

Initial fluctuations and anisotropies in heavy-ion collisions

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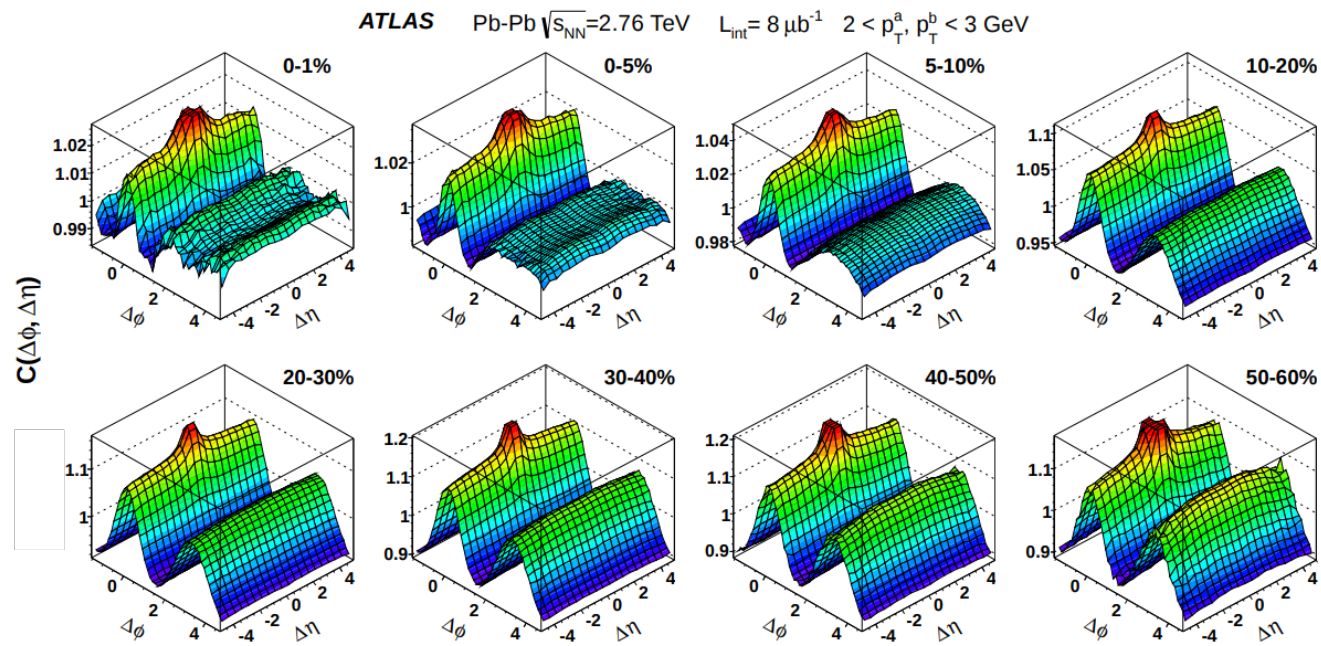
based on:

Albacete, Guerrero-Rodriguez, CM, JHEP 1901 (2019) 073
Giacalone, Guerrero-Rodriguez, Luzum, CM and Ollitrault, PRC 100 (2019) 024905
Gelis, Giacalone, Guerrero-Rodriguez, CM and Ollitrault, arXiv:1907.10948

Outline

- The fluid paradigm in heavy-ion collisions
final-state momentum anisotropies from
initial-state spatial anisotropies
- Initial eccentricities from the Glasma
how to calculate the initial spatial anisotropies:
from the energy momentum tensor and its fluctuations
- A new picture of initial-state fluctuations
ab-initio description free from the Monte Carlo Glauber Ansatz

Flow in heavy-ion collisions

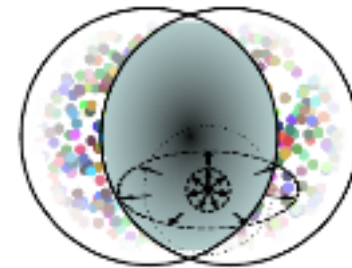


Two ingredients needed for flow

flow is an **initial spatial anisotropy** turned into a momentum anisotropy by the **hydrodynamic expansion** of the medium

$$\text{final-state harmonic} \longleftarrow v_n = \kappa_n \varepsilon_n \longrightarrow \text{initial-state harmonic}$$

v_2 has two components: a geometric one and one due to fluctuations (the geometric component vanishes in central collisions)

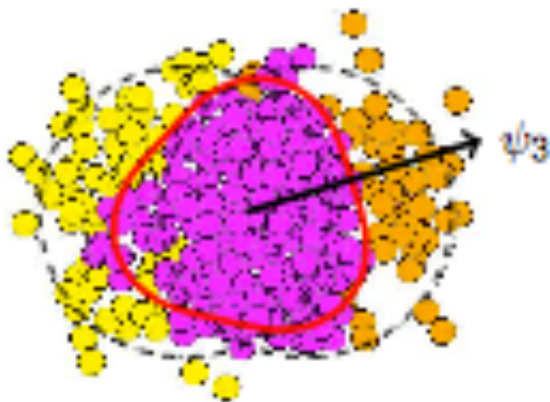
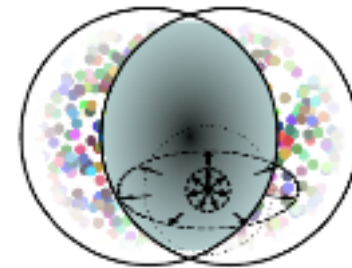


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v_3 is only due to fluctuations

The eccentricity harmonics

How do we calculate the initial anisotropy?

