

Exotic Nuclei and Astrophysics:

First spectroscopy of ^{110}Zr

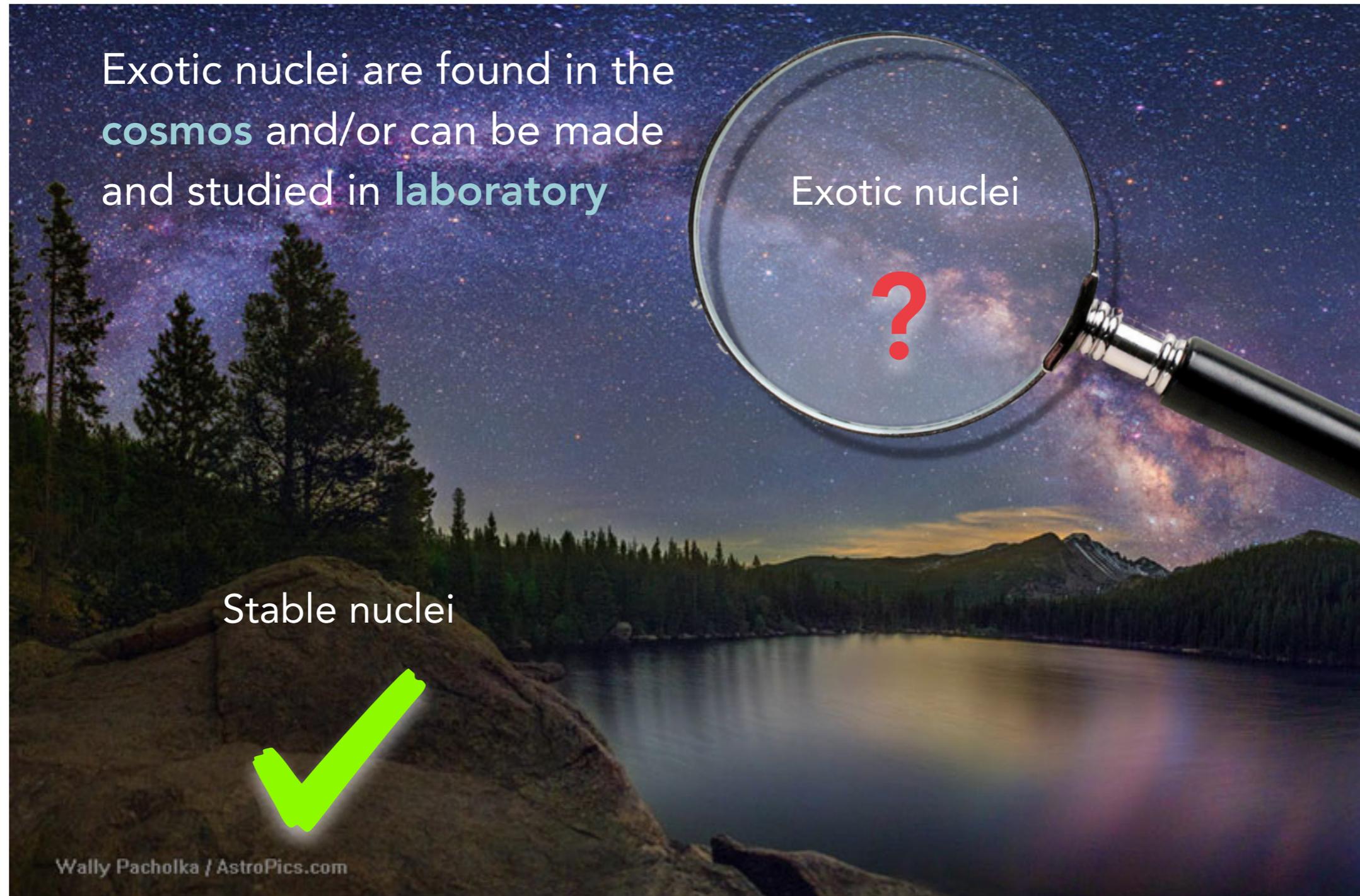
Nancy Paul

PhD. Advisors: Alexandre Obertelli and Anna Corsi

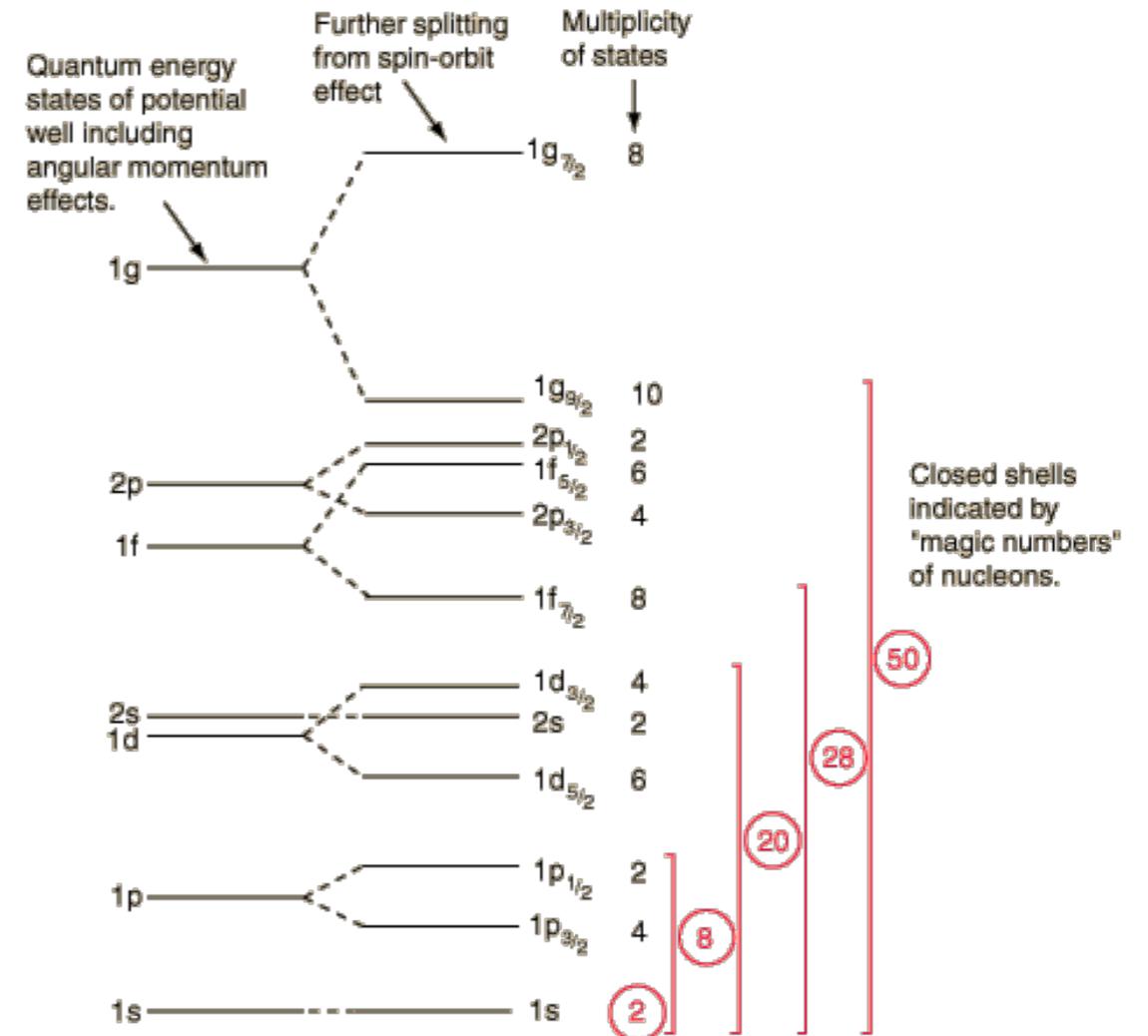
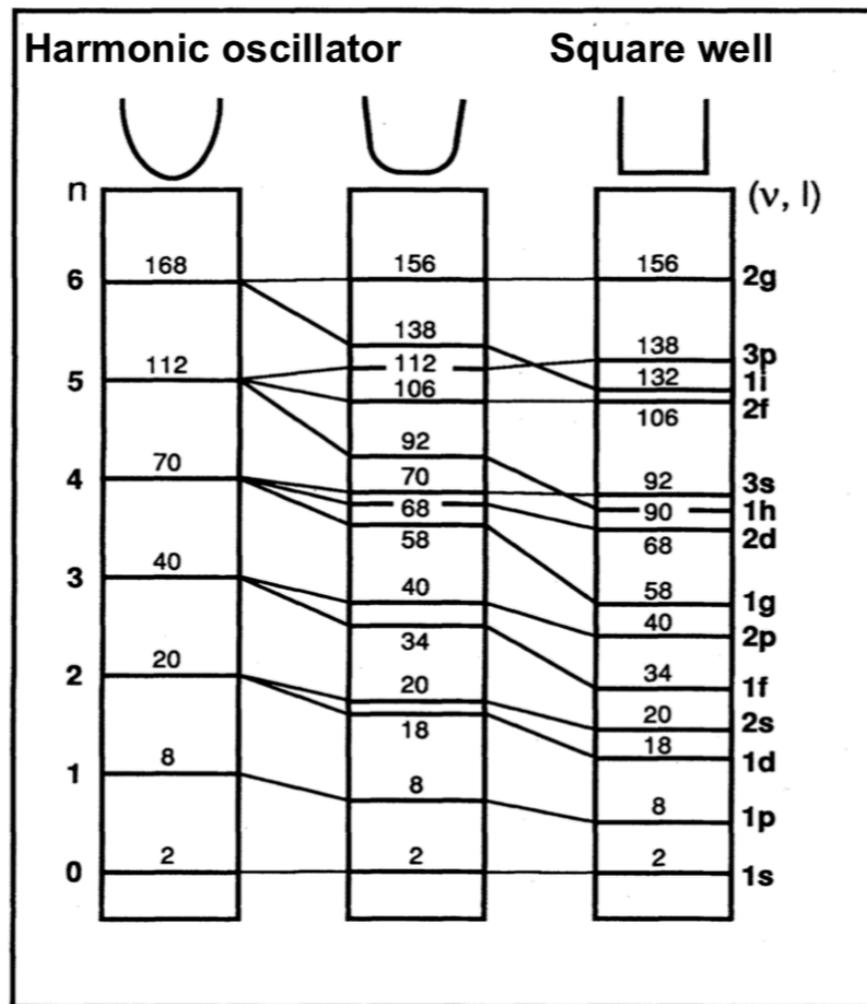
November 29, 2019





