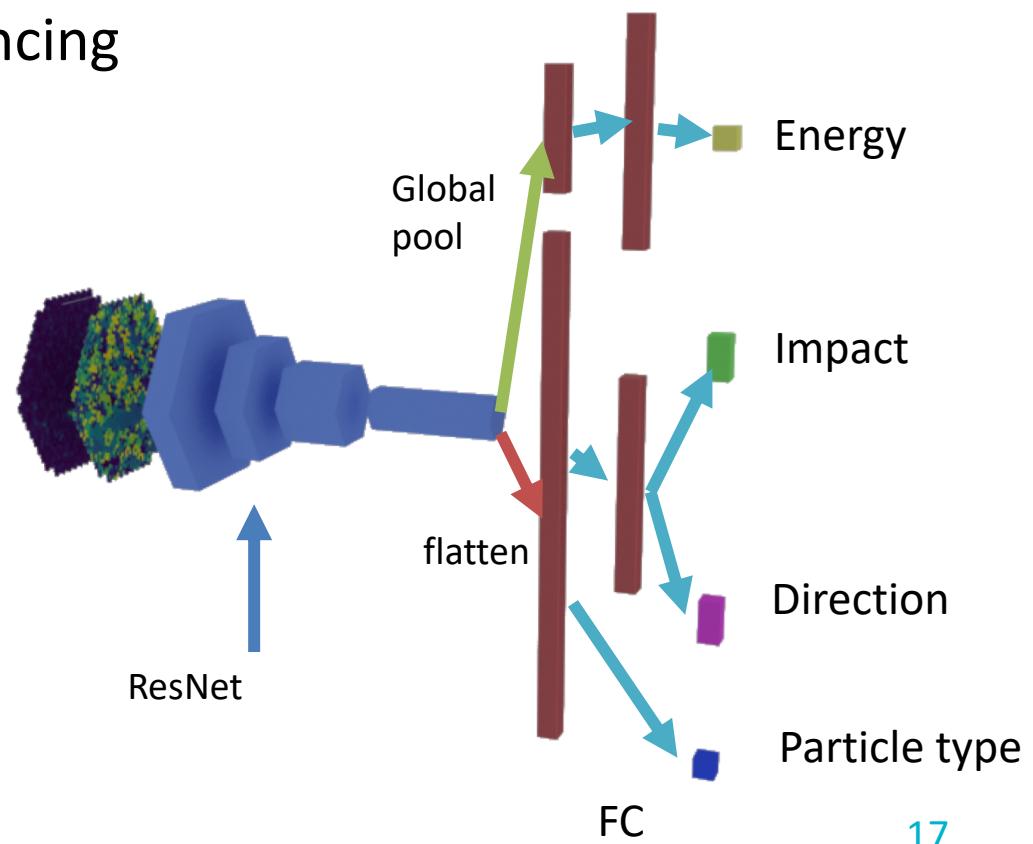


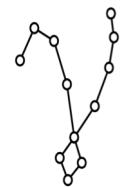
- Architectures for full event reconstruction
 - Gamma / proton separation
 - Energy
 - Direction
 - Impact point (auxiliary task)
- Focus on mono analysis for now
 - Interesting use-case
 - Preparation of the LST1 analysis
 - Aiming to apply on all telescope types
 - Incoming of real data
 - First step to the stereo analysis



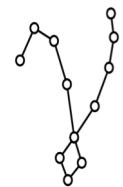
γ -PhysNet

- Physically guided architecture
- IndexedConv
- Uncertainty balancing
- Conditional loss
- Gradient penalty

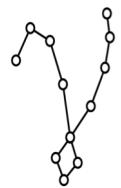




- Mono analysis with Deep MT Learning
- Deep Multi-task Learning is promising
- Very interesting results → consistent with template based analysis in H.E.S.S. but much ($\sim 1e3$) faster



- Optimize γ -PhysNet
- Use these good results in mono for the stereo reconstruction
- Test on real data → adaptation needed
- Publication

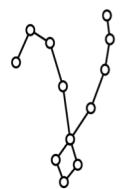


Thank you for your attention

GammaLearner framework: <https://gitlab.lapp.in2p3.fr/GammaLearn/GammaLearn>

IndexedConv package: <https://github.com/IndexedConv/IndexedConv>

Ctaphot and Gammaboard: <https://github.com/vuillaut/ctaphot>

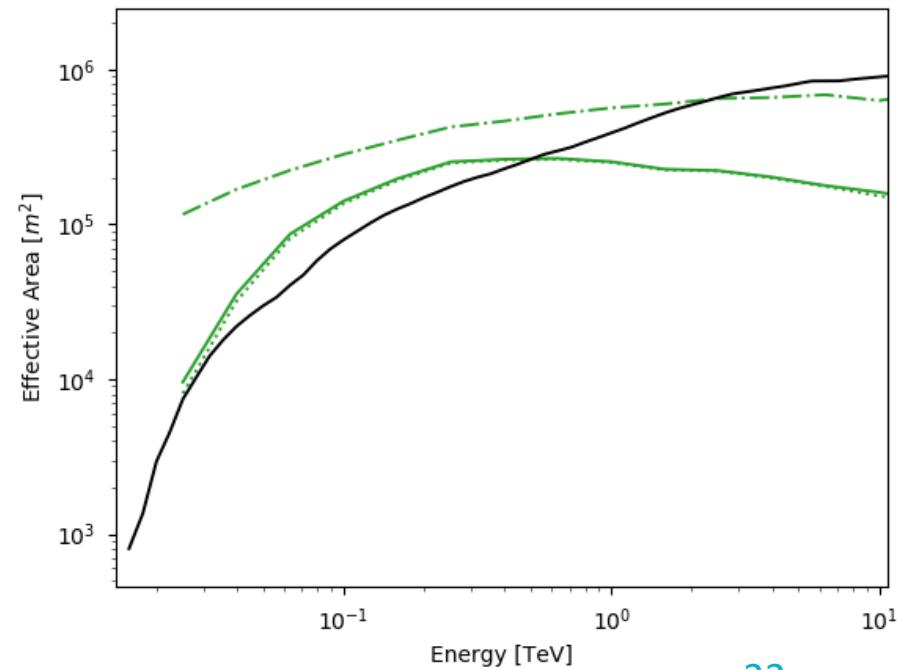


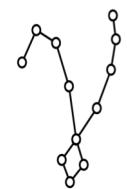
- Backup



Data

- LST4 monotrigger (gamma and proton events)
- Diffuse for training, point like for testing
- Calibrated and integrated images (625 k / 156 k)
- Mono analysis
- Cuts
 - Minimum amplitude 300
 - Leakage2 < 0,2
 - Result of cleaning





- Backup

--- γ -PhysNet
--- H.E.S.S. II mono loose cuts
--- H.E.S.S. II mono safe cuts

