

Participation of LPNHE-Neutrino group in Japan – based accelerator neutrino program (from T2K to Hyper-Kamiokande)

- Boris A. Popov
- for LPNHE-neutrino group

T2K-II Hyper-K

The T2K-II project: the second phase of the T2K experiment

Alain Blondel², Margherita Buizza Avanzini¹, Olivier Drapier¹, Jacques Dumarchez², Frank Gastaldi¹, Claudio Giganti*², Michel Gonin¹, Mathieu Guigue², Jean-Michel Lévy², Thomas Mueller¹, Boris Popov², Benjamin Quilain¹ and Marco Zito²

¹LLR Neutrino group, IN2P3/Ecole Polytechnique ²LPNHE Neutrino group, IN2P3/Sorbonne University

The Hyper-Kamiokande experiment

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LPNHE-Neutrino group members:

- Bernard Andrieu (CR)
- Claudio Giganti (CR/HDR)
- Mathieu Guigue (MdC SU)
- Jacques Dumarchez (DR)
- Jean-Michel Levy (benevole)
- Boris Popov (DR)
- + former PhD students: Laura Zambelli (2010-2013), Pierre Bartet-Friburg (2013-2016), Matej Pavin (2014-2017), Simon Bienstock (2015-2018)
- + new members starting from 2019
- Ciro Riccio (invited researcher)
- Alain Blondel (DR)
- Marco Zito (DR)
- Viet Nguyen (PhD student, 2019-)
- + a new PostDoc (Adrien Blanchet) from January 2020

The group participates in the following experiments:

T2K → T2K-II [till 2026] NA61/SHINE (for T2K/Fermilab neutrino beams, then for Hyper-Kamiokande/DUNE) [>2021] Hyper-Kamiokande [2018-]

LPNHE - ITA support (as of 2019):

- Jean-Marc Parraud (Technical Coordinator @ LPNHE)
- Eric Pierre
- Francois Toussenel
- Julien Philippe
- Yann Orain
- Diego Terront
- Stefano Russo

Support from IN2P3, ANR and SU

On top of the regular annual support from IN2P3 the group has recently obtained:

- ANR grant (CG) [2-years postdoc + PhD student grant]
- SU "Emergence" grant (MG) [1-year postdoc]

LPNHE -Neutrino group research:

Main research topics:

Precise measurements of neutrino and anti-neutrino oscillation parameters and search for CP-violation in the lepton sector

Measurements of neutrino and anti-neutrino interaction cross sections

Hadron production measurements for precise predictions of accelerator (anti-)neutrino fluxes

The group had and still has important responsibilities:

ND280-upgrade project leader (MZ → CG), T2K Oscillation Analysis convener (CG), T2K Beam group convener (BP), ND280-upgrade Simulation group convener (MG), ND280 Magnet operation, TPC operation, Publication Board members, Speakers Board members, Internal referees, NA61/SHINE Analysis coordinator, etc.

More than 50 publications during the last 5 years

Awards: Le Prix "La Recherche" Physique (2012), Breakthrough Prize in Fundamental Physics (2016)

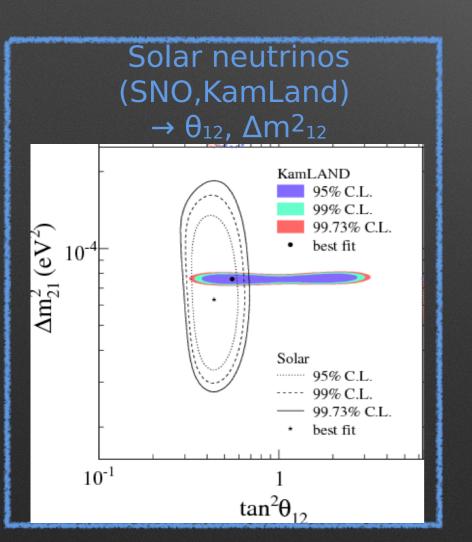
Neutrino oscillations

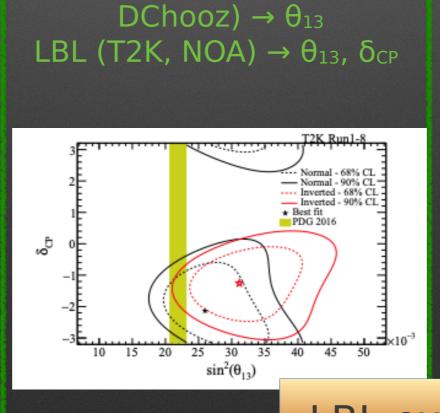
$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} \cos\theta_{12} & \sin\theta_{12} & 0 \\ -\sin\theta_{12} & \cos\theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} \cos\theta_{13} & 0 & \sin\theta_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -\sin\theta_{13}e^{-i\delta} & 0 & \cos\theta_{13} \end{pmatrix} \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos\theta_{23} & \sin\theta_{23} \\ 0 & -\sin\theta_{23} & \cos\theta_{23} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

- 3 mixing angles
- 2 independent mass differences

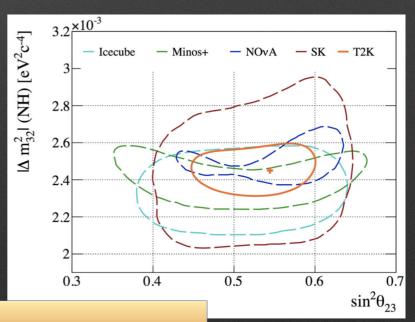
Reactors (Daya Bay, RENO,

3 1 CP violation phase → not yet measured





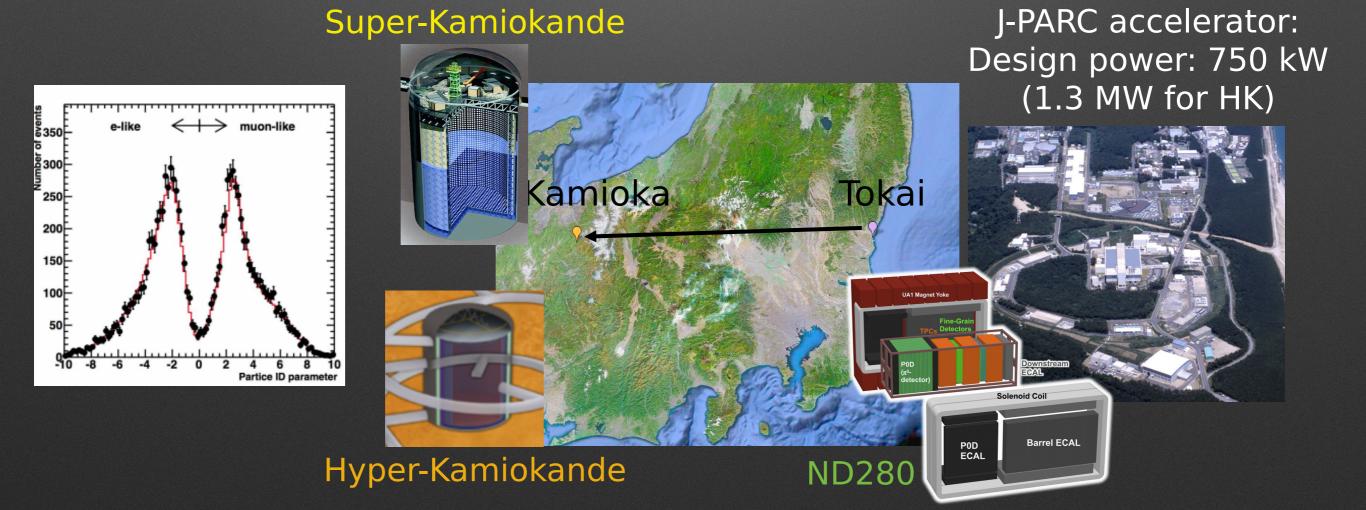
Atmospheric (SK, IceCUBE) LBL (Minos, T2K, NOA) $\rightarrow \theta_{23}$, $|\Delta m^2_{32}|$

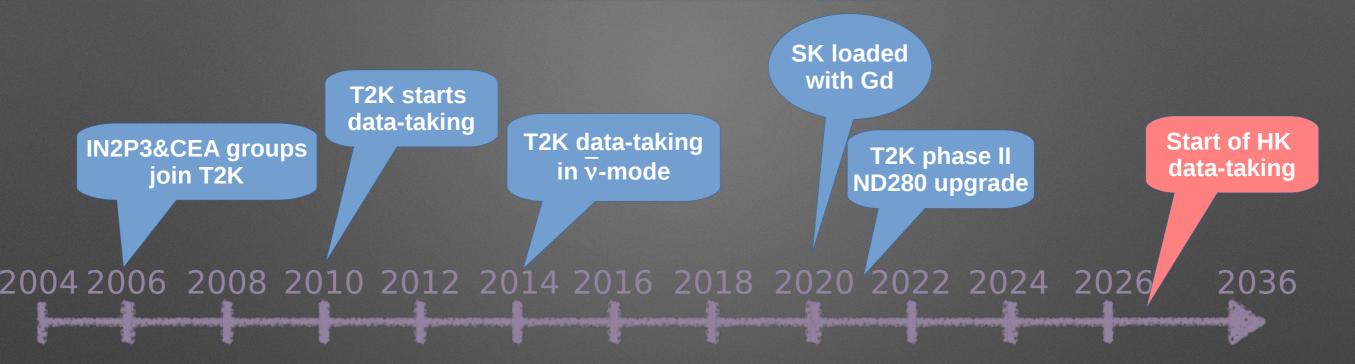


LBL experiments

Tokai-to-Kamioka (SK or HK)

- * High intensity \sim 600 MeV ν_{μ} beam produced at J-PARC (Tokai, Japan)
- * Neutrinos detected at the Near Detector (ND280) and at the Far Detector, Super-Kamiokande (Hyper-Kamiokande) 295 km from J-PARC
- * Main physics goals:
 - * Observation of v_e and \overline{v}_e appearance \rightarrow determine θ_{13} and δ_{CP}
 - * Precise measurement of v_{μ} and \overline{v}_{μ} disappearance $\rightarrow \theta_{23}$ and Δm^2_{32}





- *15 years of successful research by IN2P3 and CEA groups in Japan
- *15 exciting years to come
 - *****SK run with Gd
 - *T2K phase II and Near Detector upgrade
 - *Hyper-Kamiokande!