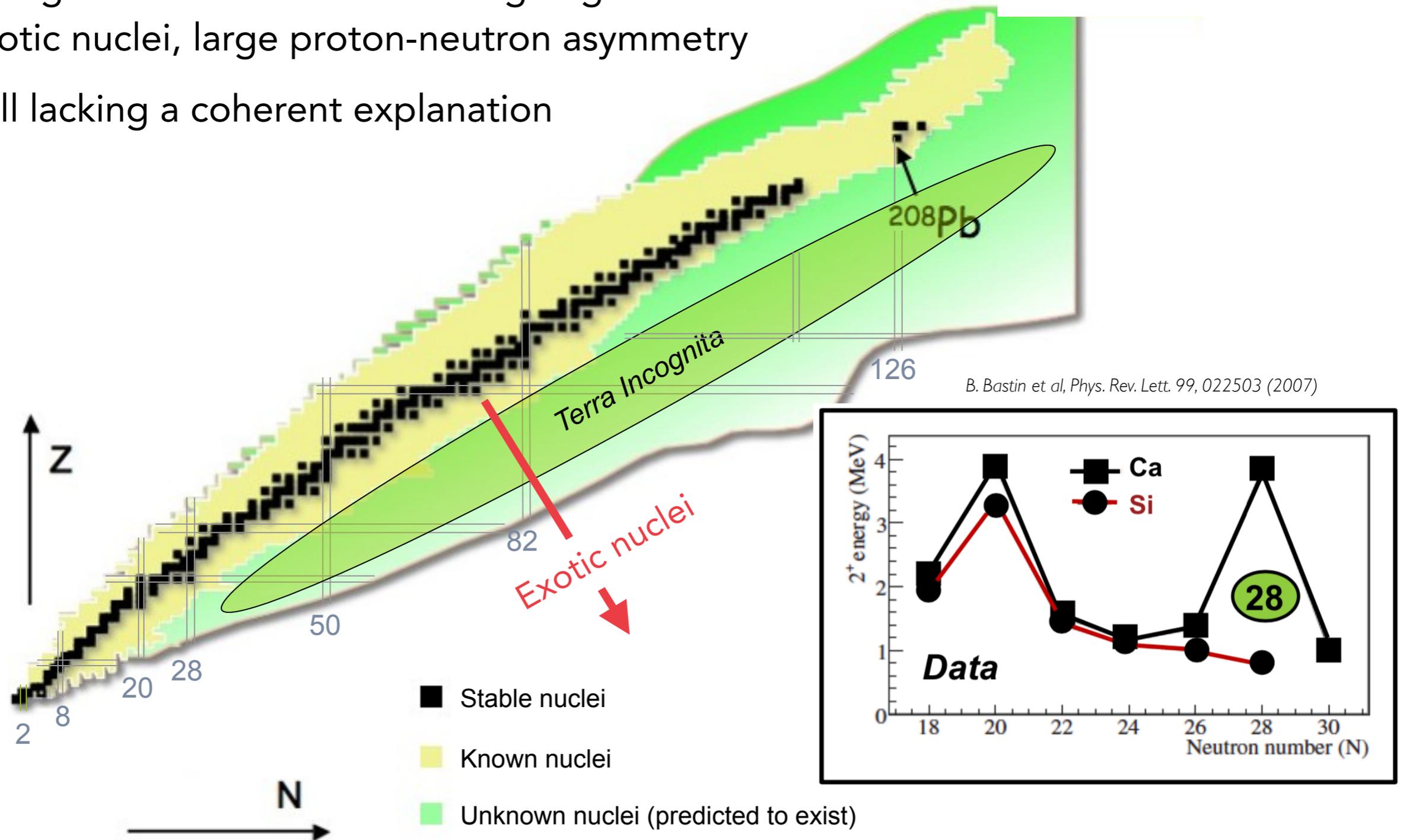


Rev. Mod. Phys. 65, 611 (1993) After Mayer and Jensen, 1955



- Basic property of finite quantum systems
- Electrons in atoms, atoms in clusters, nucleons in nuclei
- Nuclear shells evidenced in binding energies [Elasser, 1933], solar abundances [Burbidge, 1957], half-lives..
- Nuclear « magic numbers » : 2, 8, 20, 28, 50, 82, 126

- « Magic numbers » break down going to exotic nuclei, large proton-neutron asymmetry
- Still lacking a coherent explanation



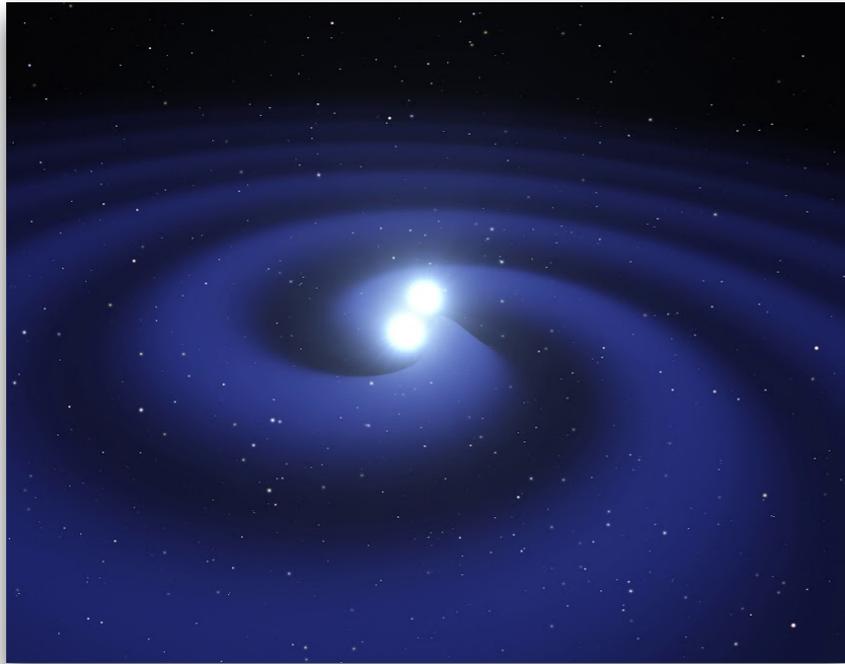
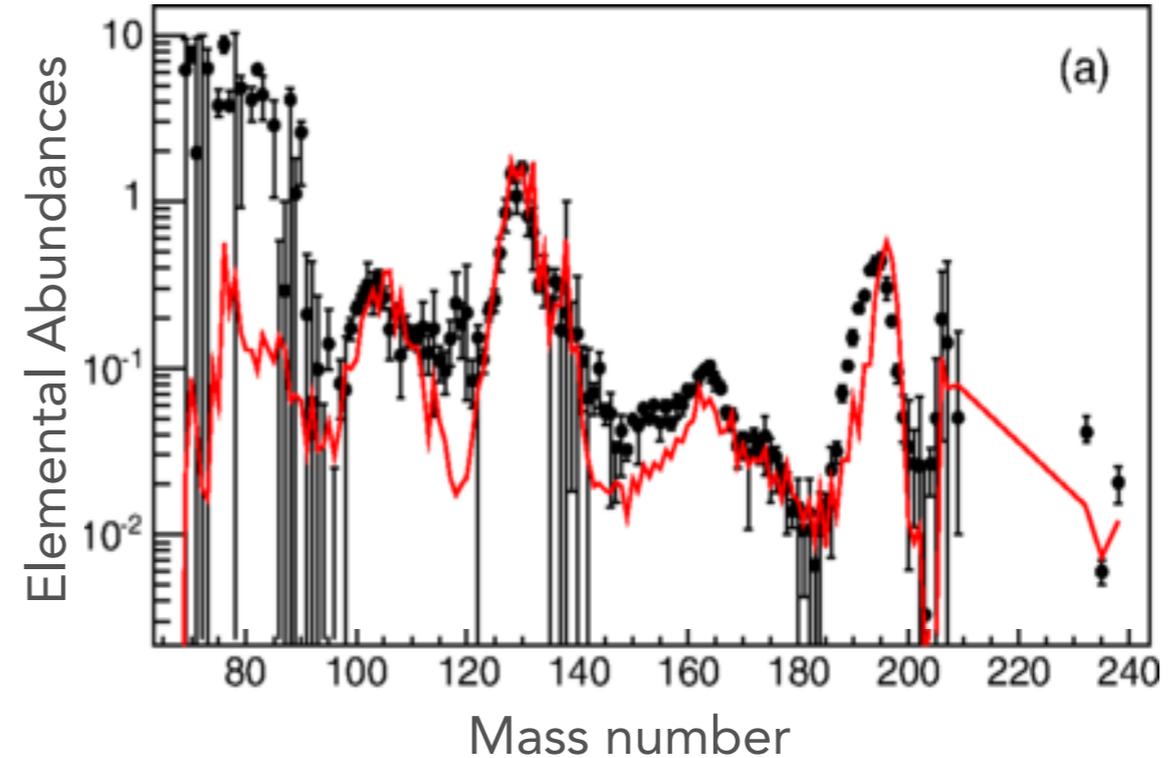
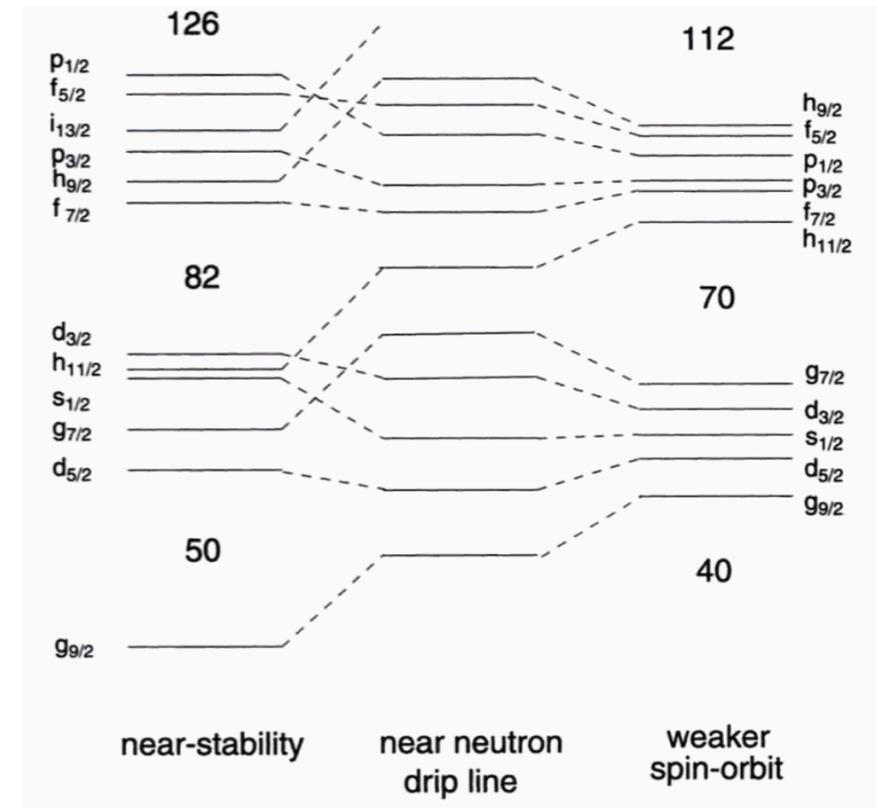
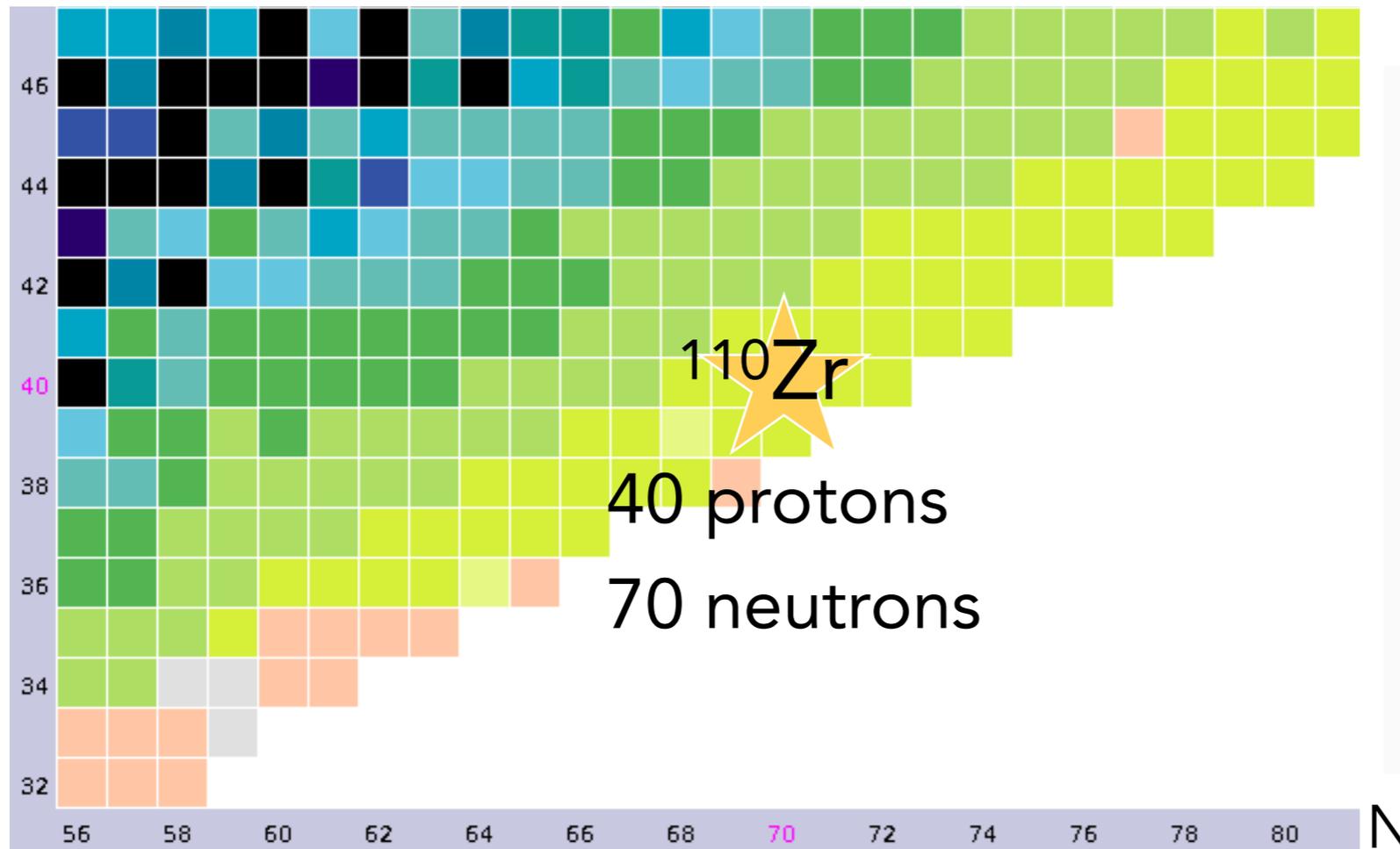