[Teaney, Yan [1010.1876](#)]

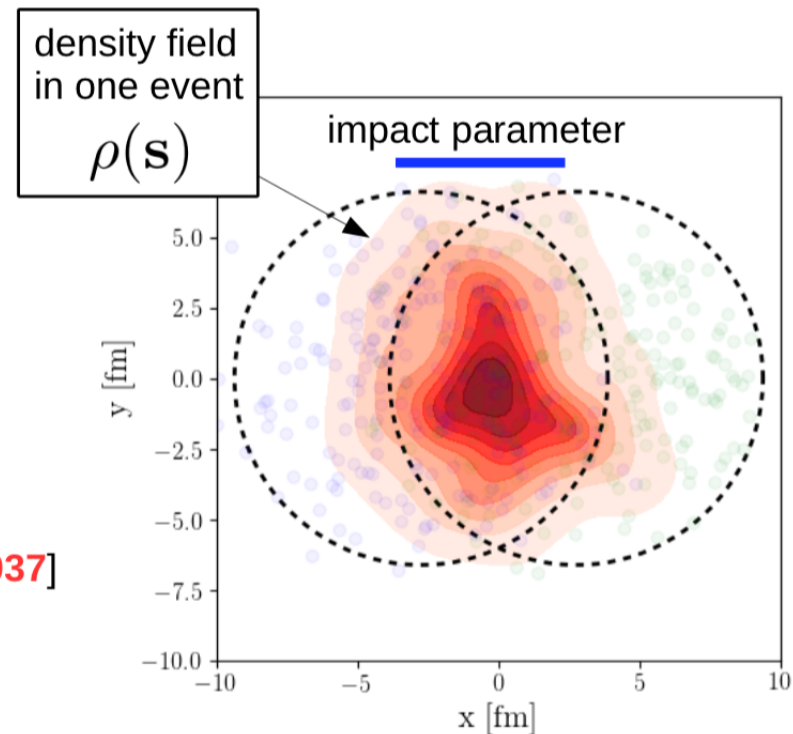
$$\varepsilon_n \equiv \frac{\int_{\mathbf{s}} \mathbf{s}^n \rho(\mathbf{s})}{\int_{\mathbf{s}} |\mathbf{s}|^n \rho(\mathbf{s})}$$

$\mathbf{s} = x + iy$

Origin of anisotropy:

n=2
elliptic flow $\rightarrow$ **geometry + fluctuations**
[PHOBOS Collaboration [nucl-ex/0610037](#)]

n=3
triangular flow $\rightarrow$ **fluctuations only**
[Alver, Roland [1003.0194](#)]



The theoretical input is a model for $\rho(\mathbf{s})$ and its fluctuations.

Relevant averaged quantities

- averaged quantities relevant for experiments:

$$\begin{array}{lll} \text{geometry + fluctuations} & \longleftarrow \sqrt{\langle v_2^2 \rangle} = v_2\{2\} = \kappa_2 \varepsilon_2\{2\} & \\ & \sqrt[4]{2\langle v_2^2 \rangle^2 - \langle v_2^4 \rangle} = v_2\{4\} = \kappa_2 \varepsilon_2\{4\} & \longrightarrow \text{geometry only} \\ \text{fluctuations only} & \longleftarrow \sqrt{\langle v_3^2 \rangle} = v_3\{2\} = \kappa_3 \varepsilon_3\{2\} & \end{array}$$

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 \end{array}$$

- a 10-year old prescription to compute $\rho(s)$:

- start with the 'Glauber Monte Carlo Ansatz'

(sample nucleons according to the Woods-Saxon distribution)

- propose some mechanism (more or less motivated by physics) to convert nucleons/partons into a smooth map of energy density (e.g. through interactions)

instead, can we describe experimental data using the correlation functions of the primordial energy density as (only) input ?

Our strategy

we follow Blaizot, Bronjowski, Ollitrault (2014)

$$\rho(\mathbf{s}) = \langle \rho(\mathbf{s}) \rangle + \delta\rho(\mathbf{s}), \quad \langle \rho(\mathbf{s}) \rangle \gg \delta\rho(\mathbf{s})$$

Connected 2-point function:
(i.e. fluctuations) $S(\mathbf{s}_1, \mathbf{s}_2) \equiv \langle \delta\rho(\mathbf{s}_1)\delta\rho(\mathbf{s}_2) \rangle = \langle \rho(\mathbf{s}_1)\rho(\mathbf{s}_2) \rangle - \langle \rho(\mathbf{s}_1) \rangle \langle \rho(\mathbf{s}_2) \rangle$

Perturbative expansion of the anisotropy: $\varepsilon_n \equiv \frac{\int_{\mathbf{s}} \mathbf{s}^n \rho(\mathbf{s})}{\int_{\mathbf{s}} |\mathbf{s}|^n \rho(\mathbf{s})}$

To first nontrivial order:

$$\langle \varepsilon_2 \varepsilon_2^* \rangle = \varepsilon_2 \{2\}^2 = \sigma^2 + \bar{\varepsilon}_2^2$$

$$\text{fluctuations } \sigma^2 = \frac{\int_{\mathbf{s}_1, \mathbf{s}_2} (\mathbf{s}_1)^2 (\mathbf{s}_2^*)^2 S(\mathbf{s}_1, \mathbf{s}_2)}{\left(\int_{\mathbf{s}} |\mathbf{s}|^2 \langle \rho(\mathbf{s}) \rangle \right)^2}$$

$$\text{geometry } \bar{\varepsilon}_2 = \frac{\int_{\mathbf{s}} \mathbf{s}^2 \langle \rho(\mathbf{s}) \rangle}{\int_{\mathbf{s}} |\mathbf{s}|^2 \langle \rho(\mathbf{s}) \rangle} \sim \varepsilon_2 \{4\}$$