IN2P3&CEA groups join T2K

T2K starts data-taking

T2K data-taking in v-mode

SK loaded with Gd

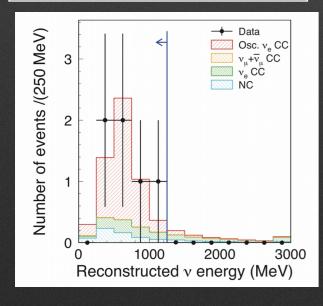
T2K phase II ND280 upgrade Start of HK data-taking

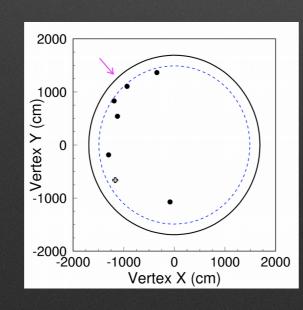
2004 2006 2008 2010 2012 2014 2016 2018 2020 2022 2024 2026

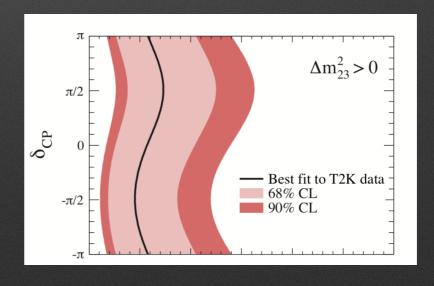
2036

Hints of ve appearance $(\theta_{13} \neq 0 @ 2.5\sigma)$

Phys.Rev.Lett. 107 (2011) 041801







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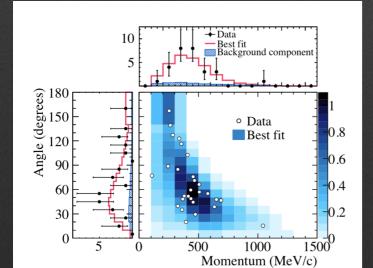
2036

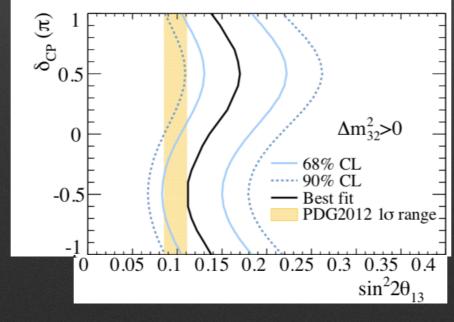
Hints of ve appearance $(\theta_{13} \neq 0 \ @ 2.5\sigma)$

Phys.Rev.Lett. 107 (2011) 041801

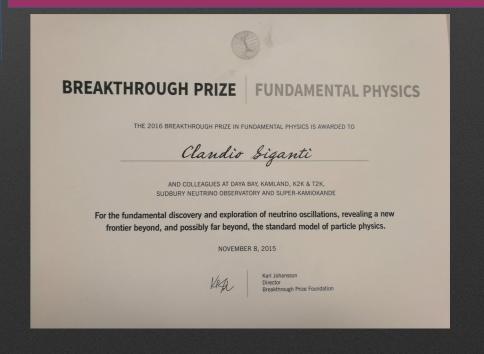
Observation of ve appearance $(\theta_{13} \neq 0 @ 7.3\sigma)$

Phys.Rev.Lett. 112 (2014) 061802





2016 Breakthrough prize in Physics to Daya Bay, Kamland, SK, SNO, K2K and T2K Collaborations



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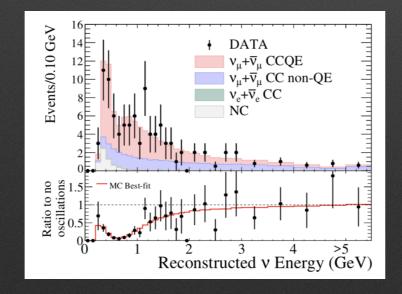
Observation of ve appearance $(\theta_{13} \neq 0 @ 7.3\sigma)$

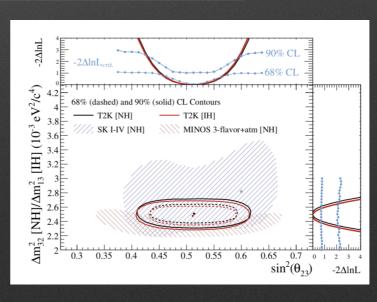
World-best measurements of Θ_{23} and Δm^2_{32}

Phys.Rev.Lett. 107 (2011) 041801

Phys.Rev.Lett. 112 (2014) 061802

Phys.Rev.Lett. 112 (2014) no.18, 181801





IN2P3&CEA groups join T2K

T2K starts data-taking

T2K data-taking in \overline{v} -mode

SK loaded with Gd

> T2K phase II ND280 upgrade

Start of HK data-taking

2004 2006 2008 2010 2012 2014 2016 2018 2020 ²022 2024 2026 2036

Hints of ve appearance $(\theta_{13} \neq 0 \ @ \ 2.5\sigma)$

Observation of ve appearance $(\theta_{13} \neq 0 @ 7.3\sigma)$

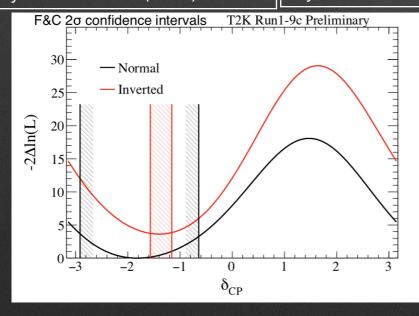
World-best measurements of Θ_{23} and Δm^2_{32}

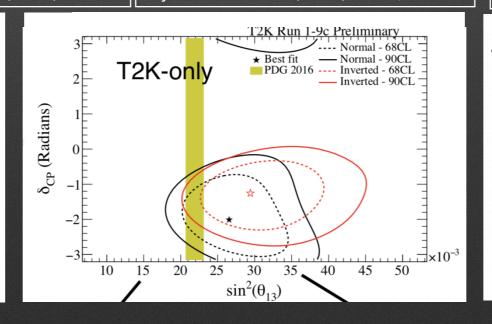
Hints of leptonic CP violation→ $sin(\delta_{CP})=0$ excluded at 95%

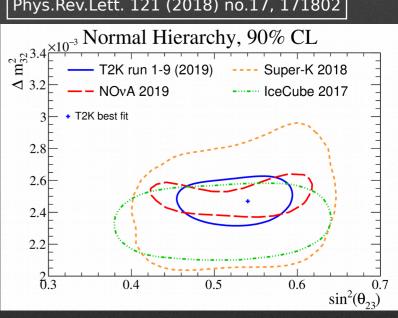
Phys.Rev.Lett. 107 (2011) 041801

Phys.Rev.Lett. 112 (2014) 061802 Phys.Rev.Lett. 112 (2014) no.18, 181801

Phys.Rev.Lett. 121 (2018) no.17, 171802







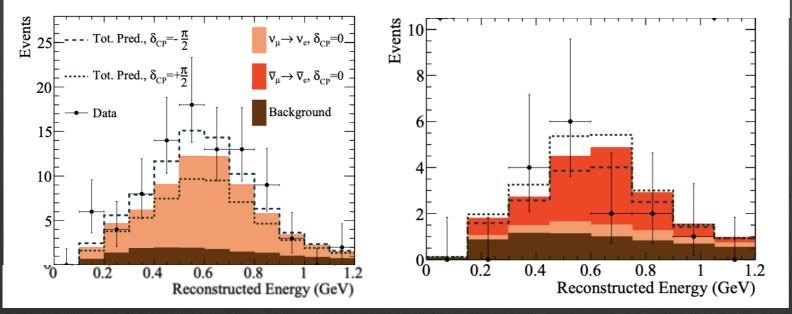
Hot of the press

- *Paper submitted to Nature:

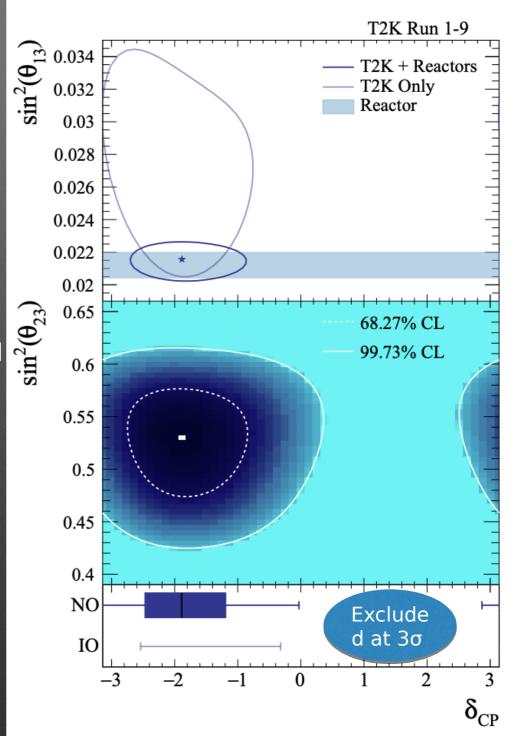
 Constraint of the matter-antimatter symmetry
- violating phase in Neutrino oscillations
- *First 3σ exclusion for 46% (65%) of the δ_{CP} values in NO (IO)
- *Need more data (and smaller systematics)!

v-mode

v-mode

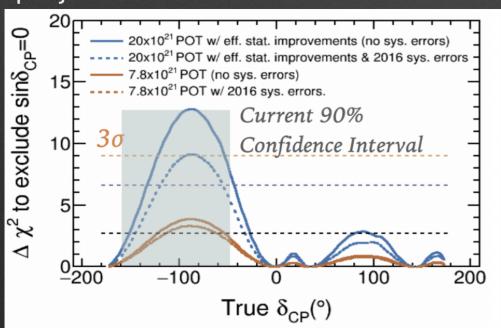


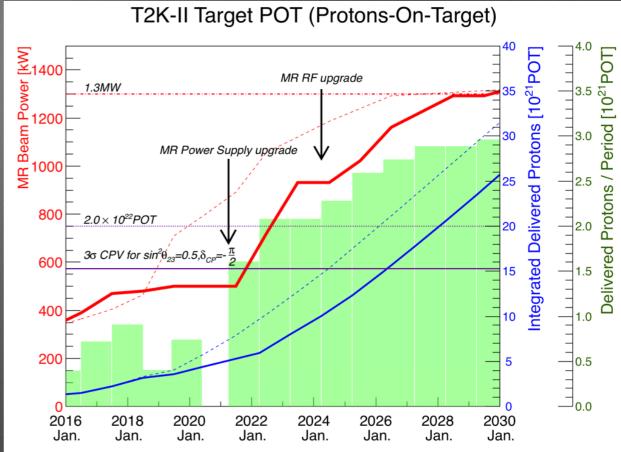
	<i>v</i> -mode	- v-mode
Observed	90	15
Expected if δ CP=- π /2	81.7	17.2
Expected if δCP=0	68.4	19.6

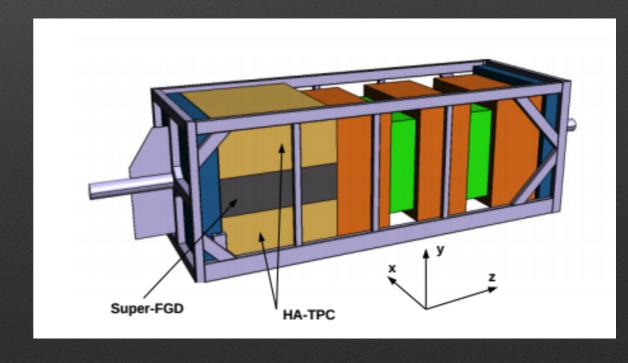


T2K phase-II

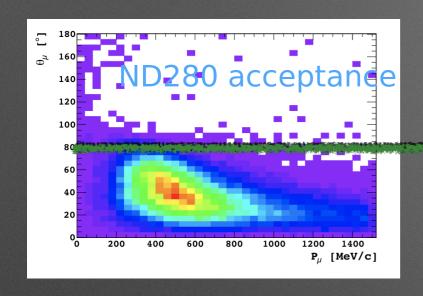
- *****Upgrade of J-PARC Main Ring (1.3 MW beam)
 - *Approved and funded, will be done in 2021
- *****Goal: collect >10x10²¹ POT by 2026 → 3σ measurement of CP violation if δ_{CP} ~-π/2
- *Near Detector upgrade to reduce systematics from ~7% to ~4%
 - *We will install the new detectors in 2021
 - *Use the ND280 Upgrade detector also as initial Near Detector for HK
 - *Funded by France (CEA+IN2P3), Italy, US, Japan, Spain, Poland, Russia, Germany, Switzerland
 - *Leading role of the LPNHE group (project coordinator)
- *Improvements of the Far Detector thanks to the SK-Gd project