Figure from the European Space Agency

Figure adapted from Lorusso et al PRL 114, 192501 (2015)



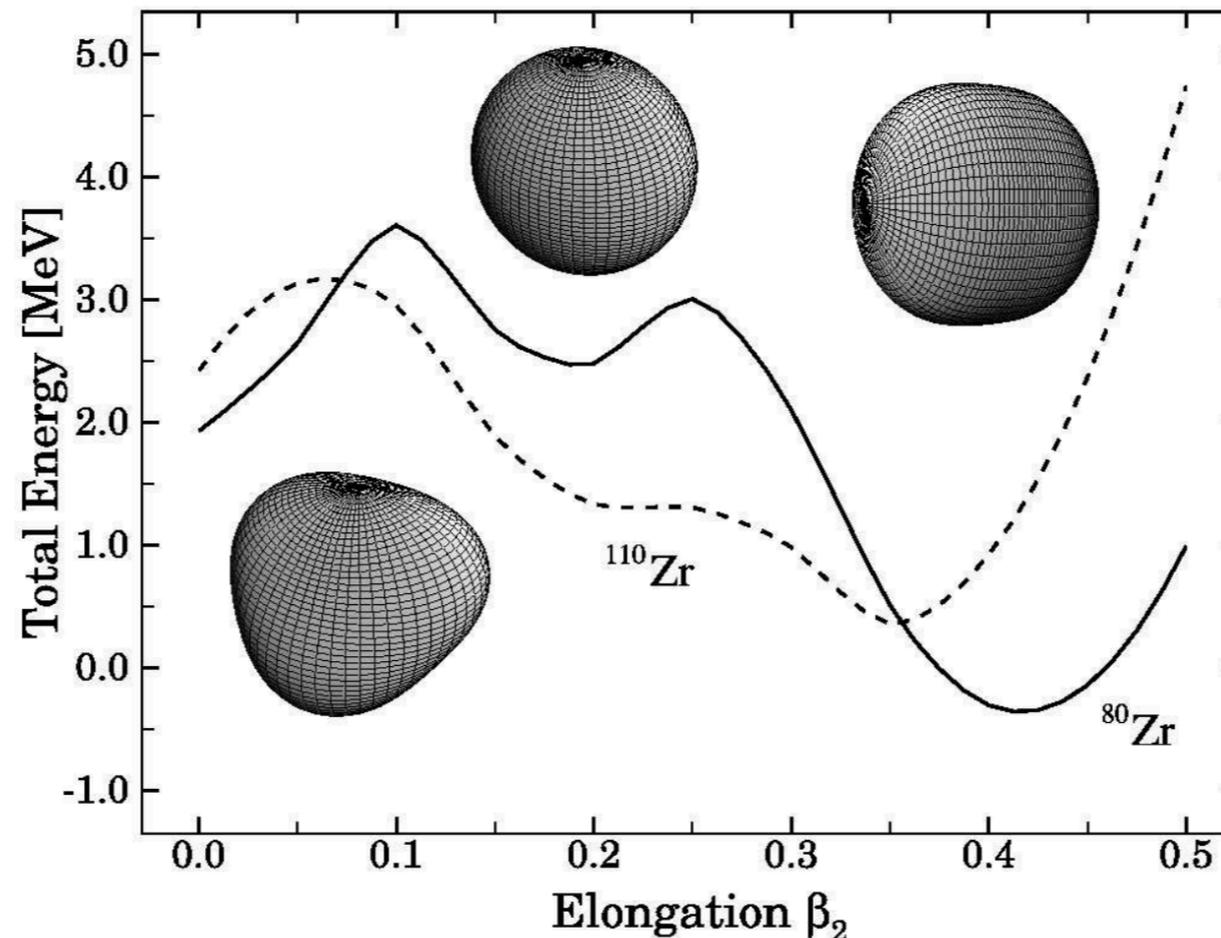
- **~50% of abundances of heavy elements (Fe-U)** thought to originate in rapid neutron capture process (r-process)
- **Astrophysical site debated** : Neutron star merger (yes, gravitational wave confirmation), Core-collapse supernova? Some other extreme environment?
- **Nuclear physics debated** : **Need to understand heavy neutron-rich nuclei !**



R. Casten, *Nuclear Structure from a Simple Perspective*, Oxford University Press (2001).

- Potential **weakening of spin-orbit coupling** in ^{110}Zr
- Could give rise to stabilisation in spherical or tetrahedral shape
 - Accumulate material near $A=110$ in the r-process, **resolve model problems?**
- Astrophysical motivation \rightarrow many theoretical predictions

Figure from Schunk and Dudek, IJMP E (2003).



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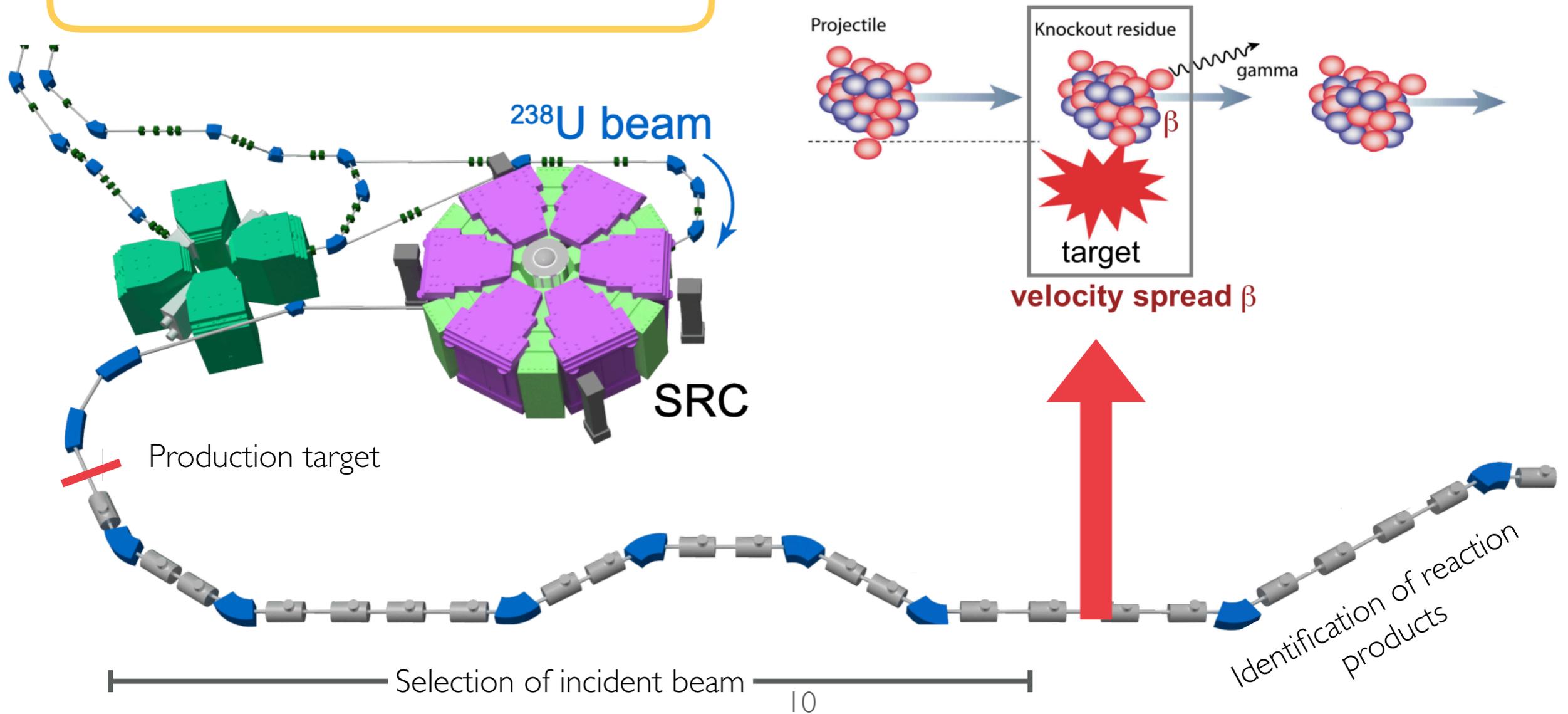
Key benchmark of exotic nuclear structure evolution