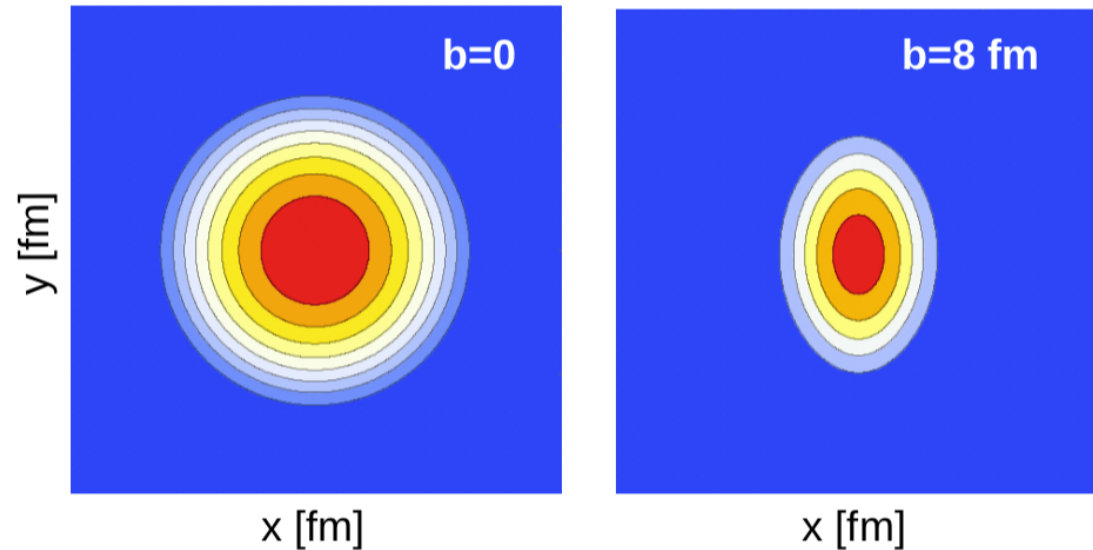
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fluctuations
only!

r.m.s. spatial anisotropy can be mapped to experimental data
we need the 1-point and 2-point correlators

What is needed ?

- $\langle \rho(s) \rangle$
The average density.



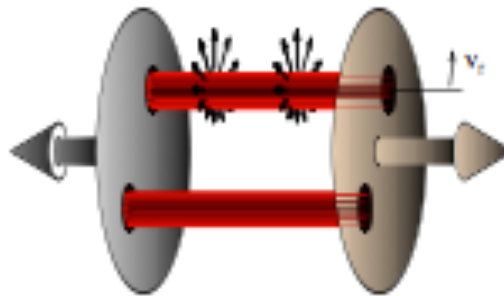
- The fluctuations around the average.

$$S(s_1, s_2) = \langle \rho(s_1) \rho(s_2) \rangle - \langle \rho(s_1) \rangle \langle \rho(s_2) \rangle$$

in particular the integral $\xi(s) = \int_{\mathbf{r}} S\left(s + \frac{\mathbf{r}}{2}, s - \frac{\mathbf{r}}{2}\right)$

→ compute the 1-point and 2-point energy correlators

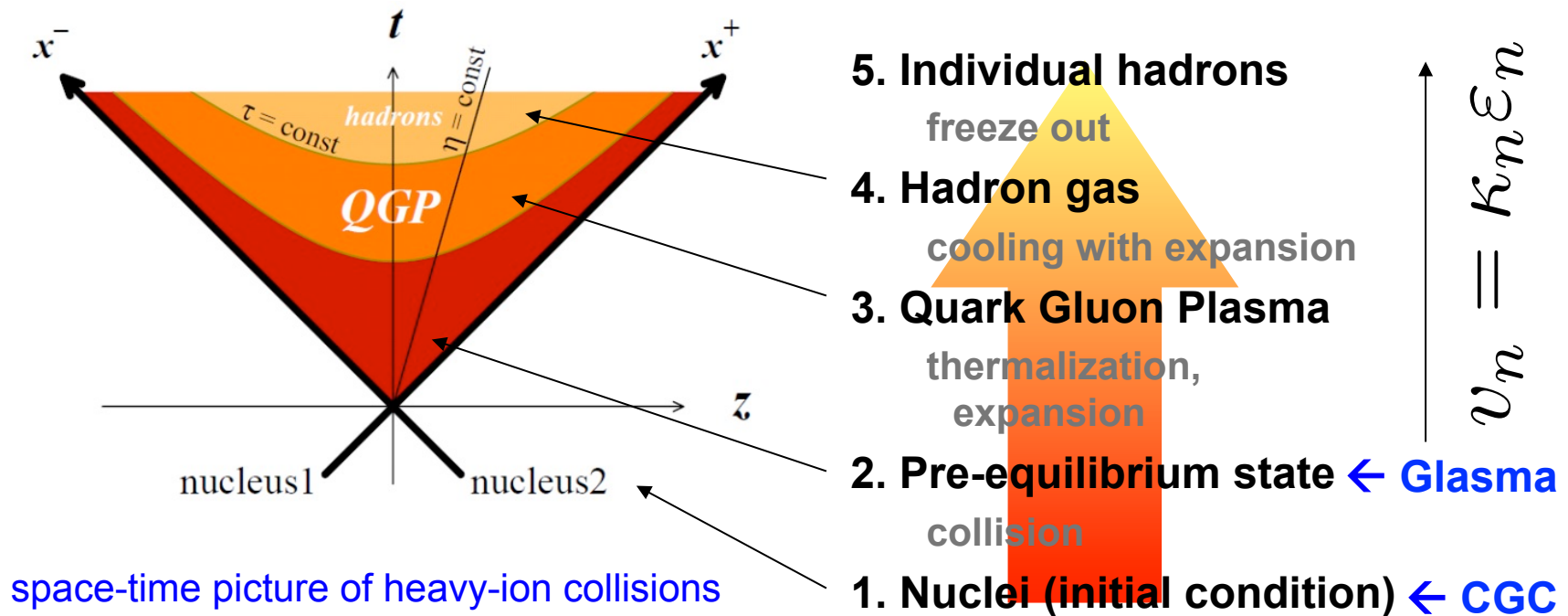
Initial eccentricities from the Glasma



Albacete, Guerrero-Rodriguez, CM, JHEP 1901 (2019) 073

What is the Glasma ?