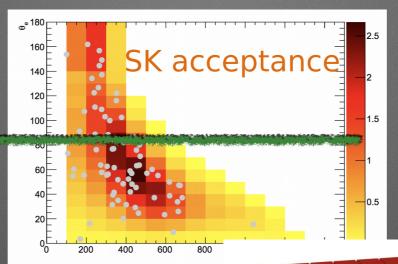


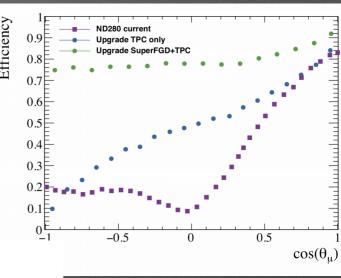




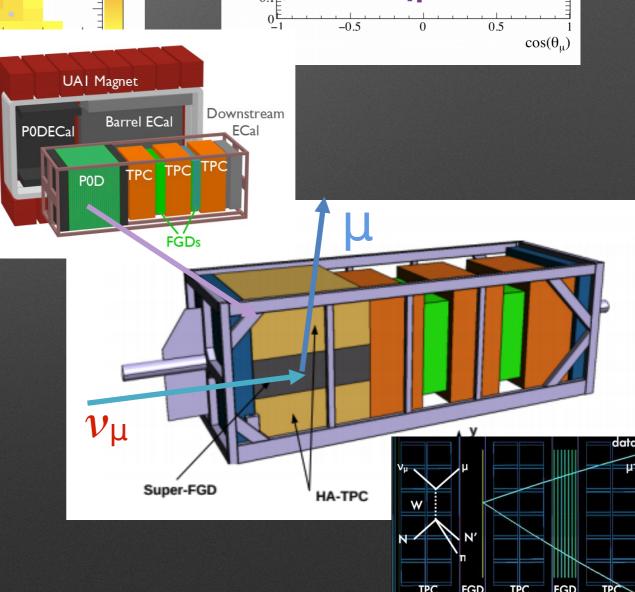
ND280 upgrade







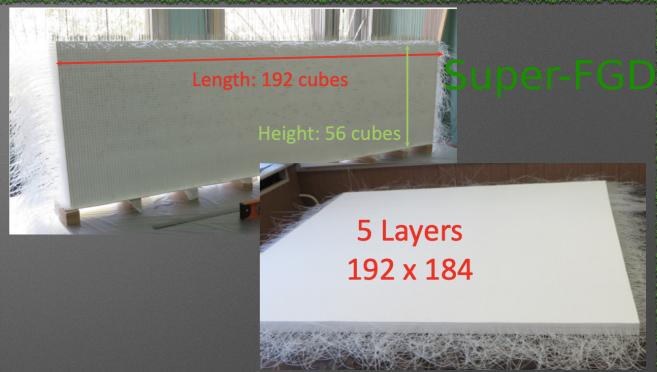
- *Main strength of ND280 : magnetized detector \rightarrow separate ν from $\overline{\nu}$ (cannot be done in SK or HK)
- *Main limitation of ND280: reduced angular acceptance → only forward going tracks are reconstructed with high efficiency
- *An analysis dedicated to select tracks with high polar angles → 20% efficiency
- *We can do better with an upgrade → Horizontal target and horizontal TPCs



New detectors

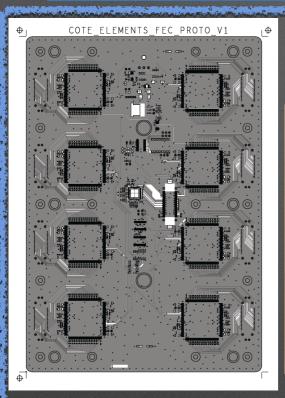


- *New TPCs instrumented with Resistive MicroMegas
- * DESY and CERN Test beams
 - **★** Spatial resolution ~200 μm
 - *dE/dx resolution ~7% for 70 cm tracks
- * First TPC expected by Summer 2020
- ***** LPNHE responsible for the Front-End electronics



- New concept of detectors, 2x10⁶
 1cm³ cubes
 - 25% of the cubes already built
 - *All produced by Dec 2020
- Each read by 3 WLS
- Improve reconstruction of the hadronic part of the interactions
- LLR responsible for the Front-End electronics using CITIROC chips

LPNHE contributions

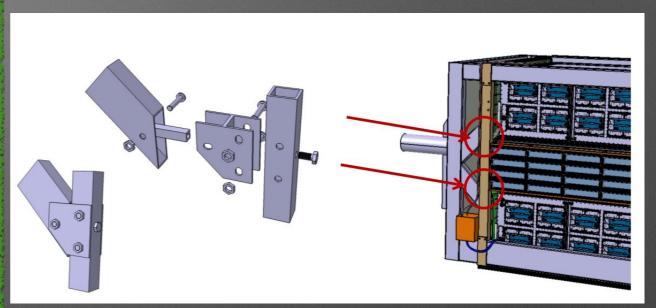


TPC electronics



- *Design and production of 80 Front-End Cards (FEC) for TPC readout
 - *FEC-mockup to validate floating connectors
 - * First full FEC design is ready, prototype is being produced
- *Test production by Summer 2020
- * Full-scale production just after
- *Installation in Japan by end 2021

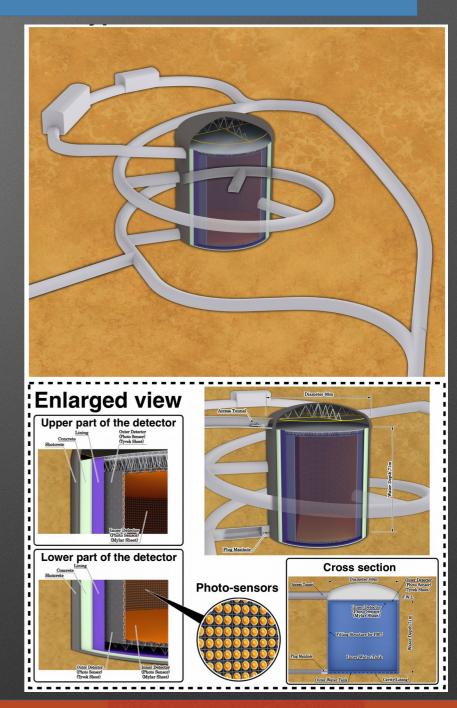
Mechanics



- Contributions to installation of new detectors to the existing basket
 - Design of basket modifications including finite element and seismic analysis
 - Global detector envelopes and service volume definitions
- FEC cooling system

Hyper-Kamiokande (HK)

- *Extremely well established Water Cherenkov technology
 - *190 kton FV (SK 22.5)
 - *Instrumented with up to 40k PMTs
- *HK will be the most sensitive observatory for rare events (proton decay, SN neutrinos, ...)
- *Search for CP violation in lepton sector
 - *Upgrade of J-PARC neutrino beam (1.3 MW)
 - *Near and Intermediate detector complex
- *August 2019 → MEXT approved HK and required budget for construction to the Ministry of Finance
- *Begin construction in 2020, start operation in 2027



MEXT Statement

In addition to the ongoing 13 largescale projects, the next-generation neutrino research project Hyper-Kamiokande, will be newly launched in FY2020

(MEXT) will start the next-generation neutrino research project "Hyper-Kamiokande" in JFY2020.

- 日本学術会議において科学的観点から策定したマスタープランを踏まえつつ、専門家等で構成される文部科学省の審議会において戦略性・緊急性等を加味し、 ロードマップを策定。
- ロードマップの中から大規模学術フロンティア促進事業として実施するプロシェクトを選定の上、国立大学法人運営費交付金等の基盤的経費により戦略的・計画 的に推進。原則、10年間の年次計画を答定し
- 現行の13プロジェクトに加え、<u>令和2年度より、ニュートリノ研究の次世代計画である「ハイパーカミオカンデ計画」に新たに着手</u>。

主众成果

- ノーベル賞受賞につながる画期的研究成果 (受賞歷: H14小柴昌俊氏、H20小林誠氏、益川敏英氏, H27梶田隆章氏)
- 年間約1万人の共同研究者が集結し、国際 共同研究を推進。このうちの半数以上が外国人

大規模学術フロンティア促進事業等の主な事業

大型電波望遠鏡「アルマ」による国際 共同利用研究の推進

新しいステージに向けた学術情報 ネットワーク(SINET)整備

(情報・システム研究機構国立情報学研究所)

Next generation of neutrino project with a 260 kton detector and the J-PARC upgrade. The project will reveal the mysteries in elementary particles and the Universe by the observation of proton decays and the neutrino researches including CP violation.

天文学·宇宙 物理学分野	論文数	Top10 %割合	国際共著 割合
すばる望遠鏡	644	18,5%	86,3%
アルマ望遠鏡	878	27,3%	89.0%
日本全体	8,938	12.9%	68.0%
世界全体	103,445	9,6%	50,6%

※ 大学共同利用機関法人自然科学研究機構がInCites | (Web of Science) に基づき 2013-2017の5か年に出版された天文学・宇宙物理学分野の論文 (article, review)

三大フラックホールの「影」の撮影に世界で初めて 学術情報基盤 成功した国際プロジェクトに参加し、高い感度の

観測機能により、その成果に大きく貢献。

<産業等への波及>

産業界と連携した最先端の研究装置開発により、 イノベーションの創出にも貢献

・【すばる望遠鏡】超高感度カメラ技術⇒医療用X線カメラへの応用

【放射光施設】加齢による毛髪のハリ・コシの低下が毛髪内の亜鉛と 関係性を解明⇒亜鉛を毛髪に浸透させる**新しいヘアケア技術**





化により、ニュートリノの検出性能を著しく向上。素 研究を通じ、新たな物理法則の発見、素粒子と学 宙の謎の解明を目指す。[ロードマップ2017掲載事業]

ハイバーカミオカンデ(HK)計画の推進

ハイバーカミオカンラ

(高エネルギー加速器研究機構)

ニュートリノビーム

HK @ IN2P3



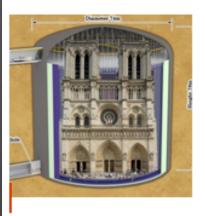
L'IN2P3

Recherche

Innovation

Formation

Médiation scientifique



A. / A.

Contact(s)

Laurent Vacavant Perrine Royole-Degleux

Partagercecontenu







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Accueil > Actualités

Le projet Hyper-Kamiokande approuvé par le MEXT

01 actabre 2019

PHYSIQUE DES NEUTR IN OS

Le Ministère de la recherche et de l'enseignement japonais (MEXT) a décidé de lancer le projet Hyper-Kamiokande. Le coût de cette infrastructure de recherche dédiée à l'étude des neutrinos est estimé à environ 550 millions d'euros et sa construction pourrait démarrer dès avril 2020, tout près du site de l'expérience actuelle Super-Kamiokande localisée dans les montagnes nippones. Les premières prises de données sont attendues pour les années 2027-2028.

Hyper-Kamiokande prendra la suite du projet T2K (Tokai to Kamiokande) actuellement en fonctionnement avec le réservoir Super-kamiokande. Il représente la nouvelle génération de détecteurs Tcherenkov à eau de très grande taille, une technique éprouvée pour la détection des neutrinos. Il est prévu que son volume d'eau utile, environ 200,000 tonnes, soit supérieur d'un facteur 10 à celui de son prédécesseur Super-Kamiokande. En détectant les neutrinos, le détecteur Hyper-Kamiokande sera à la fois un microscope exceptionnel pour observer les particules élémentaires et un télescope original pour observer le Soleil et les phénomènes cosmologiques très violents. Le projet Hyper-Kamiokande comprend un programme de physique extrêmement riche avec un potentiel de découvertes très important allant de l'étude de la violation matière antimatière dans le secteur leptonique, la recherche de désintégration du proton, l'étude des neutrinos d'origine astronomique, notamment les neutrinos émis lors d'effondrement de Supernovæ...