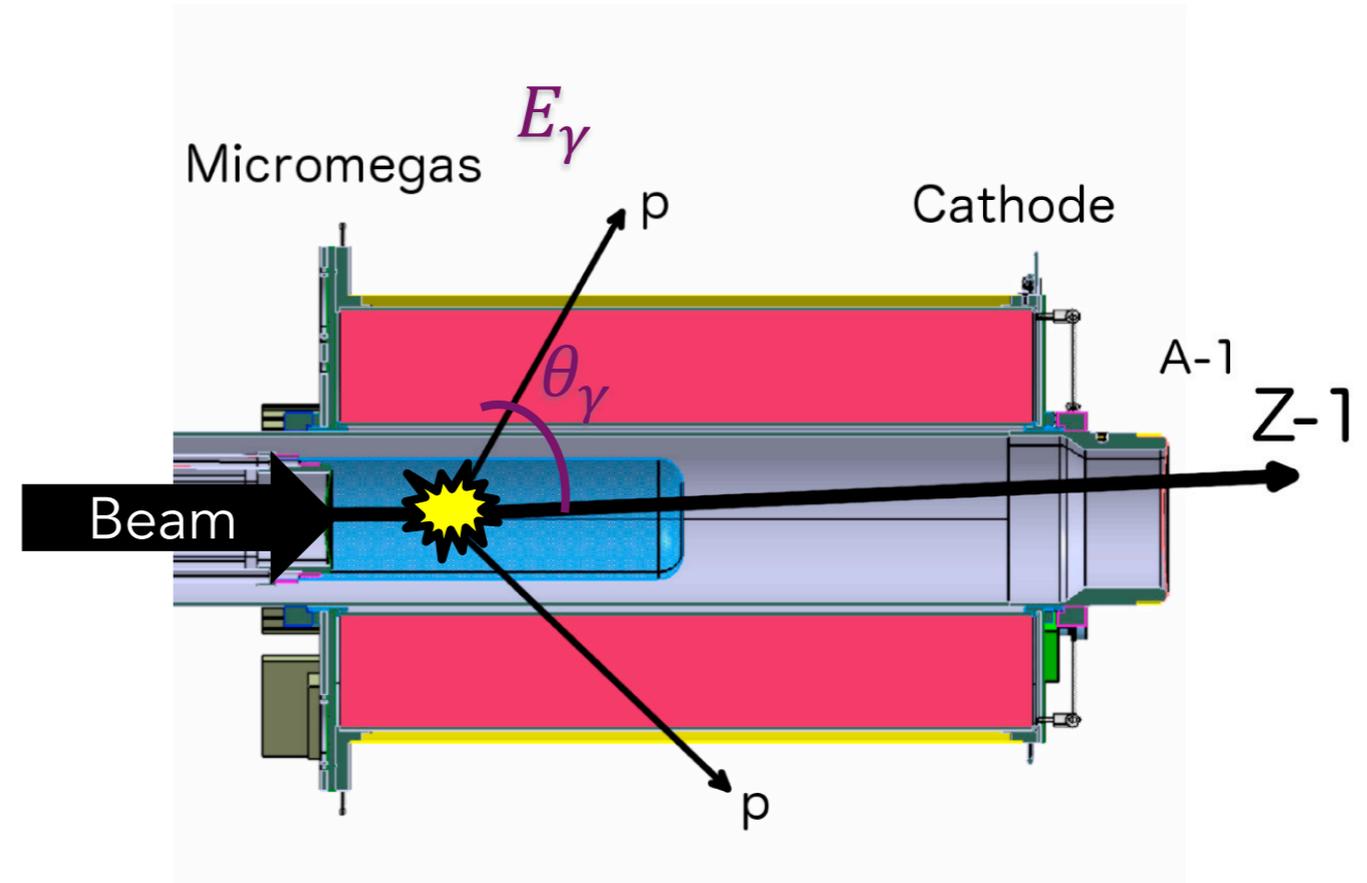
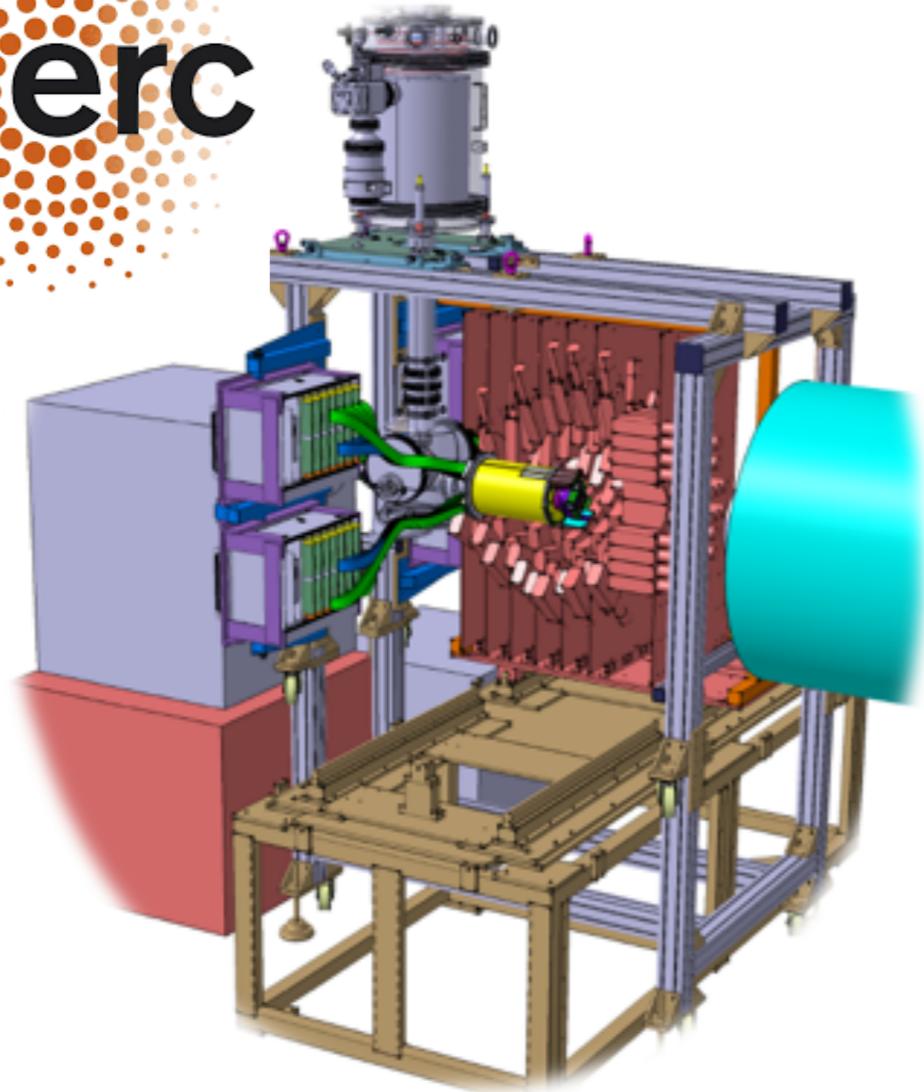
^{110}Zr produced by proton removal on a 10 cm long liquid hydrogen target
 In-beam gamma ray spectroscopy of excited nuclear states



Primary beam: ^{238}U
Energy: 345 MeV/U
Intensity: 30 pA
 β @ LH2 target: 0.6c

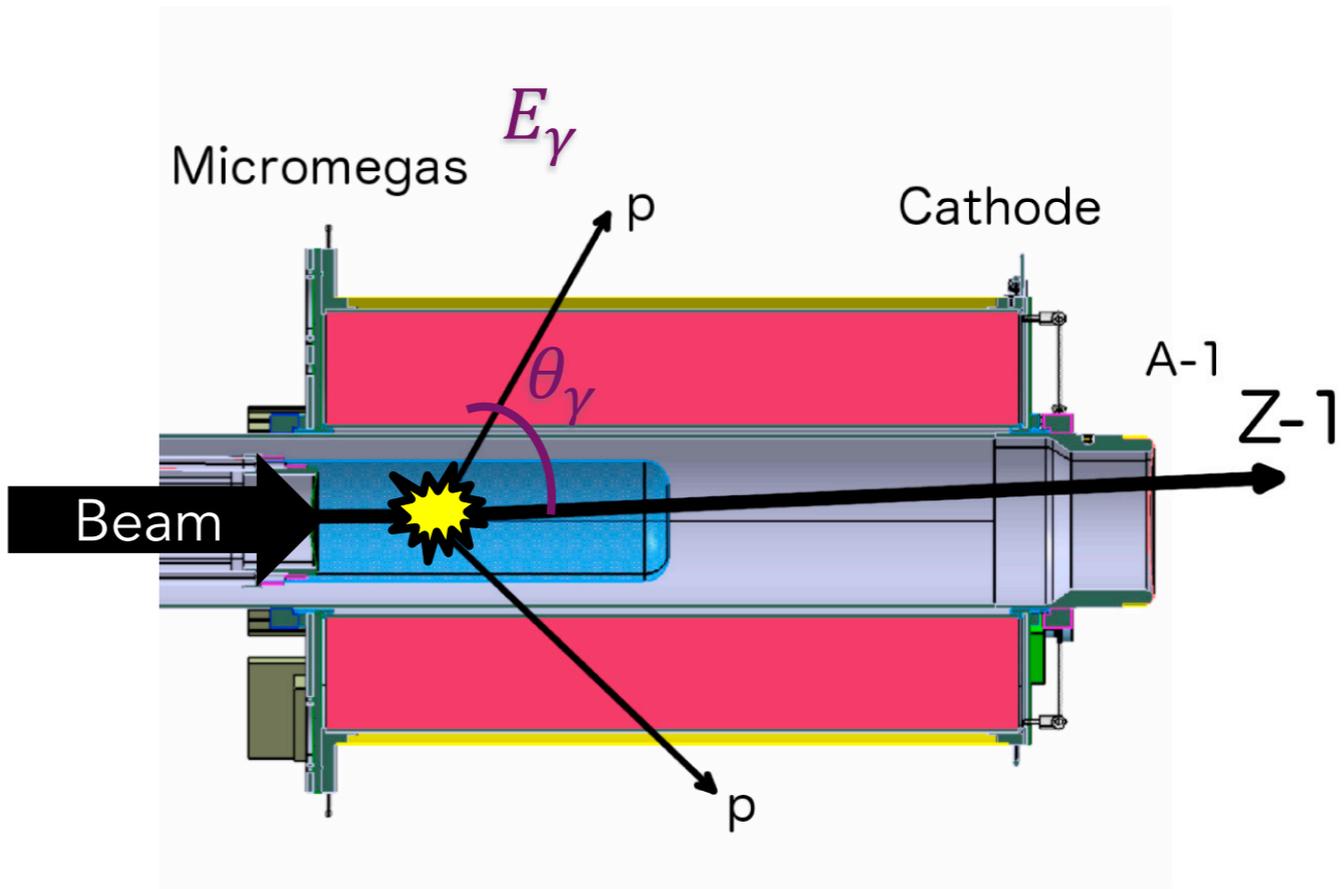
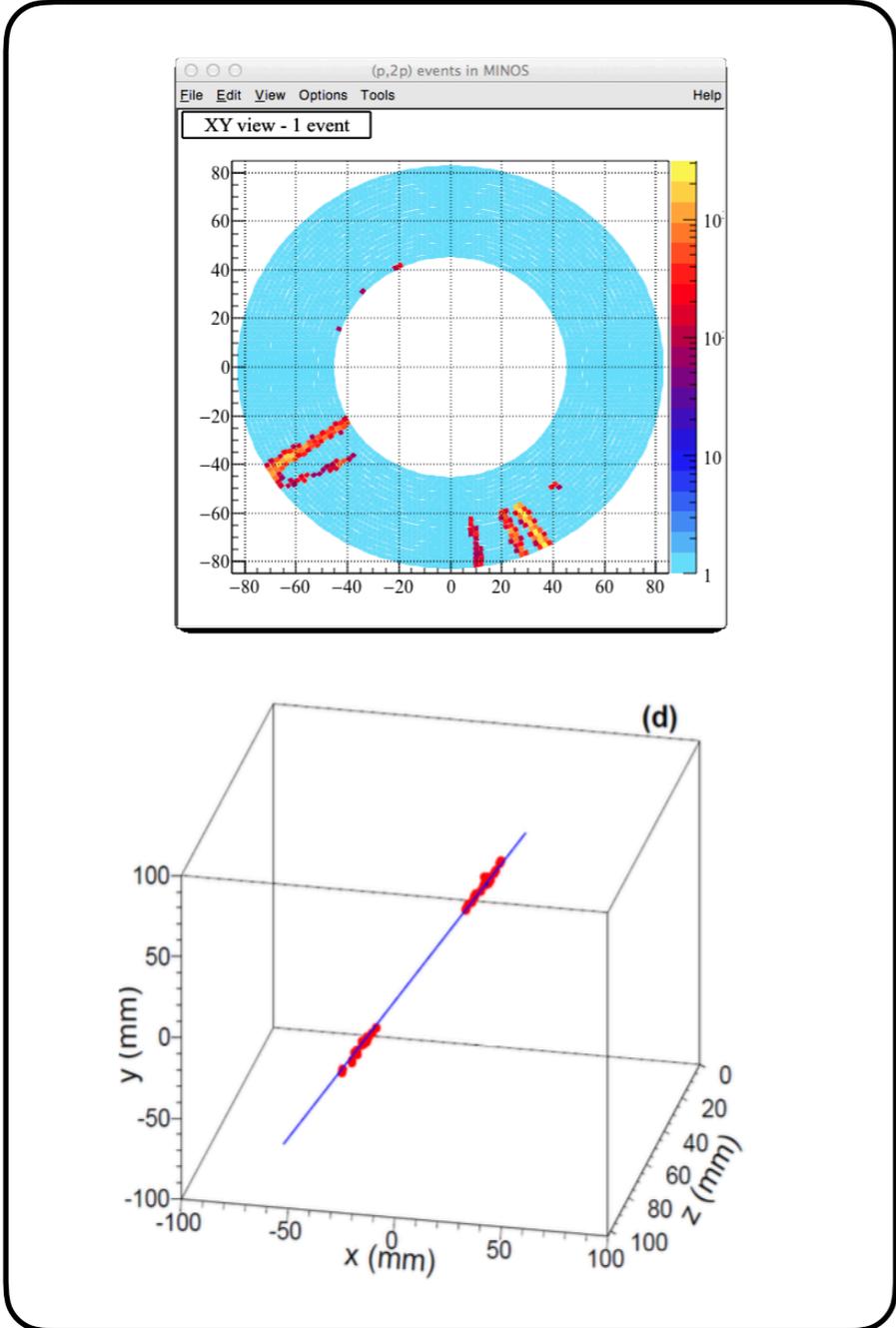
**Target thickness = tradeoff
 between luminosity and resolution**