the initial strong color field created after the collision

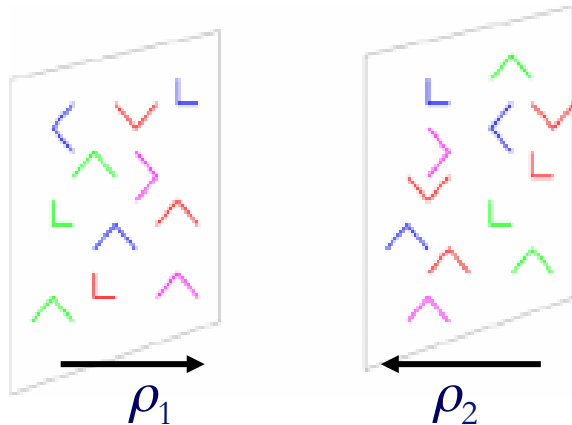


space-time picture of heavy-ion collisions

CGC = Color Glass Condensate

The collision of two CGCs

- the initial condition for the time evolution in heavy-ion collisions



before the collision:

$$J^\mu = \delta^{\mu+} \delta(x^-) \rho_1(x_\perp) + \delta^{\mu-} \delta(x^+) \rho_2(x_\perp)$$

$$\rho_1 \sim 1/g \qquad \rho_2 \sim 1/g$$

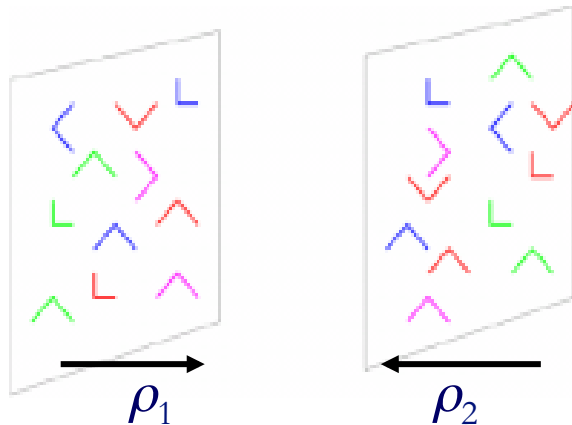
the distributions of ρ contain the small- x
evolution of the nuclear wave functions

$$|\Phi_{x_1}[\rho_1]|^2 \quad |\Phi_{x_2}[\rho_2]|^2$$

$\rho(x_\perp) = -\nabla^2 \alpha(x_\perp)$ denotes the color charge which generates the field

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- after the collision

the gluon field is a complicated function of the two classical color sources

the field decays, once it is no longer strong (classical)

a particle description is again appropriate

“strong-field” QCD factorization

- solve Yang-Mills equations

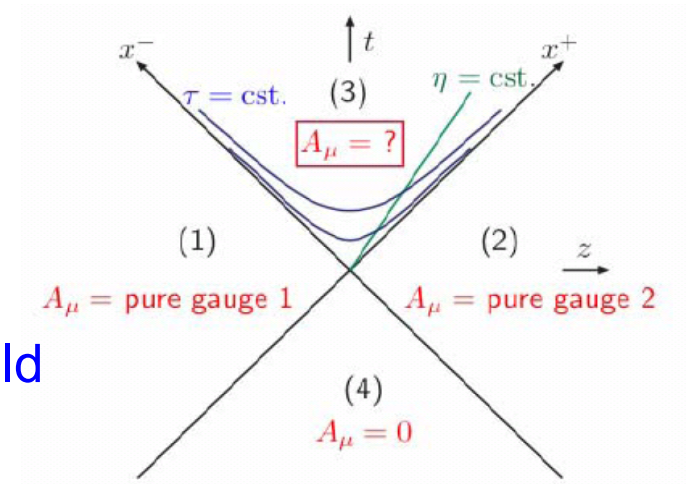
$$[D_\mu, F^{\mu\nu}] = J^\nu \longrightarrow \mathcal{A}_\mu[\rho_1, \rho_2]$$

this is done numerically (it can be done analytically in the p+A case)

- express observables in terms of the field

determine $O[\mathcal{A}_\mu]$, in general a non-linear function of the sources

e.g. for this talk
$$T^{\mu\nu} = \frac{1}{4} g^{\mu\nu} F^{\lambda\sigma} F_{\lambda\sigma} - F^{\mu\lambda} F_\lambda^\nu$$



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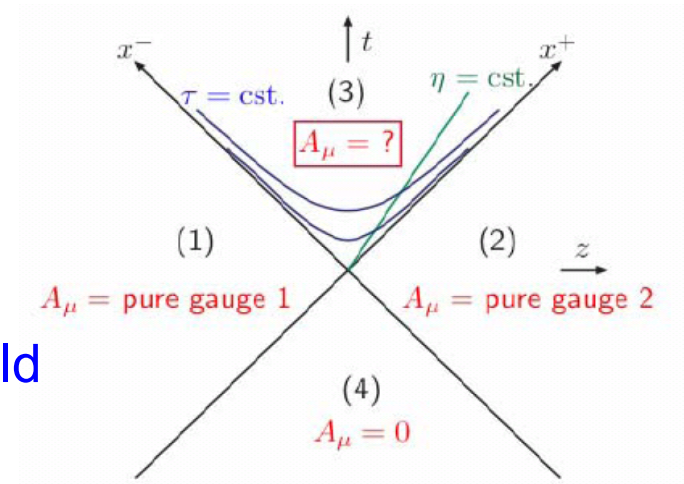
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- perform the averages over the color charge densities