La collaboration Hyper-Kamiokande est actuellement composée de physiciens provenant de 13 pays de 3 continents différents. Les laboratoires LLR et LPNHE de l'IN2P3 déjà impliqués dans l'expérience T2K réfléchissent activement à participer à ce projet.

Contact

A Laurent Vacavant

DAS Particules et Hadronique

01 44 96 49 74 ✓ vacavant@in2p3.fr Chargée de communication

Perrine Royole-Degieux

% 04 73 40 54 59

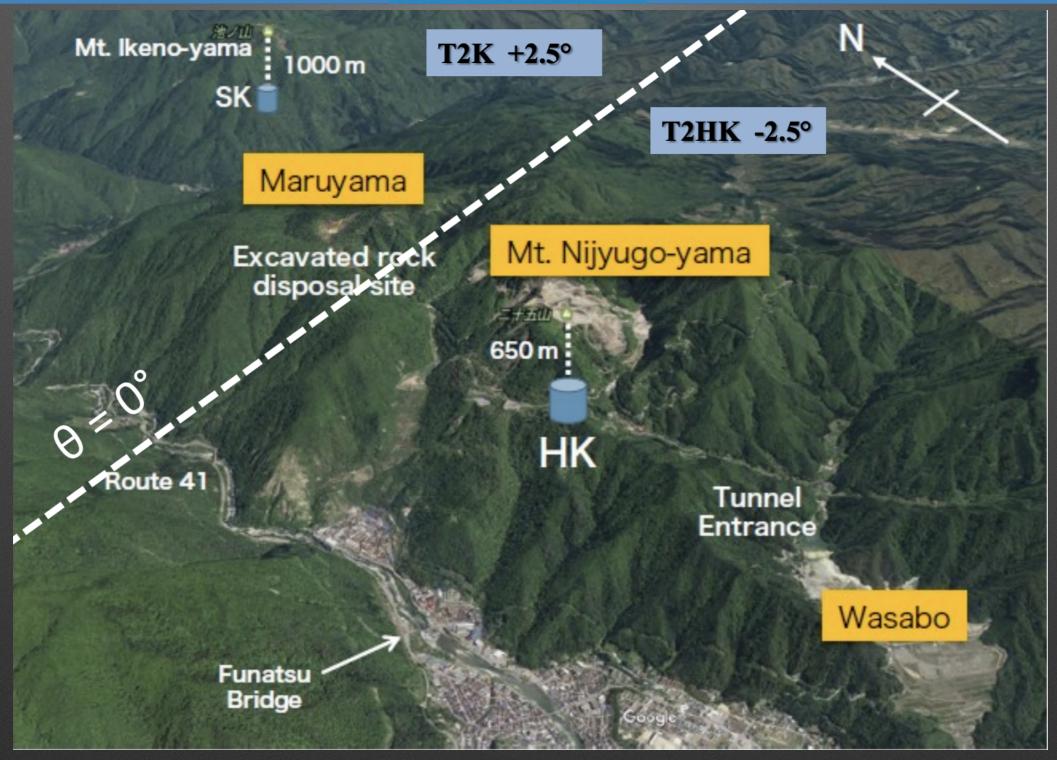
✓ royole@in2p3.fr

As announced by the IN2P3 directorate, the LPNHE physicists in close collaboration with LLR and CEA colleagues are actively thinking (and trying to take actions as well!) on participation in this project.

We have already identified some possible contributions, see below.

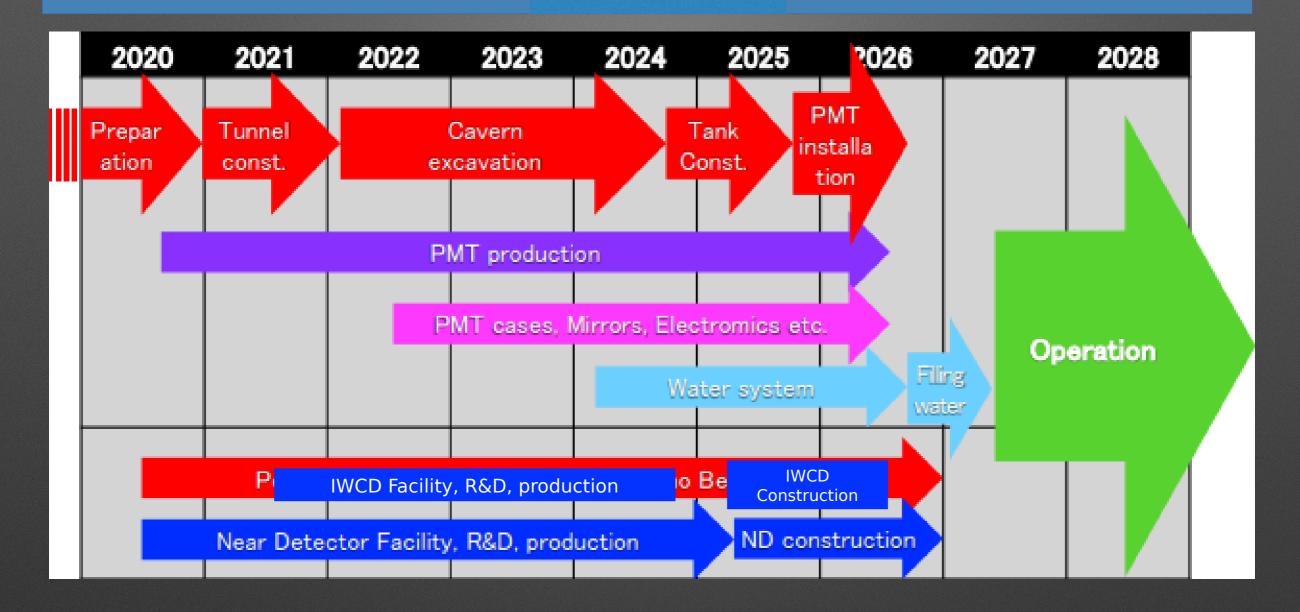
Negotiations with the IN2P3 directorate are on-going.

HK: Where



295 km from J-PARC 2.5 degrees off-axis (as SK)

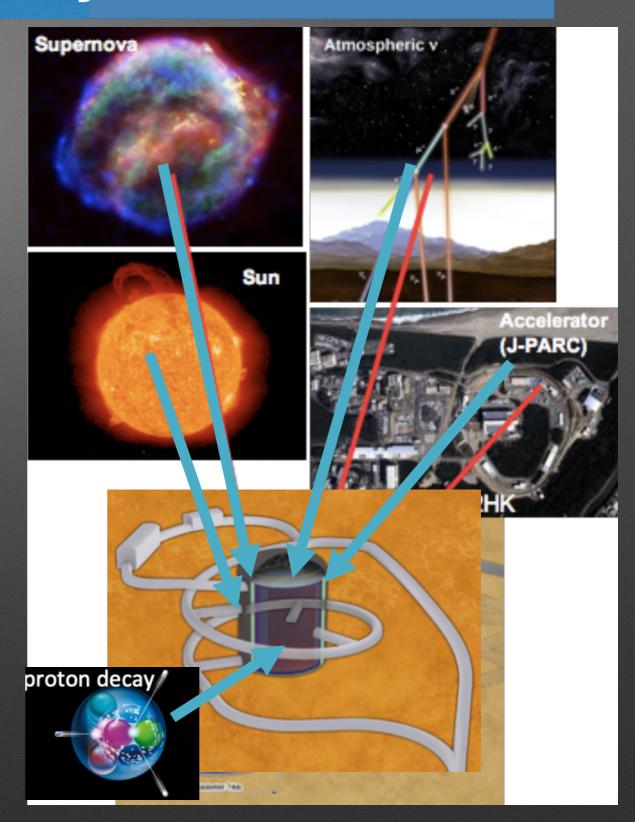
HK: When



- *Start Construction in 2020 (some preparatory work already started)
- *Start data taking in JFY 2027
- *Budget requested by MEXT to Ministry of Finance for Japanese part (~80% of the total cost of the experiment)
- *International contributions being formalized

HK: Why

- *****Neutrino oscillation → CP violation
 - *Combination of beam and atmospheric neutrinos
- *Search for nucleon decay
 - *~10 times better sensitivity than SK
- *Neutrino astrophysics
 - *Solar v
 - *Atmospheric *v*
 - *****SuperNovae burst
 - *Relic SN neutrinos
- *****Geophysics
- *****Others

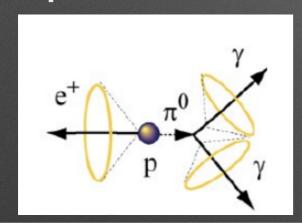


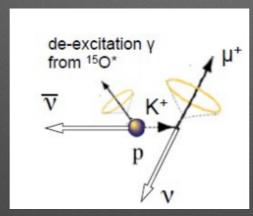
Proton-decay

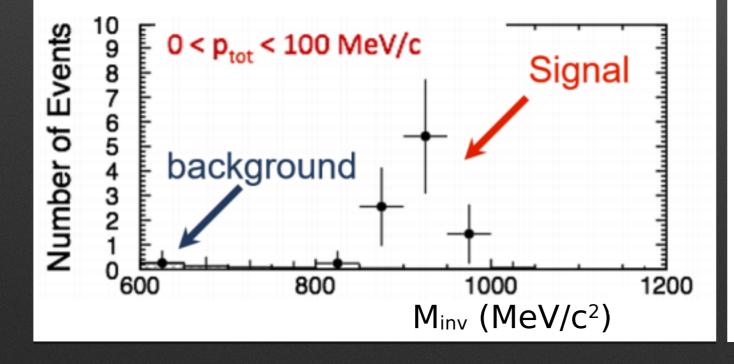
Sensitivity to many different modes Surpass SK by one order of magnitude in the leading p \rightarrow e⁺ + π^0

