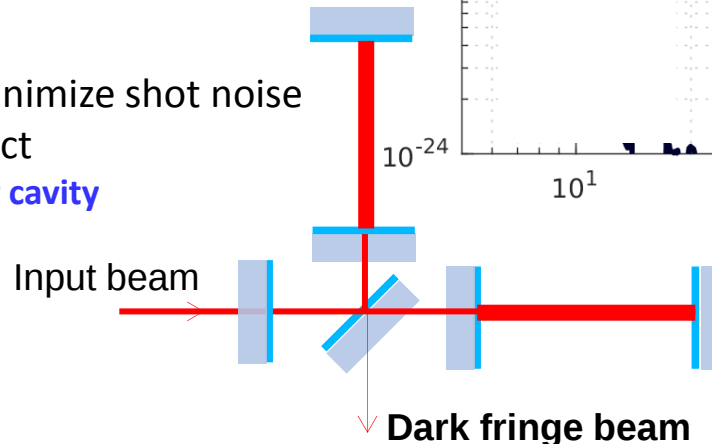


Prospectives IN2P3, GT08

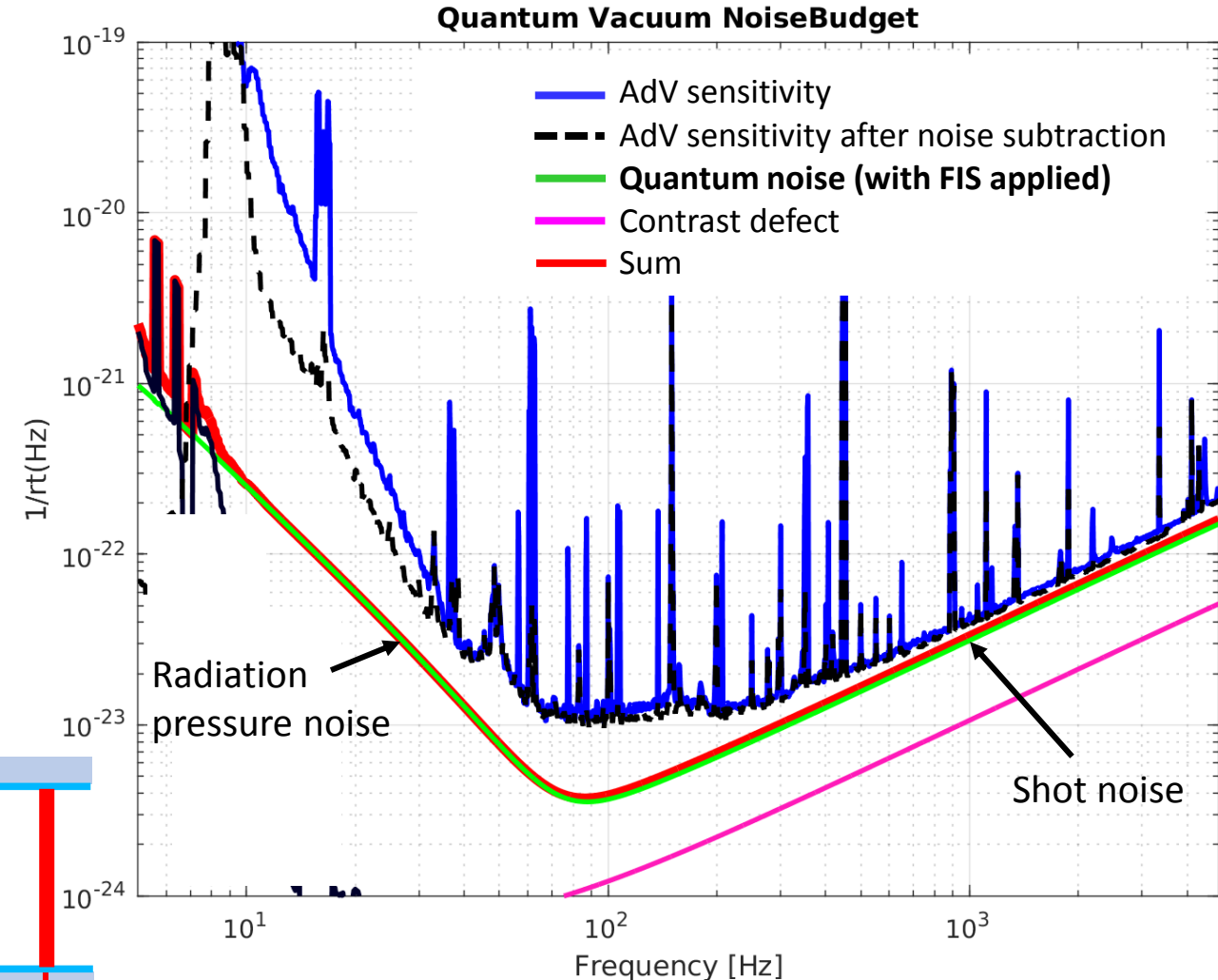
Low loss signal detection

Context

- Sensitivity of GW detector limited by quantum noise:
 - Radiation pressure noise at low frequency
 - Shot noise at high frequency
- Quantum noise can be reduced by injecting a squeezed vacuum state in the interferometer:
 - Frequency Independent Squeezing to reduce shot noise: already implemented in advanced detectors