$$E_{dop} = \frac{E_{\gamma}(1 - \beta \cos \theta_{\gamma})}{\sqrt{1 - \beta^2}}$$

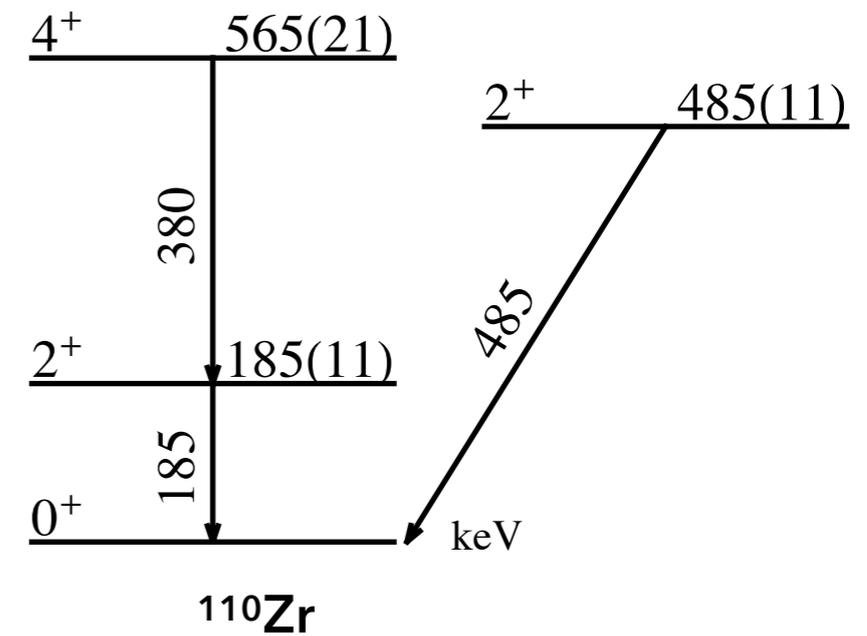
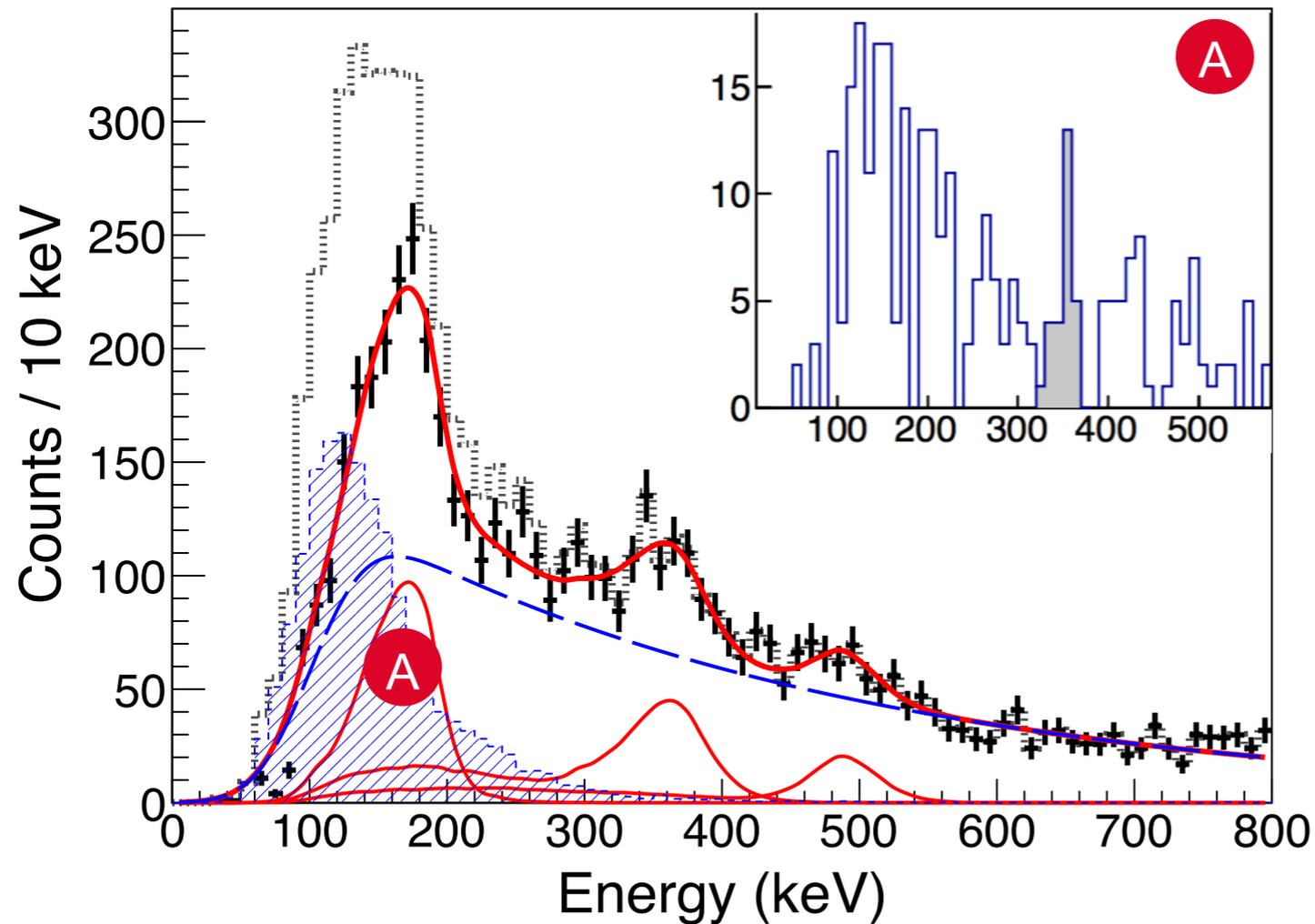
- o Track reconstruction in 3D based on combined charge and time measurements
- o Vertex position → Determine β and θ_{γ} of gamma ray detected in DALI2



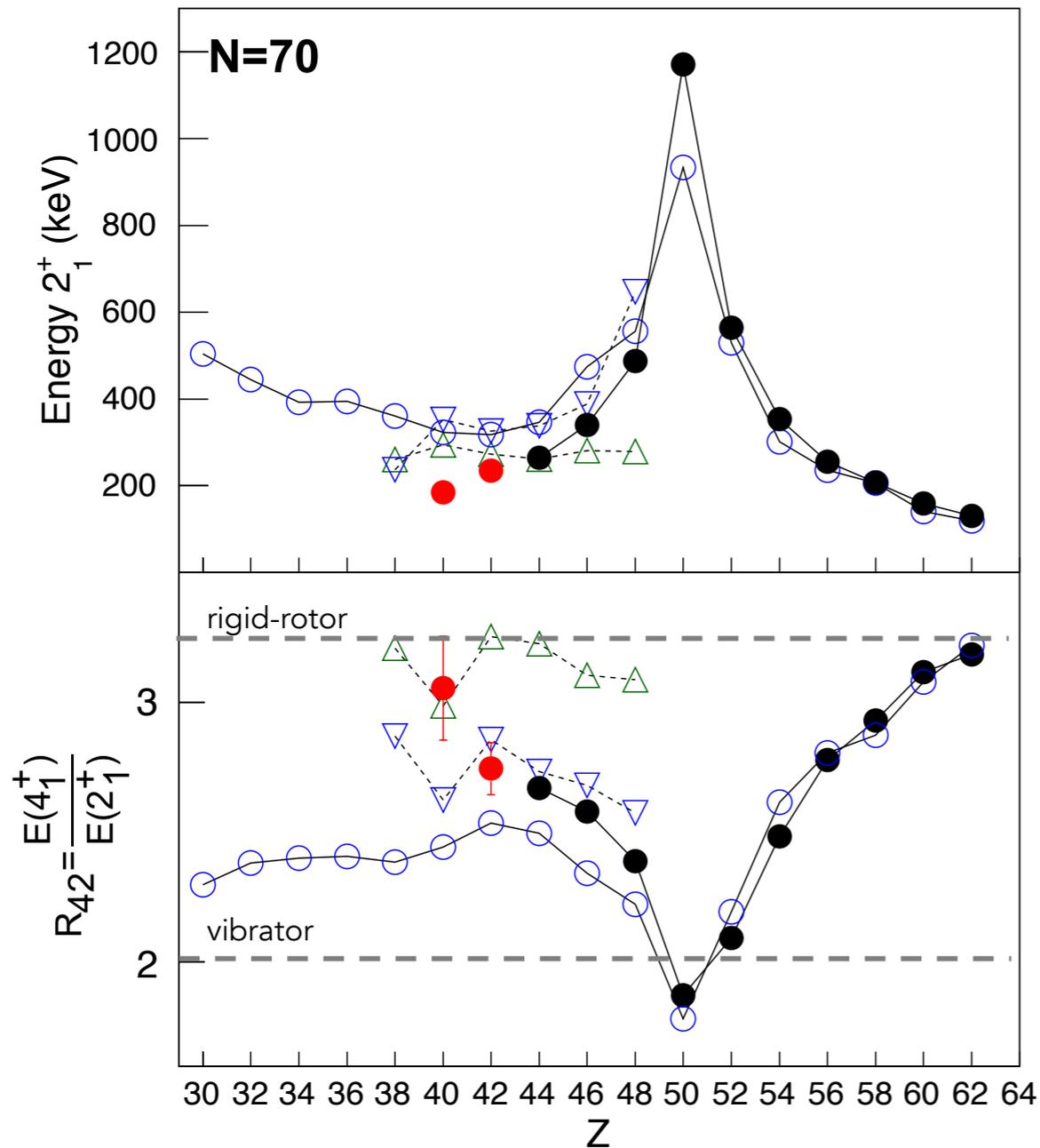
$$E_{dop} = \frac{E_{\gamma}(1 - \beta \cos \theta_{\gamma})}{\sqrt{1 - \beta^2}}$$