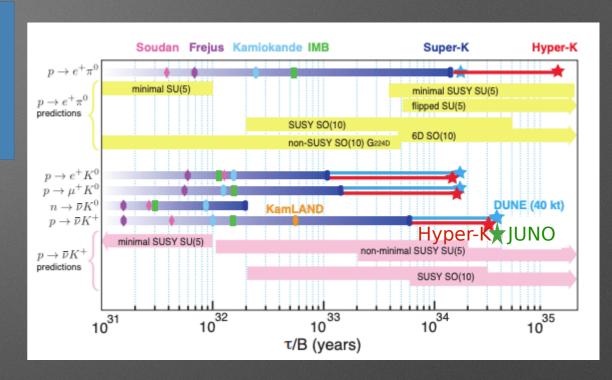
$$p\rightarrow e^++\pi^0$$

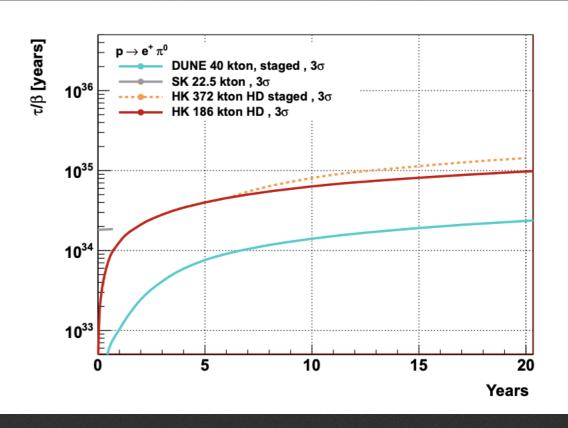




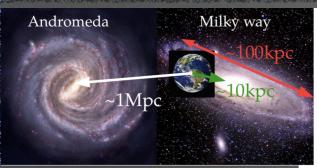




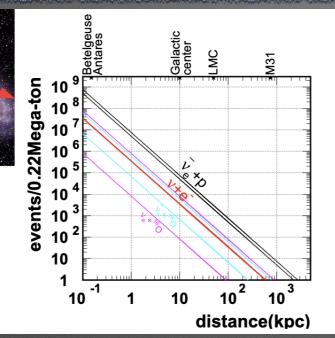


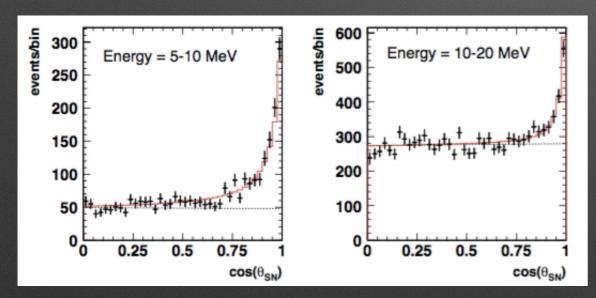


Supernovae neutrinos

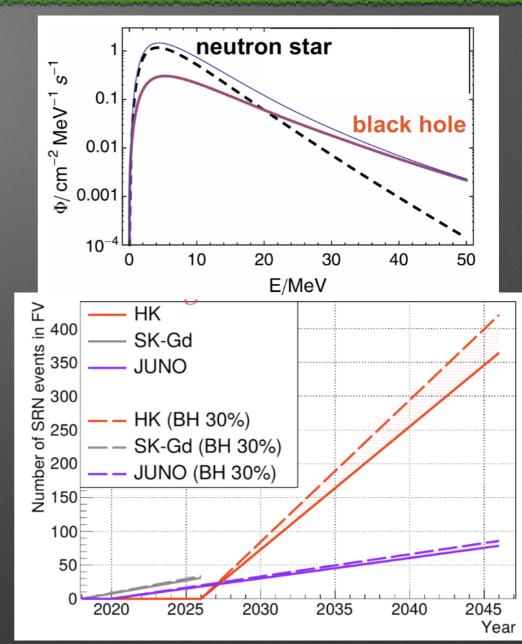


- * IBD: huge statistics → SN model
- * ES: directionality



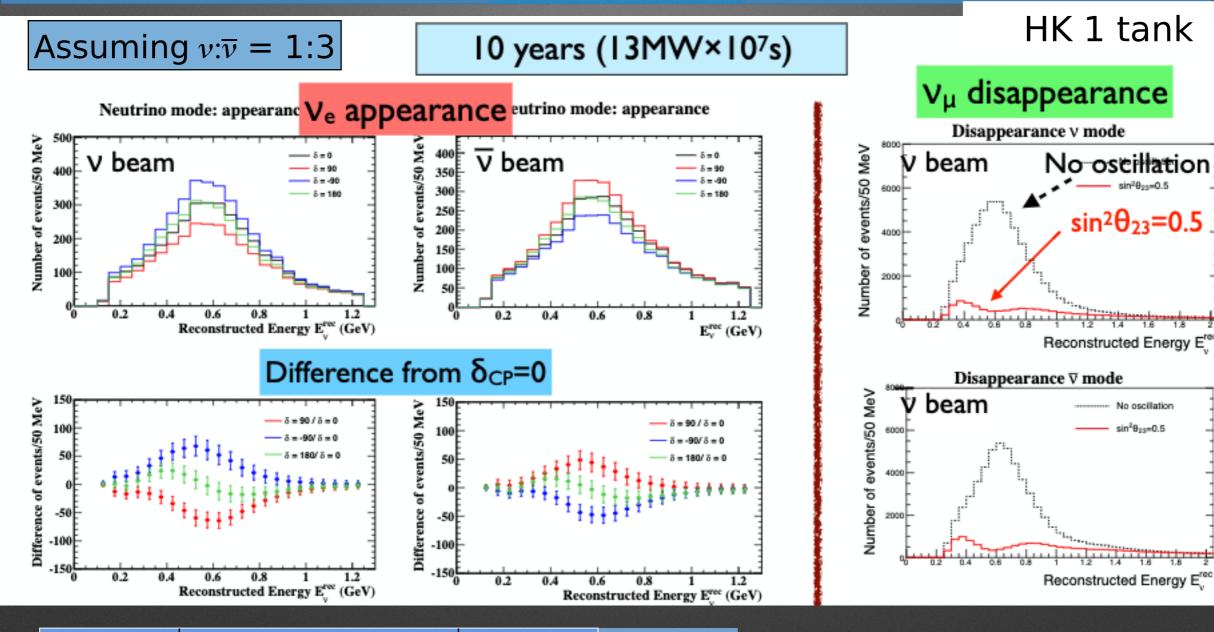


- *~80k IBD and ~3k ES for SN explosions in the galactic center
- *Sensitive also to SN explosions in Andromeda



- *If SRN will be observed in SK-Gd or JUNO we will perform precision measurement with HK
- Constraints on cosmic star history

Long-baseline physics

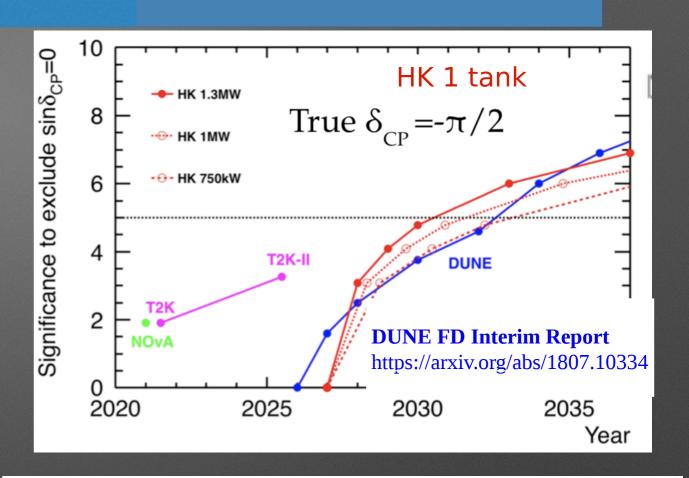


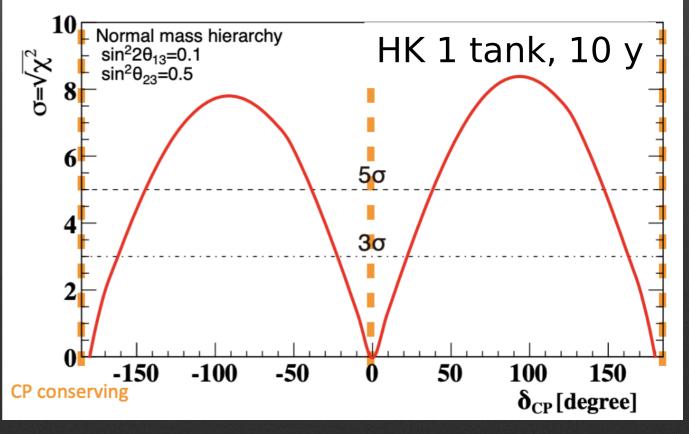
$\delta CP = -\pi/2$	Sig	nal	Total	
v-mode	1643	15	400	2058
v-mode	206	1183	517	1906

Huge statistics →
sensitivity to CP
violation
Need to control
systematics!

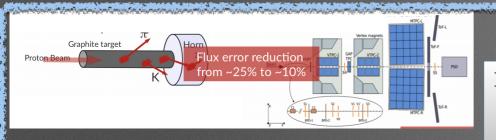
CP Violation

- *Exclusion of $sin(\delta_{CP})=0$
 - *8 σ for $\delta_{CP} \sim \pm \pi/2$
 - *>3 σ (5 σ) significance for 76% (57%) of δ_{CP} space
- *Sensitivity will be further enhanced by combination with atmospheric neutrino measurements
- *Assume systematics uncertainties of ~4% (currently 7% for T2K)

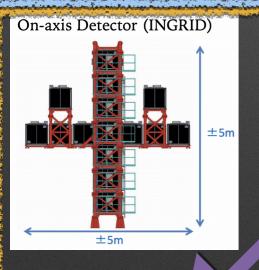


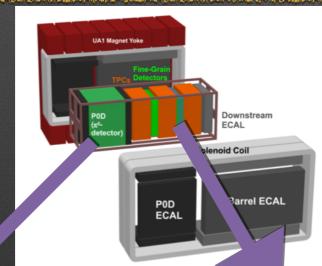


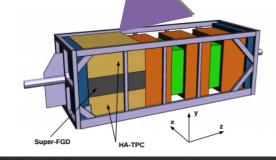
Systematics and Near Detectors

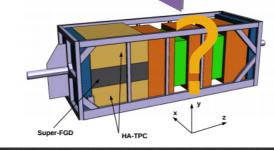


- NA61/SHINE hadron-production experiment @CERN
- T2K → uncertainties on the neutrino flux ~5% thanks to NA61/SHINE
- New measurements planned for HK



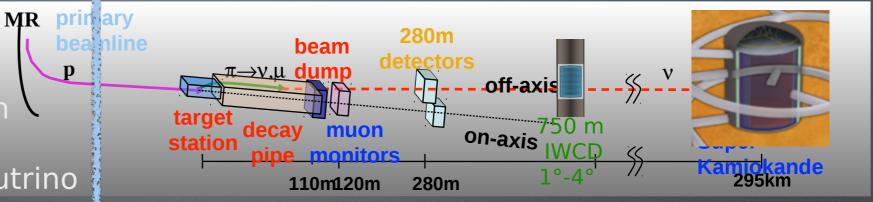


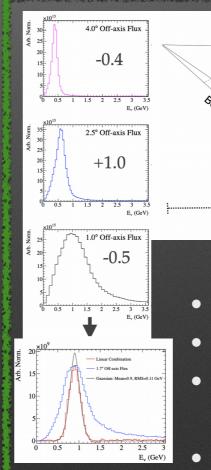


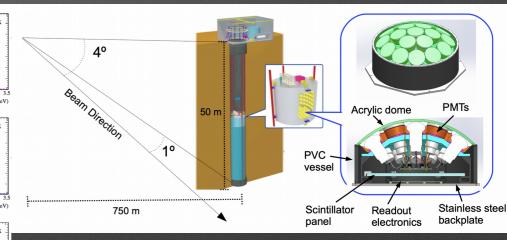


2021

2030?