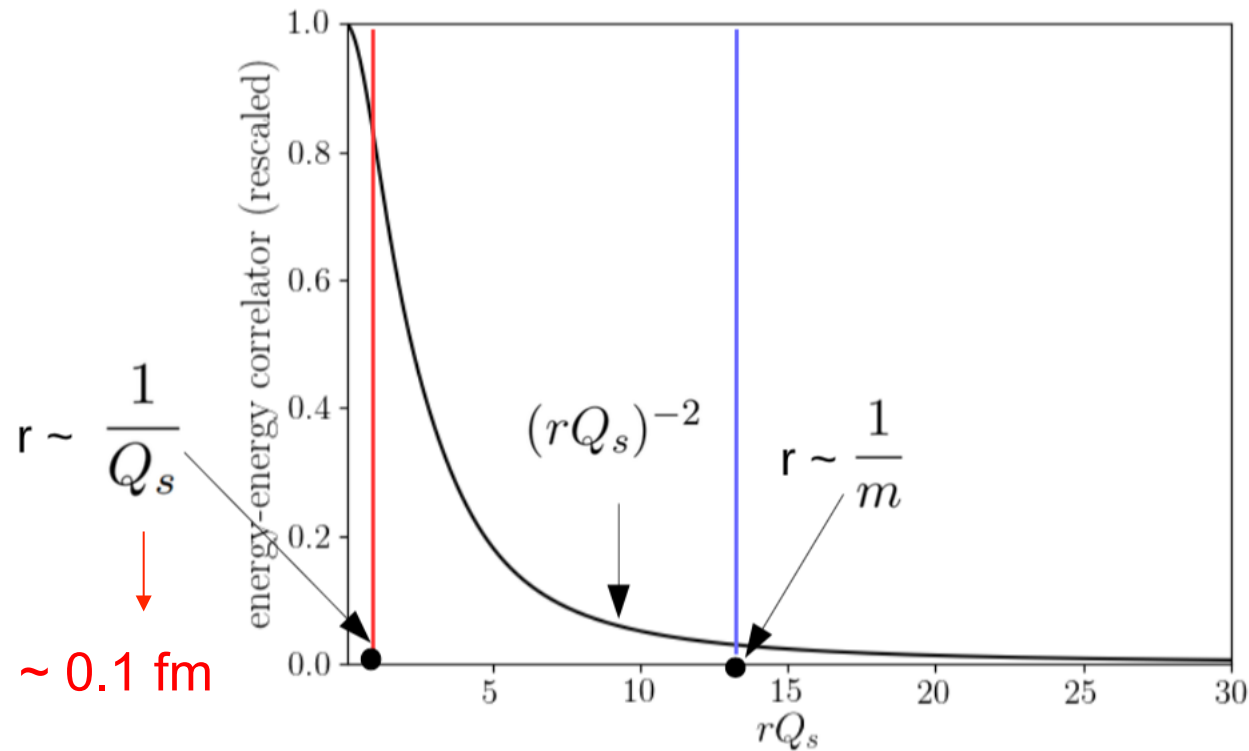
$$\langle O \rangle = \int D\rho_1 D\rho_2 |\Phi_{x_1}[\rho_1]|^2 |\Phi_{x_2}[\rho_2]|^2 O[\mathcal{A}_\mu]$$

we shall use the MV model for $|\Phi_{x_1}[\rho_1]|^2$ and $|\Phi_{x_2}[\rho_2]|^2$

→ each nucleus is characterized by its saturation scale $Q_s^2(\mathbf{s}) \propto T(\mathbf{s})$



Relevant features of $S(s_1, s_2)$

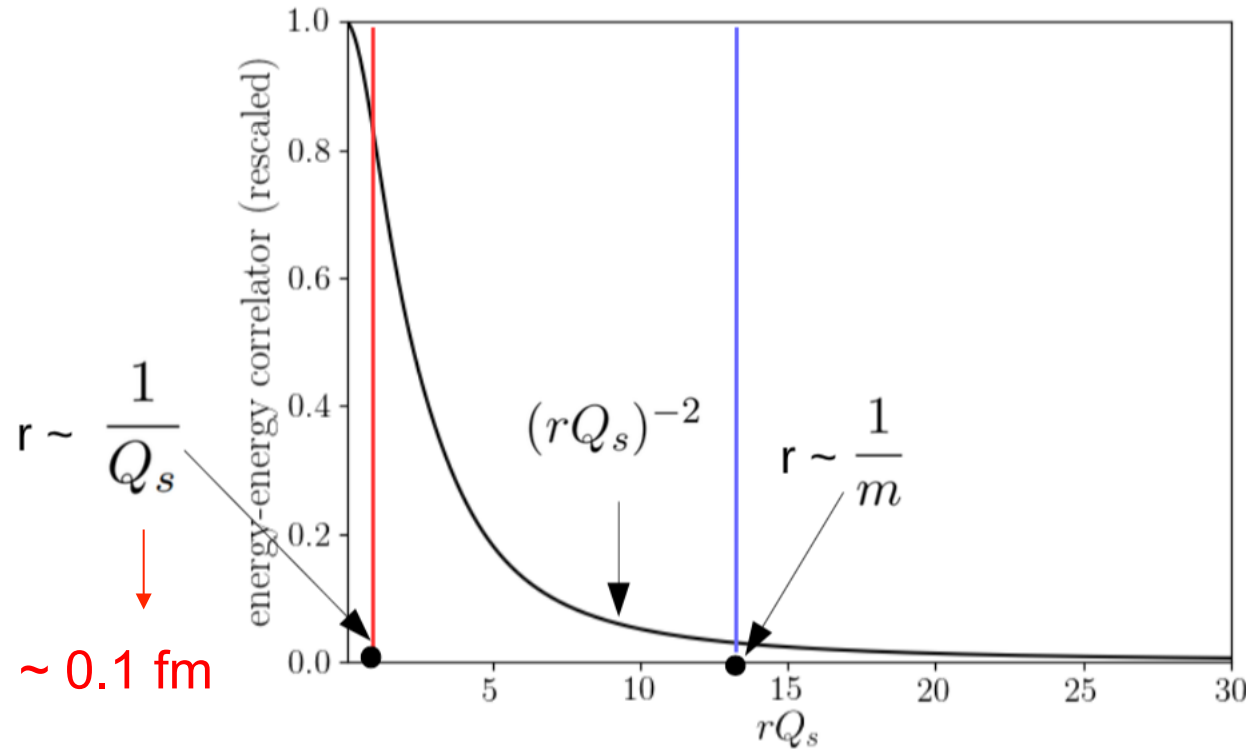


system size
 $\sim 10 \text{ fm}$

$$\frac{1}{Q_s} \ll \frac{1}{m} \ll R$$

scale of color
 neutrality
 $\sim 1 \text{ fm}$

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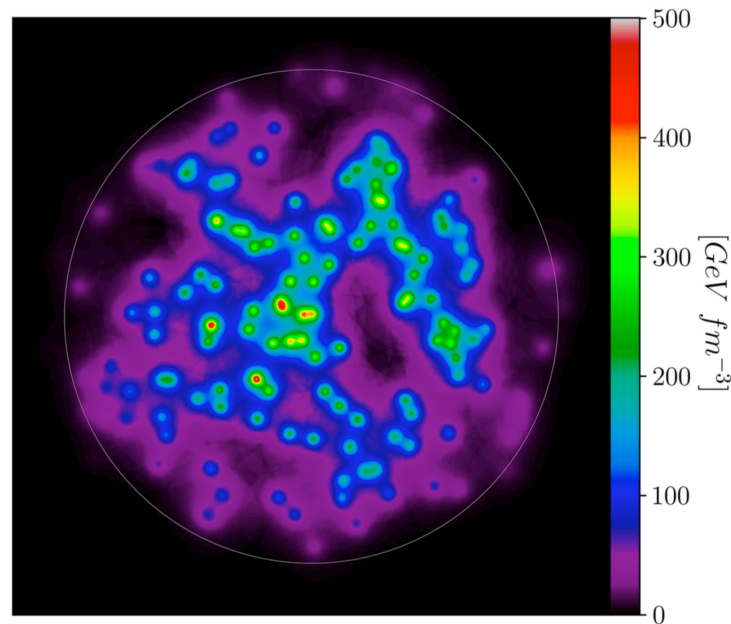
$$\frac{1}{Q_s} \ll \frac{1}{m} \ll R$$