- o Track reconstruction in 3D based on combined charge and time measurements
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Gate 150-210 keV



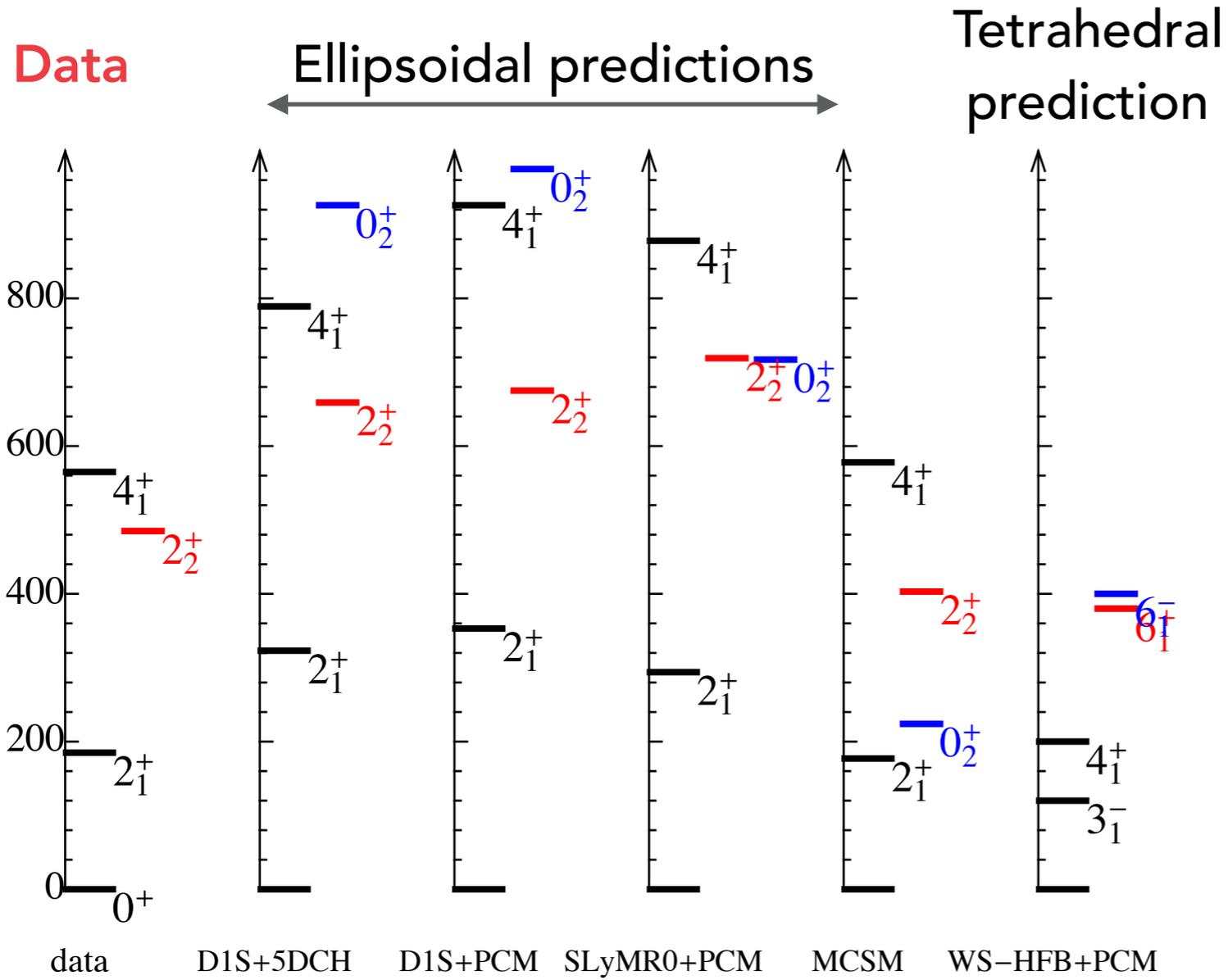
- Few stats, background subtraction essential
- Gamma-gamma coincidence allows to build level scheme

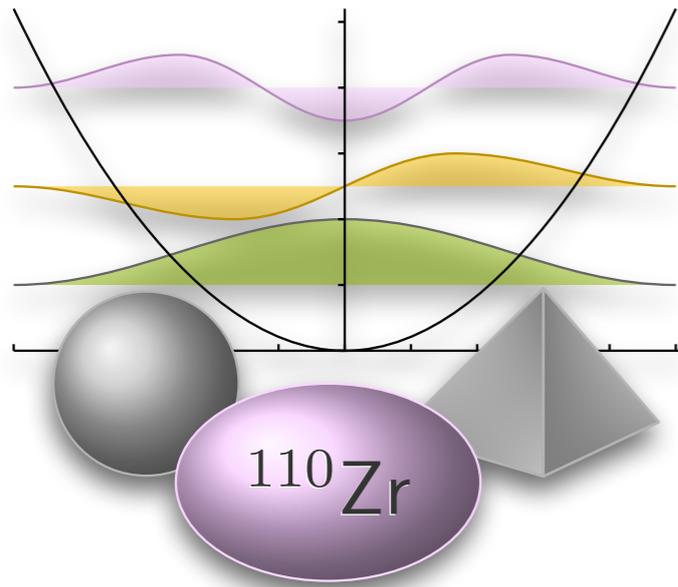


- NNDC-evaluated data
- This work
- D1S-5DCH
Delaroche, Girod, Libert (CEA)
- ▽ D1S-PCM
T. R. Rodriguez (UA Madrid)
- △ SLyMR0-PCM
Bally (CEA), Bender (IPNL),
Heenen (Université Libre de Bruxelles)

- No sign of stabilisation in first excited state systematics
- R_{42} ratio suggests well-deformed nucleus

N. Paul et al., Physical Review Letters 118, 032501 (2017)





- Spectroscopy of ^{110}Zr did not reveal a surprising quantum configuration that could solve r-process problems.
- Results show **even more deformation** than could have been expected in a « non-stabilized » case

- Demonstrates the **outstanding problem of understanding exotic quantum structure evolution**
- But there's hope ! New facilities in Europe, **FAIR** (GSI), **SPIRALII** (GANIL) and new methods (**antiprotons at CERN**) to probe the most exotic systems

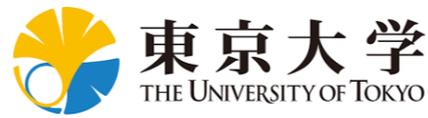


We need you !





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T. Ando, S. Momiyama, R. Taniuchi, M. Niikura, H. Sakurai, T. Saito, K. Wimmer, S. Nagamine, T. Miyazaki, S. Ota, M. Matsushita, T. Otsuka, T. Togashi, N. Shimizu



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THANK YOU



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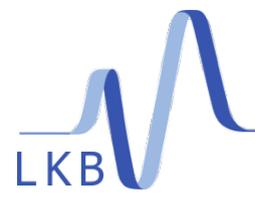
B. Ding, Z. Liu