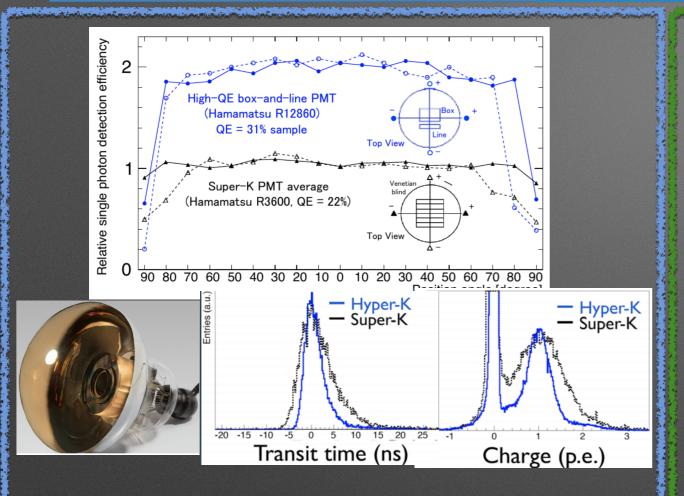






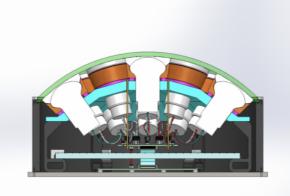
- Intermediate WC detector
- 1 kton mass
- Instrumented with ~500 multi-PMTs
- Movable position to scan different off-axis angles

Hyper-K photo-detection system



- *HK will be instrumented with "box-and-line" 20" PMTs
- *At least 20k modules
- *31% QE (2 times better than SK)
- * Better transit time spread

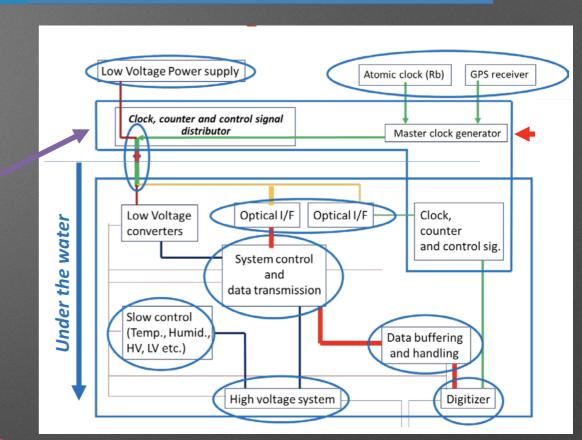
- *Array of 19 3" PMTs
- Baseline option for IWCD
- Possibility to add 5k or 10k m-PMTs in HK (depending on funding)
 - Would improve vertex reconstruction and energy resolution at low energy
 - Good opportunity for France
 - Synergies with KM3Net and with JUNO small PMTs
 - Memphyno test setup @APC

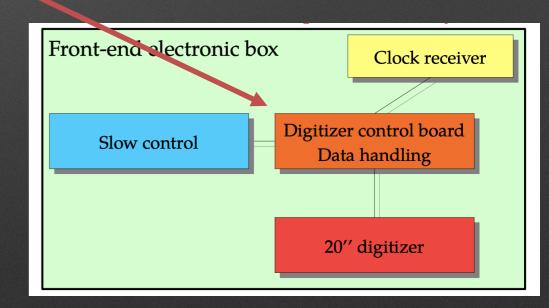




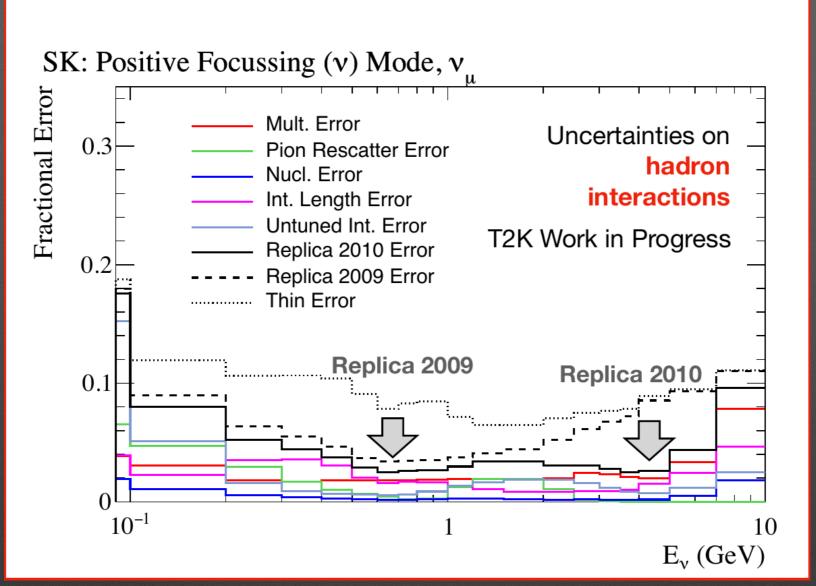
LPNHE/IN2P3 contributions

- *NA61/SHINE hadron-production measurements for HK and significant involvement in the ND280 upgrade
- *Contribution to the Far Detector centered around the electronics for the 20" PMTs
 - *Design and procurement of the clock distribution and time synchronization system for the 20" PMTs (White Rabbit or Custom Made solution)
 - *Front End digitizers (OMEGA chips) and front end boards for the 20" PMTs
- *Such contributions can eventually be extended to the Multi-PMTs in HK
 - *Testing one prototype in Memphyno@APC
 - *Test Beam experiment @CERN (LOI submitted to SPSC, ~100 mPMTs, data taking in 2022) → provide synchronization system, deploy few 20" PMTs
- *****Computing → CC Lyon Tier-1 for HK



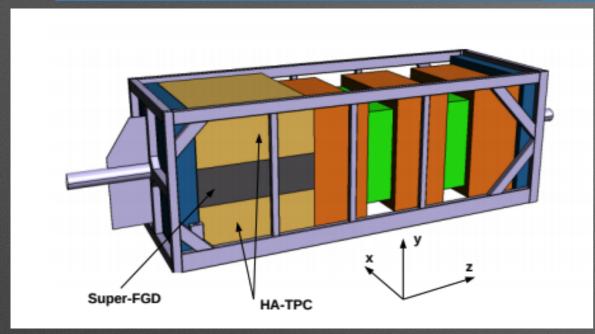


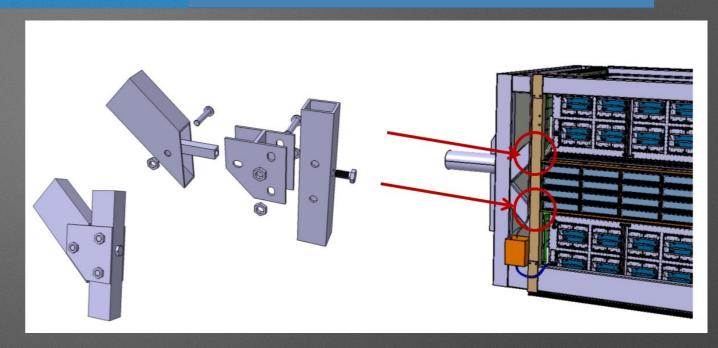
LPNHE/IN2P3 contributions (NA61)

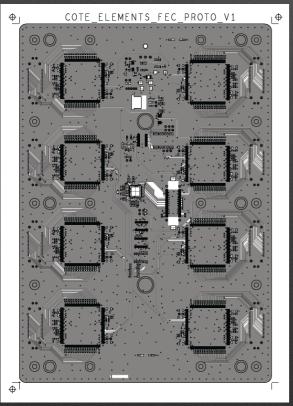


- Usage of NA61/SHINE replica-target measurements allow for further improvement in T2K (anti-)neutrino flux uncertainty (down to ~5%). Even better knowledge is desired for T2K-II and Hyper-K. New measurements are planned after the CERN LS2 (see CERN-SPSC-2018-008):
 - *Improved measurements with T2K replica target, considering alternative target material Super-Sialon ($Si_3N_4Al_2O_3$);
 - *With additional tracking detectors surrounding the long target;
 - *Hadron production with low momentum beam (<12 GeV/c).

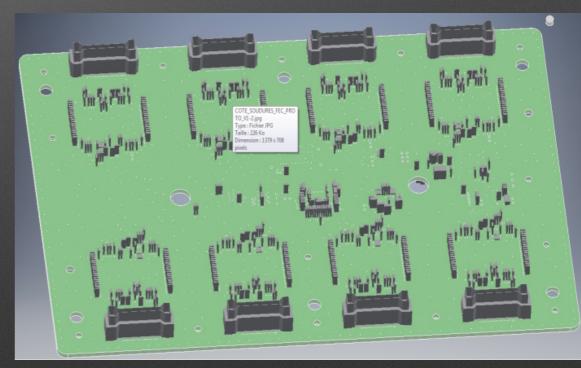
LPNHE/IN2P3 contributions (ND280)







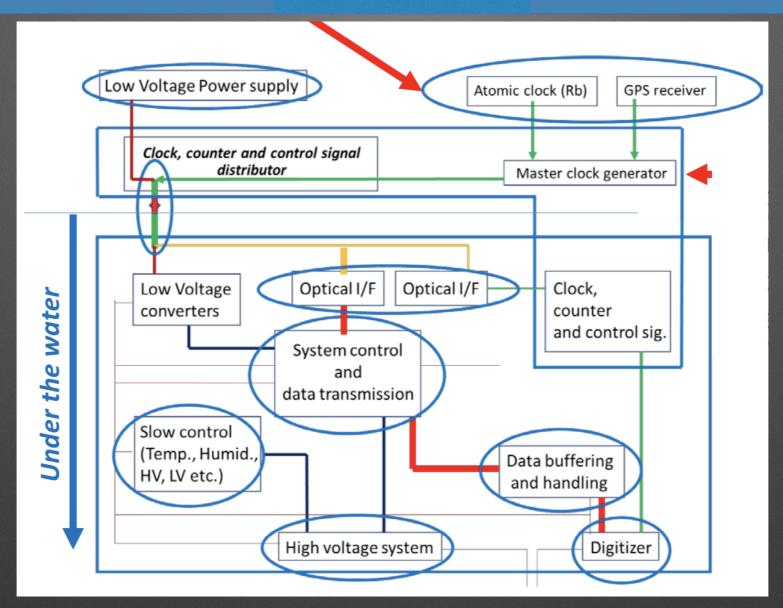




Our on-going contributions to the ND280-upgrade are also important for HK, since the upgraded detector will be taking data when the HK starts operation in 2027.

Lab resources already allocated (planned for the future)

LPNHE/IN2P3 contributions (clock)



- * Clock distribution and PMT synchronization is a critical part of the experiment: requirement is to have a timing resolution at the level of 1 ns with a maximum jitter of 100 ps RMS. This requires the distribution of a clock signal, synchronous with the GPS and a local atomic clock, to all Front-End nodes plus some signals to align local counters.
- *System structured as a data exchange protocol. The bandwidth not occupied by timing information could be used to move (physics or slow control) data.
- *Possible designs are based on White Rabbit protocol or custom solution

Estimated resources required from the lab: 0.6 FTE x 5 years

LPNHE/IN2P3 contributions (mPMT)



LPNHE/IN2P3 contributions (computing)



Conclusions

- ***LPNHE-neutrino** group continues its strong participation to the extremely successful v oscillations program in Japan
 - *T2K phase II and ND280 Upgrade → CP violation at 3σ by 2026 (support from the lab is very much appreciated!)
 - ***SK-Gd** → Observation of Supernova relic neutrinos
- *Excellent news for Hyper-Kamiokande, the next generation neutrino observatory
 - *****Experiment approved by MEXT
 - *Profit of the extremely well known Water Cherenkov technology
 - *Start data taking in 2027
 - *Leading experiment in the search for CP violation in the leptonic sector
 - *Most sensitive detector for proton decay
 - *Observatory for neutrinos from different sources (Supernova, Sun, Atmosphere, Gravitational Waves,...)
- *IN2P3 and CEA contributions are being defined → support from the lab is needed in order to participate in the Hyper-Kamiokande with a visible contribution!

Back-up

(MEXT) will start the next-generation neutrino research project "Hyper-Kamiokande" in JFY2020.