scale of color
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 $\sim 1 \text{ fm}$

$$\langle \rho(\mathbf{s}) \rangle = \frac{4}{3g^2} Q_A^2(\mathbf{s}) Q_B^2(\mathbf{s})$$

$$\xi(\mathbf{s}) = \frac{16\pi}{9g^4} Q_A^2(\mathbf{s}) Q_B^2(\mathbf{s}) \left(Q_A^2(\mathbf{s}) \ln \left(1 + \frac{Q_B^2(\mathbf{s})}{m^2} \right) + Q_B^2(\mathbf{s}) \ln \left(1 + \frac{Q_A^2(\mathbf{s})}{m^2} \right) \right)$$

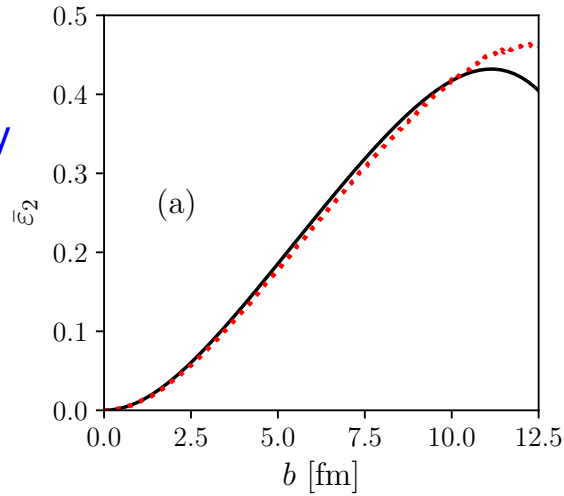
Results and data comparisons



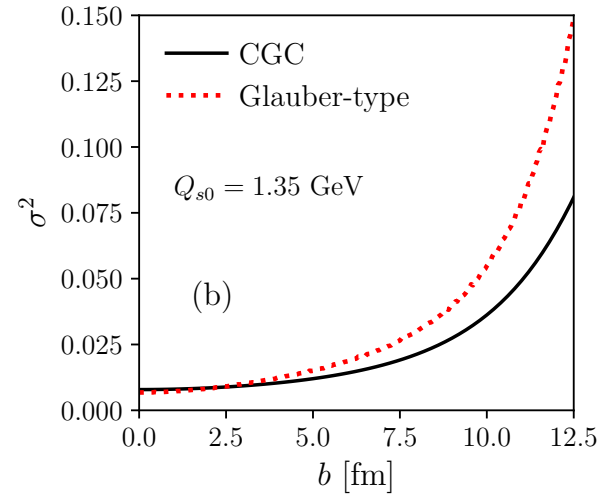
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Gelis, Giacalone, Guerrero-Rodriguez, CM and Ollitrault, arXiv:1907.10948