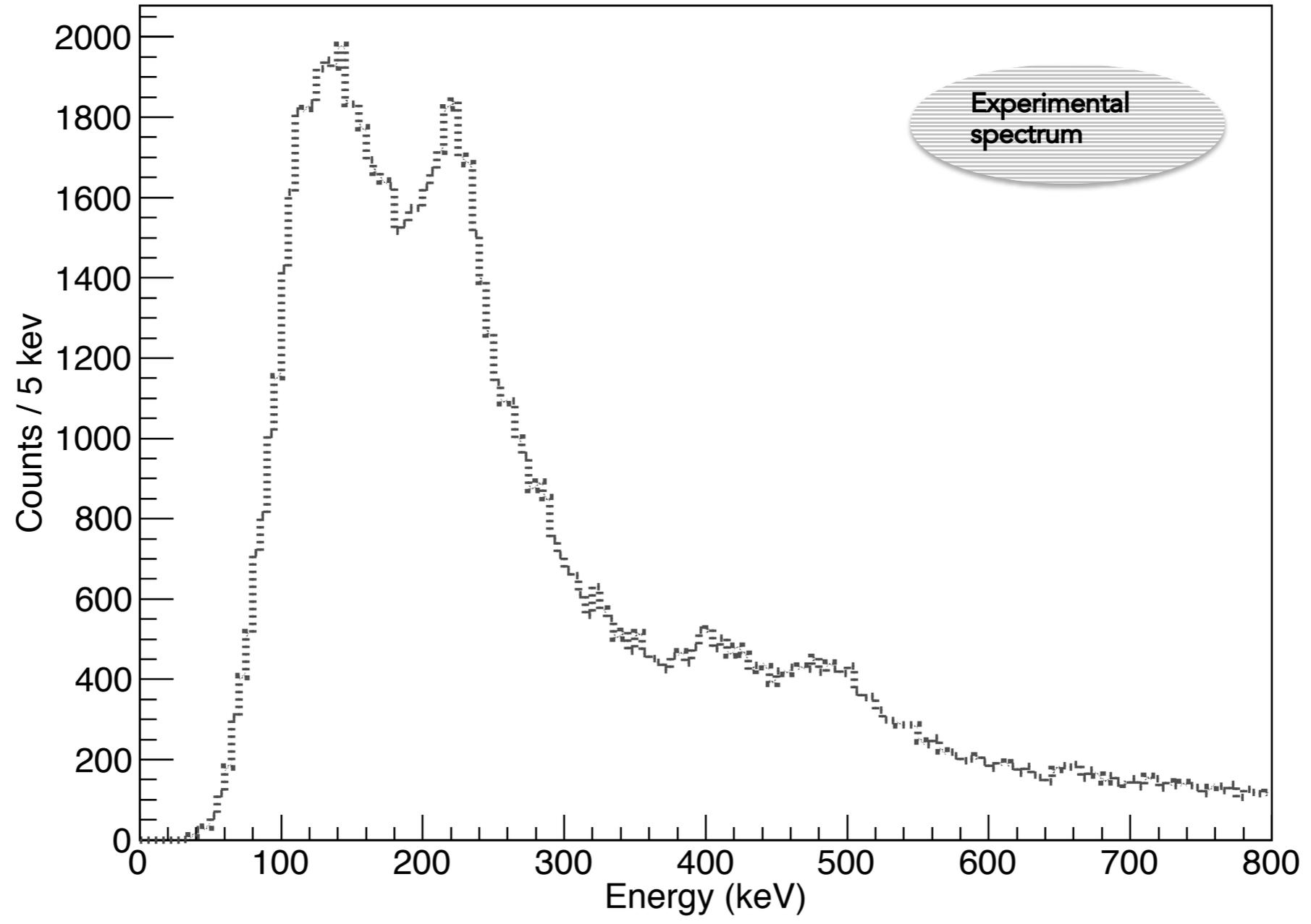
T. R. Rodriguez

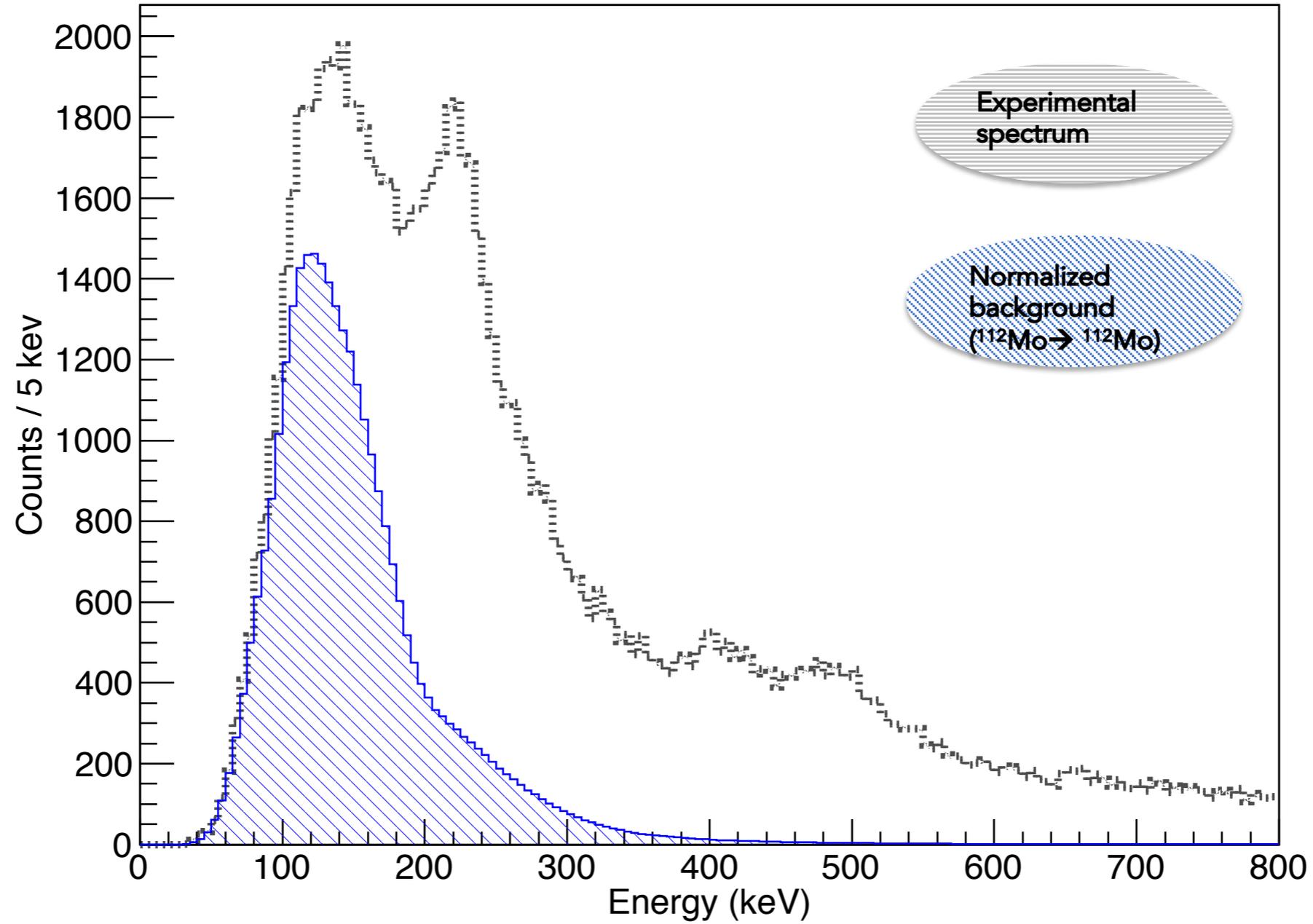


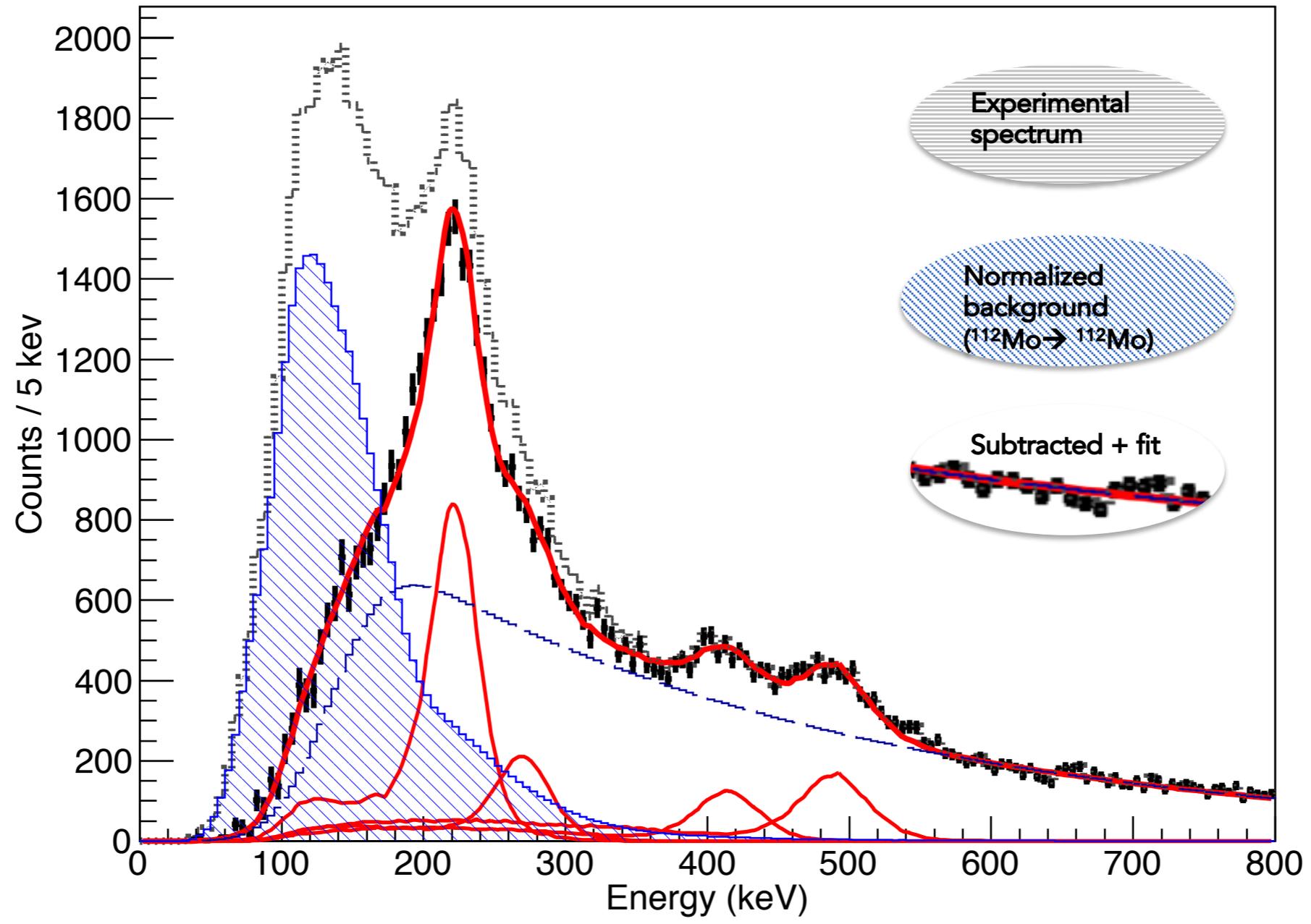
P.H. Heenen

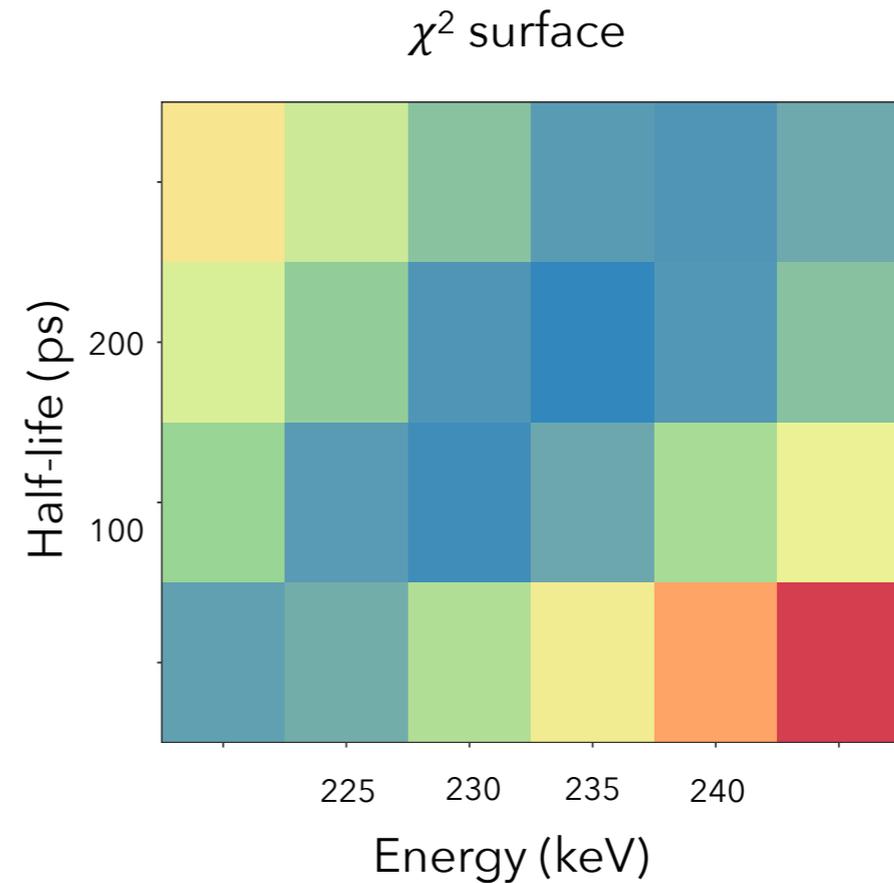
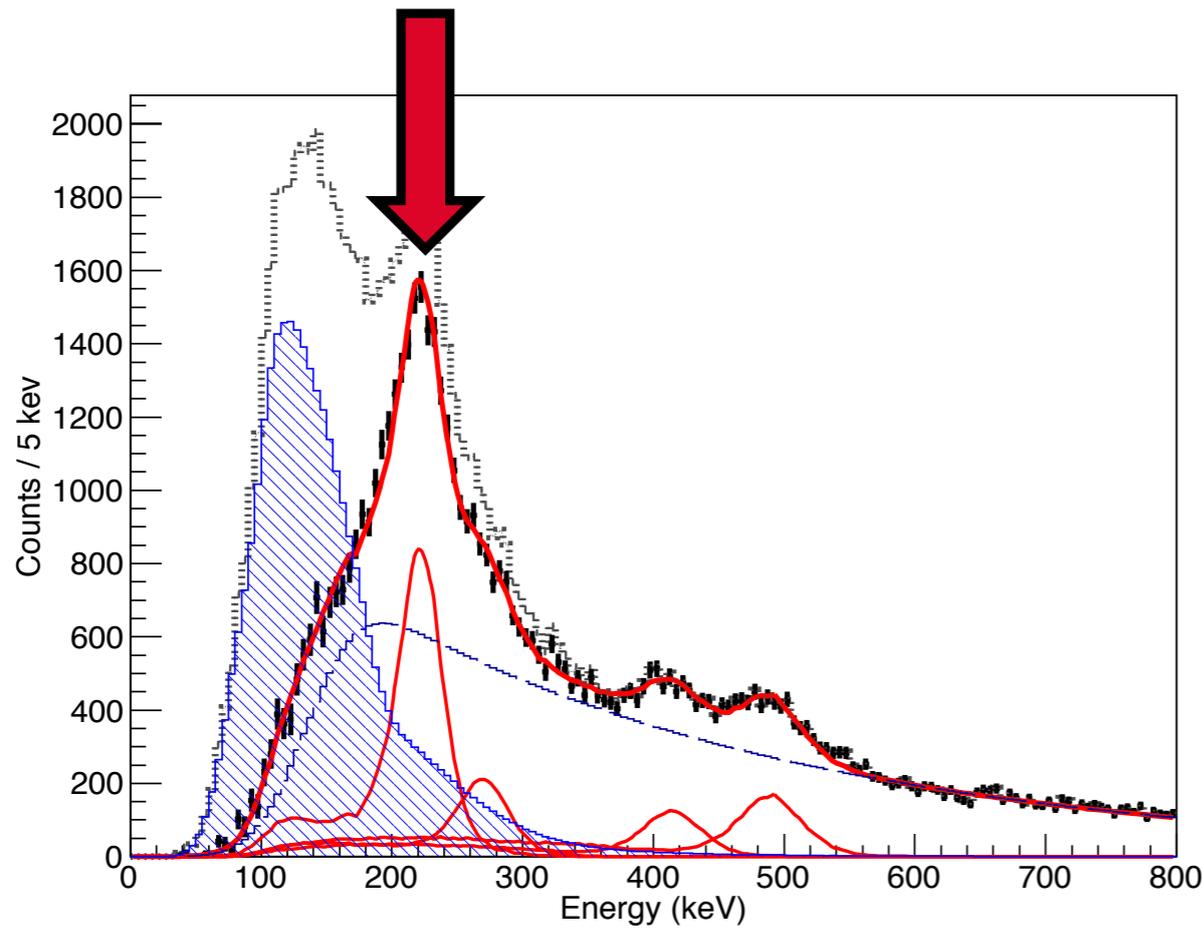