- 日本学術会議において科学的観点から策定したマスタープランを踏まえつつ、専門家等で構成される文部科学省の審議会において戦略性・緊急性等を加味し、 ロードマップを策定。
- ロードマップの中から大規模学術フロンティア促進事業として実施するプロシェクトを選定の上、国立大学法人運営費交付金等の基盤的経費により戦略的・計画 的に推進。原則、10年間の年次計画を答定し
- 現行の13プロジェクトに加え、<u>令和2年度より、ニュートリノ研究の次世代計画である「ハイパーカミオカンデ計画」に新たに着手</u>。

主众成果

- ノーベル賞受賞につながる画期的研究成果 (受賞歷: H14小柴昌俊氏、H20小林誠氏、益川敏英氏, H27梶田隆章氏)
- 年間約1万人の共同研究者が集結し、国際 共同研究を推進。このうちの半数以上が外国人

大規模学術フロンティア促進事業等の主な事業

大型電波望遠鏡「アルマ」による国際 共同利用研究の推進

新しいステージに向けた学術情報 ネットワーク(SINET)整備

(情報・システム研究機構国立情報学研究所)

Next generation of neutrino project with a 260 kton detector and the J-PARC upgrade. The project will reveal the mysteries in elementary particles and the Universe by the observation of proton decays and the neutrino researches including CP violation.

天文学·宇宙 物理学分野	論文数	Top10 %割合	国際共著 割合
すばる望遠鏡	644	18,5%	86,3%
アルマ望遠鏡	878	27,3%	89.0%
日本全体	8,938	12.9%	68.0%
世界全体	103,445	9,6%	50,6%

※ 大学共同利用機関法人自然科学研究機構がInCites | (Web of Science) に基づき 2013-2017の5か年に出版された天文学・宇宙物理学分野の論文 (article, review)

三大フラックホールの「影」の撮影に世界で初めて 学術情報基盤 成功した国際プロジェクトに参加し、高い感度の

観測機能により、その成果に大きく貢献。

<産業等への波及>

産業界と連携した最先端の研究装置開発により、 イノベーションの創出にも貢献

・【すばる望遠鏡】超高感度カメラ技術⇒医療用X線カメラへの応用

【放射光施設】加齢による毛髪のハリ・コシの低下が毛髪内の亜鉛と 関係性を解明⇒亜鉛を毛髪に浸透させる**新しいヘアケア技術**





化により、ニュートリノの検出性能を著しく向上。素 研究を通じ、新たな物理法則の発見、素粒子と学 宙の謎の解明を目指す。[ロードマップ2017掲載事業]

ハイバーカミオカンデ(HK)計画の推進

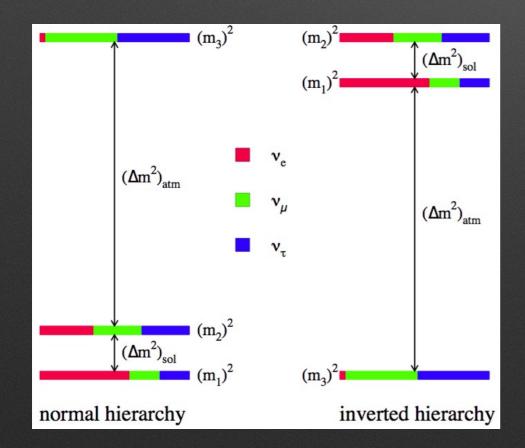
ハイバーカミオカンラ

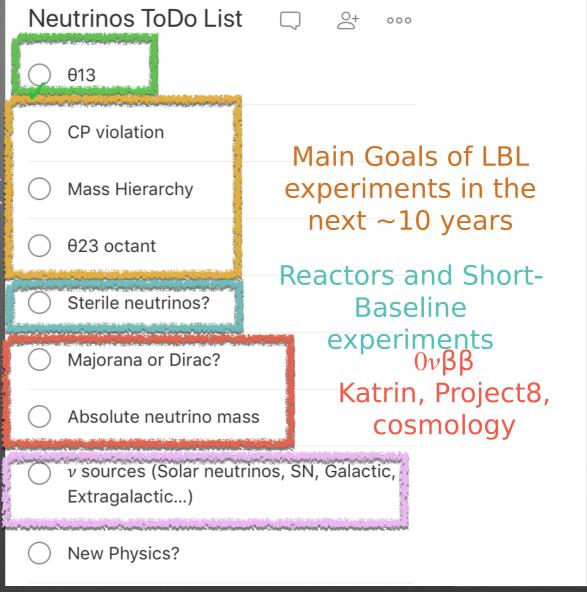
(高エネルギー加速器研究機構)

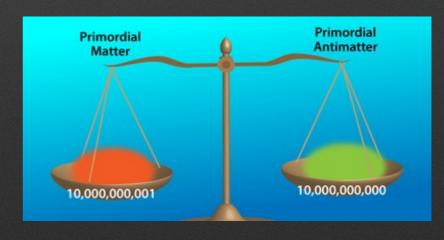
ニュートリノビーム

Open questions

- *Neutrino oscillations →
 "guaranteed" measurements for
 T2K and HK
- *Multi-messenger astronomy with neutrinos is starting now → SK, HK
- *Nature of neutrinos (Dirac or Majorana) and their mass → 0νββ experiments, Katrin, Project-8, cosmology

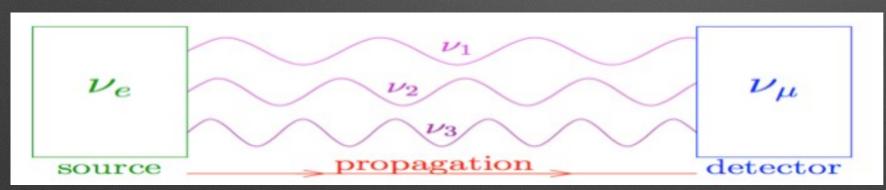




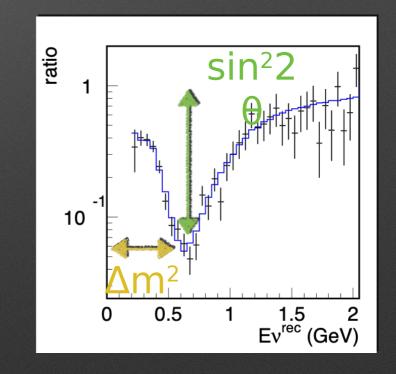


Neutrino oscillations

- *First introduced by Bruno Pontecorvo in 1957
- *Neutrinos are produced in flavor eigenstates v_e , v_μ , v_τ that are linear combination of mass eigenstates v_1 , v_2 , v_3
- *Neutrinos propagate as mass eigenstates
- *At the detection a flavor eigenstate is detected → it can be different from the one that was produced



 $v_{1,2,3}$ travel at different Different ve produced in a speed because they have mixture of $v_{1,2,3}$ mixture of $v_{1,2,3}$ different masses → $\rightarrow \mu$ from ν_{μ} is interference detected

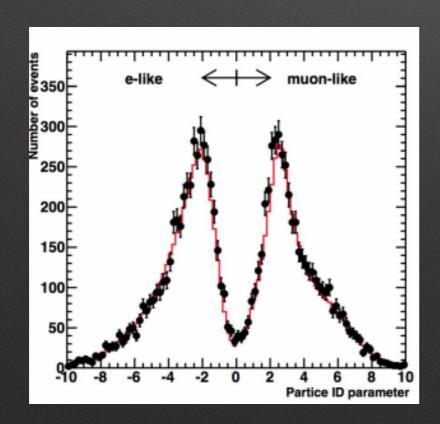


Neutrino oscillation

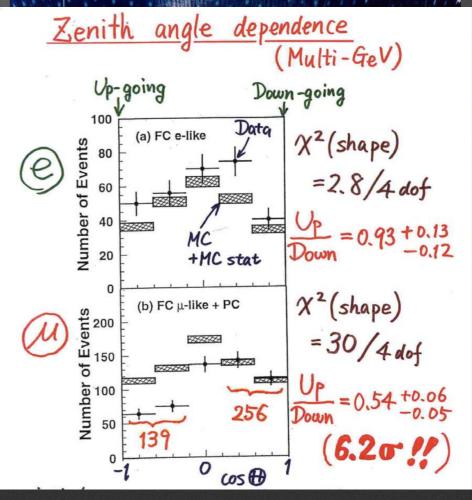
Neutrino oscillation implies massive neutrinos
$$P(\nu_e \to \nu_\mu) = \sin^2(2\theta) \sin^2(\Delta m_{12}^2 L/E)$$

Super-Kamiokande

- *50 kton Water Cherenkov detector
 *~11000 PMTs for ID, ~2000 for OD
- *1000 m underground at Kamioka mine operated since 1996
- *Very good PID capabilities to distinguish between ν_e and ν_μ thanks to shape of Cherenkov ring \rightarrow <1% misidentification probability

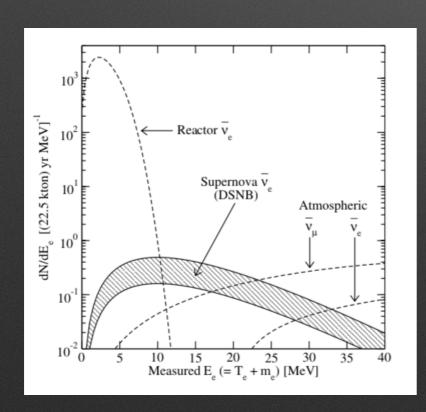


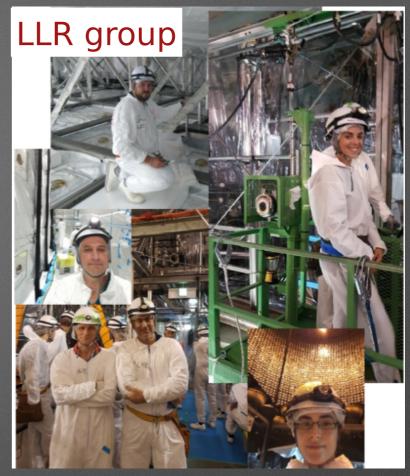


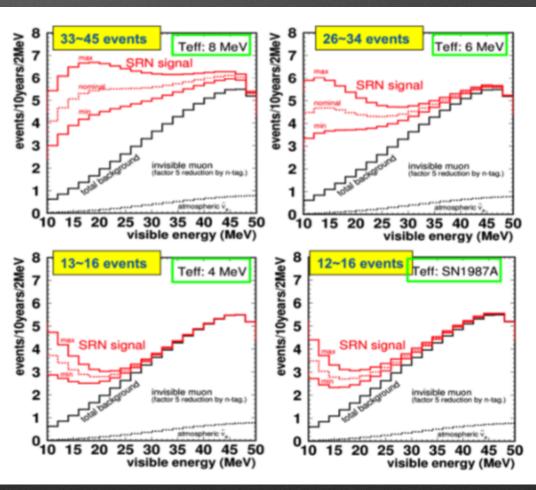


SK-Gd

- *Huge repair work in 2018 to prepare the loading of SK with Gadolinium
- *SK ready to be loaded with Gd in 2020 (0.02% → 0.2% in a second phase)
- *Enhance neutron tagging capability \rightarrow crucial to distinguish ν from $\overline{\nu} \rightarrow$ detect SN-relic antineutrinos from IBD (3-5 events per year are expected)
- *The Gd loading will also be useful for T2K

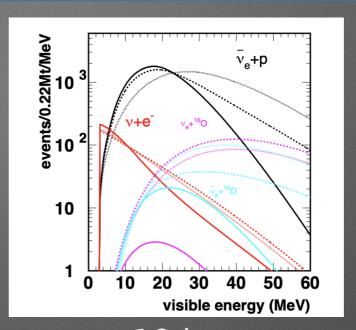


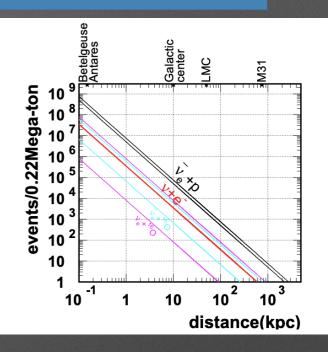




Supernovae Neutrinos

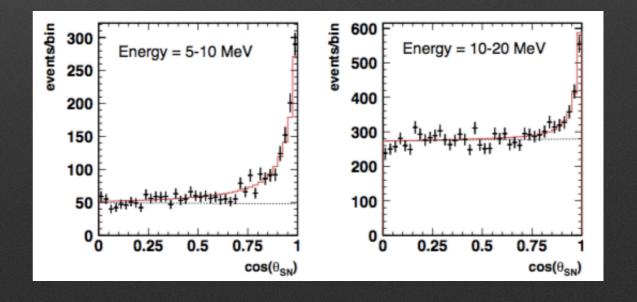
- *Neutrinos carry out ~99% of the total energy released in a SN burst
- *HK will mostly sensitive to $\overline{\nu}$ e through inverse β -decay, but also other channels can be inspected
- *Point to the SN
- *Study energy spectrum and time profile → distinguish between different models for SN explosions
- *Neutrino mass hierarchy determination?