Comparison to MC Glauber models

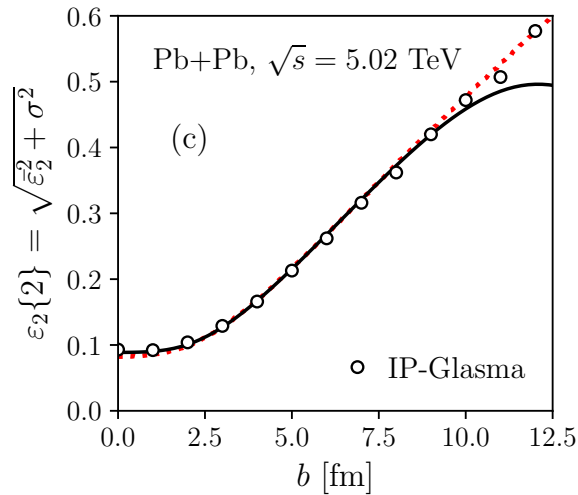
mean
anisotropy



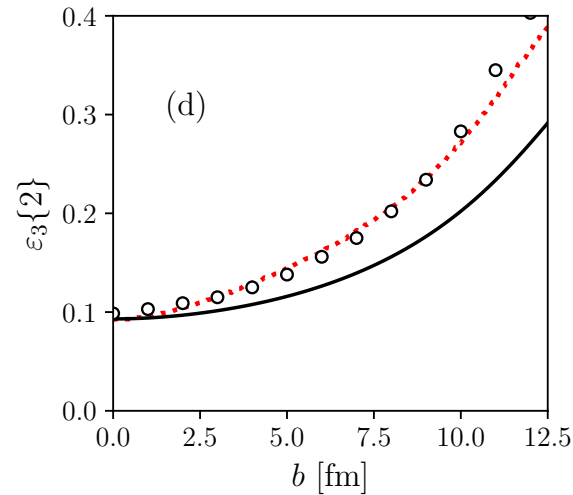
ε_2
fluctuations



rms ε_2



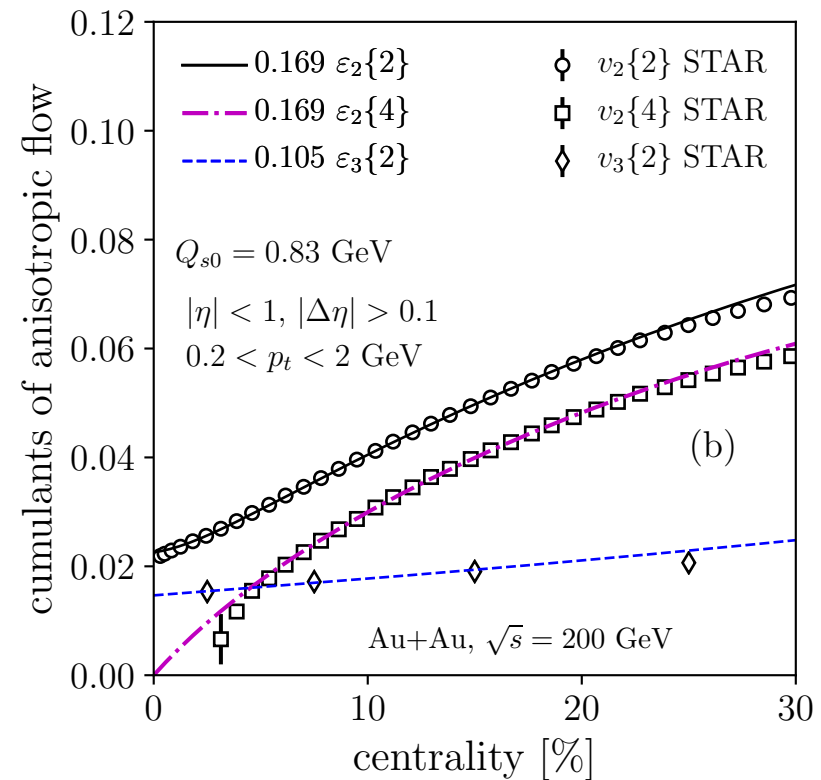
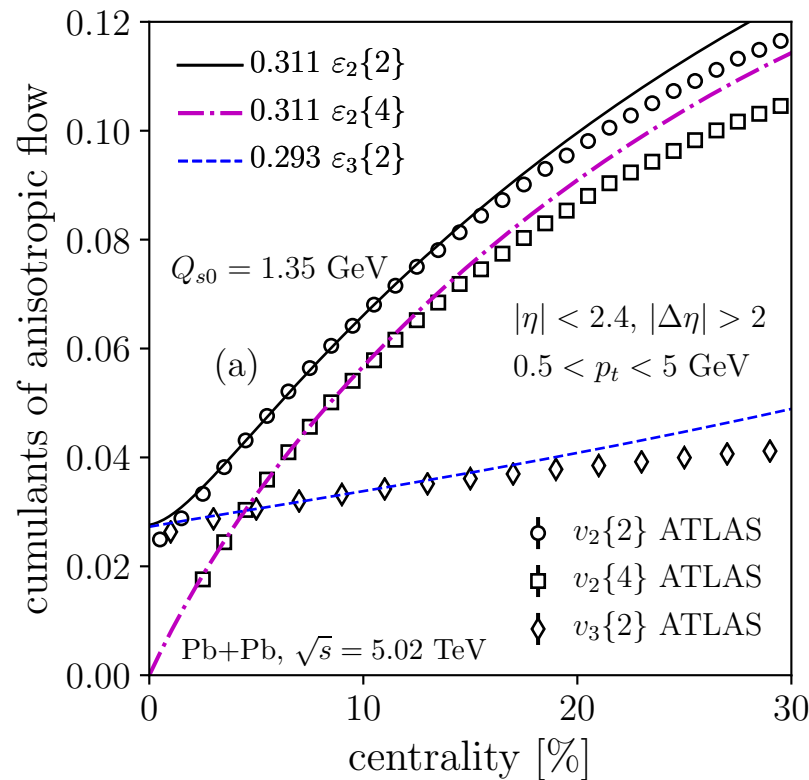
rms ε_3



IP Glasma ~ MC Glauber, its fluctuations are dominated by the nucleon position sampling
we only have fluctuations of the local color charge

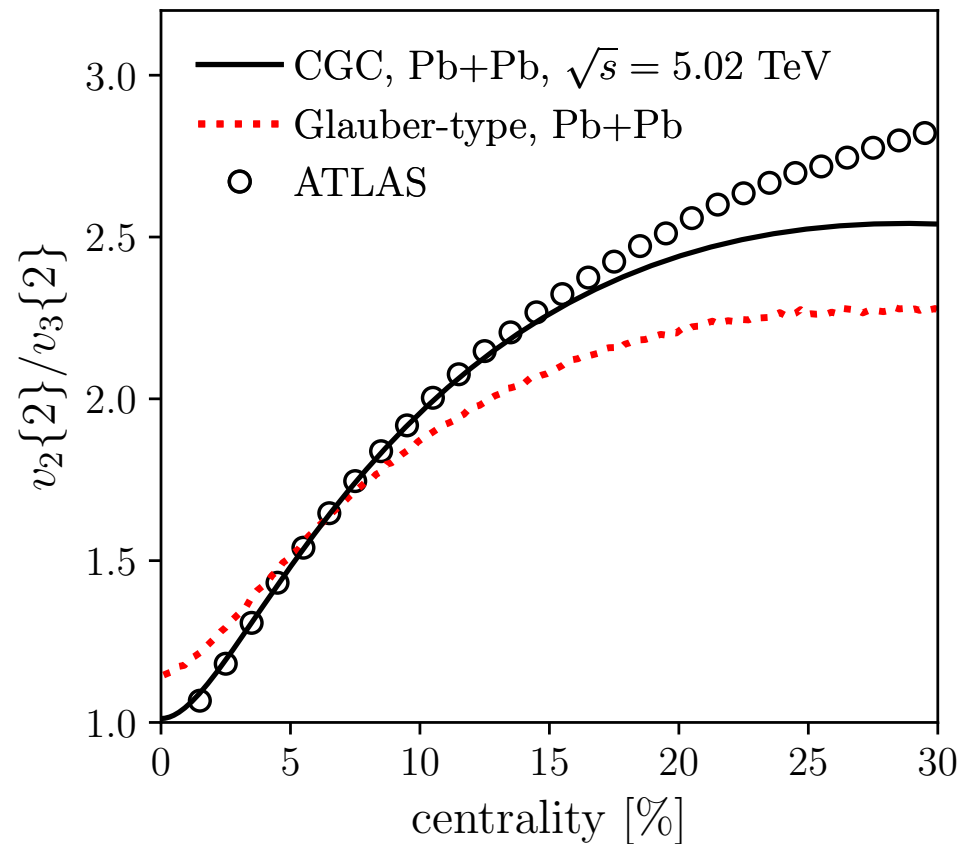
Comparison to data

- $v_2\{4\}$ simply probes the average geometry, we use it to fix K_2
- then, with a reasonable Q_s value, we can reproduce the fluctuations in $v_2\{2\}$