Backup









$$\begin{aligned}
 p(x, N) &= \prod_{i=1}^n \frac{1}{\sqrt{2\pi}\sigma_i} \exp\left(-\frac{1}{2} \frac{(x_i - M_i)^2}{\sigma_i^2}\right) \\
 &= \frac{1}{\sqrt{(2\pi)^n \prod_{i=1}^n \sigma_i}} \exp\left(-\frac{1}{2} \sum_{i=1}^n \frac{(x_i - M_i)^2}{\sigma_i^2}\right) \\
 &= \frac{1}{\sqrt{(2\pi)^n \prod_{i=1}^n \sigma_i}} \exp\left(-\frac{1}{2} \chi^2\right) \\
 &= C \exp\left(-\frac{1}{2} \chi^2\right)
 \end{aligned}$$

- χ^2 surfaces in the energy/half-life plane analysed to obtain best fit to spectrum
- Angular divided fits also considered, as well as feeding effects

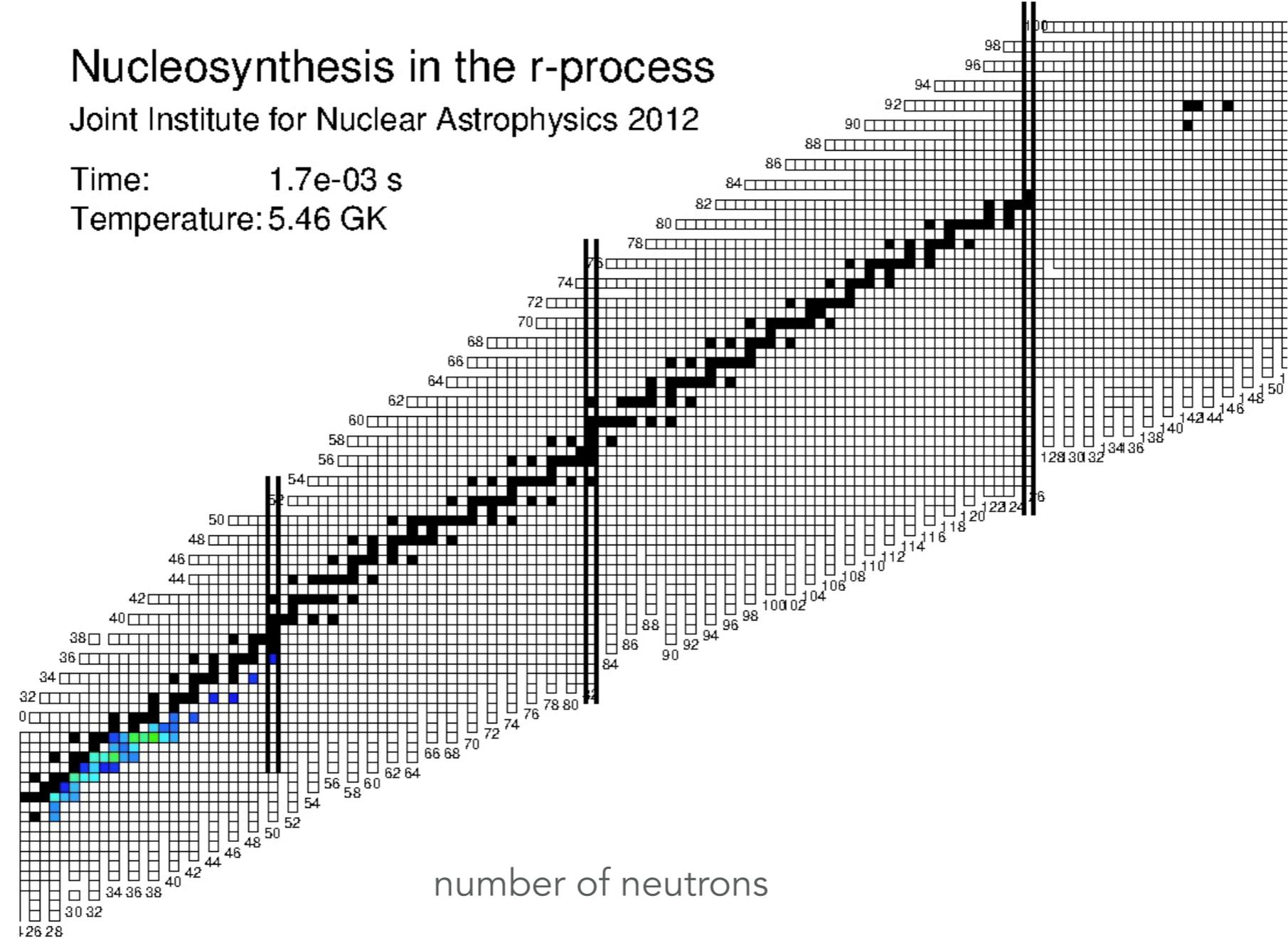
Nucleosynthesis in the r-process

Joint Institute for Nuclear Astrophysics 2012

Time: 1.7×10^{-3} s

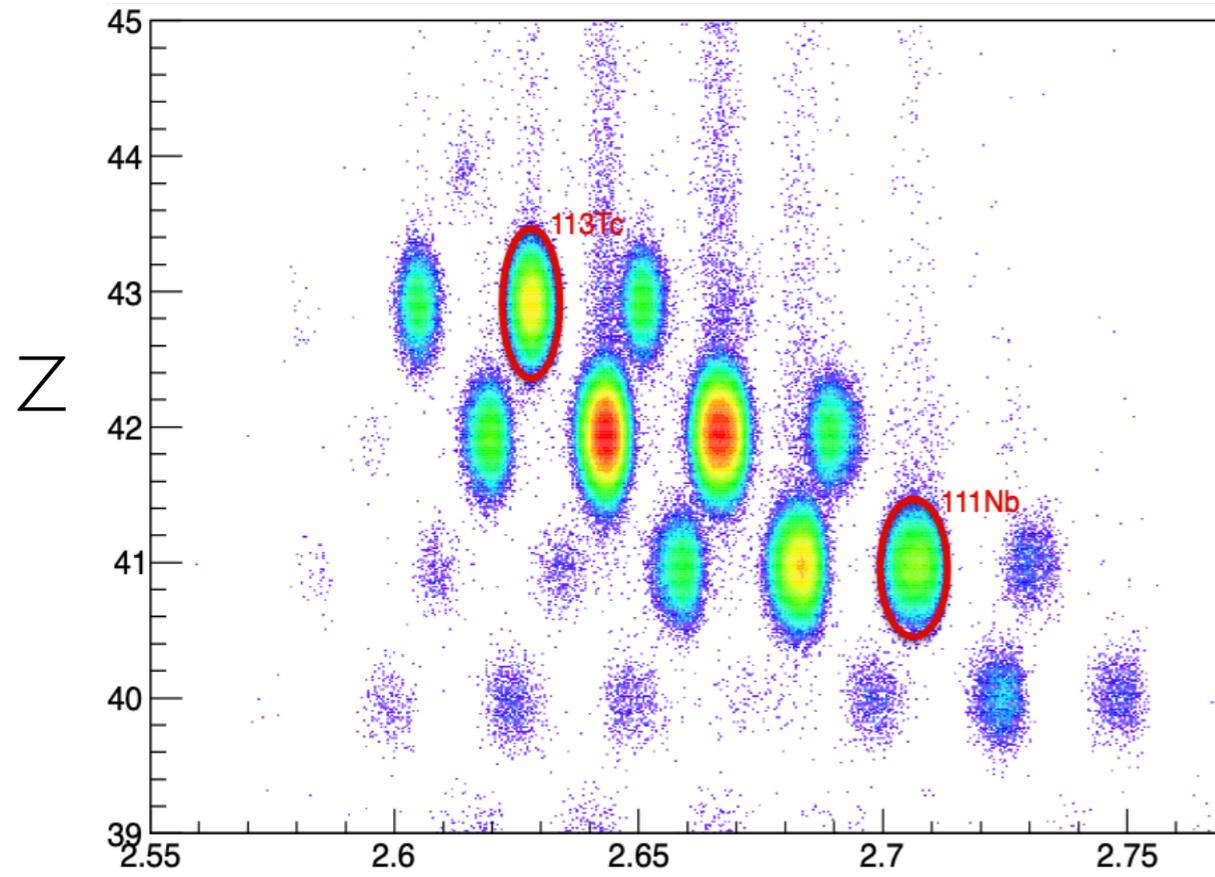
Temperature: 5.46 GK

number of protons



number of neutrons

Entrée du cible



Sortie du cible