 - Frequency Dependent Squeezing to enhance the sensitivity at all frequencies: foreseen in AdV+
- Benefit of squeezing limited by losses and signal detection noises
 - This requires:
 - Low losses optics
 - High Quantum Efficiency photodiodes
 - Low noise photodiode electronics
- GW detectors are operating at dark fringe to minimize shot noise
 - Requires an output mode cleaner filtering cavity

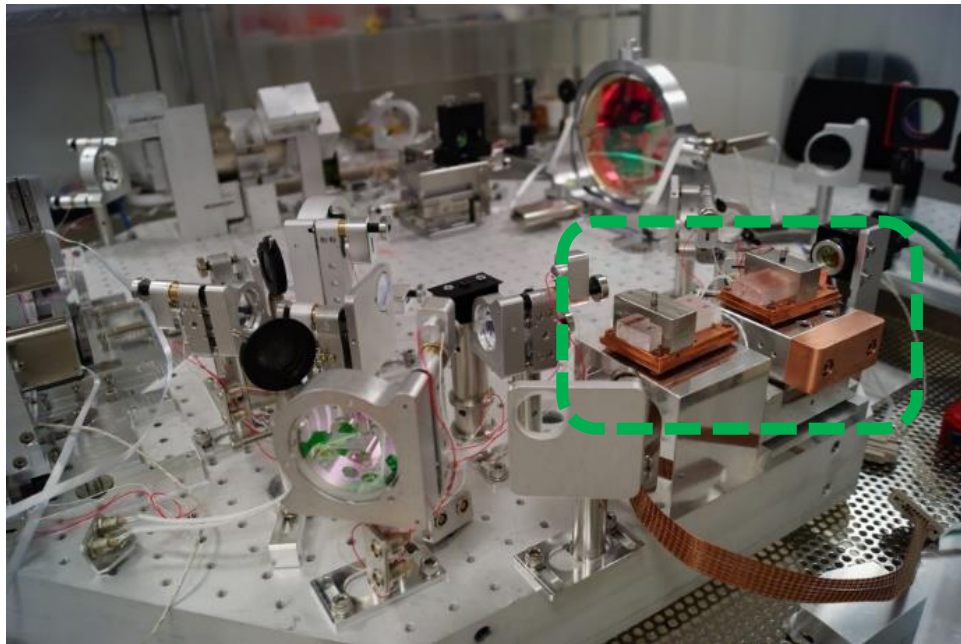


Advanced Virgo sensitivity on Jan 17th, 2020 (03)