- the response coefficients are compatible with state-of-the-art hydro simulations
- $Q_s(\text{LHC}) > Q_s(\text{RHIC})$ explains more eccentricity fluctuations at RHIC

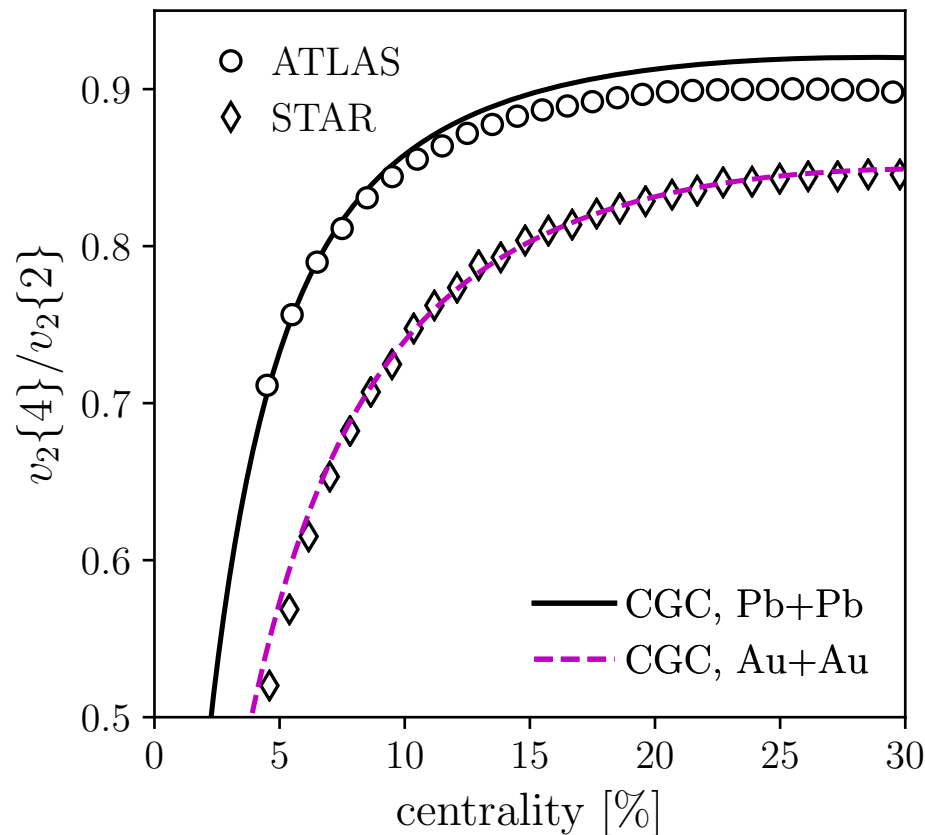
Triangular flow



we solve a long-standing problem of hydro-to-data comparisons:
the ratio $v_2\{2\} / v_3\{2\}$ grows quickly with centrality

Energy dependence

fluctuations produce the splitting between the $v_2\{4\}$ and $v_2\{2\}$
data indicate that they are larger at RHIC energies



$Q_s(\text{LHC}) \sim 1.3 \text{ GeV}$

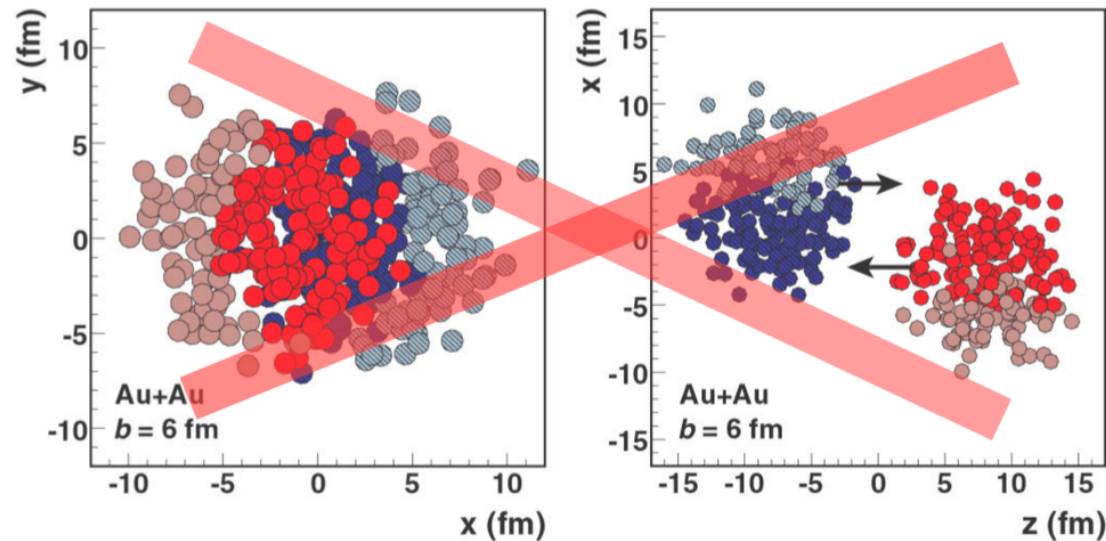
$Q_s(\text{RHIC}) \sim 0.8 \text{ GeV}$

this is compatible with what we expect
from QCD evolution towards high energies

$$\frac{Q_s(x_1)}{Q_s(x_2)} = \left(\frac{\sqrt{s_2}}{\sqrt{s_1}} \right)^{\sim 0.28}$$

by contrast, MC Glauber-type calculations
do not make any specific predictions
for the $v_2\{4\} / v_2\{2\}$ ratio

Conclusions



ab-initio description free from the Monte Carlo Glauber Ansatz:

- No random sampling of nucleons
- No ad hoc prescriptions about the deposition of energy
- Non perturbative physics only through the mass parameter