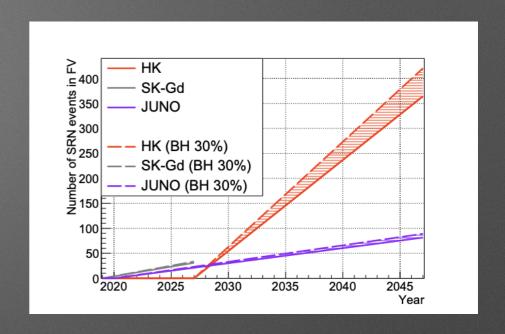
10 kpc

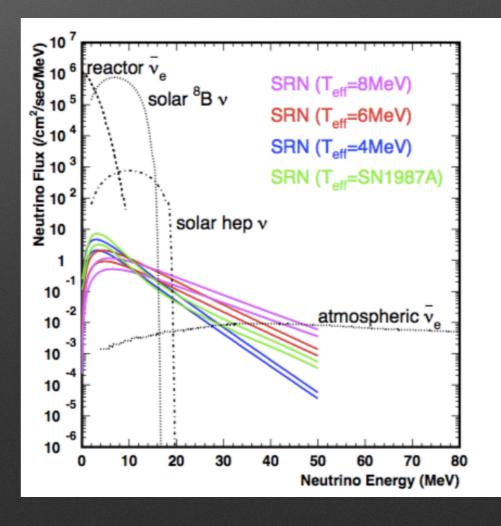
Neutrino source	Single Tank (220 kt Full Volume)	2 Tanks (440 kt Full Volume)
$\bar{\nu}_e + p$	50,000 - 75,000 events	100,000 - 150,000 events
$\nu + e^-$	3,400 - 3,600 events	6,800 - 7,200 events
$\nu_e + ^{16} O$ CC	80 - 7,900 events	160 - 11,000 events
$\bar{\nu}_e + ^{16}O$ CC	660 - 5,900 events	1,300 - 12,000 events
$\nu + e^-$ (Neutronization)	9 - 55 events	17 - 110 events
Total	54,000 - 90,000 events	109,000 - 180,000 events



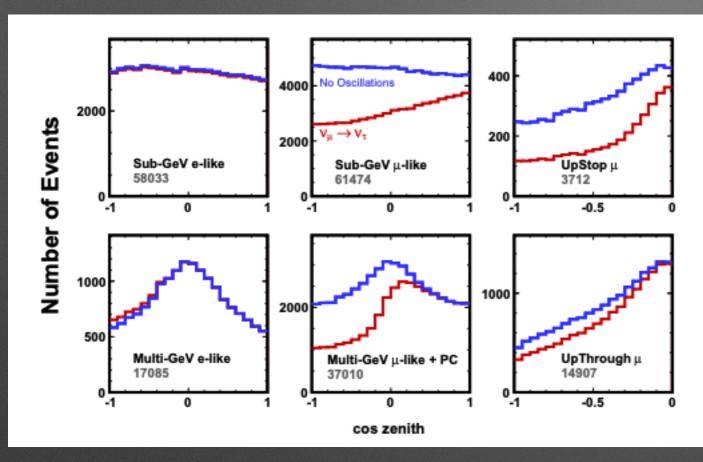
Supernovae Relic Neutrinos

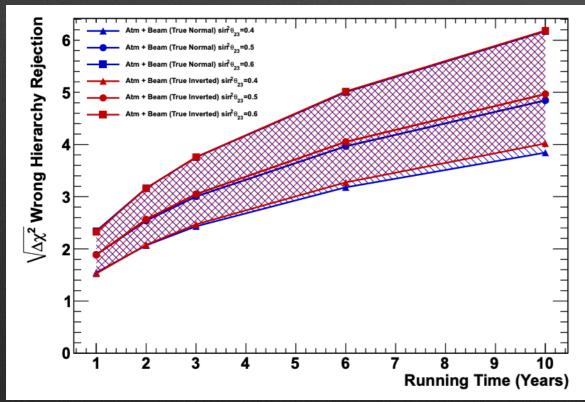
- *Neutrinos produced by all the SN since the beginning of the Universe (SRN)
- *Their detection is the main goal of the Super-Kamiokande upgrade (SK-Gd)
 - *Addition of Gd in SK to tag the neutrons and distinguish $\overline{\nu}_e + p \rightarrow e^- + n$
- *If SRN will be discovered by SK, the large size of HK will allow a detailed study of the history of the Universe through SRN

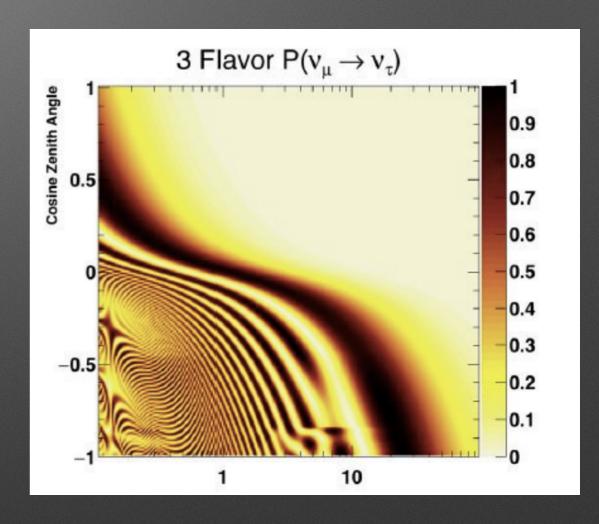




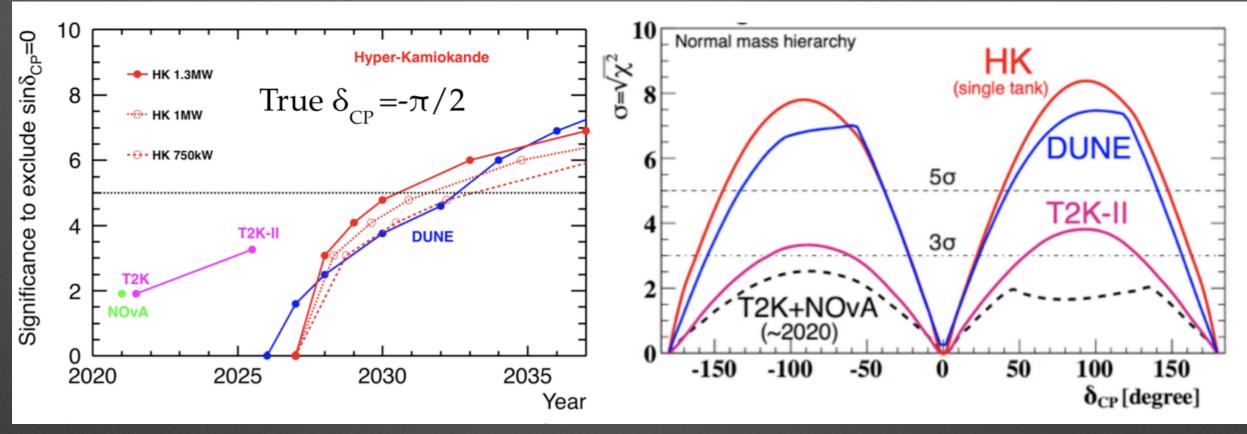
Atmospheric neutrinos

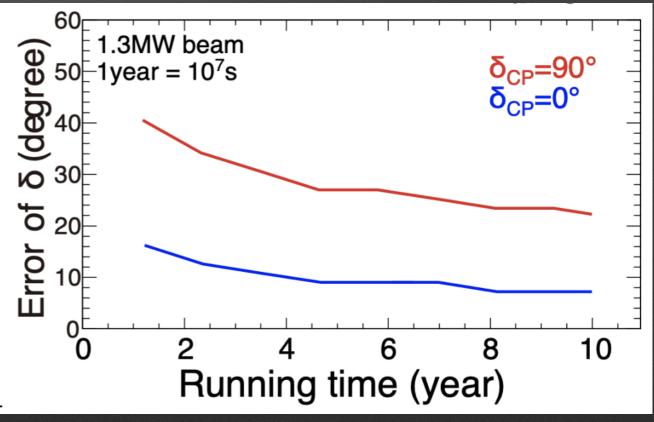




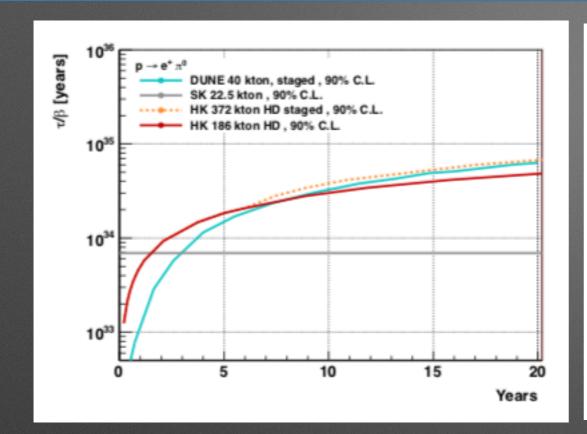


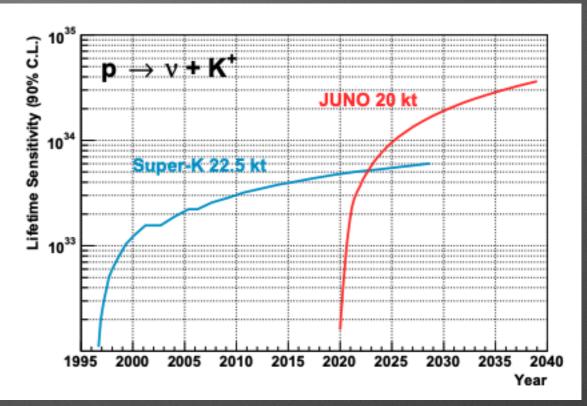
δCP Sensitivities





Sensitivities: proton decay





Trigger	self triggering for each channel
PMT impedance	50Ω
Signal reflection	<0.1%
Discriminator threshold	<0.25PE (well below 1PE)
Processing speed/hit	$<1 \mu s$
(channel dead time)	
Maximum hit rate	>1 MHz for each channel
Charge dynamic range	0.1 to 1250PE (0.2 to 2500 pC)
Charge resolution	RMS $0.05 PE$ for signals below $25 PE$
Timing LSB	$< 0.5 \mathrm{ns}$
Timing resolution	RMS < 0.3 ns at $1PE$
	RMS $< 0.2 \mathrm{ns}$ for signals above 5PE
Power consumption	<1 W per channel