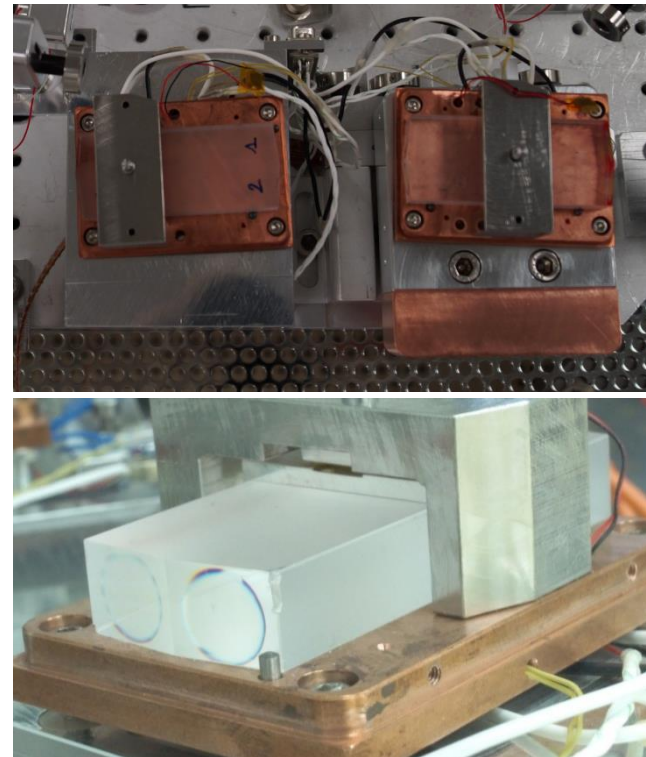


Output Mode Cleaner

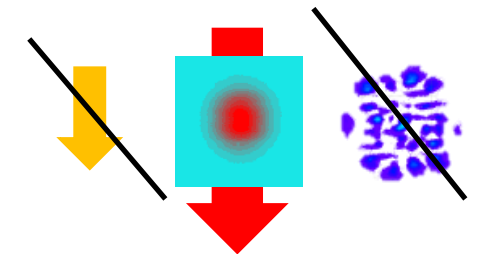
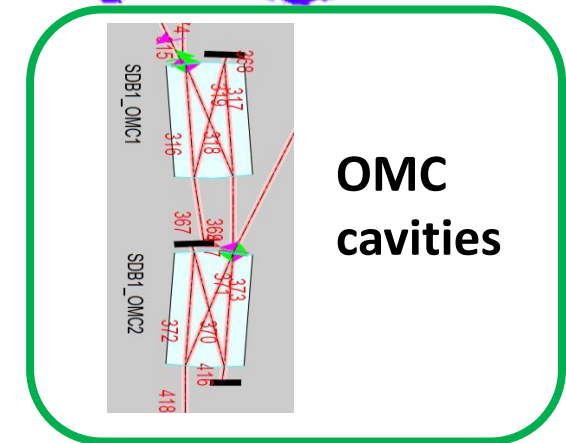
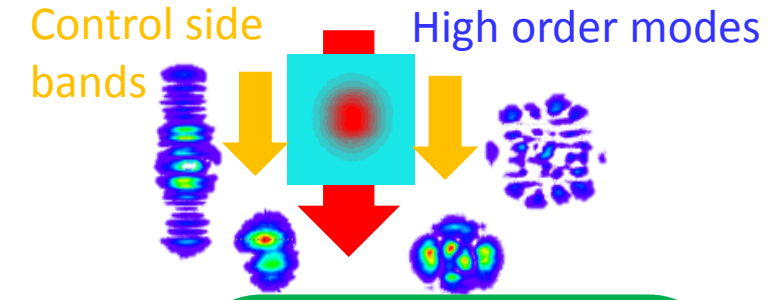
- **Goal:** Filter the Dark Fringe to keep only the component carrying information on Differential Arm Motion → **improve interferometer contrast defect**
- Adv OMC made of 2 cavities of fused silica, in serie:
 - Cavity finesse ≈ 125
 - Beam propagates inside fused silica
 - Cavity length controlled to keep it at resonance of the fundamental mode



24/01/2020



From the interferometer



To the photodiodes³

Losses induced by the OMC

- Internal losses:
 - **Scattering on the reflective surfaces**
proportional to OMC finesse and to (surface roughness)²
→ requires high quality polishing (roughness < 0.3 nm RMS for AdV OMC)
 - **Absorption in OMC substrate**
→ requires low absorption material (fused silica Suprasil 3001 for AdV OMC)
 - **Absorption in surfaces coating**
→ high quality coating performed at LMA
- Mismatching losses between incident beam and OMC cavity:
 - **Mode mismatch** → requires a telescope
 - **Misalignment**
 - **Polarization mismatch**
- Mismatching losses between the two OMC cavities

Proposed improvements for a lower loss signal detection

- Replace the 2 OMC cavities by a single OMC cavity of higher finesse (≈ 1000)
 - Will get rid of mismatching losses between the 2 cavities
 - Must learn how to control the high finesse cavity with good precision
- Improve polishing: roughness below 0.1 nm RMS → will mitigate scattering losses
- Implement more actuators on the suspended bench optics to allow a precise mode matching and polarization tuning when the bench is in vacuum
- Implement an automatic alignment based on new quadrant sensors to correct OMC alignment drifts during interferometer operation
- Explore lower absorption materials, better coating
- Explore an evolution of the OMC design towards an empty cavity (beam propagating inside vacuum)
- Develop a lower noise preamplifier for the high quantum efficiency photodiodes
- Choice of photodiode technology to be revised in view of Einstein Telescope and new wavelengths